

PHOTOGRAPHIC SURFACE PHOTOMETRY OF GALAXIES IN THE VIRGO CLUSTER

Christopher W. Fraser

A Thesis Submitted for the Degree of PhD
at the
University of St Andrews



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IN THE VIRGO CLUSTER

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DECLARATION

I hereby declare that the following Thesis is based on the results of experiments carried out by me, that the Thesis is my own composition, and that it has not previously been presented for a Higher Degree. The research was carried out at the University Observatory, St Andrews.

CERTIFICATE

I certify that Christopher W. Fraser has spent nine terms at Research Work at the University Observatory, St Andrews, that he has fulfilled the conditions of Ordinance No. 16 (St Andrews) and that he is qualified to submit the accompanying Thesis in application for the Degree of Ph.D.

PREFACE

Surface photometry is one area of extragalactic studies in which information is urgently needed by optical as well as radio astronomers. The aim of this thesis is to supply some of the photometric parameters for 48 galaxies in two international colour systems. The reader will notice that descriptions are included of many detailed aspects of the work and little attention is paid to the more basic problems. This is because the discussion of the basic problems in photographic photometry of galaxies formed the contents of an M.Sc. thesis.

Volume one is devoted to a detailed description of the methods used for the reduction techniques together with a discussion of the results, while volumes two and three contain the photometric data. In this volume, the first chapter is concerned with previous investigations and the following chapter contains detailed descriptions of the instruments and methods used in the preliminary plate reductions. The subject of isophotometry is dealt with in the third chapter, while in the fourth and fifth chapters respectively, mention is made of the reduction techniques, together with conclusions which may be deduced from the photometric data. This information is

(ii)

given in volume two for those galaxies with NGC Catalogue numbers from 4189 to 4459, while the data for objects from NGC 4461 to NGC 4762 are listed in the third volume of the thesis.

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Chapter I

1.1 Introduction. The aim of the first chapter of this thesis is to make a critical examination of the observational techniques and reduction methods adopted by other authors whose work was closely related to the current investigation on surface photometry of galaxies. No attempt has been made to include every published work on the subject in this appraisal and therefore it is only to sections of those papers which are particularly relevant to the present work that minor criticisms will be given. However, in view of the special attention devoted to the sources of error which will be revealed throughout this study of the development of surface photometry, it should become apparent in succeeding chapters that the elimination of such errors in the present investigation ought to enable results of high photometric accuracy to be achieved.

1.2 Review. Following the pioneer investigations of Reynolds (1919), in which he found that the luminosity along the major axis of M31, could be represented by the expression

$$\text{Luminosity} = \frac{\text{Constant}}{(x + 1)^{\frac{1}{2}}}$$

for radial values of x less than $7'$ from the nucleus, the first investigator to obtain detailed information on galaxies was Hubble (1930). He obtained luminosity curves for fifteen elliptical galaxies and derived analytical expressions to represent the distribution of luminosity within the objects. The account which was given not only of the observational details but also the reduction techniques contained considerably more detail than is found in some of the more recent papers. In fact, modern papers on galaxy photometry now tend to omit certain details which, although clearly appreciated by the authors involved, could highlight and illustrate by their inclusion some of the more serious difficulties of the subject to readers who are not particularly well versed in this specialized field. However, it is as a result of the inclusion of some details that two major criticisms, first suggested by de Vaucouleurs (1948a), may be put forward.

The first concerns the observational procedure, in which a series of exposures usually ranging from 15 minutes to 5 seconds was made on one photographic plate at the Newtonian focus of the 100-inch Mount Wilson telescope. Let us suppose that five exposures, t_1 to t_5 , were made for one galaxy. Under these conditions the first galaxy exposure is made up of three components:

a zero pre-exposure to the sky, an exposure t_1 to the sky plus galaxy, and a post-exposure $t_2 + t_3 + t_4 + t_5$ to the sky. The second galaxy exposure has the respective values t_1 , t_2 and $t_3 + t_4 + t_5$ for the sky pre-exposure, the sky + galaxy exposure and the sky post-exposure. It becomes clear that the five galaxy exposures were made under different conditions. Consequently, microphotometer tracings across the image of the galaxy would not give an accurate representation of the luminosity distribution, since pre- and post-exposures produce changes in the latent images which in turn affect the response of the photographic plate, particularly when the plate has been exposed to very low levels of illumination.

Secondly, errors were introduced in the calibration of the plates. Before the actual calibration exposure, Hubble first subjected the plates to an exposure which would "simulate the sky-fog on the nebular plate". In the case of the first galaxy exposure, this system would have been acceptable had no errors been introduced by this method of pre-exposure, but the sky-fog exposure prior to calibration was always a pre-exposure, while in the actual observations, it was solely a pre-exposure on only one occasion out of five.

Nevertheless, Hubble's work was still of a sufficiently high standard to enable him to re-introduce Reynold's law (1913) for the brightness B of a point at a distance r from the centre of an elliptical galaxy in the form

$$B = B_0 / (r + a)^2$$

where B_0 is the central brightness and a is a scale parameter varying from one galaxy to another.

Redman was working on a similar problem in 1931, although his first paper on the subject was not published until 1936. He proposed and used two methods for the photometry: the first required one exposure to the galaxy followed by an extrafocal exposure to standard stars, while the second method, the better from the point of view of photometric accuracy due to the difficulties encountered with the former in detecting the outer regions of a galaxy, needed a set of exposures of the galaxy in order to record more adequately the very considerable intensity range from the centre to the faint outer regions. The main criticism of this work was that the calibration spots were superimposed on the sky-fog. Redman and Shirley (1938) thought that as both the calibration and galaxy exposures were superimposed on the sky fog, they claimed that this sky fog could then be omitted in subsequent measurements.

Redman and Shirley's observing and calibration conditions were similar to the first exposure of the set taken by Hubble in his work, where the sky-fog was only a pre-exposure just the same as the calibration pre-exposure. However, what both Redman and Hubble failed to recognize was that the pre-exposure of the sky-fog on the calibration areas would affect the shape of the resulting characteristic curve.

In 1948 de Vaucouleurs made a major contribution to the subject with papers (1948, 1948a) which included a very extensive and quantitative discussion of the types of errors which are expected in galaxy photometry, together with the measurements made of three galaxies. The results which had been obtained hitherto by other observers for the total magnitudes and luminosity profiles of several galaxies did not appear to be closely inter-related. This was because of difficulties with

- a) accurate calibration techniques
- b) determination of the precise level of the sky background.

de Vaucouleurs' values of magnitudes and profiles were supplemented by several other parameters, in particular the effective dimensions and effective luminosity profiles. The advantage of these 'effective' values was that

they gave intrinsic characteristics of the galaxy which were independent of the resolving power of the telescope used for the observations. These values were precise so long as the effective radius of the galaxy was considerably greater than the effective radius of the instrumental scattering function as determined from stellar images. Elliptical isophotes were assumed after microphotometer tracings across the main axes of the object had been taken; these isophotes were used in the calculation of the total luminosity of the system. The total integrated luminosity emitted between the centre of a galaxy with circular symmetry and an isophote of distance r from the centre is $L_r = 2\pi \int_0^r I(r) r dr$ and the total luminosity is $L_T = 2\pi \int_0^\infty I(r) r dr$. The fraction of the total luminosity emitted between the curve and r is defined by $k(r) = L_r/L_T$. For elliptical objects, both Redman and de Vaucouleurs obtained the integrated luminosity by integrations which were based upon the surface luminosity distribution along the major and minor axes of these objects. The function $k(r)$ defines the relative integrated luminosity curve as a function of r , and the effective radius r_e , introduced by de Vaucouleurs (1948), is defined so that $k(r) = \frac{1}{2}$. i.e. one half of the total luminosity of the system is

emitted within this distance from the centre.
de Vaucouleurs also deduced a relationship for elliptical galaxies for which the intensity at a distance r from the centre was proportional to the quarter power of that distance, namely

$$\log J = -3.33(\rho^{\frac{1}{4}} - 1) \text{ where}$$

$J = I/I_e$ and $\rho = r/r_e$, I_e being the intensity corresponding to r_e . This formula has the advantage over Hubble's relation that neither the integrated brightness nor the central intensity is infinite.

Bigay (1951) published work on galaxies and began his paper with a review of the four methods used prior to 1951 for photometry of galaxies. They were

- 1) direct comparison between the galaxy and standard stars on the same plates obtained on telescopes with short focal lengths
- 2) comparisons between stars and galaxy images on plates taken out of focus
- 3) photo-electric aperture measurements and
- 4) magnitudes deduced from photometric profiles.

After describing these methods in detail, Bigay gave the results of measurements of the magnitudes of 175 bright galaxies by a method which was novel at the

time, viz. Fabry photometry; this work dealt with integrated rather than surface photometry. Plotting the magnitudes which he obtained against other published magnitudes, as a test of the validity and accuracy of his results, he found that the correlation exhibited a scatter of one and a half magnitudes on the average.

Evans (1951, 1952) investigated some southern elliptical nebulae in 1951 by studying isophotes which had been assembled from microphotometer tracings. A large section of the work was devoted to the determination of the size of the errors which could occur if the estimate of the level of the sky background both on the galaxy profile and on the characteristic curve was not known with sufficient accuracy. It was found that for an error of 0.02^m in the determination of the magnitude of the sky background, the subsequent error in the calculated magnitude of the galaxy could be as large as 0.15^m . Evans noted that in NGC 1291, there appeared to be a steady rotation of the major axes of successive isophotes; this effect was particularly evident on two plates microphotometered at 90° to each other. Evans highlighted certain points which are not detailed by other writers such as the quantitative determination of sky background errors,

while certain points which should have been included to retain the same consistency for detail, such as the design of a tube sensitometer, were noticeably absent. The writer has been informed on several occasions by Evans that his papers were often shortened at the suggestion of referees, and that but for these suggestions, he would have included a detailed description of a tube sensitometer.

In 1953 de Vaucouleurs published work on the distribution of mass and luminosity in elliptical galaxies. The methods used to obtain the luminosity distribution were similar to those used in 1948 and the results also confirmed the linear variation of intensity with the quarter power of distance.

The first photo-electric work on the subject came from Stebbins and Whitford (1937), who measured magnitudes of 175 galaxies brighter than fifteenth magnitude. These magnitudes were obtained using a Kunz photo-cell attached to the 60- or 100-inch Mount Wilson reflectors, together with four focal plane diaphragms giving field diameters of 42, 64, 128 and 163 arc sec on the 100-inch. These values were increased by a factor of 1.7 when used on the 60-inch telescope. When this work was carried out, the largest

aperture was considered sufficiently large to allow the total light of the majority of the chosen galaxies to be measured, this judgement being based upon the accepted measured dimensions of the objects at that time. However, it is now known from data on micro-photometer tracings of plates of galaxies obtained with modern emulsions (e.g. van den Bergh 1969) that the outer regions of galaxies extend considerably further than earlier dimensions had indicated. Consequently, the early work of Stebbins and Whitford would suffer from incomplete measurements of the total magnitudes for many of the systems. Aware of this difficulty, Stebbins and Whitford (1952) published a further set of 176 magnitudes of galaxies in the range $9.2 \leq m \leq 18.2$. This later work improved on the former in two ways, viz., the use of a 1P21 photomultiplier in place of the photo-cell, and the availability of a field diaphragm of 5.1 arc min on the 100-inch, or 8.7 arc min on the 60-inch telescope. However, even with this extra diaphragm, the possibility was still present of missing the faint outer extensions of giant elliptical systems.

Photo-electric scans were first attempted by van Houten, Oort and Hiltner in 1948, although their results were not published until 1954. A diaphragm could be

driven by a synchronous motor in an arbitrary direction in the Cassegrain focal plane of the McDonald 82-inch telescope, and tracings were made along the principal axes of the galaxy. The amplifier sensitivity was varied by the use of different grid leaks, but the determination of the relative sensitivities of these grid leaks proved difficult, due to the variation of their ratios with temperature and humidity. Furthermore, although the dark current was measured before and after each tracing, the value of the night sky for each galaxy was determined only once for each night. These difficulties lead to considerable errors in the results, and for this reason, the work could only be regarded as unsuccessful.

Further photo-electric work was published by Pettit (1954), who determined the magnitudes on the photographic scale and also the colour indices of galaxies with the 60- and 100-inch Mount Wilson telescopes between 1947 and 1954. The telescope observations were made using a refrigerated 1P21 photomultiplier, Schott BG 1 and GG 7 filters were used for the blue and yellow readings and the field apertures ranged in diameter from 0.2 to 8.6 arc min. The observational programme in 1947 and 1948 was confined to the 60-inch

telescope with the following two aims in view:

- 1) to measure galaxies of large angular diameters and
- 2) to provide standards on the international system for use on the 100-inch telescope where the polar calibration sequence was not available.

For the latter purpose, the central stars of Selected Areas in the 30° zone which passed near the zenith were chosen, enabling the extinction and reddening functions for each night to be obtained. A star was measured from a low altitude to the meridian on four occasions each night, to determine the colour equation and atmospheric extinction. The measurements of both the galaxy and the comparison star were reduced to the zenith before converting to magnitude. The magnitudes which were obtained depended on the free deflection minus the yellow (F-Y), the reduction to the zenith was made by plotting $\log (F-Y)$ against $\sec z$ for the Selected Area and the constant a determined in the extinction formula $C = a(\sec z - 1)$. Normally, the numerical value of a was approximately equal to 0.1, and when this value was substituted in the above equation, it gave a correction to C which in turn, when added to the observed $\log (F-Y)$, reduced all the deflections for the night to the zenith.

The next step in the magnitude determination involved the calibration of the instrument for the night. Since most of the objects studied were fainter than ninth magnitude, it was this value which was chosen as the basic brightness for the calibrations; this meant that $\log (F-Y)$ was determined for a ninth magnitude star at the zenith for the night. If North Polar Sequence stars were used, these would also require to be reduced to the zenith using the extinction formula, since their magnitudes were given for the zenith at Mount Wilson. The $\log (F-Y)$ for each such star, if it had been ninth magnitude, was obtained using the relation

$$\log (F-Y)_9 = \log (F-Y) + 0.4(m_{pg} - 9)$$

and so the magnitude for any object was given by the expression

$$m_{pg} = 10/4 \left[\log (F-Y)_9 - \log (F-Y) \right] + 9$$

The detailed description of this aspect of the work has been given because it highlights the contrast between some of the techniques used to obtain reasonably consistent photo-electric measurements in comparison with the more involved methods necessary

for successful photographic work, which will be described in the course of this thesis.

The accuracy of the results was checked by two methods. First, the internal agreement of repeated measurements which gave average departures of 0.07^m and secondly, agreement with other observers' measures. A comparison with the Shapley-Ames Catalogue showed that when the magnitudes were divided into four groups, the mean departures per group ranged from -0.3^m to $+0.1^m$, while comparison with Bigay's measurements showed a similar range from -0.10^m to $+0.13^m$. Such comparisons are not really justified, because the inherent difficulties of centring the galaxy in the apertures and of ensuring that the apertures are sufficiently large to allow the total luminosity of the galaxy to be measured, mean that a comparison is being made between an unknown luminosity measured photo-electrically, depending on the angular size and central brightness of an object, and a luminosity governed only by the limit to which measurements may be made above the threshold brightness on a photographic plate. Furthermore, Pettit does not state clearly how often or on what location measurements are made of the sky background. It is well known that where the galaxy and sky are

measured together, the sky can be contributing as much as 90% of the total light, and it is for this reason that measurements are made of the sky to the north, south, east and west of the galaxy, in order to minimize not only position errors but also the continual variation in the brightness of the sky background.

The results of Pettit's colour index work lead to the following formulae for the relationship between colour index and magnitude in the range $9 \leq m \leq 18$:-
For ellipticals,

$$\text{C.I.} = 0.04 (m_{pg} - 9) + 0.84$$

while the corresponding equation for spirals was

$$\text{C.I.} = 0.08 (m_{pg} - 9) + 0.50$$

thus making elliptical galaxies redder than spirals.

With the publication of Holmberg's classic paper (Holmberg 1958), there became available to other astronomers a comprehensive catalogue of magnitudes and colours of galaxies. The advantage of this work was that since each magnitude measured represented the light contained within a certain isophote corresponding to a surface luminosity which was more or less constant for the entire plate material, all the

magnitudes were comparable measurements forming a homogeneous system. The observational procedure employed was to take two exposures of equal duration on separate halves of the plate, one an in-focus exposure of the galaxy, the other an extra-focal exposure of the stars in a comparison field. Microphotometer tracings of the stars then furnished the relation between plate density and surface luminosity, hence making the calibrations "independent of artificial means such as tube sentitometers". The magnitudes were obtained from the numerical integration of parallel cross-sections across each object, together with one scan along the major axis. Since the observations dealt with surface objects, the results were almost independent of seeing conditions between the galaxy and calibration exposures. In addition, Holmberg pointed out that focussing and guiding errors in the telescope, which could lead to large errors in stellar photometry, do not arise in nebular work. Of course, this is only applicable when it is the total luminosity of the system which is being measured, and not detailed photometry of the nuclear regions, which could be affected by the instrumental scattering function. Furthermore, photo-electric observations

give good results for galaxies with a rather high surface brightness, where the relatively steep gradients make photographic work rather difficult. In contrast, with galaxies of low surface brightness, it is only by photographic methods that reliable results can be obtained, due to the difficulty of eliminating field stars in the aperture field when using photo-electric measurements. Another problem in photo-electric work, the variation in sky brightness, which is unimportant in photographic work since the galaxy plus sky are being observed simultaneously, could only be overcome successfully by an observational technique which would permit a continuous monitoring of the sky background. A twin telescope, such as the one at the Royal Observatory, Edinburgh, would be well suited to this requirement.

One minor criticism of Holmberg's work is that at the beginning of the observational programme, each plate was calibrated individually, the exposure of the galaxy being made on one half of the plate and the star field exposure on the other. However, towards the end of the work, the two plates facing each other within a sub-package were treated as one photometric unit. Although it was claimed that no noticeable

increase in the errors resulted from this procedure, it is, nevertheless, true that no two photographic emulsions can have truly identical calibration curves. The principal sources of error in the work were the following:-

- 1) A difference in the sky fog between the region surrounding the galaxy and the calibration field.
- 2) Changes in the sensitivity of the emulsion between the first and second exposures.
- 3) the smoothing effect obtained due to the use of too large a scanning aperture in the microphotometry and
- 4) incorrect extrapolation over the nebular regions for the background curves representing the sky-fog.

In comparisons with other observers' results, Holmberg pointed out one possible reason for the unexpectedly large errors which arose in Pettit's (1954) work. These errors were due partly to inaccurate magnitudes of the North Polar Sequence stars which were used to calibrate the photo-electric galaxy measurements and also to the high ultra-violet transmission of Pettit's BG 1 filter (de Vaucouleurs 1961a). Holmberg's comparison with Stebbins and Whitford's (1952) photo-electric work produced deviations which ranged from -0.29^m to $+0.17^m$ with a mean deviation of -0.03^m . A

possible source of error in the photo-electric work was the difficulty in locating the centres of the galaxies without a pronounced central core in the field apertures. Following these comparisons, Holmberg studied integrated colours of galaxies against their classification type, rather than studying colours at specific points within the galaxy.

The methods by which galaxy dimensions were measured, discussed in detail by de Vaucouleurs (1957) were standardized in a paper by the same writer published in 1959. Prior to these publications, the brightness dimensions referred to a given isophote of some specified brightness. The effective dimensions, defined previously (de Vaucouleurs 1948), had the advantage that only relative measurements of surface intensities with an arbitrary zero point were required for their calculation, whereas brightness dimensions required the absolute measurement of surface brightness in a given magnitude scale. For an elliptical system, the effective dimensions were D_e and d_e , giving an axial ratio of d_e/D_e . Photometric dimensions, which are in nature brightness dimensions, are determined by the threshold brightness \underline{I} together with the law of luminosity distribution $I(r)$ in the image, while the

effective dimensions depend only upon the latter. The two laws for the luminosity distribution were summarized (de Vaucouleurs 1961c) in the following form:-

- 1) for ellipticals, the quarter power law (de Vaucouleurs 1948) was now written in the revised form:-

$$\log J = -3.33(\rho^{\frac{1}{4}} - 1) \quad \text{and}$$

- 2) for lenticulars and spirals the relation was

$$\log J = -0.729(\rho - 1).$$

The coefficients -3.33 and -0.729 are required by the form of the respective functions.

Another type of dimension, micrometric, was a direct micrometer measurement of photographic negatives. These dimensions, of course, refer to an unspecified mean brightness threshold depending upon the properties of the object, of the foreground and of the photographic plate itself, de Vaucouleurs decided after experiments that the most significant parameters were the major diameter D and the axial ratio d/D . If many independent measurements were available, together with correspondingly good calibrations, the corrected micrometric data could be used as distance indicators. A statistical analysis of apparent diameters (D) as a function of axis ratio D/d suggested that the relationship between the face-on

diameter and the observed diameter D is

$$\log D = \log D(0) + 0.4 \log D/d$$

where $D(0)$ was the face-on diameter i.e. the diameter that would be observed if the axial ratio were equal to unity or if $\log D/d = 0$. Consequently, a galaxy which was observed edge-on with an axial ratio $R = d/D = 0.2$ appeared with a maximum diameter almost twice the face-on diameter. This confirmed the observational evidence for the apparent excess of edge-on systems in statistical surveys. Recently, de Vaucouleurs et al. (1970) have found that the coefficient of $\log D/d$ is nearer the value 0.2. The principal aim in securing more consistent data on diameters was the need to improve the extrapolation correction, $m_r - m_t$, which must be applied to the integrated magnitude m_r within a distance r from the nucleus in order to calculate the total or asymptotic magnitude m_t .

Departing from the field of photographic photometry, de Vaucouleurs (1961a) published integrated colours in the UBV system for 148 galaxies observed photo-electrically in 1957 and 1958. After having determined $(U-B)$ and $(B-V)$ colours, the face-on colours (i.e. corrected for tilt effects), $(U-B)_0$ and $(B-V)_0$ were found to vary as a function of morphological type and of the ratio

$A/D(0)$, the field aperture to the face-on diameter. One of the advantages of these photo-electric measurements is that they can be used to fix the zero point of photographic measurements. By knowing the diameter $2r$ of the field aperture used in the photo-electric work, it is possible to establish the relationship between m_r , the measured magnitude with reference to the standard stars in the UBV system, and L_r , the integrated luminosity within the same radius derived from the photographic plate, in the form $m_r = \text{constant} - 2.5 \log L_r$. This constant yields the zero point of the photographic photometry.

The colour indices quoted do not measure the local colour, $C(r)$, at a given radius r from the centre of the galaxy, but the integrated colour out to a given radius r defined by the following equation

$$|C|_0^r = \int_0^r C(r).I(r).dr / \int_0^r I(r).r.dr$$

Clearly, if detailed knowledge of the law of luminosity distribution $I(r)$ were available, it would be possible in principle, to derive $C(r)$ itself from the integrated colours, although in practice, this is not the case due to the sensitivity of integral equations to accidental errors of measurement. When the face-on colours were plotted as a function of $A/D(0)$ for several types of

galaxy, the general decrease of integrated colour indices outwards from the centre, together with the tendency to reach asymptotic values for ratios of $A/D(0) \geq 1.0$, is reversed and becomes an increase for Magellanic type spirals and irregulars. The range of values for the colour index was small for ellipticals and lenticulars and larger for the a, b and c stages of spirals. In comparison of his colours with other observational results, de Vaucouleurs pointed out when a comparison was made with Holmberg's (1958) values, asymptotic colours for which $A/D(0) > 1$ had to be used, because Holmberg's values were differences between the total photographic and photovisual magnitudes, as determined by the integration of microphotometric cross-sections.

The next two important papers to appear were by M.H. Liller (1960) and van Houten (1961). References to both papers appear to be relatively rare, and it is therefore considered prudent to make a more critical examination of these two papers in comparison with the more general background information which has been the object of this appraisal of published work in the field of galaxy photometry.

The most obvious source of error common to both papers, was the apparently indirect method of plate

calibration. Liller used two plate collections in a study of the distribution of intensity of elliptical galaxies in the Virgo cluster, one taken with the 24-36-inch Curtis Schmidt, the other taken with the 48-inch Mount Palomar Schmidt. The Curtis Schmidt plates were calibrated in a fairly conventional manner, and by a comparison between the intensities within galaxies common to both sets of plates, a calibration was obtained for the 48-inch Schmidt plates, which had previously been uncalibrated. The location in this manner of identical points within a galaxy on different plates is extremely difficult to achieve accurately, and the graph which Mrs Liller used to reinforce her argument that the calibration was satisfactory was derived from measurements of three objects. This graph was then used as a calibration standard for all the 48-inch plates. Isophotes for the galaxies were obtained with the Michigan isophotometer, and graphs were drawn of the ellipticity of the isophotes against distance from the centre. It was found that spiral galaxies exhibited a curve of increasing ellipticity from the centre outwards, while the curve for ellipticals rose to a maximum at a point approximately equal to half the value of the semi-major axis. Agreement with the quarter power law was obtained for most of the objects under investigation,

although some galaxies which did not satisfy the law were shown by de Vaucouleurs (1961c) to be S0 types.

The year 1961 saw the publication of a very extensive study by van Houten of 20 galaxies photographed by Oort from 1936 onwards at Mount Wilson. Oort's original plates had sensitometer marks, together with the additional calibration of extrafocal exposures on Selected Area stars, but in the photometric reductions, both these calibration sources produced different results. The absolute calibration was to be provided by photo-electric observations of four galaxies, together with extra-focal exposures of standard stars. However, before the absolute calibration could be used, it was necessary to obtain the relative calibrations, and these were obtained with the following rather indirect technique.

The images of the extra-focal Selected Area stars which appeared as rings on the plate were scanned in two mutually perpendicular directions. Since coma was present, it was found that only those images in the centre of the field gave four deflections suitable for averaging and which could give a value for the density scale. Even if the abscissa of the characteristic curve, the intensity, had been known

with a high degree of accuracy, the possibility of errors in the curve would, nevertheless, have been considerable for two reasons. The first was the poor density measurements and secondly, more important, the intensity scale used by Seares in his 1930 catalogue of magnitudes of Selected Area stars. Although these magnitudes were known to have a scale error for stars fainter than magnitude 14^m , van Houten neglected this error in the construction of the calibration curves which he subsequently described as "not ambiguous". The general form of the curve was established after inspection of a mean curve determined in 1940 by Kleibrink; unfortunately, the blackening of the plates had changed during the intervening years. Finally, differences in the exposure times of the galaxy and standard stars introduced errors due to reciprocity failure, and the Schwarzschild exponent had to be calculated. It is surprising that any useful information whatever resulted from the photometry of plates which seemed beset by so many calibration difficulties. In addition, one of the sets of isophotes which were constructed for each galaxy contained deformations due to the fact that the plate had been broken, the pieces laid together on the microphotometer

and the isophotes assembled from the resulting tracings! Since the apertures used for the photo-electric observations were much larger than those used for the microphotometry, dispersion conditions had to be applied to the data before the relationship between the two systems could be established.

Part of the discussion of results concerned the variation of axial ratio with different isophotes. van Houten found that for NGC 4472, 4486, 4621 and 4649, the axial ratio of all isophotes was constant, but Oort (1940) de Vaucouleurs (1948, 1959a) and Liller (1960) had shown that there were variations in the ellipticities. This would suggest that the accuracy of van Houten's isophotes was not particularly high, especially since Liller's isophotes were obtained using the University of Michigan isophotometer, which was known to produce reliable results. van Houten tested the validity of the quarter power law for ellipticals, and the intensity distributions of the four galaxies already mentioned for which the axial ratio of the isophotes remained constant could be well represented by de Vaucouleurs' (1959) relation. However, this relationship, which was known to hold only for elliptical objects, was then tested for mixed edge-on systems and spirals. Surprisingly, the only agreement

obtained was for the minor axes of the bulges of spirals but van Houten still denounced the quarter power law on the basis of these comparisons.

Around the same time as de Vaucouleurs ' (1961) photo-electric work, Tifft (1961, 1963) was conducting an investigation into the colours of galaxies. The scope of the work involved four-colour photometry of about 50 galaxies, each galaxy usually being observed with three diaphragms. An important feature of the work was that, in addition to plotting diagrams for the variation of integrated colours against $A/D(0)$, Tifft plotted diagrams for concentric ring measurements against $A/D(0)$ to study the actual colour within the galaxy. There was no appreciable difference between the two diagrams, which agreed with the results of de Vaucouleurs (1961).

Shobbrook (1965), who made a similar investigation of 53 southern galaxies, also determined their total magnitudes. In order to obtain such magnitudes directly, it would have been necessary to take measurements through field diaphragms whose diameters were at least twice the accepted diameter of the galaxies. Naturally, this was found impossible for some of the larger brighter objects, but the problem

was supposedly overcome by a rather indirect method which involved the determination of the dependence of other observers' magnitude measurements upon the ratio $A/D(0)$. Extrapolation then produced for different types of objects the magnitudes corresponding to the largest $A/D(0)$ ratios required by the project.

Two papers have been published in recent years concerning the relationship, if any, between the ellipticity and diameters of elliptical galaxies. Edelen (1965) proposed that the mean linear diameter was a decreasing function of the ellipticity, and in contrast, Rood and Sastry (1967) found no relationship between the two quantities after an examination of 176 elliptical systems.

One publication closely related to the work of this thesis is that by Markaryan, Oganessian and Arakelyan (1965). Blue and yellow observations were made of ten elliptical and lenticular galaxies in the Virgo cluster using the 40-inch Schmidt telescope at Byurakan Observatory. The reductions were performed on a microphotometer which had an attachment to permit an adjustment of 3 to 4 microns in two mutually perpendicular directions. Once more the calibrations were not of the accuracy demanded by precise

photometry, because following the galaxy exposure, another plate from the same box was exposed for the same time as the galaxy to an out-of-focus North Polar Sequence, obtained by inserting a plane-parallel piece of glass in front of the emulsion, in order to furnish the calibration. Both plates were processed simultaneously to minimize errors. This calibration technique is similar to that used by Holmberg (1958). For each galaxy under study, a chart was assembled giving the brightness of the galaxy within grid squares of size 9.7 arc seconds. This information was used to draw graphs of the surface brightness and colour index distributions along the principal axes. Much larger variations of the colour index were found when compared with other results (e.g. de Vaucouleurs 1961a). However, there was a moderate degree of uncertainty in the establishment of the zero points of the colour curves. A continuation of this work examined the surface distribution of six spiral galaxies (Markaryan et al 1966). Rood and Baum (1967) have catalogued 315 galaxies in a 0.32 square degree field in the centre of the Coma cluster. In later work (Rood and Baum 1968), the detailed

reduction procedure for deriving brightness profiles for 32 objects was described. Due to the large sampling interval used in the microphotometer digitization, errors may be introduced with possible resolution difficulties. King (1961) has attempted to obtain empirical laws for stellar density distribution in globular clusters. This density for $r > 1'$ was found to be closely proportional to $\left[(1/r) - (1/r_0) \right]^2$ where r_0 denotes the limiting radius at zero density. Later work (1962a) produced a law represented by three parameters: (1) a number factor (2) the size of the central core r_c and (3) the limiting radii r_0 . The profile is characterized by a concentration ratio (defined by r_0/r_c) which may have values from 4 to 130 for the globular clusters. This same profile fits the dwarf ellipticals, where the ratios range from 3 to 10, but larger values are found for the giant ellipticals, with their higher central brightness. For the central regions, the surface density is given by $f = f(0) / 1 + (r/r_c)^2$, where $f(0)$ denotes the central surface density (King 1962b); complete profiles for 9 globular clusters have been obtained (1966) using the above relations.

The first Atlas of photometric results for many galaxies was published by Sersic (1968); it gave photographs, descriptions and detailed surface photometry of 50

southern galaxies. Although the parameters in such a recent publication are not identical to the "standards" defined by de Vaucouleurs (1968a), there is still some similarity between the two systems. For example, if $S(m)$ denotes the area of the galaxy brighter than magnitude m , then

$$S(m) = K(m - m'')^n$$

where m'' defines the zero point and the exponent n ranges in value from 2 for late-type spirals to 8 for ellipticals; intermediate values of n denote mixtures of types. The value of n is related with the parameter m_0 by the relation

$$m_0 = m'' + 1.086 n.$$

Possible errors in the Atlas may be found in galaxies with complex structures, since some of the isophotes were drawn "with the aid of photographs".

Kron and Papiashvili (1967) and Walker and Kron (1967) used the electronic camera in an investigation of M13 and M15, and obtained good agreement with the quarter power law (de Vaucouleurs 1948, 1959). The superb advantages of the instrument are its high storage capacity, relatively high quantum efficiency, low noise level, and, more recently (e.g. Ables & Kron 1967), with a suitable choice of emulsion and development conditions, its linear

response (some difficulties were present in original designs (Walker 1965) with non-linearity of response). Since the camera has electrostatic electron optics, the photo-cathode must be curved towards the photographic plate. Due to the small size of the cathode, the largest circular field available at the present time is 35 mm diameter. Larger fields may be possible with the Griboval camera which uses magnetic focussing.

In conclusion, six recent publications offer encouragement for workers in the future to improve upon the accuracy of previous measurements. In the first place, the characteristic curve has now been expressed in a linear form by de Vaucouleurs (1968). This relationship should enable more precise calibrations to be obtained for astronomical emulsions. Secondly, the use of the IIIaJ emulsion and a Wratten 2C or 4 filter (van den Bergh, 1969, Arp and Bertola 1969, 1971) has enabled faint outer regions of galaxies to be more easily detected than in the past. Such regions have been confirmed by de Vaucouleurs (1969, 1970c) for M87 and NGC 4889 using sensitive photo-electric methods. This wealth of information still available on these faint outer regions may herald the beginnings of significant advances in theories concerned with the structure and evolution of galaxies.

Chapter II

2.1 Introduction. While the last few decades have seen rapid advances in many fields of astronomical research, detailed surface photometry remains one field in which astronomers have seemed reluctant to devote the attention necessary to keep pace with the progress in other related topics in the study of galaxies. Only Oort (1940), de Vaucouleurs (1961b, 1963b, 1963c, 1964), de Vaucouleurs and Page (1962), van Houten (1961) and Liller (1960) have published and used detailed isophotal maps in their investigations.

Two possible explanations may be offered to account for the absence of regular publications in the subject. The first is that the reduction techniques necessary to produce reliable results are extremely involved and lengthy, a fact which does not suit the American publication system where frequent small articles often seem to be more important than less frequent but more substantial contributions. Secondly, no standard system of parameters for galaxies had been in existence before the ones proposed by de Vaucouleurs and now adopted by the Working Group on Galaxy Photometry.

This Group was formed at the 1961 I.A.U. Meeting in Berkeley, and before that date, comparisons of results were almost impossible, due to the different systems chosen by each observer. The present investigation was carried out following the recommendations of the Working Group. However, before undertaking a major programme on the surface photometry of galaxies in an attempt to bridge the existing gap in extra-galactic studies, a preliminary investigation was made into all the factors which could affect the accuracy of the results, and it is these factors which will be considered in turn throughout this chapter.

2.2 The Telescope. The Cassegrain Schmidt telescope available for use at the University Observatory, St Andrews was completed in 1963, and all the plate material used in the work was obtained with this instrument. The optical system and dimensions of the James Gregory Telescope, as it is called, are shown in figure 1. The primary and secondary mirrors are spheres of diameter 37 and 18 inches and radii of curvature 116 and 112 inches respectively, giving a nearly flat Petzval surface (Linfoot 1955). The aperture stop which is situated midway between the

primary mirror and the corrector plate is formed by an iris diaphragm, adjustable from 30 to 37 inches diameter.

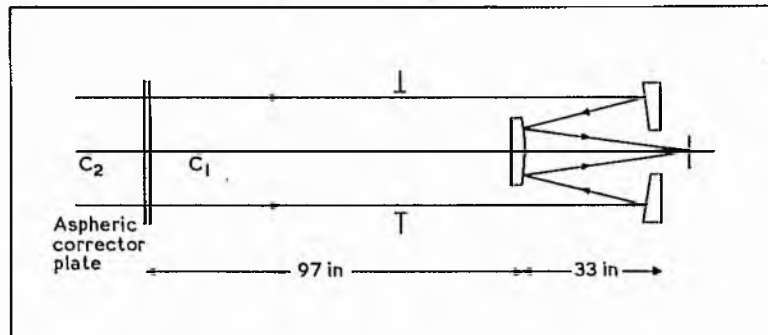


Figure 1 - Schematic diagram of James Gregory Telescope.

The large size of the secondary mirror is dictated by the requirement of an accessible focus with mirrors of similar radii, and the precise value of the Petzval curvature is chosen to balance out the field curvature due to the presence of a small amount of astigmatism (less than 1 arc sec at 2° off axis), thus giving a flat best field. The focal ratio of the system is 2.9 giving a plate scale of 1 mm = 76.4 arc sec and using the 8 inch square photographic plates, a 4° square field is produced.

After initial tests to determine the focus of the

telescope, it became apparent that a certain amount of spherical aberration was present, amounting to a circle of least confusion of under 8 arc sec, due to the corrector plate being too strong. It was pertinent to enquire what effects, if any, this would have on a programme of surface photometry. The luminosity distribution of a stellar image may be closely approximated (de Vaucouleurs 1948) by a unique Gauss function of the form

$$\frac{I(x)}{I(0)} = \exp -x^2/2\sigma^2$$

where $I(x)$ represents the intensity at a radial distance x from the centre of the image, having a central intensity $I(0)$. The dispersion of the stellar images for the James Gregory Telescope was found equal to 2.75 arc sec compared with 2.3 arc sec for the 30-inch Cambridge Telescope used by Redman and Shirley (1938). Clearly, image structure within 8 arc sec of the centre of a galaxy will suffer from distortion, but the overall luminosity distribution will not be affected. However, since surface photometry involves the determination of a mean profile, some selection effect must be employed for the present work. If at $r = 6$ arc sec, the specific intensity in the star image is $I = I(0)/10$, then for

such effects to be relatively unimportant in the smoothing of a galaxy profile, the galaxy should be detectable to at least 10 times this value, i.e. to a distance of 60 arc sec or a diameter of 2 arc min. Consequently, only those galaxies for which the value of $\log D(0)$ exceeded 1.3 in the Reference Catalogue (de Vaucouleurs and de Vaucouleurs 1964) were chosen. The comments on image structure for the James Gregory Telescope referred to the centre of the field. For wide field work the aperture stop should be set to 32 inches to eliminate vignetting, and initial tests using 4-inch square plates confirmed the acceptability of the images over the field. However, when the work was extended to the full field width of 4 degrees, it became necessary to reduce the aperture stop to its minimum value of 30 inches in order to improve the image quality at the edge of the field, and it was as this aperture that all the observational work was undertaken.

2.3 Calibration. Calibrations which under most conditions would appear to be satisfactory were the subject of some critical comments in the previous chapter. These comments were made for two important reasons. In the first place, it is believed that

the method used throughout the current work was extremely accurate, since it avoided many of the sources of uncertainty in previous work. Secondly, under conditions where a calibration exposure is made on a plate which may, but not necessarily so, have an emulsion sensitivity identical to that of another plate taken from the same box and used for the telescope exposure, it is still almost impossible to create identical processing conditions for both the telescope and calibration plates.

The calibrations in the present investigation were undertaken in the following manner. At one end of an optical bench of length four feet was a light source with a Wratten 78A filter to raise the colour temperature from 2360°K to 3200°K . At the other end, there was a plate holder assembly with several light baffles interposed between these two units to prevent reflections from the walls of the bench assembly. The mounting of the lamp was such that the plane of the filament lay along the line drawn from the wedge to the lamp filament in order to reduce to a minimum the effective area of the source as seen from the plate. The colour temperature filter was placed as close as possible to the lamp to achieve minimal variation of the intensity of illumination at the plate. For a

similar reason, the achromatic density wedge was almost in contact with the emulsion. All surfaces in the system were painted matt black.

This wedge was in the form of an achromatic prism mounted in one corner of a square metal plate. Between the clear glass strips at each end of the wedge, density steps with value 0.2, 0.4, 0.6, 0.8 and 1.0 were obtained by a rhodium plating technique. The width of each strip was 1.9 mm. Since the photographic emulsion responds to an intensity ratio far in excess of the 10:1 ratio produced by the wedge, it was necessary to increase the ratio by placing neutral density filters in the light beam. These filters were located so that they were not quite perpendicular to the beam in order to prevent the occurrence of multiple reflections, and also as close as possible to the light source to minimize any variations in the flux incident upon the wedge due to any non-uniformity in their transmission characteristics. During the telescope exposures, the four corners of the plate were masked off to leave the corners free for calibrations, which, throughout this work, were always made on the same plates as the galaxy images. The interposition of the filters in the calibration beam necessitated further exposures

in order to obtain other segments of the characteristic curve. Measurements of the calibrations performed in this way showed that the plate fog level adjacent to the wedge exposures remained extremely constant at the four corners. Since the calibration exposures were four times as long as the telescope exposure, it was decided to obtain a wedge with a larger density range so that the time devoted to calibration could be reduced without any loss of accuracy, so long as the fog level remained constant. A coating of rhodium producing a density of 1.0 and designated by the hatched area in figure 2 was made on the existing wedge assembly.

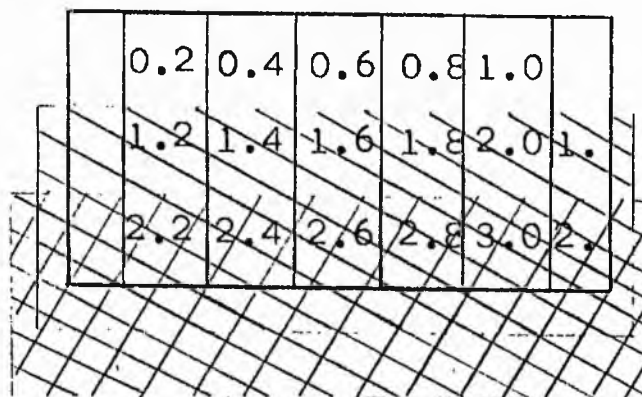


Figure 2 - Diagram of density wedge.

A further 1.0 coating, shown by crosses, was then evaporated on top of the 1.0 hatched area, thus

producing a wedge with a range of density of 3 and giving an effective transmission factor of one thousand. Using this wedge, it was possible to obtain the complete density range necessary for drawing the calibration curve with one exposure on one corner of the plate. However, both as a safeguard that the exposure and neutral density filter combination were suitably chosen to give the correct densities on the plate and also as a check on the initial exposure, a second exposure was made in the diametrically opposite corner of the plate. Following the calibrations, a time delay of at least three days was usually given to take account of latent image decay of both images. This interval concurs with, and in most cases betters a value given by de Vaucouleurs (1968) for $\Delta T/T < 0.1$ where ΔT is the age difference between the telescope and calibration exposures and T is the time delay between the latter exposure and the development. In the same paper, details were also given of the linear relationship between the intensity and the opacitance (equal to the opacity minus one) over the whole range of densities of practical interest in photographic photometry. All the characteristic curves in the work were originally drawn in the conventional H and D

S-shaped curve. However, these curves were constructed by drawing through a set of points a curve whose correct shape was not known a priori. When the writer became aware of the linearization method for drawing characteristic curves, the calibrations were carried out a second time using this new system. The remarkable differences between this method and the previous work emphasized the errors which could be introduced by drawing an arbitrary curve through a set of points. With the straight line, it was obvious whether or not any one point had an erroneous density value, whereas previously a polynomial curve was forced to fit all the data points. Another advantage of this new system is that since fewer data measurements are needed, such a refined density wedge as that used in this work is not really necessary in future calibrations.

2.4 Microphotometers. The original type of microphotometer available for use was a Hilger and Watts L470 scanning-type microphotometer. Basically, the operating principle of the instrument was as follows:- most of the available output from a light source passed through the plate and hence to a photomultiplier while the remainder passed directly to a second photo-

multiplier in order to take account of any variations in the illumination of the lamp. The net current from the two photo-tubes was fed to a Brown Recorder. The instrument thus measured transmission and not density.

Two major difficulties were present when initial scans across galaxies were made in the summer of 1965. In the first place, errors were introduced due to the fact that the drive motors on the microphotometer and the Brown Recorder did not rotate with a constant angular velocity. This resulted in distortions and scale errors in the scan abscissae. For isophotometric work, where many scans had to be taken across the object at well determined intercepts, a measuring screw was fitted to the plate table to provide the accuracy in y-spacing down to the necessary 5 micron stepping interval between the successive scans. The second and possibly the more important difficulty was the inherent drift present in the instrument. Due to the relatively slow response of the Brown Recorder, especially in regions where the density gradients are steep, individual scans had to be taken at slow driving speeds. Even in the course of one scan, it was possible to detect changes in the dark current reading before and after the scan. This change was almost certainly

caused by the non-uniformity of response of the two photo-tubes, thus making transmission measurements dependent upon intensity changes in the reference beam.

These two difficulties are overcome in the design of the Joyce Loebel microphotometer. A direct mechanical linkage between the plate and recording table ensures synchronization of table and pen positions, irrespective of the magnification ratio chosen. The ratios available range from 1 to 2000. Furthermore, the instrument incorporates a true double-beam system in which two light beams are taken from a single source, one to act as the measuring beam, the other as reference, and both beams terminate at the cathode of one photomultiplier, thus eliminating possible variations in the intensity of the source. Due to the limitations of the Hilger microphotometer, the University Observatory acquired in 1965 a Joyce Loebel microphotometer, which also incorporated the Isodensitracer attachment. A more detailed description of the instrument's isophotometric operation mode will be given in the next chapter which will concern isophotometers. For the moment, however, the microphotometric mode will now be described, together with the additions necessary for the digitization of the instrument, carried out during the summer of 1969.

All the cross-sectional scans across galaxies, which were necessary to obtain the profiles and calibrate the isophotes, were digitized and subsequent computer reduction of the information was used to obtain the luminosity profiles for each object.

The Joyce Loeb1 microphotometer is a double beam instrument of the null balancing type, balance being achieved by the servo-positioning in the reference beam of a linear density wedge. Linked directly to this wedge assembly is the recording pen. The machine will therefore measure density linearly within the range of the particular wedge in use. The instrument supplied to the University Observatory was also fitted with helical potentiometers to provide analogue outputs of both table position and density; both outputs were connected to digital voltmeters. The sampling interval was achieved in the following manner. From inspection of a typical cross-section, the noise pulses were found to occur on average every 1.5 arc sec on the plate. In order to retain all the available information, it was decided to sample every second of arc. Using the 50 times magnification arm, this would involve making 300 samples in the course of one scan. The digital voltmeter measuring the table position was a five digit ten volt full scale model, and with a

suitable choice of the voltage applied to the table potentiometer, it was possible to make the sampling interval correspond to 0.02 volts.

It was then only necessary to feed at each sample a triggering signal to the wedge digital voltmeter in order to perform the density digitization. This was achieved by fitting on a nor gate to each terminal of the digital voltmeter corresponding to an odd number in its second least significant digit. This had the effect of switching on a transistor every time an odd number was displayed, and of switching off the transistor when the display was even. When the transistor was switched on, the current supplied was used to drive the triggering signal for the digital voltmeter measuring density, thus performing the necessary digitization. At the same time, the voltage value corresponding to density was punched out on paper tape.

The basic idea of the system, although satisfactory in that it produced digitized pulses when commanded to do so, was nevertheless subject to a considerable number of spurious multiple density digitizations due to pick up of noise pulses. Let us consider in detail the effects produced by noise pulses which are ringing in nature, that is, cases where the amplitudes of the

pulses slowly decay with time. At a change-over from an even to an odd number, one digitization is produced. When noise pulses are present, the probability of the odd number going even again and producing a double digitization is highest exactly at a change-over, either from an even to an odd number or vice versa. This is because the table position is continuously moving away from or towards a change-over value, and at intermediate points the amplitudes of the pulses are not sufficiently large to swing the voltage to a change-over value. In order to overcome this problem, the average time between digitizations was calculated, and a monostable introduced to the electrical circuit in order to leave on a transistor, once switched, for the period of the duty cycle, irrespective of whether the input signal to the base of the transistor is attempting to reverse the switching, due to noise pulses. The circuit of the monostable was arranged so that the duty cycle was 80% of the sampling time interval. This arrangement largely overcame the difficulties experienced with multiple digitizations, but a further check was made during scanning, when it was a simple matter to detect audibly 50 punches per volt change in the table digital voltmeter corresponding to 50 required digitizations. Any scans in which extra pulses

happened to be produced were discarded and the scan re-run.

This description of the digitization procedure has been included because it only involved the expense of some £10 on the electronic components necessary to build the stabilized power supply for the potentiometers, the nor gate and the monostable. All the other equipment used in the system was supplied from stock already available in the University Observatory.

2.5 Processing Equipment. Basic to all successful photographic photometry is the meticulous care with which the processing of valuable plates must be undertaken. It is appropriate to consider each stage of the process in turn. In the first phase, each plate should be developed in fresh developer at a temperature of 68°F for five minutes. For a set of plates taken on different occasions, it is desirable to obtain identical development conditions. In this connexion, although time control is relatively easy to achieve, such control is not quite so easy with regard to temperature. The main requirement would be extremely accurate control over the darkroom temperature and the user should check the temperature of the solutions before commencing development. Perhaps the most suitable

system would be one in which control was made over both temperature and humidity, as suggested by Miller (1966).

Agitation of the plates also presents problems. In microphotometer scans across the initial wedge exposures, it was noticed that edge effects were present, even although the developing dish had been thoroughly rocked during development. These effects disappeared when brush development was used, and this method was used for all the plates processed in this investigation. The edge effects found in the previous plates were attributed to standing waves being set up in the developer, possibly caused by too rhythmic a rocking motion of the dish. Another source of uneven development was due to poor immersion of plates in the developing solution. For optimum results, the developer should, on immersion, make immediate contact with the entire emulsion surface and to aid this process, the plates were first immersed in water before immersion in the developer.

With regard to fixing and washing, the plates were fixed only for as long as necessary, because excessive fixing is liable to cause bleaching of the image. In a similar way, washing was accomplished in the minimum time in order to avoid the clumping of silver grains

which often occurs after long periods of washing. Throughout the present work, the plates were not washed continuously during the washing period, but this interval was subdivided into six equal washes, each of five minutes duration, interspersed with complete changes of water.

Chapter III

3.1 Introduction. One of the most important requirements for obtaining successful results in any photometric investigation of galaxies is the accuracy of the isophotal map which gives the distribution of relative intensity with the object. Isophotes may be produced in two different ways, one photographic, the other by mechanical means, and both systems will be considered in turn in the following sections.

3.2 Photographic methods. The first work using this method was done by van Hopf (1937), who constructed isophotes by sketching around the boundaries between the light and dark areas of a very high contrast print made from a galaxy negative. The high contrast removed the smooth intensity distribution inherent in most galactic images and the nett result was a fairly well-defined light galaxy on a dark background. This approximate method was superseded with the application to galaxy photometry of the Sabattier effect in 1963 by Richter and Hogner (1963). In view of the relative speed with which output may be obtained in comparison with more conventional techniques, this method appears well suited to the problems in galaxy photometry. This is true up to a point, but one important feature is the critical dependence of the

method upon the uniformity of illumination of the light source used for the exposures. A series of tests was made to assess the potentiality of the technique. The results of these tests are illustrated in figures 3 and 4, which represent the best isophotes which could be obtained for NGC 4258. However, experts in the technique seem to produce better results (e.g. Richter and Hogner 1964, 1965, 1966). The Sabattier process was used for the isophote shown in figure 3 while the variation of the original process proposed and used by Hodge and Brownlee (1966) was employed to produce figure 4. In this variation, several contact copies of a plate were made, each with a different exposure. Contact copies of copies were also made in order that pairs of negatives and positives of slightly different density may be selected and sandwiched together, when an enlargement from the composite should produce a clear sharp isophote.

Two features are apparent after an examination of figures 3 and 4. In the first place, isophotes as such are extremely difficult to produce, because the definition of the isophote seems to depend very critically on the exposure ratio between the negative/positive pair. In unfavourable cases, an area of approximately equal intensity is produced, not a clear



Figure 3 - Isophote of NGC 4258 obtained using
the Sabbatier process.

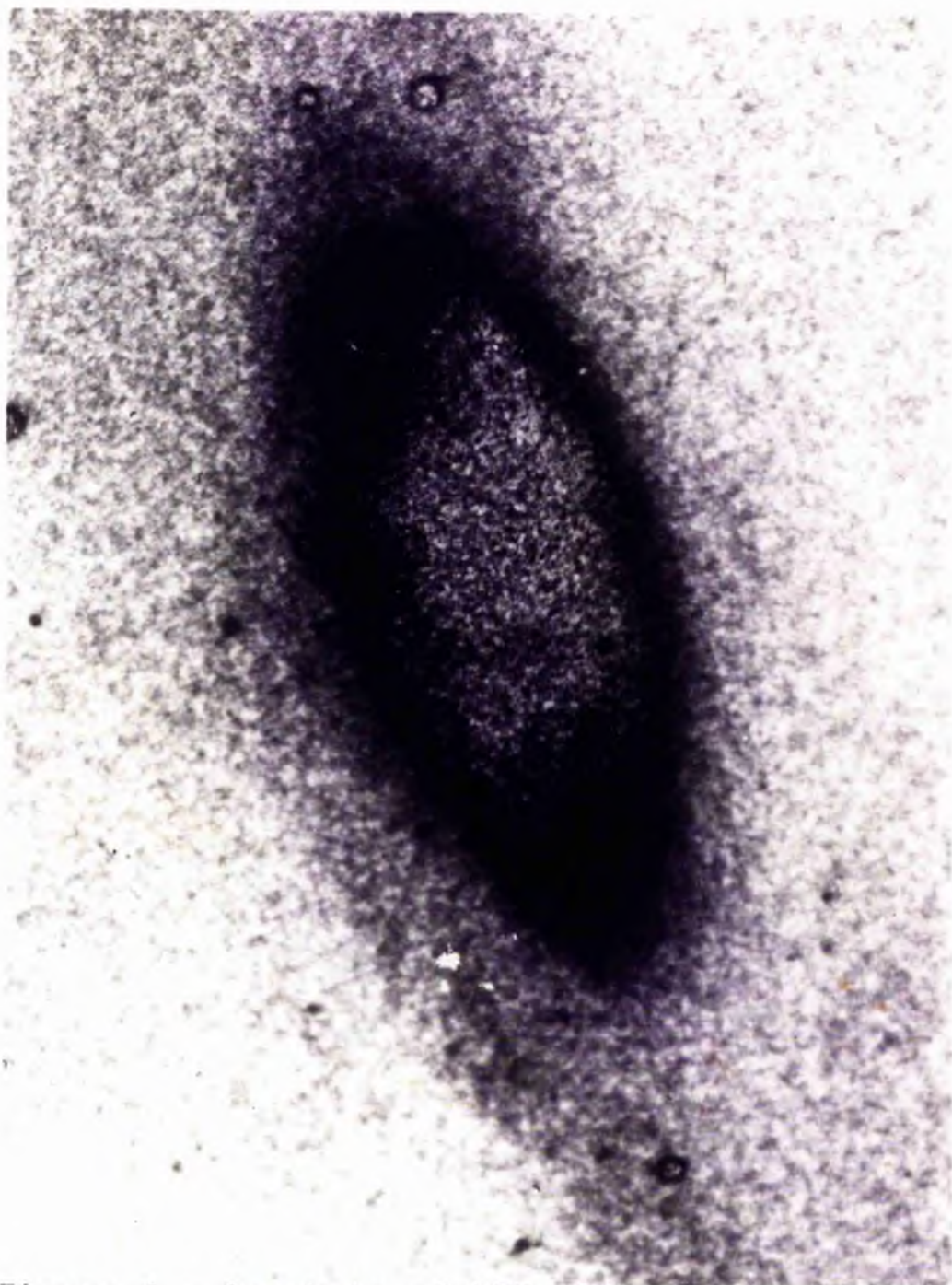


Figure 4 - Isophote of NGC 4258 obtained using the photographic method of Hodge and Brownlee.

sharp line. When high contrast photographic materials are used, and in cases when the exposure ratio is suitable for producing the required line, photographic effects then play a major role in determining the exact shape of the resulting isophote. Secondly, the magnifications which may be obtained with an enlarger are much smaller than those available on most micro-photometers. Accordingly, to achieve any degree of successful isophotometry by photographic means, it would be expedient for the original plates to be obtained on a telescope with a large plate scale. Tracings of NGC 4258 with the scanning type isophotometer at the University Observatory are shown in the figures 5 and 6. The visible distortion in the stellar images were produced by design and will be explained at a later stage in this chapter. Comparison with figures 3 and 4 should illustrate to the reader why the photographic methods were abandoned in favour of mechanical methods.

At the recent 1970 IAU General Assembly in Brighton, it was pointed out at the meeting of the Working Group on Photographic Materials that a new high contrast emulsion yielded favourable results when used in work on the Sabattier effect. In view of this new information, it is possible on the one hand that more accurate results

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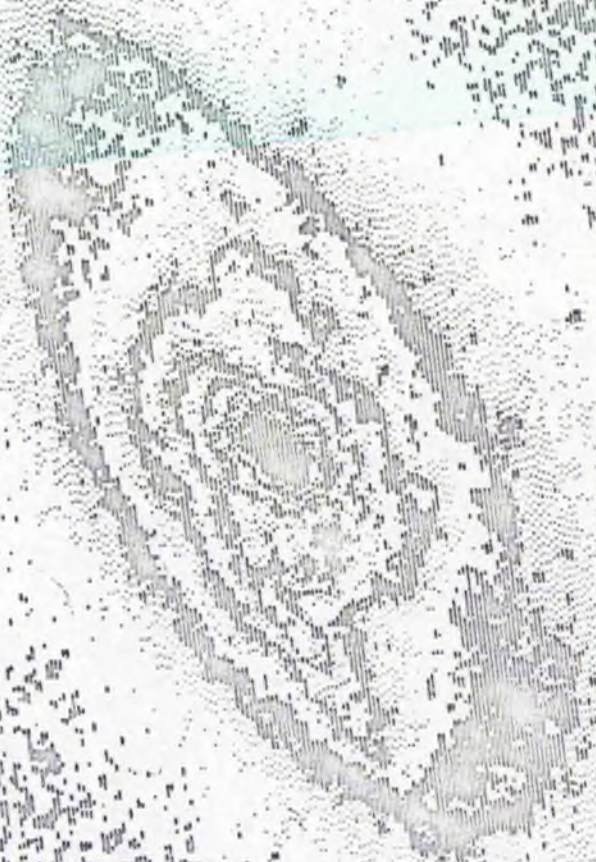


Figure 5

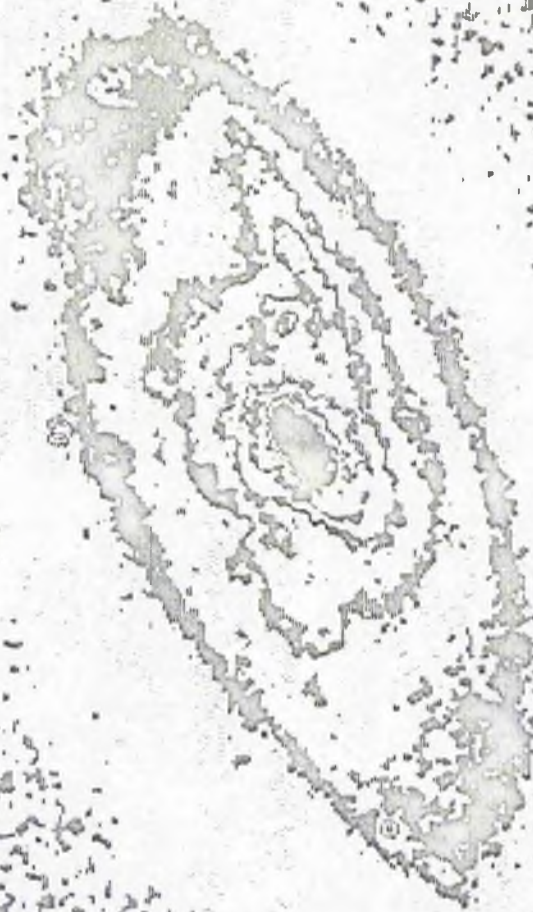


Figure 6

may be obtained in the future using photographic techniques, while on the other hand, it should be noted that all these isodensitometry techniques give curves of constant density and not the true isophotes corrected for plate errors, which can only be derived by more elaborate numerical mapping techniques (Jones et al. 1967).

3.3 Mechanical Methods. The simplest system for isophotometry using mechanical methods is the case where isophotes are constructed from microphotometric scans. For objects with fine structure, the process is extremely tedious and time consuming due to the small stepping increment required to record the details. Such a procedure was used by Holmberg (1958) and van Houten (1960). Holmberg, in particular, was successful in obtaining accurate results over the large sample of objects under study. The logical extension of this method would be complete digitization of the microphotometer output, subsequent computer reduction of the data followed by the production of isophotal contours.

As an alternative, adaption to an automatic scanning mode of a microphotometer fitted with some method for detecting and recording pre-defined densities would produce lines of equal density, which could then

readily have assigned to them their respective intensities. This scheme was used in principle by Mohler and Pierce (1957). However, the slow operating speed of their instrument due to the long time constant of the recorder combined with the large mass of the pen trip mechanism proved to be major disadvantages of this system. These difficulties were overcome by Koelbloed (1965), who designed an attachment for use with a normal microphotometer. The use of a high speed galvanometer together with teledeltos paper for the recording provided a much faster instrument than that of Mohler and Pierce. The writer used Koelbloed's instrument and considers it to have great possibilities as an isophotometer, even although it still lacks some of the refinements necessary to make it suitable for multiple applications. Further design improvements are in progress.

Perhaps the most simple and inexpensive method for obtaining isophotes is to follow, manually, points of constant deflection within an object using a microphotometer with precision x and y screws fitted to the plate carriage, noting the intermediate readings. This was done by de Vaucouleurs (1961b, 1963b, 1963c, 1964) and later adopted by Bertola (1966) in an investigation of

three late-type galaxies. It is surprising that such a method was chosen for objects with complex structure, rather than ellipticals or lenticulars, since a considerable amount of smoothing was effected in the delineation of the isophotes from the measured data points.

Automation of this principle was found in an isophotometer designed by Williams and Hiltner (1940). Its contour following mode of operation was known as hunting. Two limitations in the system were as follows:- first of all, the inability to record more than one isophote at a time and secondly, the inability to differentiate between a true isophote and such details as small islands or closed-loop isophotes within larger ones. Clearly these disadvantages only make the instrument suitable for isophotometry of systems without much fine structure. In addition, constant operator attendance would seem to be necessary.

de Vaucouleurs and Griboval (1965) have designed an isophotometer which can operate in two basic modes: (1) a contour following mode for simple closed loop isophotes and (2) a scanning mode to record more complex structures. This instrument combines the principle of the Michigan isophotometer together

with Koelbloed's design, and some illustrations are given for sample tracings. At the present time, there is a two dimensional digitized densitometer in the final stages of construction at the Department of Astronomy in the University of Texas at Austin.,

After due consideration had been given to the relative advantages of the different systems discussed in this chapter, it would appear that the basic essentials for an isophotometer are as follows:-

- (1) electronic stability
- (2) ability to scan rather than hunt in order to record all details
- (3) versatility in the choice of magnification used for a particular tracing in order to study selected areas of a galaxy rather than the complete object
- (4) fast operating speed.

These four requirements were fulfilled in the Joyce Loeb1 Microdensitometer with an Isodensitracer attachment (abbreviated in future to MDM and IDT). In view of the drift problems encountered with the Hilger microphotometer, together with the large time necessary to complete one scan, reliable construction of meaningful isophotes from the resulting tracings proved impossible. The MDM can scan fifty times more quickly than the Hilger micro-

photometer described in Section 2.4, and using the IDT attachment, a scan is produced not in the conventional y-deflection form, but as a coded x-direction line. Multiple scans, together with the appropriate table and plate increments, form an overall contour map of scanned area.

The scanning process is controlled by an electronic programmer. Upon completion of an x-scan, the table drive is reversed and at the same time two pre-selected pulse trains n_1 and n_2 are fed from the programmer, one to a stepping motor providing y-drive to the plate table and the other to a special recording pen stepping motor. This pen replaces the one linked to the linear density wedge in MDM work. The number of pulses n_1 fed into the plate table controls the scan spacing and the ratio $n_2 : n_1$ determined the y-magnification. Various magnifications may be obtained by selecting different settings of n_1 and n_2 . When the plate table has returned to its starting value and the y-stepping process has been completed, the next scan commences. A driving pulley is used to provide movement to the density wedge during the servo-positioning routine. In the IDT attachment, a commutator ring, illustrated in figure 7, is screwed on top of the pulley and two

metallic contacts are located radially at the points indicated in the figure.



Clear section denotes metallic areas.

Dark sections denote insulating areas.

Figure 7 - Diagram of commutator ring.

As the wedge assembly moves during servo-positioning, the pulley rotates and the contacts either touch one of the two metallic areas of the ring or in certain positions, no contact is made at all. This sequence of operations is called the quantization of the wedge movement. Depending upon which of the two metallic contacts are being made at any particular instant will determine the mode in which the recording pen will write - either dots or continuous lines. The density quantization thus changes a normal microphotometer tracing into a coded line. As scanning

proceeds, a two dimensional density map consisting of spaces, dots or lines is produced, a transition from spaces to dots to lines corresponding to increasing density and vice-versa. The resulting isophotes are thus uniquely defined as those lines which may be drawn through the points where the pen mode changes. By dividing the complete density range of the wedge by the number of divisions of the commutator ring formed within this range, it is possible to calculate the density increment per change of pen mode. All the isophotes therefore have integral multiples of density above some reference level. After the principal axes of an object have been determined from the isophotes, microphotometric cross sections along these axes furnish the calibration of the isophotes.

Before making any tracings, it was first necessary to check the magnification ratios so that identical x and y ratios would produce a one to one correspondence. An elaborate set of dial test indicators was fitted to the instrument. Results of test tracings showed that, contrary to expectations, it was not in the stepping motors that the errors were larger, but in the ratio arm. The 100 times magnification, used for most galaxies with angular diameters less than 3 arc min,

was found to have a magnification of 105 while the 50 times arm gave a value nearer to 51. For the y-stepping motors, the errors were only around one half of one percent.

In their paper on the merits of photographic methods for producing isophotes, Hodge and Brownlee (1966) mentioned that distortions could occur with scanning type isophotometers, and also that the interpretation of small scale structures with dimensions similar in size to the step size in y is extremely difficult. Criticism of this statement in defence of the scanning type isophotometer formed the basis of a paper published by the writer (1966). Figure 8 shows isophotes of NGC 3389 obtained with the University of Michigan isophotometer. Clearly the interpretation in this case of some details would prove difficult, and it is possible that when Hodge and Brownlee made their critical comment on this matter, they had no knowledge of the existence of the Joyce Loeb1 instrument. In figure 9, isophotes with a density increment of 0.06 are shown for the same object. They were obtained on the Joyce Loeb1 machine from a plate taken with the James Gregory Telescope. The reader should have no difficulty in assessing the increase in detail of the latter diagram over the Michigan isophotes. This additional accuracy is



Figure 8 - Isophotes of NGC 3389 obtained with
the University of Michigan isophotometer.



Figure 9 - Isophotes of NGC 3389 obtained with
the Joyce Loeb1 Isodensitracer.

possible mainly due to the close y-spacing adopted. The minimum increment which may be made in the y direction is 1.25 microns on the plate table, corresponding to 0.095 arc sec on plates taken with the Gregory Telescope. As a further comparison, a plate of NGC 5566 was taken at St Andrews to provide the necessary source from which an IDT tracing could be obtained. Hodge and Brownlee had selected this object in order to obtain the isophotes shown in figure 10 which were supposed to demonstrate the high degree of accuracy possible with the photographic methods over other methods. The IDT tracing from the St Andrews plate is shown in figure 11 and inspection of this figure with the previous one should reinforce previous remarks. The other advantages of the Joyce Loeb1 instrument over other isophotometers, in particular ones of the scanning type, are as follows:-

1. Since the y-scanning steps may be extremely small, it is possible to make detailed investigation on small scale structures without the interpretative problems described by Hodge and Brownlee (1966). Any confusion which may arise may be resolved on the basis of rotating the plate through 90° and repeating the tracing. Exact similarity of the isophotes is obtained after taking account of the slight magnification errors in the ratio arm.

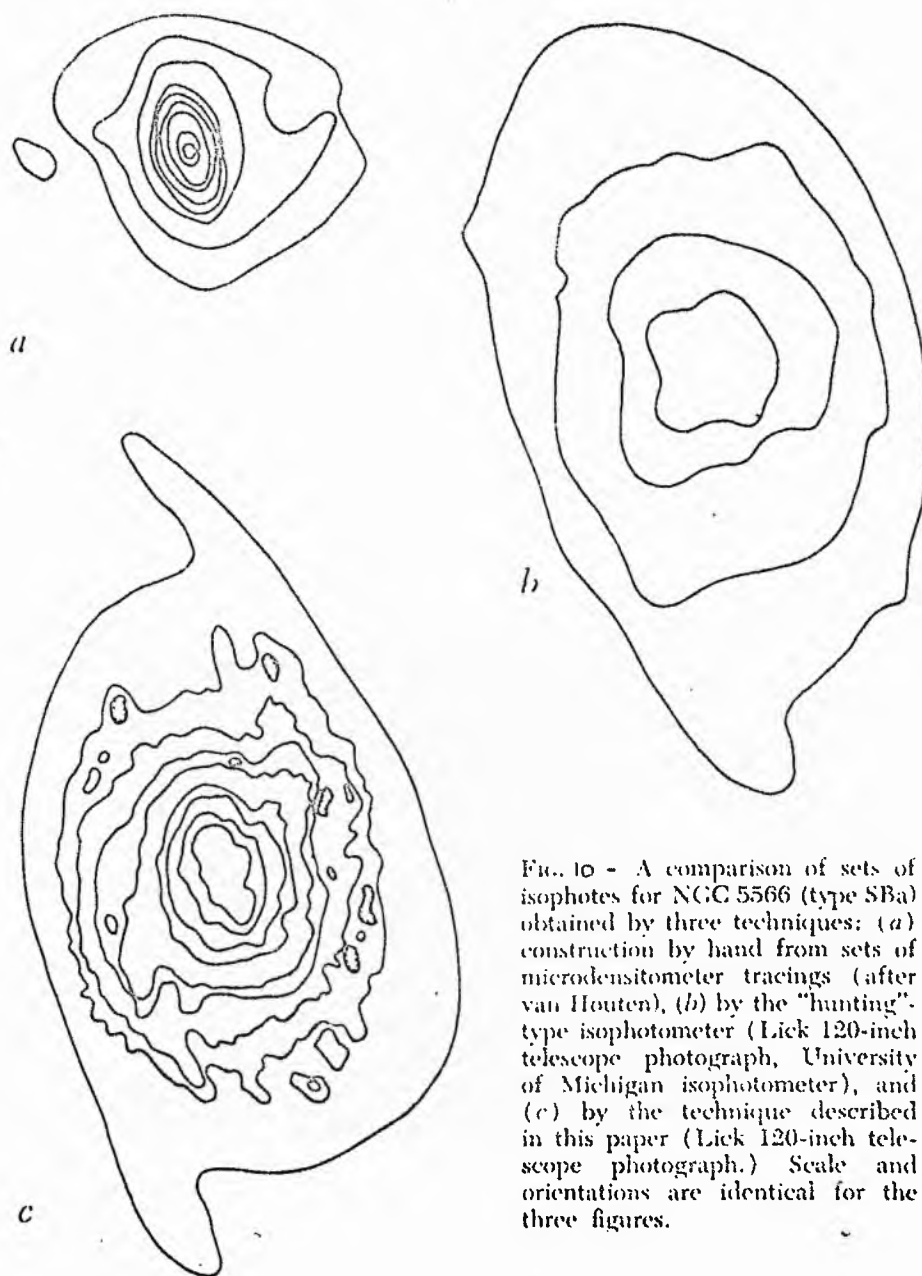


FIG. 10 - A comparison of sets of isophotes for NGC 5566 (type SBA) obtained by three techniques: (a) construction by hand from sets of microdensitometer tracings (after van Houten), (b) by the "hunting"-type isophotometer (Lick 120-inch telescope photograph, University of Michigan isophotometer), and (c) by the technique described in this paper (Lick 120-inch telescope photograph.) Scale and orientations are identical for the three figures.



Figure 11 - Isophotes of faint outer regions
of NGC 5566. The core is over-exposed.

2. Since the recording pen can write in three different modes, the difficulties formerly present in determining the sign of the density gradient are not now present. Clearly no ambiguity is found in cases where closed loop isophotes occur within larger ones.

3. The speed with which the tracings are produced is extremely rapid.

In this connexion, an examination will now be made of the distortions found in the stellar images in figures 5 and 6. The plate was rotated through 180° between each tracing. A useful design feature of the MDM is a differential control. This adjusts the speed of the recording table, which is driven by a servo motor. The control phase of this motor receives signals from an amplifier, the output of which varies inversely with the rate of change of density on the plate. Hence the recording pen will never overshoot in parts of the scan with high density gradients, since the table is then moving at its slowest speed. However, when used as an IDT, this valuable feature was inoperative, and great care was necessary in the selection of the maximum possible scanning speed consistent with the production

of distortion-free tracings. The largest density gradients encountered on a plate are in stellar images, and inspection of these images in figures 5 and 6 readily shows that the scanning speed was too fast, since the distortions are in the same direction. However, if the two tracings are superimposed, the isophotes within the galaxy match perfectly, indicating that although the scanning speed has obviously been too fast to produce symmetrical stellar images, this speed was nevertheless acceptable for accurate isophotometry of galaxies. Therefore it is proposed that so long as the stellar images are reasonably symmetrical, there is no reason for believing that asymmetry could be introduced in isophotes of galaxies. This condition formed the basis for the determination of the maximum speed at which scanning could proceed.

4. The higher accuracy in the detection of faint isophotes. Contrary to the operation of other scanning-type isophotometers where isophotes are obtained corresponding to certain pre-selected densities, the IDT provides a continuous digital record of density changes over the entire scanned area.

Initially the wedge base control is arranged so that the noise in the sky background makes the mode of the recording pen alternate between lines and spaces in approximately equal proportions.

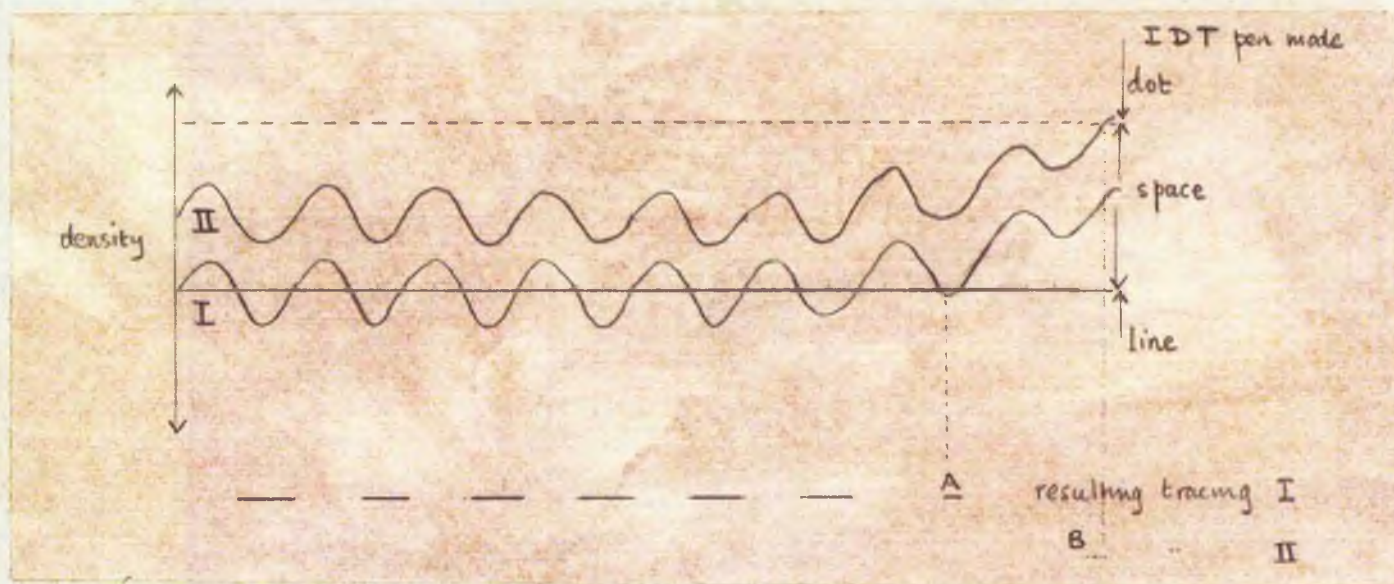


Figure 12 - Diagram illustrating base level setting for background.

In figure 12, the solid line represents the level of the background corresponding with the density transition level at which the IDT pen mode will change from lines to spaces. A typical density tracing is shown by line I while an identical tracing at a higher base control setting is indicated in line II. In case I, a density change in the back-

ground signal considerably smaller than the noise level will produce a marked change in the line to space ratio as indicated at A in the resulting tracing. Considering case II , the first indication of a density change will be shown at B when dots will appear on the tracing. Clearly the first case enables identification of a much fainter isophote level than the second. This system permits more accurate values for the contributions from the faint outer regions to the total luminosity of the galaxy under study, since fewer "average" isophotes derived from the radial scans across the object will be required.

A recent development in isophotometry by Bredohl (1970) has combined the detection of pre-selected density values with rapid speed of electronic raster scanning. A plate of a galaxy is placed in front of a television camera, and as the image is built up on a large screen, an identification mark is added on the screen every time the scanning video signal crosses a pre-selected reference level. This arrangement can produce detailed isophote maps of a galaxy in 10 minutes, but the limitation of the technique is similar to that encountered in equidensitometry, namely that regions of low density gradients tend to produce broad isophotes.

3.4 Smoothing Methods. In addition to photographic and mechanical systems isophotes may also be drawn on graph plotters after computer reduction of data. Various smoothing functions may be applied to the data in order to aid definition of the fainter isophotes and an investigation was made into the comparison between different smoothing techniques and the IDT tracings. The procedure followed for the construction of isophotes was to scan an area of the plate at least twice the apparent size of the galaxy under investigation so as to ensure that each scan extended well into the sky background, and then to sample the data at uniform intervals Δx along each scan to produce a data block for computer processing with on-line graph plotting. The plotting routine involved searching along every grid line forming the boundaries of each individual data block, locating a particular value of plate density and plotting a small straight line to the adjacent grid line thus forming a segment of the isophote corresponding to the chosen density. The shift Δy between successive scans should be less than Δy_0 , the average size of the plate grain, for maximum effectiveness in sampling. In practice, however, this would produce a prohibitive amount of data and it is unfortunately necessary to

use a value of Δy which exceeds the optimum value sometimes by as much as a factor of ten.

The object selected was the E2 type galaxy NGC 4261 with an apparent size of 2.8×3.1 arc min. The scanning aperture used for the microphotometry was 50×50 microns, or 3.8×3.8 arc sec on the plate, and the scans were sampled for data points at every 100 microns in x , with the same value for the stepping increment in y . The total area scanned was 6.3×6.3 arc min for which 2,500 data points were obtained. This preliminary study, carried out in 1967, was one project which could have benefited from the modern digitization facilities available, but the basic results illustrated in the next set of figures would only be changed in minor details, not in general character, by such an improved system. The IDT tracing of NGC 4261 with which the smoothing results will be compared is shown in figure 13. The density increment for the isophotes is 0.06, twice that used for figure 14, which shows output for the same galaxy from computer graph plotter data reduction of the microphotometric scans. The density increment is small so that although the details in the background may be retained, it was necessary to restrict the maximum density plotted in order to avoid confusion

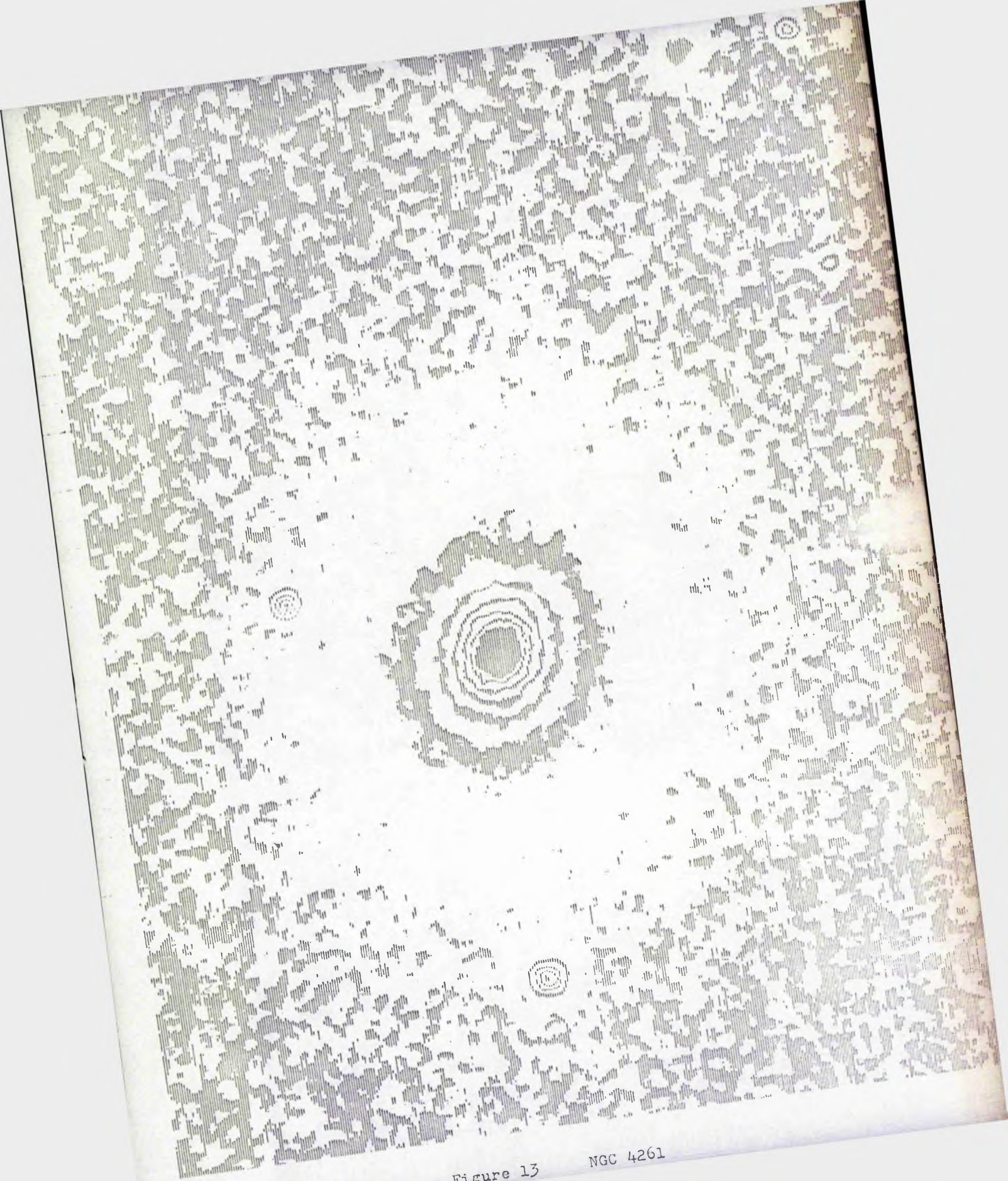
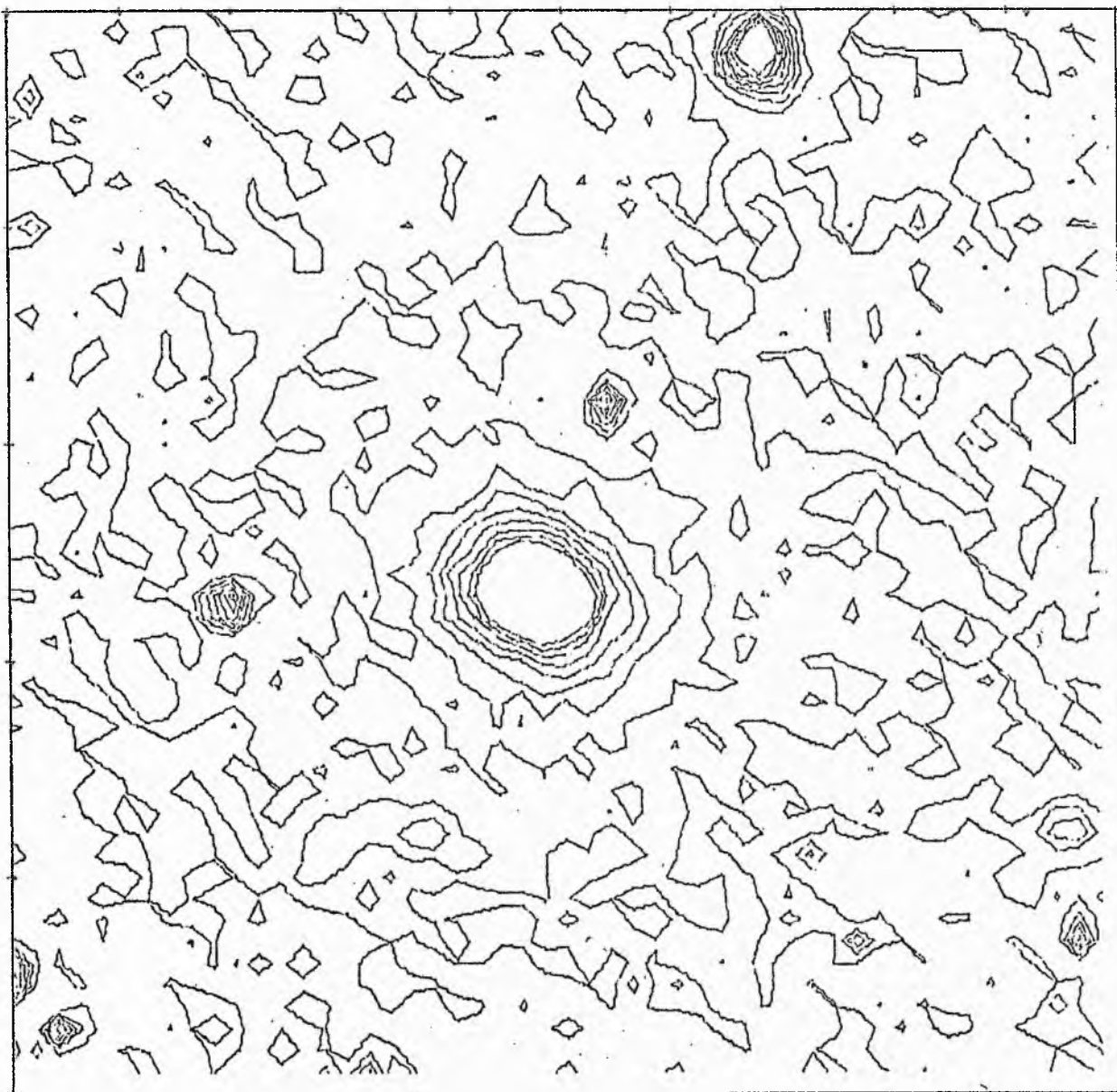


Figure 13

NGC 4261



RAW DATA

Figure 14 Computer output from microphotometric scans across NGC 4261

near the centre. It is clear from the figure that the output contains too much information on the random plate fluctuations compared with the galaxy. It is thus necessary to examine the methods of smoothing the data, especially in the outermost regions of the galaxy, which will remove the plate grain noise but yet retain information on the galaxy profile.

The array of data points in the form of a nine-point cross is shown in figure 15.

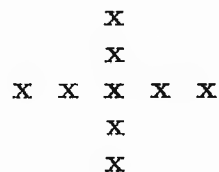


Figure 15 - Diagram of nine-point cross array of data points.

Let us consider the smoothing of the data first by rows and then by columns throughout the data block. It is assumed that the data are sufficiently close together to justify the hypothesis that the second derivative does not change much over the course of a few measurements.

Let us assume that the five horizontal data points of the cross designated by the co-ordinates $x = -2, -1, 0, +1, +2$ lie very nearly on a parabola of second order, represented theoretically by the expression

$$f(x) = a + bx + cx^2$$

By the principle of least squares, we have two sets of normal equations. These are

$$\begin{aligned} 5a + 10c &= \sum f \\ 10a + 34x &= \sum fx^2 \end{aligned}$$

It is the central value at $x = 0$ which is to be corrected, and the theoretical value at $x = 0$ is $f(x) = a$. Solution of the normal equation for a gives the following result for the smoothed value \bar{f}_0 at the centre of the row

$$70 \bar{f}_0 = 70f_0 - 6\delta_x^4 f_0$$

or
$$\bar{f}_0 = f_0 - 3/35 \delta_x^4 f_0$$

where δ_x^4 is the fourth central difference. After scanning the entire data block by rows, we may smooth the new data block of \bar{f}_0 values by columns to give

$$\bar{\bar{f}} = \bar{f}_0 - 3/35 \delta_y^4 \bar{f}_0$$

for the doubly smoothed central value of the nine-point cross. It may be shown that the procedure of smoothing by rows and then by columns is equivalent to smoothing firstly by columns and then by rows. However, this

double smoothing process is likely to be more severe than smoothing simultaneously by rows and columns. The smoothed central value for the nine-point cross for simultaneous smoothing is

$$\bar{f}_o = f_o - 3/53 (\delta_x^4 f_o + \delta_y^4 f_o)$$

Since it would appear to be rather artificial to use raw data in the form of a cross rather than a small block of data, the array shown in figure 16 in the form of a nine-point block was considered.

```

x x x
x x x
x x x

```

Figure 16 - Diagram for data array forming nine-point block.

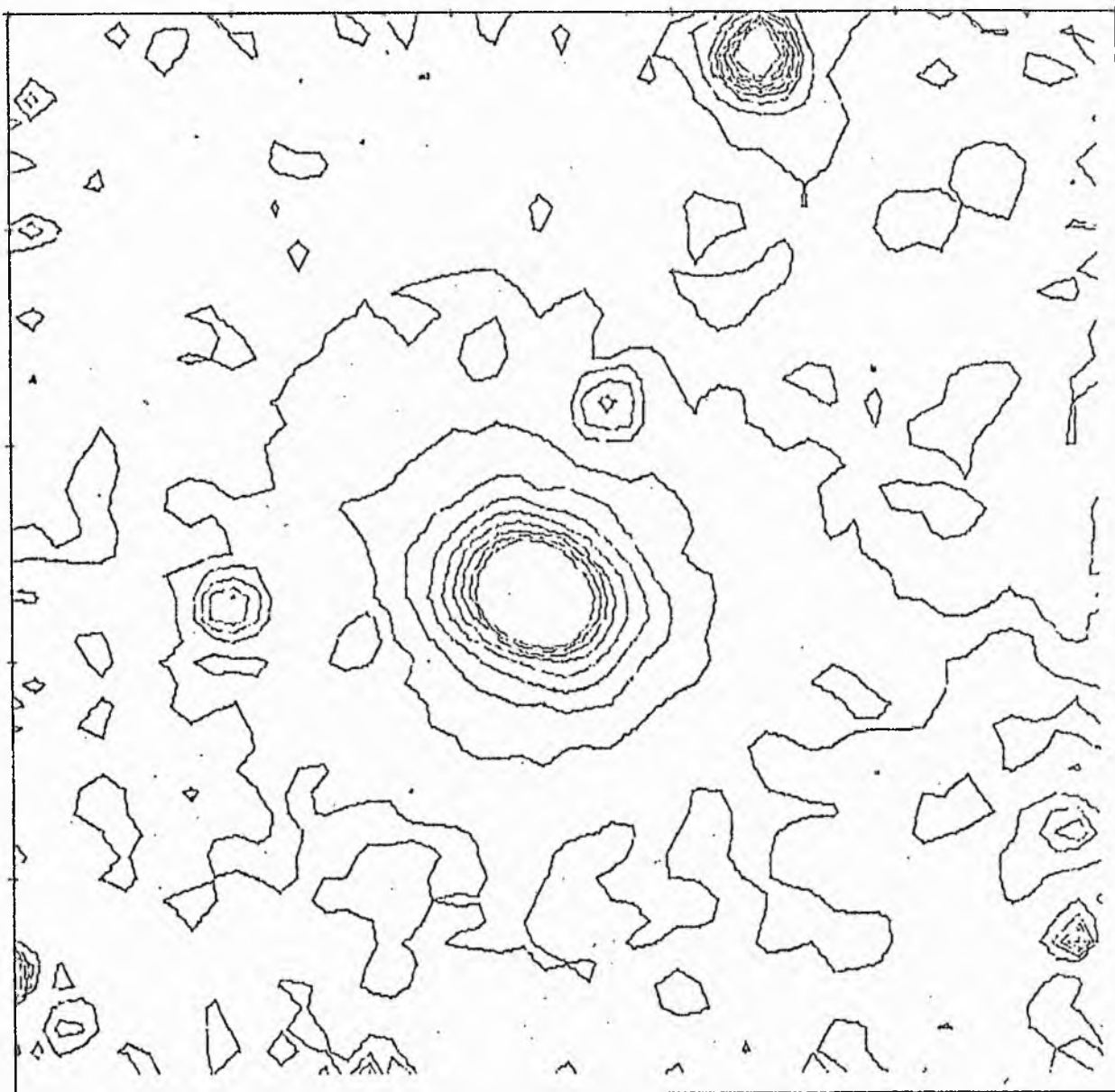
For this array the central smoothed value is given by

$$\bar{f}_o = f_o - 1/9 \delta_y^2 (\delta_x^2 f_o)$$

Figures 17, 18 and 19 show the effects of smoothing by rows then columns. It is apparent that one application of this procedure immediately removes most of the ambiguity in the outer regions of the galaxy. However, successive applications of this process could produce distortions in the nuclear regions due to

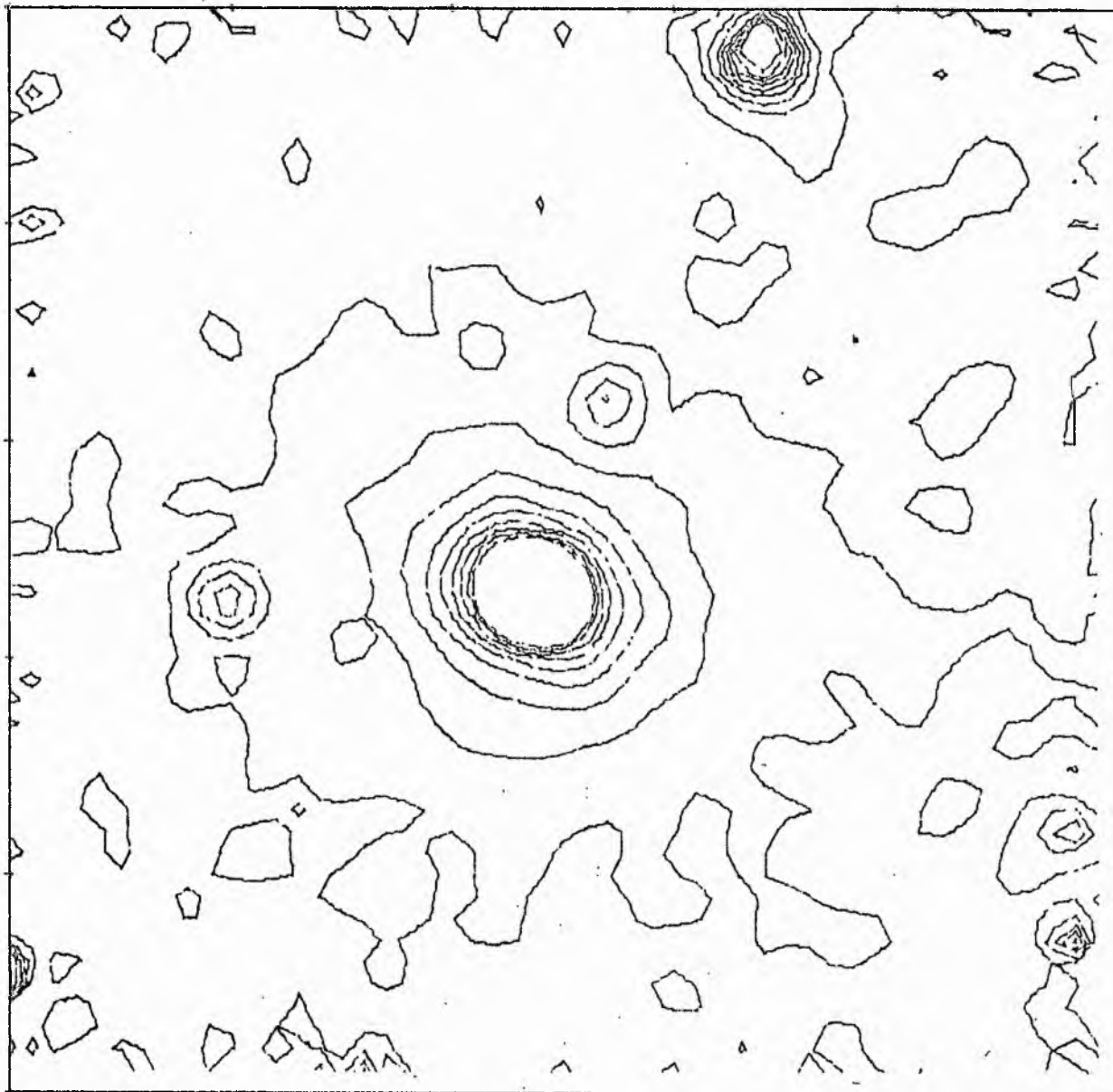
oversmoothing. The effects of simultaneous smoothing by rows and columns in the nine-point cross are illustrated in figures 20, 21 and 22. This procedure is less severe and therefore less likely to produce distortion in isophotal contours. The third case is considered in figures 23, 24 and 25 and it is apparent that quite considerable distortions are introduced by this process, especially in high density gradient areas such as those found in stellar images.

If we compare the results of all the smoothing techniques with the IDT tracing of NGC 4261 illustrated in figure 13, it is apparent that the outermost isophote runs near to, or even intersects the two stars located closest to the galaxy. This dramatic agreement between figures 13, 18, 19, 21 and 22 illustrates the advantages of the smoothing of data. There is, however, a case against smoothing of data based on the view that some important information is lost in the process. For this reason, it would be preferable to preserve all the information as it is done in the IDT tracings. This point is demonstrated in a comparison of the IDT tracings of NGC 4261 (Figure 13) on the two-dimensional map with the computed contours from the unsmoothed data (Figure 14), when it is seen how some confusion and ambiguity is



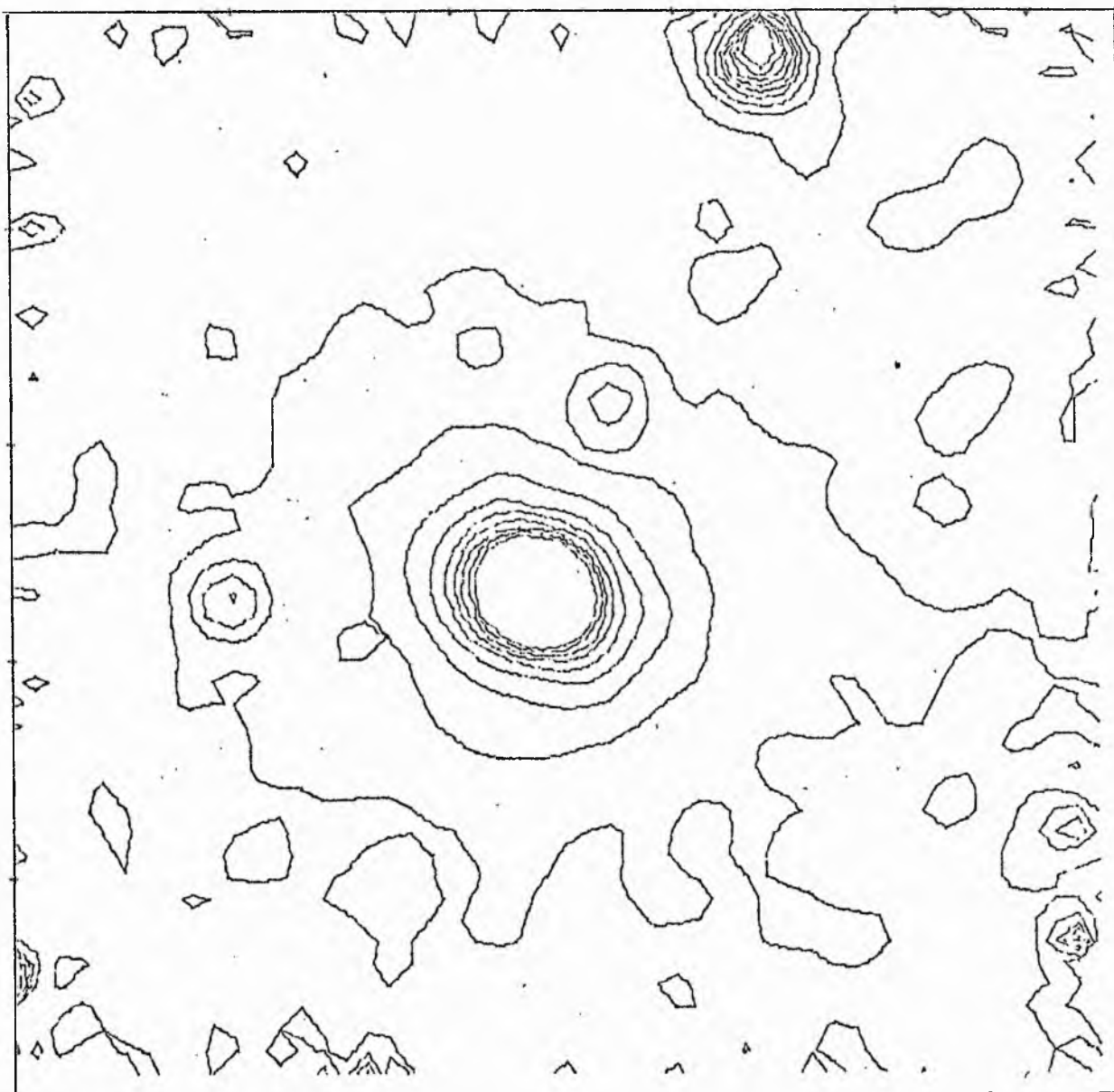
NINE POINT CROSS --- ROWS THEN COLUMNS

Figure 17 NGC 4261



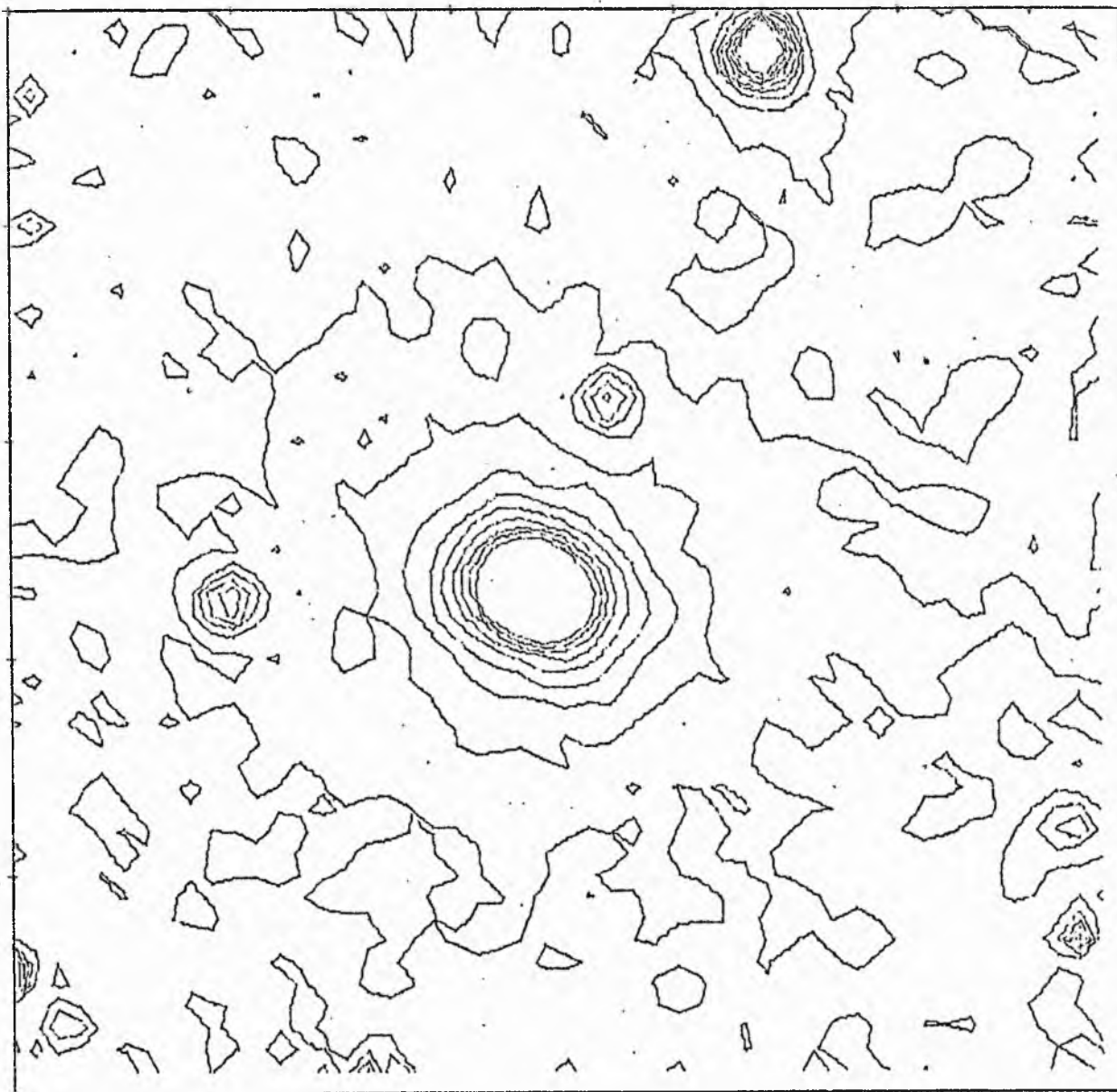
NINE POINT CROSS-ROWS THEN COLUMNS SMOOTHED 2 TIMES

Figure 18 NGC 4261



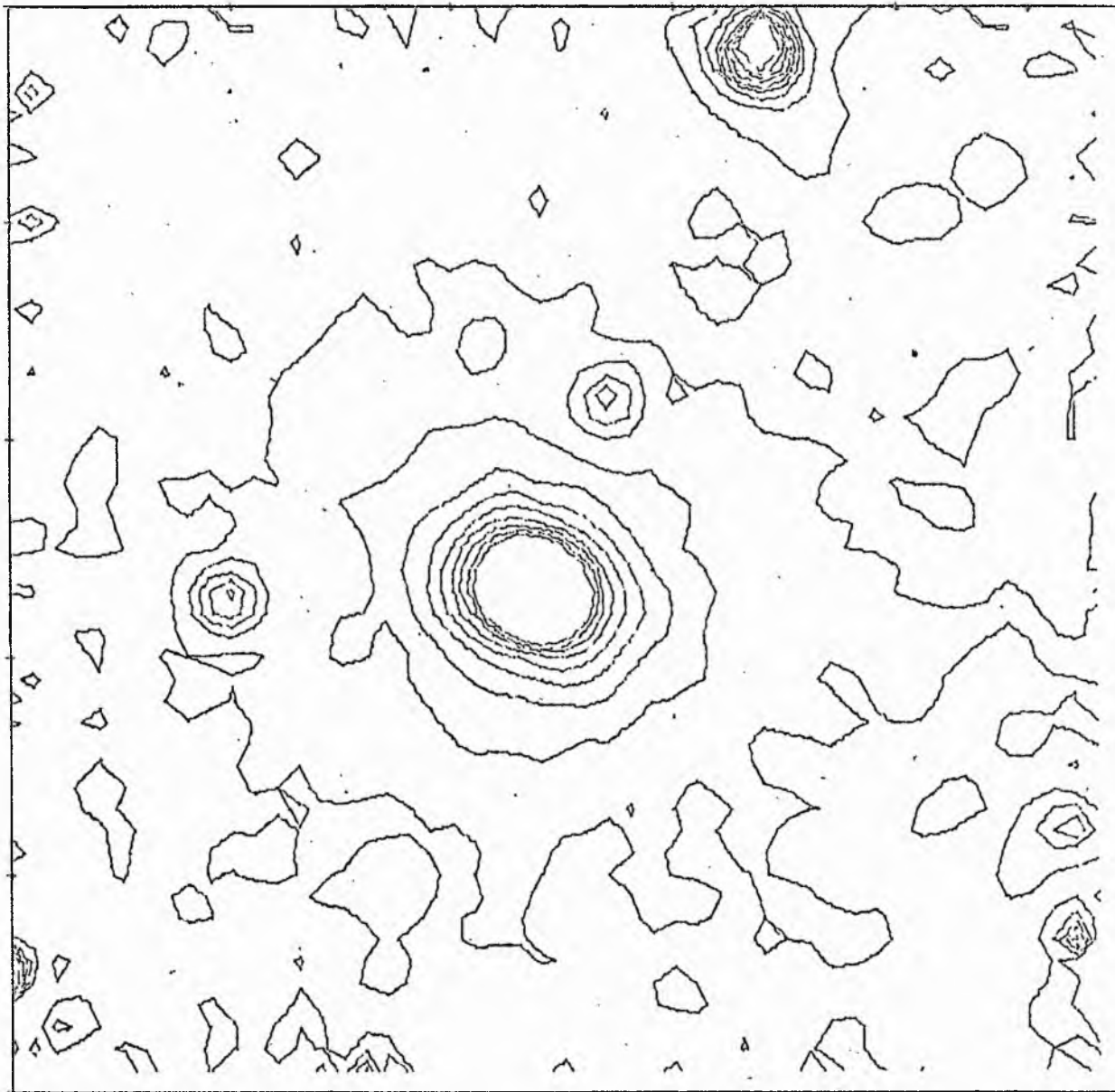
NINE POINT CROSS-ROWS THEN COLUMNS SMOOTHED 3 TIMES

Figure 19 NGC 4261



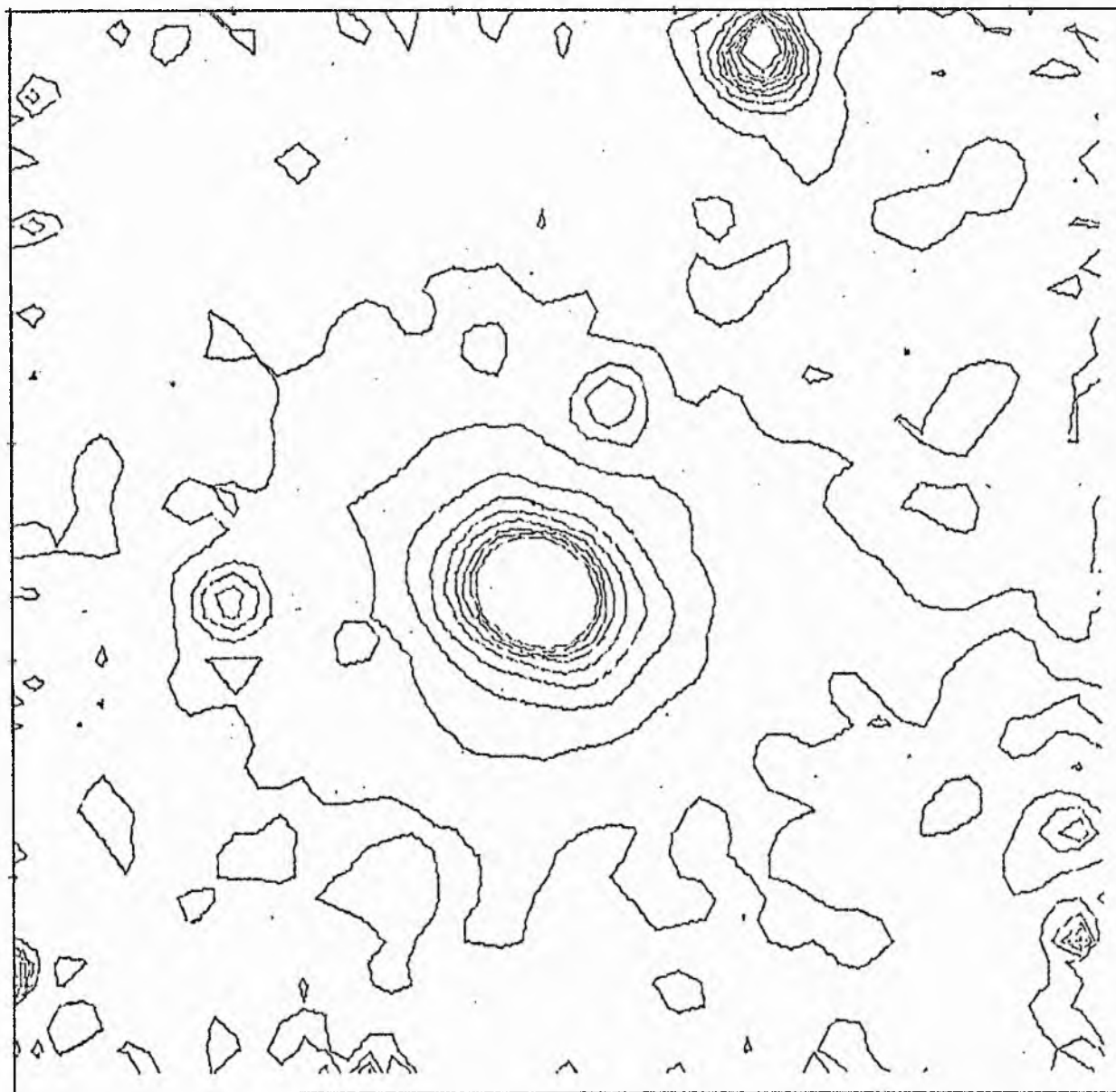
NINE POINT CROSS, SIMULTANEOUS IN X AND Y

Figure 20 NGC 4261



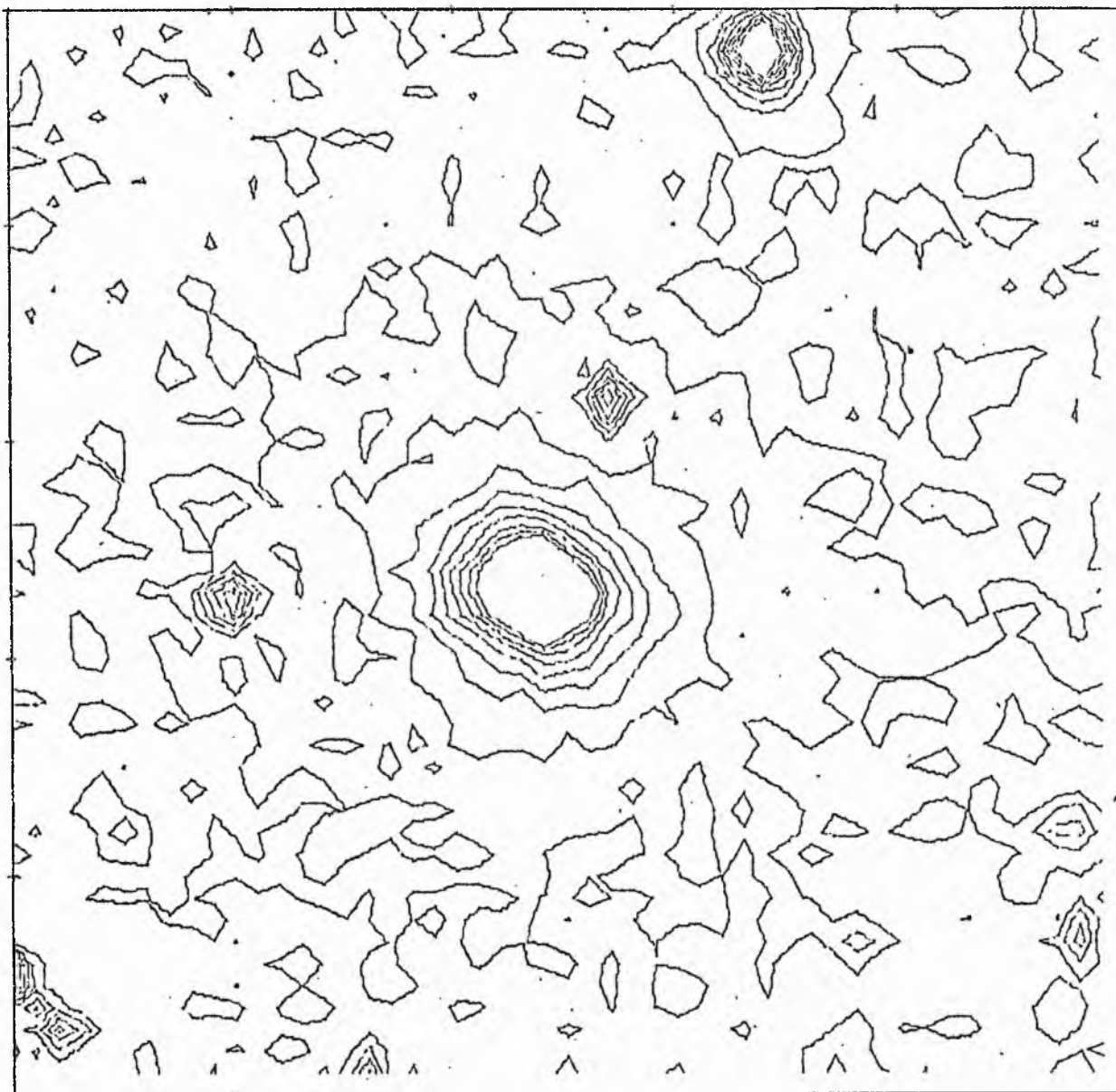
NINE POINT CROSS, SIMULTANEOUS IN X AND Y SMOOTHED 2 TIMES

Figure 21 NGC 4261



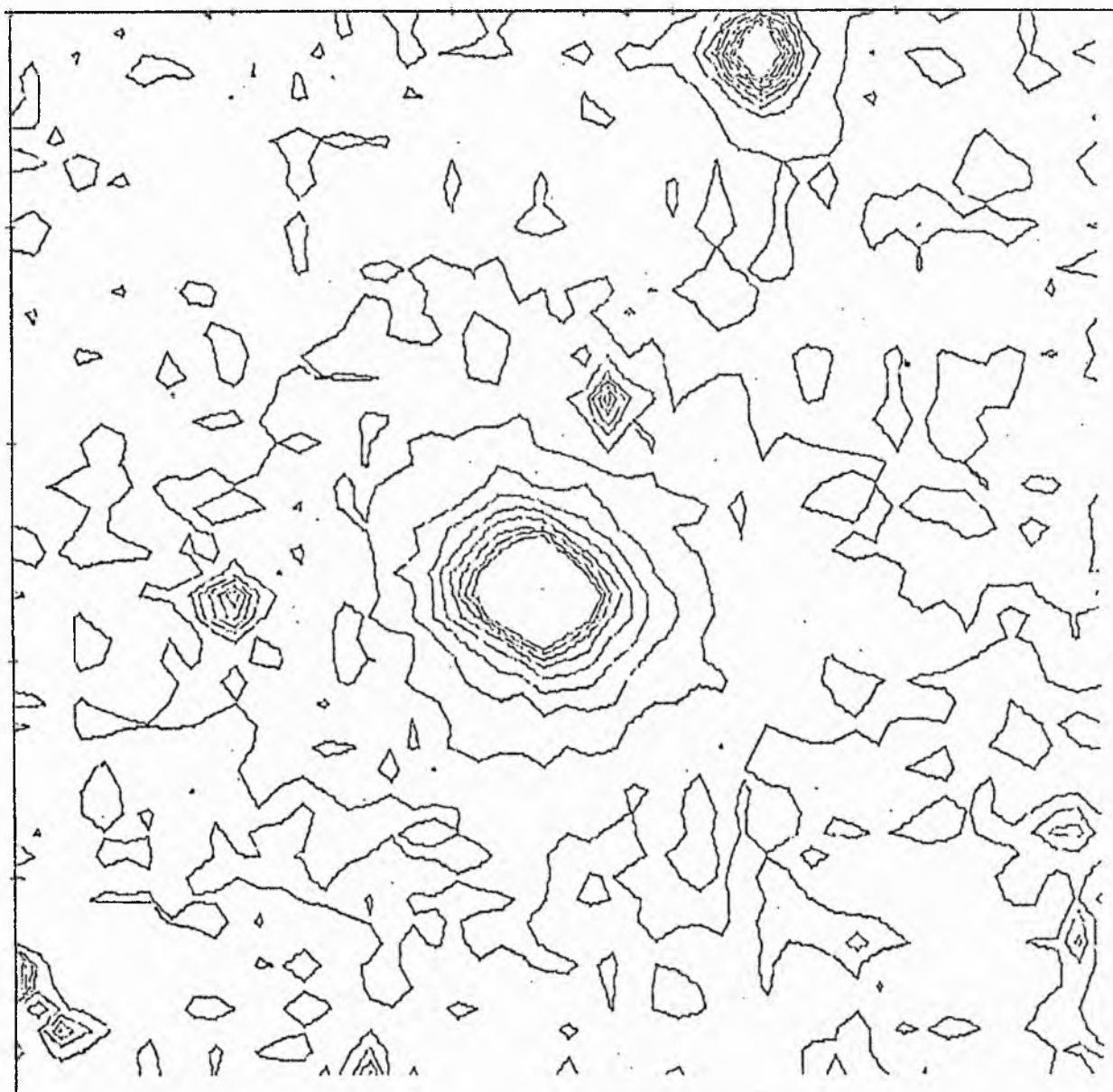
NINE POINT CROSS, SIMULTANEOUS IN X AND Y SMOOTHED 3 TIMES

Figure 22 NGC 4261



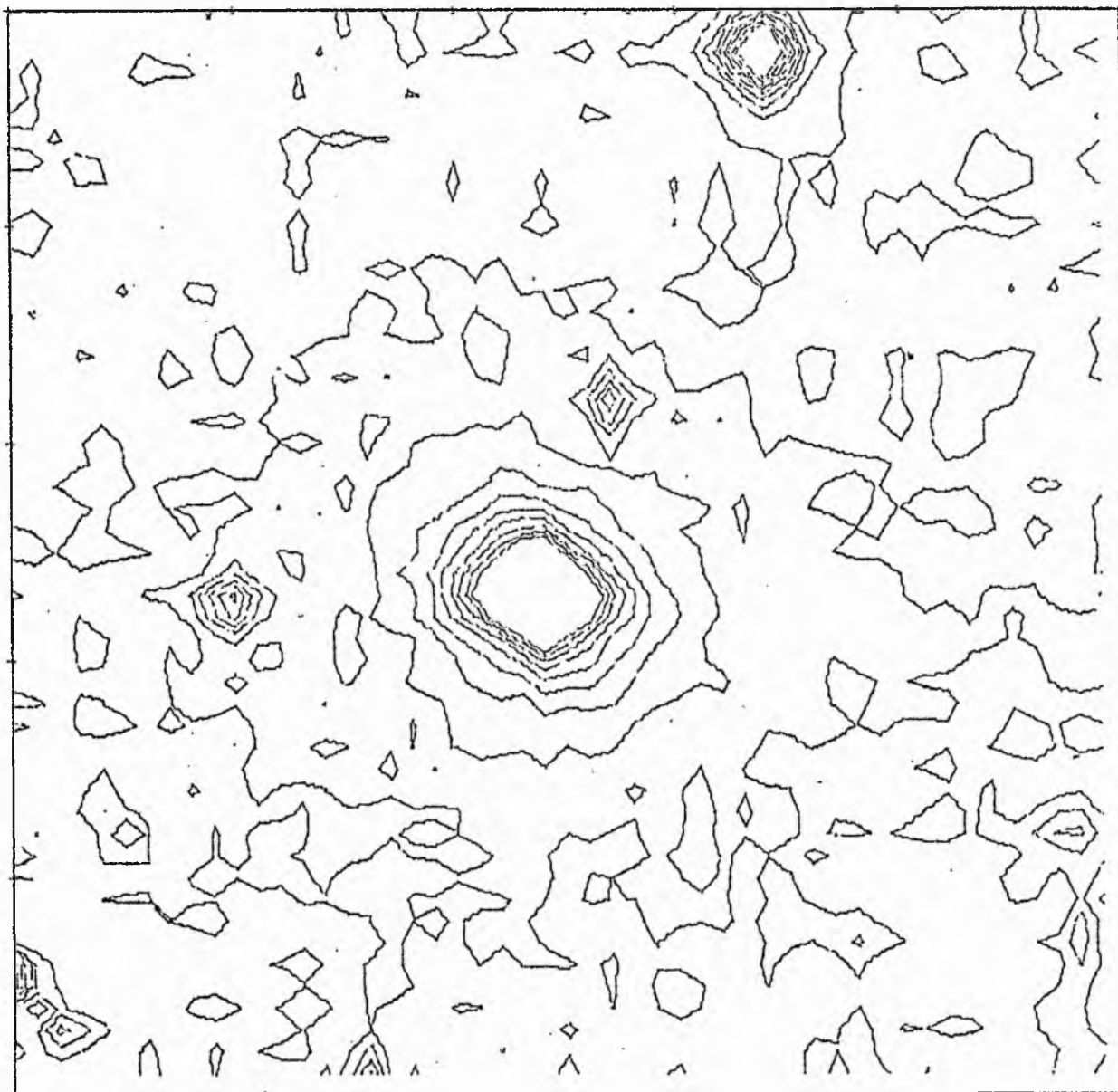
NINE POINT BLOCK

Figure 23 NGC 4261



NINE POINT BLOCK SMOOTHED 2 TIMES

Figure 24 NGC 4261



NINE POINT BLOCK SMOOTHED 3 TIMES

Figure 25 NGC 4261

present in regions of low density. However, the IDT tracing has the advantage that a far smaller stepping increment, of the order of magnitude of the plate grain, was used and all the information was retained and presented visually to the human eye, which is an excellent integrator. The mass of detail is easily averaged and the outermost isophote may easily be located at the changes in the line to space ratio of the background. The agreement with the smoother data for the nine-point cross shows that the principle of smoothing is successful when applied to the fainter isophotes so long as the remaining information is left unsmoothed in order to preserve resolution and detailed structure in the central regions. It is concluded that the coded IDT maps are vastly superior for the isophotometric studies, but in the absence of an isodensitracer, a large amount of useful information may be obtained with the application of appropriate smoothing techniques, with due consideration being given to their inherent nett gains and losses.

This work paralleled a detailed investigation by W. Jones et al. (1967) on numerical mapping techniques. These techniques involved the determination of isophotes from a set of grid measurements. The sky background was

divided into two components, one to examine the overall variations of the background and the other to study point sources which could be removed from the background, such as stars, flaws in the emulsions etc. Least squares fits of different trigonometrical functions were performed on the data, and the smoothing corrections derived in this way were only applied up to a given intensity level, consistent with the conditions indicated in the previous paragraph.

The results of this investigation would appear impressive compared with the work carried out on NGC 4261. However, it should be stressed that the two projects were entirely different in character. On the one hand, the former study attempted to obtain a complete set of isophotes for an object, while the current work at St Andrews was only concerned with the determination of the outermost isophotes.

Chapter IV

4.1 Introduction. The aim of the current work has been to obtain detailed photometric parameters in two colours for a set of galaxies of different morphological types. In the past, workers have published their own parameters which were determined arbitrarily, but with the establishment of an IAU Working Group on Galaxy Photometry, the parameters have now been standardized. It is in these standard parameters that the results of this investigation will be presented. Although some of the parameters have been described in the first chapter of this thesis, the following list gives the full definitions and descriptions of the parameters (de Vaucouleurs 1961c) obtained for each galaxy. This photometric information will be presented in Volumes 2 and 3 of the thesis.

4.2 Parameters. It is convenient to subdivide the definitions for parameters into parts, depending upon galaxy type.

Case 1. Circular Case. Let the mean radial specific intensity distribution be $I(r)$ for a system with circular symmetry, such as an EO galaxy. The integral luminosity emitted within radius \underline{r} from the centre is then

$$L_r = 2\pi \int_0^r I(r) r dr \quad (1)$$

and the total luminosity is

$$L_T = 2\pi \int_0^\infty I(r) r dr \quad (2)$$

giving a corresponding total magnitude of

$$m_T = c - 2.5 \log L_T. \quad (3)$$

The fraction of the total luminosity emitted between 0 and r is

$$k(r) = L_r / L_T \quad (4)$$

and the function $k(r)$ defines the relative integrated luminosity profile of the galaxy, either as a function of r or, through the luminosity profile $I(r)$, as a function of I . The value of L_T is calculated by extrapolation using I as the variable, as outlined by de Vaucouleurs (1960). The effective radius r_e of the galaxy is such that

$$k(r_e) = 1/2 \quad (5)$$

Case 2. General Case. For a system without circular symmetry, such as an elliptical seen edge-on, a spiral, or even an irregular system with detached "islands", an

equivalent radius r^* may be defined by the relation

$$r^* = \sqrt{A/\pi} \quad (6)$$

where A is the total area within a given isophote, including irregular structures. The equivalent luminosity profile $I(r^*)$, the equivalent relative integrated luminosity curve $k(r^*)$ and the equivalent effective radius r_e may then be defined as before.

If the unit of surface brightness of I is calibrated in photographic magnitudes per square minute, the luminosity profile may be placed on an absolute basis of specific intensity and the absolute value of specific intensity I_e along the effective isophote of area A_e corresponding to the equivalent effective radius r_e can be determined. If the absolute value of specific intensity is expressed in magnitude per unit solid angle μ , the value of μ_e corresponding to I_e may be used to calculate the average specific intensity I'_e or μ'_e inside the effective isophote from the relation

$$\begin{aligned} I'_e &= L_e/A_e \\ &= L_T/2A_e \quad \text{from (4) and (5).} \end{aligned}$$

Hence

$$I'_e = L_T/2\pi(r_e^*)^2. \quad (7)$$

The corresponding relation for μ'_e is derived from equation (3) giving

$$\begin{aligned}\mu'_e &= c - 2.5 \log I'_e \\ &= c - 2.5 \log L_T + 5 \log r_e^* + 1.995 \\ &= m_T + 5 \log r_e^* + 1.995. \quad (8).\end{aligned}$$

It is important to notice that the effective parameters are independent of the observing conditions so long as the semi-minor axis is much greater than the instrumental profile. This was one reason for making the selection of galaxies, discussed in section 2.2., for the present investigation.

The luminosity distribution of galaxies can also be expressed in terms of the reduced, or dimensionless variables, as outlined in the first chapter. If $J = I/I_e$ and $\alpha = a/a_e$ or $\rho = r^*/r_e^*$, galaxies of the same type and orientation should be represented by the same reduced luminosity profile $J(\alpha)$ or $J(\rho)$.

In addition to the effective dimension r_e^* defined by $k(r_e^*) = 1/2$, it is in terms of this function that two further dimensions, r_1^* and r_3^* , are introduced in order to define the concentration indices given by the expressions

$$C_{21} = r_e^*/r_1^* \quad \text{and}$$

$$C_{32} = r_3^*/r_e^* \quad (9)$$

where $k(r_1) = 1/4$ and $k(r_3) = 3/4$.

In addition to the characteristic parameters just described, the study of surface photometry within an object is completed with the following supplementary tables and diagrams:-

- (1) A set of isophote maps
- (2) The luminosity profile $\log I(a)$ and $\log I(b)$ along the respective major and minor axes.
- (3) The relative integrated luminosity curve $k(r^*)$ and
- (4) A table of the mean luminosity distribution $I(r)$ and the integrated luminosity $k(r)$ with arguments I and r^* .

4.3 Plate Material. The total number of galaxies with $m < 13$ for which plates have been obtained on the James Gregory Telescope was 106. The dates during which the observations were carried out are shown as follows:-

Dates	Number of Plates taken	Number of galaxies photographed.
March 1965 to May 1965	48	52
November 1965 to April 1966	71	46
November 1966 to March 1967	6	8
Total	125	106

However, for the reasons given in section 2.2, only 48 galaxies with face-on diameter $D(0) > 2$ arc min were chosen for detailed investigation. Observations were made in the international B and V system using the following plate/filter combinations:-

B system	Kodak 103a0 and Schott GG 13
V system	Kodak 103aD and Schott GG 14

All the plates were taken with Hour Angles less than 2^h from the meridian, so that over the range of declinations in the study ($16^\circ \leq \delta \leq 3^\circ$), the maximum and minimum values of $\sec z$ were 1.91 and 1.31 respectively, at the latitude $+56^\circ 20' 10''$ of the University Observatory.

In view of the small f/ratio of the telescope, it was necessary to maintain a strict check on the focus setting during observations. In this connexion, it is

interesting to note the effect on the focus setting of interposing a glass filter in front of the photographic emulsion. At the centre of the plate, the focus is pushed out by $0.2004t$ while at the edge of the field, the coefficient of t is 0.2036 , where t is the thickness of the filter in use. The value of the coefficient is based upon values of the angles of incidence i of the convergent beam, which for the James Gregory Telescope were found to lie in the range $6^\circ \leq i \leq 9^\circ$. It should be noted that with a filter in use, there is a difference in focal setting equal to $0.0032t$ at full aperture between the centre and the edge of the field. This is another reason, in addition to that outlined in section 2.2 for using the minimum aperture of 30 inches.

In order to secure photometric information that is homogeneous, two exposures of 4 and 40 minutes were used throughout the work. The galaxy profiles obtained from the 40 minute exposures were supplemented in the central regions by profile sections derived from the 4 minute plate. The longer exposure, the maximum value possible, was dictated by the blackening of the sky background on the plate due mainly to urban lighting.

Nevertheless, this exposure was still sufficiently large to detect the faint outer regions. The plates chosen for the reduction had passed a very thorough series of tests, based upon image quality, emulsion defects, uniformity of background, and density of the calibration areas etc. One of the most common causes for plate rejection was the considerable difficulty encountered with the telescope sidereal drive during the 1965/66 observing season. Frequent audible bursts of hunting of the drive motor resulted in trailed images before it was possible to end the exposure by moving from the guiding eyepiece to the plate loading assembly area in order to close the dark slide.

4.4 Isophotes. The first stage in the overall reductions involved obtaining detailed B and V isophote maps for each of the 48 galaxies under study. Most of the advantages of the Joyce Loeb1 Isodensitracer have already been mentioned in the previous chapter, but before satisfactory tracings could be produced with regularity, it was necessary to make a few modifications to the instrument. To secure some degree of uniformity, all tracings were made with a density increment of 0.12, a scanning aperture of $50\mu \times 50\mu$ or 3.8×3.8 arc sec., and depending upon the angular diameter of the object,

with a magnification of either 50 or 100 in order to fill the entire recording area. The base-line setting for the background was set up to produce approximately equal line to space ratios in order to assist with the determination of the fainter isophotes. After the initial setting sequence, it was possible to leave the machine unattended to trace the isophote maps automatically. However, three major system faults became apparent in the early tracings and they will be described at this point.

First of all, the recording pen supplied did not dot correctly and only lines and spaces could be seen on resulting tracings. The fault did not lie in the electronic circuits, since the pen was physically going through a dotting sequence, neither was it due to the ink flow of the biro pen since lines were being drawn with ease. After many tests, it was established that the quality of the dots depended upon the velocity with which the pen made contact with the recording paper. Now with no signal input present on the pen solenoid, that is, in the clear mode, and also with the IDT pen assembly lowered as far as possible down to the recording paper, there was still a large gap between the pen tip and the recording paper surface through which the pen had to travel in order to dot. The velocity acquired by the pen when contact

was made with the paper was considered far too large, and an additional spacer was inserted in the solenoid/pen-holder assembly in order to narrow the gap to about half its original value. This modification produced an immediate improvement in the dotting quality of the pen and the resulting tracings were acceptable for detailed measurements.

The second problem was only encountered after approximately the first ten tracings. For no obvious reason, the pen mode would change completely in the middle of a tracing, giving clear areas where there were previously dots, dots where there were lines and lines instead of spaces. Initially, such tracings were immediately stopped and a second trace taken. However, after this fault had occurred several times, a connexion was established between the fault and the lamp failure. The lamp life seemed to average out at about 40 hours, or the time to complete 10 tracings, and the fault was noted to occur about 5 hours before failure of the lamp.

The third problem, repeated breakages of the drive cord, arose as a consequence of the lamp failure. Should the lamp fail during an IDT tracing, the servo-positioning motor controlling the wedge movement tries to position the density wedge at infinite density. Since the maximum

density on the wedge has a finite value, the servo motor pushes the wedge to end-stops, and all the strain is then taken up on the drive cord, which fractures easily after a short time under tension. Such a problem does not arise in MDM operation because the operator is in constant attendance. The difficulty was overcome by fitting a silicon photo-sensitive diode inside the lamp-holder assembly, together with the necessary electronic circuit for switching off the servo amplifier when the diode detected the large change in light intensity at bulb failure.

Returning to the second problem it seemed strange that such sudden density jumps were present, especially since the optical system was designed to take account of any intensity changes in the reference supply. However, let us consider the effect of a slow variation in the intensity of the reference beam, while the measuring beam intensity remains constant. On IDT work, this variation would have remained unnoticed but for the obvious error in the visual display on the tracing. Of course, the sudden jump on the tracings would have been due to the magnitude of the difference in light output being by this stage large enough to move the commutator ring relative to the pen contacts, described on figure 7 in section 3.3., from the centre of the area in which they were originally

set up over to an adjacent area. However, having described in detail the effect, the cause must still be sought for the apparent variation in the reference source intensity. The clue to the solution lay in the very fact that such density jumps were found near the end of the usable life of the lamp. Figure 26 shows the plan view of the initial stages of the optical paths.

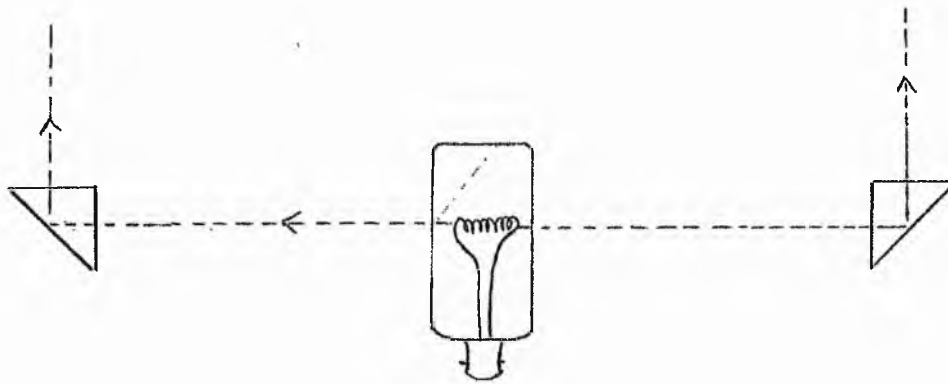


Figure 26 - Lamp and prism optical arrangement in Joyce Loeb1. Although both prisms reflecting the respective reference and measuring beams are receiving light from an almost point source, namely the ends of the lamp filament, this point source is still different for each beam in that it is separated by the length of the filament. Now the effect of a variation of light output across the filament itself would be to produce a change in the intensity of the reference beam with respect to the measuring beam, thus removing the advantages of the long-term stability

found with the single lamp system. It is proposed that it is local variations in the resistance of the lamp filament which produced the different light outputs, and such resistance variations would be expected near the end of a lamp's working life.

The lamp fitted to the Joyce Loeb1 was originally a tungsten lamp with an average working life of 40 hours. Around the time of delivery of the instrument to the University Observatory, quartz-iodine lamps were becoming increasingly popular, in view of their high light output and apparently long life. Furthermore, when used as standard fittings in motor car headlights, they must be able to withstand multiple on-off switchings, something which is not recommended for the tungsten projector lamps. The lamp housing assembly was modified to accept the slightly different mounting of a quartz-iodine lamp, which was then fitted. In view of the relatively high light output of quartz-iodine compared with a tungsten source, this new lamp was slightly under-run, and this had the effect of prolonging its working life. Even the filament of the quartz-iodine lamp had much smaller dimensions than its tungsten counterpart. Tracings obtained since this modification have been completely free from drift problems, indicating that the tungsten lamp

filament must have suffered from resistance variations. The new quartz-iodine lamp has been in use now for over 1000 hours, and although the photo-diode lamp-failure detection circuit has not been required, it is a source of encouragement that this is the case.

The isophote maps produced were all measured with a planimeter, and appropriate scaling factors were used to transform the areas of each isophote to angular measurements. The planimeter was fitted with a special magnifying tracer to facilitate accurate measurements. From the isophote map, the major and minor axes are determined by plotting the length of the radius vector for each isophote against its corresponding position angle. The points of intersection with the principal axes are measured for each isophote to determine its ellipticity. Where the isophote intersects the axis in more than one point, the average value is taken. It now remains to calibrate each isophote, and this is done by means of microphotometric cross-sections.

4.5 Cross-sections and calibrations. Microphotometer scans were made in four directions across the centre of each galaxy, two directions being along the principal axes, and the other two at 45° to these axes. In view of the recent digitization of the microphotometer, almost all the

reductions were carried out at the University Computing Laboratory, in which an IBM 360 Model 44 is available. At first, the scans were made with the same aperture as that used for the isophote maps, $50\mu \times 50\mu$, and they extended well beyond the measured major axis of each galaxy. However, after preliminary results on all the galaxies became available, some of the values for the effective radii seemed to be rather low. It became clear during discussions with Dr G. de Vaucouleurs, who visited the University Observatory prior to the 1970 IAU General Assembly in Brighton, that the scans may not have extended sufficiently far into the background and as a result, the sky background density level was too high, giving too low a value for the total luminosity of the object. Accordingly, all the original scans were re-run following the IAU meeting, but on this occasion the total length of scan was 31 arc min, giving an extension from the nucleus of 16.5 arc min into the background; the scanning aperture was $200\mu \times 50\mu$ or 14.2×3.8 arc sec on the plate. These large scans enabled a much more precise assessment to be made of the sky background level. Measurements were made of the densities above the sky background at known

radial distances greater than those for which values had originally been obtained and this information, together with the revised sky background value was used to supplement the original digitized measurements.

The relative calibration was carried out by making scans across the wedge exposures with a projected scanning aperture equal to 0.4×0.1 mm. Since the calibration "curve" was being constructed in the linear form following de Vaucouleurs (1968), it was necessary to convert the densities measured on the tracing back to opacities. The continuous density wedges supplied with the MDM are of the linear type and have known factors for determining the relationship between the density of the wedge and distance on the recording paper. These MDM wedge factors were checked on the Hilger L470 machine and were found to be accurate within the limits of the calibration graphs supplied with each wedge. It is thus only necessary to measure distance on the tracings above a fixed reference level and multiply this distance by the wedge factor to ascertain the density level.

Let us suppose a particular step of the calibration wedge gives a density above the fog level of the plate $D = \log (\Delta_o / \delta)$ where Δ_o would be the equivalent deflection produced by the fog level in a transmission

measuring system and δ is the corresponding deflection produced by scanning across the step. Then the opacitance for this particular step is given by

$$\omega = \left(\frac{\Delta}{\delta} - 1 \right)$$

A graph of $\log \omega$ versus $\log I$ for all steps on the wedge exposure yields the value of n in the following relation for the characteristic curve

$$I = A\omega^n$$

At the same time as the calibration exposures were made, scans were taken at least 20 arc min to the north, south, east and west of the galaxy in order to determine the level of the sky background relative to the fog level on the plate. The average produces the required density level of the background D_s .

The reduction procedure for the cross-sections and isophote calibration was as follows:-

- 1) For all the digitized values in each scan, it was first necessary to transform the voltage measurements to density. Results of separate tracings produced a value of 0.2884 for the conversion factor for changing voltage to density. Since the least significant digit on the digital voltmeter measuring density was one hundredth

of a volt, the accuracy of the density measurement was 0.0028, or better than one per cent in transmission.

2) Each density value in its present form was related only to the arbitrary initial voltage setting on the digital voltmeter before the digitization scan commenced. A zero point conversion had to be applied to convert these arbitrary density measures to known densities above the plate fog. The average value of the first ten digitized points of the scan was calculated. At the same time as the digitization scans were in progress, a microphotometer tracing was also produced for checking purposes in the subsequent reductions. The average density level of that part of the tracing corresponding to the first ten scan values was determined and any necessary adjustment between this level and the true background level as determined from the later scans was noted. After applying this minor adjustment, the difference was calculated between the arbitrary density value of the background found from the digitized scans and the true density level D_s determined during the calibration traces. This difference was applied to every digitized density value to make the new values correspond with correct densities above the plate fog level.

3) Every density value was converted to a corresponding intensity above the plate fog level using the linear calibration curve. The intensity value of the sky background corresponding to density D_s is equal to i_s and the intensity of a point in the cross-section with a corrected digitized density value equal to D_{g+s} above the fog level is given by i_{g+s} , made up of an intensity i_s due to the sky background plus the intensity contribution i_g from the galaxy itself.

4) The next step is to assign to each relative intensity value on the digitized scan a position value relative to the centre of the galaxy. The precise position sampling interval during digitization was known in angular measure, and it only remained to find the centre of the cross-section from the array of digitized values. The radial distances r of each point from the centre were then computed, and with information on the intensity, it was then possible to output from the computer the luminosity profiles $\log I = f(r)$ either in numerical form on the line-printer or in graphical form on the plotter.

5) The intensity in the luminosity profile is given by the expression

$$I = \frac{i_{g+s}}{i_s} - 1$$

i.e. using the sky background intensity as a unit.

6) The values for the abscissae of the points of intersection of isophotes along each axis (+a and -a for major axis, $\pm b$ for minor axis, $\pm c$ and $\pm d$ for the two axes at 45°) were punched on cards and loaded in the computer. Interpolations on the luminosity profile appropriate to the axis gave values for $\log I$ for the abscissae of each intersection of each axis with each isophote. The mean values were taken for each axis, then for each isophote and the $\overline{\log I}$ value obtained is plotted against $r^* = \sqrt{A/\pi}$ where A is the area measured by planimetry of the isophotes. A smooth curve drawn through the points yields that part of the equivalent luminosity profile within the outermost isophote. The remaining section is found after measuring on the luminosity profile graphs the abscissae where $\log I = -1.0, -1.2, -1.4, -1.6, -1.8$ and -2.0 . The equivalent radii were given by

$$r^* = \sqrt[4]{\bar{a} \cdot \bar{b} \cdot \bar{c} \cdot \bar{d}}$$

4.6 Integration. Values of r were measured from the equivalent luminosity profile for intervals in $\log I$ of 0.1 down to a threshold level of $\log I = -2.0$. The

luminosity L_r within a radius r was found from the following expression

$$L_r = \pi \sum_{i=1}^n (r_{i+1}^2 - r_i^2) I_{i+\frac{1}{2}} \quad \text{where } i \text{ denotes the isophote under consideration.}$$

This expression may also be written in the form

$$L_r = \pi \sum_{i=1}^n P_i \quad \text{where each value of } P \text{ is the luminous flux in an annulus of width } (r_{i+1} - r_i).$$

The selection of the rather small integration interval of 0.1 in $\log I$ produced a negligible error in the calibration of the total luminosity. This choice was based on results of a study by de Vaucouleurs (1948), in which it was found that the systematic errors which could be introduced in $\sum P$ increased approximately with the square of the integration interval.

The total luminosity L_T is found by computing the extrapolation correction ΔP which allows for the contribution of the outermost parts of the galaxy below the threshold. The correction ΔP is obtained by plotting $\log (P/\Delta I)$ versus $\log \bar{I}$ and this curve is extrapolated to a value of \bar{I} midway between $I_1 =$ threshold value and $I = 0$ corresponding to $r = \infty$, i.e. $I_1/2$. With a threshold value of $\log I = -2.0$, this gives $I_1 = 0.01$ and $\bar{I} = .005$. L_T is given by the expression

$$L_T = \sum P_i + \Delta P$$

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Calculations of L_T are shown in the appendix where the integration tables of the mean luminosity distribution for each galaxy have been listed.

The absolute calibration has not been made with some of the unnecessarily complicated methods used in the past, e.g. extra-focal stellar images or north polar sequence transfer. Although these methods may have been successfully used by Holmberg (1958) and Markaryan (1965), they are nevertheless subject to many sources of error, such as variable absorption and changes in temperature and humidity. With the large number of photo-electric measurements listed in the Reference Catalogue (de Vaucouleurs and de Vaucouleurs 1964), it was possible to relate measurements of the integrated magnitude of a galaxy within some arbitrary fraction of the total flux received through field apertures of diameter $2r$ placed in the focal plane of a telescope. If m_r is the tabulated magnitude in the B.G.C. with a given aperture radius r and L_r the integrated luminosity within the same radius, the relation

$$m_r = \text{constant} - 2.5 \log L_r$$

defines the conversion constant between relative and specific intensities (e.g. de Vaucouleurs et al. 1968). Several apertures were desirable to determine the constant more accurately, since both L_r and m_r are subject to errors. Values of L_r corresponding to each value of m_r and r listed in appendices II and III of the B.G.C. are obtained from the integration table of mean luminosity distribution, and for each pair of entries in the catalogue, a separate constant is produced.

Chapter V

5.1 Introduction. The photometric data obtained in the current investigation are listed in the subsequent volumes of the thesis. However, before a discussion of these results is given, it is first necessary to examine some of the more important sources of error in the work.

5.2 Errors. Errors may be divided into two categories. First of all, there are errors arising from those sections of the work concerned only with relative measurements, and secondly, there are errors arising from the photo-electric calibration process.

In the first case, many of the error sources introduced in the initial stages of the investigation by the use of the conventional H-D characteristic curve have been removed by the use of the 'linear' curve (de Vaucouleurs 1969). In addition, there are errors introduced when the sky background is not well-defined; these errors have been investigated by de Vaucouleurs (1948). Not only will the value of the total magnitude be affected by such poor determinations, but also the values of the parameters involving distance, such as the major axis at threshold, effective radius etc. Although Evans (1952) has already shown that,

depending on the magnitude of the sky background upon which the galaxy is superimposed, a background level error of $0.^m02$ could introduce an error of $0.^m15$ in the total magnitude of the galaxy, it would, nevertheless be interesting to convert the $0.^m02$ background magnitude difference into density on a typical microphotometer tracing and then compare such a density value with the typical conversions applied to the present data after the reduction of the additional large radial distance scans made following the IAU. Before proceeding, it is necessary to know the density value of the background to which the $0.^m02$ difference is to be applied. The usual density for the sky background on a 40^m exposure plate was about $0.^45$ above the plate fog level. This corresponds to $\omega = 1.818$, and $I = 0.1291$ if $n = 0.5$ in the equation

$$I = A \omega^n$$

A magnitude difference of $0.^m02$ leads to a new value of $\log I = 0.1321$ giving $\omega = 1.864$, equivalent to a density above the plate fog of $0.^457$. Thus a magnitude variation in the galaxy of $0.^m15$ would be produced by a density difference of only 0.007 .

Another approach to the problem may be taken by

using numerical values for the intensities obtained from the integration tables of two luminosity distributions. With the largest density wedge supplied with the MDM, the density of 0.007 is equivalent to 0.54 mm on the chart. Tables 1 and 2 show intermediate computer reductions for the mean luminosity distribution of NGC 4486 prior to carrying out the photo-electric calibration. The first table gives the results obtained in the summer of 1970 while table 2 shows the final results for the luminosity distribution using the revised value for the sky background. The ratio of the respective total luminosities L_T in tables 2 and 1 is 1.64, corresponding to a magnitude difference of 0.41. With the large scanning aperture used in the additional cross-sections, it is a simple matter to measure, on the partly smoothed profile for those radial distances at which the erroneous background was taken, the density above the revised background level; in the case of NGC 4486, the density difference was found to be 0.18. Using the magnitude difference for this galaxy of 0.41^m from tables 1 and 2, equivalent to a background magnitude difference of 2.74 times Evans' value, the corresponding density difference should be 0.19. Although the actual value was 0.18 or 1.39 mm., the agreement

Table 1

MEAN LUMINOSITY DISTRIBUTION IN NGC 4486
B COLOR

LOG I	I	MEAN I	N	AREA	DEL A	P	SIGMA P	K(R)	RO	LOG J
1.47	29.512	27.315	0.0	0.0	00.81	1879.5303	C.C	0.0	0.0	1.593
1.40	25.115	22.534	4.68	68.81	81.20	1829.8232	1879.530	C.030	0.077	1.523
1.30	19.953	17.971	6.61	150.61	85.60	1532.3118	3709.354	C.060	0.114	1.423
1.20	15.845	14.219	8.66	235.61	108.78	1546.7246	5241.664	C.084	0.142	1.323
1.10	12.585	11.295	10.47	344.38	206.33	2330.4055	6788.387	C.109	0.172	1.223
1.00	10.000	8.972	13.24	550.71	215.79	1935.9453	9118.789	0.147	0.218	1.123
0.90	7.943	7.126	15.62	766.50	172.66	1230.4458	11054.734	C.178	0.257	1.023
0.80	6.310	5.661	17.29	939.16	332.60	1882.7546	12285.180	0.198	0.284	0.923
0.70	5.012	4.446	20.12	1271.76	401.72	1806.3184	14167.934	0.228	0.331	0.823
0.60	3.981	3.572	23.08	1673.48	532.70	1902.6340	15974.250	0.257	0.379	0.723
0.50	3.162	2.837	26.50	2206.18	683.79	1939.9692	17876.883	0.287	0.436	0.623
0.40	2.512	2.254	30.33	2889.98	644.10	1451.5198	19816.852	C.319	0.499	0.523
0.30	1.995	1.790	33.54	3534.07	1437.33	2572.9270	21268.371	C.342	0.551	0.423
0.20	1.585	1.422	35.78	4971.41	1710.93	2432.1786	23841.297	0.383	0.654	0.323
0.10	1.259	1.129	46.12	6682.34	1966.78	2221.3994	26274.074	C.422	0.758	0.223
-0.00	1.000	0.897	52.47	8649.12	2316.44	2078.2103	28495.473	0.458	0.863	0.123
-0.10	0.794	0.713	55.08	10965.55	3775.58	2690.6335	30573.687	0.492	0.971	0.023
-0.20	0.631	0.566	68.50	14741.14	5713.37	3234.1687	33264.320	C.535	1.126	-0.077
-0.30	0.501	0.450	80.69	20454.50	5515.30	2479.9338	36498.488	0.587	1.327	-0.177
-0.40	0.398	0.357	90.92	25969.80	6039.50	2157.1091	38978.422	0.627	1.495	-0.277
-0.50	0.316	0.284	100.94	32009.31	6524.09	1650.9351	41135.527	0.661	1.659	-0.377
-0.60	0.251	0.225	110.75	38533.40	6962.24	1568.9890	42986.461	0.691	1.821	-0.477
-0.70	0.200	0.179	120.34	45495.64	7311.76	1308.8594	44555.449	C.716	1.978	-0.577
-0.80	0.158	0.142	125.65	52807.40	7629.58	1084.8550	45864.309	0.737	2.131	-0.677
-0.90	0.126	0.113	138.70	60436.98	19485.89	2200.8579	46949.160	0.755	2.280	-0.777
-1.00	0.100	0.090	155.50	79922.87	13234.25	1187.3303	49150.016	0.790	2.622	-0.877
-1.10	0.079	0.071	172.20	93157.12	19658.25	1400.9341	50337.344	0.809	2.831	-0.977
-1.20	0.063	0.057	189.50	112815.37	13477.44	762.9226	51738.277	0.832	3.115	-1.077
-1.30	0.050	0.045	200.50	126292.81	33741.44	1517.1807	52501.199	C.844	3.296	-1.177
-1.40	0.040	0.036	225.70	160334.25	31320.87	1118.6860	54018.379	0.869	3.710	-1.277
-1.50	0.032	0.028	248.80	191355.12	38516.00	1092.7385	55137.062	C.887	4.057	-1.377
-1.60	0.025	0.023	270.50	229871.12	46313.37	1043.7156	56225.801	C.904	4.447	-1.477
-1.70	0.020	0.018	296.50	276184.50	46520.69	832.7649	57273.516	0.921	4.874	-1.577
-1.80	0.016	0.014	320.50	322705.19	44748.06	636.2839	58106.277	0.934	5.269	-1.677
-1.90	0.013	0.011	342.00	367453.25	43808.25	546.5945	58742.559	0.945	5.622	-1.777
-2.00	0.010		375.00	451261.50			59689.152	C.960	6.231	-1.877
							62189.000	(1)		

Table 2

REVISED MEAN LUMINOSITY DISTRIBUTION IN NGC 4486
B COLOUR

LOG I	I	MEAN I	R	AREA	DEL A	P	SIGMA P	K(R)	RO	LOG J
1.43	26.915		0.0	0.0			0.0	0.0	0.0	1.833
1.40	25.119	26.017	4.33	58.90	58.90	1532.4402	1532.440	0.02	0.04	1.803
1.30	19.953	22.536	7.19	162.41	103.51	2332.5947	3865.035	0.04	0.07	1.703
1.20	15.849	17.901	9.32	272.89	110.48	1977.6392	5842.672	0.06	0.09	1.603
1.10	12.589	14.219	11.69	429.32	156.43	2224.3042	8066.973	0.09	0.11	1.503
1.00	10.000	11.295	14.10	624.58	195.26	2205.4043	10272.375	0.11	0.14	1.403
0.90	7.943	8.972	16.05	809.28	184.70	1657.0632	11929.437	0.13	0.16	1.303
0.80	6.310	7.126	18.60	1086.86	277.58	1978.1682	13907.605	0.15	0.18	1.203
0.70	5.012	5.661	20.78	1356.57	269.70	1526.6997	15434.305	0.17	0.20	1.103
0.60	3.981	4.496	23.90	1794.51	437.94	1969.1885	17403.492	0.19	0.23	1.003
0.50	3.162	3.572	27.26	2334.54	540.03	1928.8147	19332.305	0.21	0.27	0.903
0.40	2.512	2.837	31.17	3052.27	717.73	2036.2593	21368.562	0.23	0.31	0.803
0.30	1.995	2.254	35.99	4069.24	1016.97	2291.8044	23660.363	0.26	0.35	0.703
0.20	1.585	1.790	42.46	5663.82	1594.58	2854.4138	26514.773	0.29	0.42	0.603
0.10	1.259	1.422	48.87	7502.99	1839.17	2615.1267	29129.898	0.32	0.48	0.503
-0.00	1.000	1.129	54.83	9444.66	1941.66	2193.0305	31322.926	0.34	0.54	0.403
-0.10	0.794	0.897	64.37	13017.18	3572.52	3205.1267	34528.051	0.38	0.63	0.303
-0.20	0.631	0.713	77.23	18737.93	5720.75	4076.8423	38604.891	0.42	0.76	0.203
-0.30	0.501	0.566	89.36	25086.26	6348.33	3593.6016	42198.492	0.46	0.88	0.103
-0.40	0.398	0.450	101.39	32295.35	7209.09	3241.5413	45440.031	0.50	1.00	0.003
-0.50	0.316	0.357	114.88	41460.89	9165.54	3273.6248	48713.652	0.53	1.13	-0.097
-0.60	0.251	0.284	129.65	52807.40	11346.51	3219.0918	51932.742	0.57	1.27	-0.197
-0.70	0.200	0.225	139.33	60987.25	8179.85	1843.3857	53776.125	0.59	1.37	-0.297
-0.80	0.158	0.179	151.62	72220.87	11233.62	2010.9021	55787.023	0.61	1.49	-0.397
-0.90	0.126	0.142	166.98	87594.87	15374.00	2186.0396	57973.062	0.64	1.64	-0.497
-1.00	0.100	0.113	185.06	107590.75	19995.87	2258.4583	60231.520	0.66	1.82	-0.597
-1.10	0.079	0.090	206.69	134211.19	26620.44	2388.2922	62619.809	0.69	2.03	-0.697
-1.20	0.063	0.071	237.49	177190.50	42979.31	3062.8967	65682.687	0.72	2.33	-0.797
-1.30	0.050	0.057	309.96	301828.94	124638.44	7055.4570	72738.125	0.80	3.04	-0.897
-1.40	0.040	0.045	350.30	385504.62	83675.69	3762.4695	76500.562	0.84	3.44	-0.997
-1.50	0.032	0.036	376.05	444263.44	58758.81	2098.6855	78594.187	0.86	3.69	-1.097
-1.60	0.025	0.028	394.67	489348.06	45084.62	1279.0972	79878.250	0.88	3.88	-1.197
-1.70	0.020	0.023	424.90	567182.75	77834.69	1754.0781	81632.312	0.90	4.17	-1.297
-1.80	0.016	0.018	460.50	666206.94	99024.19	1772.6282	83404.937	0.92	4.52	-1.397
-1.90	0.013	0.014	494.56	768400.25	102193.31	1453.1125	84858.000	0.93	4.86	-1.497
-2.00	0.010	0.011	514.75	832420.19	64019.94	723.0901	85581.062	0.94	5.06	-1.597
							91099.000	(1)		

between the current work and Evans' investigation emphasizes the importance of precise background determination, and also demonstrates the basic difficulties in measuring accurately on noisy tracings, distances of 0.2 mm. The determination of the background level has also been discussed qualitatively and quantitatively by de Vaucouleurs (1948). The two tables also show the dramatic difference in the values derived for r_e , viz., 61".83 and 101".4, indicating an increase of 64% after revision of the sky background level.

For the photo-electric calibration of the photographic measurements, remarks were passed in the previous chapter (section 4.6) on the convenience of such a system, since it removes some of the inherent difficulties present in the more indirect methods. However, it was found in the present work that some pairs of photo-electric results in the Reference Catalogue showed occasional large differences depending upon the observer. Let us consider the case of NGC 4649, whose photometric calibration is shown in table 3, where the magnitude data are taken from appendix I of the Reference Catalogue (G. and A. de Vaucouleurs 1964).

Table 3

Zero point of photographic photometry : NGC 4649

r	B(r)	L_r	Const.	Source of B(r)
32.2 arc sec	11.10	22614	21.99	A
122.2 " "	10.27	48313	21.98	B
122.2 " "	10.06	48313	21.77	C
286.5 " "	9.95	64318	21.97	G

Mean adopted constant in mag.sec.⁻² 21.98

The sources of the magnitudes, all transformed to the B system, are as follows:-

A - Stebbins & Whitford (1957); B - Stebbins & Whitford (1952); C - Pettit (1954); G - Holmberg (1958).

Clearly the value for the adopted constant should be averaged among A, B and G, since source C, using the same aperture as B, is obviously in error by a fifth of a magnitude. These errors in Pettit's work, detected by de Vaucouleurs (1963), have already been mentioned elsewhere in the thesis. For NGC 4649, the determination of the constant has been easy, but for many other galaxies, where there is a large scatter of constants about a mean, it is necessary to refer to the Introduction to the Reference Catalogue and also to de Vaucouleurs'

work (1963) in order to assess the accuracy of individual magnitude sources. Consider, for example, the following pair of values for NGC 4267. For an aperture radius of 65.6 arc sec, Bigay's (1951) magnitude transformed to the B system is 11.74 while at a radius of 72.1 arc sec, the B magnitude derived from Bigay's et al. work (1953, 1954, 1955) is 12.06. Before these measurements may be used in a zero point determination, there is already a magnitude difference of at least $0.^m32$, and although it is known that Bigay's magnitudes were poor at large apertures (de Vaucouleurs 1963), it is still difficult to decide which of these two measurements is the more likely to be accurate. In addition, if the luminosity profile of the galaxy is not correctly determined, the differences will be even greater. For more than half the galaxies studied, no more than two values appear in appendix II for the B-V colour and for 14 objects, there is only one value. Although for some galaxies it may be possible to construct colour curves within the galaxy if both the B and V calibrations are accurate, no attempt will be made in this thesis to construct colour curves of B-V within galaxies, since precise determination of the V

magnitude is impossible for many of the objects, due to inadequate calibration data. Further work to obtain good photo-electric measurements on galaxies is urgently required and this is one field where scope for research is still virtually unlimited.

Some of the calibration problems may well be removed with the introduction of the concept of "standard galaxies", in which well-determined profiles may be used to provide an internal photometric calibration, irrespective of whether the plate carries additional relative calibration areas. There are not yet sufficient photometric profiles of galaxies determined by direct photo-electric scanning methods to justify the establishment of such galaxies as standard objects, although unpublished data on NGC 3115, 3379 and 4486 are now available from de Vaucouleurs. However, using the integrated magnitude $B(0)$ within an aperture of diameter equal to the face-on value $D(0)$, and also the integrated luminosity law for normal ellipticals described in the Introduction to the Reference Catalogue, it is possible to derive the brightness distribution $I(r)$ in magnitudes per square second. This method was outlined by G.de Vaucouleurs at the 1970 IAU meeting of the Working

Group on Galaxy Photometry and full details have now been circulated to members of the Group (de Vaucouleurs 1970).

The integrated luminosity law may be expressed as

$$B(0) - B(X) = \beta X(2X_0 - X)$$

where for ellipticals, $\beta = 0.75$ and $X_0 = +1$. $B(X)$ is the integrated magnitude within a circular aperture of diameter $A = 2r$ and X is equal to $\log A/D(0)$. The surface brightness $\mu(X)$ at X is given by

$$\begin{aligned} \mu(X) &= -2.5 \log P.Q \quad \text{where} \\ P &= (0.8) 10^{-0.4B(0)} / \pi D(0)^2 \quad \text{and} \\ Q &= 2 \beta (X_0 - X) / 10^{2X - 0.4 \beta' X(X_0 - X)}. \end{aligned}$$

The quantity P , derived from values of $B(0)$ and $D(0)$ listed in columns 16 and 12 respectively of the Reference Catalogue, fixes the zero point of the stellar magnitude scale and Q gives the shape of the luminosity distribution law.

This method was used to check measurements for NGC 4649, the galaxy whose photometric zero-point calibration was shown in table 3. Results of the calculations are shown in the following table.

Table 4

Use of integrated magnitude to calculate surface brightness
in NGC 4649

X	r	P	Q	$\mu(X)_{\text{calc.}}$	μ_{obs}	O-C
-1.00	9.5 [✓] arc sec	1.93x10 ⁻⁶ [✓]	75.36	18.48 ³	18.80 ³³	+0.32
0.11	122.2 [✓] "	"	0.860	23.35	23.49 ⁴⁶	+0.14
0.30	189.2 [✓] "	"	0.303	24.47	24.73 ⁷⁰	+0.26
0.60	378 [✓] "	"	0.044	26.57	26.75 ⁹⁶	+0.18

It is noted that even with the good calibration constants obtained in table 3 for the zero points of the photographic magnitudes, the surface brightness within the galaxy are about 0.^m22 faint compared with the quantities derived in the above calculation. No error was present in the profile since a graph of values of log I against $\alpha^{\frac{1}{4}}$ agreed closely with the theoretical straight line equation for ellipticals. It is possible that the differences in O-C for NGC 4649 may be due in part to the overlap of this galaxy with NGC 4647. In cases where galaxies overlap, the isophotes of both galaxies are obtained on the same tracing and the corresponding areas within contours are measured on the assumption that the isophotes would have been symmetrical had the objects been isolated. It is thus a problem of deciding at what points the

isophotes depart from symmetry, although in this work, the presence of a detailed isophote map is a distinct advantage (cf. work of Holmberg 1958). The isophote calibration is established in a similar way, by using only one real half of a cross-sectional scan and assuming symmetry.

The calculation was repeated for NGC 4486, an isolated galaxy compared with the previous example, where $B(0) = 10.3$ and $D(0) = 3.715$, and the results are shown in table 5.

Table 5

Calculation of distribution of surface brightness in

NGC 4486

X	r	P	Q	$\mu(X)_{\text{calc.}}$	μ_{obs}	O-C
-0.87	15 arc sec	1.4×10^{-6}	50.0	19.27	19.53	+0.26
-0.18	75 " "	"	3.30	22.23	22.46	+0.23
+0.13	150 " "	"	.78	23.80	23.97	+0.17
+0.35	250 " "	"	.24	25.13	25.08	-0.05
+0.55	400 " "	"	.06	26.51	26.19	-0.32
+0.65	500 " "	"	.03	27.29	26.86	-0.43

The systematic difference in the values of O-C with increasing radial distance from the centre of this galaxy can easily be explained in the following way. At small

values of r , the observed distribution of luminosity is faint due partly to the smoothing of the true profile by both the telescope and microphotometer instrumental functions. It is to be expected that the central depletion of intensity is replaced in the outer regions of the system, but not to the extent of a half magnitude nett difference from centre to edge. The major contribution to the observed increase in luminosity at large radial distances must be due to the detection and measurement of part of the outer corona which is now known to surround NGC 4486 (de Vaucouleurs 1969). For example, at 500 arc sec from the centre, the theoretical calculation for μ gives a value of 27.79, the observed value is 26.86 and where the photometry of the corona is complete, its contribution to the light of the galaxy raises the observed distribution at this distance, measured off de Vaucouleurs graph (1969) to a surface brightness of $25.3 \text{ mag sec}^{-2}$. In the Working Group Circular (de Vaucouleurs 1970), mention was not made of the range of values of X satisfying the integrated luminosity law; it is only applicable for $-1 < X < +1$. However, although this method will prove extremely

useful in the future where elliptical galaxies are present on uncalibrated wide field plate collections, it must be stressed that recent developments in the detection of coronas around giant elliptical systems, such as M87, mean that results derived in this way should be checked carefully in case the unknown luminosity distribution contains light measurements from such outer extensions.

yes

5.3 Results. Complete photometric data for the galaxies are listed in standard format, following the presentation adopted by de Vaucouleurs (1961b), and X from the contents of volumes 2 and 3 of the thesis. In a large scale investigation of this nature, where many types of galaxies have been studied, it should be possible to derive relationships between some of the photometric parameters and the revised morphological type (de Vaucouleurs 1963).

First of all, let us compare and examine the data for the radial distribution of surface brightness $I(r)$ in NGC 4535 (type SAB(s)c) with similar data by Sersic (1968) for M83, and also data for NGC 4459 (type SA(r)0) with measurements made by Liller (1960). These distributions, shown in figure 27, are typical

for the whole family of disk-like galaxies, including spirals and lenticulars (Sandage, Freeman and Stokes 1970). The measurements from the current work are indicated in the figure by crosses. de Vaucouleurs (1959a) has shown that the luminosity distribution for these objects consists of two main components, an inner spheroidal component which follows closely the law for ellipticals

$$\log I \propto r^{\frac{1}{4}}$$

and an outer exponential disk component

$$I(r) = I_0 e^{-\alpha r}$$

containing a large fraction of the total luminosity. I_0 is the central intensity of the disk component extrapolated to $r = 0$. This two-component nature of the disk systems distinguishes them from ellipticals, which have a single component given by the quarter power law (de Vaucouleurs 1959a). At large radial distances, the quarter power law distribution appears to be approximately exponential (Freeman 1970), so that the $I(r)$ law is qualitatively similar to that shown in figure 27 for objects like NGC 4459, but the main difference between the distribution for elliptical

and disk galaxies is that the apparently exponential and non-exponential sections of the profiles for ellipticals are not independent.

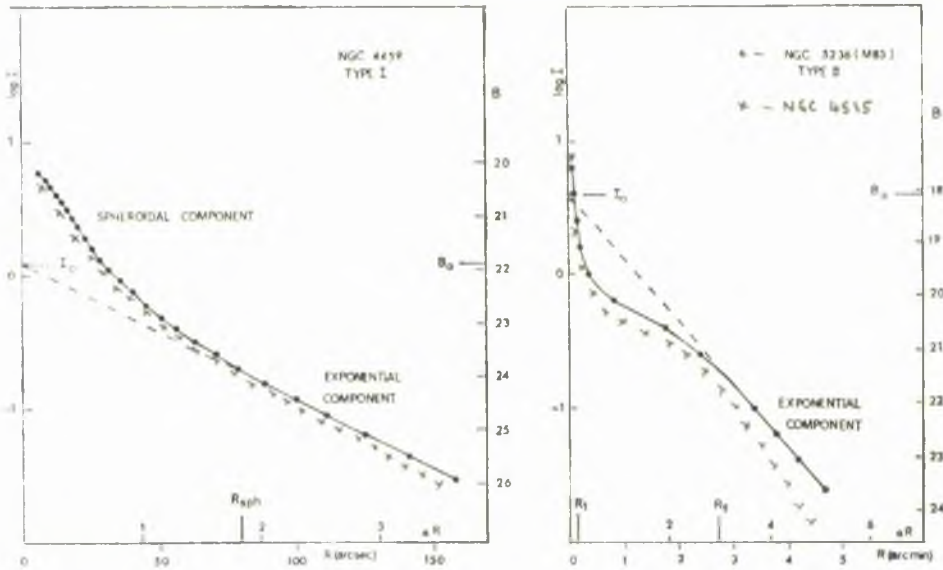


FIG. 27.—Radial luminosity distributions for NGC 4459 and M83. I is the surface brightness. Ordinates are $\log I$ and I in B -mag per square second of arc. R is distance from the nucleus along the major axis; the dimensionless radius aR is also shown. I_0 , B_0 are the surface-brightness scale for the exponential disk, uncorrected for inclination and galactic absorption. R_1 , R_2 , and R_{ph} are defined in § III. Filled circles, observed points; crosses, St. Andrews work.

A study was undertaken of the variation of the central surface brightness $B(0)_c$ with morphological type, since Freeman found, for 28 galaxies out of 36, that $B(0)_c$ was nearly constant at 21.65^m along the entire spiral sequence and showed no apparent dependence on morphological type. After extrapolating

the exponential disk component to $r = 0$ in order to derive the apparent intensity scale given by

$$B(0) = -2.5 \log I_0 + \text{const.},$$

the intrinsic intensity scale is obtained by correcting $B(0)$ for the inclination to the line of sight and also for galactic absorption. The corrected central brightness is derived from the relation

$$B(0)_c = B(0) + 2.5 \log R - 0.2 \operatorname{cosec} b^{\text{II}}$$

where R is the ratio of major to minor diameter of the disk, as derived from the isophotes and cross-sections of the galaxy; the coefficient of $\operatorname{cosec} b^{\text{II}}$ is that obtained by de Vaucouleurs and Malik (1969) while the values of b^{II} are obtained from the B.G.C. (G. & A. de Vaucouleurs 1964). Only those galaxies for which the exponential components were easily defined have been included in this study, and the results are given in table 6.

Figure 28 shows $B(0)_c$ against morphological type. Statistical analysis of the data is made possible by

TABLE 6
^A
 DATE FOR 19 GALAXIES

NGC	Classification stage	B(0)	log R	B(0) _c
4235	a	21.20	.49	22.32
4293	O/a	21.76	.22	22.22
4321	bc	21.08	.15	21.36
4382	O ⁺	21.96	.08	22.08
4394	bc	21.76	.05	21.80
4450	ab	21.50	.16	21.82
4477	O ⁰	22.02	.01	21.96
4501	b	22.50	.07	22.56
4535	c	21.46	.02	21.43
4536	bc	21.93	.14	22.34
4548	b	21.50	.00	21.51
4567	bc	22.30	.08	22.42
4569	ab	21.18	.31	21.84
4571	cd	21.62	.03	21.63
4578	O	22.40	.14	22.64
4579	b	20.86	.08	20.95
4698	ab	22.40	.10	22.56
4754	O ⁻	22.21	.11	22.37
4762	O ⁰	22.32	.44	22.40

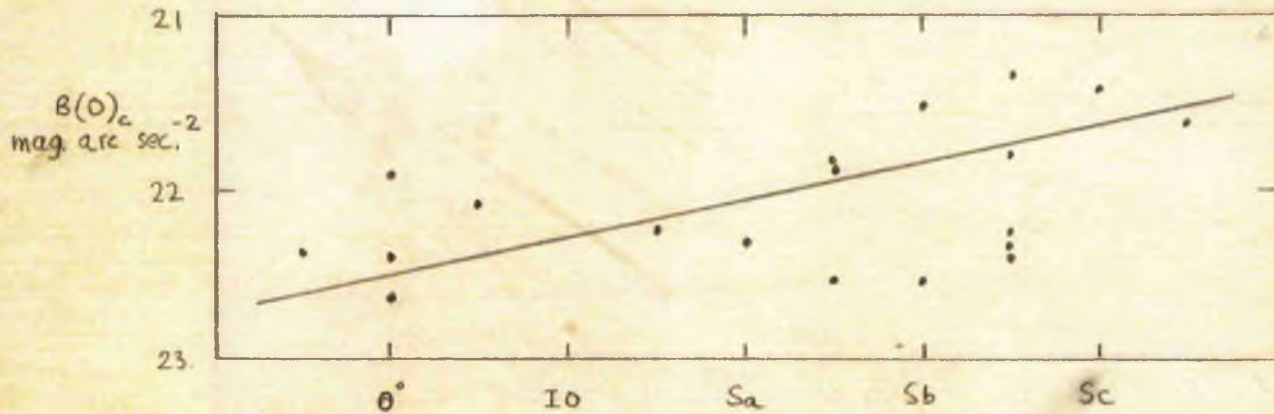


Figure 28 - $B(0)_c$ against morphological type.

assigning to each classification stage the following numerical scale

\sim	L^-	L^0	L^+	IO	SO/a	Sa	Sab	Sb	Sbc	Sc	Scd
n	1	2	3	4	5	6	7	8	9	10	11

This scale is similar to one quoted by de Vaucouleurs (1970b). Note that in Freeman's paper (1970), the scale given throughout for morphological type departs considerably from the above scheme, being much more condensed for all the lenticular stages. Least squares fitting of the data for figure 28 produces the

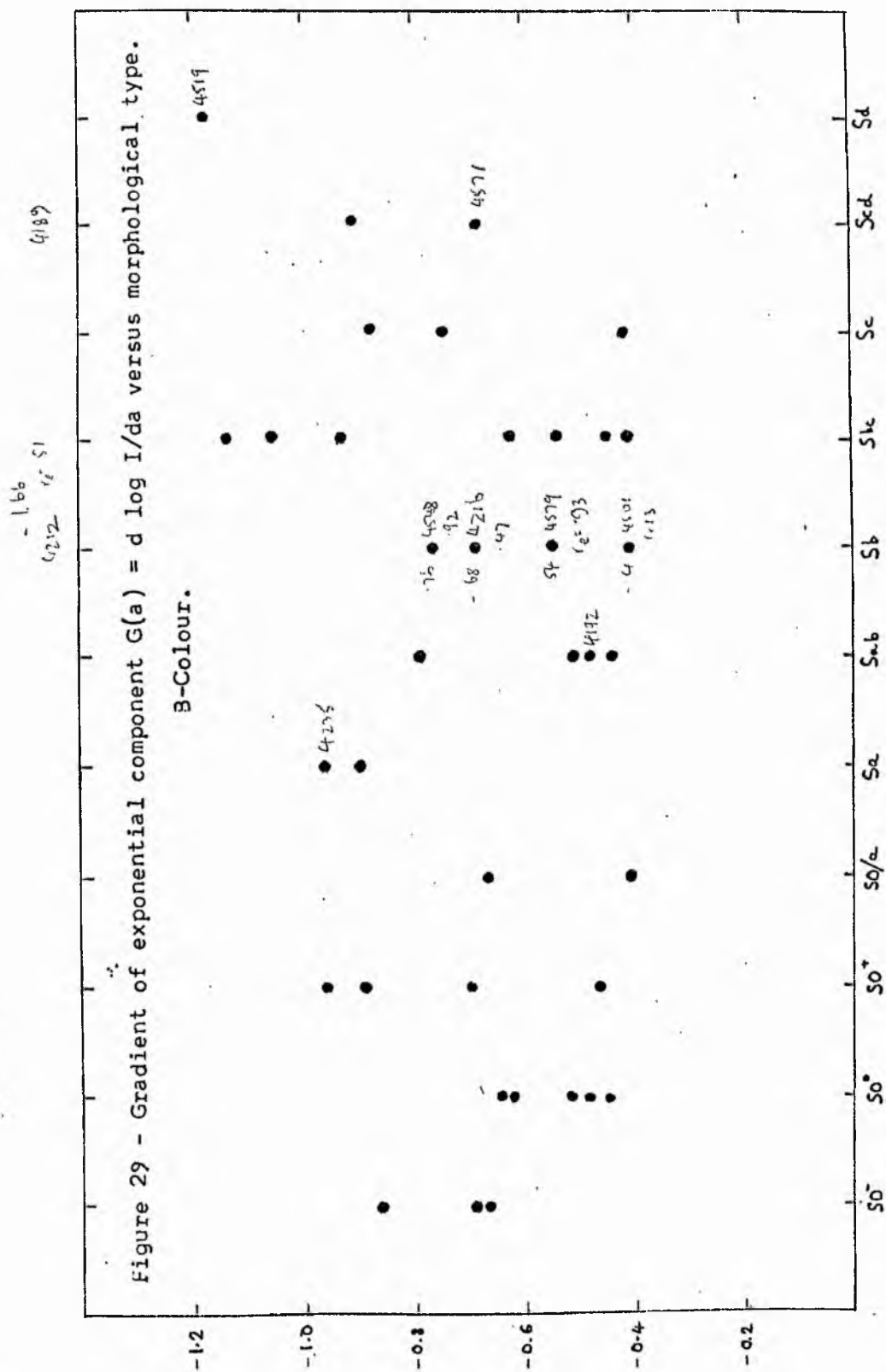
equation

$$B(0)_c = 22.76 - 0.11 n,$$

where n denotes the Hubble type. This expression indicates that there is a systematic variation of $B(0)_c$ with morphological type, contrary to Freeman's (1970) results. The difference is greatest around the lenticular stage, and it may be due in part to possible errors in Liller's photometry (1960, 1966), which provided the source of Freeman's data on $B(0)$ for lenticulars. A comparison of Liller's work with the writer's investigation will be given at a later stage.

The gradient $G(a) = d \log I/da$ has been plotted against revised morphological type in figure 29, and it is seen that the data exhibit the usual scatter since $G(a)$ is proportional to the angular size of each object. A more meaningful relation may be obtained by normalizing the length scale to r_e and setting $y = \frac{r}{r_e}$. In figure 30, the gradient $G(y) = d \log I/dy$ is seen to vary with morphological type according to the relation

$$G(y) = -.22 + .05n .$$



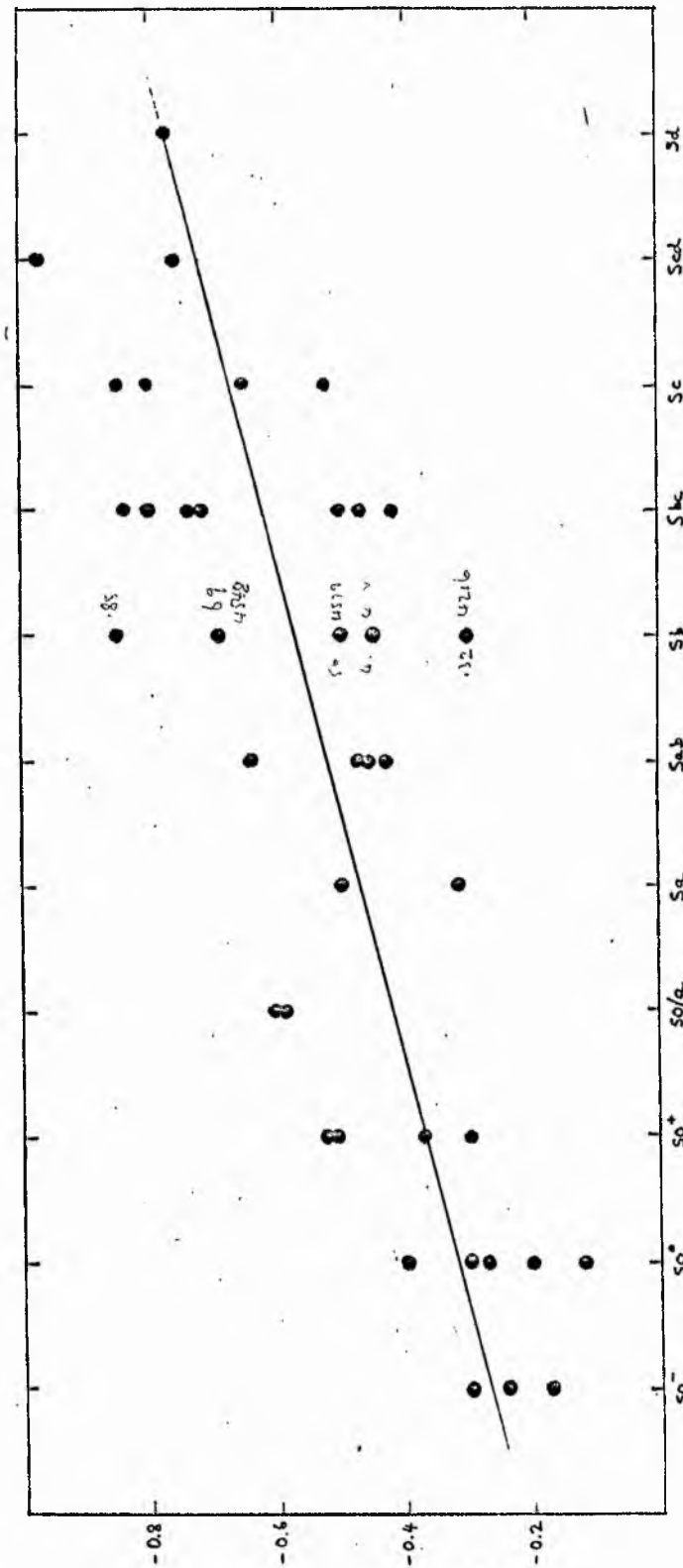


Figure 30 - Gradient of exponential component $G(y) = d \log I/dy$ versus morphological type
 where $y = r^*/r_e$. B-Colour.

This result will be verified after the concentration indices are plotted against morphological types, but for the present, the above relation agrees with the work of Freeman (1970) who found a mixture of values for $G(a)$ in early types but predominantly high values for later types. However, in this investigation, two stage cd galaxies and one stage d object were the latest types to be observed, so it is not possible to verify completely Freeman's results for the late-type objects, although extrapolation of the least squares fit in figure 30 would indicate agreement in principle. Originally, plots were also made on this figure to test whether there was any correlation of gradient with families within each stage, but no correlation was noted.

A comparison of results in the current work with that of Mrs Liller (1960, 1966) can be made by using her values for either m_T , the total integrated magnitude including an extrapolation correction, or a_{25} , the ~~smei~~ major axis of the isophote at a surface brightness of 25.0 mag.per square second of arc. Table 7 gives data for all the galaxies common to both investigations, while the filled circles in figure 31 illustrate graphically the respective values. The slope of the

Table 7

Comparison of different measurements of semi-major axis
in arc min. at $\mu = 25 \text{ mag. sec}^{-2}$.

NGC	a'_{25} (Liller)	a/a	$a(25)$ (Fraser)	NGC	a'_{25}	a/a	$a(25)$
4365	3.33	1.03	3.22 3.06	4486	4.28	1.10	3.90 5.36
4374	3.60	1.23	2.92 2.78	4552	2.81	1.13	2.48 2.62
4406	6.15	1.52	4.05 4.20	4649	4.63	1.13	4.08 3.81
4459	2.10	0.95	2.20 2.06	4526	3.78	1.16	3.25 3.55
4472	6.33	1.26	5.03 5.05	4578	1.66	0.99	1.67 1.78

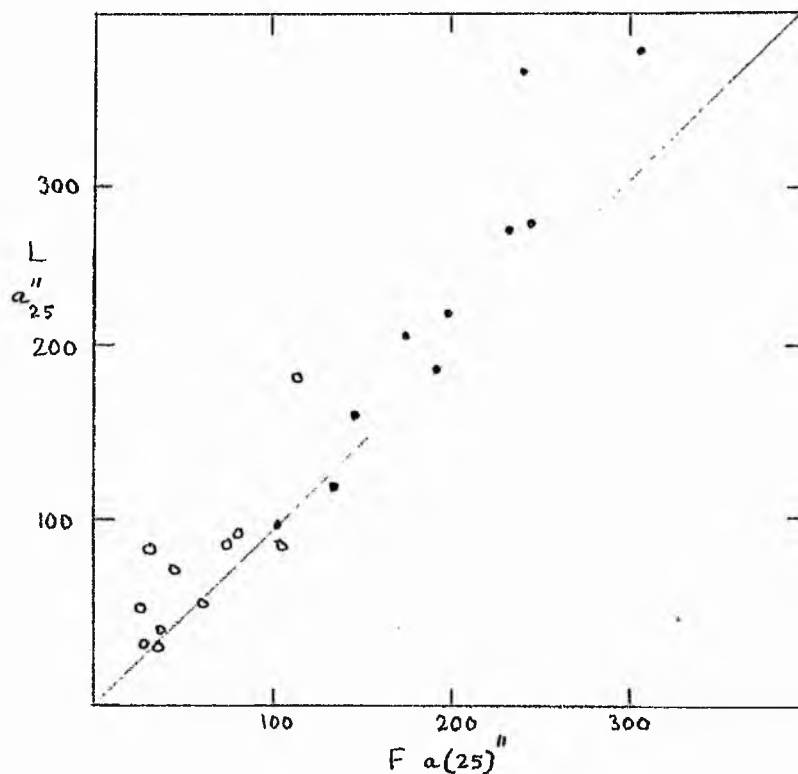


Figure 31 - Graphed Output of Table 7

X
give proper
caption!

no, there is a npt. cur.
plot a/a vs a/a

graph departs only slightly from 45° but in general the data show good agreement. The large discrepancy in the two values for NGC 4406 may be due to the fact that Liller finds its m_T to be exceptionally bright at 9.61, compared with Holmberg's and the writer's respective values of 10.1 and 10.04. The open circles in figure 31 show values for Liller's a_e and the writer's equivalent a_e ; this quantity was derived from values of r_e and the measured axial ratio $q = b/a$ of the isophotes at r_e using the relation

$$a_e = r_e \sqrt{1/q_e}$$

Out of the 11 galaxies for which comparisons could be made, NGC 4459, 4461, 4486 and 4552 show particularly good agreement. In view of the fact that such good agreement was obtained for these four objects, for which the results were published by Mrs Liller in 1966, and since the measuring procedure in the current work has been uniform throughout, it is suggested that there may be errors in some of Mrs Liller's results, supported by the following three reasons.

First of all, it is possible that errors could have been introduced in the extrapolation procedure. Since the isophote levels reached down to a few per cent

of the sky background, it was necessary to extrapolate to an infinite radius for the undetected contributions of the galaxy, and in some cases such corrections could be significant. Furthermore, Liller (1966) pointed out that an error of only $0.^m1$ in the extrapolation yielding m_T would lead to an error of more than 10% in a_e . In addition, if m_T is erroneous, as may well have been the case for NGC 4406, the resultant error in a_e may be considerable.

Secondly, in the 1960 investigation, Mrs Liller smoothed all the isophotes and assumed them to be ellipses whereas in 1966, corrections were applied to any isophote which departed from an elliptical form.

Thirdly, Mrs Liller (1966) has pointed out that some of the 1960 calibrations were known to have zero point inaccuracies of as much as $0.^m5$ after check measurements were made on NGC 4486 using unpublished photo-electric data of Baum.

For these reasons, it would appear that the 1960 results would be more likely to have errors compared with the 1966 investigation. From the writer's work on the central brightness $B(0)_c$ of the disk component (see figure 28), it was found that the value for lenticulars would lie around $22.^m4$. Using Liller's 1960 data,

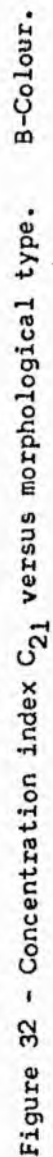
Freeman finds for 8 lenticular galaxies an average $B(0)_c$ of 21.51 while 4 objects in the 1966 investigation produced an average for $B(0)_c$ of 21.89. It is proposed that if the writer's values for lenticulars err on the faint side, then Freeman's results for the same stages will err on the bright side, thus contributing to the apparent constant value for $B(0)_c$ along the entire spiral sequence. If Liller's 1960 work is inaccurate and if the average $B(0)_c$ derived by Freeman from this data should in fact have been closer to 21.9, this would have the effect of producing a variation in $B(0)_c$ along the morphological sequence, similar in slope to the relation derived by the writer in figure 28.

The apparent ellipticity of the isophotes increases from the nucleus outwards beyond which a very slow decrease takes place in the outer regions; this is confirmation of results given by de Vaucouleurs (1959a, 1970a). Part of the observed decrease towards the centre is due to the instrumental smoothing which tends to produce circular isophotes irrespective of their true shapes. However, recent work by Arp and Bertola (1969) has shown that for M87, the ellipticity of outer isophotes is 0.5 compared with zero for the central isophotes.

Confirmation of such results should be obtained in future work using not only the IIIaJ emulsions now available (van den Bergh 1969), (Arp and Bertola 1969), but also with photo-electric techniques, such as those employed by de Vaucouleurs (1969). It is important to point out that Arp and Bertola, although they obtained plates with IIIaJ emulsion and published detailed isodensitracer photographs (1971) which [they claimed could] enable faint outer isophotes to be detected, this method has been employed by the writer in St Andrews for the past six years and its advantages in drawing isophotes were described in a paper published in 1967 (Fraser 1967).

One point concerning the ellipticities of isophotes is that there appears to be some ambiguity in the published literature about the definition of axial ratio and the resulting interpretation of ellipticity. This confusion originally lead to some difficulties in linking the writer's observed ellipticities with the published data.

The variation of concentration indices with morphological type has been analysed in both colours and the results shown in figures 32 to 35. It is seen that in all cases there is a systematic dependence of index along the Hubble sequence, the smaller indices



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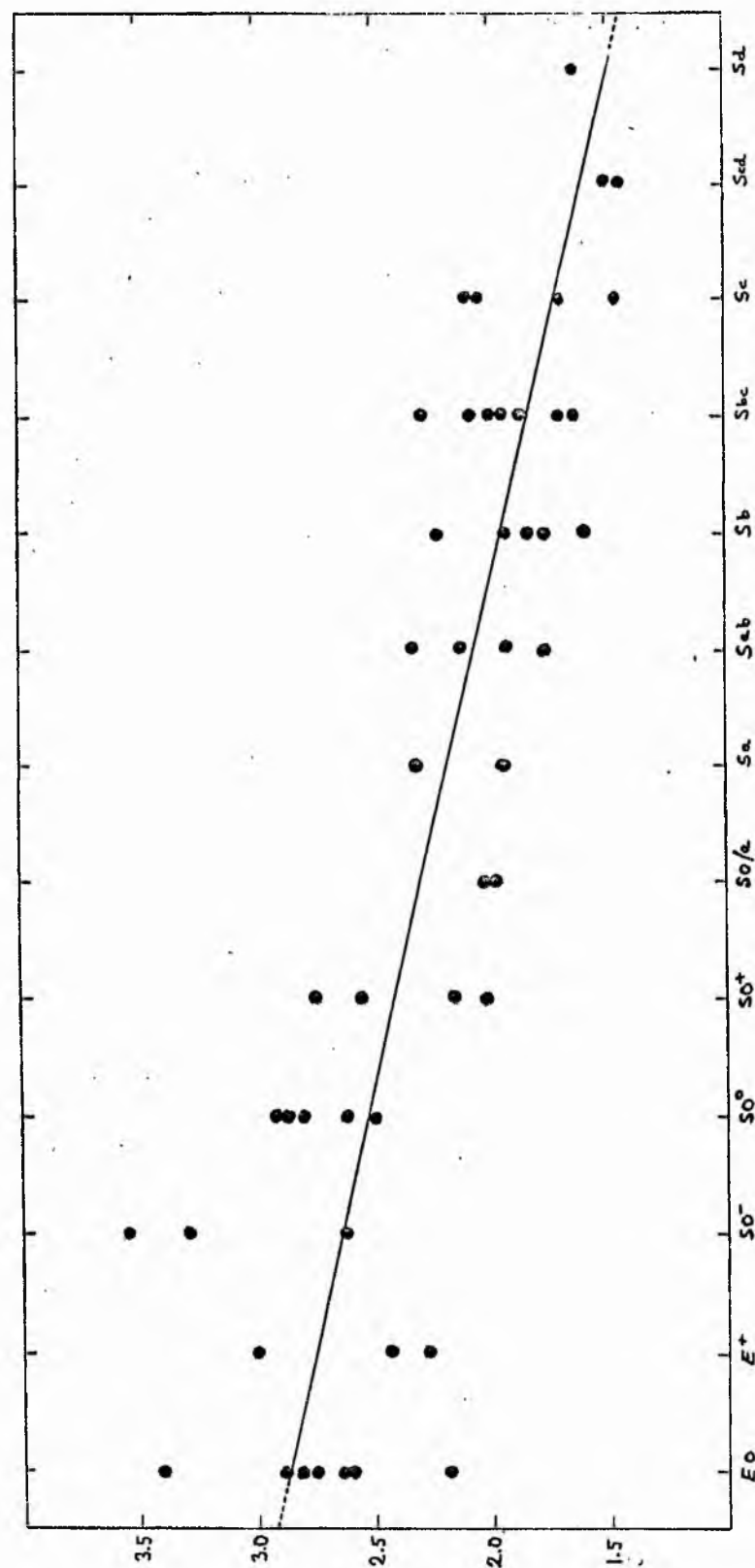


Figure 33 - Concentration index C_{32} versus morphological type. B-Colour.

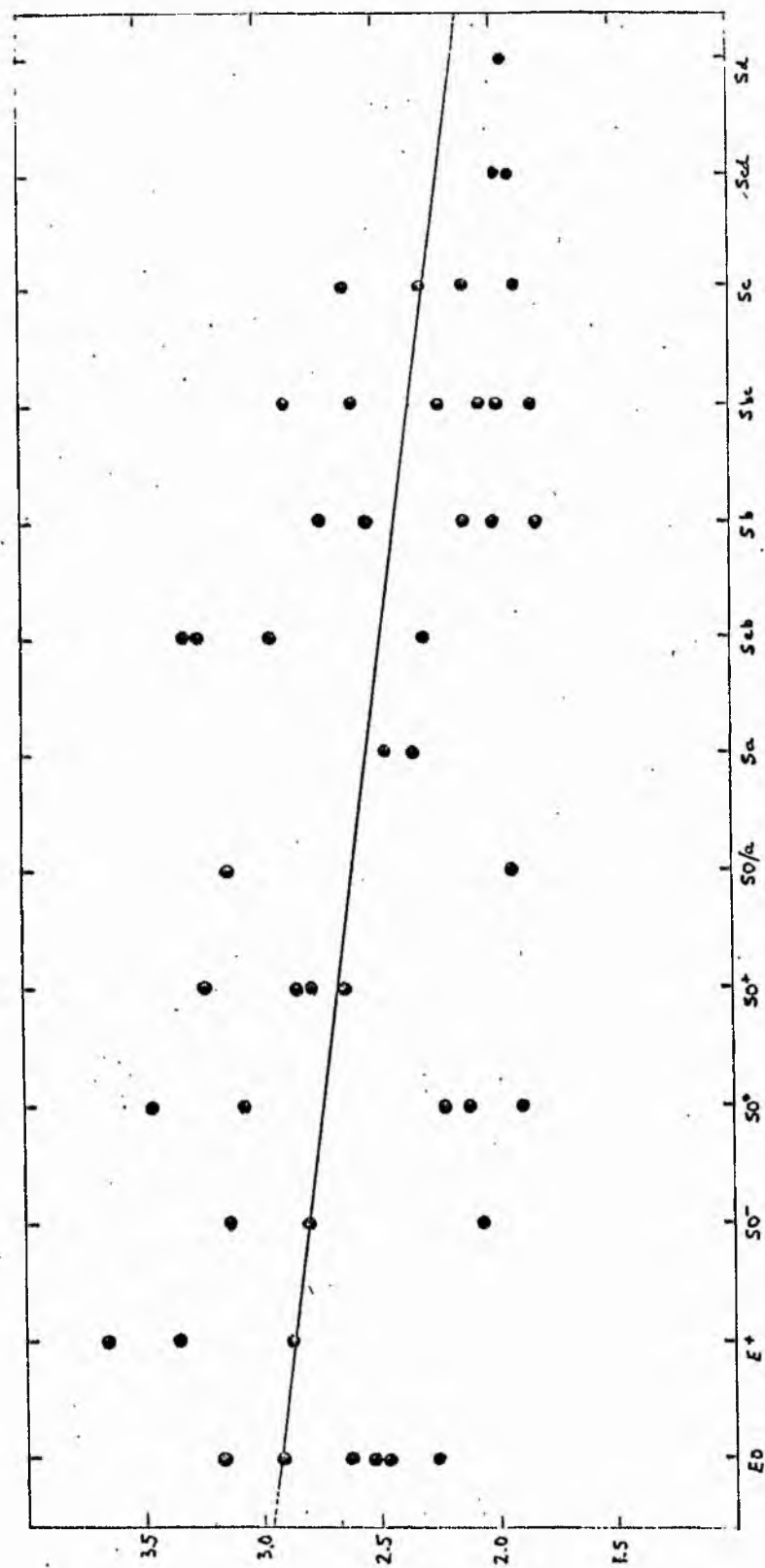


Figure 34 - Concentration index C_{21} versus morphological type. V-Colour.

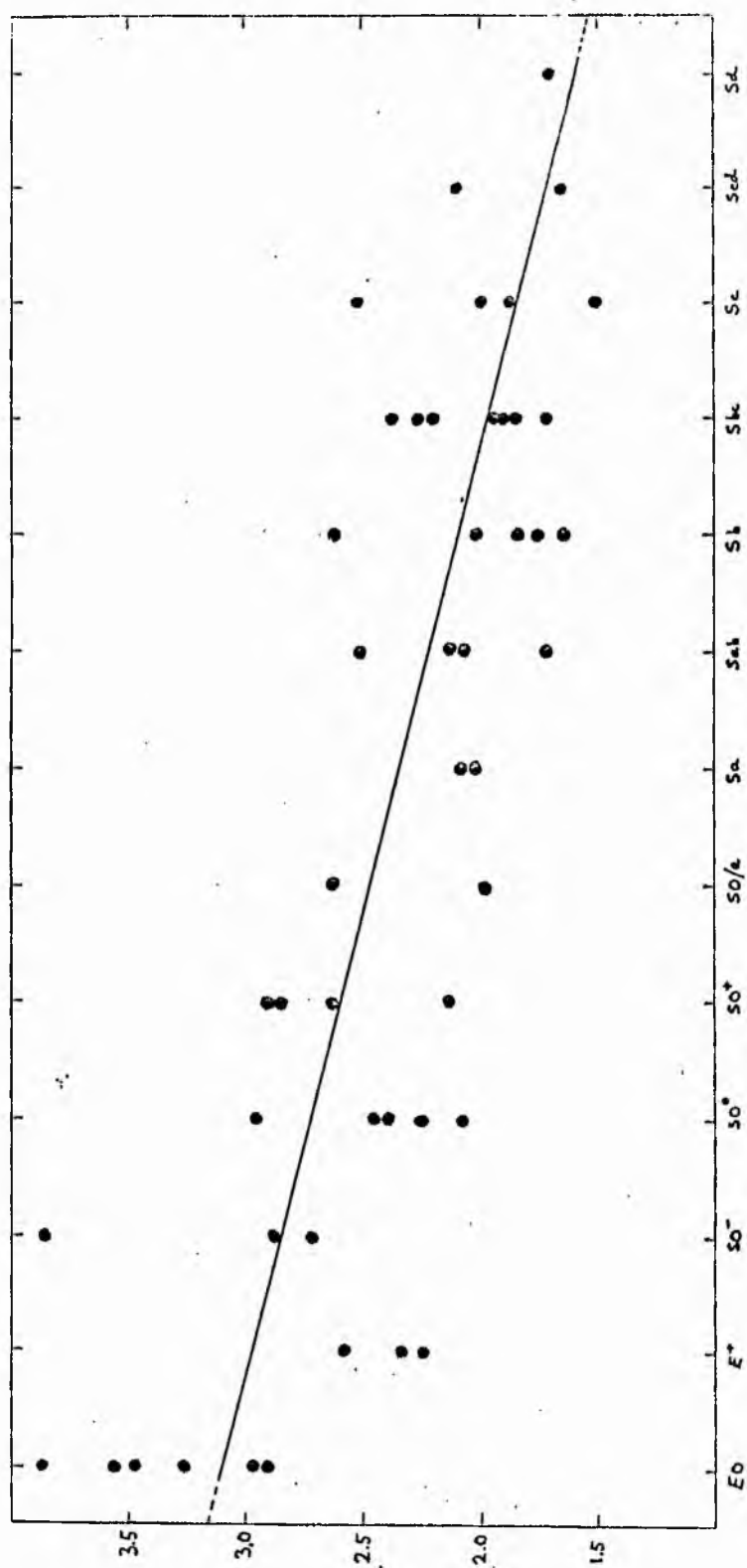


Figure 35 - Concentration index C_{32} versus morphological type. V-Colour.

being found with the late types and the SO and E galaxies having all but the lowest values of concentration index. This is to be expected since the late types have compact nuclei and their boundaries are well defined by the spiral arms, thus making the ratios defined by the concentration indices all fairly small, while for ellipticals, which merge imperceptibly into the background at large radial distances, r_3 , defined by $k(r_3) = \frac{3}{4}$, is much farther out in the ellipticals compared with r_e than it is in late spirals. Since the morphological sequence includes ellipticals on this occasion, the numerical values assigned to each stage are as follows:

compute
theoret.
values

∞	E	E ⁺	L ⁻	L	L ⁺	SO/a	Sa	Sab	Sb	Sbc	Sc	Scd	Sd
n	1	2	3	4	5	6	7	8	9	10	11	12	13

use
our #
system

A least squares fit to each set of data on figures 32 and 33 produced the following relationships for the B system:

$$\begin{array}{ll} C_{21} & \text{C.I.} = 3.6 - 0.140 n \\ C_{32} & \text{C.I.} = 3.2 - 0.124 n \end{array}$$

where n denotes, as before, the Hubble type. Analyses for the data shown in figures 34 and 35 produced for the V system the following equations:

$$C_{21}: \quad C.I. = 2.96 - 0.061 n$$

$$C_{32}: \quad C.I. = 3.20 - 0.124 n.$$

The results (shown in figures 32 to 35) for the concentration indices may be compared with the trend shown in figure 30 for the gradients. For the lower gradients found in the outer envelopes of lenticulars, the concentration index should be numerically large compared with later type objects; similarly, low concentration indices should be produced by galaxies having high gradients. These inverse relationships are confirmed in the results shown in figures 30, 32, 33, 34 and 35. This classification work on luminosity is similar to Freeman's work (1970) on the angular sizes of different sections of a galaxy. Freeman's parameter gives the ratio of the apparent radius R_{sph} (defined in figure 27) of the spheroidal component to the radius of the disk at some consistent surface brightness. However, this ratio is shown to be only very weakly correlated with galaxy type, increasing from late to early types. The present work on the concentration indices, shown in figures 32 to 35, gives more dramatic relationships.

In case inclination effects were important in.

determining the location of points in the four figures under discussion, an analysis was made of the relationship between concentration index and galactic inclination. For a galaxy, the inclination i of the axis of rotation to the line of sight is given by

$$\cos^2 i = (q^2 - q_0^2) / (1 - q_0^2) \quad \text{where } q = b/a.$$

The value of q_0 has, in the past, been taken as 0.2 (Holmberg 1958), but it has recently been shown (de Vaucouleurs 1970a) that the value of q_0 varies systematically along the Hubble sequence. However, it suffices to remark that q may be used as a measure of the inclination of a galaxy. Figure 36 shows the concentration indices versus varieties, families and apparent ellipticities within stages for each of the points plotted in figure 32. No systematic effects are noted, except perhaps a tendency for more of the larger ellipticities to be associated with the later types, indicating a decrease in concentration index with increase in angle of inclination. Note the cluster of SAB(rs) types at stage bc.

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(15.9)

Just as the luminosity profile $I(r)$ may be divided up into distinct components, it is also

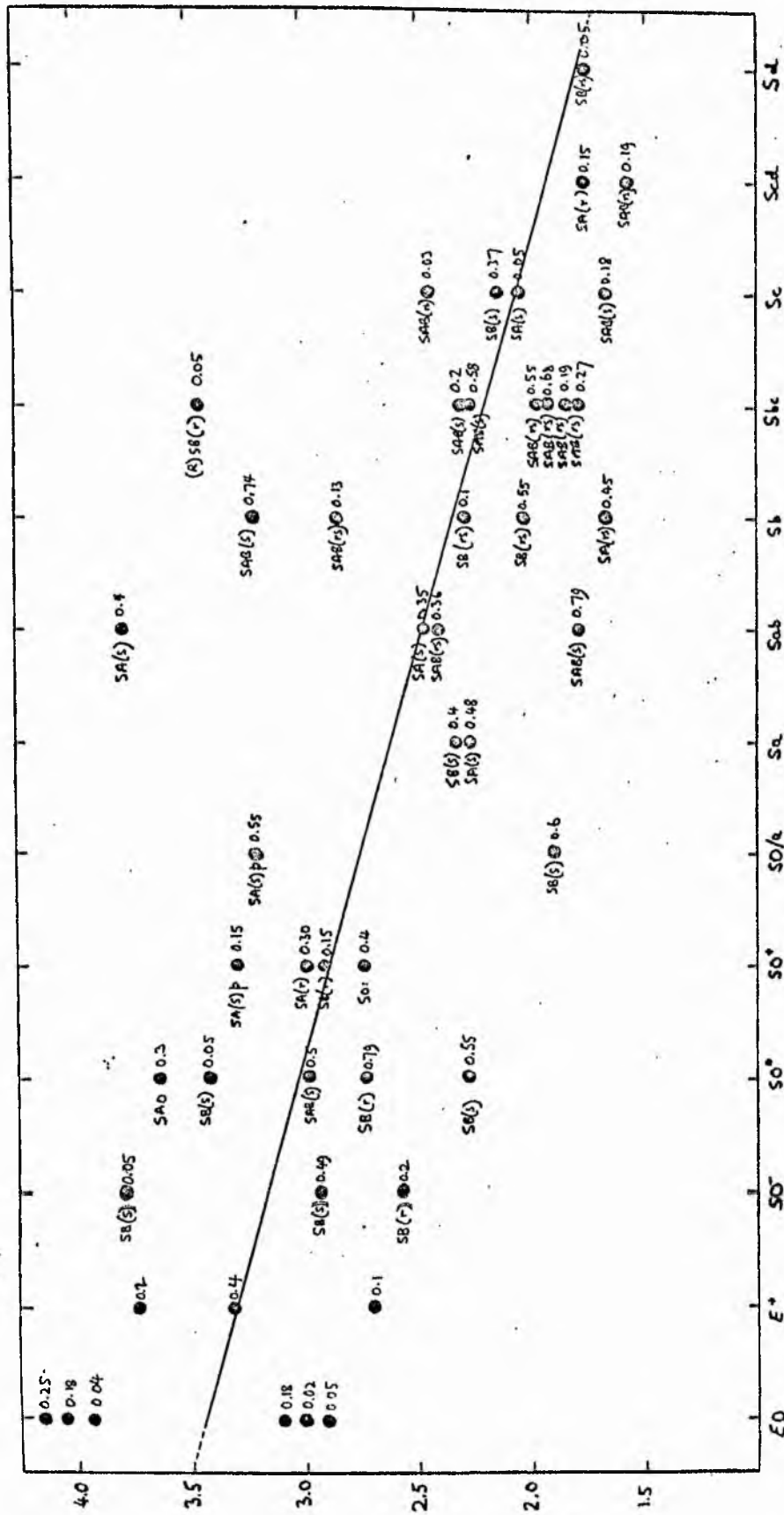


Figure 36 - Concentration index C_{21} (B-Colour), varieties, families and ellipticities versus stages of revised morphological sequence.

possible to subdivide the reduced luminosity profile. Such a profile for the B system photometry of NGC 4477 is shown in figure 37. It is seen that the graph may be divided into three components, a central, an intermediate and an outer. de Vaucouleurs (1959a) has already given equations for the reduced laws for ellipticals and the exponential law observed for late-type spirals, and pointed out that the ellipsoidal law seems to apply to the spheroidal component of early spirals. Instead of attempting to fit a truly exponential component to the central regions where the scale in ρ is small, the reduced law was plotted in terms of $\rho^{\frac{1}{4}}$, where a straight line fit was obtained for the first 16 data points, neglecting the first two which suffer from resolution difficulties and effects of smoothing functions. The equation for the line fit to the data was

$$\log J = 3.43 - 3.39 \rho^{\frac{1}{4}},$$

indicating good agreement with de Vaucouleurs' (1959a) results. For the second component, the same straight line fitting technique was applied to the data and the result, shown in figure 37, produced for this second component the following equation

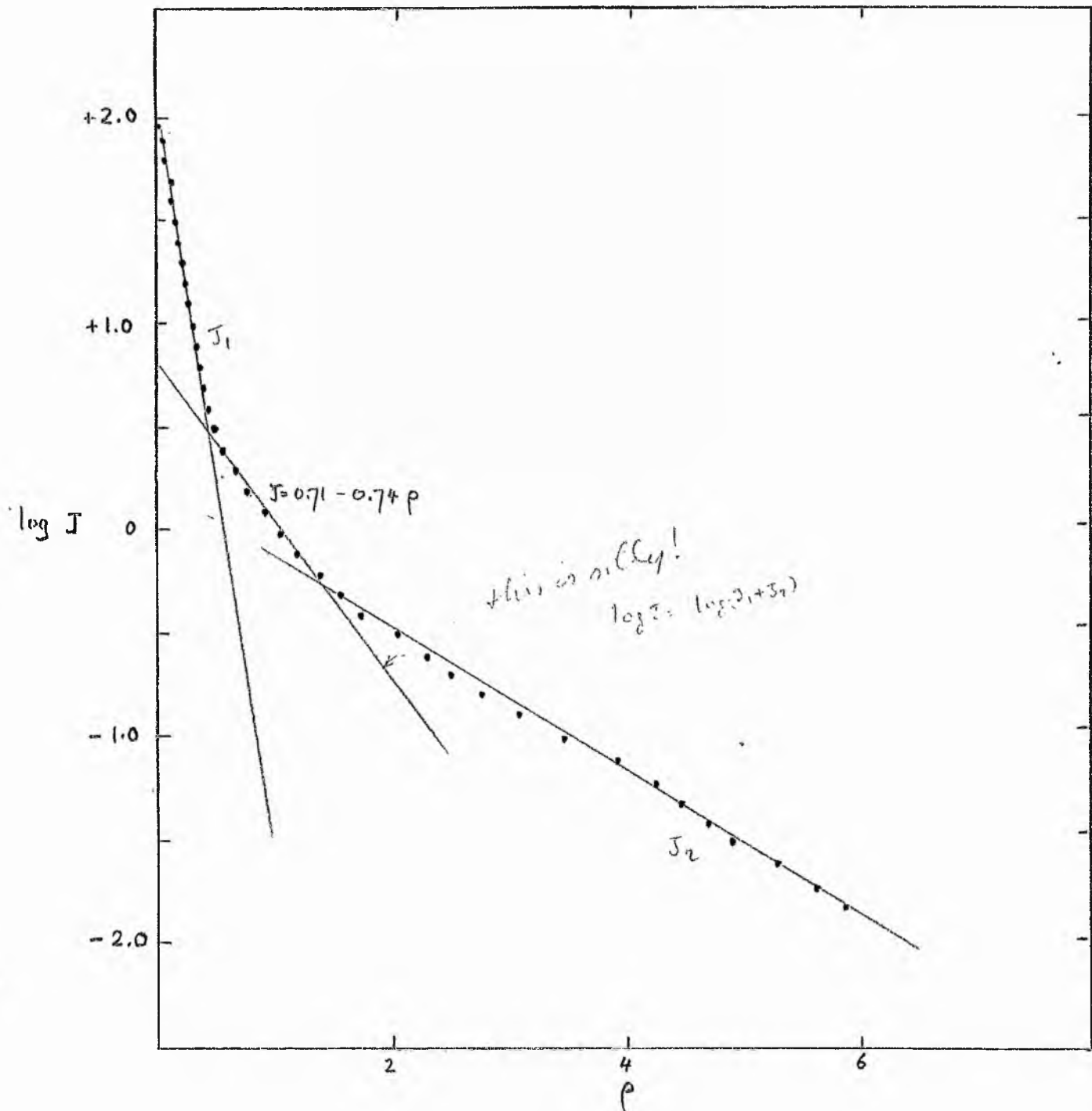


Figure 37 - Reduced luminosity profile of NGC 4477.

wrong
 direction

$$\log J = 0.71 - 0.74 \rho.$$

This result is consistent with de Vaucouleurs' relation (1959a) which differs from the above only in the sense that each coefficient is equal to .729. This type of analysis has been carried out for all the galaxies in the work; the results for the measurements of the gradient for each component of the reduced luminosity profile are shown in figure 38. Examination of the data reveals one important feature, also exhibited in figure 37. The profile in most cases seems to consist, not of two distinct components, but three. As a general rule, the first or central component is consistent with the quarter power law for ellipticals, the second concurs with the exponential law for late-type spirals and the third component satisfies an equation of the type

$$\log J = -b(\rho - 1) \text{ where } b \neq 0.3.$$

In cases where only two components are detected, the second of these obeys the exponential law for late-type systems. The detection of the third component in such galaxies as NGC 4216, 4293, 4382, 4461, 4477, 4501, 4519 and 4527 may be due to the large radial distances scanned on the plate. These distances are consistent with the measured principal axes at threshold being systematically greater than Holmberg's (1958) values

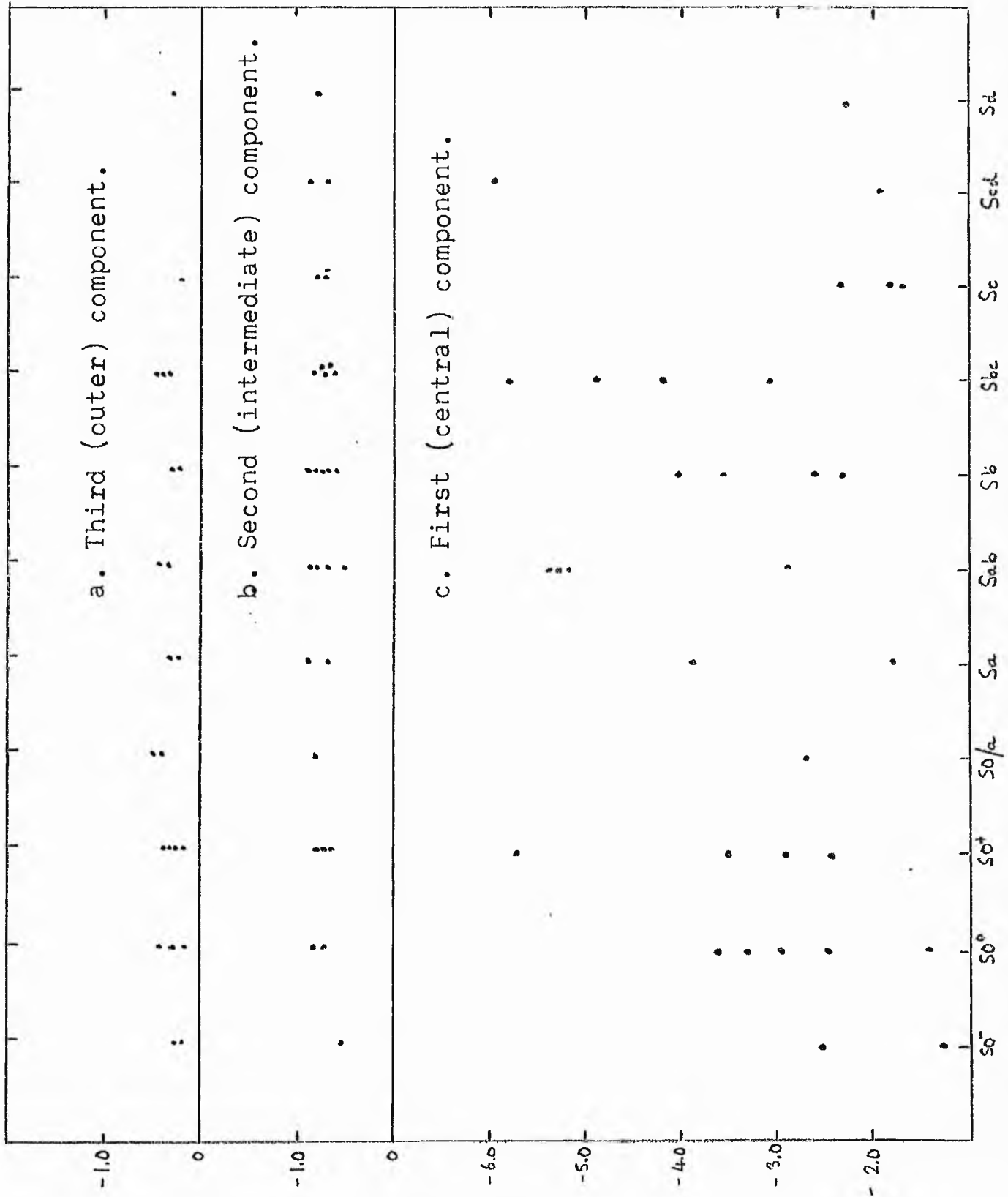


Figure 38.

Gradient of exponential components of reduced luminosity profile versus morphological type.

for NGC's 4216, 4382, 4501, 4519 and 4527; these are the only galaxies from the previous list common to both investigations. It may be possible that this third component could be more widely measured in future investigations. The question arose as to whether the presence of the third component in most, but not all, of the galaxies studied could be a selection effect, appearing only in the objects with small apparent magnitudes. However, a graph of the number of components versus the apparent magnitude showed that any correlation was extremely weak, if not altogether absent. The ratio of the number of galaxies with three components to those objects with two was about 2:1, all the component numbers being uniformly distributed in apparent magnitude. Comparison of the reduced luminosity profiles with other observers' results for galaxies in common is not possible, since previous investigations have not presented the photometric data in the "standard" form.

Heidmann (1967) has investigated the relationship between the intrinsic luminosity L of a galaxy and its diameter A up to a faint isophote and finds that

$$L \propto A^q$$

The exponent of A was found for spirals to be 2.8 using Holmberg's (1964) data and for ellipticals the corresponding q value was 1.9, obtained from Liller's (1960, 1966) data. The difference in the values of q was not due to different values used for the brightness of the limiting isophote (26.5 and 26.0 mag.sec.⁻² respectively), since the same q value was obtained from Liller's data using either the 25 or 24 mag.sec.⁻² limiting isophote. In a similar investigation, Heidmann (1969) found a value for q of 1.8 for Rood and Baum's (1968) photometric data for galaxies in the Coma cluster.

Such a study was also possible for the present work, since one of the tabulated standard parameters was the major axis $2a(25)$ at $\mu = 25.0$ mag.sec.⁻²; the results of this investigation are shown in figure 39. Analysis of the data for all galaxy types grouped together produced the line, drawn in the figure, which has a slope $q = 1.71$. Since there were 27 spirals included, one should have expected the slope of the line to have been closed to the value of 2.8 found by Heidmann from Holmberg's (1964) data on spirals. After analysing the data on 27 spirals and 7 ellipticals separately, the resulting q values were 1.62 and 1.89 respectively.

a must be corrected to "face-on", also m.

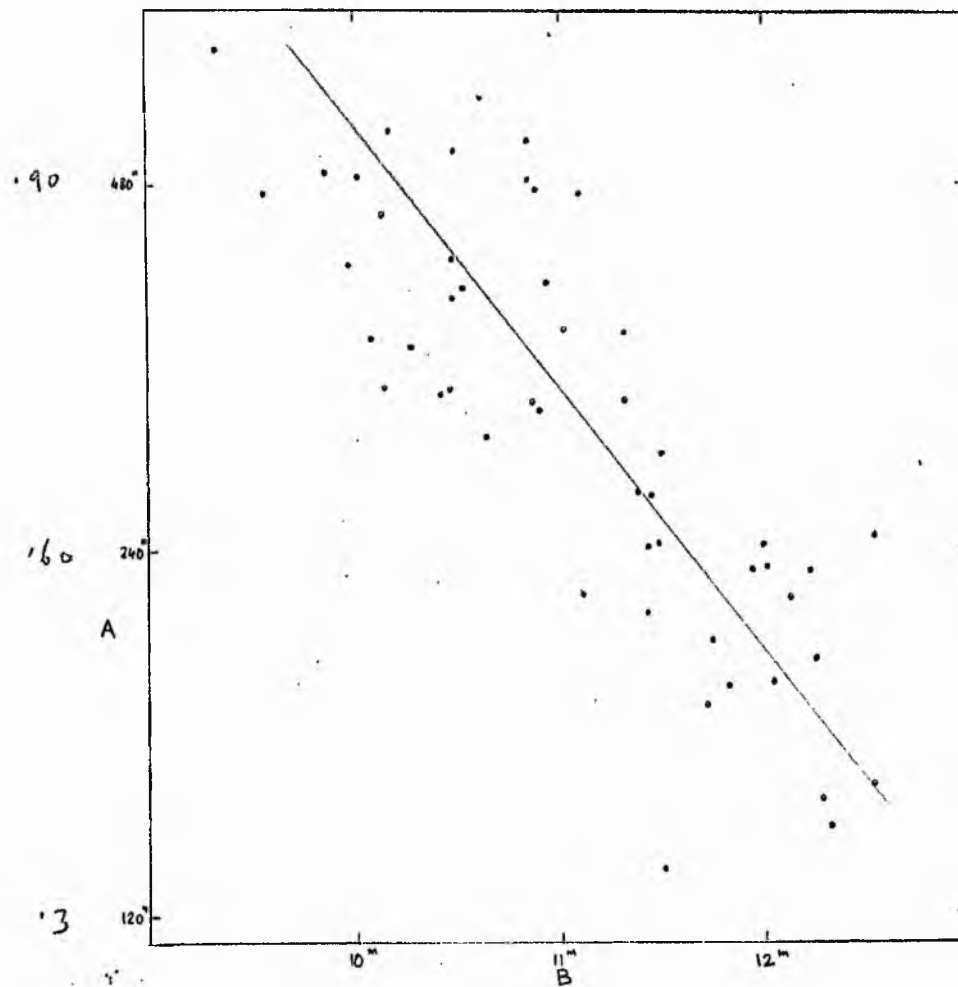


Figure 39 - Diameter A in arc sec of isophote 25 B mag sec⁻² as a function of apparent total magnitude B.

The value for ellipticals is consistent with Heidmann's values for both the Virgo ellipticals and the Coma cluster galaxies, but there is disagreement in the results for spirals. It was mentioned in the previous section on the reduced luminosity profiles that the third component could be due to the large radial distances to which galaxies had been detected. In this connexion, some of the writer's diameters were found to be larger than Holmberg's 1958 values, from which Heidmann derived the q value for spirals of 2.8. It would only require a reduction of 20% in the diameter of the brighter galaxies (i.e. those around 10^m) on figure 39 to change the q value from 1.71 to 2.65. Accordingly, the observed slope of the data in the figure may be due to increased diameter measurements, but the study emphasizes the sensitivity of the q value in the luminosity-diameter relation to the methods in which the diameters are determined.

Several of the previous comparisons have depended upon the accuracy of Holmberg's photometry. For this reason it is interesting to compare the writer's r^* at $\mu_B = 26.5$, interpolated from the integration tables, with the corresponding quantity \sqrt{ab} from Holmberg's

(1958) photometry. The results are shown in figure 40, where the mean systematic difference for 7 ellipticals and lenticulars is $\log r^* - \log \sqrt{ab} = +0.05$;

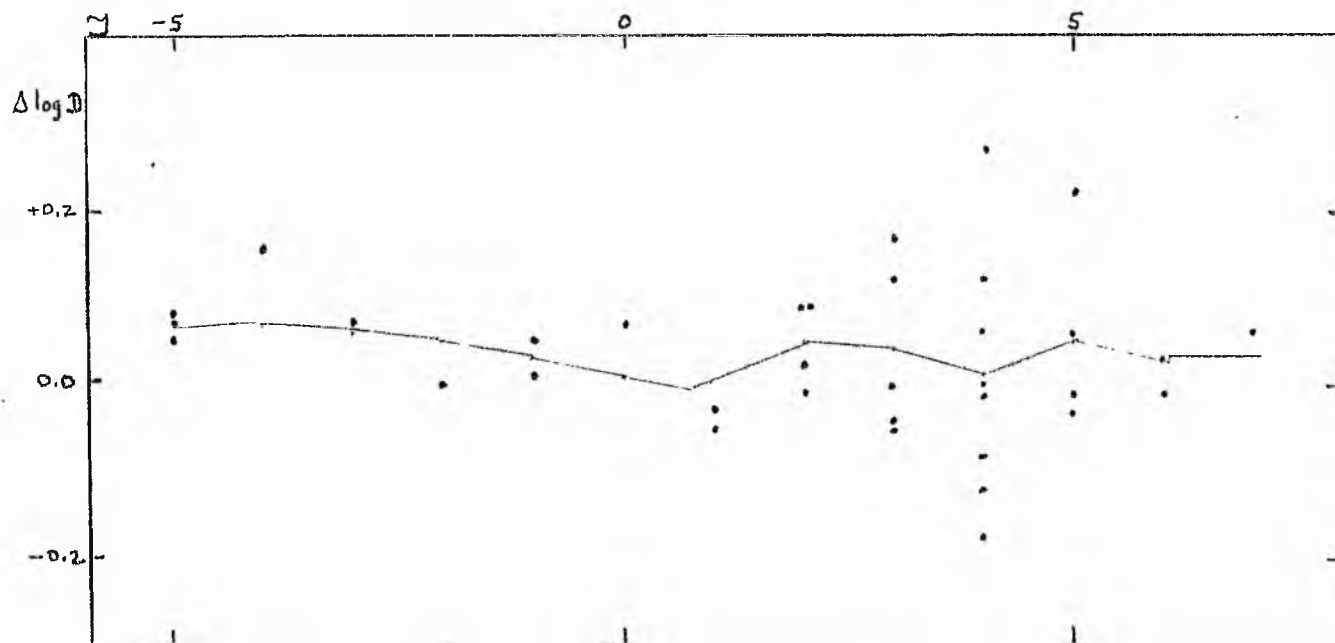


Figure 40 - Comparison with Holmberg's data
at $\mu = 26.5 \text{ mag. sec}^{-2}$.

the corresponding quantity for 27 spirals is $+0.03$, almost independent of morphological type. Near $\mu = 26.5$ the average slope of the luminosity curve is $\Delta \log D = 0.06 \Delta \mu$, so that either Holmberg's statement that his diameters refer to 26.5 is incorrect or there are errors in the present photometric work at this faint level.

In order to check Holmberg's photometry at $\mu = 26.5$, de Vaucouleurs' method (1970), which used the integrated magnitudes listed in the Reference Catalogue (already

described on page 128) has been applied to all the elliptical galaxies with axial ratios close to unity originally measured by Holmberg (1958). The calculated surface brightness for each galaxy listed in table 8 at $\mu = 26.5$ is shown in the third column of the table, while in the fourth Holmberg's value of \sqrt{ab} is given; the writer's measurements for r^* at $\mu = 26.5$ are indicated in column five for objects in common.

Table 8

Values of radial distances at $\mu = 26.5 \text{ mag. sec.}^{-2}$

NGC	E type	r''_{calc}	r'_c/r_H	$r''(H)$	r^*	$r^* (F)$
4278	1-2	208	1.118	186	-	-
4283	0	91	1.338	68	-	-
4472	2	396	1.165	340	431	410
4486	0-1	(399)	1.243	321	439	460
4649	2	370	1.298	285	292	346
		Mean	1.232	$\rightarrow \Delta \log r = 0.090 \rightarrow \Delta \mu =$		

It is noted that in each case, Holmberg's values are systematically too low by an average of 22%. For three objects, the writer's average values are only 3% high; the difference may be due in the case of NGC 4649 to overlap with NGC 4647 and for NGC 4486, the large value for r^* could result from partial detection of

the outer corona. a point already discussed elsewhere in the thesis. However, there seems little doubt that Holmberg's measurements are in error; his assertion that the diameters refer to an approximate background level of $26^m.5$ for the (1958, 1964) 103a0 plates was based upon the assumption of a constant low density being produced by the sky background on standard exposure plates. In fact, since the background density of the standard photographic exposure "reduced the sky fog to a minimum", the background measurements were being made right at the toe of the calibration curve, in the area in which accurate density measurements are exceedingly difficult to achieve. Furthermore, when exposing each plate to the standard 10^m exposure, no effect was applied to take account of possible variations in the sky brightness throughout the entire investigation. In addition, the accurate measurement of the minor axis depends upon whether one of the 15 tracings was actually scanning the true axis itself or scanning parallel to it.

The present data has^{ve} been used by de Vaucouleurs (1971) to calibrate the Reference Catalogue diameter i.e. to find the value of μ which statistically corresponds to the log D values. The results agree at $\mu = 24$ to 25 mag.sec.⁻² with de Vaucouleurs' previous calibration using

the Mount Stromlo photometry.

In the future, it may be possible to derive further relationships between the parameters, morphological types and colours of galaxies following a more accurate calibration of the magnitudes. There is little doubt that many more investigations should be carried out using surface photometry to obtain standard parameters, since the results of such studies will assist our understanding of the structure and evolution of galaxies.

CONCLUDING REMARKS

This thesis is the first attempt to obtain detailed parameters for a large number of objects in two colours. One result of such an investigation is the streamlining of the photometric reductions for one galaxy. All the reductions are performed on the University of St Andrews IBM 360/44 computer and detailed parameters for one galaxy can be produced in five minutes. In the future, data reduction at the University Observatory will be greatly assisted by the use of a Honeywell 316 data processor, which will enable scanned information in IDT work to be read straight from the Joyce Loeb1 into the 360. Hence detailed maps of colour may be obtained if calibration standards were accurately set up.

Further work at St Andrews will use some plates of the Coma cluster of galaxies to be taken in 1971 with the 98" Isaac Newton Telescope, when it is hoped that the better plate scale, together with the improved computational facilities available, will enable more detailed results to be obtained for many elliptical and lenticular galaxies with a wide selection of luminosities. Of course, the need still exists for photo-electric observers to obtain further precise measurements to supplement former photographic studies.

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REFERENCES

- Ables, H.D. and Kron, G.E. 1967, P.A.S.P., 79, 423.
- Arp, H. and Bertola, F. 1969. Ap.Lett., 4, 23.
- . 1971, Ap.J., 163, 195.
- Bergh, S. van den 1969, Ap.Lett., 4, 117.
- Bigay, J.H. 1951, Ann d'ap., 14, 319.
- Bigay, J.H., Dumont, R., Lenouvel, F. and Lunel, M.
1953, Ann d'ap., 16, 133.
- Bigay, J.H. and Dumont, R. 1954, Ann d'ap., 17, 78.
- . 1955, Ann d'ap., 18, 141.
- Bredohl, H.B. 1970, Astron. and Astroph., 7, 167.
- Edelen, D.G.B. 1965, A.J., 70, 747.
- Evans, D.S. 1951, M.N.R.A.S., 111, 526.
- . 1952, M.N.R.A.S., 112, 606.
- Fraser, C.W. 1967, Observatory, 87, 29.
- Freeman, K.C. 1970, Ap.J., 160, 811.
- Heidmann, J. 1967, C.R.Acad.Sci., 265B, 866.
- . 1969, Ap.Lett., 3, 19.
- Hodge, P.W. and Brownlee, D.E. 1966, P.A.S.P., 78, 125.
- Holmberg, E. 1958, Medd.Lund.Obs., Ser.II, No. 136.
- . 1964, Ark.Astr., 3, 387.
- Hopf, A.van 1937, Z.Ap., 14, 104.
- Houten, C.J.van 1961, B.A.N., 16, 1.

Hubble, E. 1930, Ap.J., 71, 231.

Jones, W.B., Obitts, D.L., Gallet, R.M. and Vaucouleurs,
G.de 1967, Publ.Dept. of Astron. Univ.Texas,
Series II, Vol.1, No. 8.

Liller, M.H. 1960, Ap.J., 132, 306.

———. 1966, Ap.J., 146, 28.

Linfoot, E.H. 1955, Recent Advances in Optics (Oxford: O.U.P.).

King, I.R. 1961, A.J., 66, 47.

———. 1962a, A.J., 67, 224.

———. 1962b, A.J. 67, 471.

———. 1966, A.J., 71, 276.

Koelbloed, D. 1965, B.A.N., 18, 62.

Kron, G.E. and Papiashvili, I.I. 1967, P.A.S.P., 79, 9.

Markaryan, B.E., Oganesyan, E. and Arakelyan, S.N. 1965,
Astrophysics, 1, 38.

Miller, W.C. 1966, P.A.S.P., 78, 228.

Mohler, O.C. and Pierce, A.K. 1957, Ap.J., 125, 285.

Oort, J.H. 1940, Ap.J., 91, 273.

Redman, R.O. and Shirley, E.C. 1936, M.N.R.A.S., 96, 588.

———. 1938, M.N.R.A.S., 98, 613,

Reynolds, J.H. 1913, M.N.R.A.S., 74, 132.

Richter, N. and Hogner, W. 1963, A.N. 287, 261.

———. 1964, Mitt Karl-Schwarzchild Obs., Tautenburg, No.10.

———. 1965, Mitt Karl-Schwarzchild Obs., Tautenburg, No. 23.

———. 1966, Mitt Karl-Schwarzchild Obs., Tautenburg, No. 32.

Rood, H.J. and Baum, W.A. 1967, A.J., 72, 398.

———. 1968, A.J., 73, 442.

Rood, H.J. and Sastry, C.V. 1967, A.J., 72, 233.

Sandage, A., Freeman, K.C. and Stokes, N.R. 1970,
Ap.J., 160, 831.

Sersic, J.L. 1968, 'Atlas de Galaxias Australes', Cordoba.

Shobbrook, R.R. 1965, M.N.R.A.S., 131, 293.

Stebbins, J. and Whitford, A.E. 1937, Ap.J., 86, 237.

———. 1952, Ap.J., 115, 284.

Tifft, W.G. 1961, A.J., 66, 390.

———. 1963, A.J., 68, 302.

Vaucouleurs, G.de 1948, Ann.d'ap., 11, 247.

———. 1948a, J.des Observ., 31, 113.

———. 1957, Ann.Obs.Houga, II, Part 1.

———. 1959, A.J., 64, 397.

———. 1959a, Hdb.d. Phys., 53, 275.

———. 1960, Ap.J., 131, 574.

———. 1961a, Ap.J. Supp., 5, 233.

———. 1961b, A.J., 133, 405.

———. 1961c, Problems of Extra-Galactic Research:
I.A.U. Symp. No. 15.

———. 1963, Ap.J. Supp., 8, 31.

———. 1963b, Ap.J., 137, 720.

———. 1963c, Ap.J., 138, 934.

———. 1964, Ap.J., 139, 899.

Vaucouleurs, G.de 1968, Applied Optics, 7, 1513.

———. 1968a, I.A.U. Comm. 28, Circular No. 6, Supp. I.

———. 1969, Ap.Lett., 4, 17.

———. 1970, I.A.U. Comm. 28, Circular No. 7a.

———. 1970a, Bull. A.A.S., 2, 308.

———. 1970b, Stars and Stellar Systems, Vol.9, Chap.17
(Chicago: Univ.Chicago Press).

———. 1970c, Ap.Lett., 5, 219.

———. 1971, private communication.

Vaucouleurs, G.de, Angione, R. and Fraser, C.W. 1968,
Ap. Lett., 2, 141.

Vaucouleurs, G.de, Griboval, P. and White, T. 1965, P.A.S.P.,
77, 115.

Vaucouleurs, G. de and Heidmann, J. and N. 1970, private
communication.

Vaucouleurs, G.de and Page, J. 1962, Ap.J., 136, 107.

Vaucouleurs, G.de and Malik, G. 1969, M.N.R.A.S., 142, 387.

Vaucouleurs, G.de and A. de 1964, Reference Catalogue of
Bright Galaxies (Austin: Univ.Texas Press).

Walker, M.F. 1965, informal meeting organized by Dr Merle
Walker at Third Symposium on Photo-electronic Devices
held in London.

Walker, M.F. and Kron, G.E. 1967, P.A.S.P., 79, 551.

Williams, R. and Hiltner, W.A. 1940, Pub.Univ.Mich.Obs.,
8, 45.

PHOTOGRAPHIC SURFACE PHOTOMETRY OF GALAXIES

IN THE VIRGO CLUSTER

by

Christopher W. Fraser

VOLUME TWO



Tn 5840

PHOTOMETRIC DATA FOR GALAXIES

NGC 4189 to NGC 4559

EXPLANATORY NOTES FOR USE WITH THE PHOTOMETRIC DATA

The data for each galaxy occupy twelve pages, divided equally between the B and the V colours. The galaxies are listed in increasing order of NGC identification numbers and each page may be described as follows:-

1. Luminosity profile of major axis. The y-axis denotes $\log I$, with the value of $\log I = -2.00$ at the origin. Each division on the y-axis represents a value of $+1.00$ in $\log I$. For the x-axis, which denotes distance r , the centre of the galaxy is plotted at the centre of the axis. Each division represents 50 arc sec.
2. Luminosity profile of minor axis. The units for the x and y axes are as before.
3. Equivalent luminosity profile. The y-axis denotes increasing $\log I$, each division representing an integer value from a minimum at the origin of $\log I = -2.00$. The distance r is given on the x-axis, from zero at the origin and increasing by 50 arc sec per division.
4. Relative integrated luminosity curve $k(r)$ is plotted as ordinate, each positive division in y

representing a value of 0.1, and r^* , as in case 3, denotes increasing distance r^* , in 50 arc sec steps.

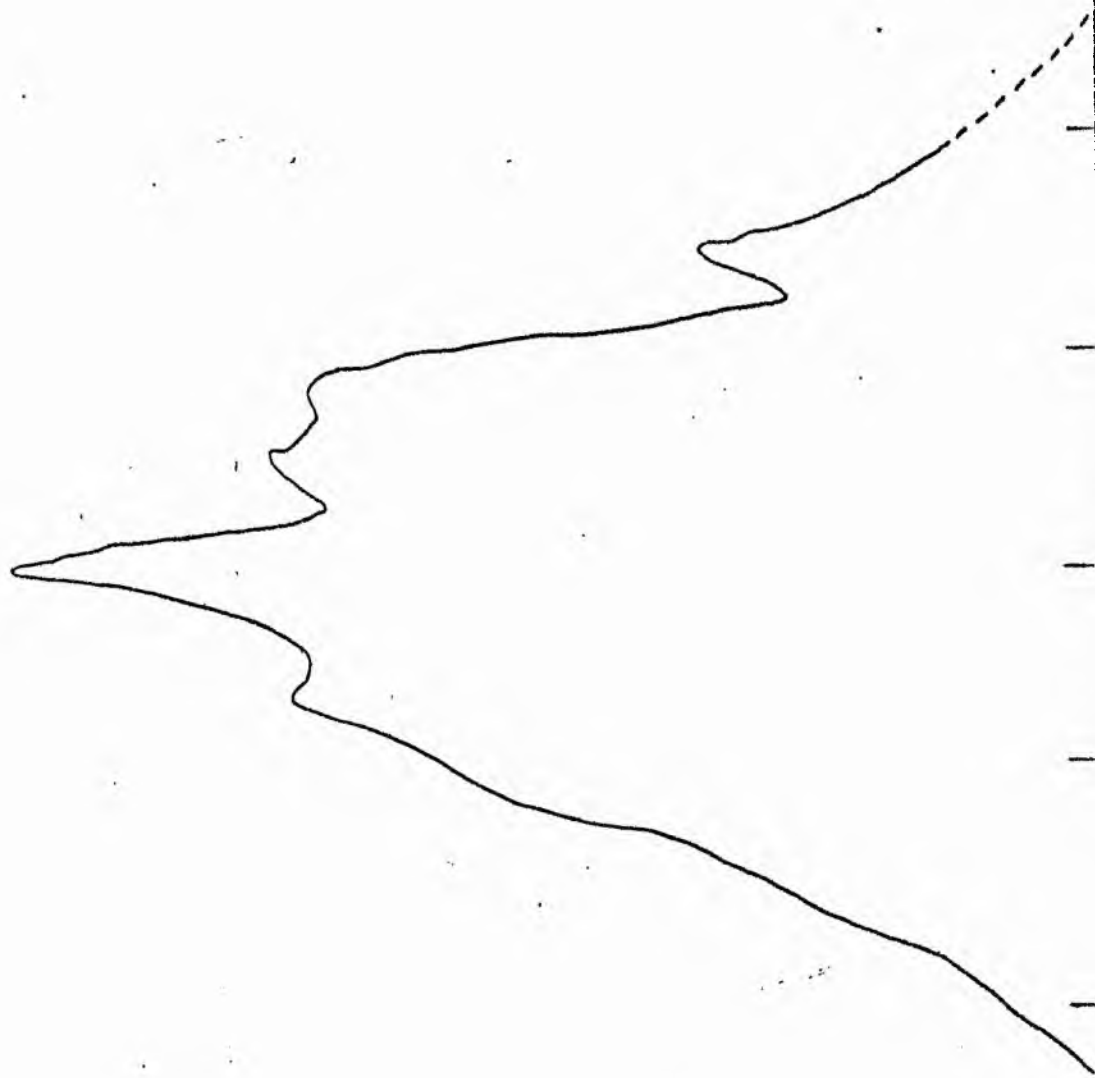
5. Table of mean luminosity distribution. The areas in column 5 are expressed in square seconds of arc, and in the last column, μ is in mag. sec^{-2} .

6. Table of photometric parameters. The units are as follows:-

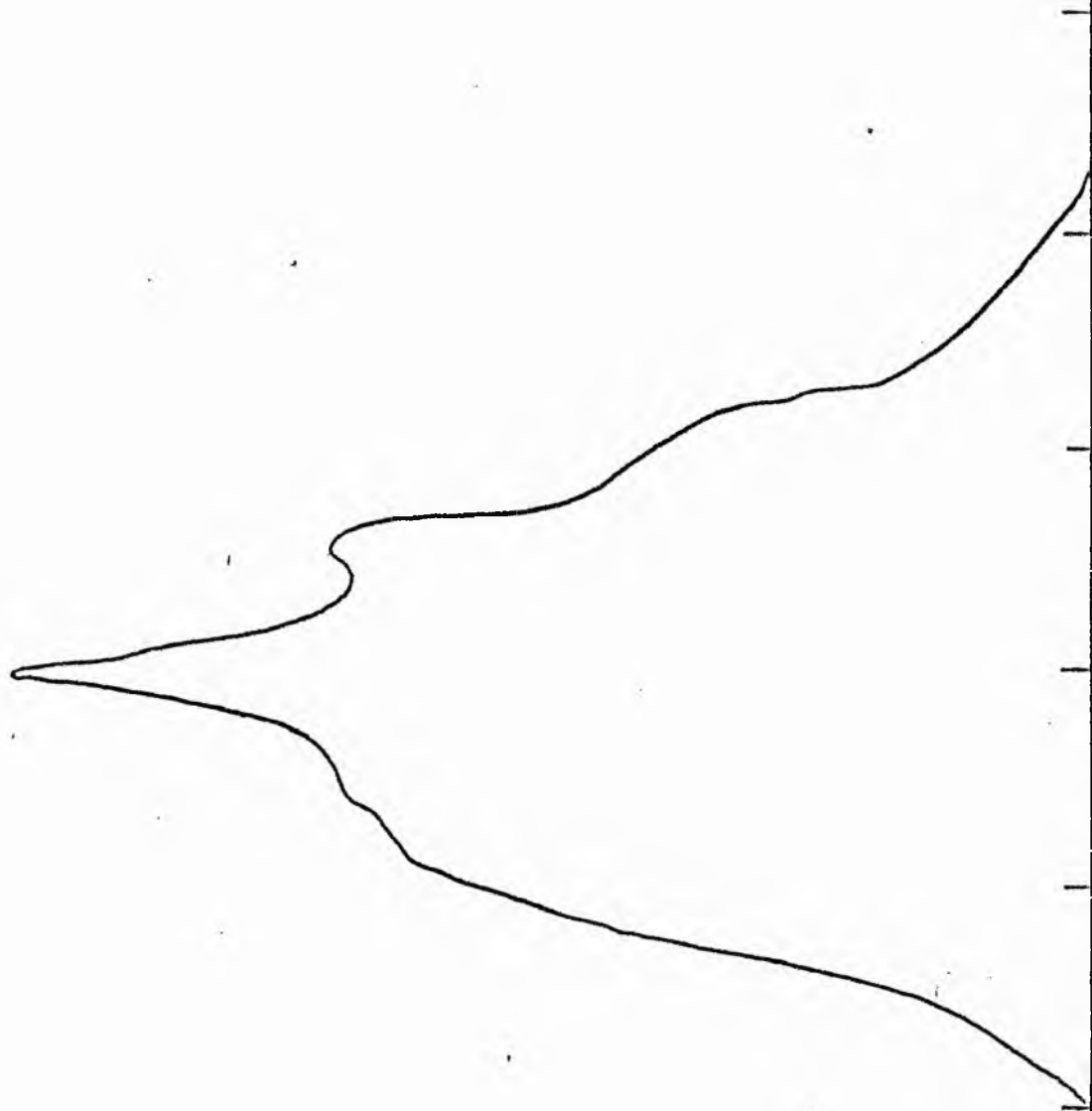
- a. The total luminosity is expressed in sky units per square minute.
- b. The apparent central surface brightness is in mag. per square second, as are the surface brightnesses at $k = \frac{1}{4}, \frac{1}{2}$ and $\frac{3}{4}$.
- c. The major axis and minor axis at threshold together with the major axis at $\mu = 25.0 \text{ mag. sec}^{-2}$, and the semi-major axis measurements at $k = \frac{1}{4}, \frac{1}{2}$ and $\frac{3}{4}$, are expressed in minutes of arc.
- d. The gradient of the exponential component $G(a) = d \log I/da$ is expressed per minute of arc, as is the equivalent gradient of the exponential component $G(r^*) = d \log I/dr^*$.
- e. The mean surface brightness is in magnitudes per square minute.

NGC 4189
B-Filter
Axis 1

Luminosity profile

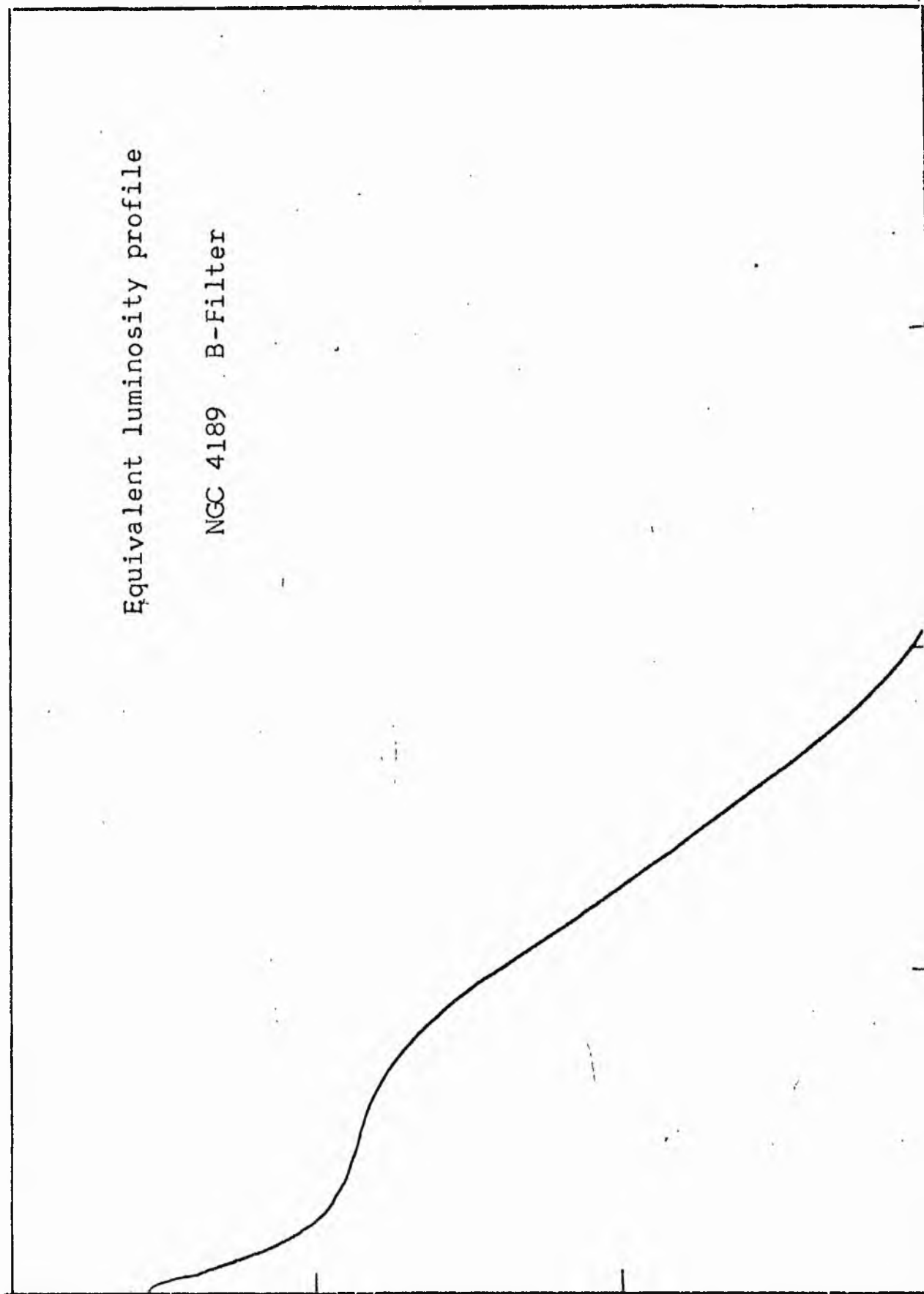


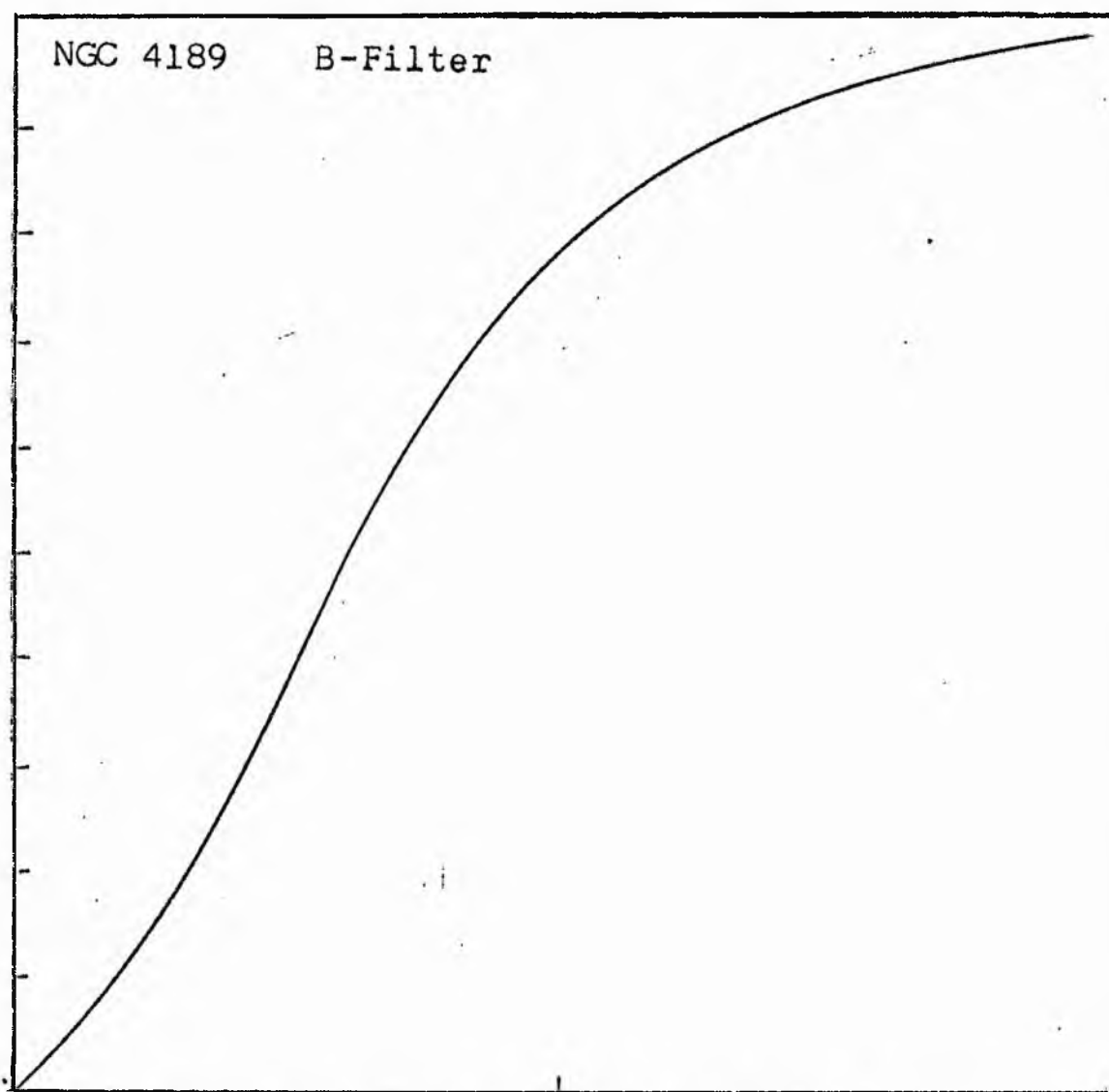
NGC 4189
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4189 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4189
B CLUSTER

LOG I	I	T	z	AREA	A	P	ΣP	K(R)	P	LOG J	μ
0.48	3.020	2.765	0.6	1.5	20.27	56.0604	0.0	0.0	0.0	0.713	20.73
0.40	2.512	2.254	2.44	20.27	41.11	92.6379	56.06	0.01	0.07	0.633	20.93
0.30	1.995	1.790	4.42	01.38	71.77	128.4659	148.73	0.03	0.13	0.533	21.18
0.20	1.585	1.422	6.51	133.14	75.13	117.5176	277.16	0.05	0.19	0.433	21.43
0.10	1.259	1.129	8.22	212.27	93.77	105.9118	389.68	0.07	0.24	0.333	21.68
0.00	1.000	0.897	9.87	316.04	716.36	642.6917	495.59	0.08	0.29	0.233	21.93
-0.10	0.794	0.713	12.04	1022.40	1931.48	1305.1909	1138.28	0.19	0.53	0.133	22.18
-0.20	0.631	0.566	30.14	2853.68	1426.05	807.2495	2443.48	0.42	0.88	0.033	22.43
-0.30	0.501	0.452	36.91	4279.94	550.02	247.3145	3250.73	0.55	1.08	-0.067	22.68
-0.40	0.398	0.357	39.21	4929.96	1068.34	381.5176	3498.04	0.59	1.15	-0.167	22.93
-0.50	0.316	0.284	43.33	5858.30	1032.61	292.9609	3879.62	0.66	1.27	-0.267	23.18
-0.60	0.251	0.225	46.97	6930.91	1767.73	398.3796	4172.58	0.71	1.37	-0.367	23.43
-0.70	0.200	0.179	52.62	8698.64	1090.16	195.1466	4570.95	0.78	1.54	-0.467	23.68
-0.80	0.158	0.142	55.82	9788.80	1180.47	167.8527	4766.09	0.81	1.63	-0.567	23.93
-0.90	0.126	0.113	59.09	10969.27	1279.02	144.4607	4933.94	0.84	1.73	-0.667	24.18
-1.00	0.100	0.090	62.44	12248.25	1378.48	123.6725	5078.40	0.86	1.83	-0.767	24.43
-1.10	0.079	0.071	65.86	13626.17	1705.50	121.5418	5202.07	0.88	1.93	-0.867	24.68
-1.20	0.063	0.057	69.86	15332.27	1377.17	77.9581	5323.61	0.90	2.04	-0.967	24.93
-1.30	0.050	0.045	72.93	16709.45	1709.59	76.8716	5401.57	0.92	2.13	-1.067	25.18
-1.40	0.040	0.036	76.57	18419.04	1833.18	65.4756	5478.44	0.93	2.24	-1.167	25.43
-1.50	0.032	0.028	80.29	20252.21	1957.09	55.5247	5543.91	0.94	2.35	-1.267	25.68
-1.60	0.025	0.023	84.08	22209.30	2006.01	47.0101	5599.43	0.95	2.46	-1.367	25.93
-1.70	0.020	0.018	87.94	24245.31	2225.79	39.8438	5646.44	0.96	2.57	-1.467	26.18
-1.80	0.016	0.014	91.88	26521.11	2371.52	33.7213	5686.29	0.97	2.69	-1.567	26.43
-1.90	0.013	0.011	95.90	28992.63	2517.00	28.4290	5720.00	0.97	2.81	-1.667	26.68
-2.00	0.010		99.99	31409.63			5748.43	0.98	2.92	-1.767	26.93
-∞							5884.00	(1)			∞

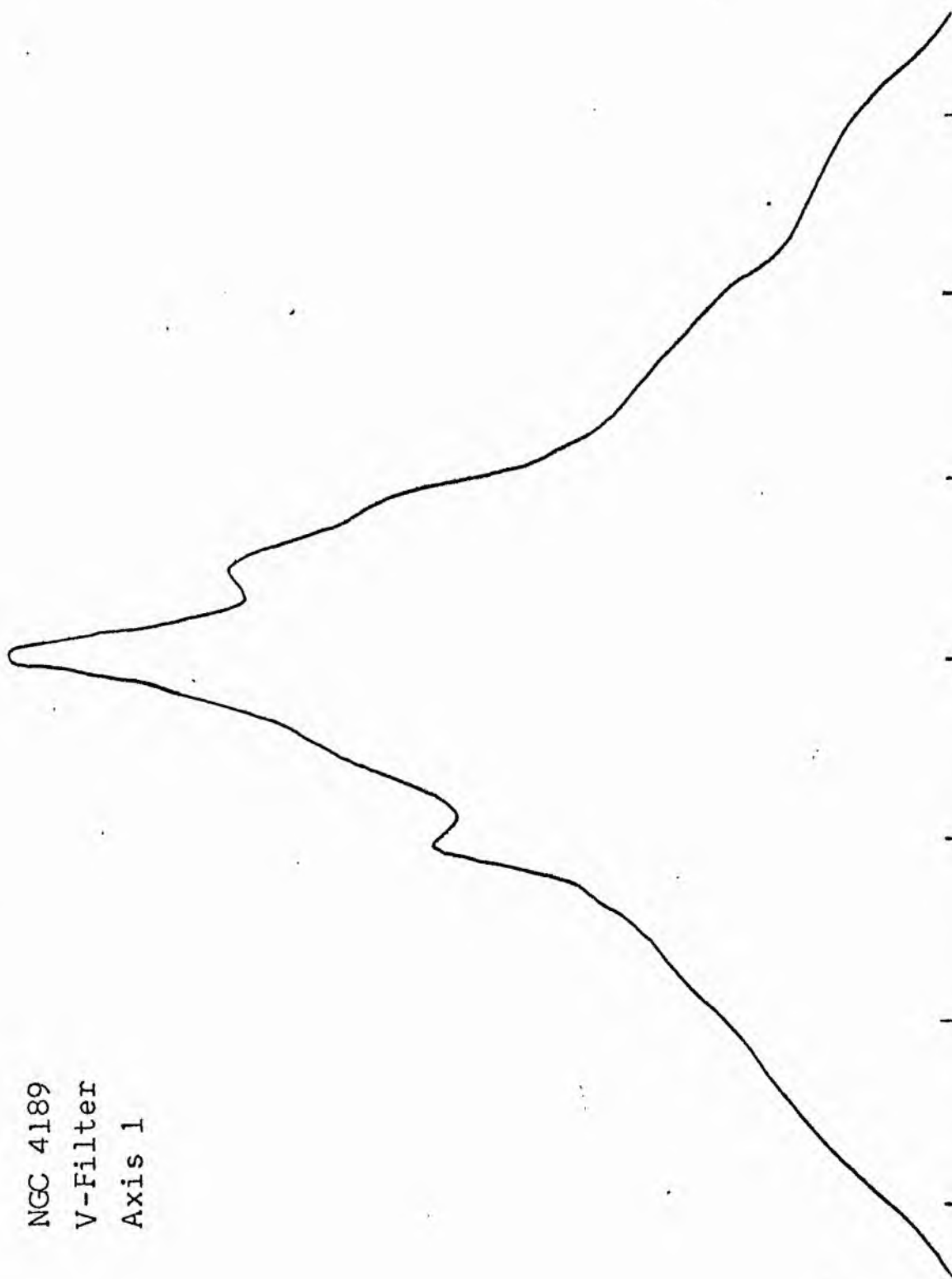
52

26

PHOTOMETRIC PARAMETERS OF NGC 4189

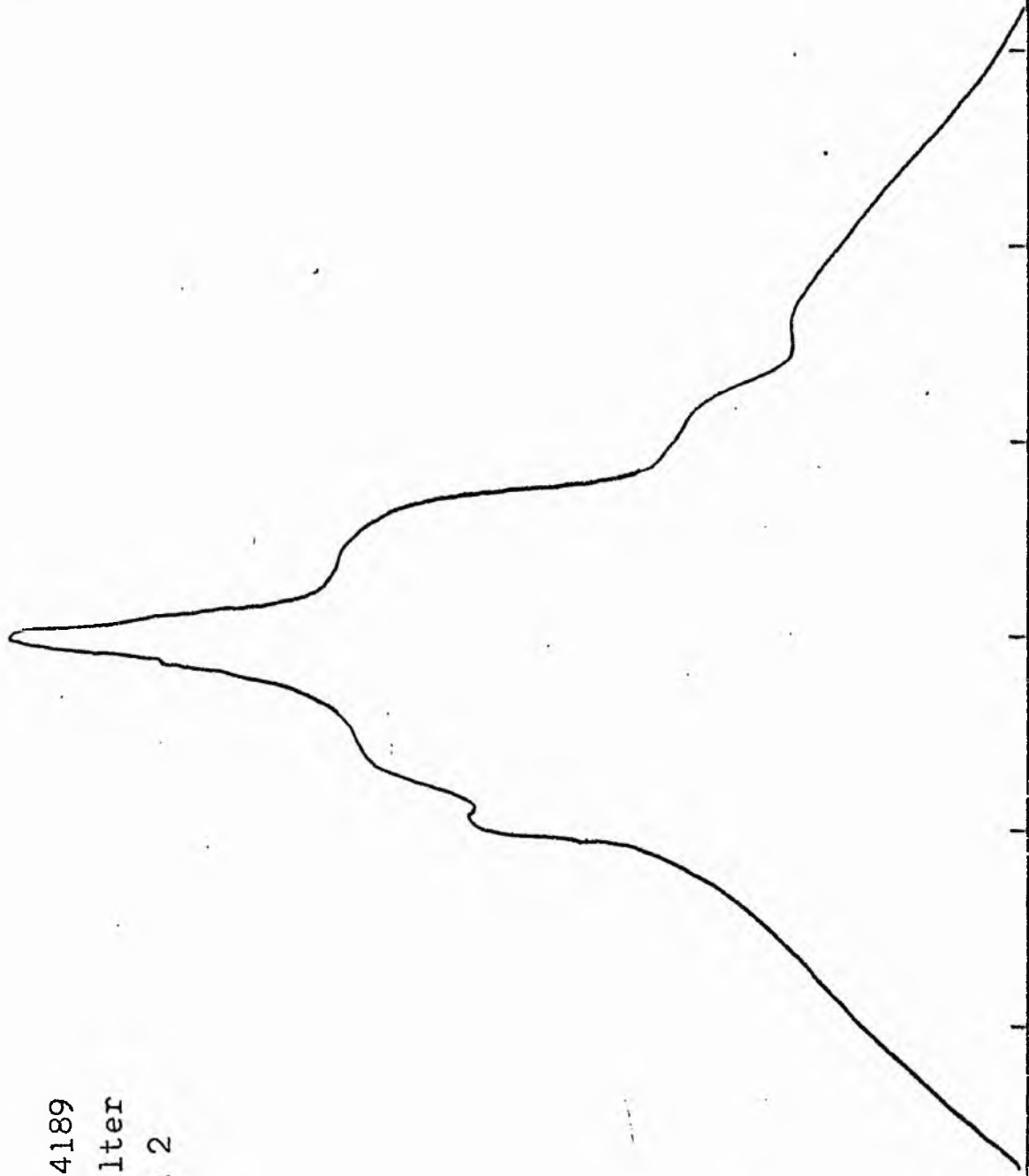
B-FILTER

Total luminosity	L_T	= 1.63
Total apparent magnitude	m_T	= 12.53
Apparent central surface brightness	μ_0	= 20.73
Major axis at threshold	$2a_m$	= 4.01
Minor axis at threshold	$2b_m$	= 3.38
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 2.56
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.91
Gradient of exponential component	$G(a)$	= -1.72
Equivalent gradient of exponential comp.	$G(r^*)$	= -1.57
Equivalent gradient of reduced exp. comp.	$G(p)$	= -0.89
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.32
Axis ratio	b/a	= 0.81
Equivalent radius	r_1^*	= 0.36
Surface brightness	μ_1	= 22.24
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.71
Axis ratio	b/a	= 0.78
Equivalent radius	r_e^*	= 0.57
Surface brightness	μ_e	= 22.55
Mean surface brightness	μ_e'	= 13.32
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.93
Axis ratio	b/a	= 0.78
Equivalent radius	r_3^*	= 0.83
Surface brightness	μ_3	= 23.57
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.58 \\ = 1.47 \end{matrix}$



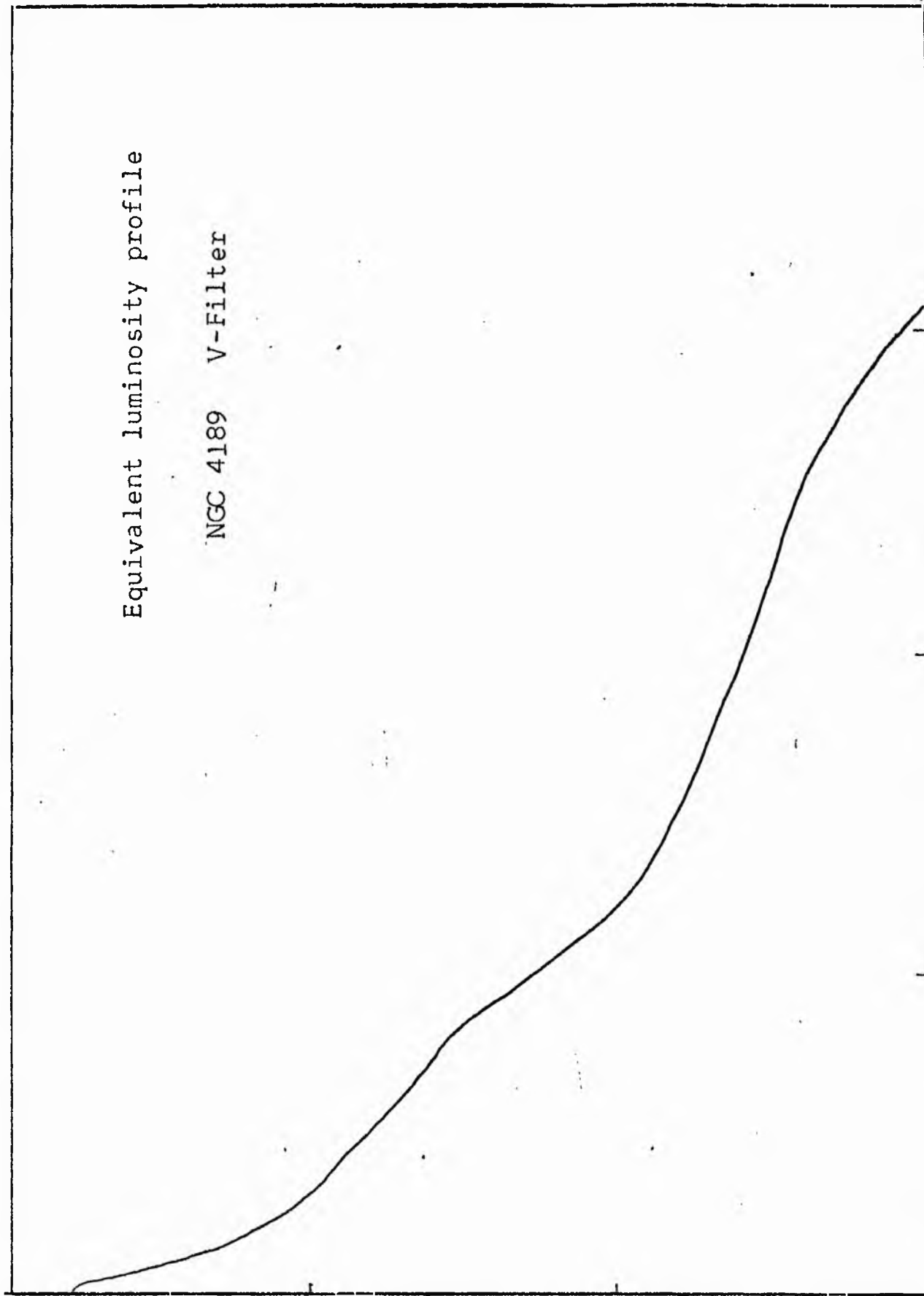
NGC 4189
V-Filter
Axis 1

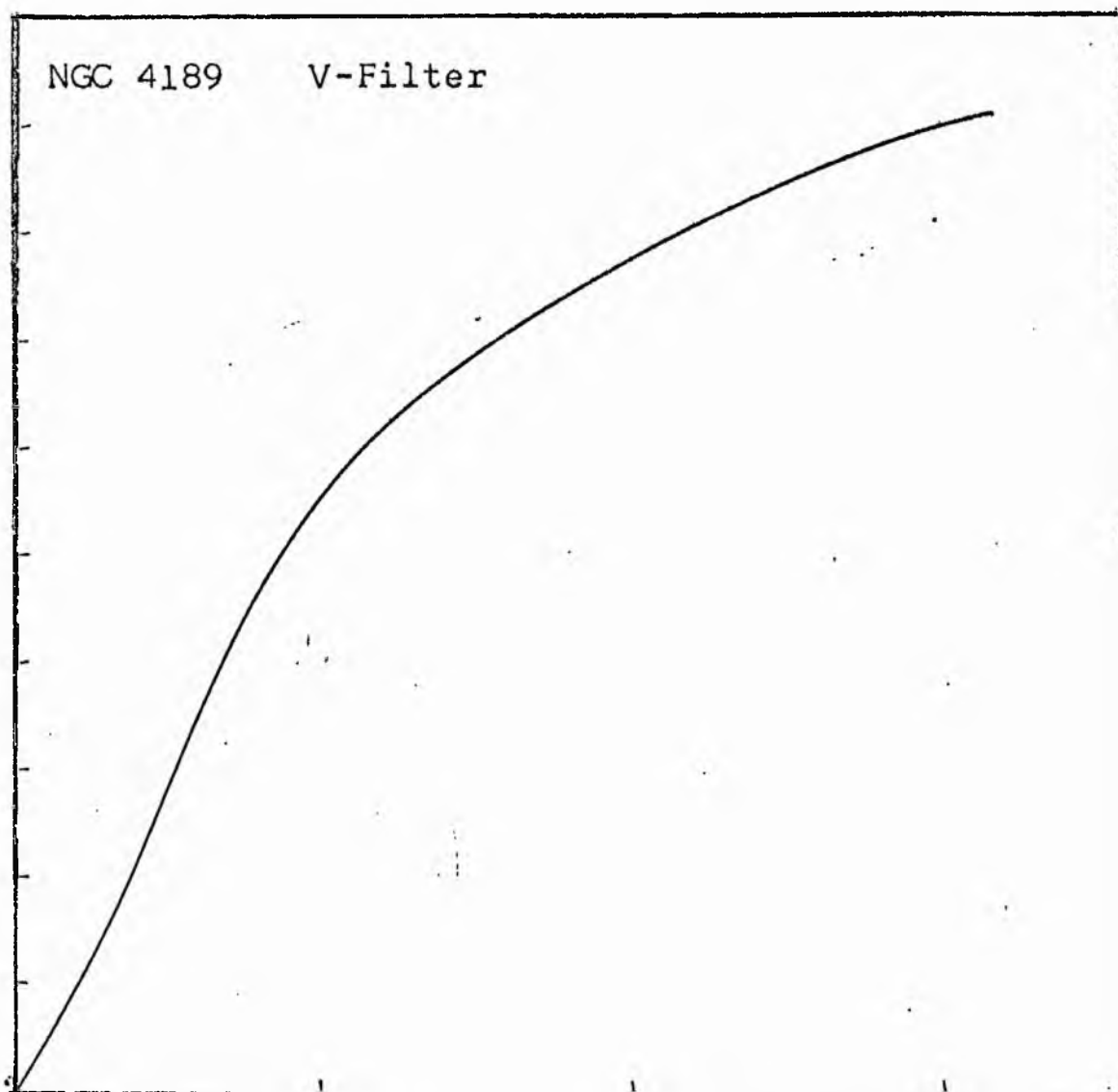
NGC 4189
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4189 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4189
V COLOUR

LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.72	5.248		0.0	0.0			0.0	0.0	0.0	1.262	19.46
0.70	5.012	5.130	1.87	10.99	10.99	56.3570	56.36	0.01	0.04	1.242	19.51
0.60	3.981	4.496	2.78	24.28	13.29	59.7749	116.13	0.02	0.07	1.142	19.76
0.50	3.162	3.572	4.20	55.42	31.14	111.2153	227.35	0.03	0.10	1.042	20.01
0.40	2.512	2.837	5.67	101.00	45.58	129.3171	356.66	0.05	0.13	0.942	20.26
0.30	1.995	2.254	7.11	158.81	57.82	130.2913	486.96	0.06	0.17	0.842	20.51
0.20	1.585	1.790	9.13	261.07	103.06	104.4840	671.44	0.09	0.21	0.742	20.76
0.10	1.259	1.422	11.91	445.63	183.76	261.2839	932.72	0.12	0.28	0.642	21.01
0.00	1.000	1.129	16.39	843.93	398.30	449.8684	1382.59	0.18	0.39	0.542	21.26
-0.10	0.794	0.897	20.41	1300.69	464.75	416.9600	1799.55	0.23	0.48	0.442	21.51
-0.20	0.631	0.713	25.00	1963.50	654.81	466.6462	2266.2	0.29	0.59	0.342	21.76
-0.30	0.501	0.566	29.85	2799.23	835.73	473.0847	2739.28	0.36	0.70	0.242	22.01
-0.40	0.398	0.450	35.58	3977.05	1177.83	529.6062	3268.89	0.43	0.84	0.142	22.26
-0.50	0.316	0.357	40.56	5168.27	1191.22	425.4656	3694.35	0.48	0.95	0.042	22.51
-0.60	0.251	0.284	45.05	6375.86	1207.59	342.6035	4036.96	0.52	1.06	-0.058	22.76
-0.70	0.200	0.225	49.90	7822.59	1446.73	326.0317	4362.79	0.57	1.17	-0.158	23.01
-0.80	0.158	0.179	51.07	8193.73	371.13	66.4358	4429.42	0.58	1.20	-0.258	23.26
-0.90	0.126	0.142	54.30	9262.95	1069.22	152.0340	4581.45	0.60	1.28	-0.358	23.51
-1.00	0.100	0.113	59.60	11159.43	1896.48	214.2012	4795.65	0.62	1.40	-0.458	23.76
-1.10	0.079	0.090	67.80	14441.39	3281.95	294.4456	5090.10	0.66	1.60	-0.558	24.01
-1.20	0.063	0.071	76.50	18385.39	3744.00	281.0674	5371.16	0.70	1.80	-0.658	24.26
-1.30	0.050	0.057	89.20	24996.51	6611.12	374.2395	5745.40	0.75	2.10	-0.758	24.51
-1.40	0.040	0.045	101.44	32327.21	7330.70	329.6250	6075.03	0.79	2.39	-0.858	24.76
-1.50	0.032	0.036	114.70	41331.07	9003.86	321.5411	6396.62	0.83	2.70	-0.958	25.01
-1.60	0.025	0.028	125.06	49765.14	8434.07	239.2838	6635.90	0.86	2.96	-1.058	25.26
-1.70	0.020	0.023	135.11	57348.86	7583.72	170.9067	6806.80	0.89	3.18	-1.158	25.51
-1.80	0.016	0.018	142.47	63767.09	6418.23	114.8928	6921.70	0.90	3.35	-1.258	25.76
-1.90	0.013	0.014	147.86	68683.25	4916.16	69.9042	6991.60	0.91	3.48	-1.358	26.01
-2.00	0.010	0.011	155.30	75769.19	7085.94	80.0342	7071.63	0.92	3.66	-1.458	26.26
-∞							7691.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4189

V-FILTER

2

Total luminosity	L_T	= 2.13
Total apparent magnitude	m_T	= 11.54
Apparent central surface brightness	μ_0	= 19.46
Major axis at threshold	$2a_m$	= 5.8
Minor axis at threshold	$2b_m$	= 3.83
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.1
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.68
Equivalent gradient of exponential comp.	$G(r^*)$	= -0.80
Equivalent gradient of reduced exp. comp.	$G(\rho)$	= -0.66
Parameters at $k = \frac{1}{4}$:		

Semi-major axis	a_1	= 0.383
Axis ratio	b/a	= 0.69
Equivalent radius	r_1^*	= 0.36
Surface brightness	μ_1	= 21.59

Parameters at $k = \frac{1}{2}$ (effective) :

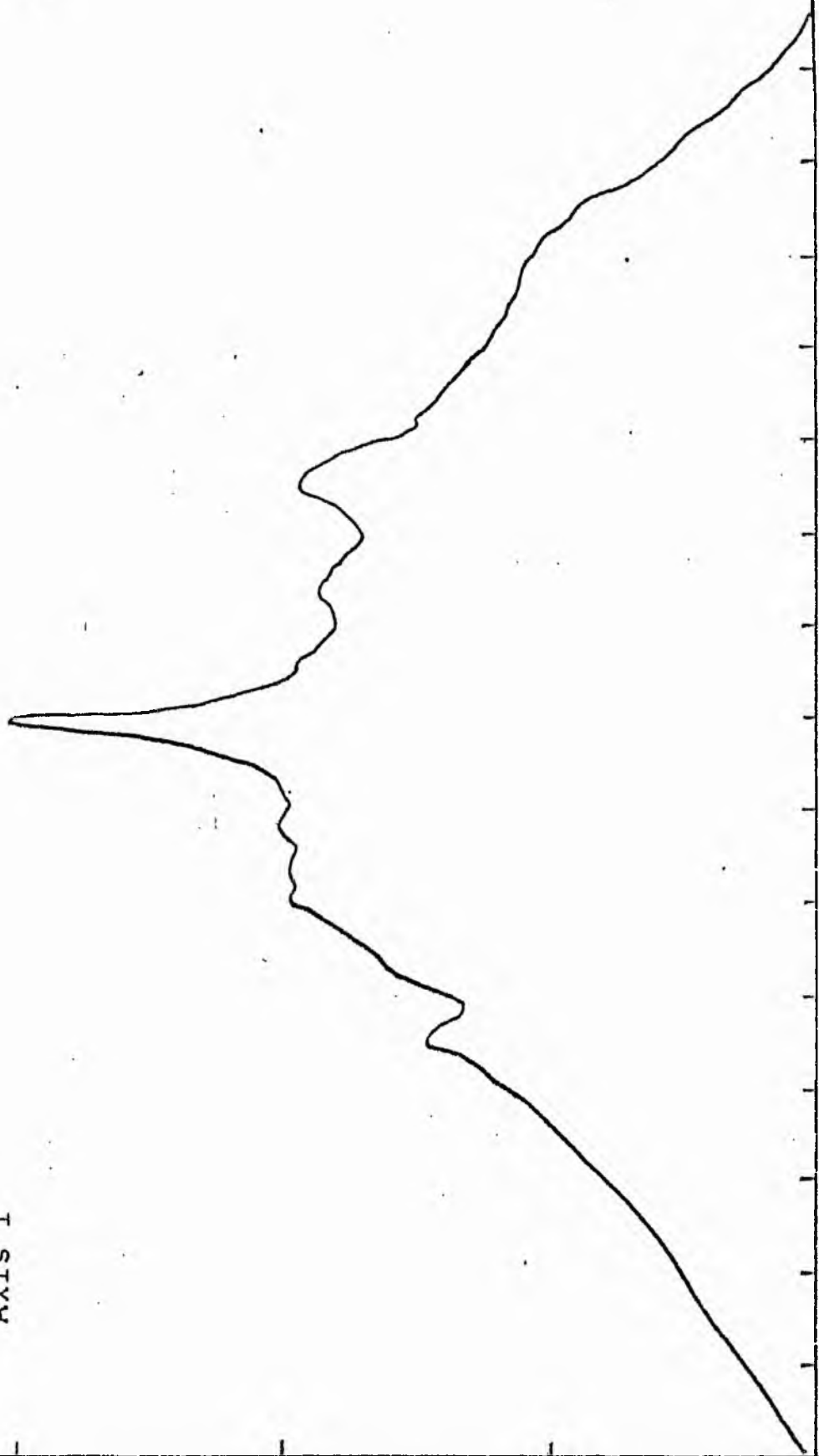
Semi-major axis	a_e	= 0.82
Axis ratio	b/a	= 0.87
Equivalent radius	r_e^*	= 0.70
Surface brightness	μ_e	= 22.63
Mean surface brightness	μ_e'	= 12.76

Parameters at $k = \frac{3}{4}$:

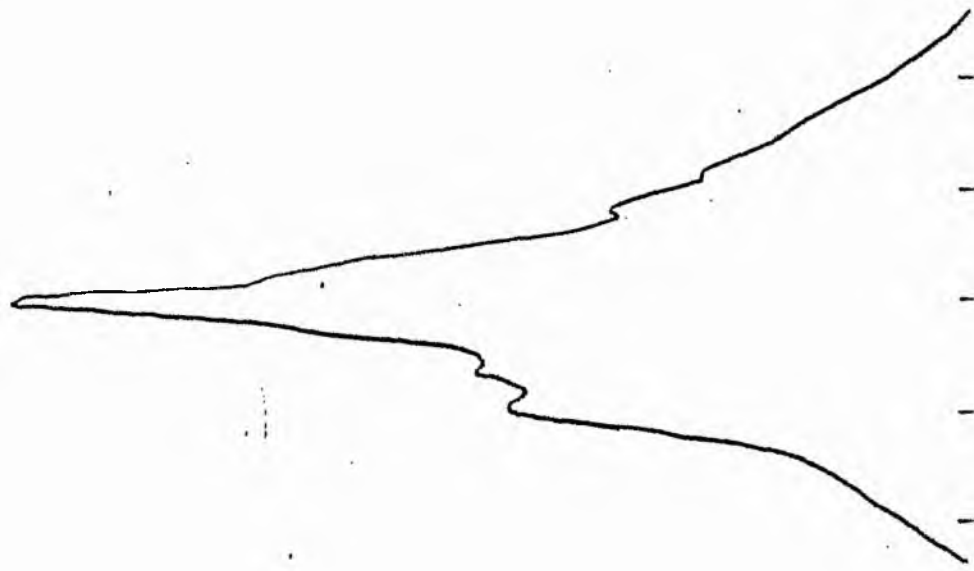
Semi-major axis	a_3	= 1.75
Axis ratio	b/a	= 0.76
Equivalent radius	r_3^*	= 1.50
Surface brightness	μ_3	= 24.51

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.97 \\ = 2.12 \end{cases}$
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NGC 4192
B-Filter
Axis 1

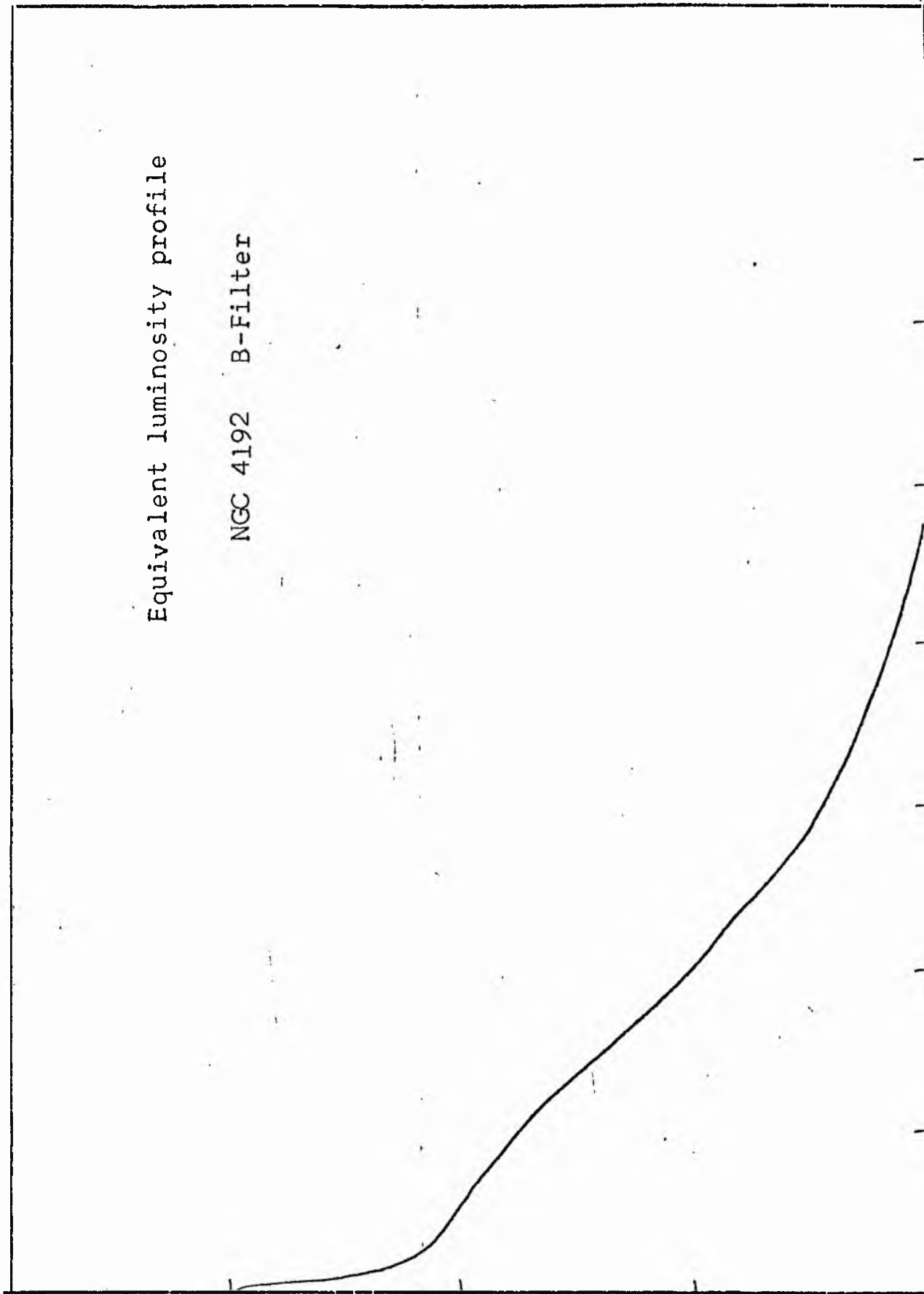


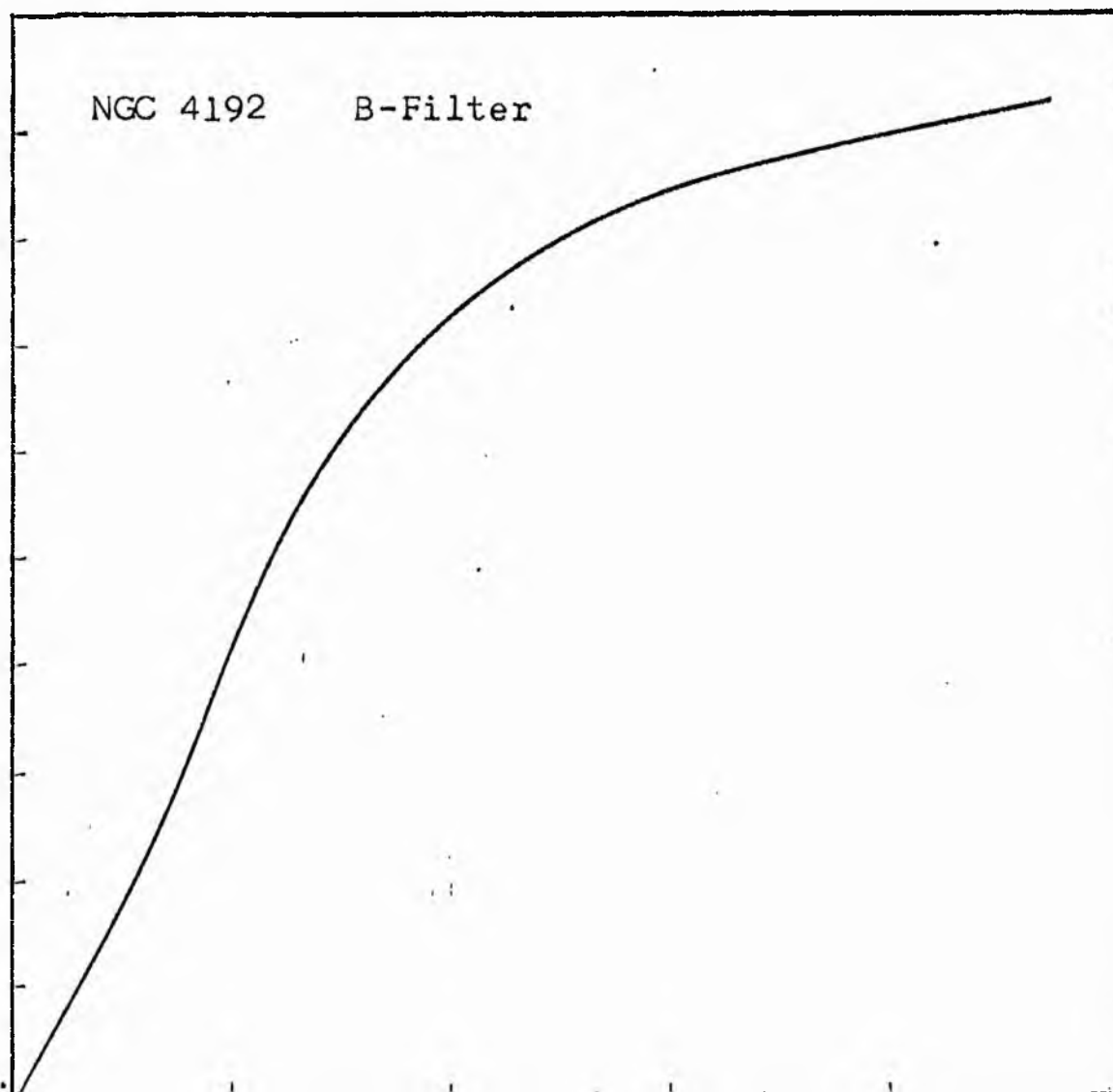
NGC 4192
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4192 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4192
& COLOUR

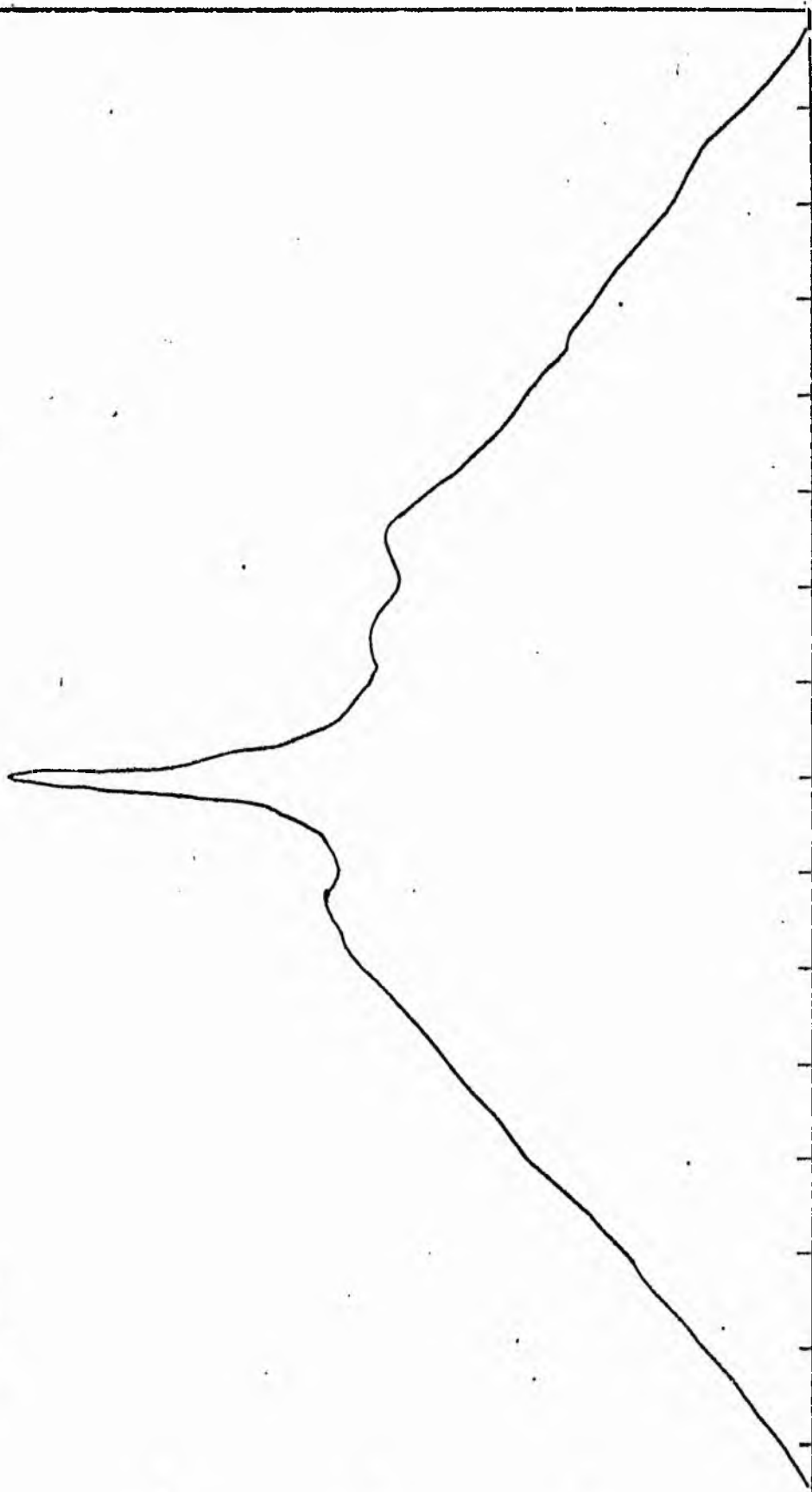
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R-I)	ρ	LOG J	μ
0.99	9.772		0.0	0.0			0.0	0.0	0.0	1.323	19.14
0.90	7.943	8.858	1.96	12.07	12.07	106.9026	106.90	0.01	0.03	1.233	19.37
0.80	6.310	7.126	3.47	37.83	25.76	183.5683	290.47	0.01	0.06	1.133	19.62
0.70	5.012	5.661	4.88	74.82	36.99	209.3760	499.85	0.02	0.08	1.033	19.87
0.60	3.981	4.496	7.16	161.06	86.24	387.7781	887.62	0.04	0.12	0.933	20.12
0.50	3.162	3.572	8.39	221.14	60.09	214.6134	1102.24	0.05	0.14	0.833	20.37
0.40	2.512	2.837	9.13	261.87	40.73	115.5549	1217.79	0.06	0.15	0.733	20.62
0.30	1.995	2.254	11.16	391.27	129.40	291.6084	1509.40	0.07	0.19	0.633	20.87
0.20	1.585	1.790	14.97	704.03	312.76	559.8687	2069.27	0.10	0.25	0.533	21.12
0.10	1.259	1.422	16.96	903.65	199.62	283.8379	2353.11	0.11	0.28	0.433	21.37
0.00	1.000	1.129	28.21	2500.09	1596.44	1803.1194	4156.23	0.20	0.47	0.333	21.62
-0.10	0.794	0.897	39.40	4876.88	2376.79	2132.3708	6288.59	0.31	0.66	0.233	21.87
-0.20	0.631	0.713	49.10	7573.78	2696.90	1921.9260	8210.52	0.40	0.82	0.133	22.12
-0.30	0.501	0.566	57.70	10459.27	2885.49	1633.3975	9843.91	0.48	0.96	0.033	22.37
-0.40	0.398	0.450	63.70	12747.61	2288.34	1028.9446	10872.86	0.53	1.06	-0.067	22.62
-0.50	0.316	0.357	70.80	15747.66	3000.06	1071.5244	11944.38	0.58	1.18	-0.167	22.87
-0.60	0.251	0.284	77.67	18952.05	3204.39	909.1118	12853.49	0.63	1.30	-0.267	23.12
-0.70	0.200	0.225	84.53	22447.69	3495.64	787.7688	13641.25	0.67	1.41	-0.367	23.37
-0.80	0.158	0.179	89.40	25108.73	2661.04	476.3474	14117.60	0.69	1.49	-0.467	23.62
-0.90	0.126	0.142	95.60	28712.13	3603.41	512.3723	14629.97	0.71	1.60	-0.567	23.87
-1.00	0.100	0.113	106.30	35499.00	6786.87	766.5532	15396.52	0.75	1.77	-0.667	24.12
-1.10	0.079	0.090	113.40	40399.50	4900.49	439.6558	15836.17	0.77	1.89	-0.767	24.37
-1.20	0.063	0.071	118.70	44264.06	3864.57	275.4065	16111.58	0.79	1.98	-0.867	24.62
-1.30	0.050	0.057	127.03	50694.68	6430.62	364.0212	16475.60	0.80	2.12	-0.967	24.87
-1.40	0.040	0.045	135.67	57825.25	7130.57	320.6262	16796.22	0.82	2.26	-1.067	25.12
-1.50	0.032	0.036	146.27	67214.06	9388.81	335.3406	17131.56	0.84	2.44	-1.167	25.37
-1.60	0.025	0.028	159.63	80053.19	12839.12	364.2598	17495.82	0.85	2.66	-1.267	25.62
-1.70	0.020	0.023	173.42	94481.81	14428.62	325.1631	17820.98	0.87	2.89	-1.367	25.87
-1.80	0.016	0.018	198.48	123760.87	29279.06	524.1243	18345.10	0.90	3.31	-1.467	26.12
-1.90	0.013	0.014	216.70	147525.69	23764.81	337.9187	18683.02	0.91	3.62	-1.567	26.37
-2.00	0.010	0.011	234.70	173051.75	25526.06	288.3115	18971.33	0.93	3.92	-1.667	26.62
-∞							20471.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4192

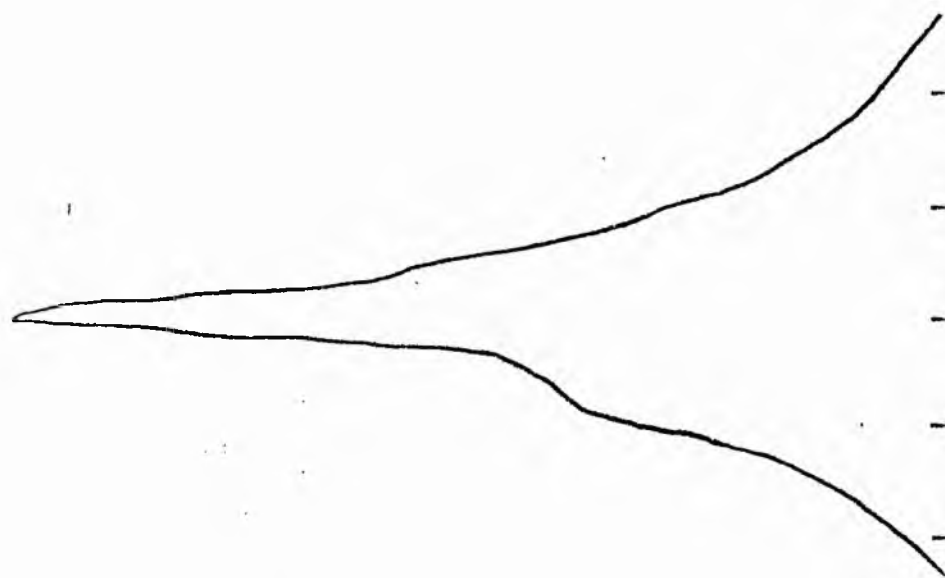
B - FILTER

Total luminosity	L_T	= 5.68
Total apparent magnitude	m_T	= 10.86
Apparent central surface brightness	μ_0	= 19.14
Major axis at threshold	$2a_m$	= 13.3
Minor axis at threshold	$2b_m$	= 4.5
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.06
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.81
Gradient of exponential component	$G(a)$	= -0.47
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.85
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.84
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 1.46
Axis ratio	b/a	= 0.25
Equivalent radius	r_1^*	= 0.56
Surface brightness	μ_1	= 21.74
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 2.25
Axis ratio	b/a	= 0.22
Equivalent radius	r_e^*	= 0.99
Surface brightness	μ_e	= 22.47
Mean surface brightness	μ_e'	= 12.83
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 4.03
Axis ratio	b/a	= 0.24
Equivalent radius	r_3^*	= 1.76
Surface brightness	μ_3	= 24.12
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.79 \\ = 1.76 \end{cases}$

NGC 4192
V-Filter
Axis 1

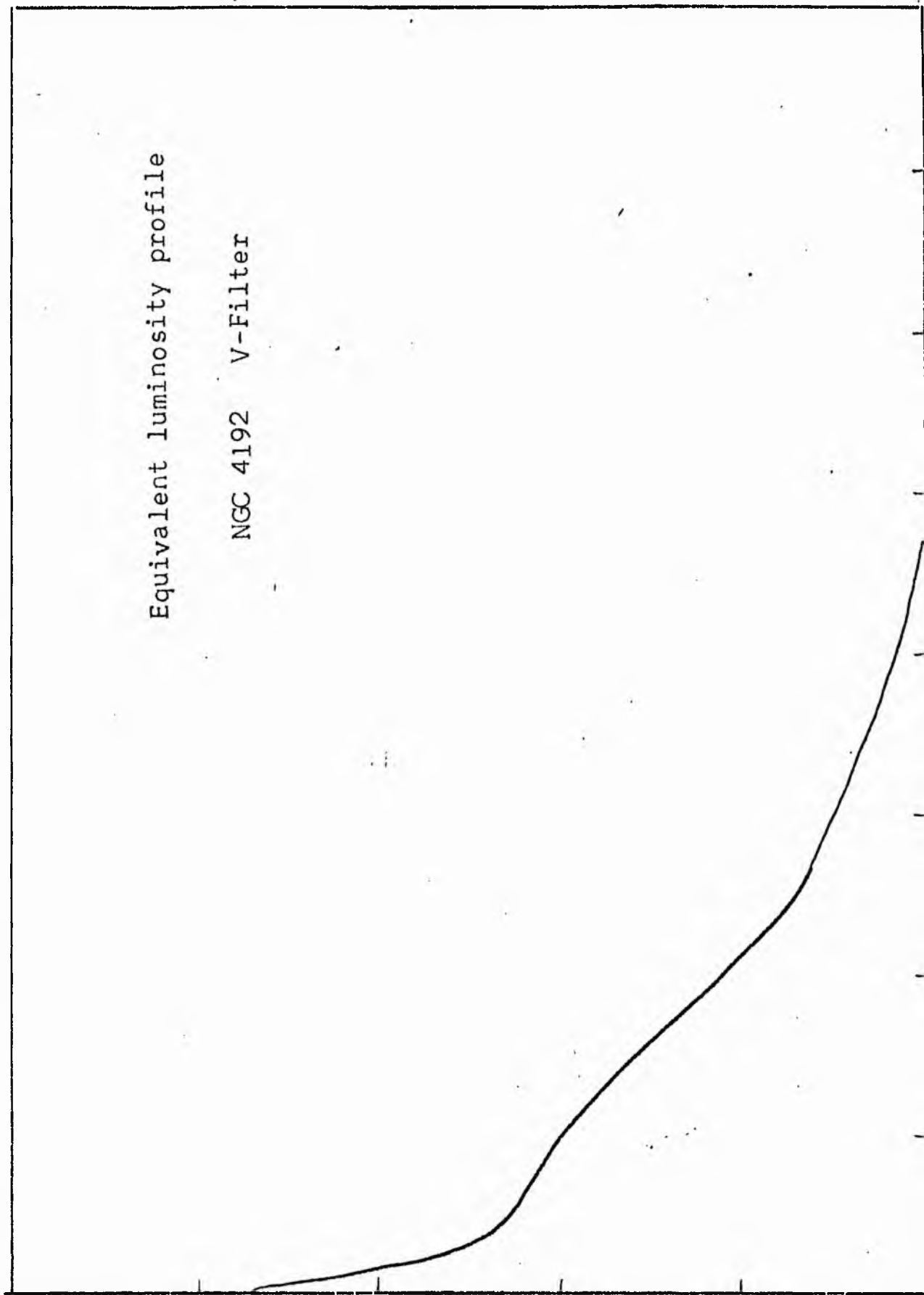


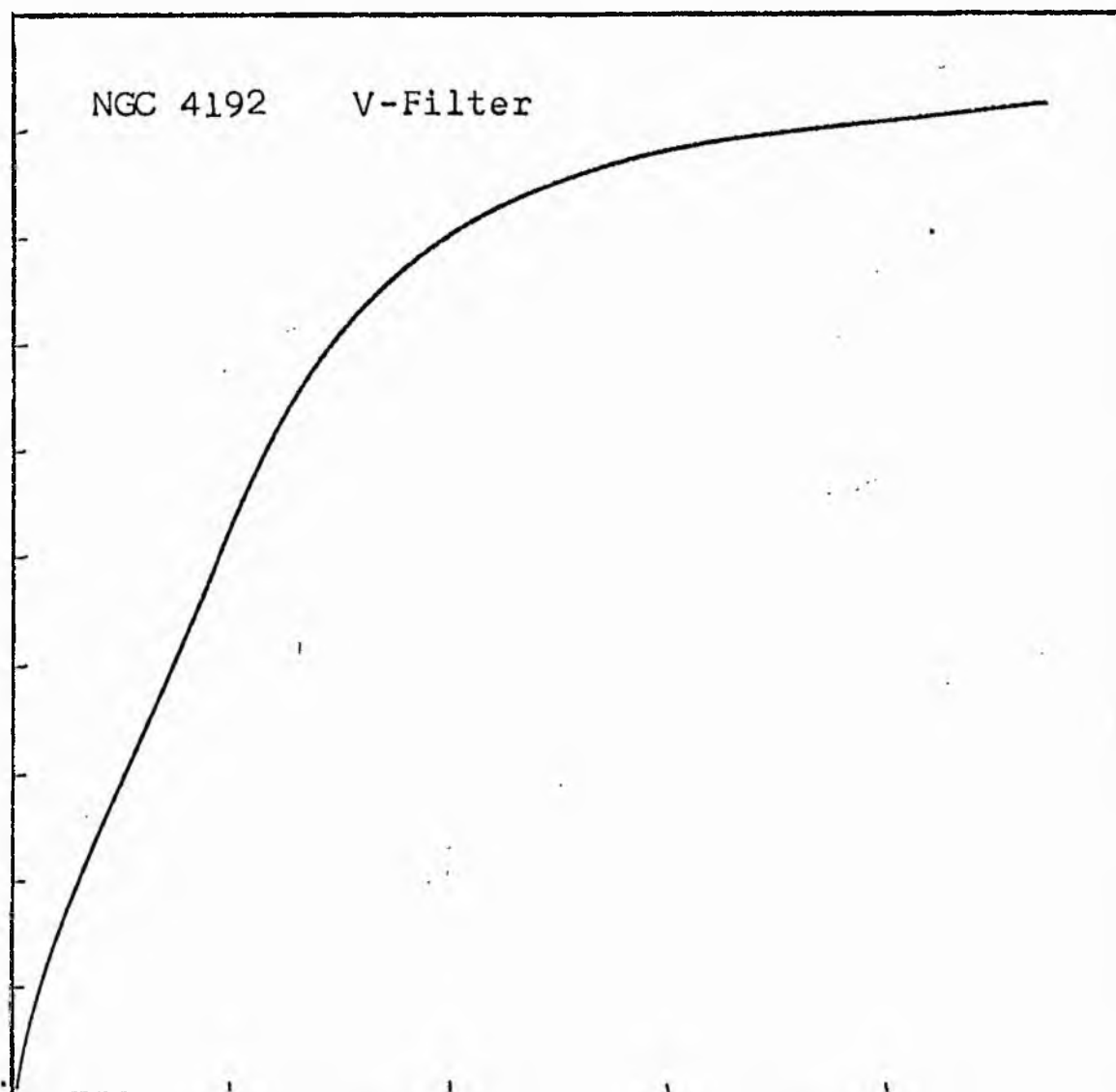
NGC 4192
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4192 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4192
V COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.68	47.863	43.837	0.0	0.0	8.04	352.5559	0.0	0.0	0.0	1.683	17.18
1.60	39.811	35.717	1.60	8.04	16.59	592.4534	352.56	0.01	0.03	1.603	17.38
1.50	31.623	28.371	2.80	24.63	16.08	456.3423	945.01	0.03	0.06	1.503	17.63
1.40	25.119	22.536	3.60	40.72	22.34	503.4004	1401.35	0.05	0.08	1.403	17.88
1.30	19.953	17.901	4.48	63.05	21.90	391.9475	1904.75	0.06	0.10	1.303	18.13
1.20	15.849	14.219	5.20	84.95	47.78	679.4363	2296.70	0.07	0.11	1.203	18.38
1.10	12.589	11.295	6.50	132.73	85.26	962.9661	2976.14	0.10	0.14	1.103	18.63
1.00	10.000	8.972	8.33	217.99	34.78	312.0645	3939.10	0.13	0.18	1.003	18.88
0.90	7.943	7.126	8.97	252.78	16.03	114.2158	4251.16	0.14	0.19	0.903	19.13
0.80	6.310	5.661	9.25	268.80	94.25	533.5073	4365.38	0.14	0.20	0.803	19.38
0.70	5.012	4.496	10.75	363.05	116.89	525.5867	4898.88	0.16	0.23	0.703	19.63
0.60	3.981	3.572	12.36	479.94	156.21	557.9294	5424.47	0.18	0.26	0.603	19.88
0.50	3.162	2.837	14.23	636.15	240.01	680.9194	5982.39	0.19	0.30	0.503	20.13
0.40	2.512	2.254	16.70	876.16	469.98	1059.1333	6663.31	0.22	0.36	0.403	20.38
0.30	1.995	1.790	20.70	1346.14	1198.46	2145.3232	7722.45	0.25	0.44	0.303	20.63
0.20	1.585	1.422	28.46	2544.60	1960.87	2788.1697	9867.77	0.32	0.61	0.203	20.88
0.10	1.259	1.129	37.87	4505.47	2369.49	2676.2339	12655.93	0.41	0.81	0.103	21.13
-0.00	1.000	0.897	46.78	6874.96	2669.87	2395.2969	15332.16	0.50	0.99	0.003	21.38
-0.10	0.794	0.713	55.12	9544.83	2632.95	1876.3430	17727.46	0.58	1.17	-0.097	21.63
-0.20	0.631	0.566	62.26	12177.78	1710.95	968.5139	19603.80	0.64	1.32	-0.197	21.88
-0.30	0.501	0.450	66.49	13888.72	2419.91	1088.1008	20572.31	0.67	1.41	-0.297	22.13
-0.40	0.398	0.357	72.05	16308.63	2594.65	926.7200	21660.41	0.70	1.53	-0.397	22.38
-0.50	0.316	0.284	77.57	18903.28	2765.21	784.5110	22587.13	0.73	1.65	-0.497	22.63
-0.60	0.251	0.225	83.05	21668.50	2937.24	661.9268	23371.64	0.76	1.77	-0.597	22.88
-0.70	0.200	0.179	88.50	24605.74	3094.34	553.9077	24033.56	0.78	1.88	-0.697	23.13
-0.80	0.158	0.142	93.90	27700.07	3252.61	462.4893	24587.47	0.80	2.00	-0.797	23.38
-0.90	0.126	0.113	99.26	30952.68	3406.84	384.7881	25049.96	0.81	2.11	-0.897	23.63
-1.00	0.100	0.090	104.58	34359.52	3563.95	319.7444	25434.74	0.83	2.22	-0.997	23.88
-1.10	0.079	0.071	109.87	37923.47	3703.61	263.9343	25754.48	0.84	2.34	-1.097	24.13
-1.20	0.063	0.057	115.11	41627.07	3845.89	217.7052	26018.42	0.84	2.45	-1.197	24.38
-1.30	0.050	0.045	120.31	45472.96	13353.94	600.4575	26236.12	0.85	2.56	-1.297	24.63
-1.40	0.040	0.036	136.84	58826.90	7416.54	264.8955	26836.58	0.87	2.91	-1.397	24.88
-1.50	0.032	0.028	145.21	66243.44	10504.62	298.0264	27101.47	0.88	3.09	-1.497	25.13
-1.60	0.025	0.023	156.30	76748.06	12237.19	275.7759	27399.50	0.89	3.32	-1.597	25.38
-1.70	0.020	0.018	168.30	88985.25	13936.44	249.4750	27675.27	0.90	3.58	-1.697	25.63
-1.80	0.016	0.014	181.00	102921.69	35226.94	500.8994	27924.74	0.91	3.85	-1.797	25.88
-1.90	0.013	0.011	209.70	138148.62	34019.44	384.2407	28425.64	0.92	4.46	-1.897	26.13
-2.00	0.010		234.10	172168.06			28809.88	0.94	4.98	-1.997	26.38
-∞							30809.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NGC 4192

V-FILTER

Total luminosity	L_T	= 8.55
Total apparent magnitude	m_T	= 10.16
Apparent central surface brightness	μ_0	= 17.18
Major axis at threshold	$2a_m$	= 13.00
Minor axis at threshold	$2b_m$	= 4.45'
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 10.7
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.88
Gradient of exponential component	$G(a)$	= -0.44
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.03
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.87

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.75
Axis ratio	b/a	= 0.51
Equivalent radius	r_1^*	= 0.34
Surface brightness	μ_1	= 20.63

Parameters at $k = \frac{1}{2}$ (effective) :

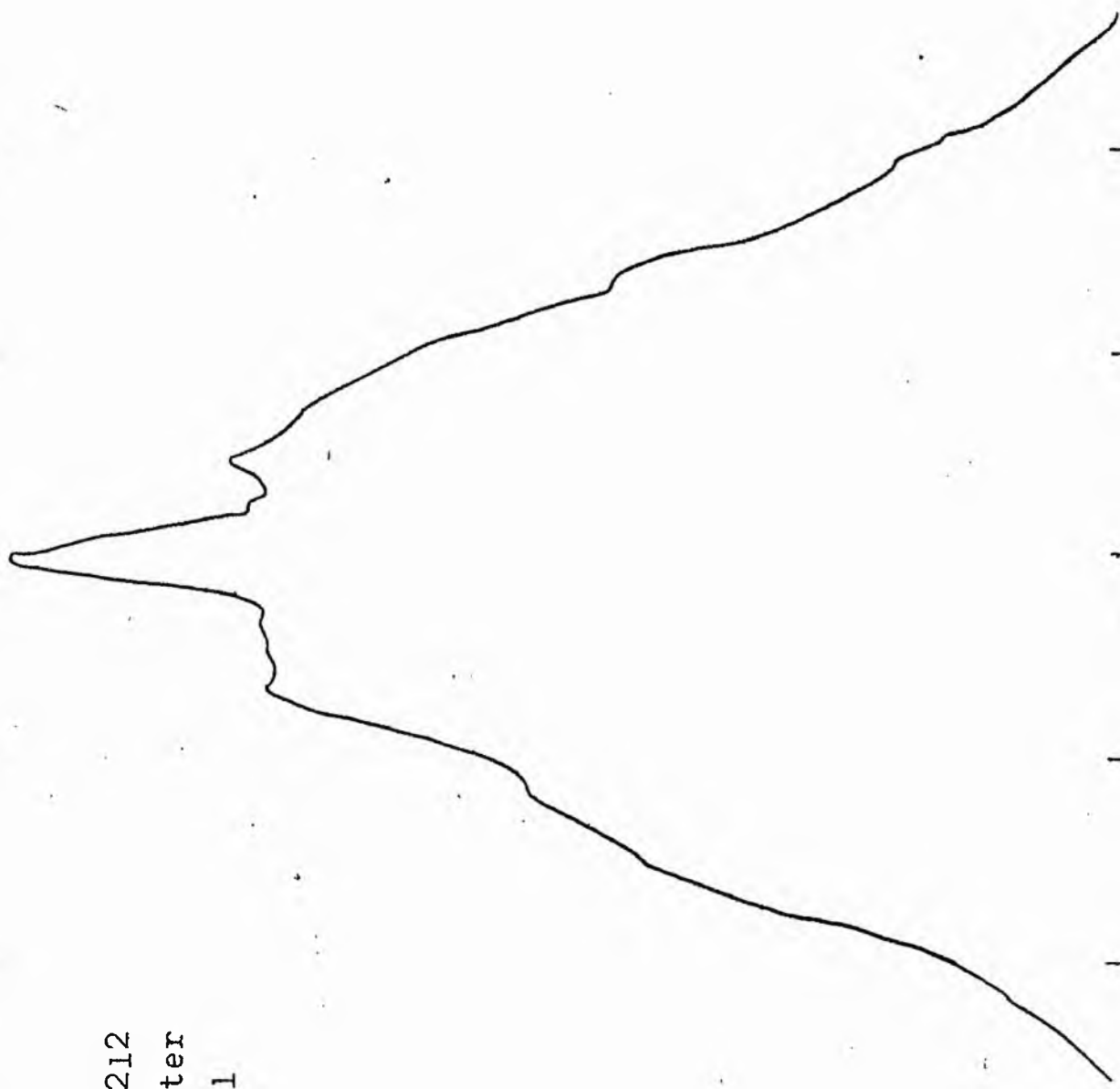
Semi-major axis	a_e	= 1.70
Axis ratio	b/a	= 0.21
Equivalent radius	r_e^*	= 0.78
Surface brightness	μ_e	= 21.38
Mean surface brightness	μ_e'	= 11.61

Parameters at $k = \frac{3}{4}$:

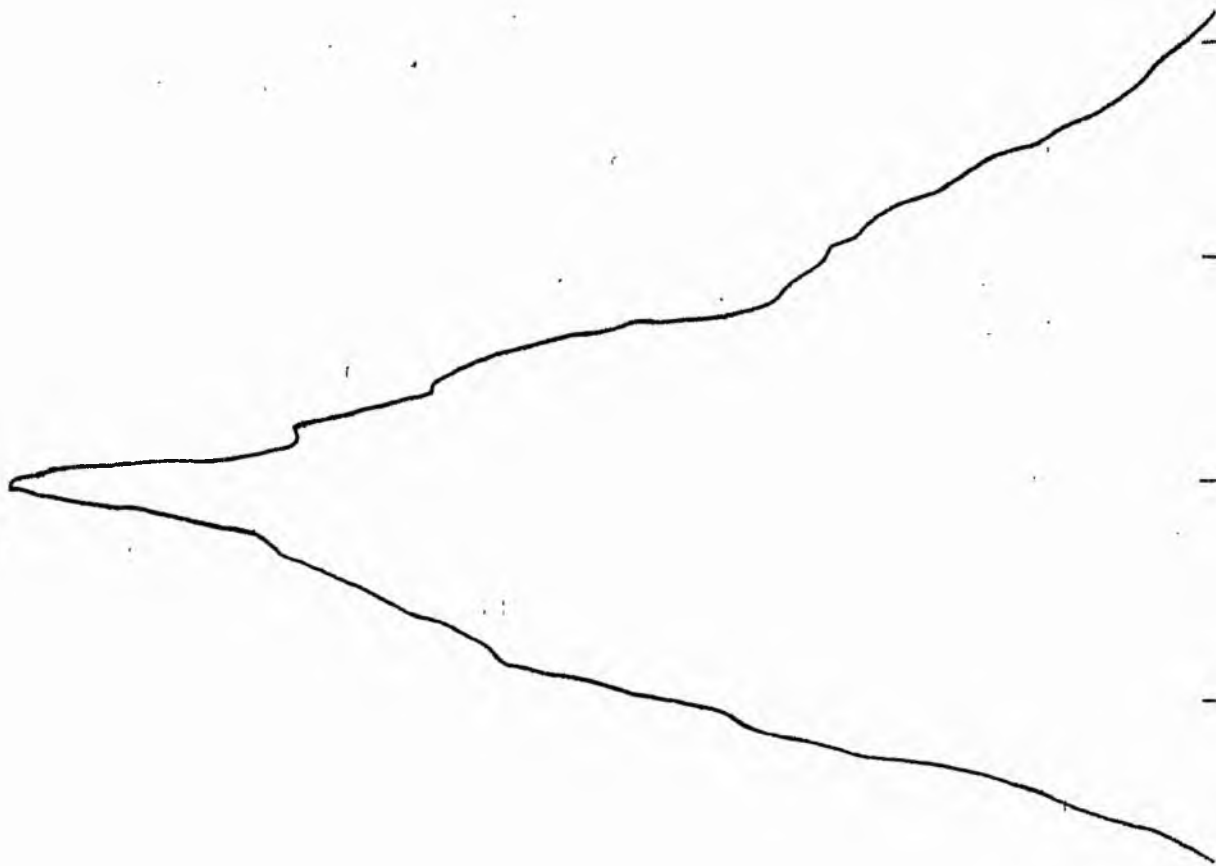
Semi-major axis	a_3	= 3.07
Axis ratio	b/a	= 0.20
Equivalent radius	r_3^*	= 1.35
Surface brightness	μ_3	= 22.78

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.28 \\ = 1.72 \end{cases}$
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NGC 4212
B-Filter
Axis 1

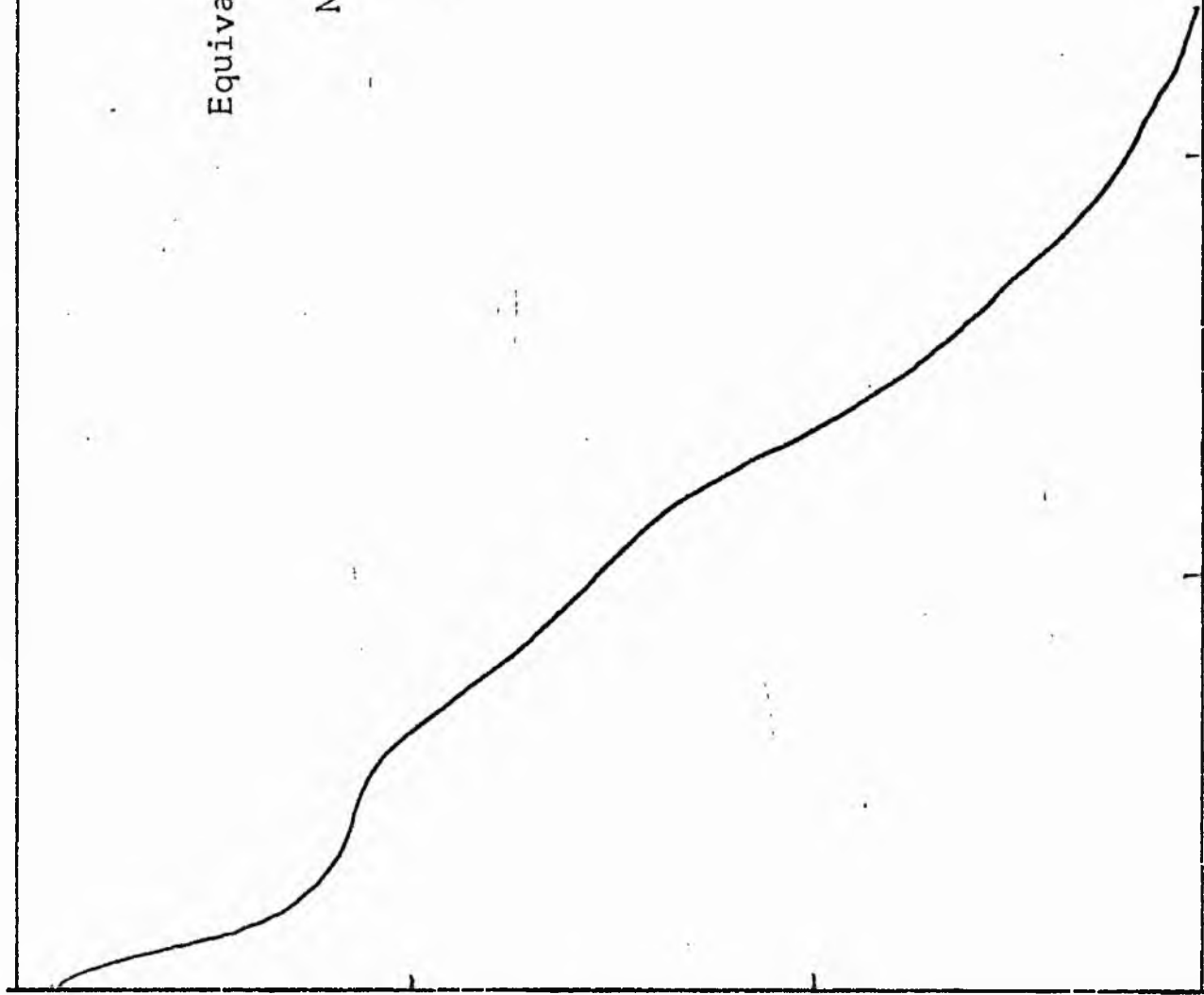


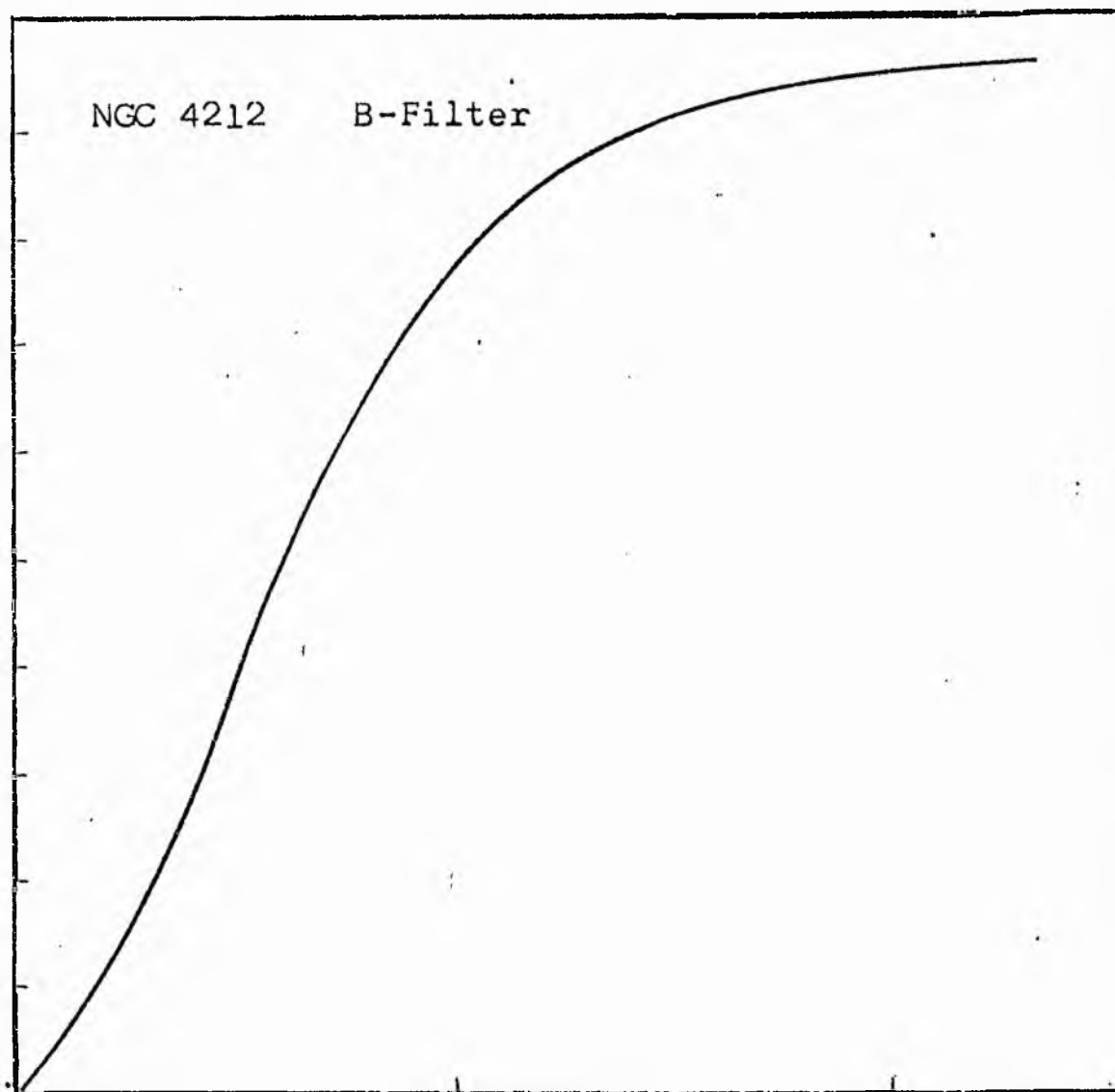
NGC 4212
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4212 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4212
B COLOUR

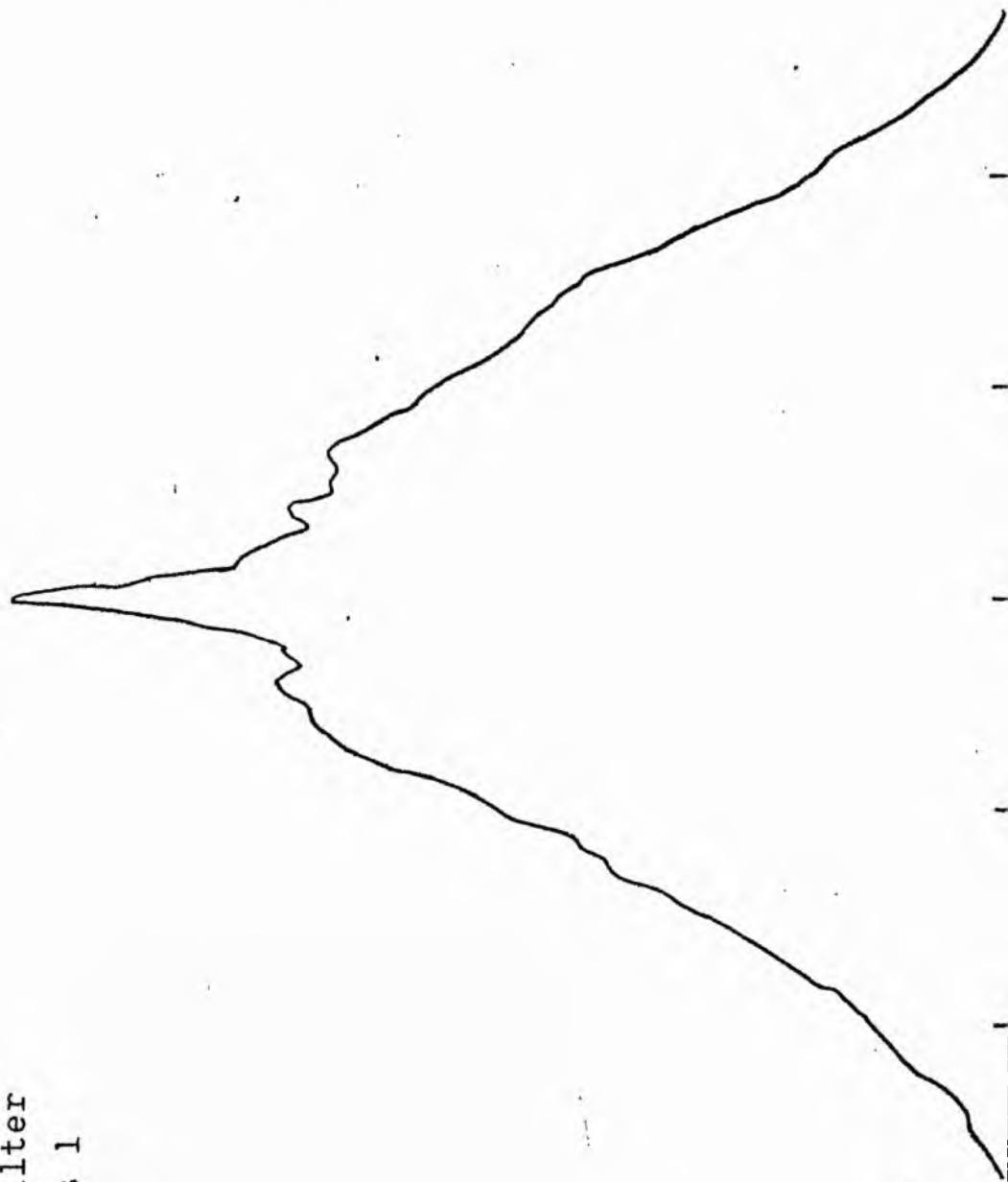
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.89	7.762		0.0	0.0			0.0	0.0	0.0	0.942	19.52
		7.036			10.29	72.4159					
0.80	6.310		1.81	10.29			72.42	0.01	0.06	0.852	19.75
		5.661			28.19	159.5889					
0.70	5.012		3.50	38.48			232.00	0.03	0.11	0.752	20.00
		4.496			43.23	194.3745					
0.60	3.981		5.10	81.71			426.38	0.05	0.17	0.652	20.25
		3.572			76.65	273.7861					
0.50	3.162		7.10	158.37			700.17	0.08	0.23	0.552	20.50
		2.837			42.69	121.1272					
0.40	2.512		8.00	201.06			821.29	0.09	0.26	0.452	20.75
		2.254			106.85	240.7838					
0.30	1.995		9.90	307.91			1062.08	0.11	0.32	0.352	21.00
		1.790			417.93	748.1206					
0.20	1.585		15.20	725.83			1810.20	0.20	0.50	0.252	21.25
		1.422			1301.00	1849.8987					
0.10	1.259		25.40	2026.83			3660.10	0.40	0.83	0.152	21.50
		1.129			560.87	633.4807					
0.00	1.000		28.70	2587.70			4293.57	0.46	0.94	0.052	21.75
		0.897			689.89	618.9463					
-0.10	0.794		32.30	3277.59			4912.52	0.53	1.06	-0.048	22.00
		0.713			726.34	517.6201					
-0.20	0.631		35.70	4003.93			5430.14	0.59	1.17	-0.148	22.25
		0.566			964.98	546.2476					
-0.30	0.501		39.77	4968.91			5976.38	0.65	1.30	-0.248	22.50
		0.450			1353.29	608.5029					
-0.40	0.398		44.86	6322.20			6584.88	0.71	1.47	-0.348	22.75
		0.357			1285.56	459.1602					
-0.50	0.316		49.21	7607.75			7044.04	0.76	1.61	-0.448	23.00
		0.284			1267.00	359.4587					
-0.60	0.251		53.15	8874.75			7403.50	0.80	1.74	-0.548	23.25
		0.225			1225.12	276.0896					
-0.70	0.200		56.70	10099.87			7679.59	0.83	1.85	-0.648	23.50
		0.179			1149.62	205.7914					
-0.80	0.158		59.84	11249.49			7885.37	0.85	1.96	-0.748	23.75
		0.142			1057.72	150.3982					
-0.90	0.126		62.59	12307.21			8035.77	0.87	2.05	-0.848	24.00
		0.113			937.44	105.8802					
-1.00	0.100		64.93	13244.65			8141.65	0.88	2.12	-0.948	24.25
		0.090			1039.55	93.2646					
-1.10	0.079		67.43	14284.20			8234.91	0.89	2.20	-1.048	24.50
		0.071			1197.70	85.3532					
-1.20	0.063		70.20	15481.89			8320.26	0.90	2.29	-1.148	24.75
		0.057			1954.73	110.6527					
-1.30	0.050		74.50	17436.62			8430.91	0.91	2.43	-1.248	25.00
		0.045			2319.23	104.2843					
-1.40	0.040		79.30	19755.86			8535.20	0.92	2.59	-1.348	25.25
		0.036			2200.60	78.5990					
-1.50	0.032		83.60	21956.46			8613.79	0.93	2.73	-1.448	25.50
		0.028			2593.70	73.5862					
-1.60	0.025		88.40	24550.16			8687.38	0.94	2.89	-1.548	25.75
		0.023			2679.93	60.3948					
-1.70	0.020		93.10	27230.09			8747.77	0.94	3.04	-1.648	26.00
		0.018			4500.79	80.5685					
-1.80	0.016		100.50	31730.87			8828.34	0.95	3.28	-1.748	26.25
		0.014			4506.64	64.0812					
-1.90	0.013		107.40	36237.51			8892.42	0.96	3.51	-1.848	26.50
		0.011			6254.68	70.6453					
-2.00	0.010		116.30	42492.19			8963.06	0.97	3.80	-1.948	26.75
-∞							9263.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4212

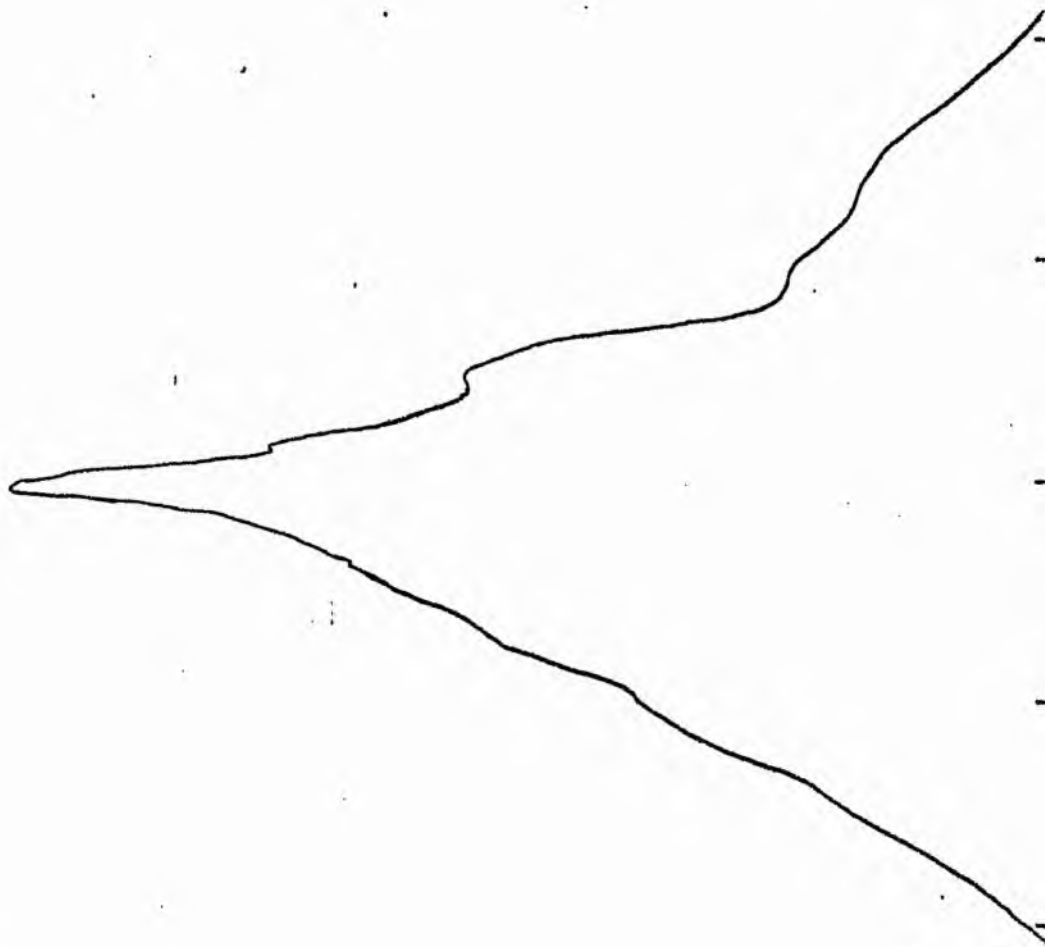
B-FILTER

Total luminosity	L_T	= 2.57
Total apparent magnitude	m_T	= 11.83
Apparent central surface brightness	μ_0	= 19.52
Major axis at threshold	$2a_m$	= 4.41
Minor axis at threshold	$2b_m$	= 3.03
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 3.08
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.91
Gradient of exponential component	$G(a)$	= -1.66
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.53
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.83
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.53
Axis ratio	b/a	= 0.62
Equivalent radius	r_1^*	= 0.30
Surface brightness	μ_1	= 21.31
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.74
Axis ratio	b/a	= 0.53
Equivalent radius	r_e^*	= 0.51
Surface brightness	μ_e	= 21.89
Mean surface brightness	μ_e'	= 12.36
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.05
Axis ratio	b/a	= 0.60
Equivalent radius	r_3^*	= 0.80
Surface brightness	μ_3	= 22.93
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.66 \\ = 1.58 \end{cases}$

NGC 4212
V-Filter
Axis 1

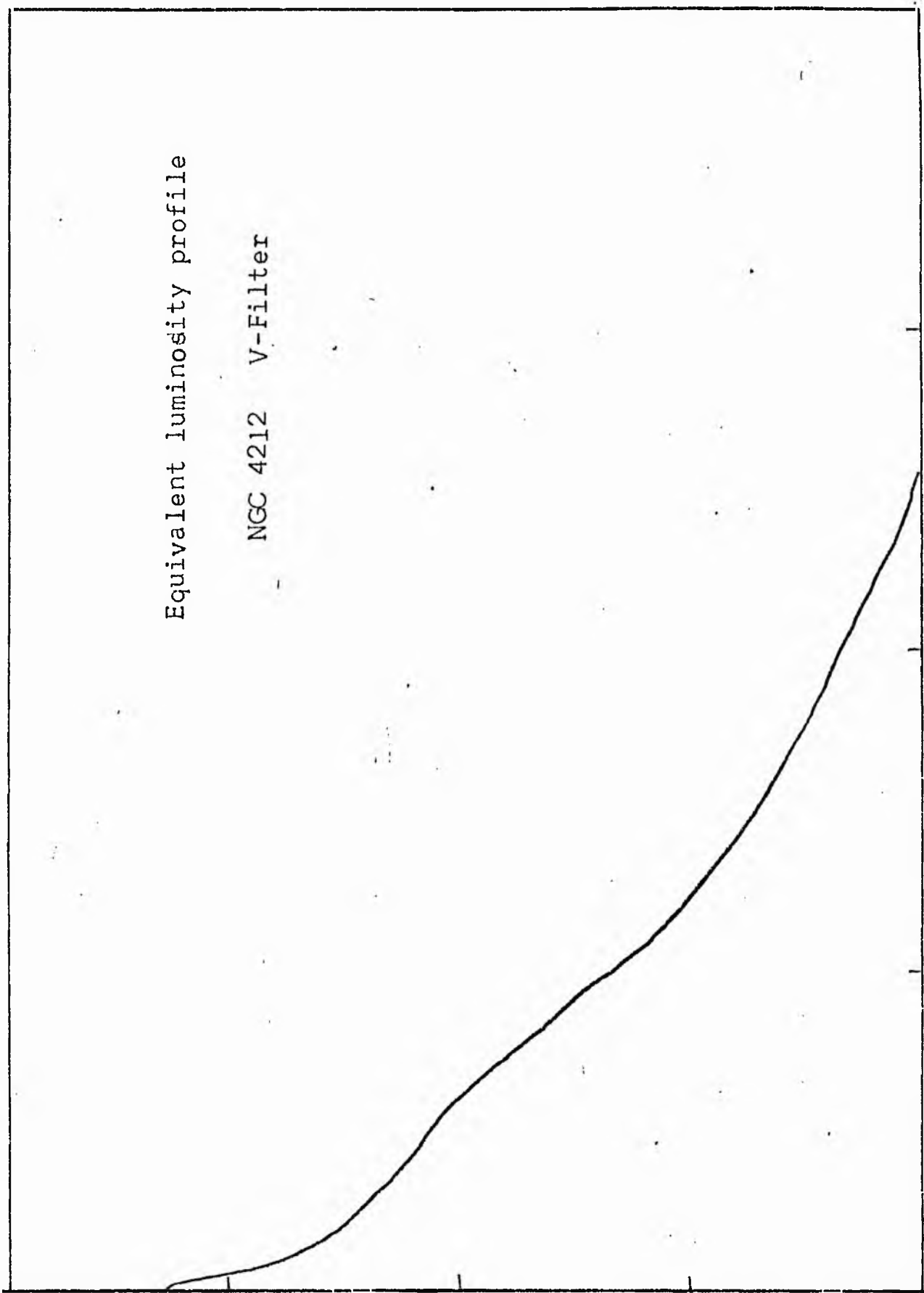


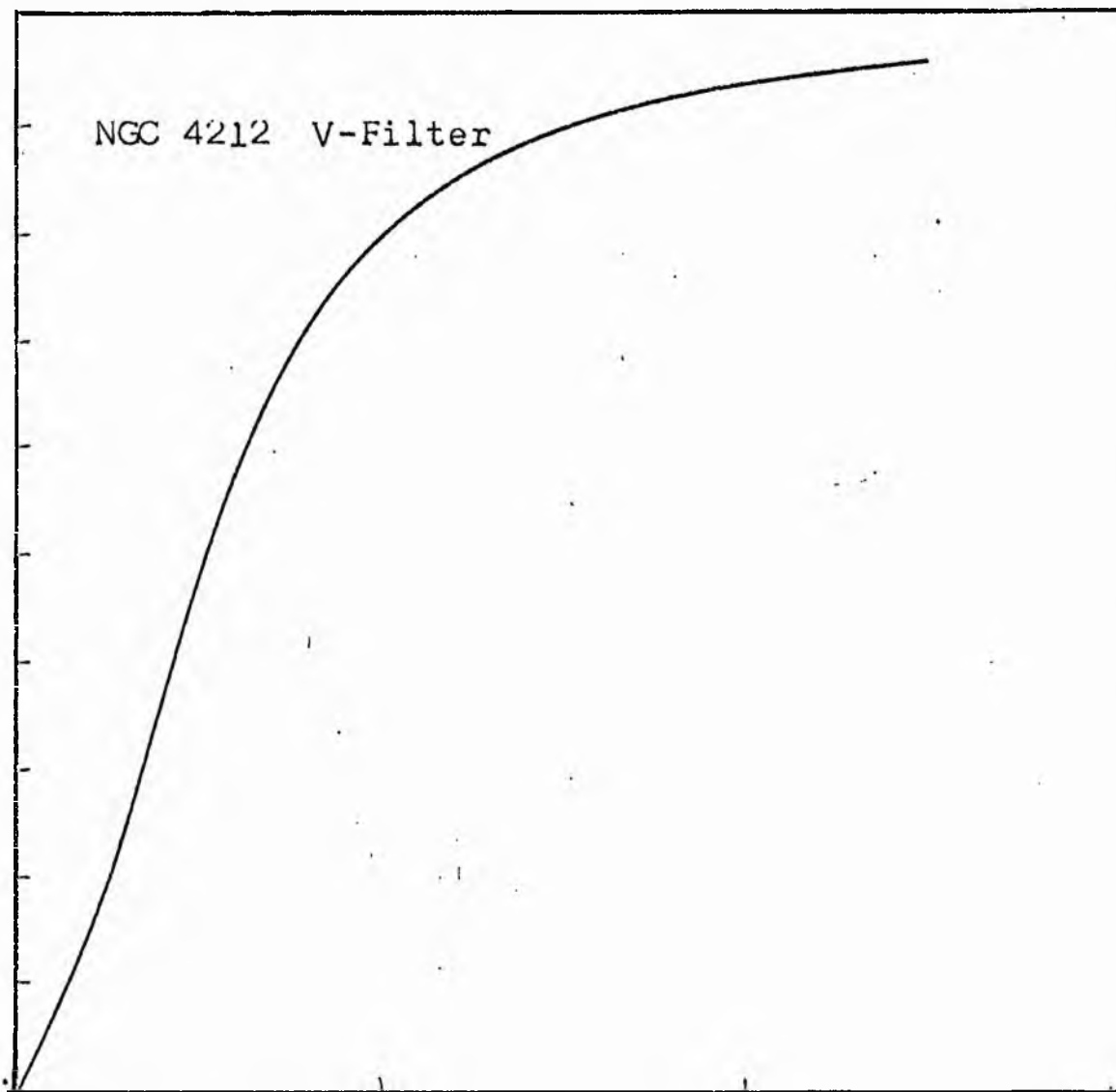
NGC 4212
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4212 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

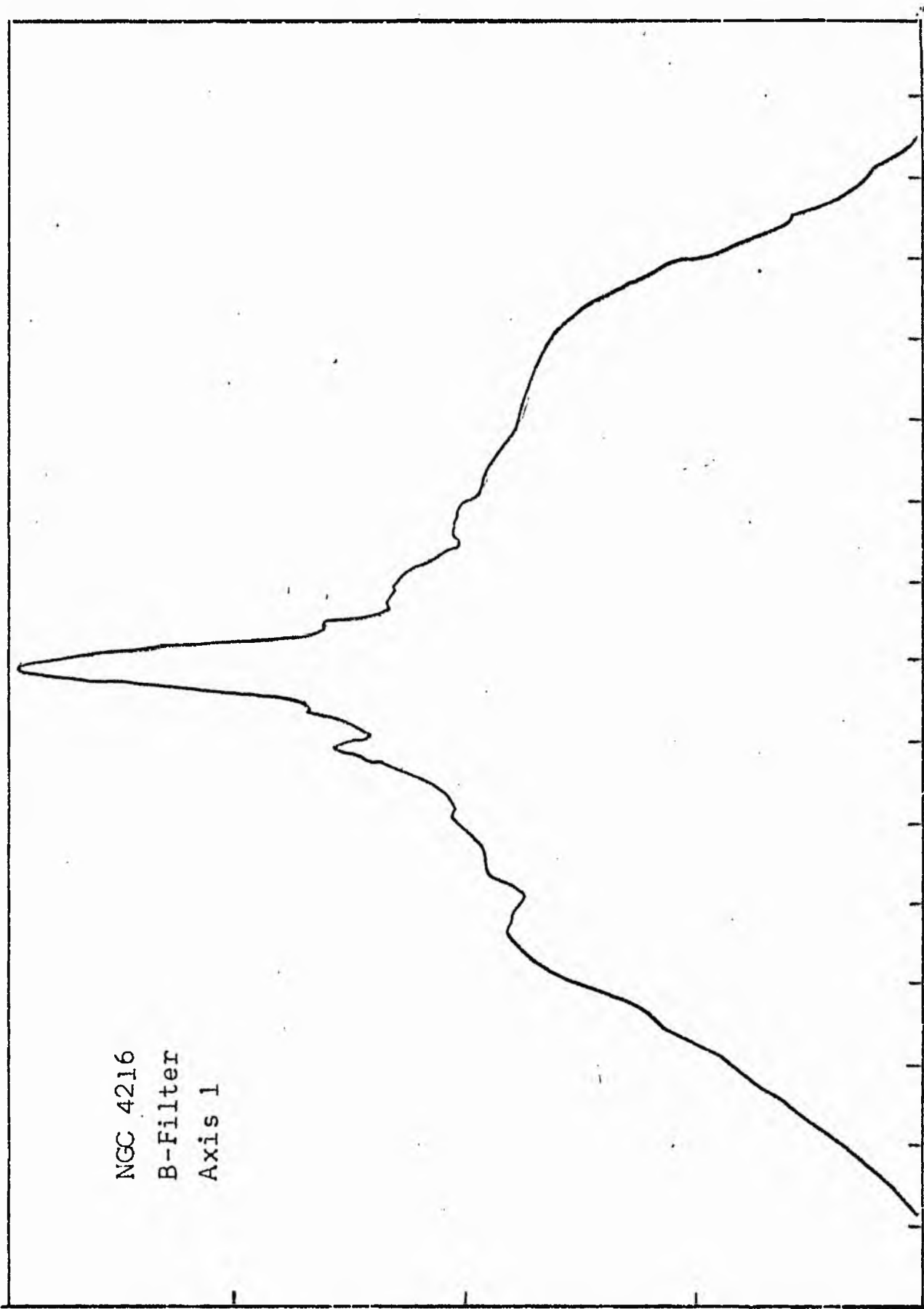
MEAN LUMINOSITY DISTRIBUTION IN NGC 4212
V COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.27	18.621		0.0	0.0			0.0	0.0	0.0	1.117	18.33
		17.235			10.87	187.3199					
1.20	15.849		1.86	10.87			187.32	0.02	0.07	1.047	18.21
		14.219			2.46	35.0216					
1.10	12.589		2.06	13.33			222.34	0.02	0.08	0.947	18.46
		11.295			19.04	215.0446					
1.00	10.000		3.21	32.37			437.39	0.04	0.12	0.847	18.71
		8.972			25.18	225.8845					
0.90	7.943		4.28	57.55			663.27	0.06	0.16	0.747	18.96
		7.126			16.35	116.5110					
0.80	6.310		4.85	73.90			779.78	0.07	0.18	0.647	19.21
		5.661			54.78	310.1025					
0.70	5.012		6.40	128.68			1089.88	0.10	0.24	0.547	19.46
		4.496			35.09	157.7655					
0.60	3.981		7.22	163.77			1247.65	0.12	0.27	0.447	19.71
		3.572			137.34	490.5217					
0.50	3.162		9.79	301.10			1738.17	0.16	0.37	0.347	19.96
		2.837			189.77	538.3943					
0.40	2.512		12.50	490.87			2276.57	0.21	0.47	0.247	20.21
		2.254			385.28	868.2668					
0.30	1.995		16.70	876.16			3144.83	0.29	0.63	0.147	20.46
		1.790			791.53	1416.8955					
0.20	1.585		23.04	1667.69			4561.73	0.43	0.87	0.047	20.71
		1.422			755.02	1073.5723					
0.10	1.259		27.77	2422.71			5635.30	0.53	1.05	-0.053	20.96
		1.129			290.91	328.5728					
-0.00	1.000		29.39	2713.62			5963.87	0.56	1.11	-0.153	21.21
		0.897			621.04	557.1748					
-0.10	0.794		32.58	3334.66			6521.04	0.61	1.24	-0.253	21.46
		0.713			864.50	616.0762					
-0.20	0.631		36.56	4199.16			7137.11	0.67	1.39	-0.353	21.71
		0.566			802.29	454.1511					
-0.30	0.501		39.90	5001.44			7591.26	0.71	1.51	-0.453	21.96
		0.450			804.66	361.8113					
-0.40	0.398		42.99	5806.10			7953.07	0.75	1.63	-0.553	22.21
		0.357			726.40	259.4473					
-0.50	0.316		45.60	6532.50			8212.52	0.77	1.73	-0.653	22.46
		0.284			805.59	228.5536					
-0.60	0.251		48.33	7338.09			8441.07	0.79	1.83	-0.753	22.71
		0.225			801.17	180.5501					
-0.70	0.200		50.90	8139.27			8621.61	0.81	1.93	-0.853	22.96
		0.179			886.40	158.6723					
-0.80	0.158		53.60	9025.66			8780.29	0.82	2.03	-0.953	23.21
		0.142			721.09	102.5330					
-0.90	0.126		55.70	9746.76			8882.82	0.83	2.11	-1.053	23.46
		0.113			1487.70	168.0297					
-1.00	0.100		59.80	11234.45			9050.84	0.85	2.27	-1.153	23.71
		0.090			1714.05	153.7786					
-1.10	0.079		64.20	12948.50			9204.62	0.86	2.43	-1.253	23.96
		0.071			3473.50	247.5368					
-1.20	0.063		72.30	16422.00			9452.16	0.89	2.74	-1.353	24.21
		0.057			1915.34	108.4224					
-1.30	0.050		76.40	18337.34			9560.58	0.90	2.90	-1.453	24.46
		0.045			3357.26	150.9590					
-1.40	0.040		83.10	21694.60			9711.54	0.91	3.15	-1.553	24.71
		0.036			3752.30	134.0207					
-1.50	0.032		90.00	25446.90			9845.55	0.92	3.41	-1.653	24.96
		0.028			3265.23	92.6380					
-1.60	0.025		95.60	28712.13			9938.19	0.93	3.63	-1.753	25.21
		0.023			5071.57	114.2927					
-1.70	0.020		103.70	33783.71			10052.48	0.94	3.93	-1.853	25.46
		0.018			3131.76	56.0615					
-1.80	0.016		108.40	36915.47			10108.54	0.95	4.11	-1.953	25.71
		0.014			2986.81	42.4702					
-1.90	0.013		112.70	39902.28			10151.01	0.95	4.27	-2.053	25.96
		0.011			9185.11	103.7437					
-2.00	0.010		125.00	49087.39			10254.75	0.96	4.74	-2.153	26.21
-∞							10674.00	(1)			∞

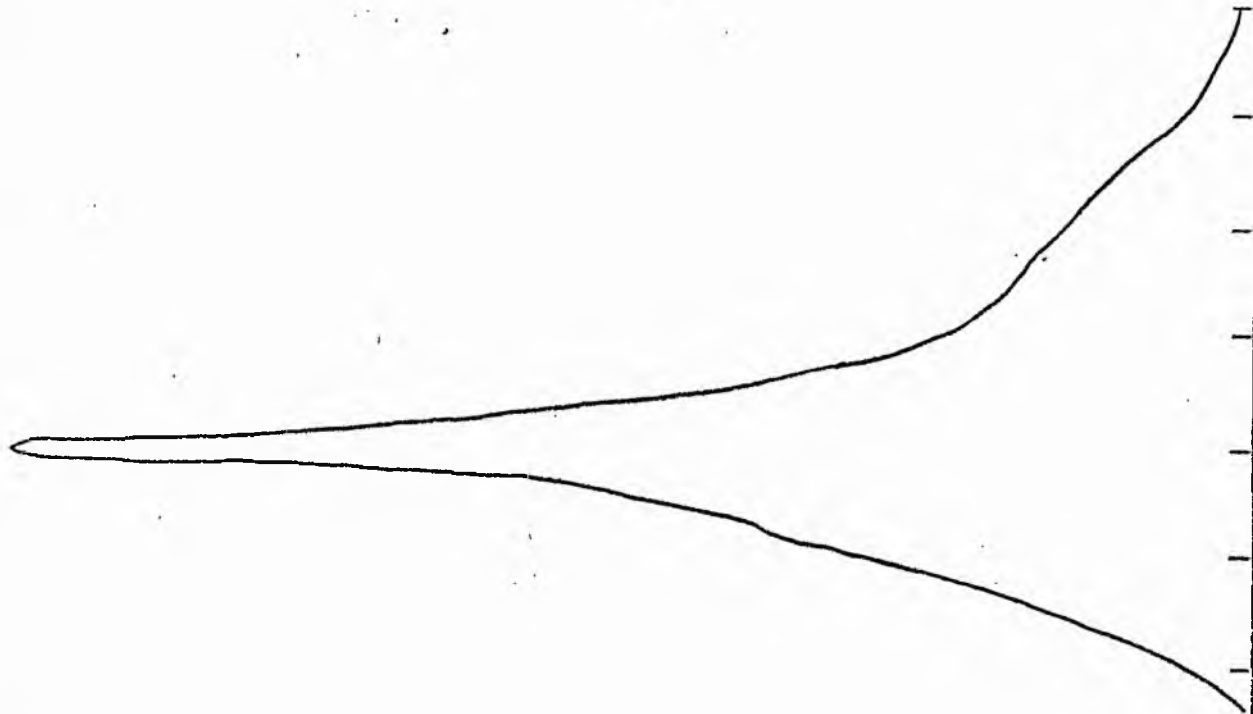
PHOTOMETRIC PARAMETERS OF NGC 4212

V-FILTER

Total luminosity	L_T	= 2.96
Total apparent magnitude	m_T	= 11.14
Apparent central surface brightness	μ_o	= 18.03
Major axis at threshold	$2a_m$	= 4.61
Minor axis at threshold	$2b_m$	= 3.65
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.56
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.92
Gradient of exponential component	$G(a)$	= -1.36
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.66
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.87
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.36
Axis ratio	b/a	= 0.54
Equivalent radius	r_1^*	= 0.24
Surface brightness	μ_1	= 20.33
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.63
Axis ratio	b/a	= 0.51
Equivalent radius	r_e^*	= 0.44
Surface brightness	μ_e	= 20.88
Mean surface brightness	μ_e'	= 11.33
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.97
Axis ratio	b/a	= 0.60
Equivalent radius	r_3^*	= 0.72
Surface brightness	μ_3	= 22.21
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.83 \\ = 1.65 \end{cases}$

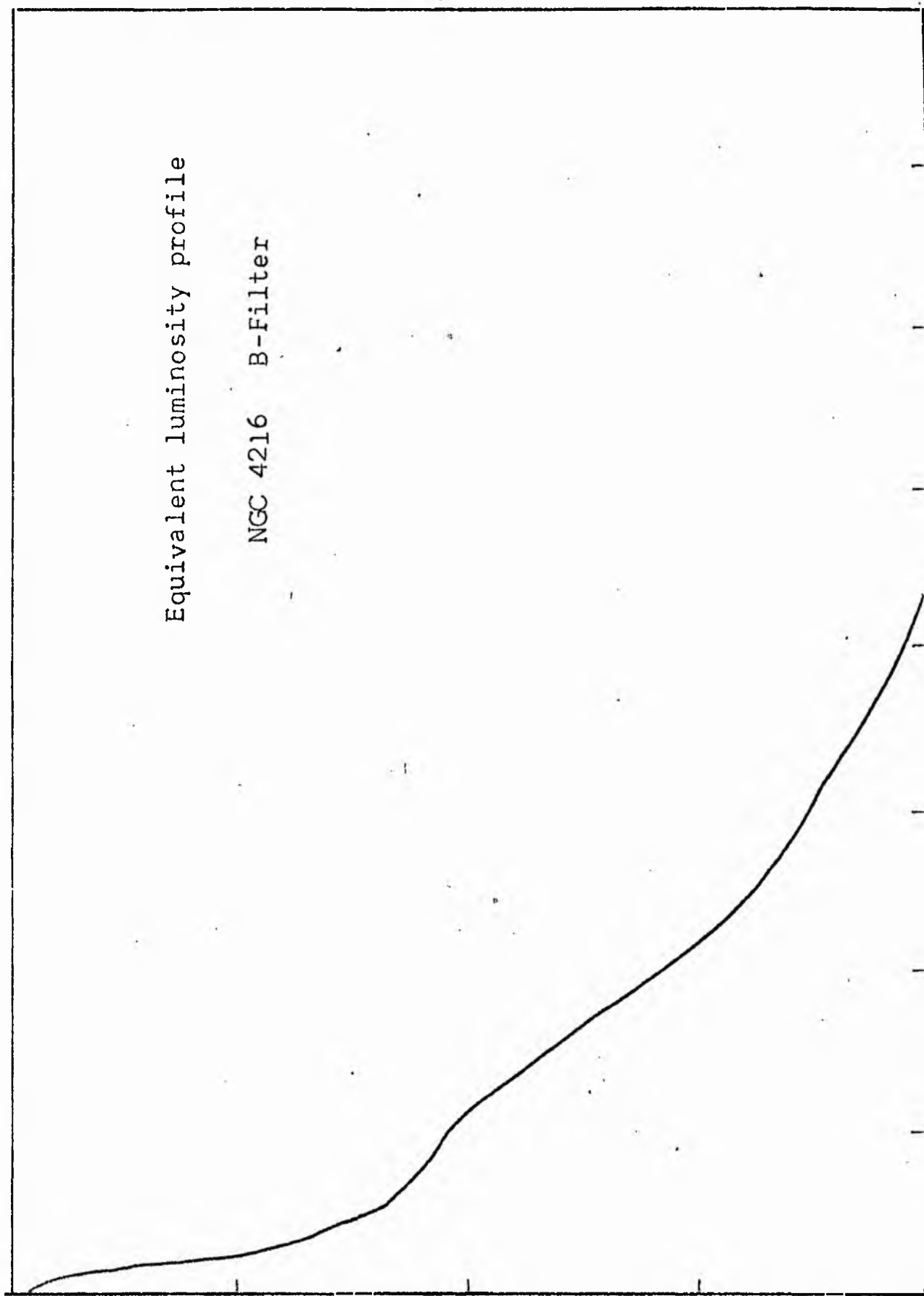


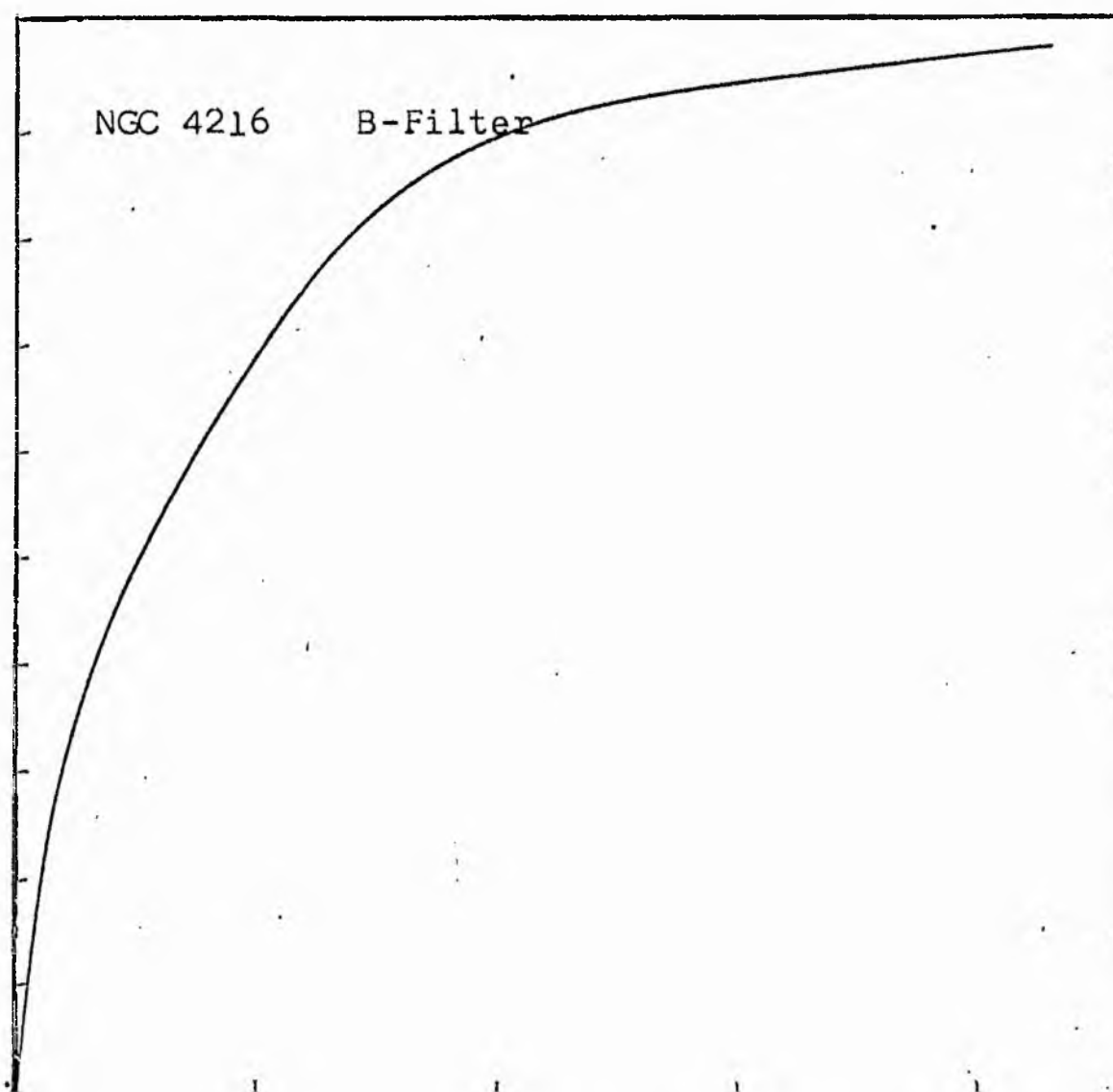
NGC 4216
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4216 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4216
B COLOUR

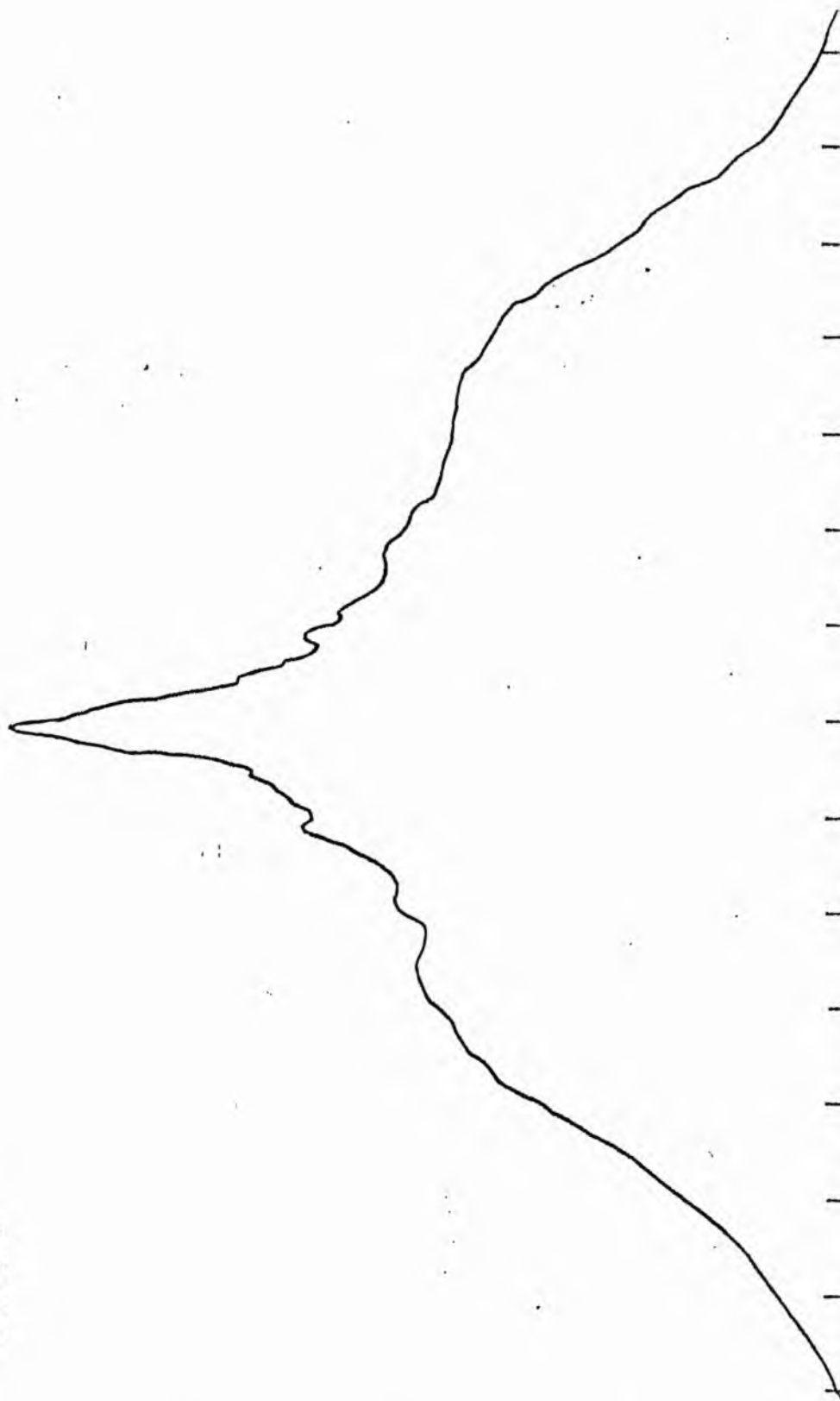
LOG I	I	\bar{I}	R	AREA	ΔA	P.	ΣP	K(R)	ρ	LOG J	μ
1.95	89.125		0.0	0.0			0.0	0.0	0.0	1.669	17.69
		84.279			38.48	3243.4304					
1.90	79.433		3.50	38.48			3243.43	0.07	0.12	1.619	17.62
		71.264			14.33	1020.9026					
1.80	63.096		4.10	52.81			4264.33	0.09	0.15	1.519	18.07
		56.607			28.90	1636.0937					
1.70	50.119		5.10	81.71			5900.43	0.12	0.18	1.419	18.32
		44.965			55.13	2479.1204					
1.60	39.811		6.60	136.85			8379.54	0.18	0.23	1.319	18.57
		35.717			79.58	2842.2036					
1.50	31.623		8.30	216.42			11221.75	0.24	0.30	1.219	18.82
		28.371			15.93	451.8033					
1.40	25.119		8.60	232.35			11673.63	0.25	0.31	1.119	19.07
		22.536			45.24	1019.4063					
1.30	19.952		9.40	277.59			12693.11	0.27	0.33	1.019	19.32
		17.901			62.20	1113.4807					
1.20	15.849		10.40	339.79			13806.59	0.29	0.37	0.919	19.57
		14.219			68.49	973.8125					
1.10	12.589		11.40	408.28			14780.41	0.31	0.41	0.819	19.82
		11.295			90.48	1021.9036					
1.00	10.000		12.60	498.76			15802.31	0.33	0.45	0.719	20.07
		8.972			65.35	586.2488					
0.90	7.943		13.40	564.10			16388.55	0.35	0.48	0.619	20.32
		7.126			114.76	817.8394					
0.80	6.310		14.70	678.87			17206.39	0.36	0.52	0.519	20.57
		5.661			176.43	998.7236					
0.70	5.012		16.50	855.30			18205.11	0.39	0.59	0.419	20.82
		4.496			439.32	1975.3665					
0.60	3.981		20.30	1294.62			20180.48	0.43	0.72	0.319	21.07
		3.572			455.12	1625.5378					
0.50	3.162		23.60	1749.74			21806.01	0.46	0.84	0.219	21.32
		2.837			325.25	922.7539					
0.40	2.512		25.70	2074.99			22728.77	0.48	0.91	0.119	21.57
		2.254			335.52	756.1191					
0.30	1.995		27.70	2410.51			23484.88	0.50	0.99	0.019	21.82
		1.790			1843.95	3300.7947					
0.20	1.585		36.80	4254.46			26785.68	0.57	1.31	-0.081	22.07
		1.422			2252.27	3202.5049					
0.10	1.259		45.51	6506.74			29988.18	0.63	1.62	-0.181	22.32
		1.129			2378.04	2685.8026					
-0.00	1.000		53.18	8884.77			32674.06	0.69	1.89	-0.281	22.57
		0.897			2342.17	2101.2974					
-0.10	0.794		59.78	11226.95			34775.36	0.74	2.13	-0.381	22.82
		0.713			1044.90	744.6340					
-0.20	0.631		62.50	12271.84			35519.99	0.75	2.22	-0.481	23.07
		0.566			2902.83	1643.2607					
-0.30	0.501		69.50	15174.68			37163.19	0.79	2.47	-0.581	23.32
		0.450			3548.70	1595.6516					
-0.40	0.398		77.20	18723.37			38758.84	0.82	2.75	-0.681	23.57
		0.357			1990.52	710.9441					
-0.50	0.316		81.20	20713.09			39469.78	0.83	2.89	-0.781	23.82
		0.284			1770.98	502.4373					
-0.60	0.251		84.60	22404.87			39972.21	0.85	3.01	-0.881	24.07
		0.225			3131.95	705.8027					
-0.70	0.200		90.30	25616.82			40678.01	0.86	3.21	-0.981	24.32
		0.179			2497.79	447.1208					
-0.80	0.158		94.60	28114.61			41125.13	0.87	3.37	-1.081	24.57
		0.142			4378.53	622.5835					
-0.90	0.126		101.70	32493.14			41747.71	0.88	3.62	-1.181	24.82
		0.113			4626.94	522.5930					
-1.00	0.100		108.70	37120.08			42270.30	0.89	3.87	-1.281	25.07
		0.090			4138.95	371.3306					
-1.10	0.079		114.60	41259.03			42641.63	0.90	4.08	-1.381	25.32
		0.071			4281.99	305.1519					
-1.20	0.063		120.40	45541.02			42946.78	0.91	4.28	-1.481	25.57
		0.057			9613.57	544.1968					
-1.30	0.050		132.50	55154.59			43490.97	0.92	4.71	-1.581	25.82
		0.045			8818.57	396.5247					
-1.40	0.040		142.70	63973.16			43887.50	0.93	5.08	-1.681	26.07
		0.036			13266.77	473.8467					
-1.50	0.032		156.80	77239.94			44361.34	0.94	5.58	-1.781	26.32
		0.028			13979.81	396.6196					
-1.60	0.025		170.40	91219.75			44757.96	0.95	6.06	-1.881	26.57
		0.023			14680.00	330.8259					
-1.70	0.020		183.60	105899.75			45088.78	0.95	6.53	-1.981	26.82
		0.018			16517.87	295.6843					
-1.80	0.016		197.40	122417.62			45384.46	0.96	7.02	-2.081	27.07
		0.014			19311.37	274.5920					
-1.90	0.013		212.40	141729.00			45659.05	0.97	7.56	-2.181	27.32
		0.011			5932.87	67.0102					
-2.00	0.010		216.80	147661.87			45726.06	0.97	7.71	-2.281	27.57
-∞							47276.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4216

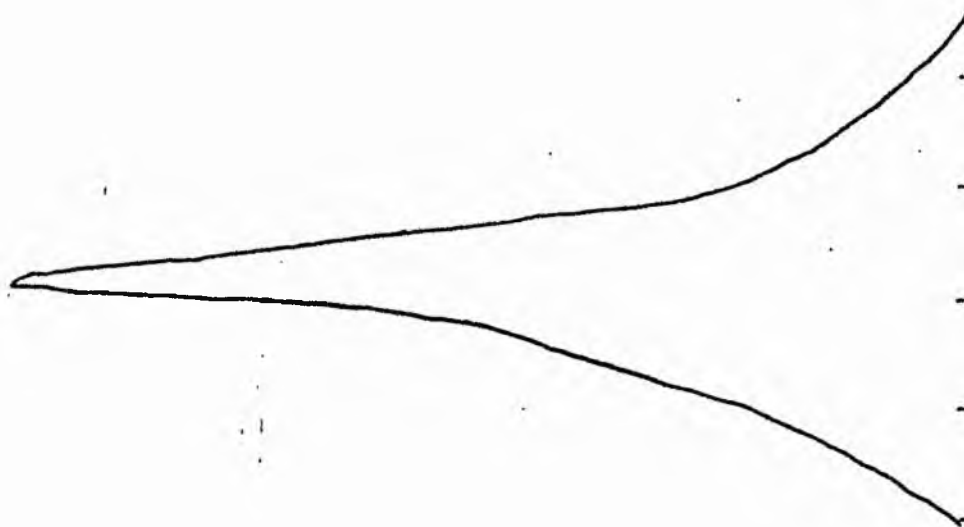
B-FILTER

Total luminosity	L_T	= 13.13
Total apparent magnitude	m_T	= 10.88
Apparent central surface brightness	μ_0	= 17.69
Major axis at threshold	$2a_m$	= 11.03
Minor axis at threshold	$2b_m$	= 5.02
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 7.95
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.68
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.03
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.51
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 0.70
Equivalent radius	r_1^*	= 0.14
Surface brightness	μ_1	= 19.07
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.92
Axis ratio	b/a	= 0.33
Equivalent radius	r_e^*	= 0.47
Surface brightness	μ_e	= 21.82
Mean surface brightness	μ_e'	= 11.22
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.2
Axis ratio	b/a	= 0.24
Equivalent radius	r_3^*	= 1.03
Surface brightness	μ_3	= 23.07
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.23 \\ = 2.21 \end{cases}$

NGC 4216
V-Filter
Axis 1

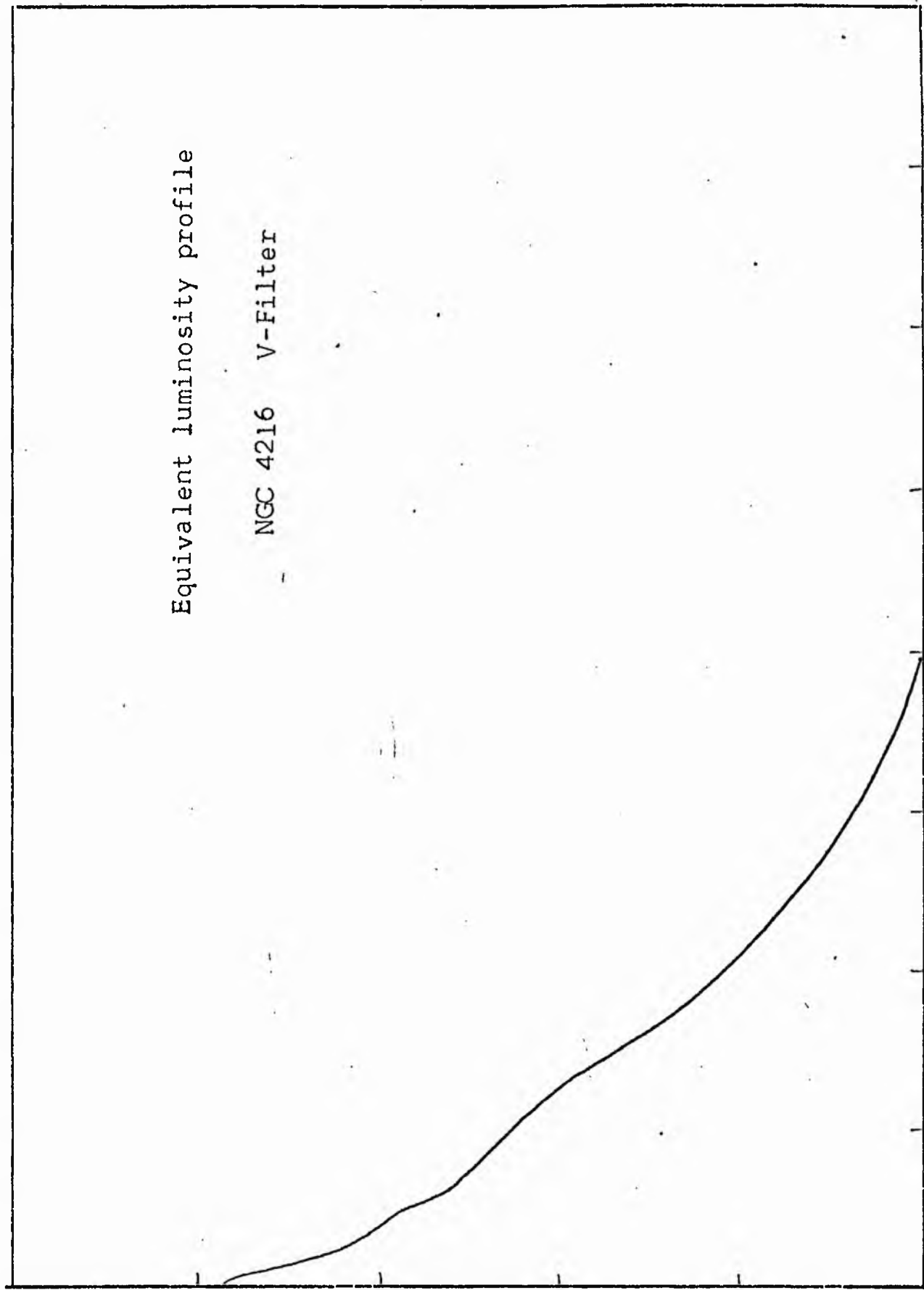


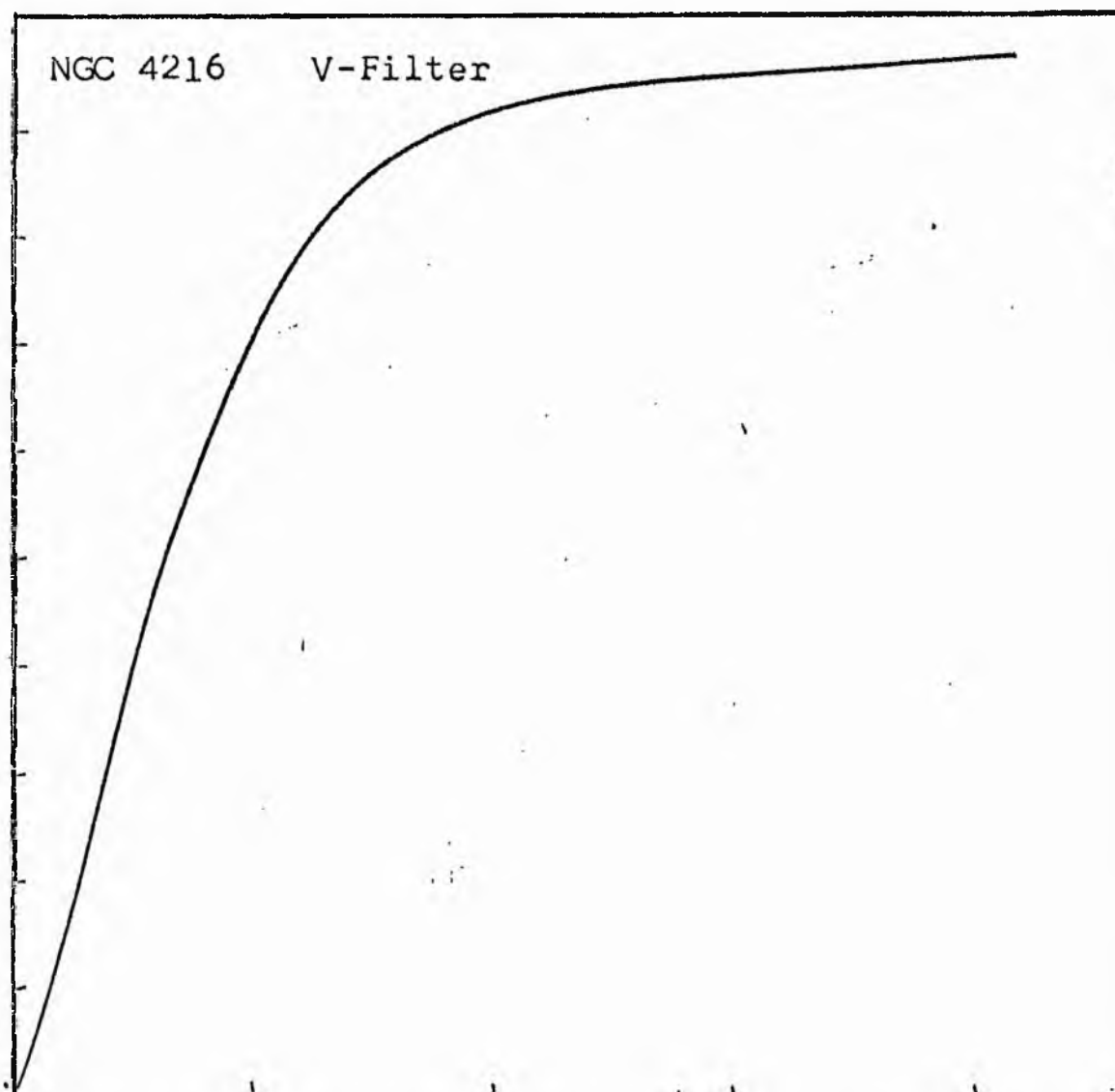
NGC 4216
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4216 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4216
V COLOUR

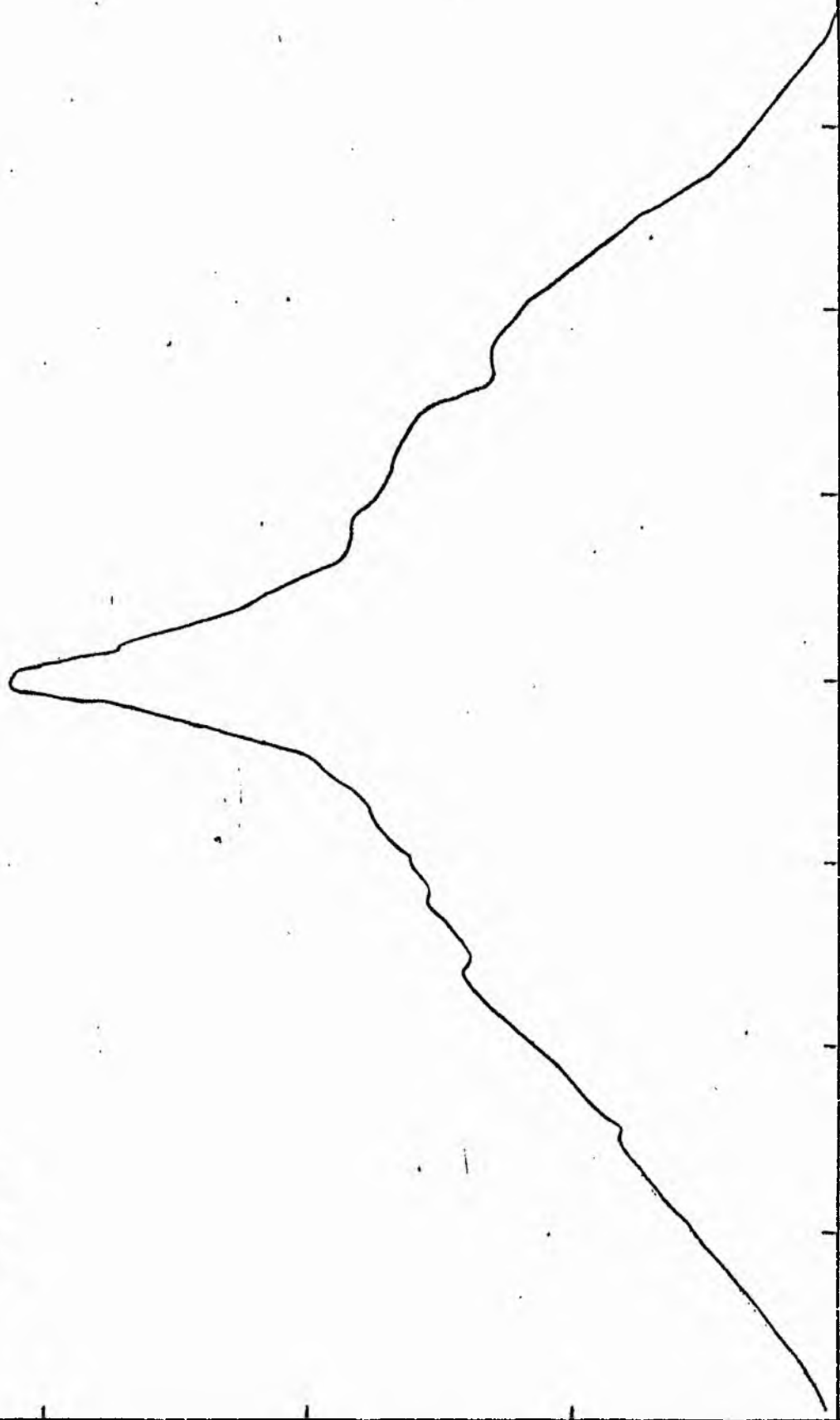
LOG I	I	T	R	AREA	ΔA	P	ΣP	KIRI	ρ	LOG J	μ
1.84	69.183		0.0	0.0			0.0	0.0	0.0	1.301	17.34
		66.139			4.52	299.2068					
1.80	63.096		1.20	4.52			299.21	0.01	0.04	1.261	17.44
		56.607			20.11	1138.1516					
1.70	50.119		2.80	24.63			1437.36	0.03	0.10	1.161	17.69
		44.965			28.18	1267.1038					
1.60	39.811		4.10	52.81			2704.46	0.05	0.14	1.061	17.94
		35.717			64.09	2289.0237					
1.50	31.623		6.10	116.90			4993.48	0.09	0.21	0.961	18.19
		28.371			45.96	1303.9592					
1.40	25.119		7.20	162.86			6297.44	0.11	0.25	0.861	18.44
		22.536			69.49	1566.0427					
1.30	19.952		8.60	232.35			7863.48	0.14	0.29	0.761	18.69
		17.901			114.01	2040.0257					
1.20	15.849		10.50	346.36			9904.30	0.18	0.36	0.661	18.94
		14.219			76.37	1085.9326					
1.10	12.589		11.60	422.73			10990.23	0.20	0.39	0.561	19.19
		11.295			190.38	2150.2888					
1.00	10.000		13.97	613.12			13140.52	0.24	0.48	0.461	19.44
		8.972			637.24	5717.0820					
0.90	7.943		19.95	1250.36			18857.60	0.34	0.68	0.361	19.69
		7.126			840.81	5991.8945					
0.80	6.310		25.80	2091.17			24849.50	0.45	0.88	0.261	19.94
		5.661			148.44	840.2800					
0.70	5.012		26.70	2239.61			25689.77	0.46	0.91	0.161	20.19
		4.496			258.71	1163.2756					
0.60	3.981		28.20	2498.32			26853.05	0.49	0.96	0.061	20.44
		3.572			501.30	1790.4778					
0.50	3.162		30.90	2999.62			28643.52	0.52	1.05	-0.039	20.69
		2.837			1074.14	3047.4358					
0.40	2.512		36.01	4073.76			31690.93	0.57	1.23	-0.139	20.94
		2.254			1781.06	4013.7212					
0.30	1.995		43.17	5854.82			35704.64	0.65	1.47	-0.239	21.19
		1.790			1905.19	3410.4160					
0.20	1.585		49.70	7760.02			39115.06	0.71	1.69	-0.339	21.44
		1.422			2197.85	3125.1226					
0.10	1.259		56.30	9957.87			42240.18	0.76	1.92	-0.439	21.69
		1.129			1465.24	1654.9243					
-0.00	1.000		60.30	11423.11			43895.10	0.79	2.05	-0.539	21.94
		0.897			1727.87	1550.1692					
-0.10	0.794		64.70	13150.98			45445.27	0.82	2.20	-0.639	22.19
		0.713			1590.16	1133.2083					
-0.20	0.631		68.50	14741.14			46578.48	0.84	2.33	-0.739	22.44
		0.566			1229.74	696.1155					
-0.30	0.501		71.30	15970.87			47274.59	0.85	2.43	-0.839	22.69
		0.450			1606.45	722.3286					
-0.40	0.398		74.80	17577.32			47996.92	0.87	2.55	-0.939	22.94
		0.357			1389.38	496.2390					
-0.50	0.316		77.70	18966.70			48493.16	0.88	2.64	-1.039	23.19
		0.284			1240.14	351.8369					
-0.60	0.251		80.20	20206.85			48844.99	0.88	2.73	-1.139	23.44
		0.225			2812.72	633.8630					
-0.70	0.200		85.60	23019.57			49478.05	0.89	2.91	-1.239	23.69
		0.179			2654.02	475.0872					
-0.80	0.158		90.40	25673.59			49953.94	0.90	3.08	-1.339	23.94
		0.142			3703.09	526.5430					
-0.90	0.126		96.70	29376.68			50480.48	0.91	3.29	-1.439	24.19
		0.113			4276.85	483.0513					
-1.00	0.100		103.50	33653.53			50963.53	0.92	3.52	-1.539	24.44
		0.090			4636.70	415.9861					
-1.10	0.079		110.40	38290.22			51379.52	0.93	3.76	-1.639	24.69
		0.071			3764.67	268.2859					
-1.20	0.063		115.70	42054.89			51647.80	0.93	3.94	-1.739	24.94
		0.057			6875.52	389.2036					
-1.30	0.050		124.80	48930.42			52037.00	0.94	4.25	-1.839	25.19
		0.045			4817.95	216.6380					
-1.40	0.040		130.80	53748.37			52253.64	0.94	4.45	-1.939	25.44
		0.036			5561.00	198.6211					
-1.50	0.032		137.40	59309.37			52452.26	0.95	4.68	-2.039	25.69
		0.028			11093.94	314.7449					
-1.60	0.025		149.70	70403.31			52767.00	0.95	5.10	-2.139	25.94
		0.023			9319.25	210.0171					
-1.70	0.020		159.30	79722.56			52977.02	0.96	5.42	-2.239	26.19
		0.018			18922.94	338.7371					
-1.80	0.016		177.20	98645.50			53315.75	0.96	6.03	-2.339	26.44
		0.014			17528.31	249.2385					
-1.90	0.013		192.30	116173.81			53564.99	0.97	6.55	-2.439	26.69
		0.011			16367.31	184.8642					
-2.00	0.010		205.40	132541.12			53749.86	0.97	6.99	-2.539	26.94
-∞							55350.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4216

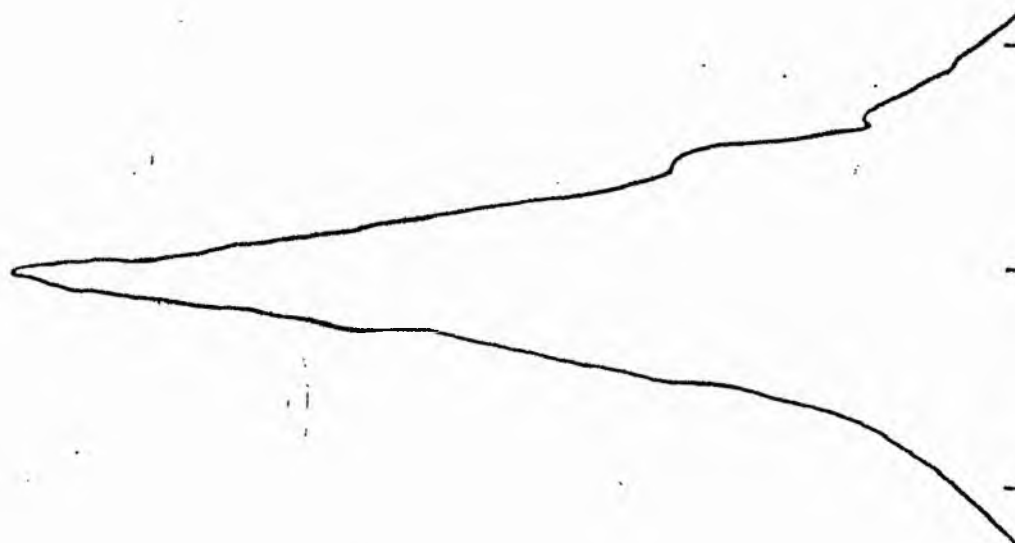
V-FILTER

Total luminosity	L_T	= 15.37
Total apparent magnitude	m_T	= 10.09
Apparent central surface brightness	μ_0	= 17.34
Major axis at threshold	$2a_m$	= 12.7
Minor axis at threshold	$2b_m$	= 4.23
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.67
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.93
Gradient of exponential component	$G(a)$	= -0.54
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.06
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.72
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.30
Axis ratio	b/a	= 0.66
Equivalent radius	r_1^*	= 0.24
Surface brightness	μ_1	= 19.46
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.63
Axis ratio	b/a	= 0.39
Equivalent radius	r_e^*	= 0.49
Surface brightness	μ_e	= 20.52
Mean surface brightness	μ_e'	= 10.53
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.36
Axis ratio	b/a	= 0.32
Equivalent radius	r_3^*	= 0.91
Surface brightness	μ_3	= 21.64
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.00 \\ = 1.86 \end{cases}$

NGC 4235
B-Filter
Axis 1

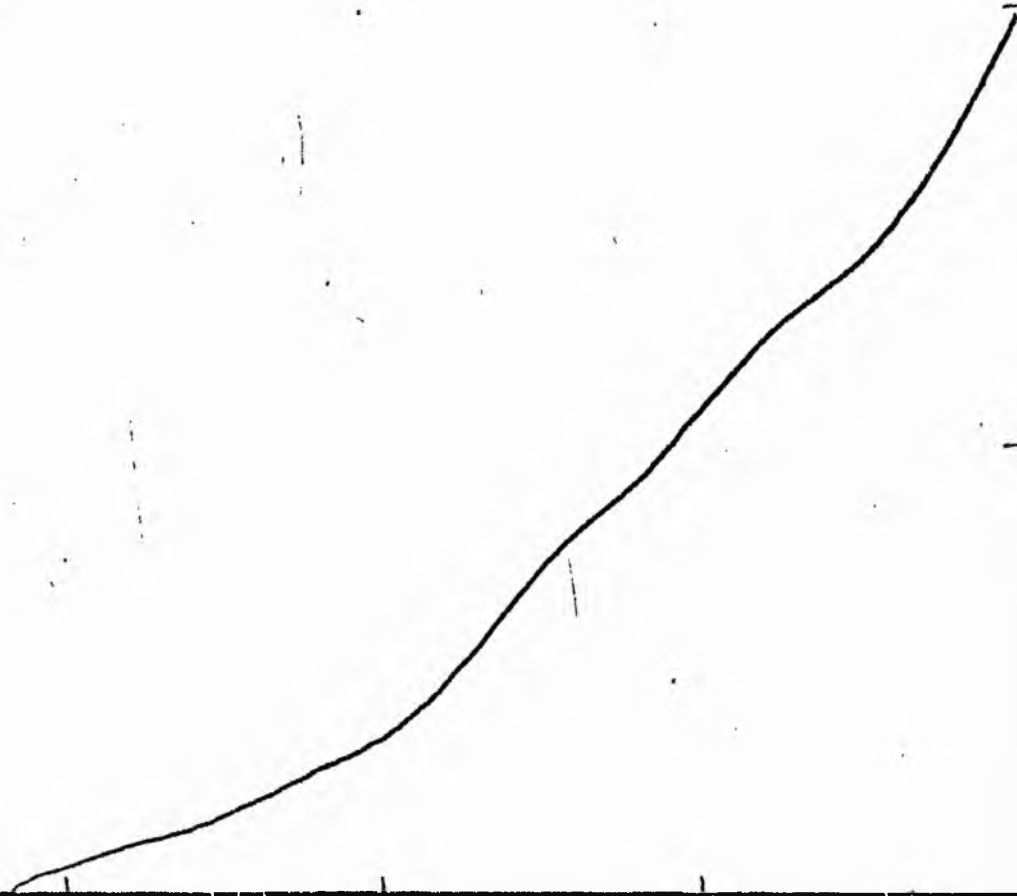


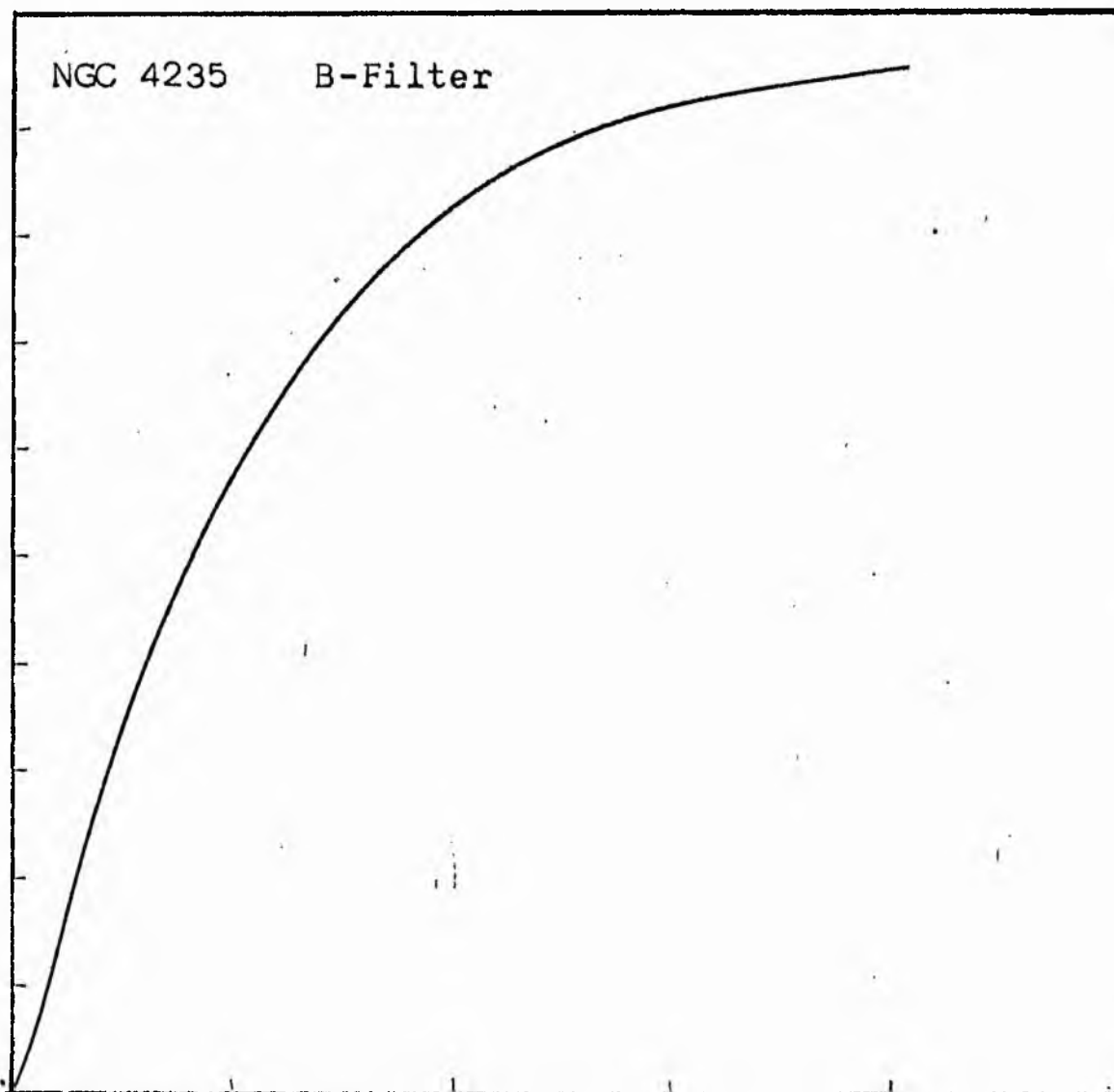
NGC 4235
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4235 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4235
B COLOUR

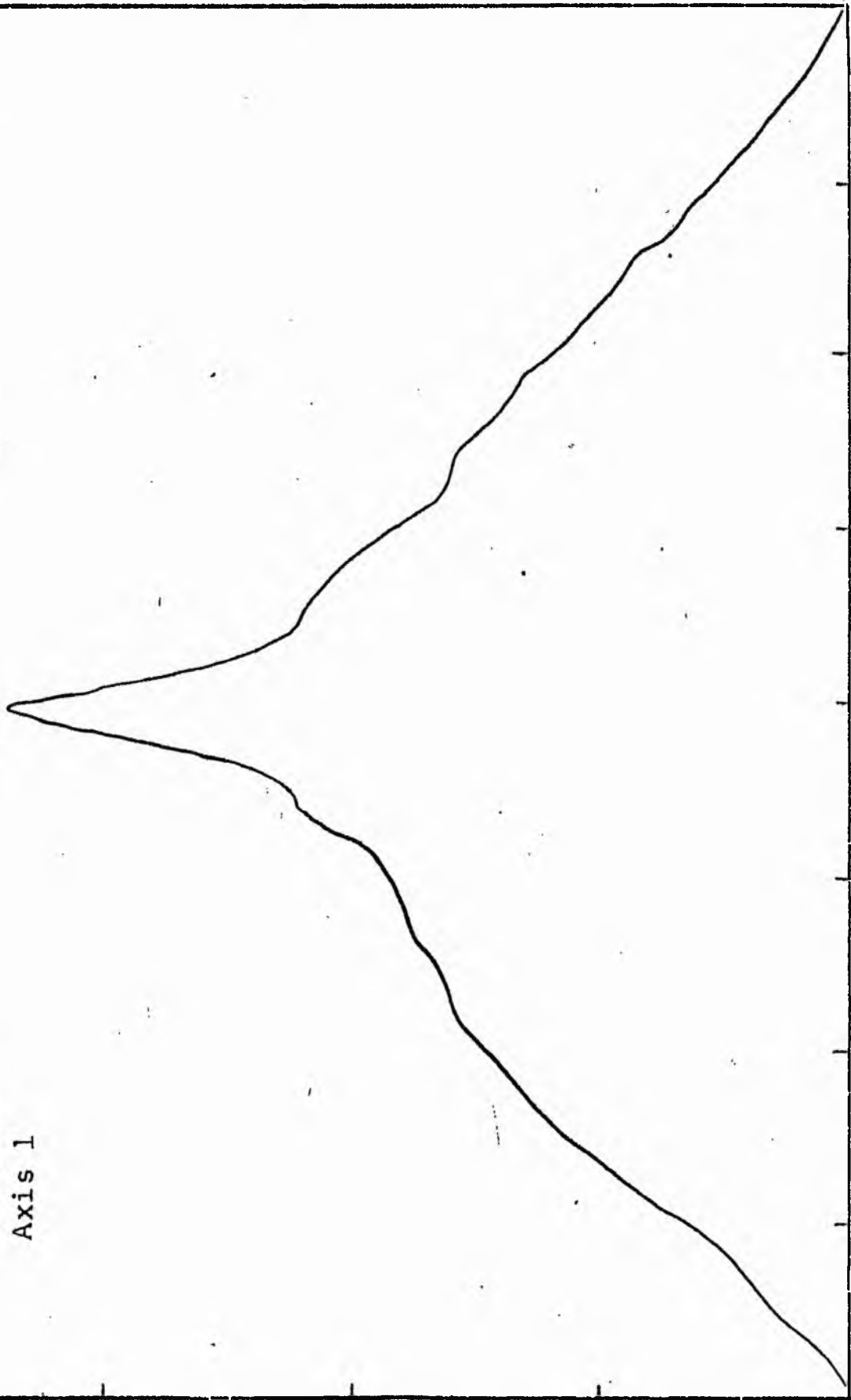
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.14	13.804		0.0	0.0			0.0	0.0	0.0	1.221	19.32
1.10	12.589	13.197	2.28	16.33	16.33	215.5156	215.52	0.03	0.11	1.181	19.42
1.00	10.000	11.295	2.90	26.42	10.09	113.9573	329.47	0.05	0.14	1.081	19.67
0.90	7.943	8.972	5.40	91.61	65.19	584.8425	914.32	0.14	0.26	0.981	19.92
0.80	6.310	7.126	6.10	116.90	25.29	180.2258	1094.54	0.16	0.30	0.881	20.17
0.70	5.012	5.661	6.80	145.27	28.37	160.5861	1255.13	0.19	0.33	0.781	20.42
0.60	3.981	4.496	6.90	149.57	4.30	19.3528	1274.48	0.19	0.34	0.681	20.67
0.50	3.162	3.572	9.90	307.91	158.34	565.5239	1840.00	0.27	0.48	0.581	20.92
0.40	2.512	2.837	11.40	408.28	100.37	284.7688	2124.77	0.31	0.55	0.481	21.17
0.30	1.995	2.254	12.80	514.72	106.44	239.8635	2364.64	0.35	0.62	0.381	21.42
0.20	1.585	1.790	14.70	678.87	164.15	293.8381	2658.47	0.39	0.71	0.281	21.67
0.10	1.259	1.422	16.30	834.69	155.82	221.5636	2880.04	0.43	0.79	0.181	21.92
-0.00	1.000	1.129	18.40	1063.62	228.93	258.5654	3138.60	0.46	0.89	0.081	22.17
-0.10	0.794	0.897	21.10	1398.67	335.05	300.5950	3439.20	0.51	1.03	-0.019	22.42
-0.20	0.631	0.713	24.50	1885.74	487.07	347.1096	3786.31	0.56	1.19	-0.119	22.67
-0.30	0.501	0.566	28.60	2569.70	683.95	387.1675	4173.47	0.62	1.39	-0.219	22.92
-0.40	0.398	0.450	32.50	3318.31	748.61	336.6111	4510.08	0.67	1.58	-0.319	23.17
-0.50	0.316	0.357	37.40	4394.33	1076.02	384.3210	4894.40	0.72	1.82	-0.419	23.42
-0.60	0.251	0.284	40.80	5229.62	835.29	236.9774	5131.38	0.76	1.98	-0.519	23.67
-0.70	0.200	0.225	43.60	5972.04	742.42	167.3103	5298.69	0.78	2.12	-0.619	23.92
-0.80	0.158	0.179	46.50	6792.91	820.87	146.9417	5445.63	0.81	2.26	-0.719	24.17
-0.90	0.126	0.142	49.80	7791.27	998.36	141.9583	5587.59	0.83	2.42	-0.819	24.42
-1.00	0.100	0.113	56.70	10099.87	2308.60	260.7480	5848.33	0.86	2.76	-0.919	24.67
-1.10	0.079	0.090	60.20	11385.25	1285.38	115.3200	5963.65	0.88	2.93	-1.019	24.92
-1.20	0.063	0.071	63.50	12667.69	1282.43	91.3919	6055.04	0.90	3.09	-1.119	25.17
-1.30	0.050	0.057	66.50	13892.91	1225.22	69.3564	6124.39	0.91	3.23	-1.219	25.42
-1.40	0.040	0.045	68.70	14827.34	934.43	42.0165	6166.41	0.91	3.34	-1.319	25.67
-1.50	0.032	0.036	72.50	16513.00	1685.66	60.2066	6226.61	0.92	3.53	-1.419	25.92
-1.60	0.025	0.028	75.80	18050.45	1537.45	43.6191	6270.23	0.93	3.69	-1.519	26.17
-1.70	0.020	0.023	81.20	20713.89	2663.45	60.0233	6330.25	0.94	3.95	-1.619	26.42
-1.80	0.016	0.018	88.40	24550.16	3836.27	68.6728	6398.92	0.95	4.30	-1.719	26.67
-1.90	0.013	0.014	93.70	27582.20	3032.04	43.1134	6442.04	0.95	4.56	-1.819	26.92
-2.00	0.010	0.011	101.20	32174.43	4592.22	51.8681	6493.90	0.96	4.92	-1.919	27.17
-∞							6763.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4235

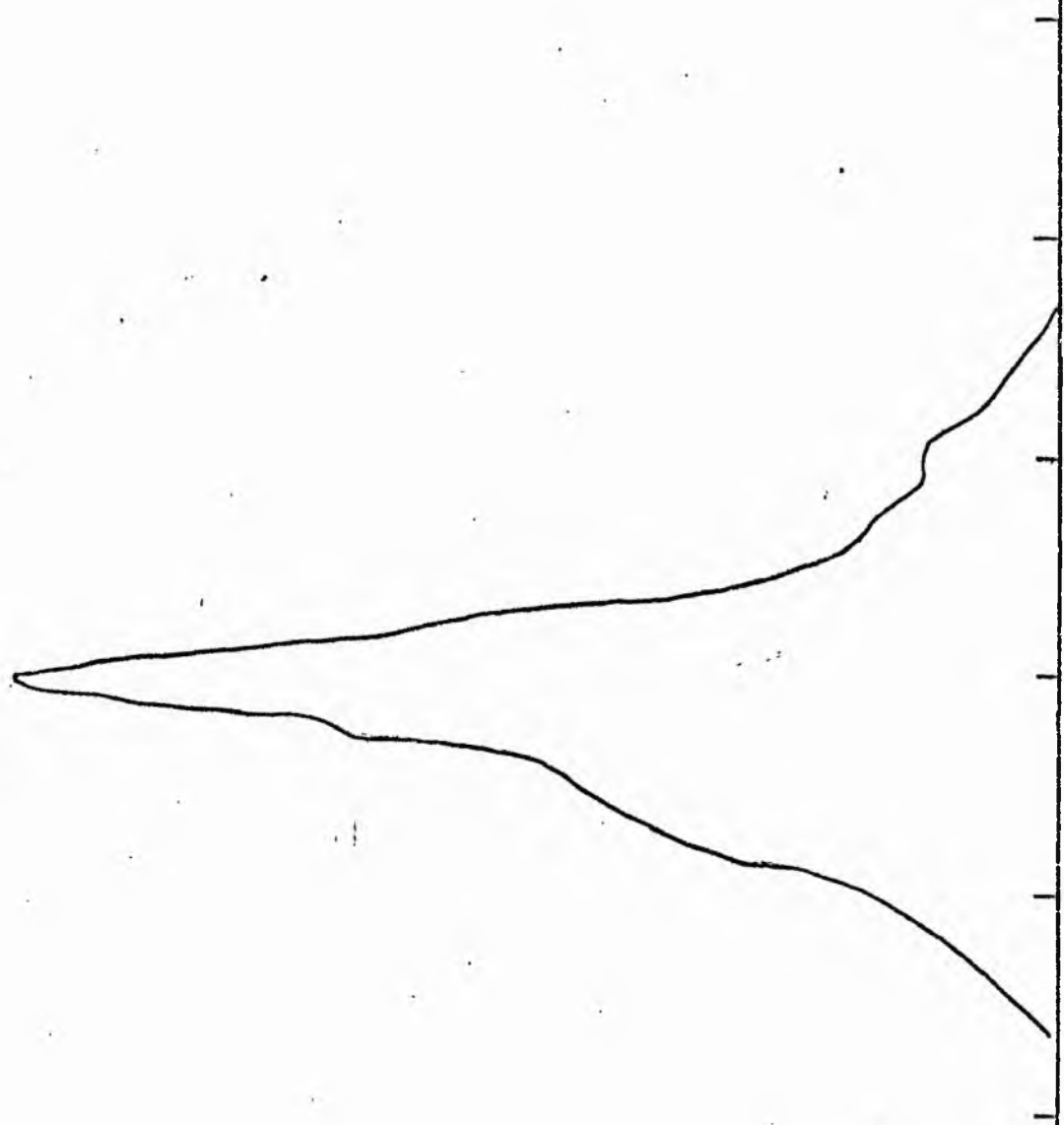
B-FILTER

Total luminosity	L_T	= 1.87
Total apparent magnitude	m_T	= 12.6
Apparent central surface brightness	μ_0	= 19.32
Major axis at threshold	$2a_m$	= 6.13
Minor axis at threshold	$2b_m$	= 1.96
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.1
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.95
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.62
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.62
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.19
Axis ratio	b/a	= 0.65
Equivalent radius	r_1^*	= 0.15
Surface brightness	μ_1	= 20.88
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.46
Axis ratio	b/a	= 0.55
Equivalent radius	r_e^*	= 0.34
Surface brightness	μ_e	= 22.37
Mean surface brightness	μ_e'	= 12.24
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.37
Axis ratio	b/a	= 0.23
Equivalent radius	r_3^*	= 0.66
Surface brightness	μ_3	= 23.61
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.26 \\ = 1.94 \end{matrix}$

NGC 4235
V-Filter
Axis 1

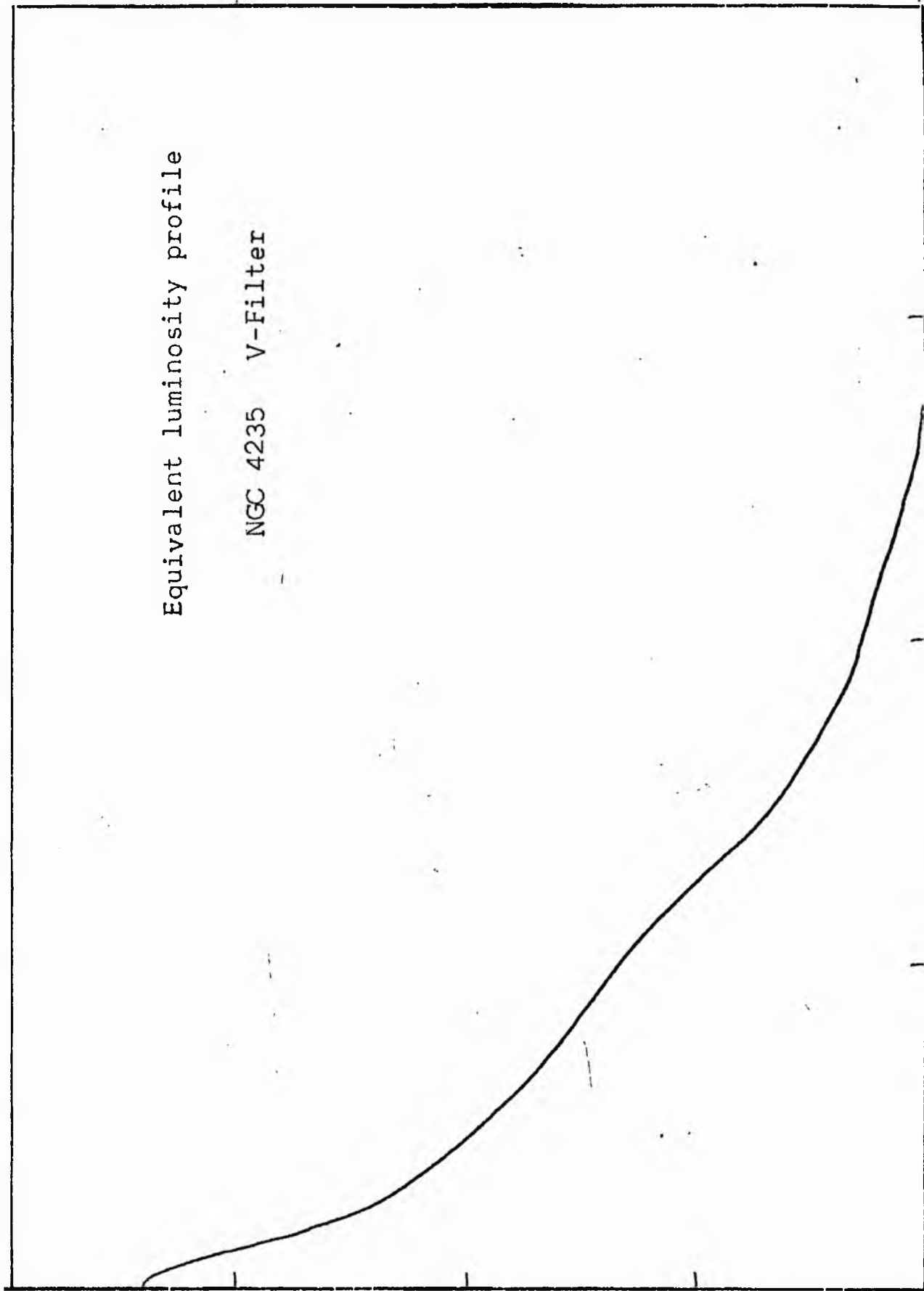


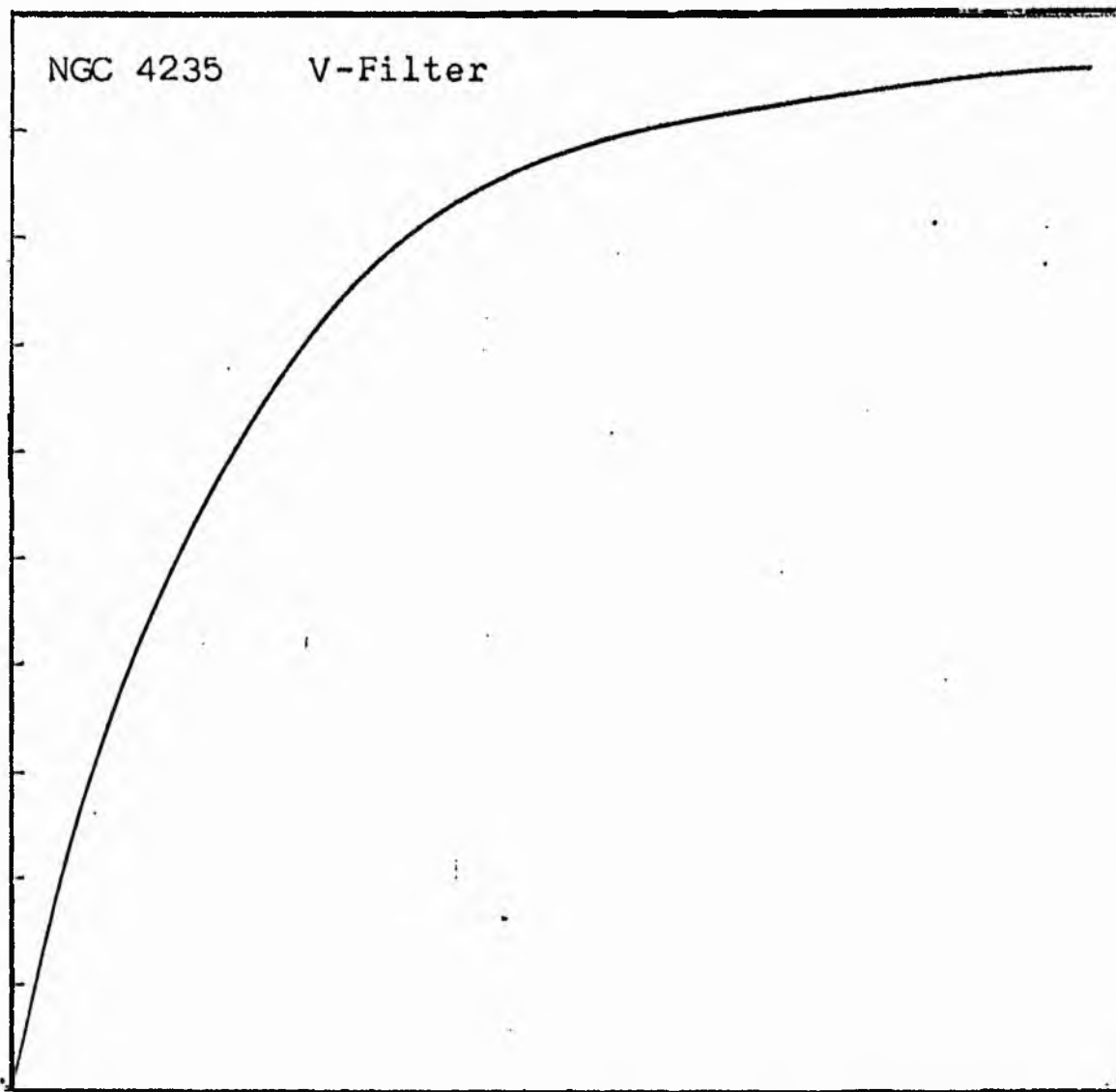
NGC 4235
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4235 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4235
V COLOUR

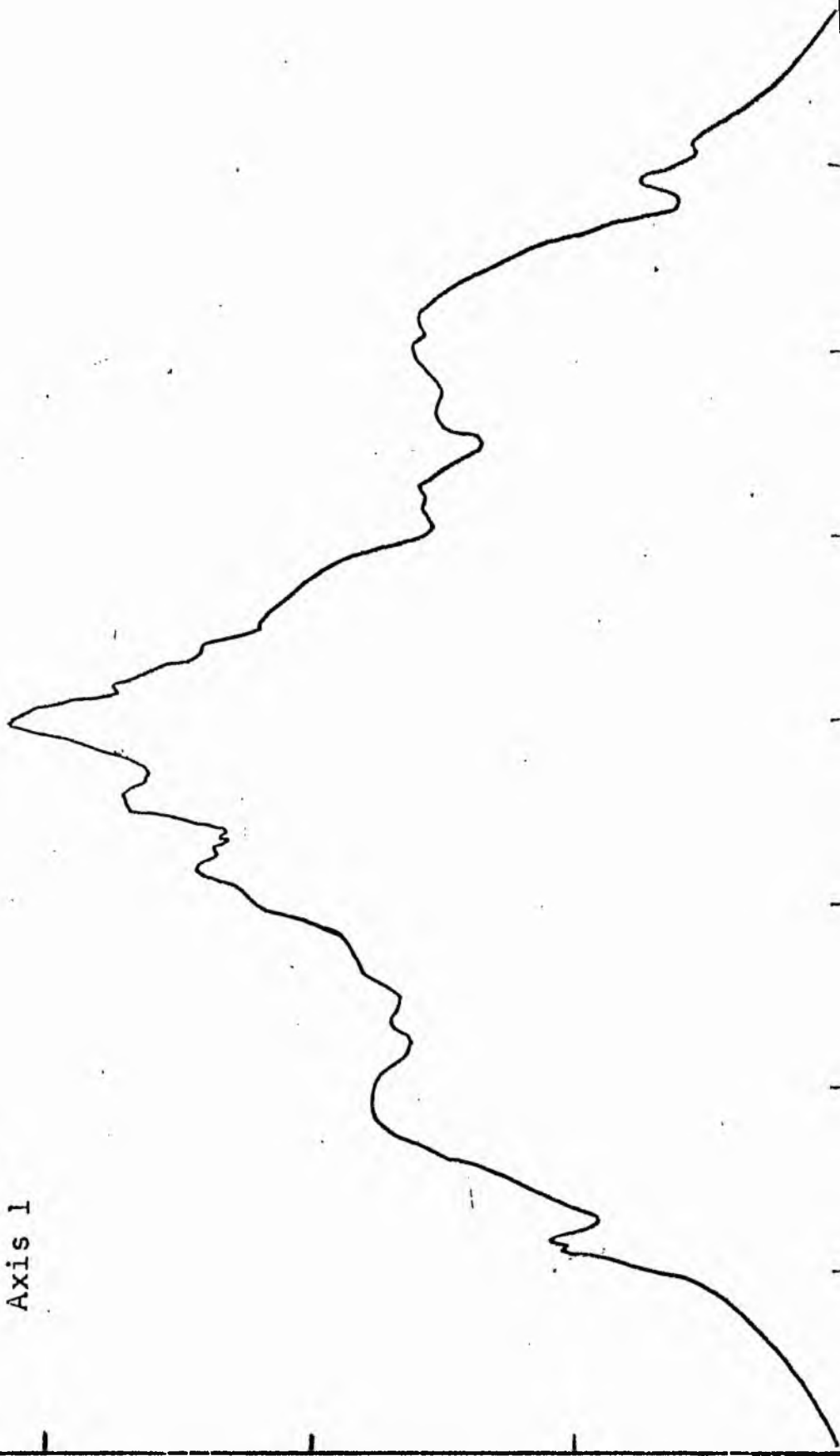
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.32	20.893	20.423	0.0	0.0	18.70	381.9824	0.0	0.0	0.0	1.323	18.25
1.30	19.953	17.901	2.44	18.70	22.69	406.2148	381.98	0.04	0.11	1.303	18.30
1.20	15.849	14.219	3.63	41.40	24.22	344.3196	788.20	0.08	0.16	1.203	18.55
1.10	12.589	11.295	4.57	65.61	40.07	452.5886	1132.52	0.11	0.21	1.103	18.80
1.00	10.000	8.972	5.80	105.68	37.46	336.0374	1585.11	0.15	0.26	1.003	19.05
0.90	7.943	7.126	6.75	143.14	21.54	153.4734	1921.14	0.19	0.30	0.903	19.30
0.80	6.310	5.661	7.24	164.67	86.97	492.3379	2074.62	0.20	0.33	0.803	19.55
0.70	5.012	4.496	8.95	251.65	88.15	396.3401	2566.95	0.25	0.40	0.703	19.80
0.60	3.981	3.572	10.40	339.79	68.49	244.6118	2963.29	0.29	0.47	0.603	20.05
0.50	3.162	2.837	11.40	408.28	98.43	279.2417	3207.91	0.31	0.51	0.503	20.30
0.40	2.512	2.254	12.70	506.71	144.73	326.1665	3487.15	0.34	0.57	0.403	20.55
0.30	1.995	1.790	14.40	651.44	435.42	779.4387	3813.31	0.37	0.65	0.303	20.80
0.20	1.585	1.422	18.60	1086.86	132.36	188.1982	4592.75	0.44	0.84	0.203	21.05
0.10	1.259	1.129	19.70	1219.22	329.08	371.6843	4780.95	0.46	0.88	0.103	21.30
-0.00	1.000	0.897	22.20	1548.30	542.87	487.0376	5152.63	0.50	1.00	0.003	21.55
-0.10	0.794	0.713	25.80	2091.17	661.37	471.3184	5639.66	0.55	1.16	-0.097	21.80
-0.20	0.631	0.566	29.60	2752.54	760.50	430.4941	6110.98	0.59	1.33	-0.197	22.05
-0.30	0.501	0.450	33.44	3513.03	1131.50	508.7761	6541.47	0.63	1.50	-0.297	22.30
-0.40	0.398	0.357	38.45	4644.54	1319.29	471.2046	7050.25	0.68	1.73	-0.397	22.55
-0.50	0.316	0.284	43.57	5963.82	1067.80	302.9441	7521.45	0.73	1.96	-0.497	22.80
-0.60	0.251	0.225	47.31	7031.62	1584.56	357.0913	7824.39	0.76	2.12	-0.597	23.05
-0.70	0.200	0.179	52.37	8616.18	1123.57	201.1281	8181.48	0.79	2.35	-0.697	23.30
-0.80	0.158	0.142	55.68	9739.76	974.83	138.6115	8382.61	0.81	2.50	-0.797	23.55
-0.90	0.126	0.113	58.40	10714.59	1557.26	175.8864	8521.21	0.83	2.62	-0.897	23.80
-1.00	0.100	0.090	62.50	12271.84	1330.11	119.3327	8697.10	0.84	2.81	-0.997	24.05
-1.10	0.079	0.071	65.80	13601.95	1355.17	96.5753	8816.43	0.85	2.95	-1.097	24.30
-1.20	0.063	0.057	69.00	14957.12	1830.31	103.6090	8913.00	0.86	3.10	-1.197	24.55
-1.30	0.050	0.045	73.10	16787.43	1636.41	73.5812	9016.61	0.87	3.28	-1.297	24.80
-1.40	0.040	0.036	76.58	18423.84	3848.91	137.4712	9090.19	0.88	3.44	-1.397	25.05
-1.50	0.032	0.028	84.20	22272.75	3628.54	102.9454	9227.66	0.89	3.78	-1.497	25.30
-1.60	0.025	0.023	90.80	25901.29	5703.41	128.5316	9330.60	0.90	4.08	-1.597	25.55
-1.70	0.020	0.018	100.30	31604.70	6720.22	120.2984	9459.13	0.92	4.50	-1.697	25.80
-1.80	0.016	0.014	110.45	38324.92	7960.49	113.1922	9579.43	0.93	4.96	-1.797	26.05
-1.90	0.013	0.011	121.38	46285.41	7380.82	83.3646	9692.62	0.94	5.45	-1.897	26.30
-2.00	0.010		130.70	53666.22			9775.98	0.95	5.87	-1.997	26.55
-∞							10325.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4235

V-FILTER

Total luminosity	L_T	= 2.86
Total apparent magnitude	m_T	= 11.52
Apparent central surface brightness	μ_o	= 18.25
Major axis at threshold	$2a_m$	= 6.23
Minor axis at threshold	$2b_m$	= 2.98
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.81
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.88
Gradient of exponential component	$G(a)$	= -0.78
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.46
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.61
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.26
Axis ratio	b/a	= 0.56
Equivalent radius	r_1^*	= 0.15
Surface brightness	μ_1	= 19.80
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.63
Axis ratio	b/a	= 0.36
Equivalent radius	r_e^*	= 0.37
Surface brightness	μ_e	= 21.55
Mean surface brightness	μ_e'	= 11.37
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.52
Axis ratio	b/a	= 0.26
Equivalent radius	r_3^*	= 0.77
Surface brightness	μ_3	= 22.97
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.47 \\ = 2.08 \end{cases}$

NGC 4254
B-Filter
Axis 1

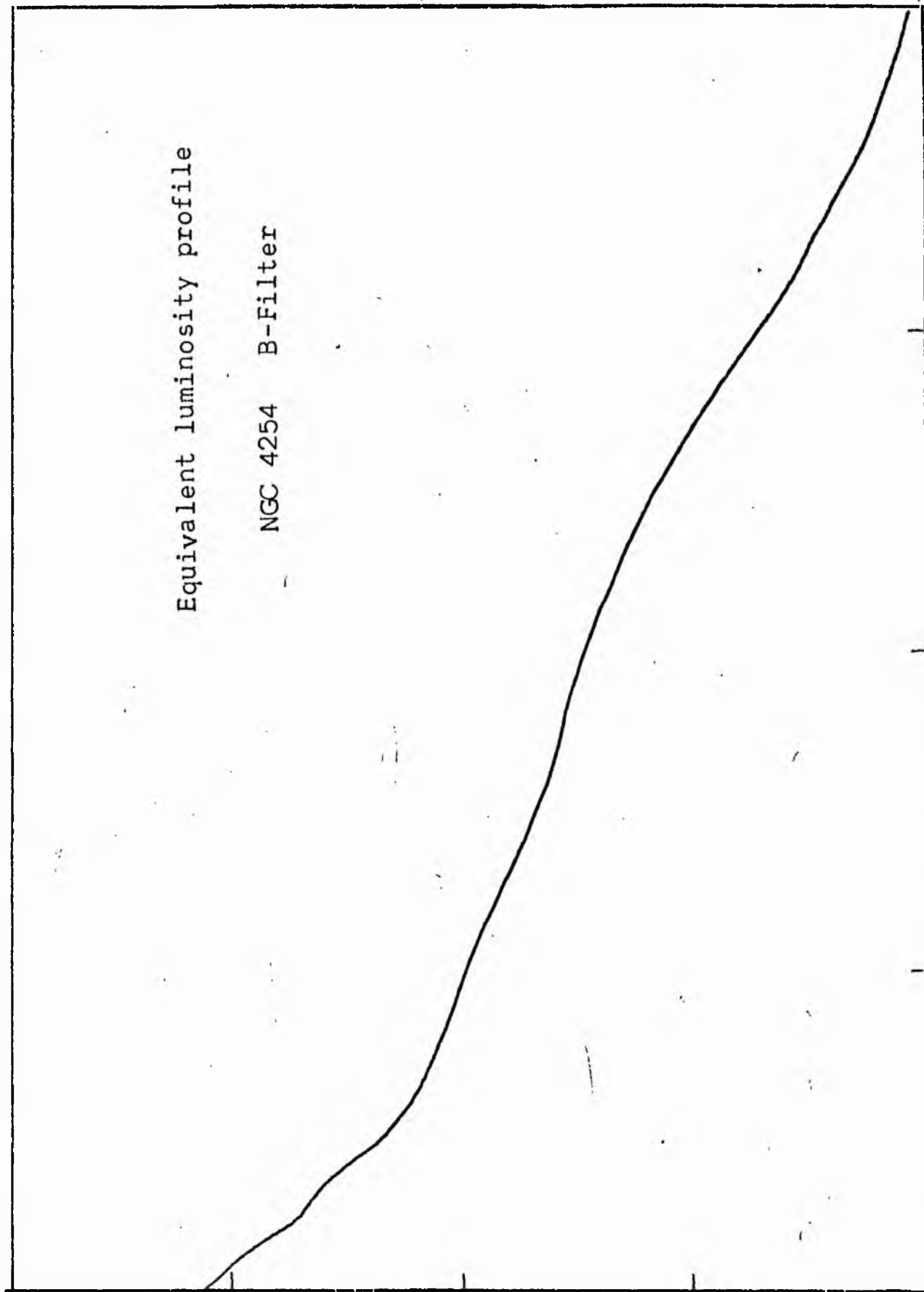


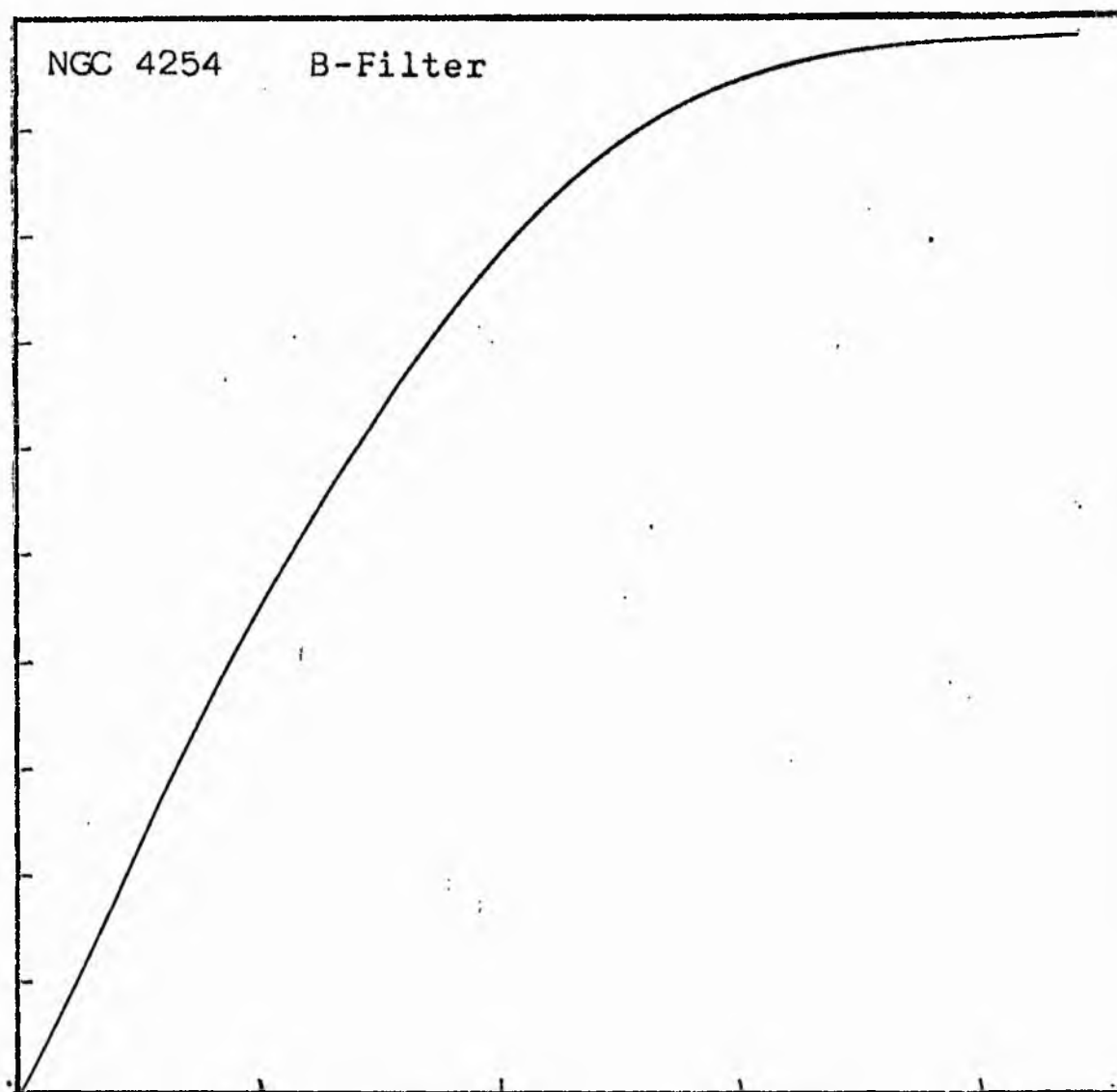
NGC 4254
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4254 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4254
O COLOR

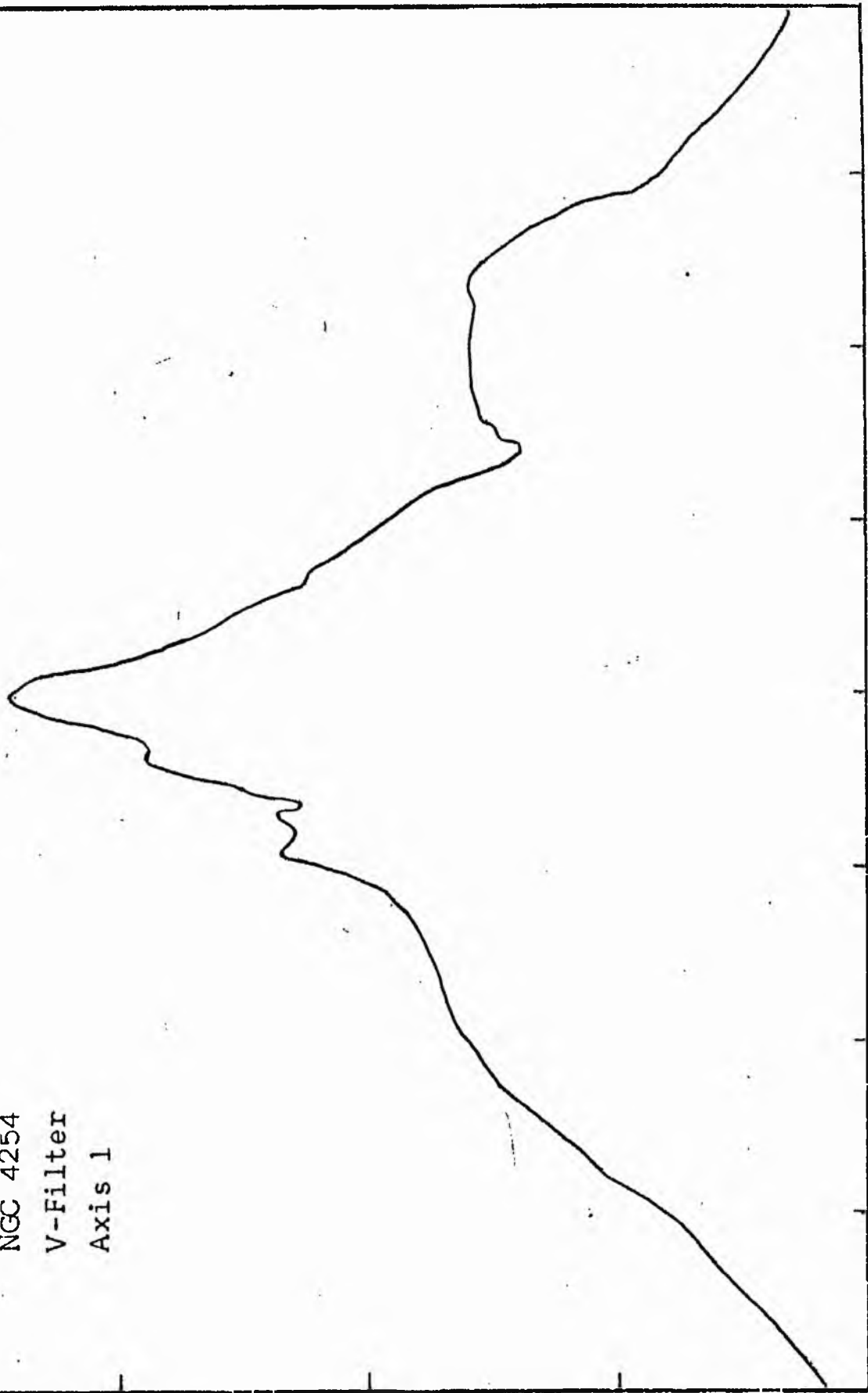
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.06	11.482	10.741	0.0	0.0	37.83	406.2466	0.0	0.0	0.0	1.171	19.14
1.00	10.060	8.972	3.47	37.83	82.94	744.0642	406.30	0.01	0.06	1.111	18.29
0.90	7.943	7.126	6.20	120.76	145.14	1034.3401	1150.36	0.03	0.11	1.011	17.54
0.80	6.310	5.661	9.20	265.90	80.46	455.4407	2184.70	0.06	0.17	0.911	16.79
0.70	5.012	4.496	10.50	346.36	508.94	2798.4250	2643.14	0.08	0.19	0.811	20.04
0.60	3.981	3.572	16.50	855.30	476.60	1521.6626	4928.57	0.14	0.30	0.711	20.29
0.50	3.162	2.837	20.20	1281.89	308.54	875.3437	6452.23	0.19	0.37	0.611	20.54
0.40	2.512	2.254	22.50	1579.43	204.08	459.9009	7327.57	0.21	0.41	0.511	20.79
0.30	1.995	1.793	23.90	1794.51	976.66	1748.2952	7787.47	0.23	0.43	0.411	21.04
0.20	1.585	1.422	29.70	2771.17	2587.41	3679.0674	9535.76	0.28	0.54	0.311	21.29
0.10	1.259	1.127	41.30	5358.58	1522.26	1719.3340	13214.83	0.38	0.75	0.211	21.54
0.00	1.000	0.897	46.80	6880.84	2382.11	2137.1482	14934.16	0.43	0.85	0.111	21.79
-0.10	0.794	0.713	54.30	9262.95	2503.70	1784.2412	17071.30	0.47	0.99	0.011	22.04
-0.20	0.631	0.566	61.20	11766.64	1892.17	2203.2490	18855.54	0.55	1.11	-0.089	22.29
-0.30	0.501	0.450	70.60	15658.81	7414.58	3333.9480	21058.79	0.61	1.28	-0.189	22.54
-0.40	0.398	0.357	85.70	23071.39	6364.07	2273.0420	24342.73	0.71	1.56	-0.289	22.79
-0.50	0.316	0.284	96.80	29437.46	6195.27	1757.6516	26665.77	0.77	1.76	-0.389	23.04
-0.60	0.251	0.225	106.50	35632.73	5770.43	1350.4126	28423.42	0.82	1.93	-0.489	23.29
-0.70	0.200	0.179	114.80	41463.16	5203.10	931.3975	29723.83	0.86	2.08	-0.589	23.54
-0.80	0.158	0.142	121.80	46606.27	4464.25	634.7776	30655.23	0.89	2.21	-0.689	23.79
-0.90	0.126	0.113	127.50	51070.52	3254.69	367.6060	31290.00	0.91	2.32	-0.789	24.04
-1.00	0.100	0.090	131.50	54325.21	4381.38	343.0825	31657.61	0.92	2.39	-0.889	24.29
-1.10	0.079	0.071	136.70	58706.59	5087.38	362.5498	32050.69	0.93	2.48	-0.989	24.54
-1.20	0.063	0.057	142.50	63793.97	6986.09	345.4656	32413.24	0.94	2.59	-1.089	24.79
-1.30	0.050	0.045	150.10	70780.06	8442.81	379.6311	32808.70	0.95	2.73	-1.189	25.04
-1.40	0.040	0.036	158.80	79222.87	4964.50	177.3172	33188.33	0.96	2.88	-1.289	25.29
-1.50	0.032	0.028	163.70	84187.37	5539.62	157.1652	33365.65	0.97	2.97	-1.389	25.54
-1.60	0.025	0.023	169.00	89727.00	5716.06	128.8171	33522.81	0.97	3.07	-1.489	25.79
-1.70	0.020	0.018	174.30	95443.06	3983.31	71.3053	33651.63	0.97	3.17	-1.589	26.04
-1.80	0.016	0.014	177.90	99426.37	22248.25	316.3542	33722.93	0.98	3.23	-1.689	26.29
-1.90	0.013	0.011	196.80	121674.62	6514.87	73.5842	34039.29	0.99	3.57	-1.789	26.54
-2.00	0.010		202.00	128189.50			34112.87	0.99	3.67	-1.889	26.79
-∞							34552.00	111			∞

PHOTOMETRIC PARAMETERS OF NGC 4254

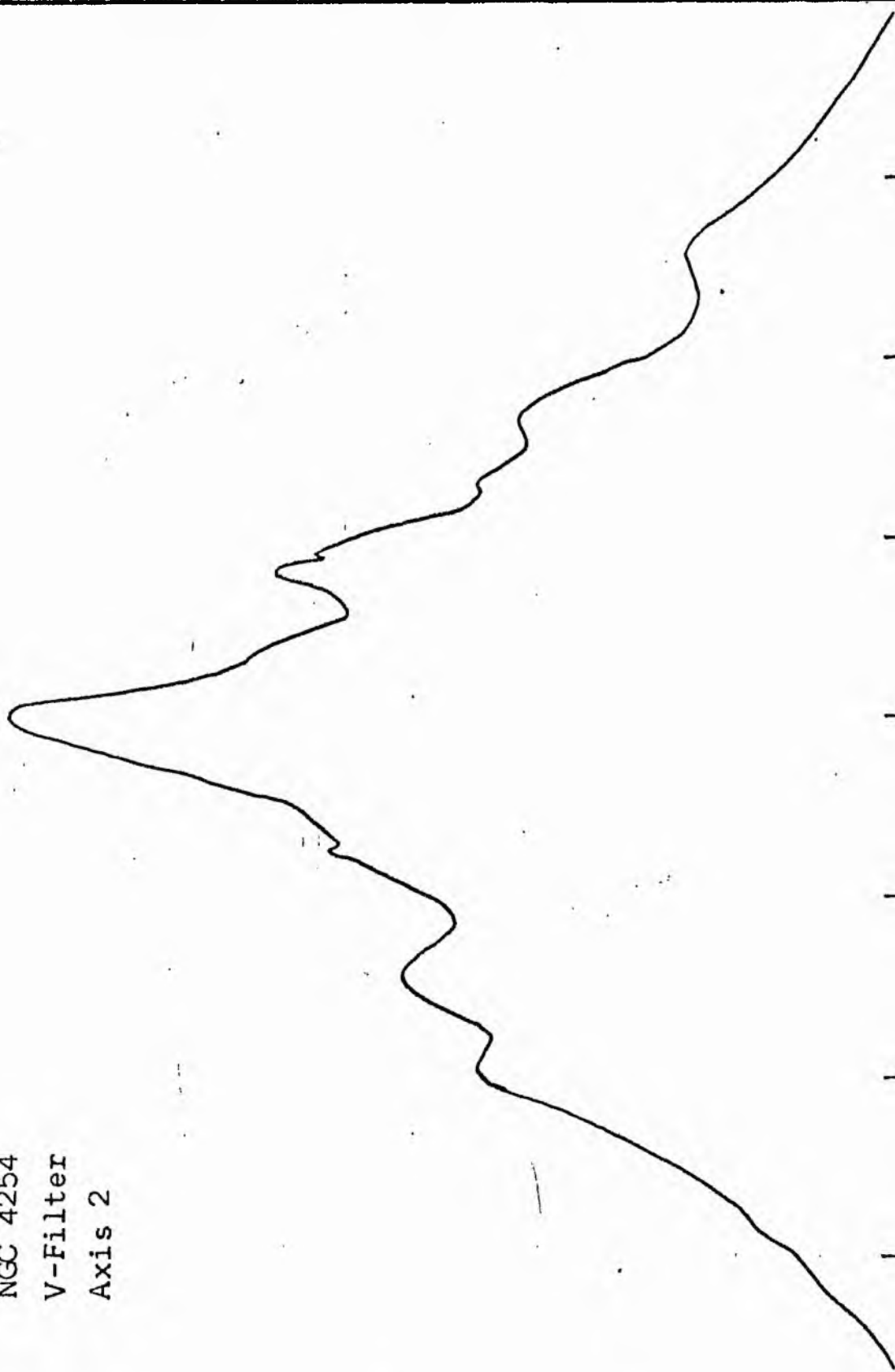
B-FILTER

Total luminosity	L_T	= 9.60
Total apparent magnitude	m_T	= 10.46
Apparent central surface brightness	μ_o	= 19.14
Major axis at threshold	$2a_m$	= 6.92
Minor axis at threshold	$2b_m$	= 6.58
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.20
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.95
Gradient of exponential component	$G(a)$	= -0.87
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.66
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.82
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.62
Axis ratio	b/a	= 0.65
Equivalent radius	r_1^*	= 0.44
Surface brightness	μ_1	= 21.14
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.93
Axis ratio	b/a	= 0.98
Equivalent radius	r_e^*	= 0.92
Surface brightness	μ_e	= 22.08
Mean surface brightness	μ_e'	= 12.25
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.62
Axis ratio	b/a	= 0.93
Equivalent radius	r_3^*	= 1.55
Surface brightness	μ_3	= 22.96
Concentration indices	$\{C_{21}$	= 2.07
	C_{32}	= 1.69

NGC 4254
V-Filter
Axis 1

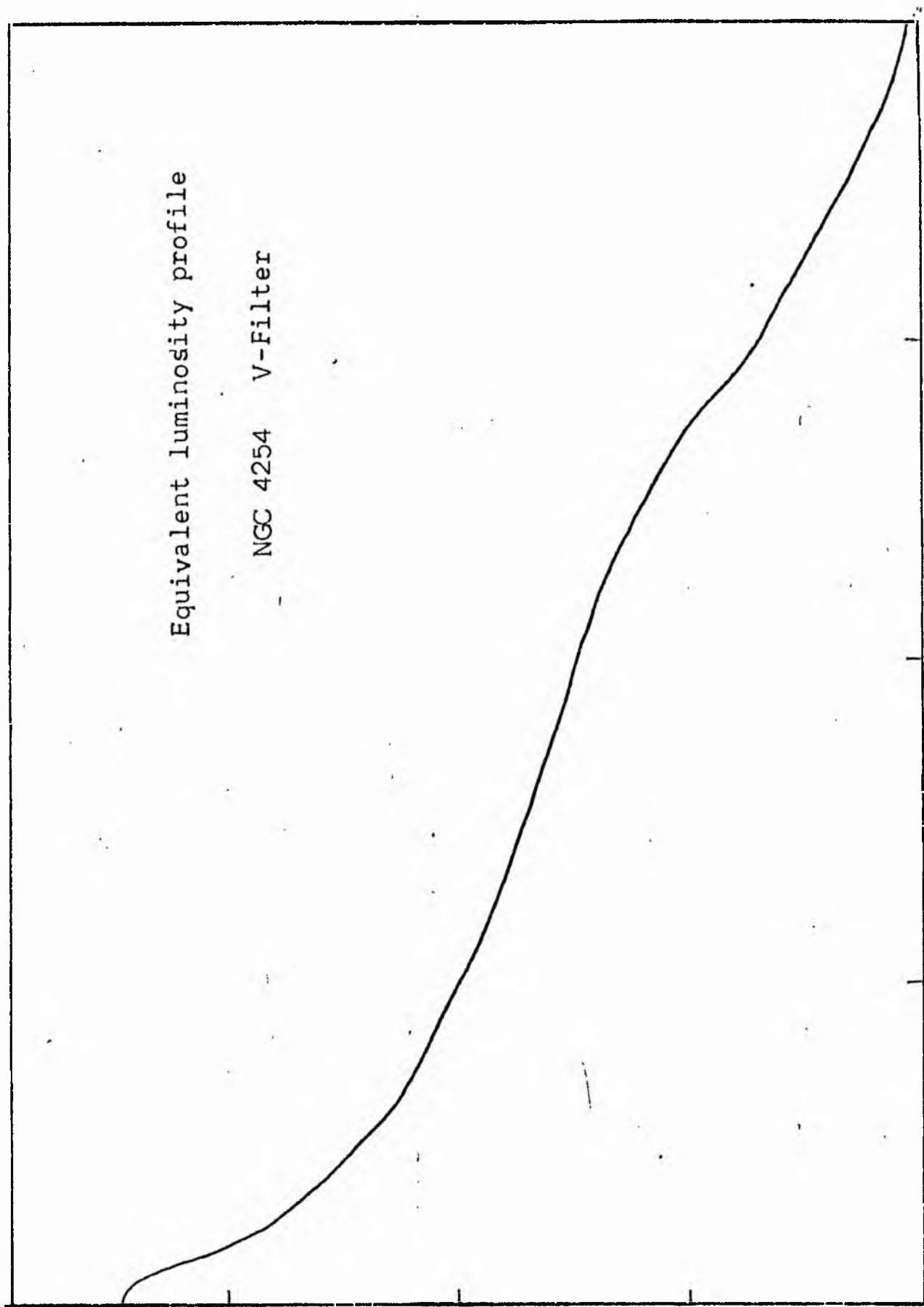


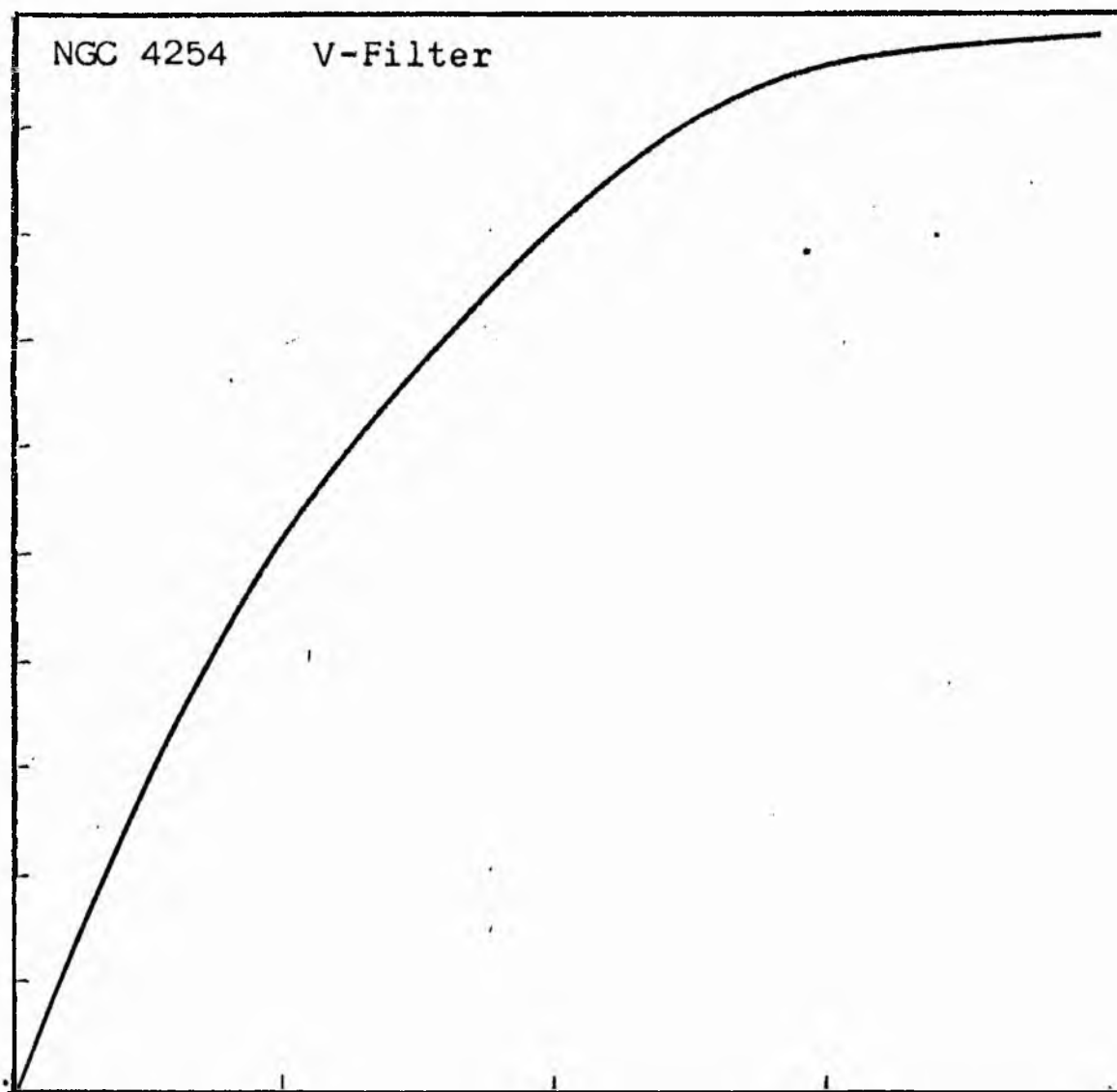
NGC 4254
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4254 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4254
V COLOUR

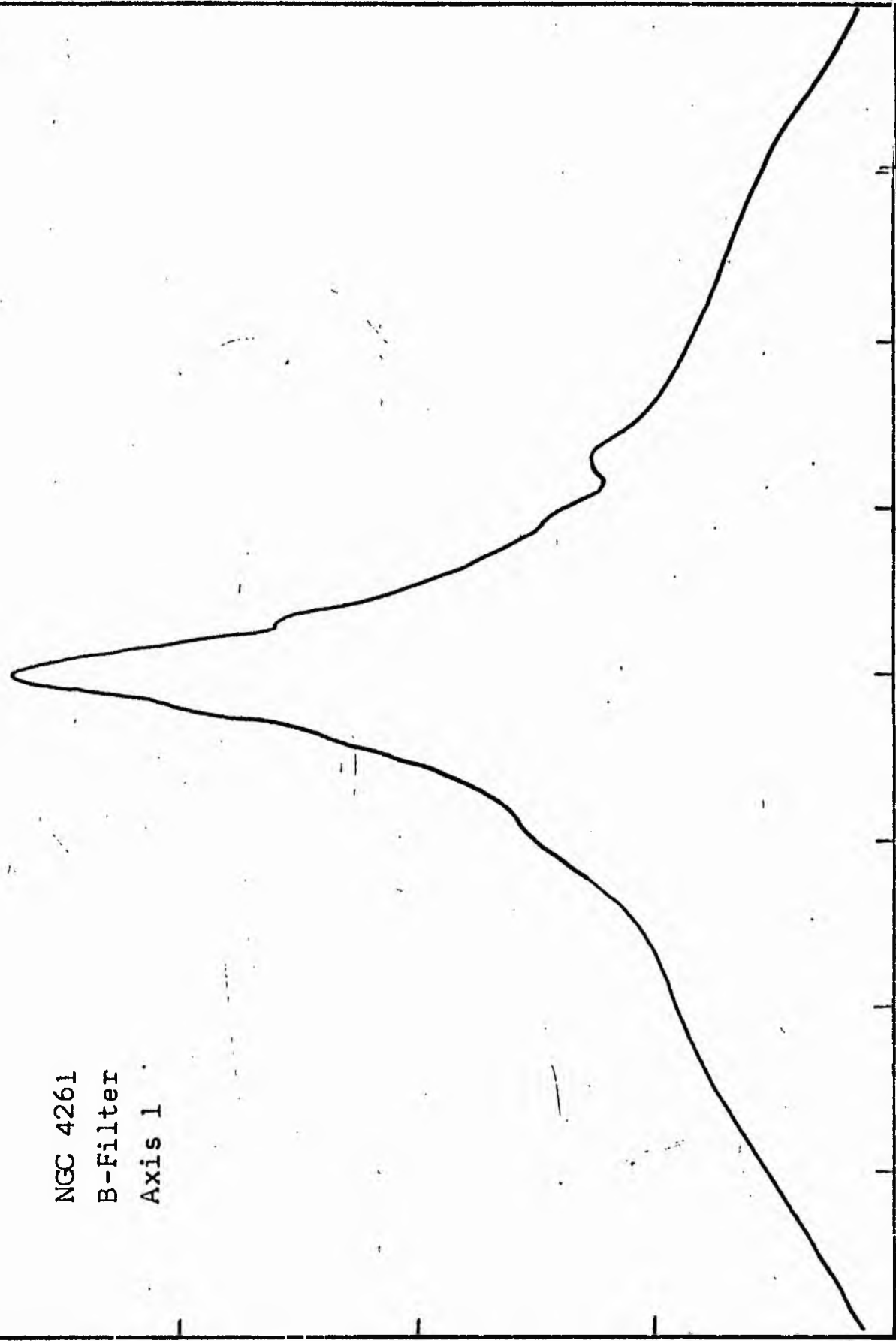
LOG I	I	I	R	AREA	A	P	ΣP	K(R)	ρ	LOG J	μ
1.42	26.303		0.0	0.0	26.42	679.2971	0.0	0.0	0.0	1.328	18.00
1.40	25.119	25.711	2.90	26.42	61.83	1393.3030	679.30	0.02	0.06	1.308	18.05
1.30	19.953	22.536	5.30	88.25	48.60	869.9832	2072.60	0.05	0.11	1.208	18.30
1.20	15.849	17.931	6.60	136.85	12.72	180.9152	2942.58	0.07	0.14	1.108	18.55
1.10	12.589	14.219	6.90	149.57	128.02	1445.9304	3123.50	0.07	0.15	1.008	18.80
1.00	10.000	11.295	9.40	277.59	152.46	1367.8232	4569.43	0.11	0.20	0.908	19.05
0.90	7.943	8.972	11.70	430.05	239.61	1797.5527	5937.25	0.14	0.25	0.808	19.30
0.80	6.310	7.126	14.60	669.66	325.72	1843.7944	7644.80	0.18	0.31	0.708	19.55
0.70	5.012	5.661	17.80	905.38	416.58	1873.1174	9488.59	0.22	0.38	0.608	19.80
0.60	3.981	4.496	21.20	1411.96	308.25	110.9746	11361.71	0.26	0.45	0.508	20.05
0.50	3.162	3.572	23.40	1720.21	519.40	1473.5742	12462.68	0.29	0.50	0.408	20.30
0.40	2.512	2.837	26.70	2239.61	276.46	623.0183	13936.25	0.32	0.57	0.308	20.55
0.30	1.995	2.254	28.30	2516.07	2435.36	4359.4727	14559.27	0.34	0.60	0.208	20.80
0.20	1.585	1.790	39.70	4951.43	1870.72	2659.9890	18918.74	0.44	0.85	0.108	21.05
0.10	1.259	1.422	46.60	6822.15	567.66	641.1445	21578.73	0.50	1.00	0.008	21.30
-0.00	1.000	1.129	48.50	7389.81	2217.46	1989.4204	22219.87	0.51	1.04	-0.092	21.55
-0.10	0.794	0.897	55.30	9607.27	3462.54	2467.5464	24209.29	0.56	1.18	-0.192	21.80
-0.20	0.631	0.713	64.50	13069.81	4696.02	2658.2783	26676.84	0.62	1.38	-0.292	22.05
-0.30	0.501	0.566	75.20	17765.83	7964.60	3581.2524	29335.11	0.68	1.61	-0.392	22.30
-0.40	0.398	0.450	90.50	25730.43	6571.29	2347.0466	32916.36	0.76	1.93	-0.492	22.55
-0.50	0.316	0.357	101.40	32301.72	8240.40	2337.8638	35263.41	0.81	2.17	-0.592	22.80
-0.60	0.251	0.284	113.60	40542.12	5529.97	1246.2214	37601.27	0.87	2.43	-0.692	23.05
-0.70	0.200	0.225	121.10	46072.11	5239.01	937.8218	38847.49	0.90	2.59	-0.792	23.30
-0.80	0.158	0.179	127.80	51311.12	4763.04	677.2593	39785.31	0.92	2.73	-0.892	23.55
-0.90	0.126	0.142	133.60	56074.16	4275.71	482.9248	40462.57	0.93	2.86	-0.992	23.80
-1.00	0.103	0.113	138.60	60349.86	1137.42	102.0452	40945.49	0.95	2.96	-1.092	24.05
-1.10	0.079	0.093	139.90	61487.28	4291.66	305.8423	41047.53	0.95	2.99	-1.192	24.30
-1.20	0.063	0.071	144.70	65778.94	4761.19	229.8932	41353.37	0.96	3.09	-1.292	24.55
-1.30	0.050	0.057	149.10	69840.12	4859.50	218.5070	41583.26	0.96	3.19	-1.392	24.80
-1.40	0.040	0.045	154.20	74649.62	6027.00	215.2661	41801.77	0.97	3.30	-1.492	25.05
-1.50	0.032	0.036	160.30	80726.62	7730.62	219.3258	42017.03	0.97	3.43	-1.592	25.30
-1.60	0.025	0.028	167.80	88457.25	4916.37	110.7952	42236.36	0.98	3.59	-1.692	25.55
-1.70	0.020	0.023	172.40	93373.62	6724.56	120.3761	42347.15	0.98	3.68	-1.792	25.80
-1.80	0.016	0.018	178.50	100078.19	9760.12	138.7817	42467.52	0.98	3.81	-1.892	26.05
-1.90	0.013	0.014	187.00	109858.31	11692.69	132.0662	42606.30	0.98	4.00	-1.992	26.30
-2.00	0.010	0.011	196.70	121551.00			42738.37	0.99	4.20	-2.092	26.55
-∞							43208.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4254

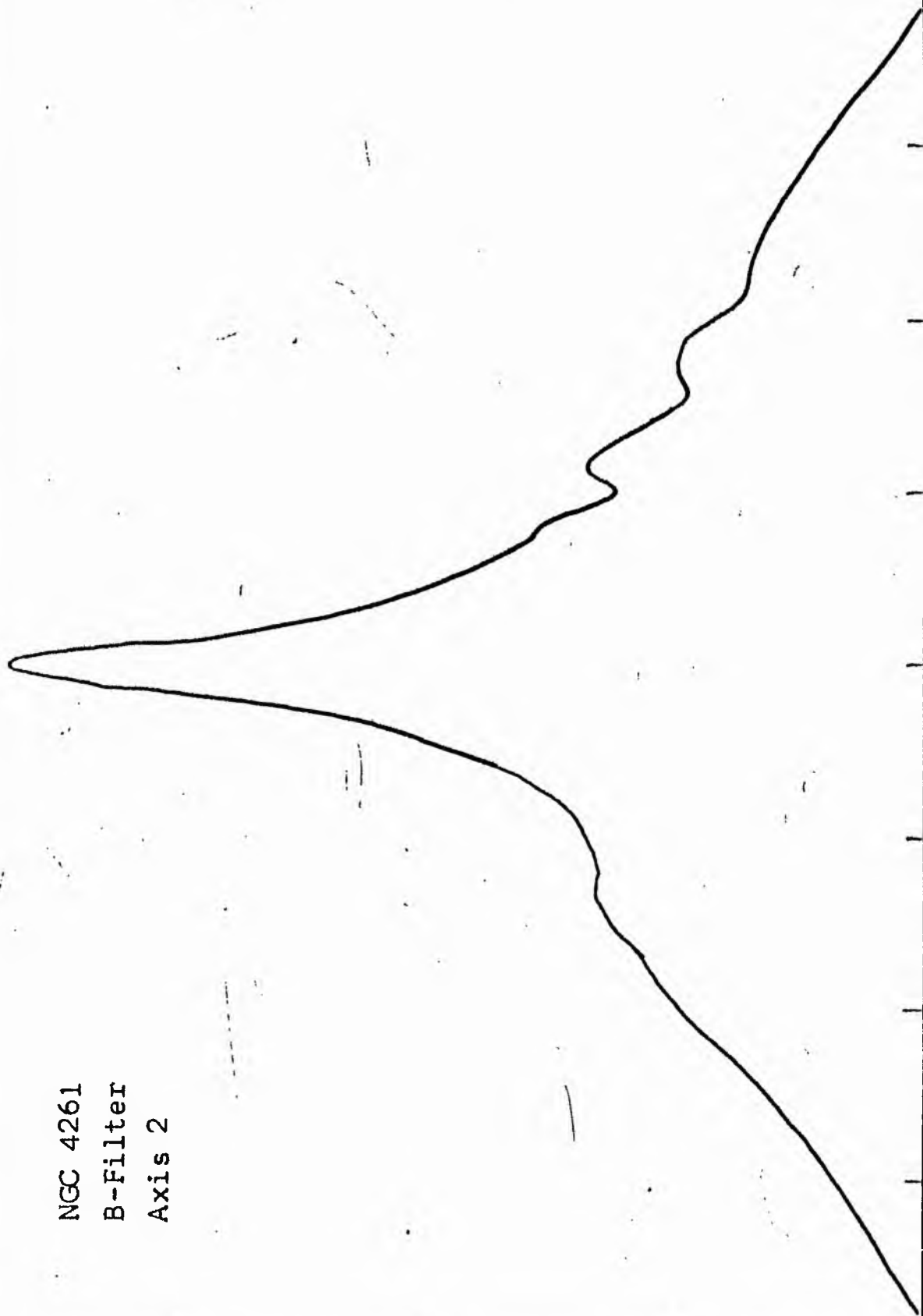
V-FILTER

Total luminosity	L_T	= 12.01
Total apparent magnitude	m_T	= 9.96
Apparent central surface brightness	μ_o	= 18.00
Major axis at threshold	$2a_m$	= 7.02
Minor axis at threshold	$2b_m$	= 6.26
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.96
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.96
Gradient of exponential component	$G(a)$	= -0.89
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.68
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.71
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 0.88
Equivalent radius	r_1^*	= 0.34
Surface brightness	μ_1	= 19.99
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.73
Axis ratio	b/a	= 0.95
Equivalent radius	r_e^*	= 0.78
Surface brightness	μ_e	= 21.30
Mean surface brightness	μ_e'	= 11.41
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.42
Axis ratio	b/a	= 1.12
Equivalent radius	r_3^*	= 1.47
Surface brightness	μ_3	= 22.52
Concentration indices	$\{C_{21}$	= 2.32
	C_{32}	= 1.89

NGC 4261
B-Filter
Axis 1

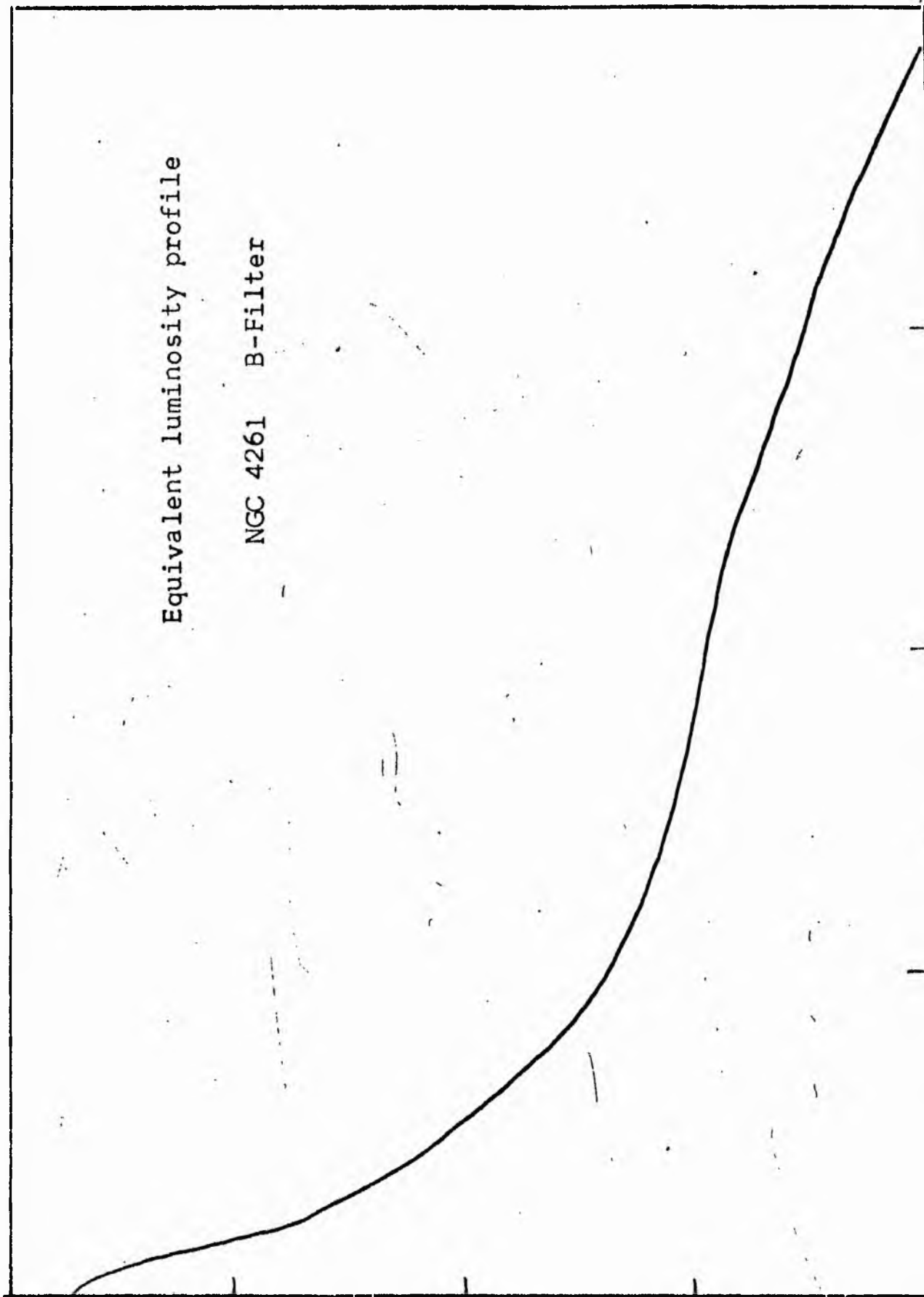


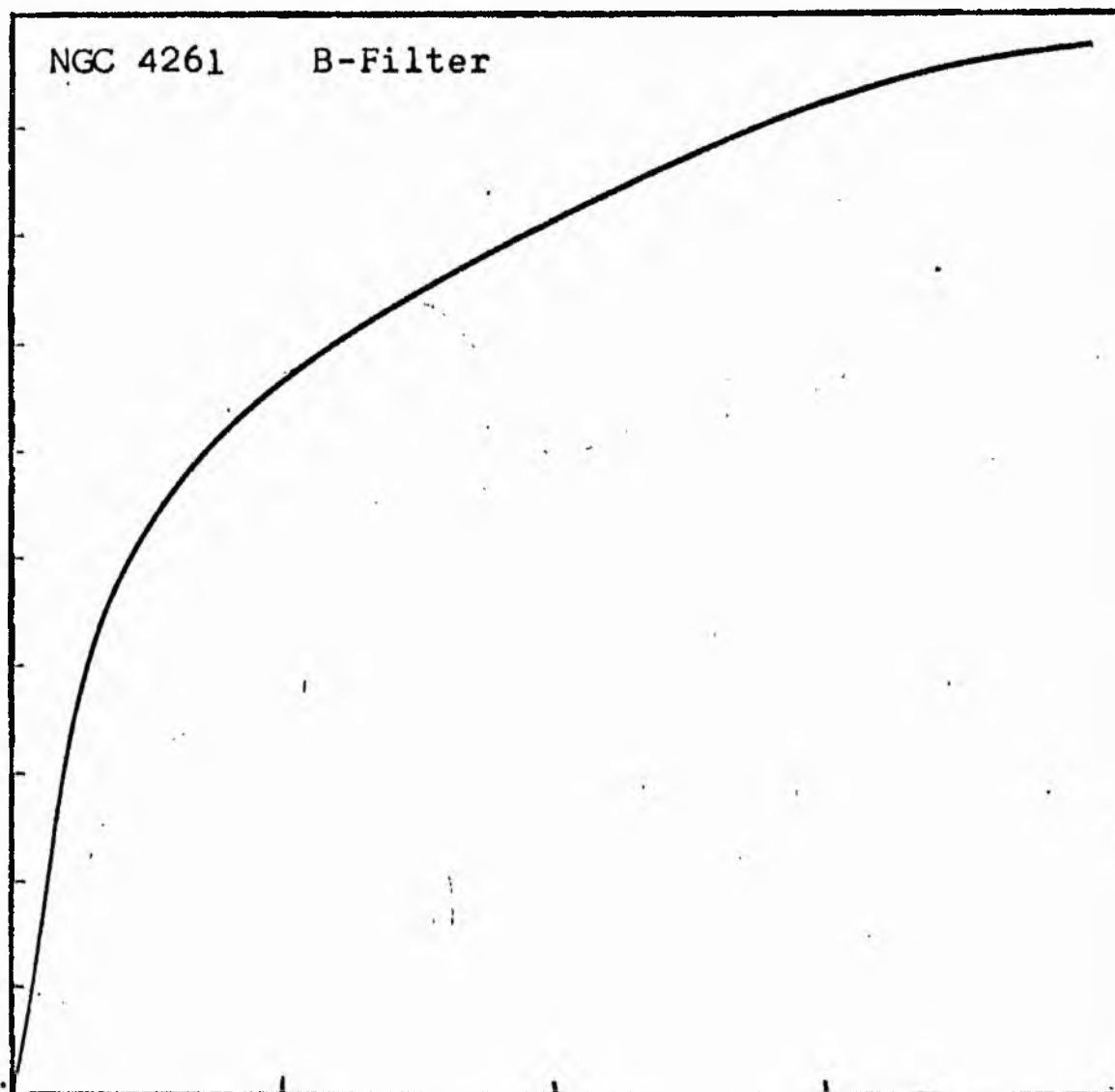
NGC 4261
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4261 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4261
B COLOUR

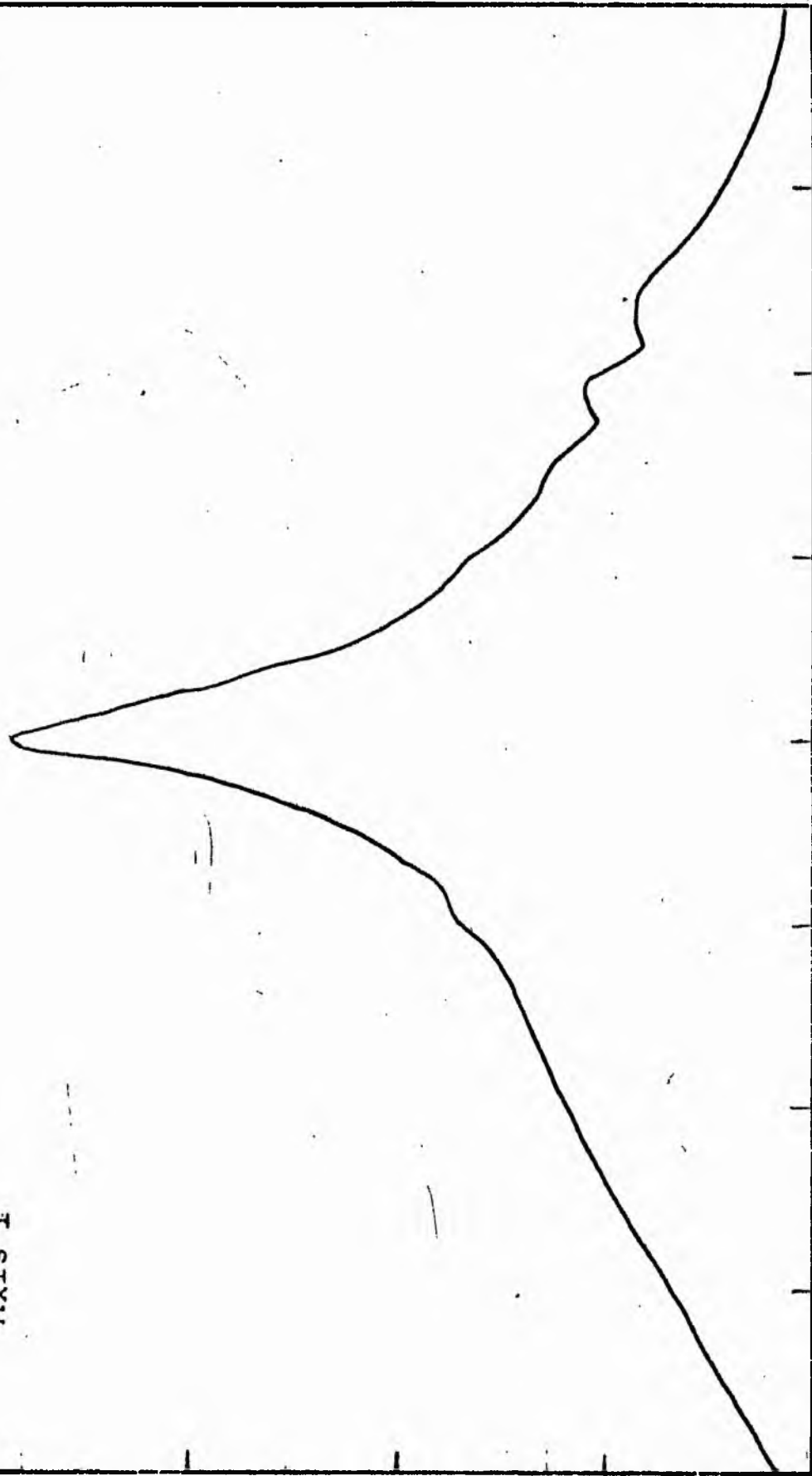
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.88	47.863		0.0	0.0	36.32	1592.0120	0.0	0.0	0.0	1.509	18.01
1.60	39.811	43.837	3.40	36.32	19.10	682.2185	1592.01	0.08	0.15	1.429	18.21
1.50	31.623	35.717	4.20	55.42	36.19	1026.7700	2274.23	0.11	0.19	1.329	18.46
1.40	25.119	28.371	5.40	91.61	29.15	657.0039	3301.00	0.16	0.24	1.229	18.71
1.30	19.953	22.536	6.20	120.76	42.10	753.5723	3958.00	0.20	0.28	1.129	18.96
1.20	15.849	17.901	7.20	162.86	53.56	761.6287	4711.57	0.23	0.32	1.029	19.21
1.10	12.589	14.219	8.30	216.42	49.48	558.8547	5473.20	0.27	0.37	0.929	19.46
1.00	10.000	11.295	9.20	265.90	42.00	376.8325	6032.05	0.30	0.41	0.829	19.71
0.90	7.943	8.972	9.90	307.91	58.53	417.0935	6408.88	0.32	0.45	0.729	19.96
0.80	6.310	7.126	10.80	366.44	71.00	401.9089	6825.97	0.34	0.49	0.629	20.21
0.70	5.012	5.661	11.80	437.44	109.96	494.4102	7227.88	0.36	0.53	0.529	20.46
0.60	3.981	4.496	13.20	547.39	178.44	637.3342	7722.29	0.38	0.59	0.429	20.71
0.50	3.162	3.572	15.20	725.83	203.57	577.5549	8359.62	0.41	0.68	0.329	20.96
0.40	2.512	2.837	17.20	929.41	169.17	381.2449	8937.18	0.44	0.77	0.229	21.21
0.30	1.995	2.254	18.70	1098.58	340.14	608.8745	9318.42	0.46	0.84	0.129	21.46
0.20	1.585	1.790	21.40	1438.72	416.35	592.0144	9927.29	0.49	0.96	0.029	21.71
0.10	1.259	1.422	24.30	1855.08	401.34	453.2942	10519.30	0.52	1.09	-0.071	21.96
-0.00	1.000	1.129	26.80	2256.42	552.20	495.4102	10972.59	0.54	1.21	-0.171	22.21
-0.10	0.794	0.897	29.90	2808.61	571.23	407.0840	11468.00	0.57	1.35	-0.271	22.46
-0.20	0.631	0.713	32.80	3379.85	737.02	417.2039	11875.09	0.59	1.48	-0.371	22.71
-0.30	0.501	0.566	36.20	4116.87	395.75	177.9453	12292.29	0.61	1.63	-0.471	22.96
-0.40	0.398	0.450	37.90	4512.61	1215.42	434.1057	12470.23	0.62	1.71	-0.571	23.21
-0.50	0.316	0.357	42.70	5728.03	1661.78	471.4578	12934.34	0.64	1.92	-0.671	23.46
-0.60	0.251	0.284	48.50	7389.81	2252.24	507.5566	13375.80	0.66	2.18	-0.771	23.71
-0.70	0.200	0.225	55.40	9642.05	2945.96	527.3484	13883.35	0.69	2.49	-0.871	23.96
-0.80	0.158	0.179	63.30	12588.01	4708.46	669.4973	14410.70	0.71	2.85	-0.971	24.21
-0.90	0.126	0.142	74.20	17296.47	7812.26	882.3625	15080.20	0.75	3.34	-1.071	24.46
-1.00	0.100	0.113	89.40	25108.73	9791.72	878.4763	15962.55	0.79	4.03	-1.171	24.71
-1.10	0.079	0.090	105.40	34900.45	10564.96	752.9031	16841.03	0.83	4.75	-1.271	24.96
-1.20	0.063	0.071	120.30	45465.40	8777.21	496.8538	17543.93	0.87	5.42	-1.371	25.21
-1.30	0.050	0.057	131.40	54242.61	9551.36	429.4751	18090.78	0.90	5.92	-1.471	25.46
-1.40	0.040	0.045	142.50	63793.97	10325.47	368.7935	18520.25	0.92	6.42	-1.571	25.71
-1.50	0.032	0.036	153.60	74119.44	8736.12	247.8524	18889.05	0.94	6.92	-1.671	25.96
-1.60	0.025	0.028	162.40	82855.56	8792.94	198.1567	19136.90	0.95	7.31	-1.771	26.21
-1.70	0.020	0.023	170.80	91648.50	8898.81	159.2969	19335.05	0.96	7.69	-1.871	26.46
-1.80	0.016	0.018	178.90	100547.31	8607.19	122.3875	19494.35	0.97	8.06	-1.971	26.71
-1.90	0.013	0.014	186.40	109154.50	10549.69	119.1560	19616.74	0.97	8.39	-2.071	26.96
-2.00	0.010	0.011	195.20	119704.19			19735.89	0.98	8.79	-2.171	27.21
-∞							20195.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4261

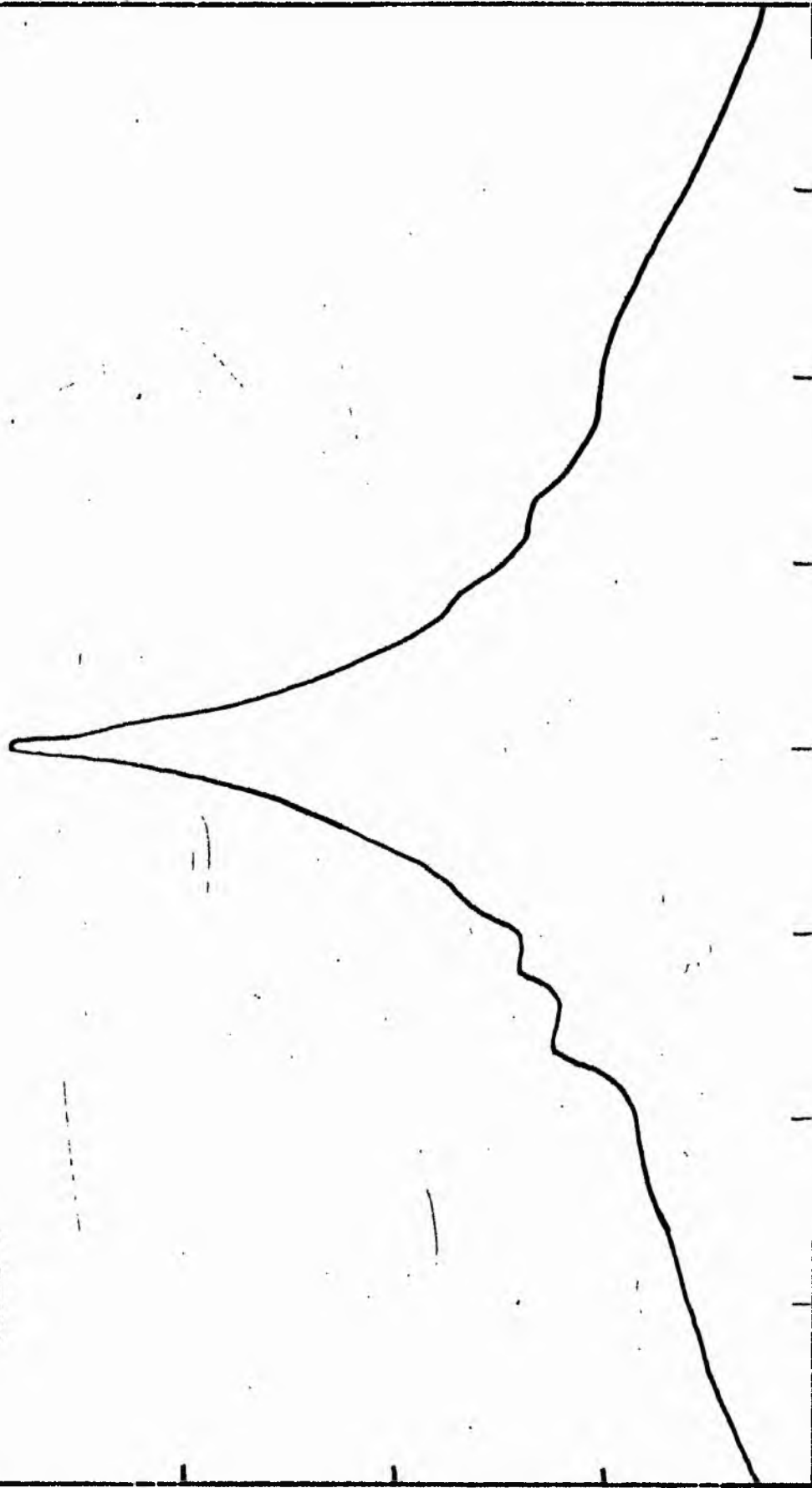
B-FILTER

Total luminosity	L_T	= 5.61
Total apparent magnitude	m_T	= 11.45
Apparent central surface brightness	μ_0	= 18.01
Major axis at threshold	$2a_m$	= 6.80
Minor axis at threshold	$2b_m$	= 6.32
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 4.01
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.84
Gradient of exponential component	$G(a)$	= -0.54
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.51
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.18
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.16
Axis ratio	b/a	= 1.20
Equivalent radius	r_1^*	= 0.13
Surface brightness	μ_1	= 19.33
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.36
Axis ratio	b/a	= 0.95
Equivalent radius	r_e^*	= 0.37
Surface brightness	μ_e	= 21.79
Mean surface brightness	μ_e'	= 11.30
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.23
Axis ratio	b/a	= 1.07
Equivalent radius	r_3^*	= 1.26
Surface brightness	μ_3	= 24.46
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.89 \\ = 3.39 \end{cases}$

NGC 4261
V-Filter
Axis I

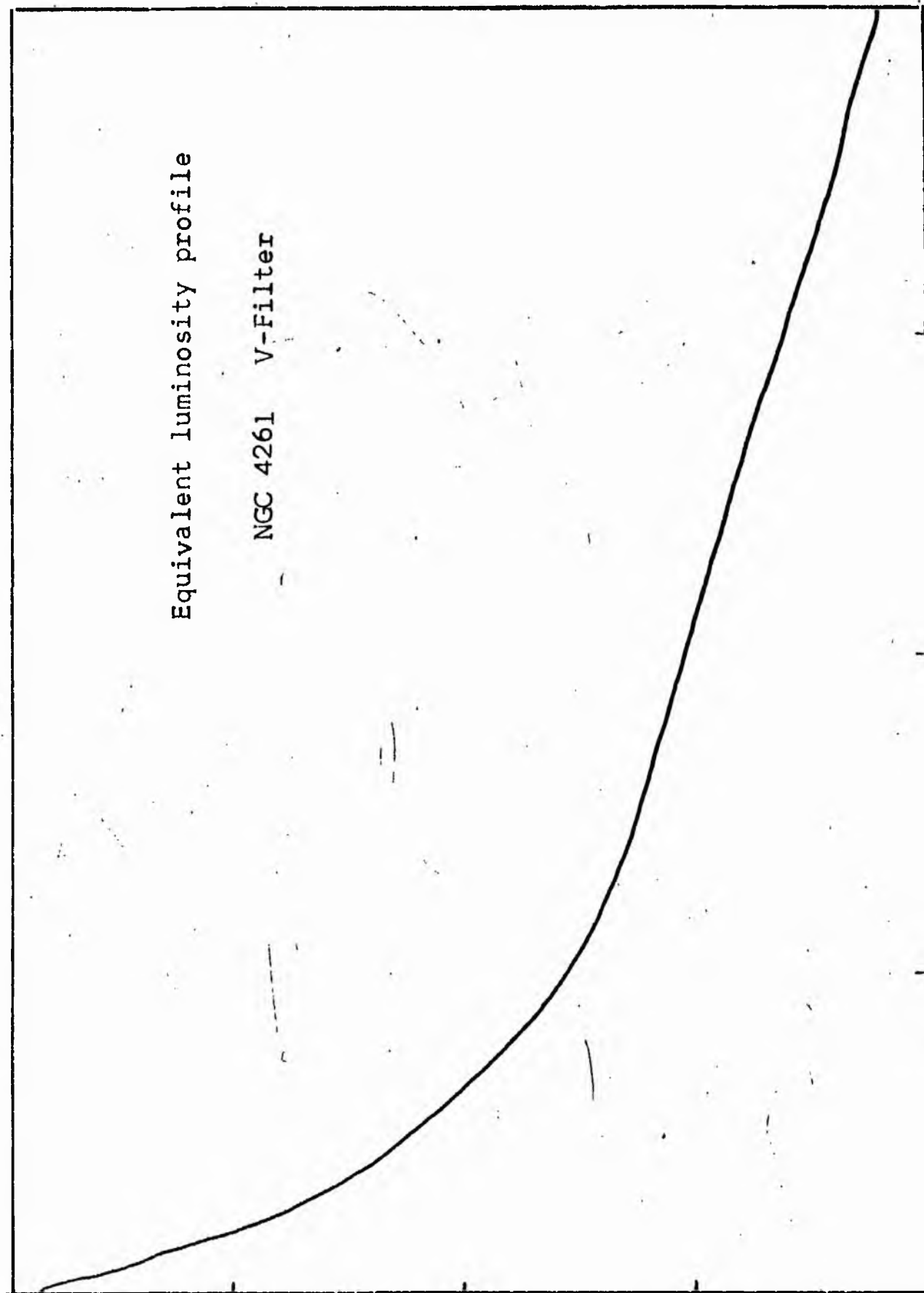


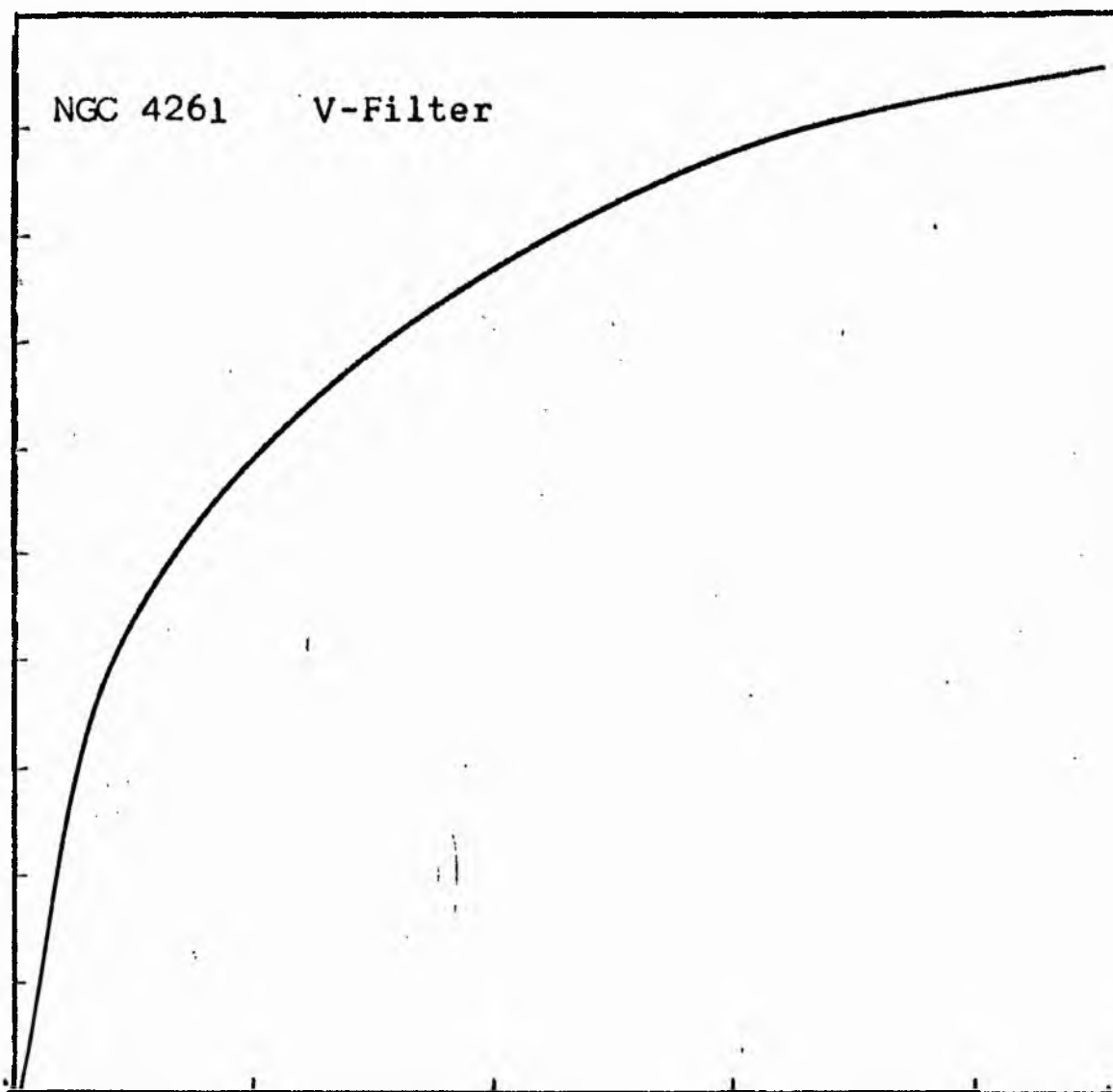
NGC 4261
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4261 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4261
V COLOUR

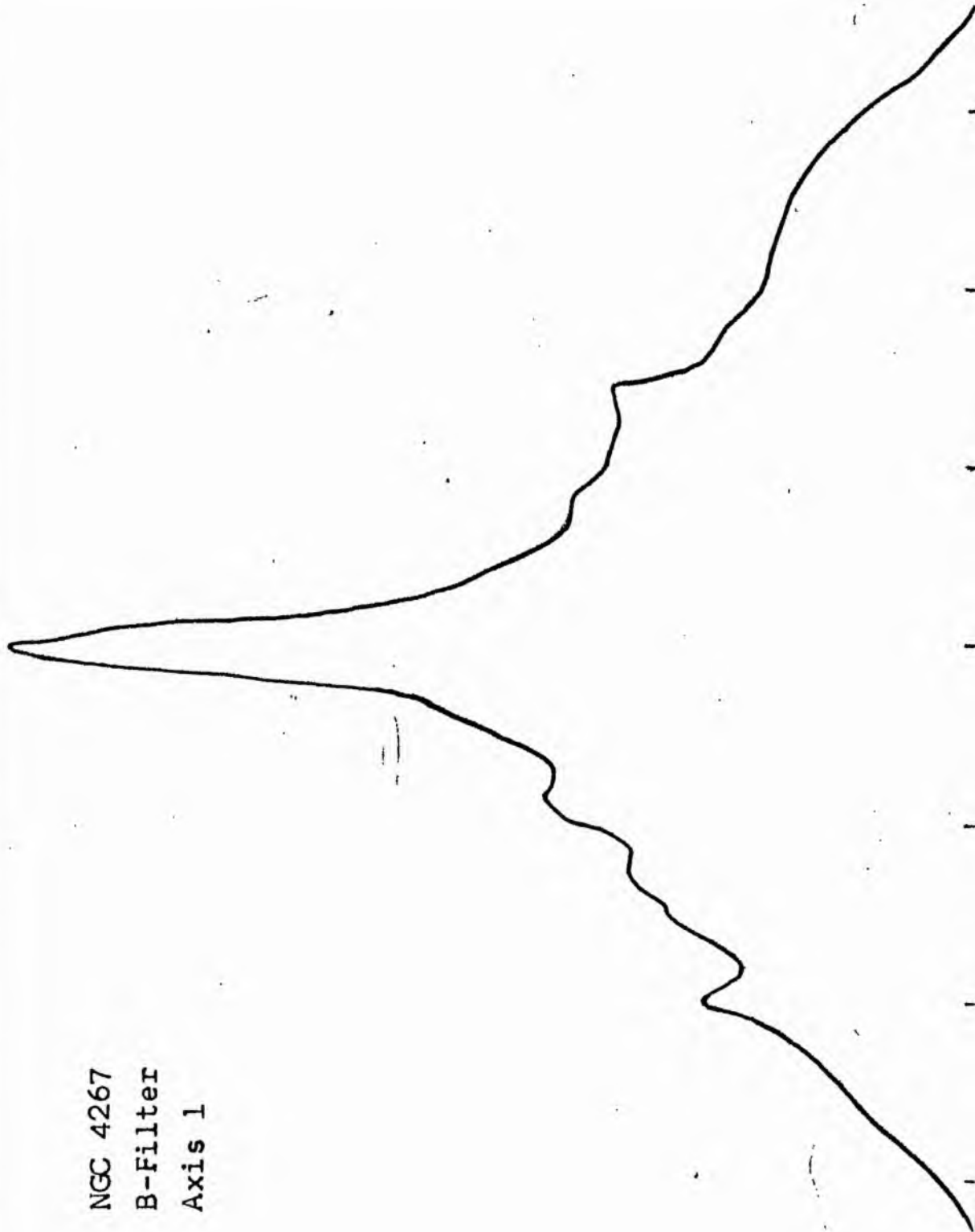
LOG I	I	T	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
1.86	72.444		0.0	0.0			0.0	0.0	0.0	1.856	16.76
1.80	63.096		0.30	0.28			19.16	0.00	0.01	1.796	16.91
1.70	50.119		0.90	2.54			147.20	0.01	0.03	1.696	17.16
1.60	39.811		1.40	6.16			309.65	0.01	0.04	1.596	17.41
1.50	31.623		2.10	13.85			584.56	0.02	0.07	1.496	17.66
1.40	25.119		4.70	69.40			2160.36	0.08	0.15	1.396	17.91
1.30	19.952		6.40	128.68			3476.31	0.13	0.20	1.296	18.16
1.20	15.849		7.80	191.13			4614.29	0.18	0.25	1.196	18.41
1.10	12.589		8.60	232.35			5200.36	0.20	0.27	1.096	18.66
1.00	10.000		9.50	283.53			5778.37	0.22	0.30	0.996	18.91
0.90	7.943		10.70	359.68			6461.58	0.25	0.34	0.896	19.16
0.80	6.310		12.30	475.29			7285.46	0.28	0.39	0.796	19.41
0.70	5.012		13.50	572.56			7836.04	0.30	0.43	0.696	19.66
0.60	3.981		15.15	721.07			8503.81	0.33	0.48	0.596	19.91
0.50	3.162		17.33	943.51			9298.29	0.36	0.55	0.496	20.16
0.40	2.512		19.20	1158.12			9907.14	0.38	0.61	0.396	20.41
0.30	1.995		22.28	1559.48			10811.64	0.41	0.70	0.296	20.66
0.20	1.585		25.39	2025.23			11645.36	0.45	0.80	0.196	20.91
0.10	1.259		28.51	2553.55			12396.57	0.47	0.90	0.096	21.16
-0.00	1.000		31.79	3174.90			13098.36	0.50	1.00	-0.004	21.41
-0.10	0.794		35.18	3888.14			13738.24	0.53	1.11	-0.104	21.66
-0.20	0.631		36.74	4240.61			13989.43	0.54	1.16	-0.204	21.91
-0.30	0.501		41.19	5330.07			14606.14	0.56	1.30	-0.304	22.16
-0.40	0.398		46.63	6830.94			15280.99	0.58	1.47	-0.404	22.41
-0.50	0.316		53.05	8841.39			15999.05	0.61	1.68	-0.504	22.66
-0.60	0.251		60.45	11480.01			16747.64	0.64	1.91	-0.604	22.91
-0.70	0.200		68.84	14887.82			17515.61	0.67	2.17	-0.704	23.16
-0.80	0.158		82.13	21191.09			18643.94	0.71	2.59	-0.804	23.41
-0.90	0.126		99.54	31127.55			20056.80	0.77	3.14	-0.904	23.66
-1.00	0.100		111.21	38854.15			20929.48	0.80	3.51	-1.004	23.91
-1.10	0.079		123.29	47753.54			21727.90	0.83	3.89	-1.104	24.16
-1.20	0.063		134.36	56713.94			22366.45	0.86	4.24	-1.204	24.41
-1.30	0.050		144.26	65379.52			22856.98	0.87	4.56	-1.304	24.66
-1.40	0.040		155.80	76257.81			23346.12	0.89	4.92	-1.404	24.91
-1.50	0.032		169.11	89843.81			23831.37	0.91	5.34	-1.504	25.16
-1.60	0.025		179.80	101501.50			24163.80	0.92	5.68	-1.604	25.41
-1.70	0.020		188.79	111971.56			24398.40	0.93	5.96	-1.704	25.66

PHOTOMETRIC PARAMETERS OF NGC 4261

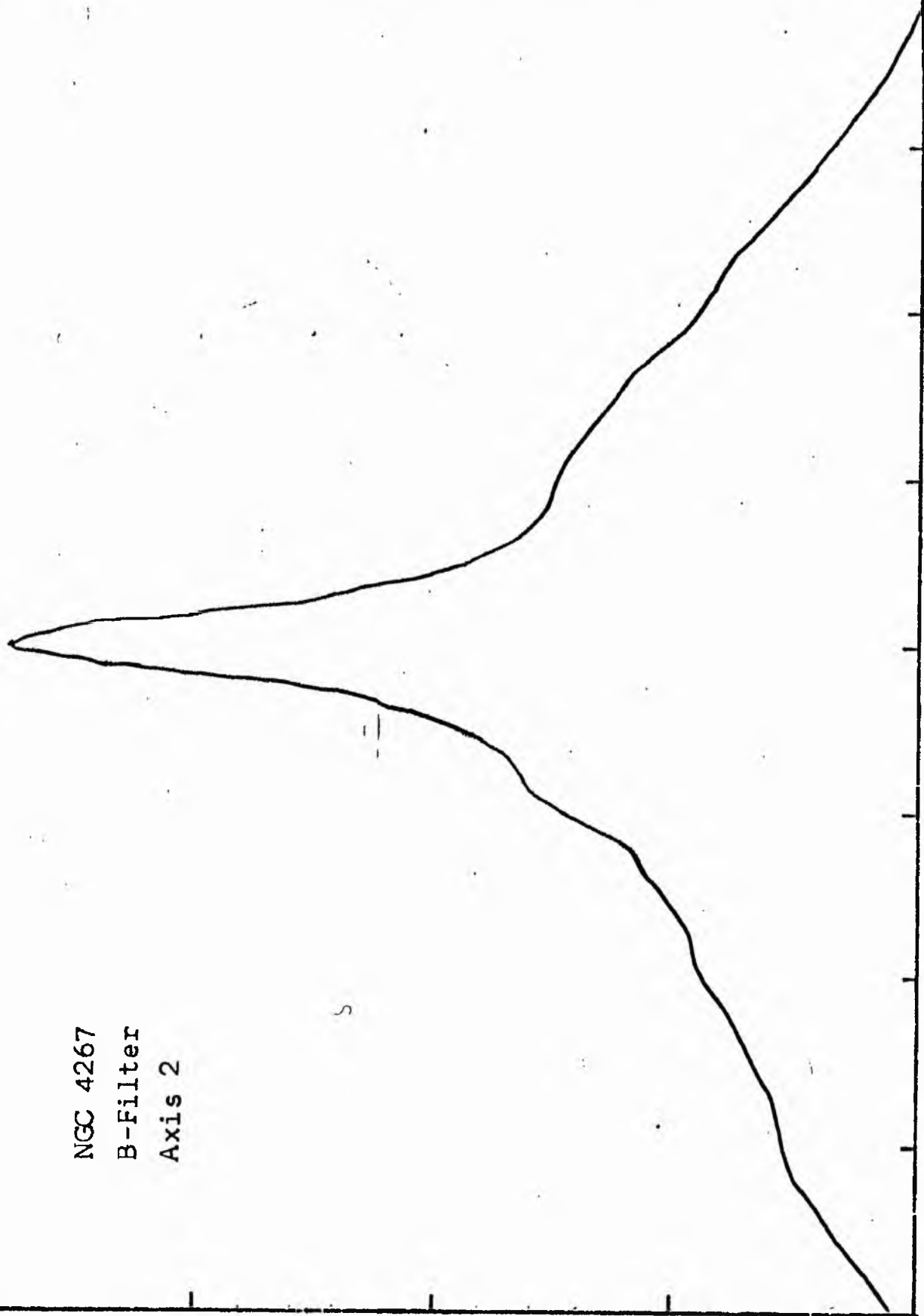
V-FILTER

Total luminosity	L_T	= 7.26
Total apparent magnitude	m_T	= 10.37
Apparent central surface brightness	μ_o	= 16.76
Major axis at threshold	$2a_m$	= 7.3
Minor axis at threshold	$2b_m$	= 7.22
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.2
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.90
Gradient of exponential component	$G(a)$	= -0.53
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.53
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.27
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.21
Axis ratio	b/a	= 0.96
Equivalent radius	r_1^*	= 0.18
Surface brightness	μ_1	= 19.16
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.53
Axis ratio	b/a	= 0.87
Equivalent radius	r_e^*	= 0.53
Surface brightness	μ_e	= 21.41
Mean surface brightness	μ_e'	= 10.87
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.59
Axis ratio	b/a	= 0.96
Equivalent radius	r_3^*	= 1.56
Surface brightness	μ_3	= 23.58
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.92 \\ = 2.96 \end{cases}$

NGC 4267
B-Filter
Axis 1

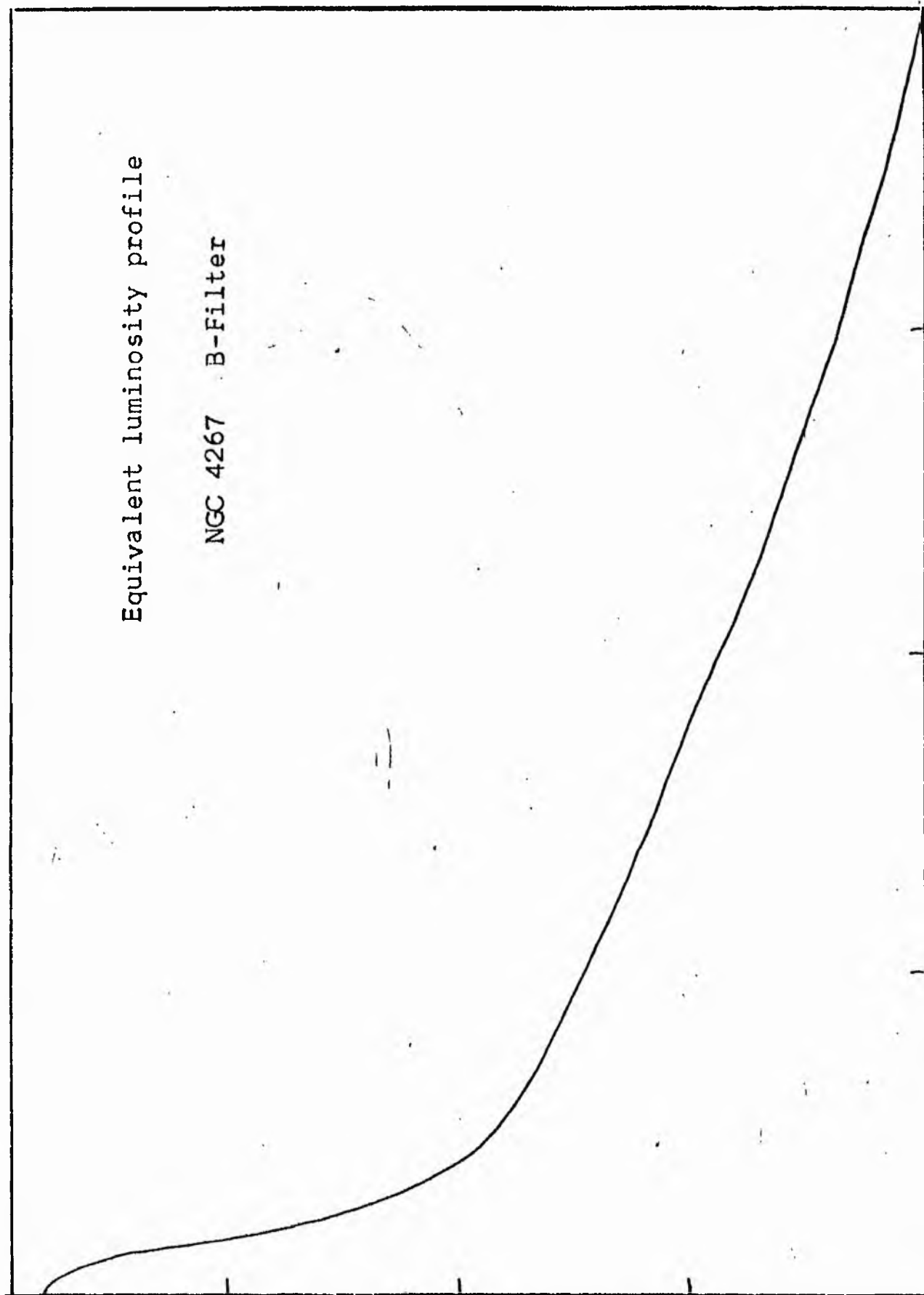


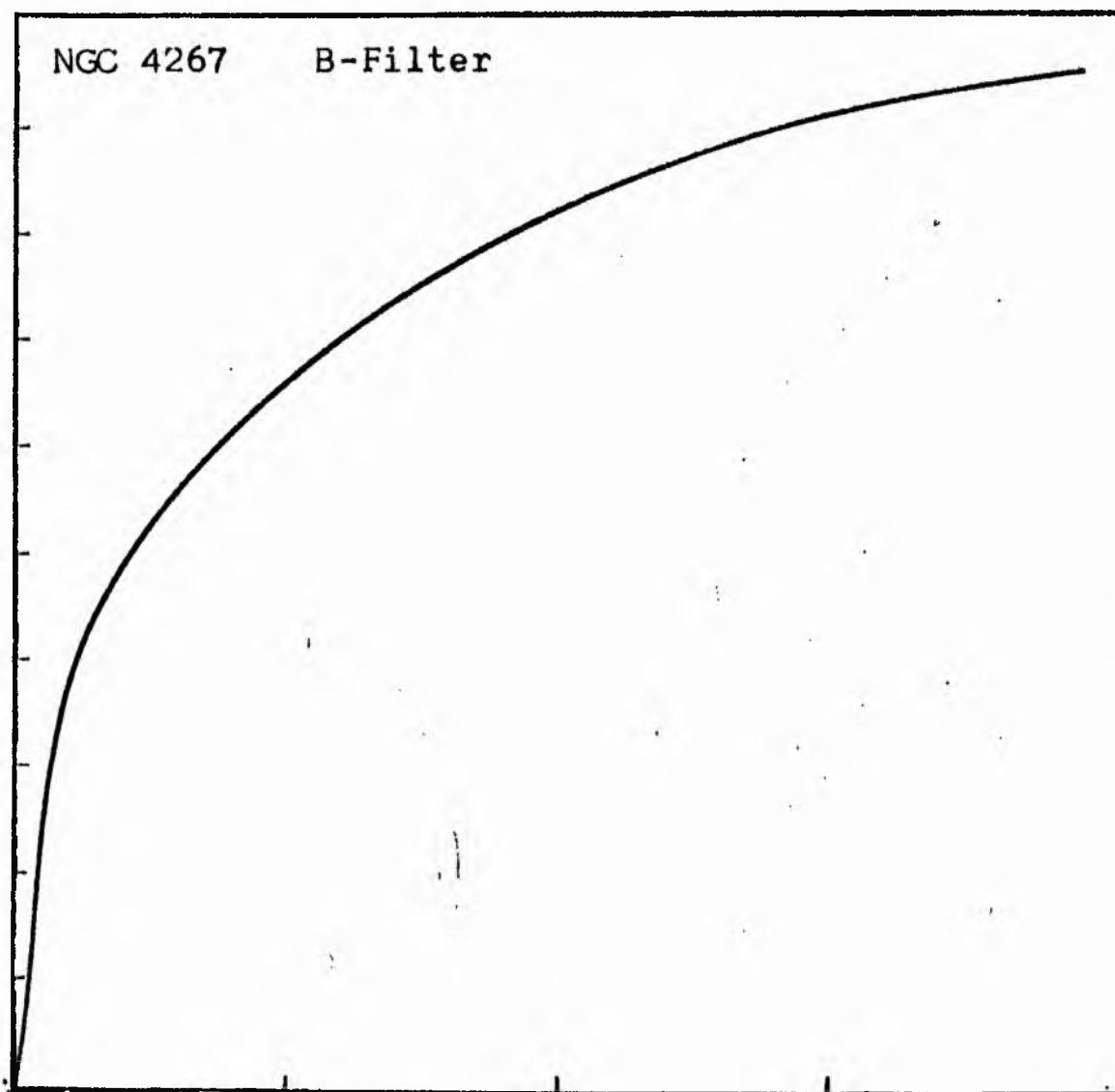
NGC 4267
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4267 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4267
B COLOUR

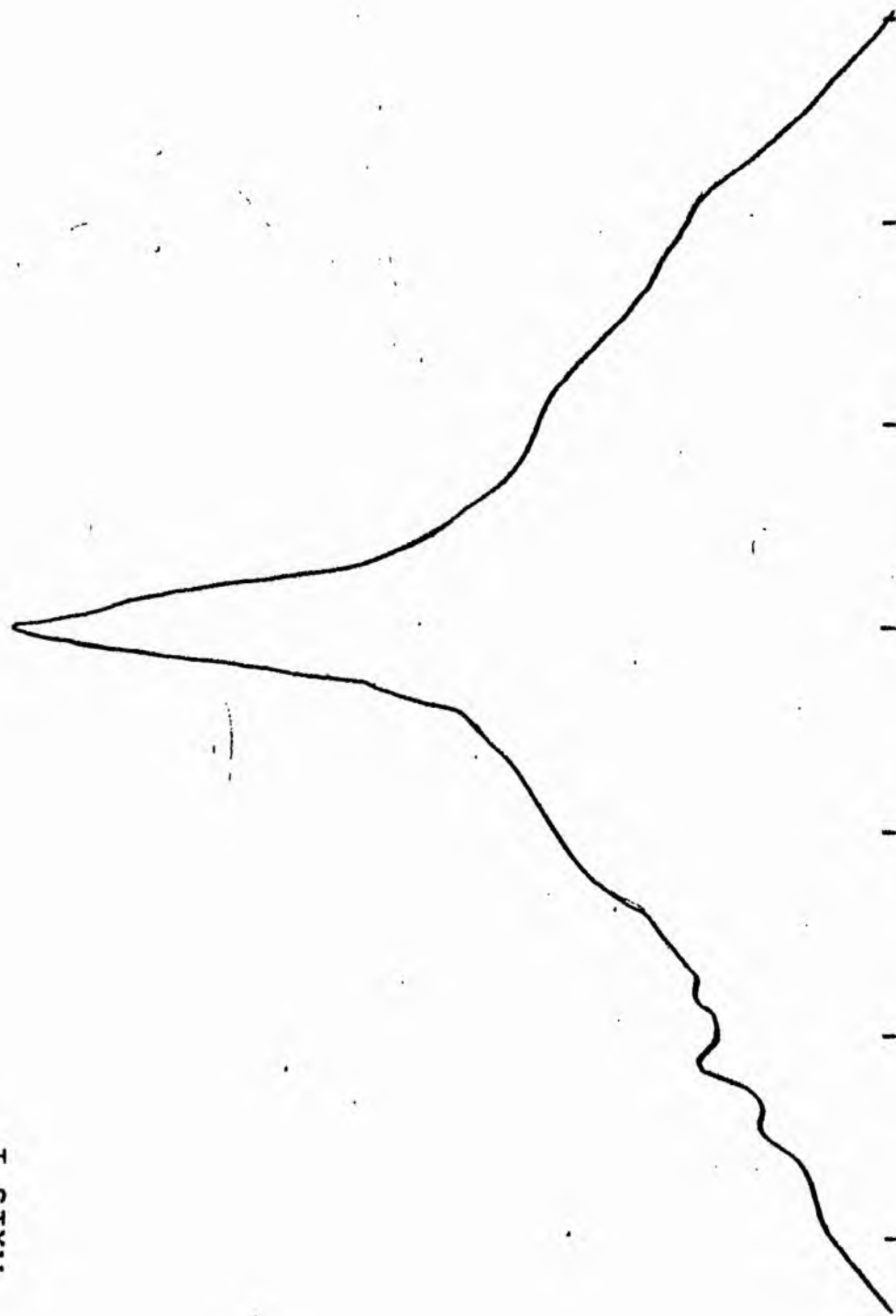
LOG I	I	I	R	AREA	ΔA	P	ΣP	KIRI	ρ	LOG J	μ
1.84	69.183		0.0	0.0			0.0	0.0	0.0	1.854	17.92
1.80	63.096	66.139	2.74	23.59	23.59	1559.9487	1559.95	0.07	0.13	1.814	18.02
1.70	50.119	56.607	3.92	48.27	24.69	1397.5798	2957.53	0.14	0.18	1.714	18.27
1.60	39.811	44.965	5.02	79.17	30.89	1309.1543	4346.68	0.21	0.23	1.614	18.52
1.50	31.623	35.717	5.65	100.29	21.12	754.2656	5100.95	0.24	0.26	1.514	18.77
1.40	25.119	28.371	6.22	121.54	21.26	603.0383	5703.98	0.27	0.29	1.414	19.02
1.30	19.952	22.536	6.49	132.32	10.78	242.9564	5946.93	0.29	0.30	1.314	19.27
1.20	15.849	17.901	7.56	179.55	47.23	845.4319	6792.36	0.33	0.35	1.214	19.52
1.10	12.589	14.219	7.80	191.13	11.58	164.6723	6957.04	0.33	0.36	1.114	19.77
1.00	10.000	11.295	8.70	237.79	46.65	526.9197	7483.95	0.36	0.40	1.014	20.02
0.90	7.943	8.972	9.30	271.72	33.93	304.3955	7788.35	0.37	0.43	0.914	20.27
0.80	6.310	7.126	9.70	295.59	23.88	170.1507	7958.50	0.38	0.45	0.814	20.52
0.70	5.012	5.661	10.00	366.44	70.84	401.0193	8359.51	0.40	0.50	0.714	20.77
0.60	3.981	4.496	11.50	415.48	49.04	220.5064	8580.02	0.41	0.53	0.614	21.02
0.50	3.162	3.572	12.50	490.87	75.40	269.2957	8849.31	0.42	0.58	0.514	21.27
0.40	2.512	2.837	13.90	606.99	116.11	329.4204	9178.73	0.44	0.64	0.414	21.52
0.30	1.995	2.254	15.10	716.31	109.33	246.3756	9425.10	0.45	0.70	0.314	21.77
0.20	1.585	1.790	16.66	871.97	155.65	278.6250	9703.73	0.47	0.77	0.214	22.02
0.10	1.259	1.422	18.24	1045.20	173.23	246.3200	9950.04	0.48	0.84	0.114	22.27
-0.00	1.000	1.129	21.17	1407.96	362.77	409.7266	10359.77	0.50	0.98	0.014	22.52
-0.10	0.794	0.897	25.25	2002.96	595.00	533.8069	10893.57	0.52	1.17	-0.086	22.77
-0.20	0.631	0.713	30.90	2999.62	996.66	710.2583	11603.83	0.56	1.43	-0.186	23.02
-0.30	0.501	0.566	36.70	4231.38	1231.76	697.2578	12301.09	0.59	1.70	-0.286	23.27
-0.40	0.398	0.450	43.31	5892.86	1661.48	747.0750	13048.16	0.63	2.00	-0.386	23.52
-0.50	0.316	0.357	50.57	8034.07	2141.21	764.7659	13812.93	0.66	2.34	-0.486	23.77
-0.60	0.251	0.284	58.70	10824.95	2790.88	791.7913	14604.72	0.70	2.71	-0.586	24.02
-0.70	0.200	0.225	66.38	13842.00	3017.85	680.0901	15284.81	0.73	3.07	-0.686	24.27
-0.80	0.158	0.179	74.20	17296.47	3453.67	618.2295	15903.04	0.76	3.43	-0.786	24.52
-0.90	0.126	0.142	82.06	21154.98	3858.51	548.6418	16451.68	0.79	3.79	-0.886	24.77
-1.00	0.100	0.113	88.10	24383.80	3228.82	364.6816	16816.36	0.81	4.07	-0.986	25.02
-1.10	0.079	0.090	96.70	29376.68	4992.87	447.9409	17264.29	0.83	4.47	-1.086	25.27
-1.20	0.063	0.071	106.80	35833.74	6457.06	460.1565	17724.45	0.85	4.94	-1.186	25.52
-1.30	0.050	0.057	118.40	44040.59	8206.85	464.5662	18189.01	0.87	5.47	-1.286	25.77
-1.40	0.040	0.045	129.30	52522.67	8482.08	381.3945	18570.41	0.89	5.98	-1.386	26.02
-1.50	0.032	0.036	140.40	61927.57	9404.90	335.9128	18906.32	0.91	6.49	-1.486	26.27
-1.60	0.025	0.028	151.70	72297.06	10369.49	294.1917	19200.51	0.92	7.01	-1.586	26.52
-1.70	0.020	0.023	164.80	85322.62	13025.56	293.5417	19494.05	0.94	7.62	-1.686	26.77
-1.80	0.016	0.018	175.80	97092.87	11770.25	210.6978	19704.74	0.95	8.13	-1.786	27.02
-1.90	0.013	0.014	184.70	107172.56	10079.69	143.3250	19848.07	0.95	8.54	-1.886	27.27
-2.00	0.010	0.011	191.90	115691.06	8518.50	96.2141	19944.28	0.96	8.87	-1.986	27.52
-∞							20844.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4267

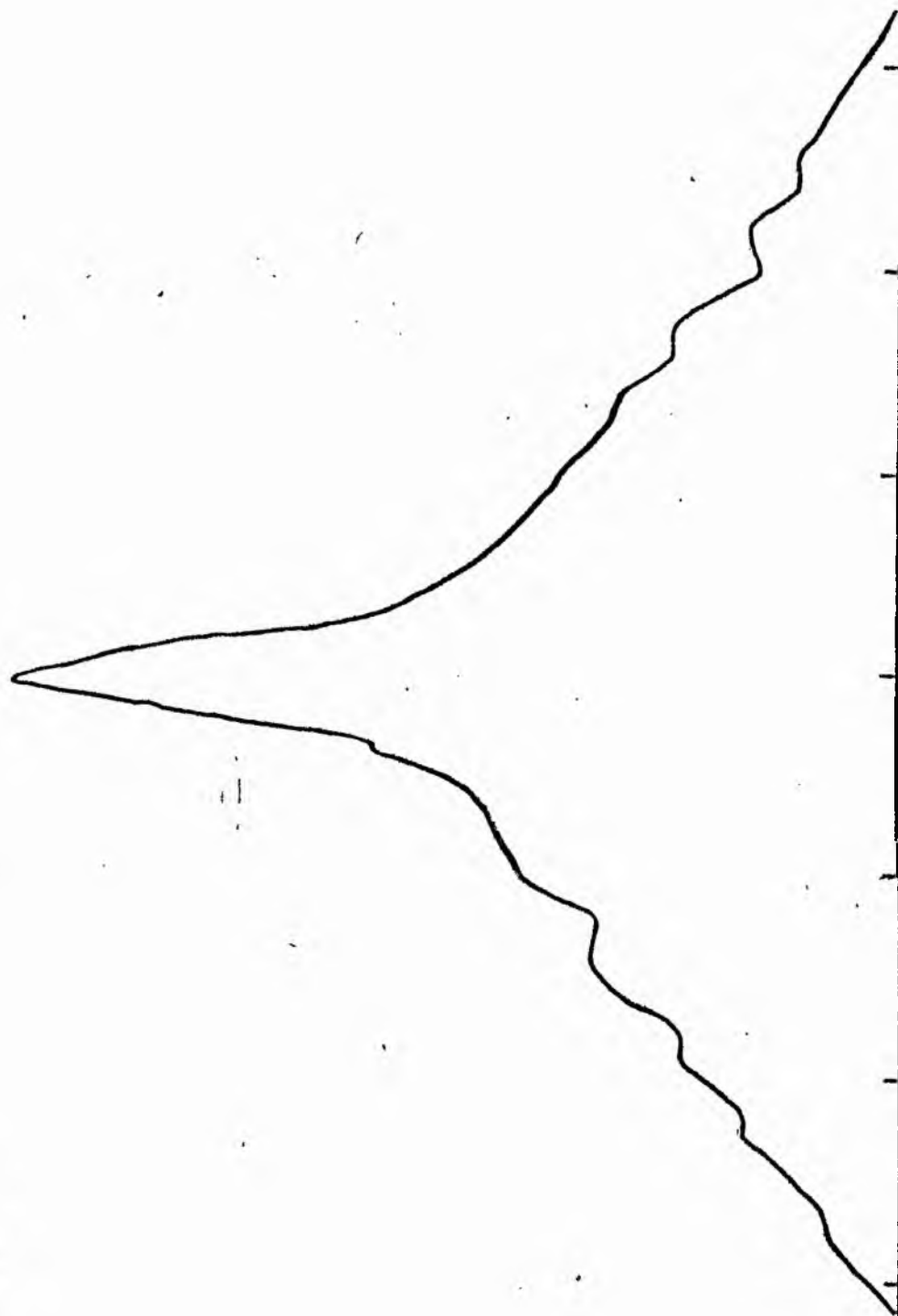
B-FILTER

Total luminosity	L_T	= 5.79
Total apparent magnitude	m_T	= 11.72
Apparent central surface brightness	μ_0	= 17.92
Major axis at threshold	$2a_m$	= 5.93
Minor axis at threshold	$2b_m$	= 6.04
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 2.98
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.81
Gradient of exponential component	$G(a)$	= -0.68
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.59
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.24
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.13
Axis ratio	b/a	= 0.96
Equivalent radius	r_1^*	= 0.09
Surface brightness	μ_1	= 18.94
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.32
Axis ratio	b/a	= 1.12
Equivalent radius	r_e^*	= 0.36
Surface brightness	μ_e	= 22.52
Mean surface brightness	μ_e'	= 11.52
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.27
Axis ratio	b/a	= 0.86
Equivalent radius	r_3^*	= 1.17
Surface brightness	μ_3	= 24.46
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 3.76 \\ = 3.27 \end{matrix}$

NGC 4267
V-Filter
Axis 1

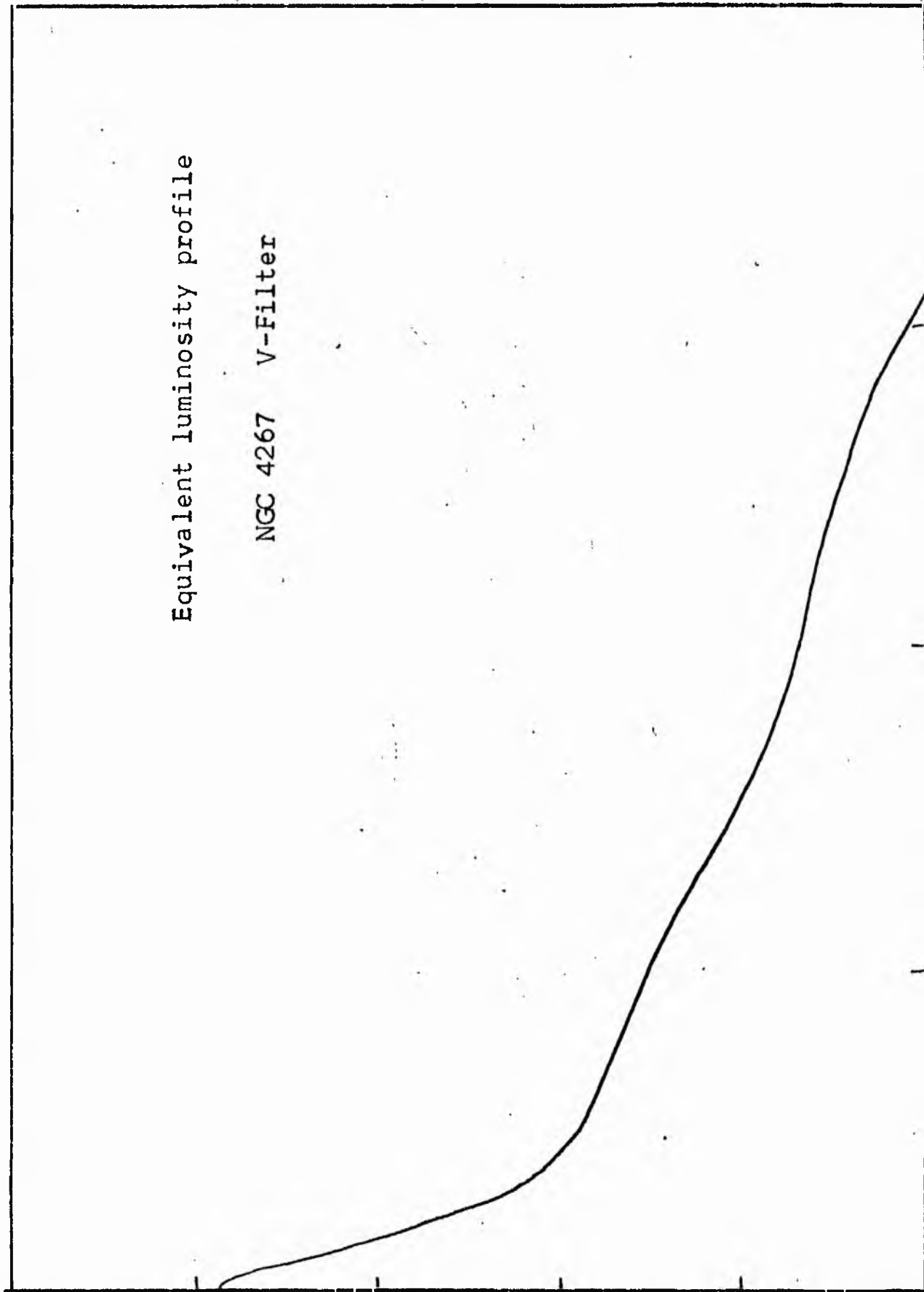


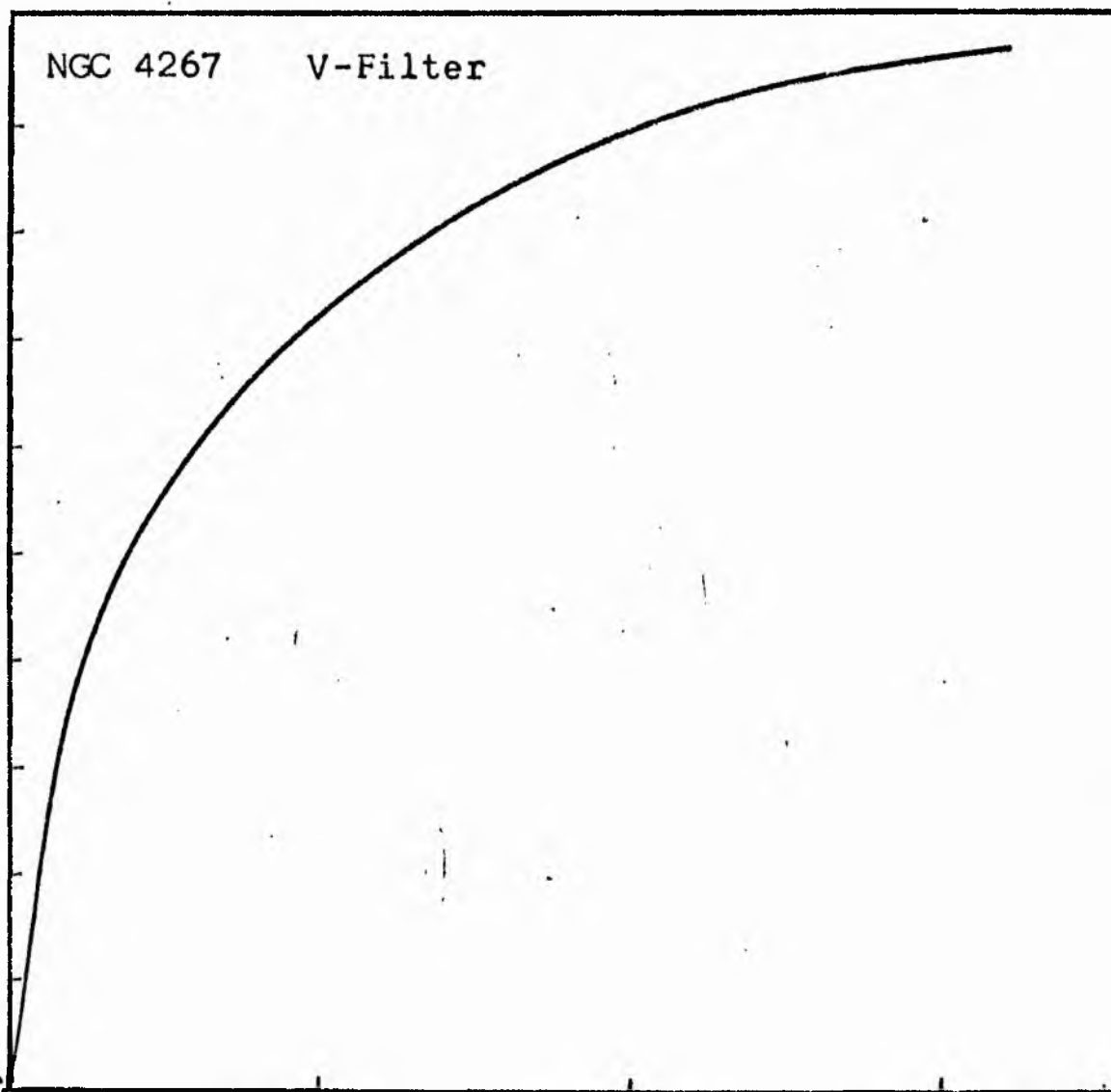
NGC 4267
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4267 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4267
V COLOUR

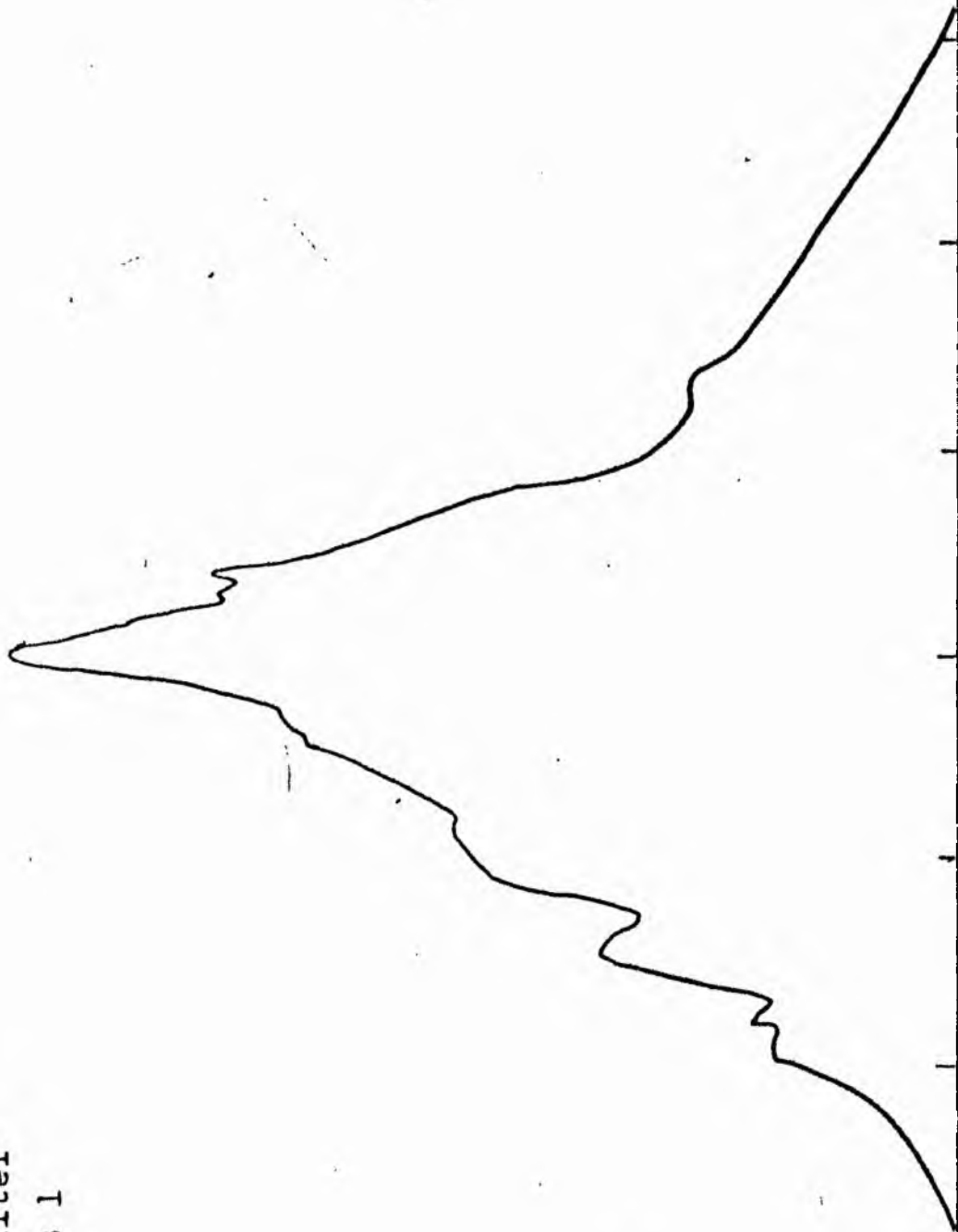
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.81	64.565		0.0	0.0			0.0	0.0	0.0	1.782	16.91
		63.830			1.54	98.2594					
1.80	63.096		0.70	1.54	12.32	697.1182	98.26	0.01	0.04	1.772	16.94
		56.607									
1.70	50.119		2.10	13.85	22.46	1010.0115	795.38	0.05	0.11	1.672	17.19
		44.965									
1.60	39.811		3.40	36.32	19.10	682.2170	1805.39	0.11	0.17	1.572	17.44
		35.717									
1.50	31.623		4.20	55.42	23.12	655.9910	2487.61	0.15	0.21	1.472	17.69
		28.371									
1.40	25.119		5.00	78.54	23.53	530.2739	3143.60	0.19	0.25	1.372	17.94
		22.536									
1.30	19.952		5.70	102.07	34.78	622.5383	3673.87	0.23	0.29	1.272	18.19
		17.901									
1.20	15.849		6.60	136.85	44.61	634.3171	4296.41	0.26	0.34	1.172	18.44
		14.219									
1.10	12.589		7.60	181.46	56.33	636.2068	4930.72	0.30	0.39	1.072	18.69
		11.295									
1.00	10.000		8.70	237.79	28.70	257.4438	5566.93	0.34	0.44	0.972	18.94
		8.972									
0.90	7.943		9.21	266.48	29.11	207.4459	5824.37	0.36	0.47	0.872	19.19
		7.126									
0.80	6.310		9.70	295.59	77.66	439.6089	6031.81	0.37	0.49	0.772	19.44
		5.661									
0.70	5.012		10.90	373.25	35.03	157.5048	6471.42	0.40	0.56	0.672	19.69
		4.496									
0.60	3.981		11.40	408.28	67.01	239.3369	6628.92	0.41	0.58	0.572	19.94
		3.572									
0.50	3.162		12.30	475.29	114.35	324.4292	6868.26	0.42	0.63	0.472	20.19
		2.837									
0.40	2.512		13.70	589.65	89.22	201.0651	7192.68	0.44	0.70	0.372	20.44
		2.254									
0.30	1.995		14.70	678.87	155.82	278.9302	7393.75	0.45	0.75	0.272	20.69
		1.790									
0.20	1.585		16.30	834.69	138.45	196.8622	7672.68	0.47	0.83	0.172	20.94
		1.422									
0.10	1.259		17.60	973.14	347.12	392.0520	7869.54	0.48	0.90	0.072	21.19
		1.129									
-0.00	1.000		20.50	1320.25	643.24	577.0886	8261.59	0.51	1.04	-0.028	21.44
		0.897									
-0.10	0.794		25.00	1963.50	1094.66	780.0928	8838.67	0.54	1.27	-0.128	21.69
		0.713									
-0.20	0.631		31.20	3058.15	1598.47	904.8450	9618.76	0.59	1.59	-0.228	21.94
		0.566									
-0.30	0.501		38.50	4656.62	2253.65	1013.3416	10523.61	0.65	1.96	-0.328	22.19
		0.450									
-0.40	0.398		46.90	6910.28	1101.57	393.4409	11536.95	0.71	2.39	-0.428	22.44
		0.357									
-0.50	0.316		50.50	8011.84	1353.74	384.0652	11930.38	0.73	2.57	-0.528	22.69
		0.284									
-0.60	0.251		54.60	9365.59	1793.85	404.2544	12314.45	0.76	2.78	-0.628	22.94
		0.225									
-0.70	0.200		59.60	11159.43	2154.66	385.6985	12718.70	0.78	3.04	-0.728	23.19
		0.179									
-0.80	0.158		65.10	13314.09	2836.48	403.3186	13104.39	0.80	3.32	-0.828	23.44
		0.142									
-0.90	0.126		71.70	16150.57	2913.89	329.1113	13507.71	0.83	3.65	-0.928	23.69
		0.113									
-1.00	0.100		77.90	19064.46	3847.66	345.1973	13836.82	0.85	3.97	-1.028	23.94
		0.090									
-1.10	0.079		85.40	22912.12	4787.95	341.2087	14182.02	0.87	4.35	-1.128	24.19
		0.071									
-1.20	0.063		93.90	27700.07	5113.35	289.4519	14523.22	0.89	4.78	-1.228	24.44
		0.057									
-1.30	0.050		102.20	32813.42	5615.66	252.5066	14812.67	0.91	5.21	-1.328	24.69
		0.045									
-1.40	0.040		110.60	38429.08	6508.75	232.4719	15065.18	0.92	5.64	-1.428	24.94
		0.036									
-1.50	0.032		119.60	44937.84	6293.03	178.5388	15297.64	0.94	6.09	-1.528	25.19
		0.028									
-1.60	0.025		127.70	51230.86	5770.47	130.0423	15476.18	0.95	6.51	-1.628	25.44
		0.023									
-1.70	0.020		134.70	57001.33	4222.52	75.5868	15606.22	0.96	6.86	-1.728	25.69
		0.018									
-1.80	0.016		139.60	61223.85	5650.59	80.3468	15681.80	0.96	7.11	-1.828	25.94
		0.014									
-1.90	0.013		145.90	66874.44	6091.44	68.8011	15762.15	0.97	7.43	-1.928	26.19
		0.011									
-2.00	0.010		152.40	72965.87			15830.95	0.97	7.77	-2.028	26.44
-∞							16290.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4267

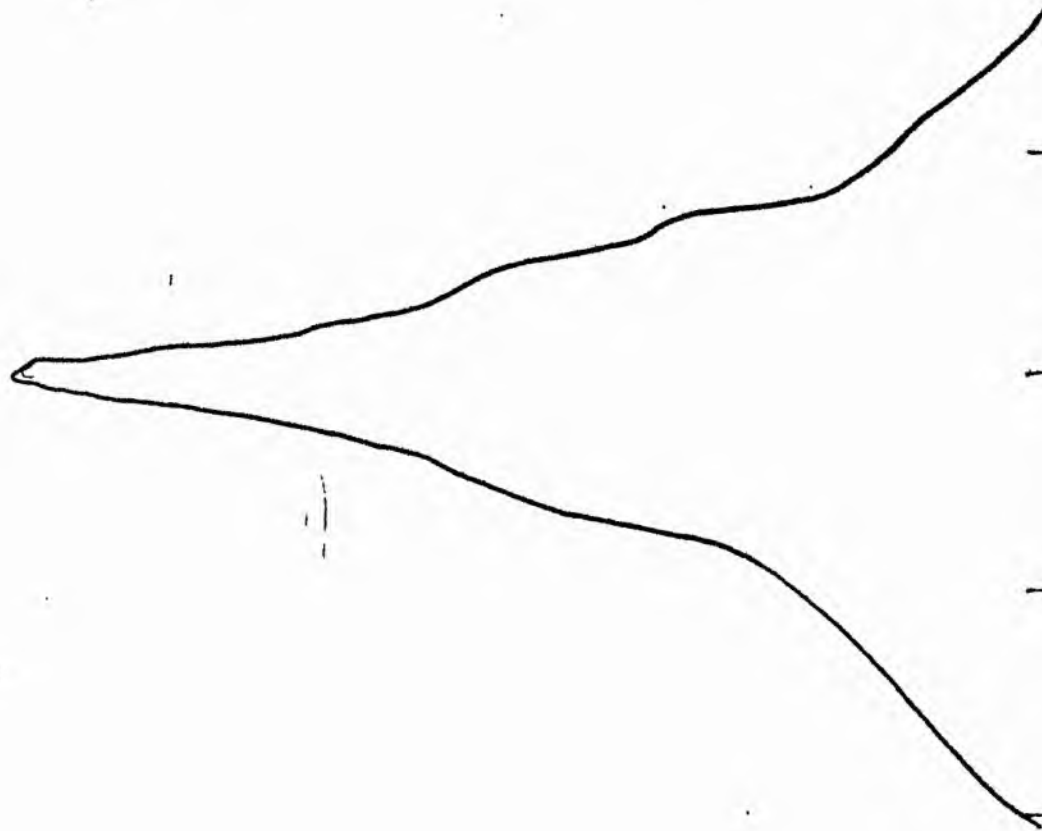
V-FILTER

Total luminosity	L_T	= 4.53
Total apparent magnitude	m_T	= 10.91
Apparent central surface brightness	μ_0	= 16.91
Major axis at threshold	$2a_m$	= 5.56
Minor axis at threshold	$2b_m$	= 5.57
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.1
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.93
Gradient of exponential component	$G(a)$	= -0.97
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.89
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.30
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.16
Axis ratio	b/a	= 1.04
Equivalent radius	r_1^*	= 0.10
Surface brightness	μ_1	= 18.36
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.38
Axis ratio	b/a	= 0.96
Equivalent radius	r_e^*	= 0.33
Surface brightness	μ_e	= 21.36
Mean surface brightness	μ_e'	= 10.51
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.02
Axis ratio	b/a	= 0.92
Equivalent radius	r_3^*	= 0.89
Surface brightness	μ_3	= 22.86
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.13 \\ = 2.73 \end{cases}$

NGC 4273
B-Filter
Axis 1

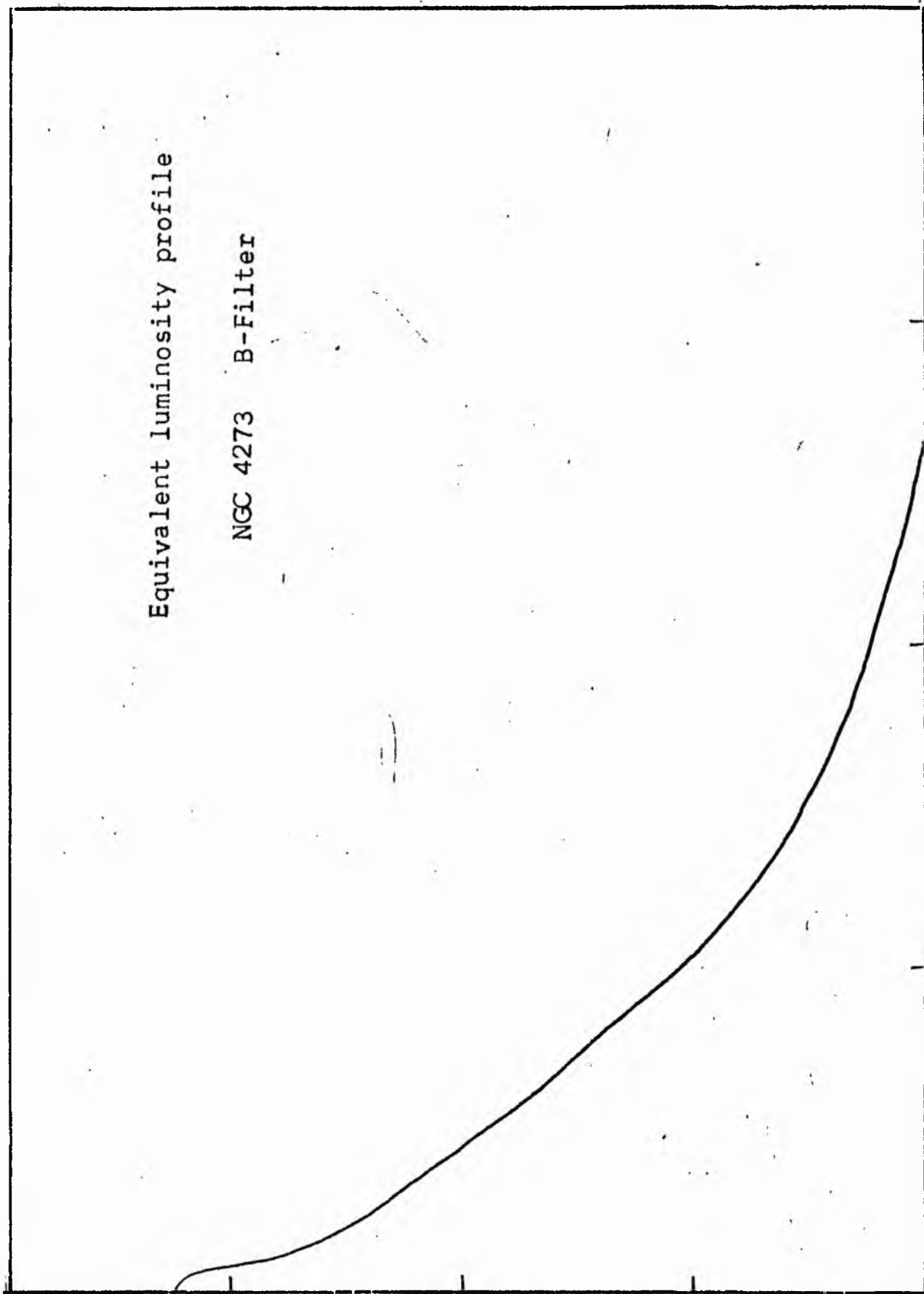


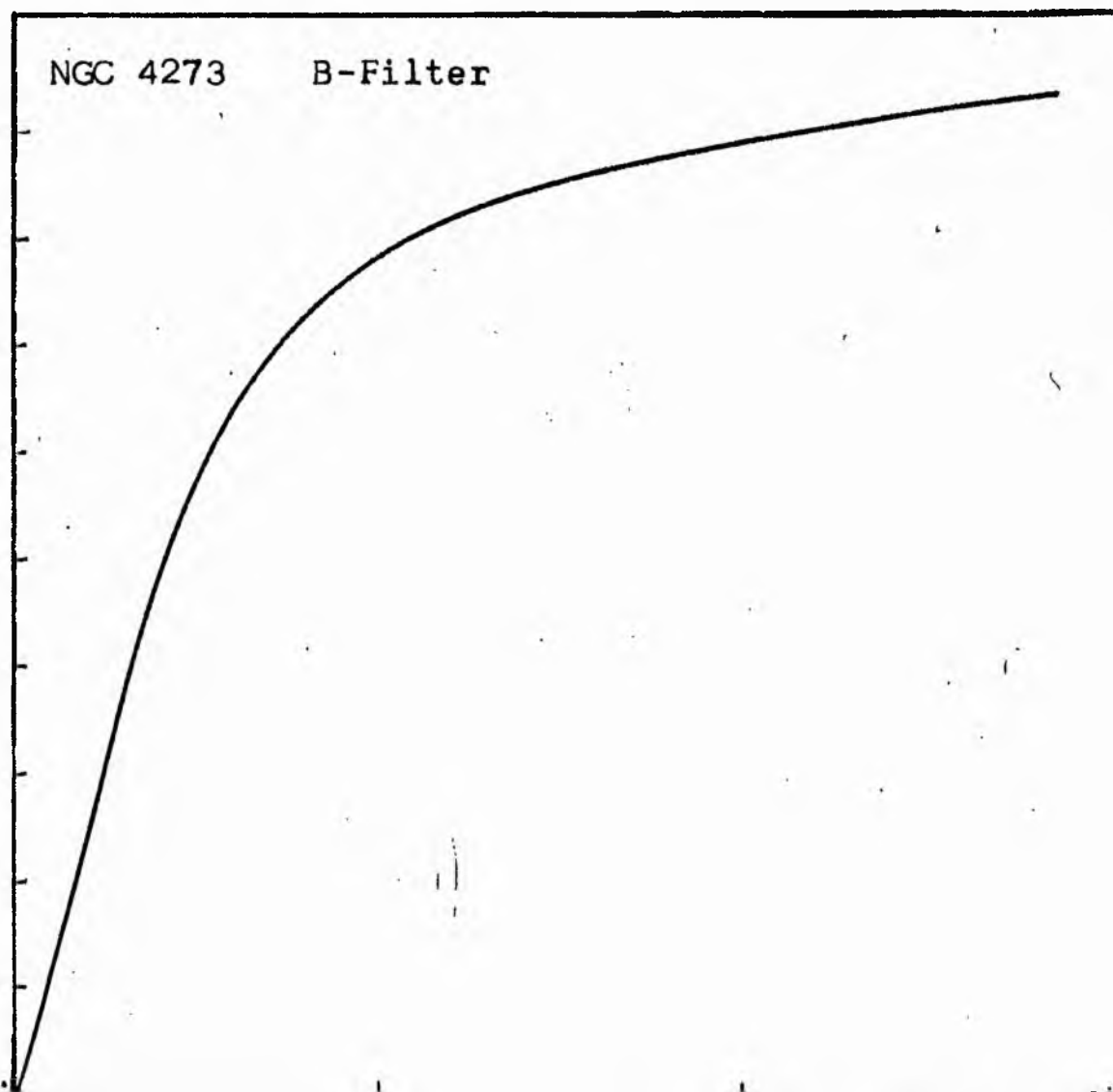
NGC 4273
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4273 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4273
B COLOUR

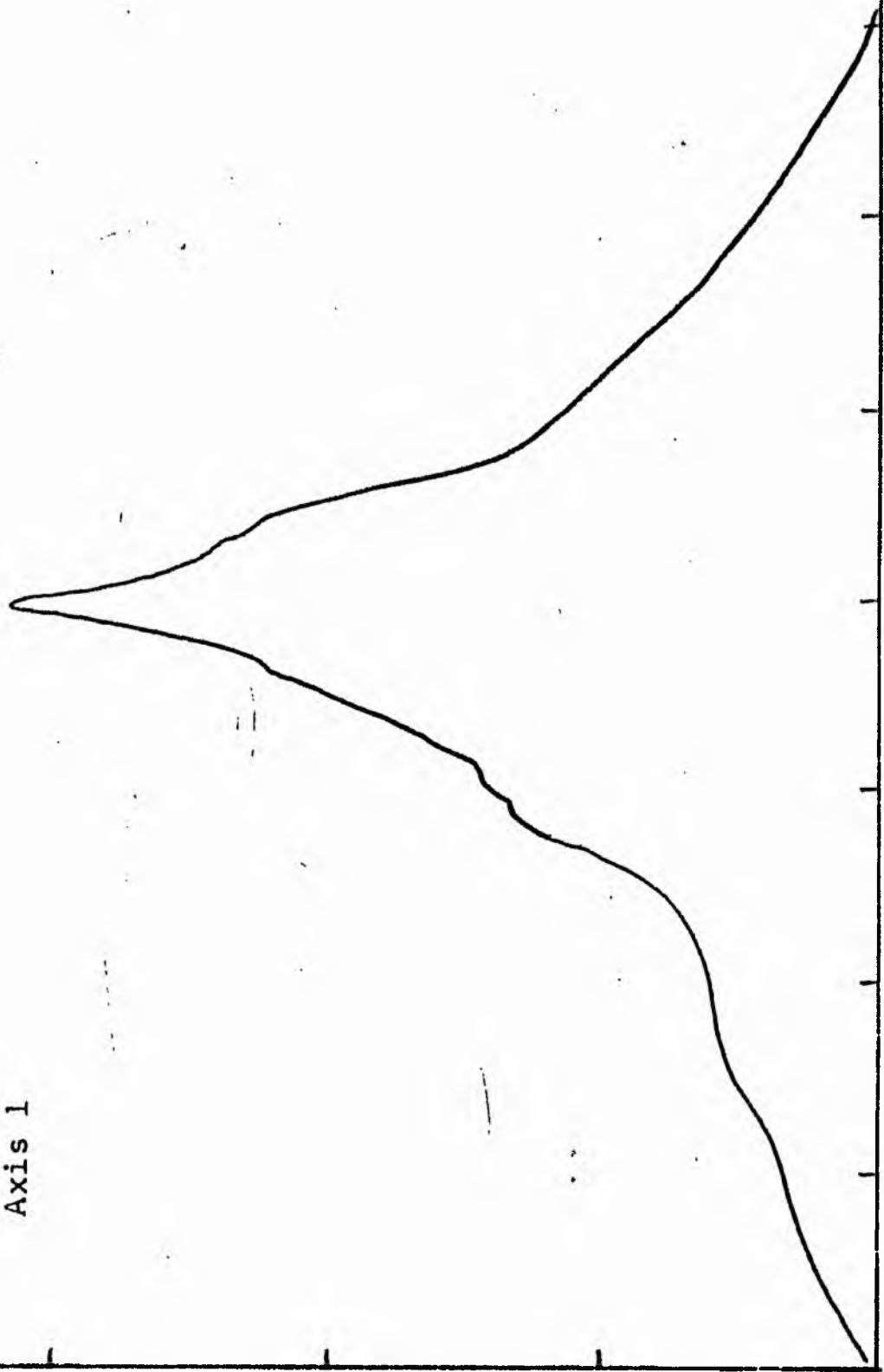
LOG I	I	I	R	AREA	ΔA	P	ZIP	KIRI	ρ	LOG J	μ
1.29	19.498		0.0	0.0			0.0	0.0	0.0	1.217	18.89
1.20	15.849	17.674	2.97	27.71	27.71	489.7666	409.77	0.06	0.14	1.127	19.12
1.10	12.589	14.219	3.13	30.78	3.07	43.5984	533.36	0.06	0.15	1.027	19.37
1.00	10.000	11.295	4.18	54.09	24.11	272.3494	805.71	0.10	0.20	0.927	19.62
0.90	7.943	8.972	4.93	76.36	21.46	192.5755	948.29	0.12	0.24	0.827	19.87
0.80	6.310	7.126	5.78	104.96	28.60	203.8120	1202.10	0.14	0.28	0.727	20.12
0.70	5.012	5.661	7.50	176.71	71.76	406.2073	1608.31	0.19	0.36	0.627	20.37
0.60	3.981	4.496	8.97	252.78	76.06	342.0042	1950.31	0.23	0.44	0.527	20.62
0.50	3.162	3.572	10.38	338.49	85.71	306.1401	2256.45	0.27	0.56	0.427	20.87
0.40	2.512	2.837	12.40	483.05	144.56	410.1350	2666.59	0.32	0.60	0.327	21.12
0.30	1.995	2.254	15.01	707.80	224.75	506.4897	3173.08	0.38	0.73	0.227	21.37
0.20	1.585	1.790	18.04	1022.40	314.60	563.1624	3736.24	0.45	0.87	0.127	21.62
0.10	1.259	1.422	20.07	1265.45	243.04	345.5862	4081.83	0.49	0.97	0.027	21.87
-0.00	1.000	1.129	22.23	1552.49	287.04	324.2021	4406.03	0.53	1.08	-0.073	22.12
-0.10	0.794	0.897	23.85	1787.01	234.52	210.4010	4616.43	0.55	1.16	-0.173	22.37
-0.20	0.631	0.713	26.50	2206.18	419.18	298.7227	4915.15	0.59	1.24	-0.273	22.62
-0.30	0.501	0.566	29.58	2748.82	542.63	307.1694	5222.32	0.63	1.43	-0.373	22.87
-0.40	0.398	0.450	33.95	3621.01	872.19	392.1772	5614.44	0.67	1.65	-0.473	23.12
-0.50	0.316	0.357	37.30	4370.86	749.86	267.8245	5882.32	0.71	1.81	-0.573	23.37
-0.60	0.251	0.284	40.48	5147.91	777.04	220.4536	6102.77	0.73	1.96	-0.673	23.62
-0.70	0.200	0.225	43.39	5914.65	766.74	172.7911	6275.56	0.75	2.10	-0.773	23.87
-0.80	0.158	0.179	46.35	6749.15	834.50	149.3820	6424.94	0.77	2.25	-0.873	24.12
-0.90	0.126	0.142	48.70	7450.88	701.73	99.7803	6524.71	0.78	2.36	-0.973	24.37
-1.00	0.100	0.113	50.20	7916.94	466.05	52.6392	6577.35	0.79	2.43	-1.073	24.62
-1.10	0.079	0.090	54.70	9399.93	1482.99	133.0485	6710.40	0.80	2.65	-1.173	24.87
-1.20	0.063	0.071	58.80	10861.86	1461.94	104.1842	6814.58	0.82	2.85	-1.273	25.12
-1.30	0.050	0.057	65.30	13396.02	2534.16	143.4520	6958.03	0.83	3.17	-1.373	25.37
-1.40	0.040	0.045	70.20	15481.89	2085.87	93.7910	7051.82	0.85	3.40	-1.473	25.62
-1.50	0.032	0.036	76.80	18529.85	3047.96	108.8639	7160.68	0.86	3.72	-1.573	25.87
-1.60	0.025	0.028	83.00	21642.43	3112.58	88.3072	7248.99	0.87	4.03	-1.673	26.12
-1.70	0.020	0.023	95.10	28412.59	6770.16	152.5719	7401.56	0.89	4.61	-1.773	26.37
-1.80	0.016	0.018	107.20	36102.67	7690.08	137.6598	7539.21	0.90	5.20	-1.873	26.62
-1.90	0.013	0.014	116.80	42858.35	6755.68	96.0607	7635.27	0.92	5.66	-1.973	26.87
-2.00	0.010	0.011	128.40	51794.04	8935.70	100.9266	7736.20	0.93	6.23	-2.073	27.12
-∞							8336.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4273

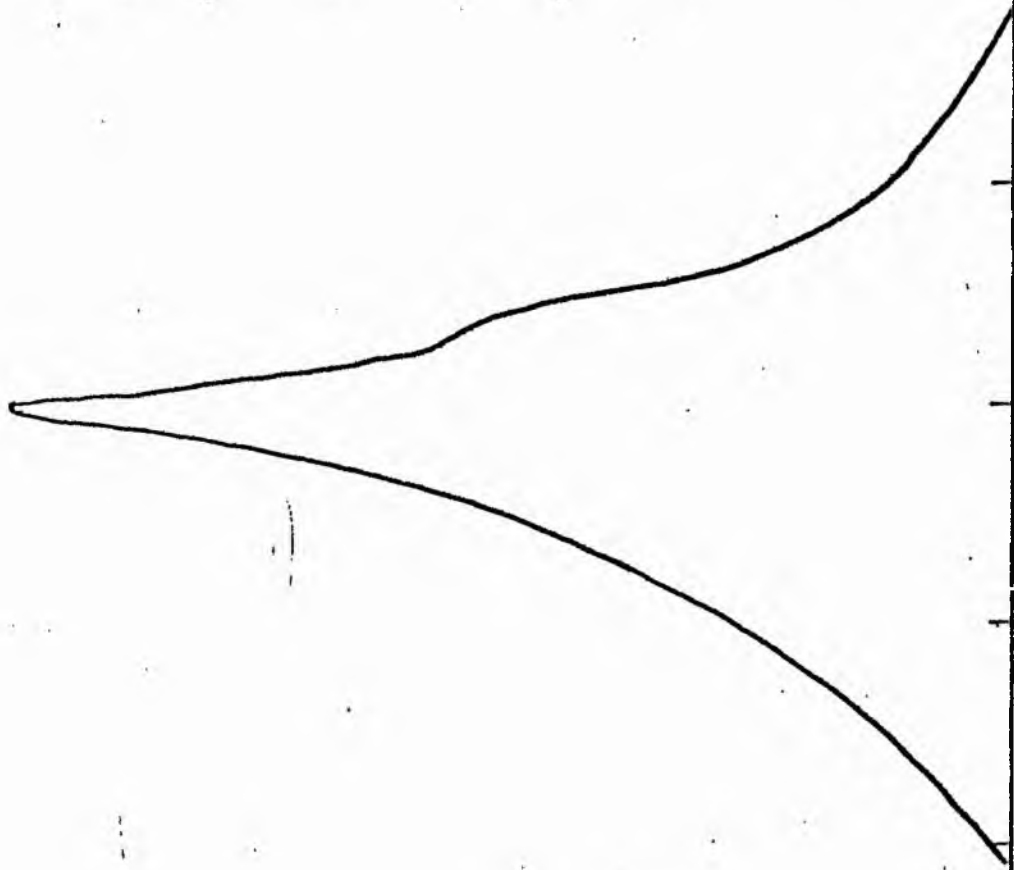
B-FILTER

Total luminosity	L_T	= 2.31
Total apparent magnitude	m_T	= 12.32
Apparent central surface brightness	μ_0	= 18.89
Major axis at threshold	$2a_m$	= 5.13
Minor axis at threshold	$2b_m$	= 3.90
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 2.35
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.81
Gradient of exponential component	$G(a)$	= -1.58
Equivalent gradient of exponential comp....	$G(r^*)$	= -2.08
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.74
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 0.61
Equivalent radius	r_1^*	= 0.16
Surface brightness	μ_1	= 20.74
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.43
Axis ratio	b/a	= 0.58
Equivalent radius	r_e^*	= 0.34
Surface brightness	μ_e	= 21.95
Mean surface brightness	μ_e'	= 11.97
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.99
Axis ratio	b/a	= 0.57
Equivalent radius	r_3^*	= 0.72
Surface brightness	μ_3	= 23.87
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.15 \\ = 2.08 \end{matrix}$

NGC 4273
V-Filter
Axis 1

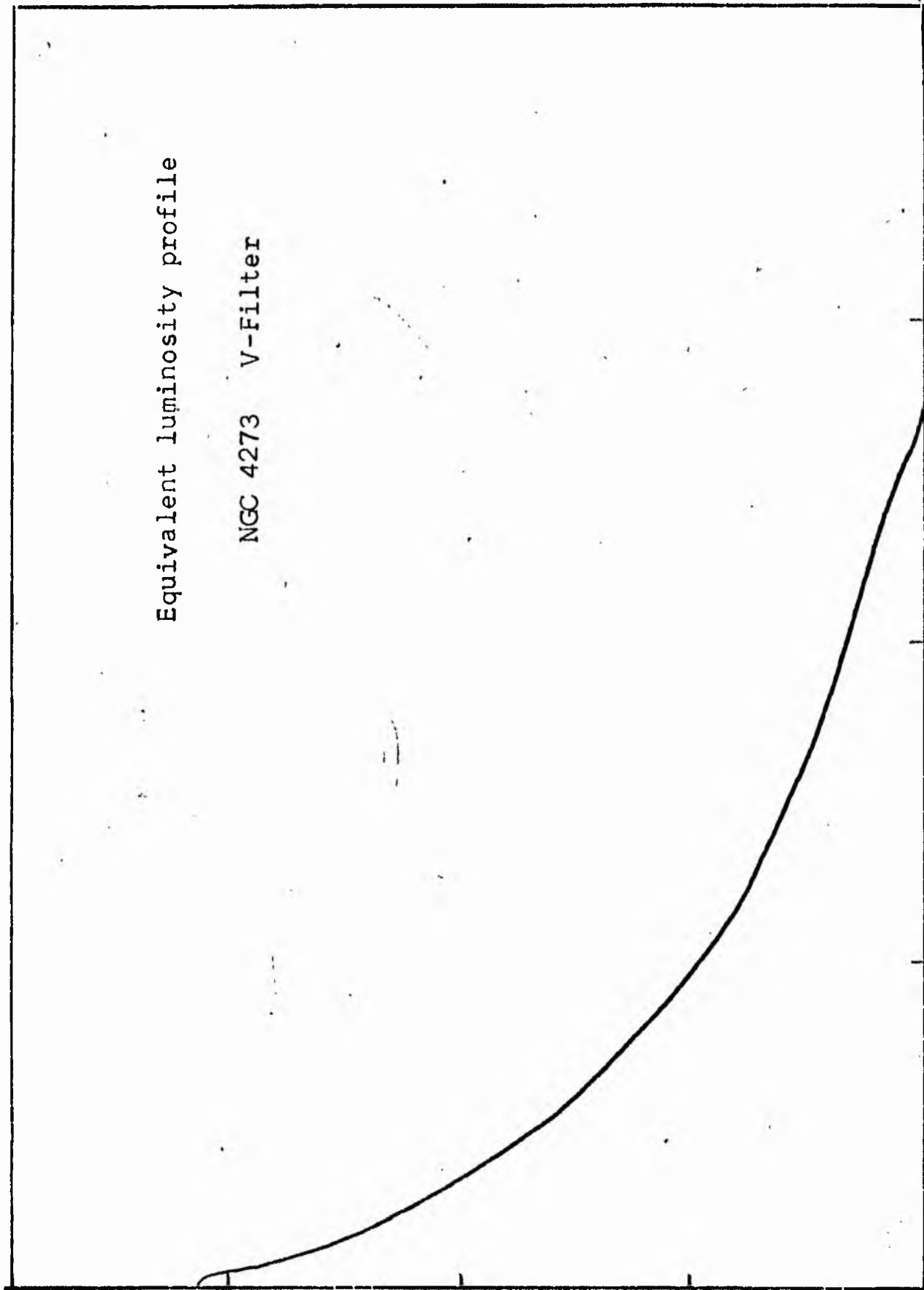


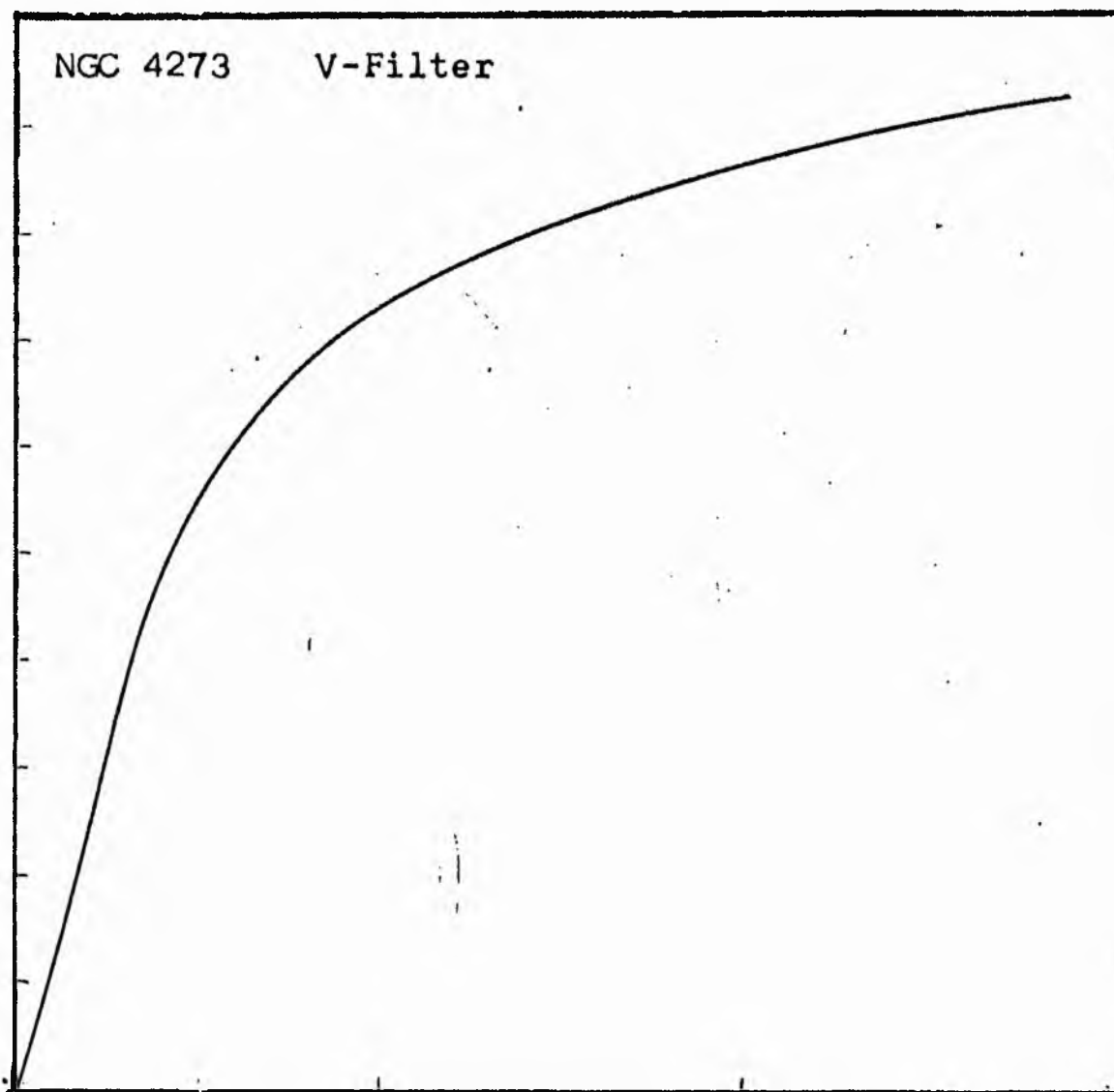
NGC 4273
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4273 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4273
V COLOUR

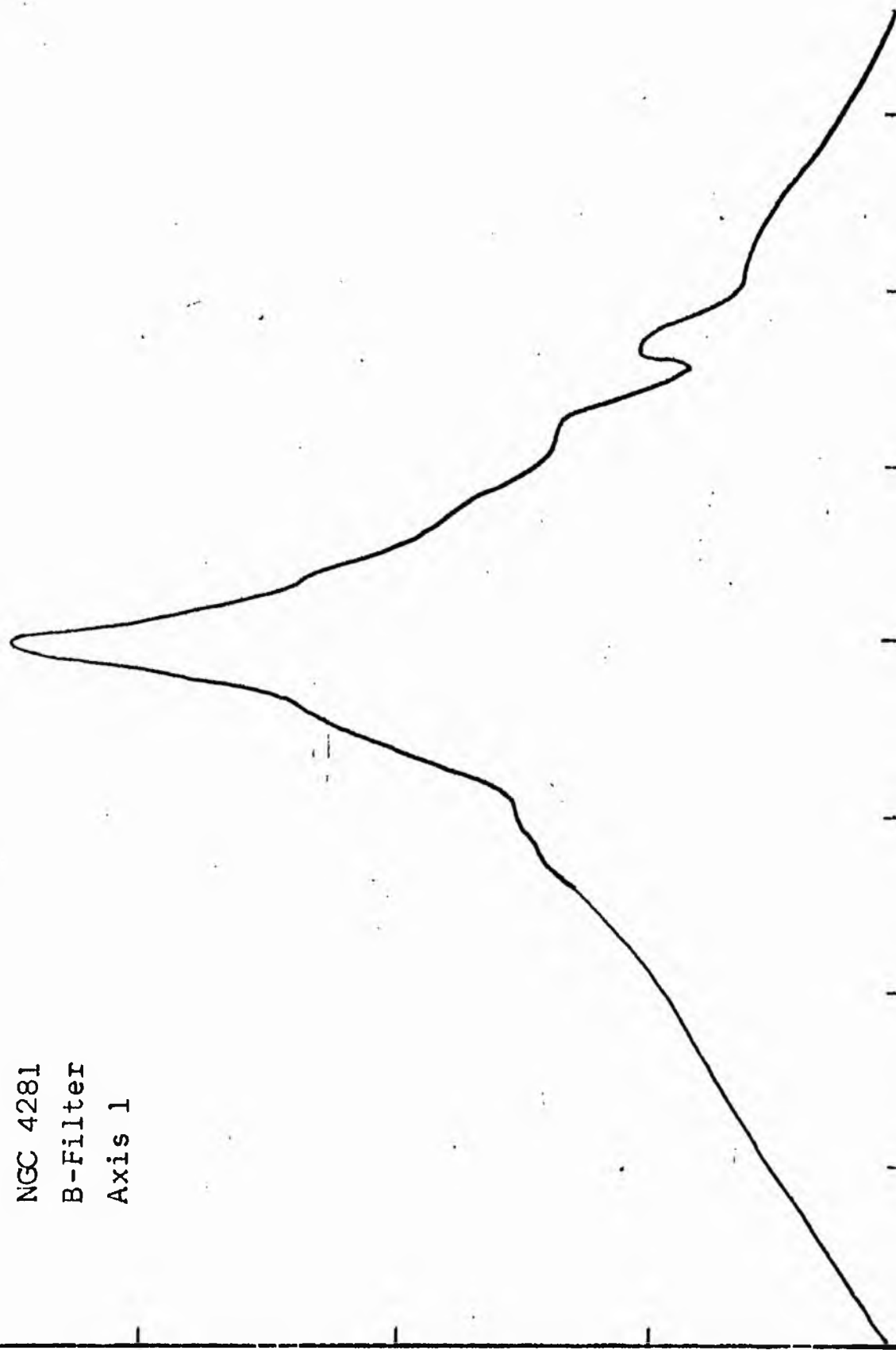
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.17	14.791		0.0	0.0			0.0	0.0	0.0	1.237	18.40
		13.690			12.07	165.2225					
1.10	12.589		1.96	12.07			165.22	0.02	0.09	1.167	18.58
		11.295			16.96	191.6084					
1.00	10.000		3.04	29.03			356.83	0.05	0.14	1.067	18.83
		8.972			31.79	285.1885					
0.90	7.943		4.40	60.82			642.02	0.09	0.21	0.967	19.08
		7.126			20.89	148.8820					
0.80	6.310		5.10	81.71			790.90	0.11	0.24	0.867	19.33
		5.661			23.97	135.6895					
0.70	5.012		5.80	105.68			926.59	0.13	0.27	0.767	19.58
		4.496			66.35	298.3423					
0.60	3.981		7.40	172.03			1224.93	0.18	0.35	0.667	19.83
		3.572			105.56	377.0164					
0.50	3.162		9.40	277.59			1601.95	0.23	0.44	0.567	20.08
		2.837			118.60	336.4895					
0.40	2.512		11.23	396.20			1938.44	0.28	0.53	0.467	20.33
		2.254			152.03	342.6001					
0.30	1.995		13.21	548.22			2281.04	0.33	0.62	0.367	20.58
		1.790			261.06	467.3174					
0.20	1.585		16.05	809.28			2748.36	0.39	0.75	0.267	20.83
		1.422			219.94	312.7278					
0.10	1.259		18.10	1029.22			3061.08	0.44	0.85	0.167	21.08
		1.129			226.16	255.4424					
-0.00	1.000		19.99	1255.38			3316.53	0.48	0.93	0.067	21.33
		0.897			241.74	216.8811					
-0.10	0.794		21.83	1497.12			3533.41	0.51	1.02	-0.033	21.58
		0.713			137.44	97.9439					
-0.20	0.631		22.81	1634.56			3631.35	0.52	1.07	-0.133	21.83
		0.566			338.37	191.5435					
-0.30	0.501		25.06	1972.93			3822.89	0.55	1.17	-0.233	22.08
		0.450			455.01	204.5961					
-0.40	0.398		27.80	2427.95			4027.49	0.58	1.30	-0.333	22.33
		0.357			595.02	212.5215					
-0.50	0.316		31.02	3022.96			4240.01	0.61	1.45	-0.433	22.58
		0.284			969.75	275.1272					
-0.60	0.251		35.65	3992.72			4515.14	0.65	1.67	-0.533	22.83
		0.225			688.12	155.0740					
-0.70	0.200		38.60	4680.84			4670.21	0.67	1.81	-0.633	23.08
		0.179			644.05	115.2908					
-0.80	0.158		41.17	5324.90			4785.50	0.69	1.93	-0.733	23.33
		0.142			1102.02	156.6973					
-0.90	0.126		45.23	6426.92			4942.19	0.71	2.12	-0.833	23.58
		0.113			1486.86	167.9357					
-1.00	0.100		50.19	7913.78			5110.12	0.73	2.35	-0.933	23.83
		0.090			1125.36	100.9636					
-1.10	0.079		53.64	9039.14			5211.09	0.75	2.51	-1.033	24.08
		0.071			1096.38	78.1331					
-1.20	0.063		56.80	10135.53			5289.22	0.76	2.66	-1.133	24.33
		0.057			3425.12	193.8872					
-1.30	0.050		65.70	13560.65			5483.11	0.79	3.07	-1.233	24.58
		0.045			2770.62	124.5809					
-1.40	0.040		72.10	16331.27			5607.68	0.80	3.37	-1.333	24.83
		0.036			6153.60	219.7877					
-1.50	0.032		84.60	22484.87			5827.47	0.84	3.96	-1.433	25.08
		0.028			3645.13	103.4161					
-1.60	0.025		91.20	26130.00			5930.88	0.85	4.27	-1.533	25.33
		0.023			5430.60	122.3837					
-1.70	0.020		100.23	31560.60			6053.27	0.87	4.69	-1.633	25.58
		0.018			6938.01	124.1971					
-1.80	0.016		110.70	38498.61			6177.46	0.89	5.18	-1.733	25.83
		0.014			7042.41	100.1378					
-1.90	0.013		120.40	45541.02			6277.60	0.90	5.63	-1.833	26.08
		0.011			8199.13	92.6072					
-2.00	0.010		130.79	53740.15			6370.20	0.91	6.12	-1.933	26.33
-∞							6970.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4273

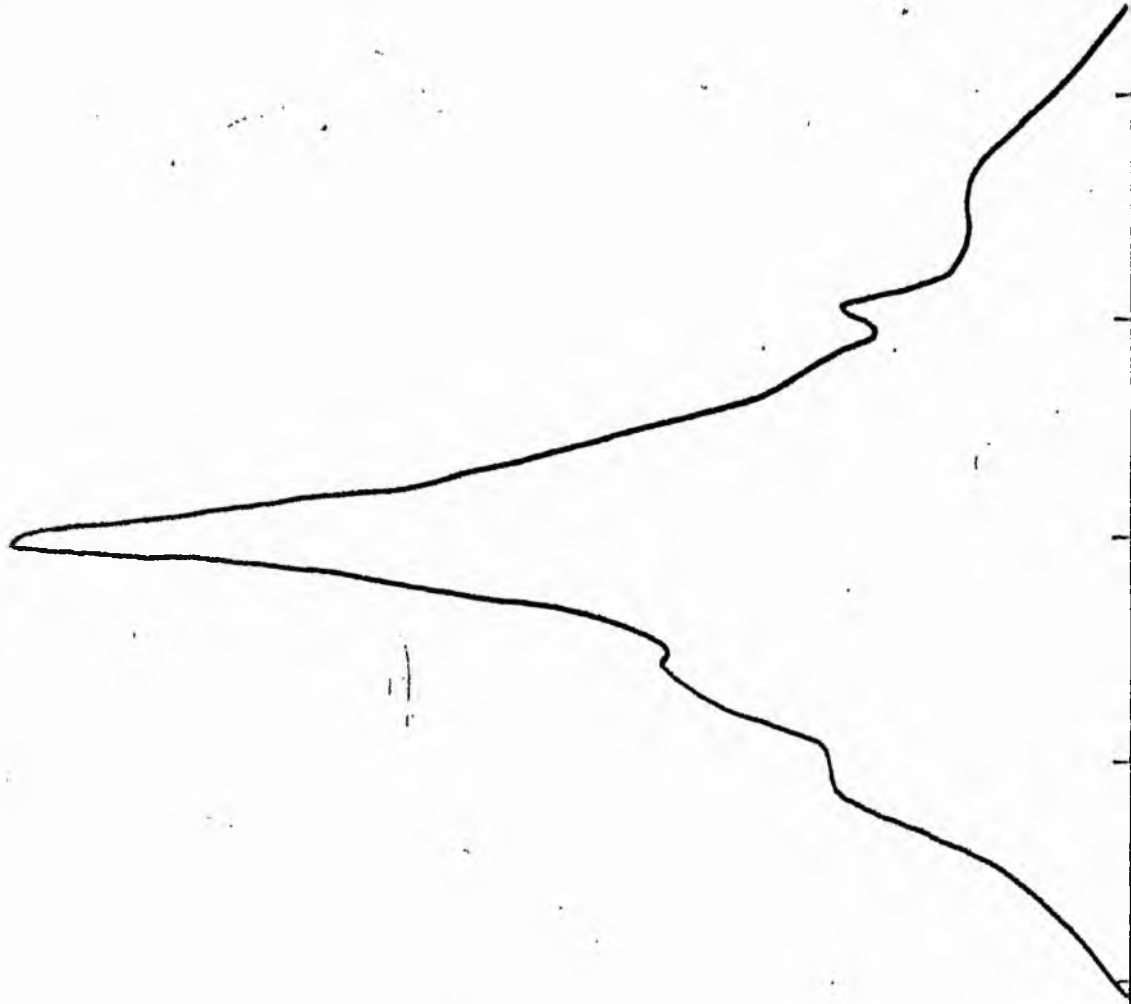
V - FILTER

Total luminosity	L_T	= 1.94
Total apparent magnitude	m_T	= 11.72
Apparent central surface brightness	μ_o	= 18.40
Major axis at threshold	$2a_m$	= 6.05
Minor axis at threshold	$2b_m$	= 3.02
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.52
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -1.31
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.71
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.79
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.23
Axis ratio	b/a	= 0.57
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 20.18
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.45
Axis ratio	b/a	= 0.59
Equivalent radius	r_e^*	= 0.36
Surface brightness	μ_e	= 21.50
Mean surface brightness	μ_e'	= 11.51
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.12
Axis ratio	b/a	= 0.65
Equivalent radius	r_3^*	= 0.90
Surface brightness	μ_3	= 24.08
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.10 \\ = 2.54 \end{cases}$

NGC 4281
B-Filter
Axis 1

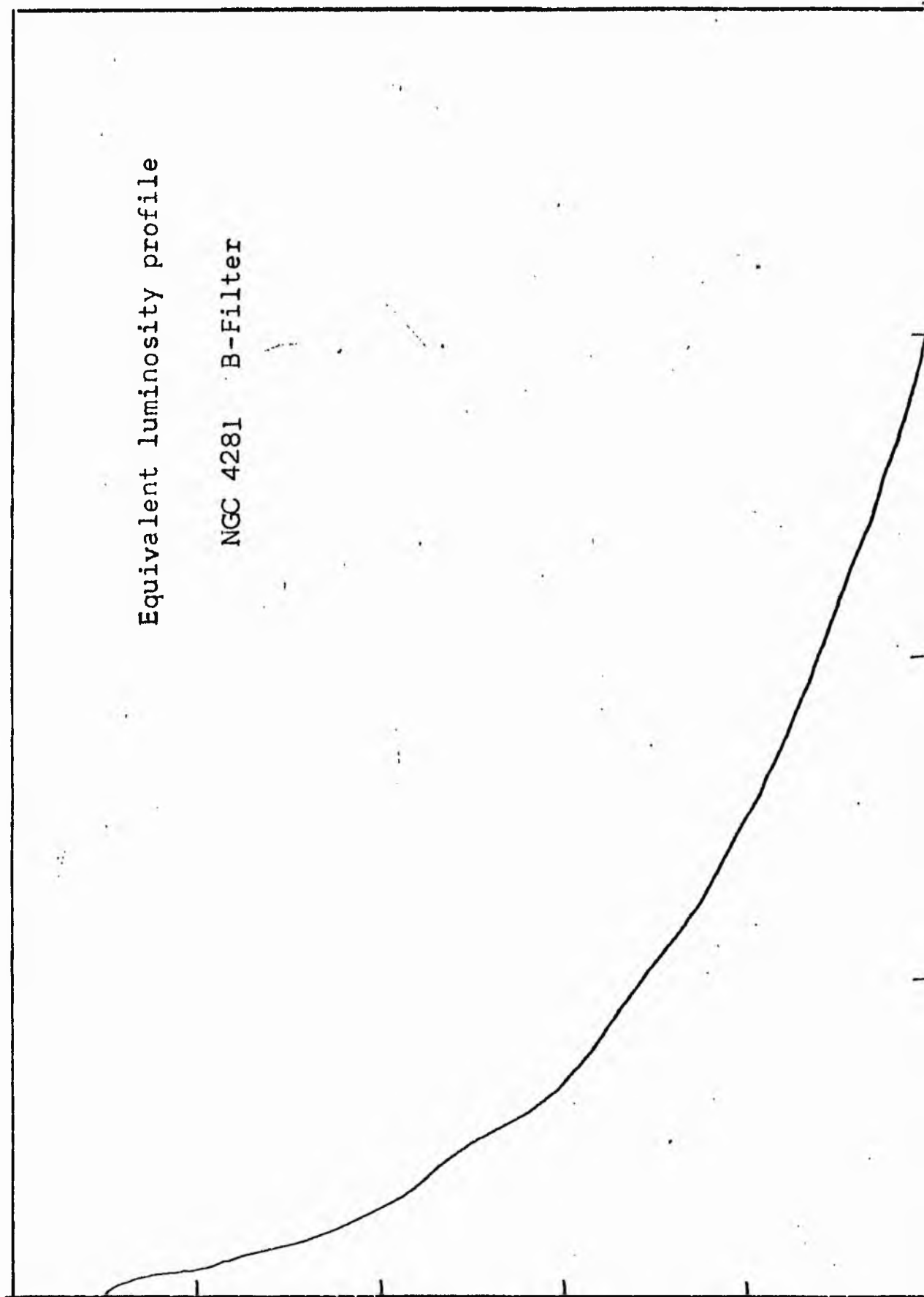


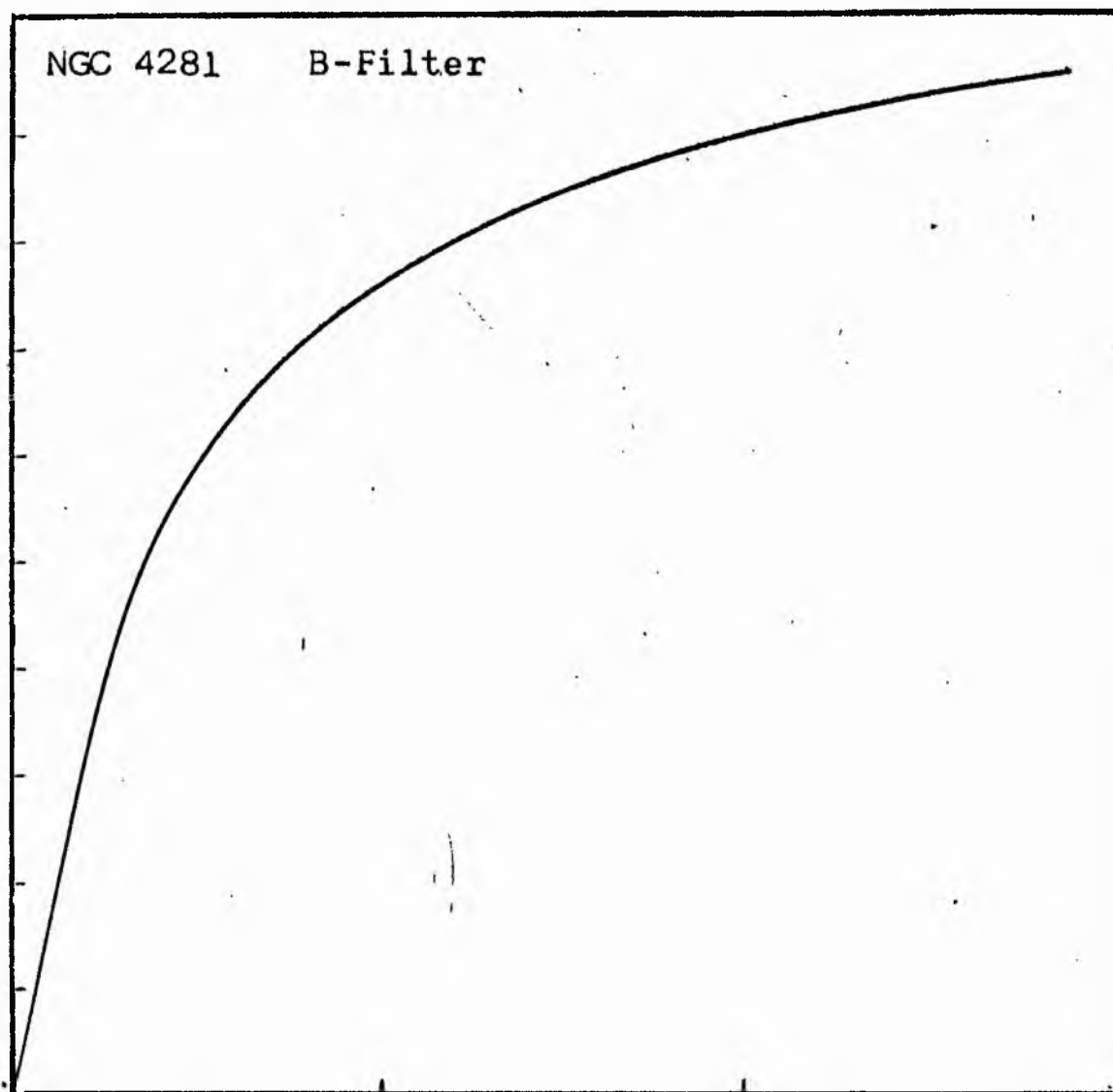
NGC 4281
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4281 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4281
Q COLOUR

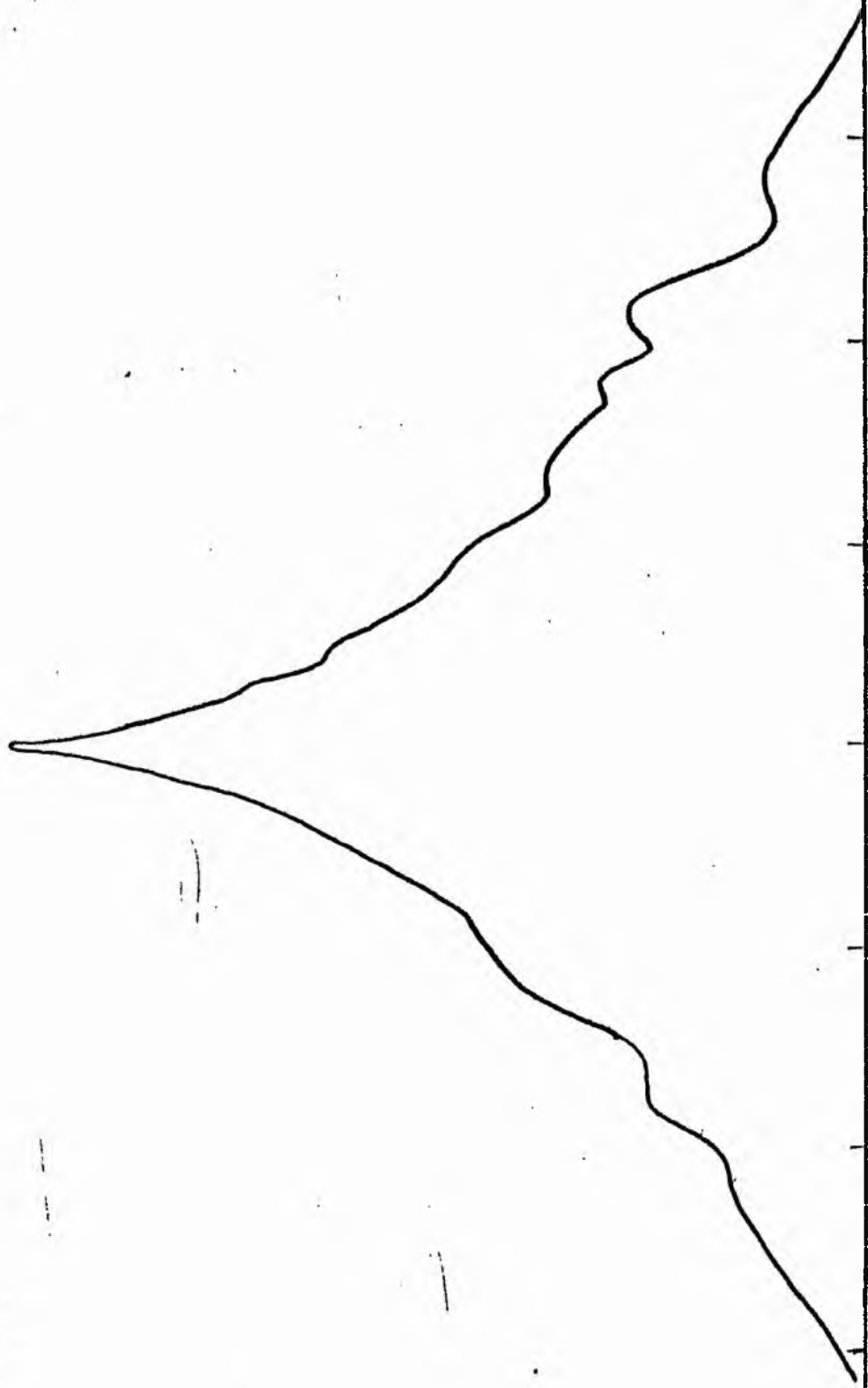
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.54	34.674		0.0	0.0			0.0	0.0	0.0	1.400	18.36
1.50	31.623	33.148	2.30	16.62	16.62	550.8901	550.89	0.05	0.12	1.360	18.46
1.40	25.119	28.371	3.30	34.21	17.59	499.1250	1050.02	0.10	0.17	1.260	18.71
1.30	19.953	22.536	3.80	45.36	11.15	251.3326	1301.35	0.12	0.20	1.160	18.96
1.20	15.849	17.901	4.50	63.62	18.25	326.7358	1628.08	0.16	0.23	1.060	19.21
1.10	12.589	14.219	5.20	84.95	21.33	303.3118	1931.40	0.18	0.27	0.960	19.46
1.00	10.000	11.295	6.10	116.90	31.95	360.0616	2292.26	0.22	0.32	0.860	19.71
0.90	7.943	8.972	6.80	145.27	28.37	254.5112	2546.77	0.24	0.35	0.760	19.96
0.80	6.310	7.126	7.56	179.55	34.29	244.3367	2791.10	0.27	0.39	0.660	20.21
0.70	5.012	5.661	8.68	236.70	57.14	323.4626	3114.57	0.30	0.45	0.560	20.46
0.60	3.981	4.496	10.02	315.42	78.72	353.9695	3468.54	0.33	0.52	0.460	20.71
0.50	3.162	3.572	11.24	396.90	81.48	291.0337	3759.57	0.36	0.58	0.360	20.96
0.40	2.512	2.837	13.04	534.20	137.30	389.5308	4149.10	0.40	0.68	0.260	21.21
0.30	1.995	2.254	14.38	649.63	115.43	260.1306	4409.23	0.42	0.75	0.160	21.46
0.20	1.585	1.790	17.01	908.99	259.36	464.2654	4873.49	0.46	0.88	0.060	21.71
0.10	1.259	1.422	20.50	1320.25	411.27	584.7817	5458.27	0.52	1.07	-0.040	21.96
-0.00	1.000	1.129	23.14	1682.20	361.94	408.7976	5867.07	0.56	1.20	-0.140	22.21
-0.10	0.794	0.897	24.89	1946.25	264.06	236.9034	6103.97	0.58	1.29	-0.240	22.46
-0.20	0.631	0.713	27.43	2363.75	417.49	297.5229	6401.49	0.61	1.43	-0.340	22.71
-0.30	0.501	0.566	29.98	2823.66	459.92	260.3447	6661.83	0.63	1.56	-0.440	22.96
-0.40	0.398	0.450	31.23	3064.03	240.37	108.0818	6769.91	0.64	1.62	-0.540	23.21
-0.50	0.316	0.357	34.77	3798.03	734.00	262.1602	7032.07	0.67	1.81	-0.640	23.46
-0.60	0.251	0.284	39.13	4810.27	1012.23	287.1775	7319.25	0.70	2.04	-0.740	23.71
-0.70	0.200	0.225	44.30	6165.34	1355.07	305.3752	7624.62	0.73	2.30	-0.840	23.96
-0.80	0.158	0.179	50.29	7945.35	1780.01	318.6345	7943.25	0.76	2.62	-0.940	24.21
-0.90	0.126	0.142	54.70	9399.93	1454.58	206.8275	8150.08	0.78	2.85	-1.040	24.46
-1.00	0.100	0.113	58.70	10824.95	1425.03	160.9514	8311.03	0.79	3.05	-1.140	24.71
-1.10	0.079	0.090	66.30	13809.46	2984.51	267.7595	8578.79	0.82	3.45	-1.240	24.96
-1.20	0.063	0.071	73.10	16787.43	2977.97	212.2233	8791.01	0.84	3.80	-1.340	25.21
-1.30	0.050	0.057	80.20	20206.85	3419.42	193.5643	8984.57	0.86	4.17	-1.440	25.46
-1.40	0.040	0.045	86.40	23451.86	3245.01	145.9115	9130.48	0.87	4.49	-1.540	25.71
-1.50	0.032	0.036	95.70	28772.24	5320.39	190.0280	9320.51	0.89	4.98	-1.640	25.96
-1.60	0.025	0.028	105.40	34900.45	6120.20	173.8634	9494.37	0.90	5.48	-1.740	26.21
-1.70	0.020	0.023	115.20	41692.20	6791.76	153.0587	9647.43	0.92	5.99	-1.840	26.46
-1.80	0.016	0.018	125.40	49402.04	7709.04	138.0135	9785.44	0.93	6.52	-1.940	26.71
-1.90	0.013	0.014	135.70	57850.82	8448.79	120.1354	9905.57	0.94	7.06	-2.040	26.96
-2.00	0.010	0.011	145.30	66325.56	8474.74	95.7202	10001.29	0.95	7.56	-2.140	27.21
-∞							10501.00	(1)		∞	

PHOTOMETRIC PARAMETERS OF NGC 4281

B-FILTER

Total luminosity	L_T	= 2.92
Total apparent magnitude	m_T	= 12.26
Apparent central surface brightness	μ_o	= 18.36
Major axis at threshold	$2a_m$	= 6.24
Minor axis at threshold	$2b_m$	= 3.65
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.28
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.95
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.02
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.79
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.18
Axis ratio	b/a	= 0.67
Equivalent radius	r_1^*	= 0.12
Surface brightness	μ_1	= 20.04
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.54
Axis ratio	b/a	= 0.41
Equivalent radius	r_e^*	= 0.32
Surface brightness	μ_e	= 21.88
Mean surface brightness	μ_e'	= 11.75
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.18
Axis ratio	b/a	= 0.47
Equivalent radius	r_3^*	= 0.82
Surface brightness	μ_3	= 24.13
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.73 \\ = 2.55 \end{matrix}$

NGC 4281
V-Filter
Axis 1

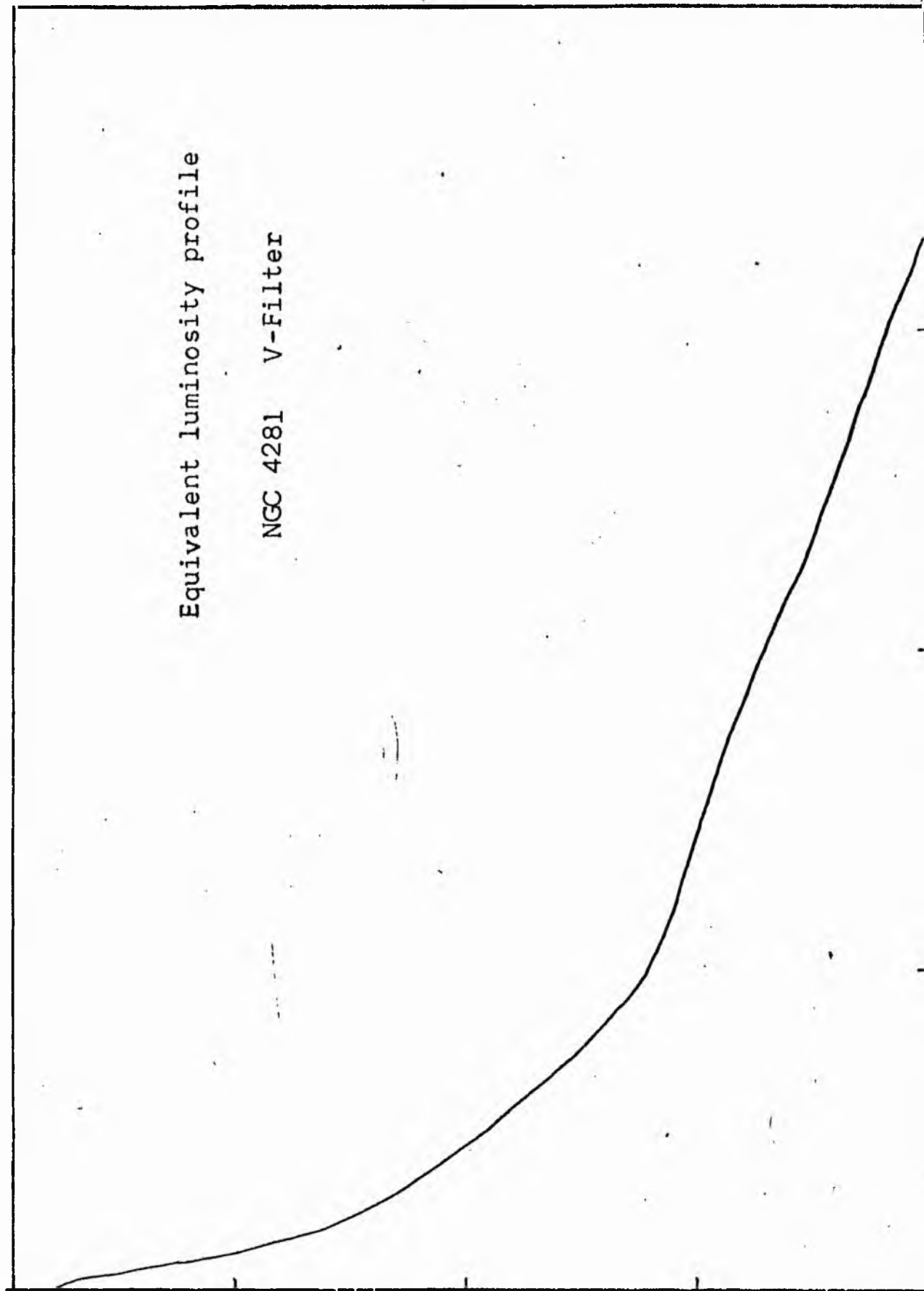


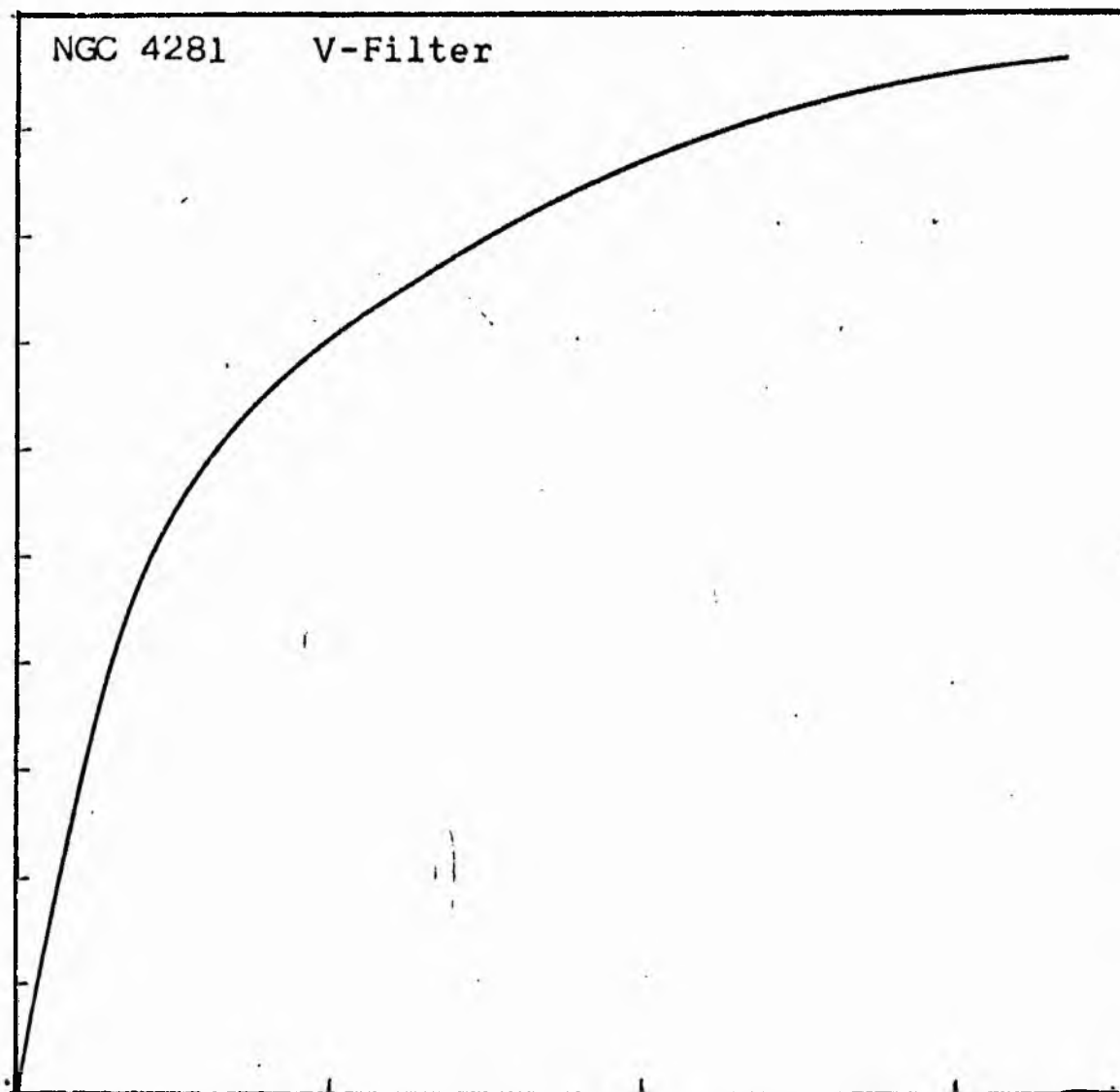
NGC 4281
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4281 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4281
V COLOUR

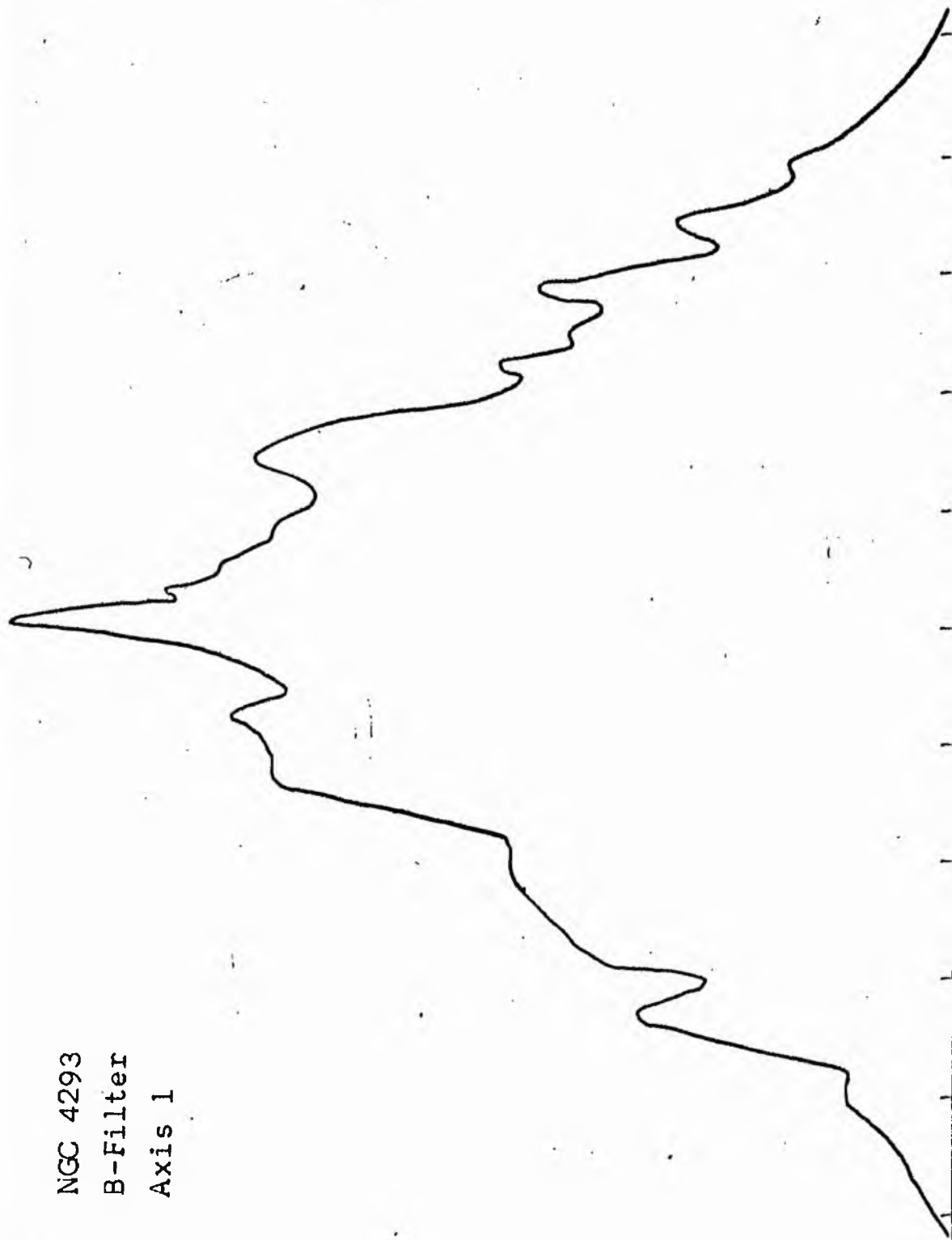
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.70	60.256		0.0	0.0			0.0	0.0	0.0	1.665	17.07
		55.187			3.14	173.3760					
1.70	50.119		1.00	3.14			173.38	0.01	0.05	1.585	17.27
		44.965			5.94	266.9824					
1.60	39.811		1.70	9.08			440.36	0.03	0.08	1.485	17.52
		35.717			15.55	555.4253					
1.50	31.623		2.80	24.63			995.78	0.07	0.13	1.385	17.77
		28.371			13.85	393.0608					
1.40	25.119		3.50	38.48			1388.84	0.10	0.16	1.285	18.02
		22.536			14.33	322.8369					
1.30	19.953		4.10	52.81			1711.68	0.13	0.19	1.185	18.27
		17.901			10.81	193.4551					
1.20	15.849		4.50	63.62			1905.14	0.14	0.21	1.085	18.52
		14.219			45.74	650.3994					
1.10	12.589		5.90	109.36			2555.54	0.19	0.27	0.985	18.77
		11.295			40.65	459.0815					
1.00	10.000		6.91	150.01			3014.62	0.22	0.32	0.885	19.02
		8.972			33.85	303.6768					
0.90	7.943		7.65	183.85			3318.29	0.25	0.35	0.785	19.27
		7.126			41.53	295.9358					
0.80	6.310		8.47	225.38			3614.23	0.27	0.39	0.685	19.52
		5.661			51.62	292.2053					
0.70	5.012		9.39	277.00			3906.44	0.29	0.43	0.585	19.77
		4.496			100.37	451.3186					
0.60	3.981		10.96	377.37			4357.75	0.32	0.50	0.485	20.02
		3.572			126.15	450.5549					
0.50	3.162		12.66	503.52			4808.31	0.36	0.58	0.385	20.27
		2.837			166.14	471.3540					
0.40	2.512		14.60	669.66			5279.66	0.39	0.67	0.285	20.52
		2.254			227.61	512.9270					
0.30	1.995		16.90	897.27			5792.59	0.43	0.78	0.185	20.77
		1.790			301.00	538.8093					
0.20	1.585		19.53	1198.27			6331.39	0.47	0.90	0.085	21.02
		1.422			350.03	497.7134					
0.10	1.259		22.20	1548.30			6829.11	0.51	1.02	-0.015	21.27
		1.129			430.93	486.7183					
-0.00	1.000		25.10	1979.23			7315.82	0.54	1.15	-0.115	21.52
		0.897			396.60	355.8101					
-0.10	0.794		27.50	2375.83			7671.63	0.57	1.26	-0.215	21.77
		0.713			395.34	281.7324					
-0.20	0.631		29.70	2771.17			7953.36	0.59	1.36	-0.315	22.02
		0.566			385.79	218.3828					
-0.30	0.501		31.70	3156.95			8171.74	0.60	1.45	-0.415	22.27
		0.450			824.57	370.7649					
-0.40	0.398		35.60	3981.53			8542.50	0.63	1.63	-0.515	22.52
		0.357			723.60	258.4460					
-0.50	0.316		38.70	4705.13			8800.95	0.65	1.78	-0.615	22.77
		0.284			627.54	178.0362					
-0.60	0.251		41.20	5332.66			8978.98	0.66	1.89	-0.715	23.02
		0.225			1228.52	276.8547					
-0.70	0.200		45.70	6561.18			9255.83	0.68	2.10	-0.815	23.27
		0.179			2163.93	387.3586					
-0.80	0.158		52.70	8725.11			9643.19	0.71	2.42	-0.915	23.52
		0.142			4425.86	629.3152					
-0.90	0.126		64.70	13150.98			10272.50	0.76	2.97	-1.015	23.77
		0.113			3774.54	426.3186					
-1.00	0.100		73.40	16925.52			10698.82	0.79	3.37	-1.115	24.02
		0.090			5825.92	522.6797					
-1.10	0.079		85.10	22751.44			11221.50	0.83	3.90	-1.215	24.27
		0.071			5185.13	369.5139					
-1.20	0.063		94.30	27936.57			11591.01	0.86	4.33	-1.315	24.52
		0.057			3605.15	204.0776					
-1.30	0.050		100.20	31541.71			11795.08	0.87	4.60	-1.415	24.77
		0.045			6264.48	281.6814					
-1.40	0.040		109.70	37806.20			12076.76	0.89	5.03	-1.515	25.02
		0.036			6085.74	217.3638					
-1.50	0.032		118.20	43891.94			12294.12	0.91	5.42	-1.615	25.27
		0.028			8630.73	244.8623					
-1.60	0.025		129.30	52522.67			12538.98	0.93	5.93	-1.715	25.52
		0.023			6183.92	139.3601					
-1.70	0.020		136.70	58706.59			12678.34	0.94	6.27	-1.815	25.77
		0.018			6618.55	118.4781					
-1.80	0.016		144.20	65325.14			12796.82	0.95	6.62	-1.915	26.02
		0.014			13300.23	189.1190					
-1.90	0.013		158.20	78625.37			12985.94	0.96	7.26	-2.015	26.27
		0.011			6180.25	69.8043					
-2.00	0.010		164.30	84805.62			13055.74	0.97	7.54	-2.115	26.52
-∞							13515.00	(1)		∞	

PHOTOMETRIC PARAMETERS OF NGC 4281

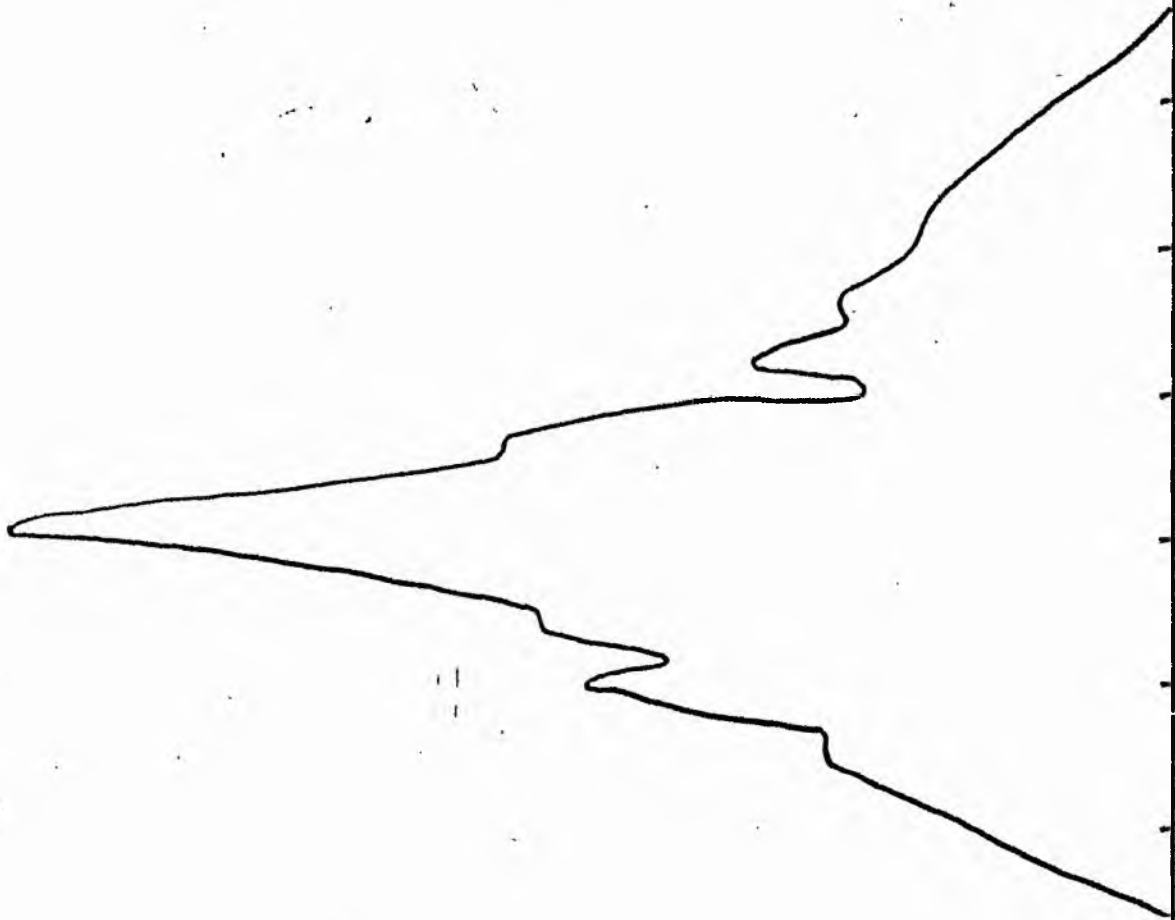
V-FILTER

Total luminosity	L_T	= 3.75
Total apparent magnitude	m_T	= 11.19
Apparent central surface brightness	μ_o	= 17.07
Major axis at threshold	$2a_m$	= 5.96
Minor axis at threshold	$2b_m$	= 3.95
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.80
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.88
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.76
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.80
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.19
Axis ratio	b/a	= 0.69
Equivalent radius	r_1^*	= 0.13
Surface brightness	μ_1	= 19.27
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.37
Axis ratio	b/a	= 0.73
Equivalent radius	r_e^*	= 0.36
Surface brightness	μ_e	= 21.21
Mean surface brightness	μ_e'	= 10.98
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.26
Axis ratio	b/a	= 0.47
Equivalent radius	r_3^*	= 1.03
Surface brightness	μ_3	= 23.72
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.79 \\ = 2.85 \end{matrix}$

NGC 4293
B-Filter
Axis 1

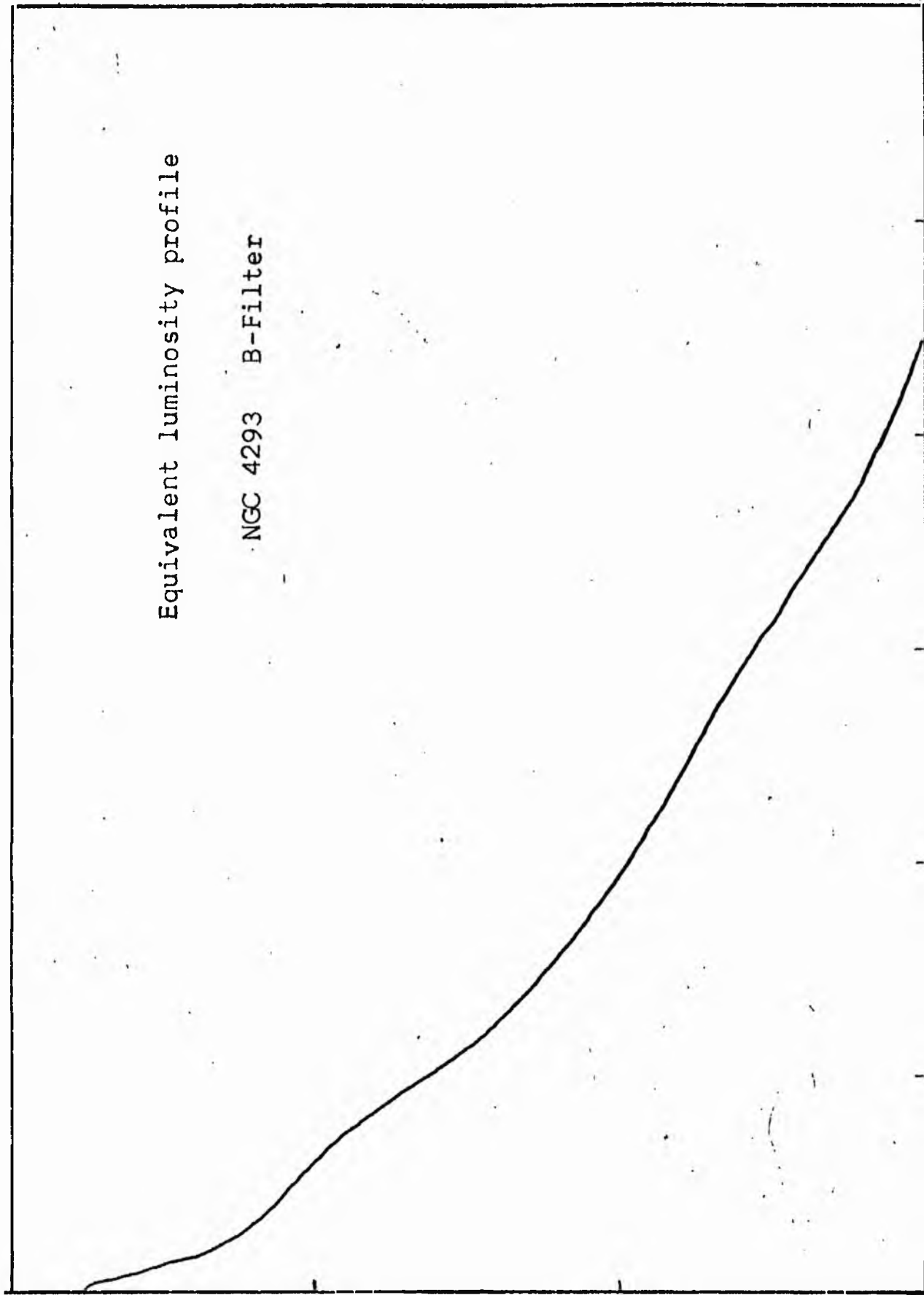


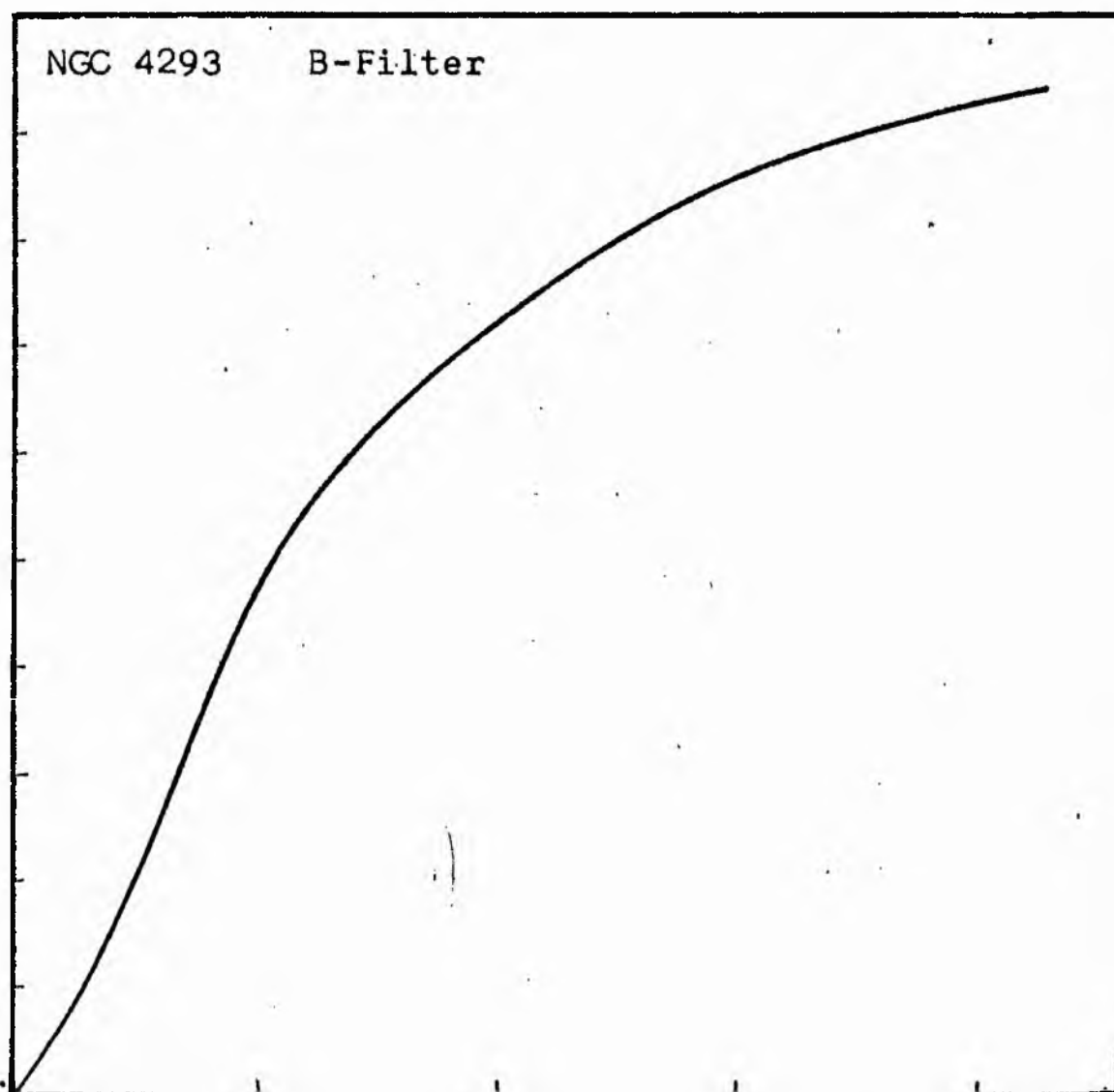
NGC 4293
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4293 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4293
B COLOUR

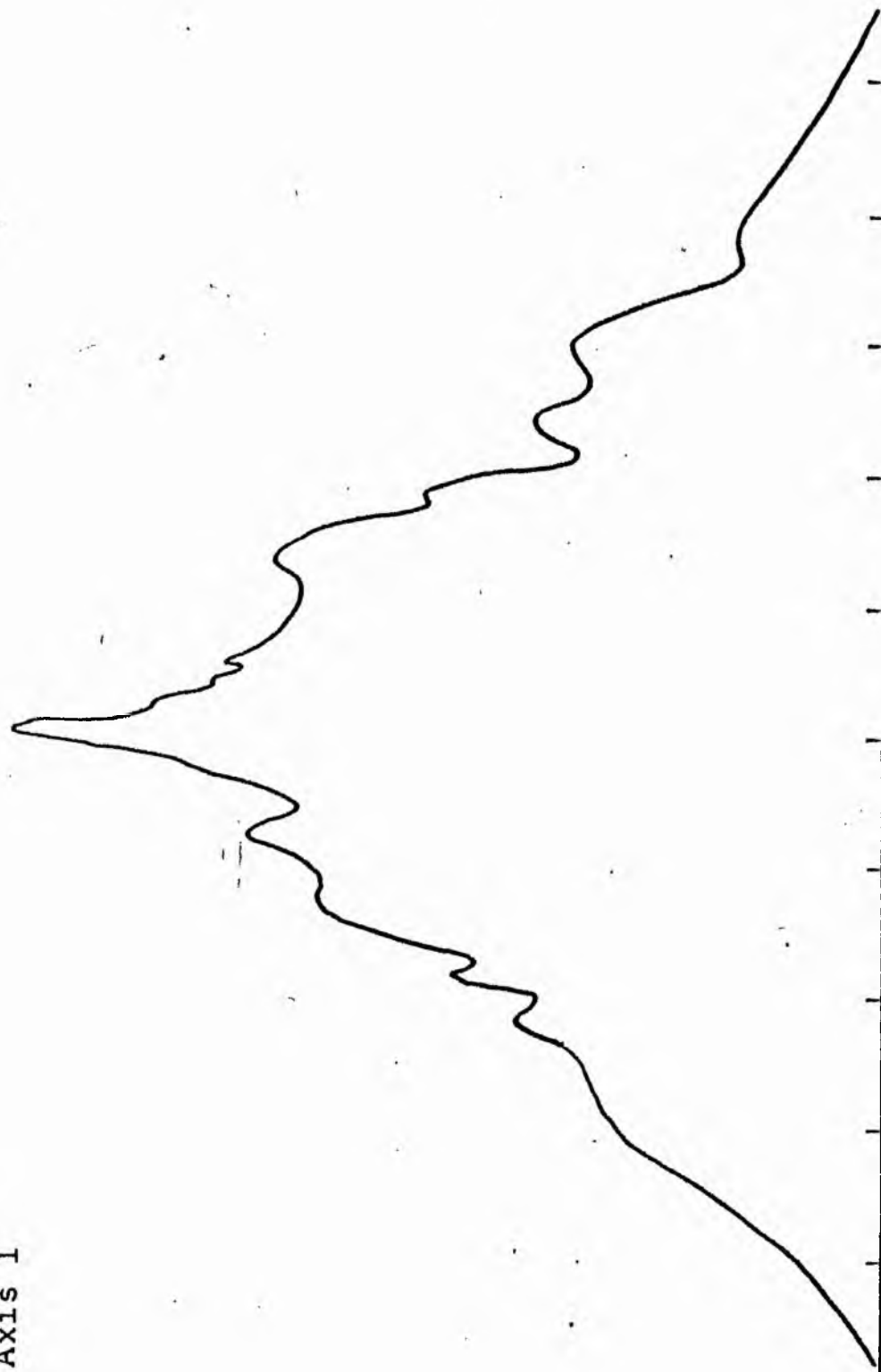
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.79	6.166	5.589	0.0	0.0	22.23	124.2338	0.0	0.0	0.0	1.288	19.58
0.70	5.012	4.496	2.66	22.23	29.81	134.0465	124.23	0.01	0.05	1.198	19.81
0.60	3.981	3.572	4.07	52.04	75.44	269.4321	258.28	0.02	0.08	1.098	20.06
0.50	3.162	2.837	6.37	127.48	25.58	72.5826	527.71	0.03	0.12	0.998	20.31
0.40	2.512	2.254	6.98	153.06	173.79	391.6519	600.29	0.04	0.13	0.898	20.56
0.30	1.995	1.790	10.20	326.85	171.91	307.7265	991.95	0.06	0.19	0.798	20.81
0.20	1.585	1.422	12.60	498.76	1035.62	1472.5649	1299.68	0.08	0.23	0.698	21.06
0.10	1.259	1.129	22.10	1534.38	1332.77	1505.3157	2772.24	0.17	0.41	0.598	21.31
0.00	1.000	0.897	30.21	2867.15	1357.31	1217.7275	4277.55	0.27	0.56	0.498	21.56
-0.10	0.794	0.713	36.67	4224.46	1269.91	904.9897	5495.28	0.34	0.68	0.398	21.81
-0.20	0.631	0.566	41.82	5494.37	1345.36	761.5730	6400.27	0.40	0.78	0.298	22.06
-0.30	0.501	0.450	46.66	6839.73	638.71	287.1965	7161.84	0.45	0.87	0.198	22.31
-0.40	0.398	0.357	48.79	7478.45	1689.22	803.3359	7449.04	0.46	0.91	0.098	22.56
-0.50	0.316	0.284	54.02	9167.67	2255.44	639.8879	8052.37	0.50	1.00	-0.002	22.81
-0.60	0.251	0.225	60.30	11423.11	2933.21	661.0210	8692.26	0.54	1.12	-0.102	23.06
-0.70	0.200	0.179	67.60	14356.32	3770.41	674.9343	9353.28	0.58	1.25	-0.202	23.31
-0.80	0.158	0.142	75.96	18126.73	4763.93	677.3889	10028.21	0.62	1.41	-0.302	23.56
-0.90	0.126	0.113	85.36	22890.66	5045.90	569.9167	10705.60	0.67	1.58	-0.402	23.81
-1.00	0.100	0.090	94.30	27936.57	9183.51	823.9133	11275.51	0.70	1.75	-0.502	24.06
-1.10	0.079	0.071	108.70	37120.08	7540.14	537.3450	12099.42	0.75	2.02	-0.602	24.31
-1.20	0.063	0.057	119.23	44660.22	6330.21	358.3376	12636.77	0.79	2.21	-0.702	24.56
-1.30	0.050	0.045	127.40	50990.43	11733.62	527.6023	12995.10	0.81	2.36	-0.802	24.81
-1.40	0.040	0.036	141.30	62724.05	11395.39	407.0093	13522.70	0.84	2.62	-0.902	25.06
-1.50	0.032	0.028	153.60	74119.44	13181.87	373.9841	13929.71	0.87	2.85	-1.002	25.31
-1.60	0.025	0.023	166.70	87301.31	9129.94	205.7522	14303.69	0.89	3.09	-1.102	25.56
-1.70	0.020	0.018	175.20	96431.25	12606.12	225.6622	14509.44	0.90	3.25	-1.202	25.81
-1.80	0.016	0.014	186.30	109037.37	17129.44	243.5684	14735.10	0.92	3.46	-1.302	26.06
-1.90	0.013	0.011	200.40	126166.81	13302.56	150.2497	14978.67	0.93	3.72	-1.402	26.31
-2.00	0.010		210.70	139469.37			15128.91	0.94	3.91	-1.502	26.56

PHOTOMETRIC PARAMETERS OF NGC 4293

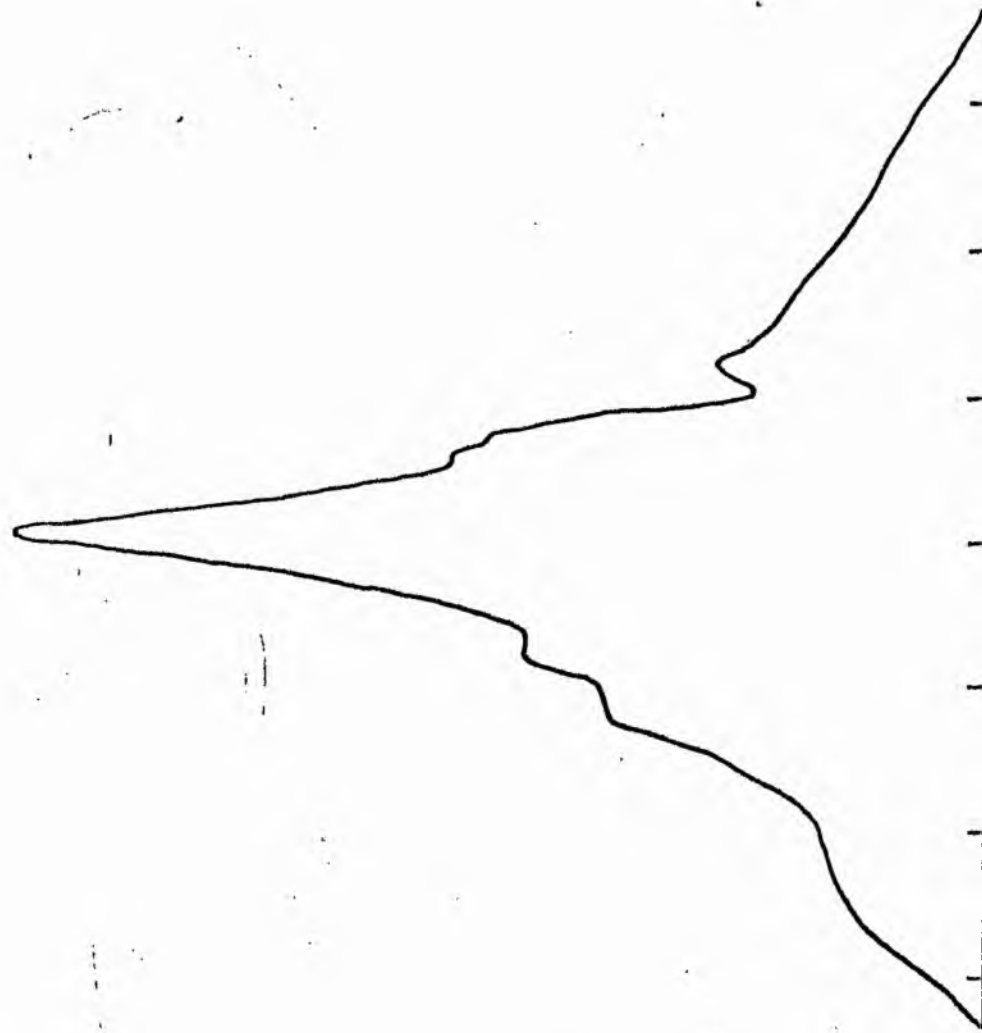
B-FILTER

Total luminosity	L_T	= 4.47
Total apparent magnitude	m_T	= 11.04
Apparent central surface brightness	μ_o	= 19.58
Major axis at threshold	$2a_m$	= 8.20
Minor axis at threshold	$2b_m$	= 4.96
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.04
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.66
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.57
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.75
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.80
Axis ratio	b/a	= 0.26
Equivalent radius	r_1^*	= 0.48
Surface brightness	μ_1	= 21.51
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.70
Axis ratio	b/a	= 0.27
Equivalent radius	r_e^*	= 0.90
Surface brightness	μ_e	= 22.81
Mean surface brightness	μ_e'	= 12.80
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.70
Axis ratio	b/a	= 0.43
Equivalent radius	r_3^*	= 1.80
Surface brightness	μ_3	= 24.31
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.87 \\ = 2.00 \end{cases}$

NGC 4293
V-Filter
Axis 1

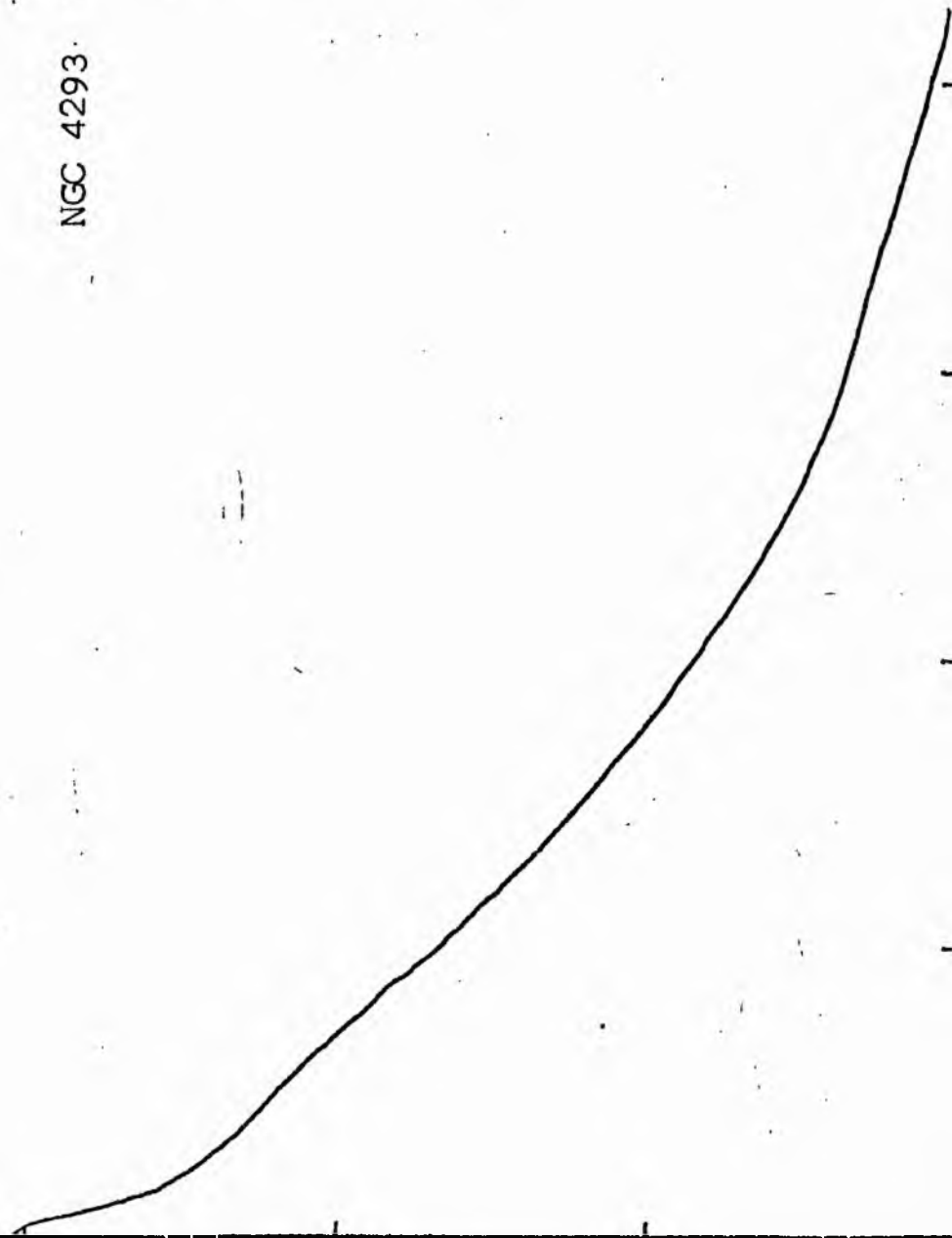


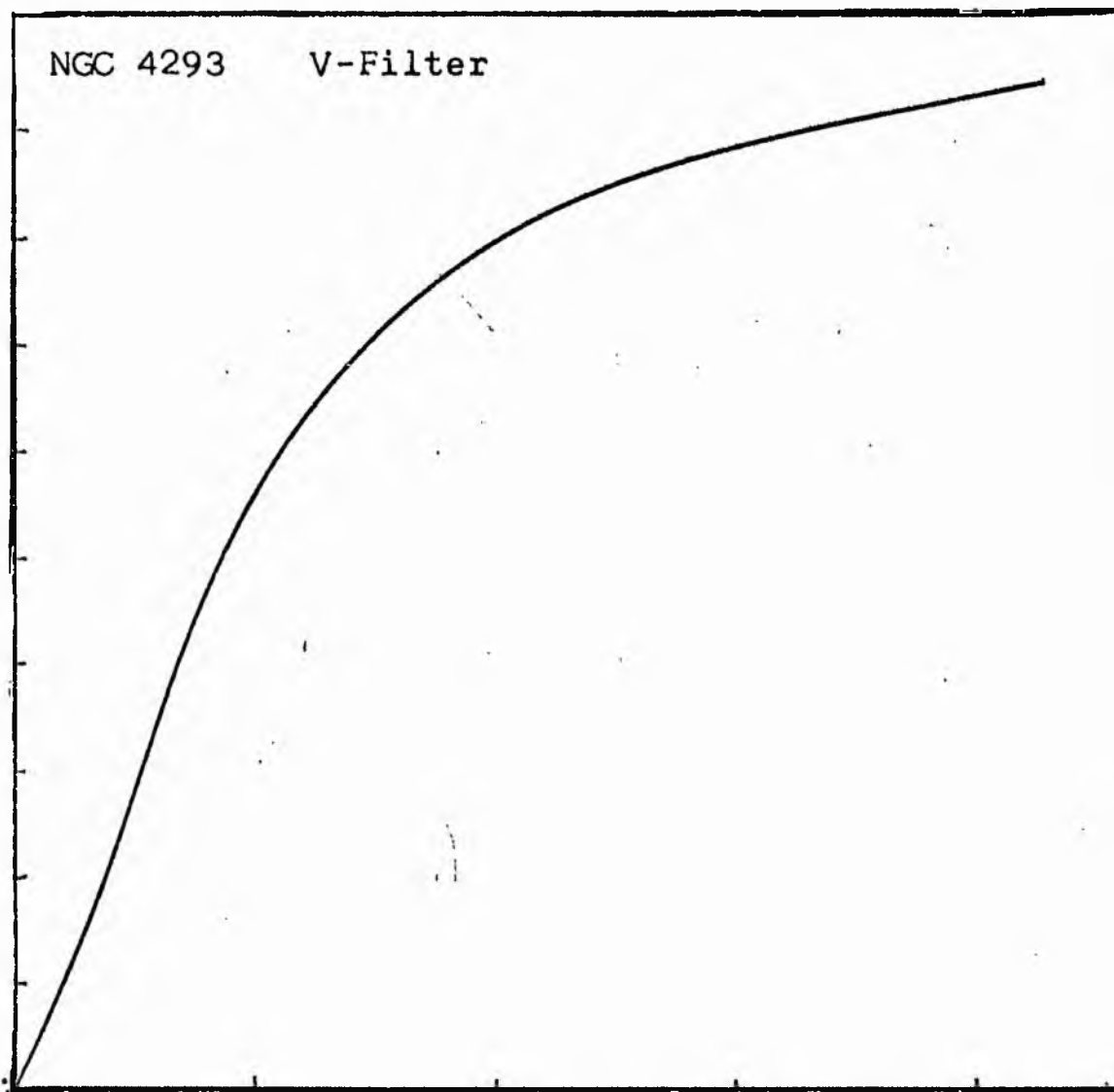
NGC 4293
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4293 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4293
V COLOUR

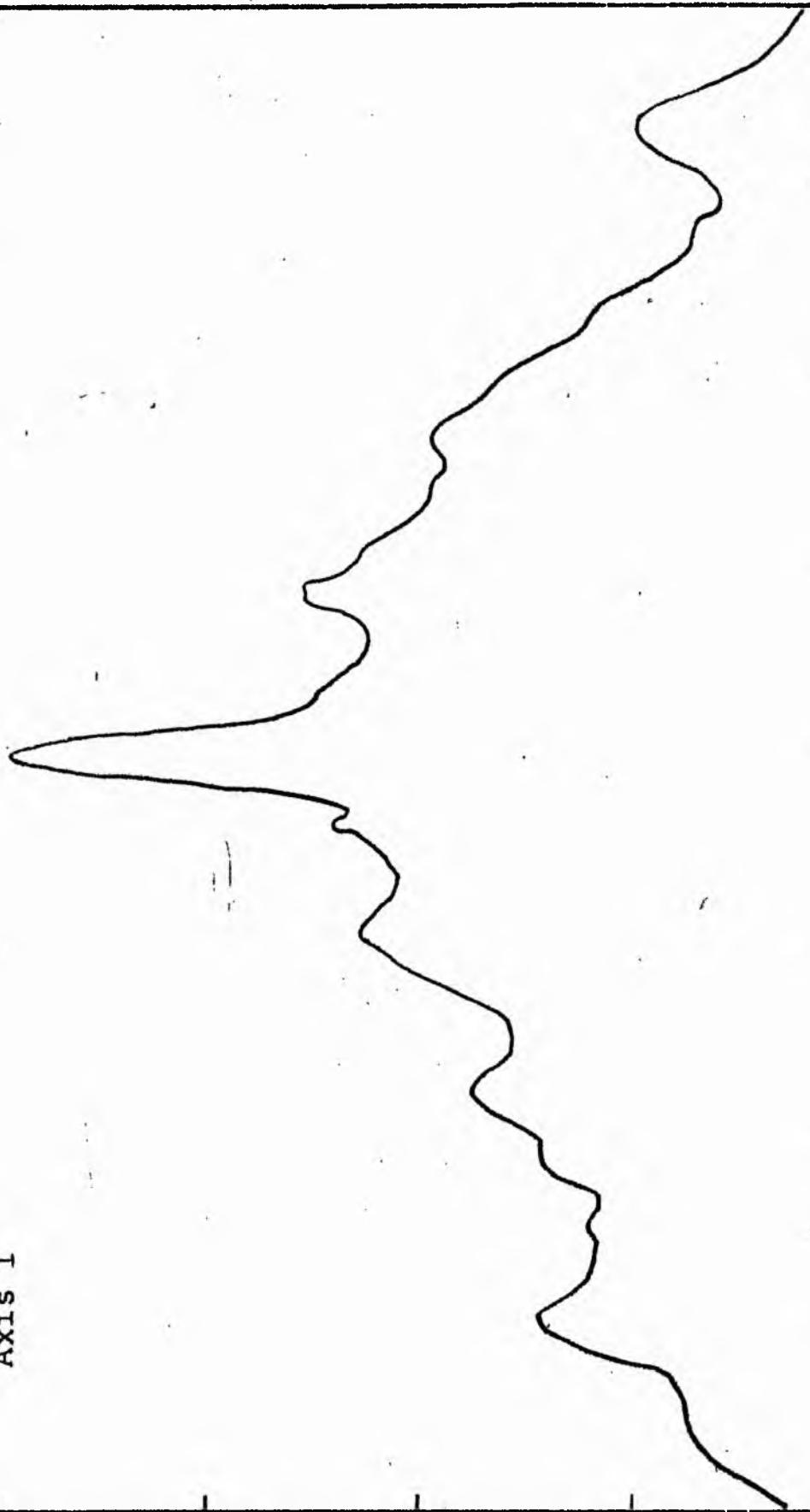
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J
1.04	10.965	10.482	0.0	0.0	19.63	205.8211	0.0	0.0	0.0	1.206
1.00	10.000	8.972	2.50	19.63	33.69	302.2693	205.82	0.01	0.06	1.166
0.90	7.943	7.126	4.12	53.33	65.11	464.0022	508.09	0.03	0.10	1.066
0.80	6.310	5.661	6.14	118.44	42.17	238.7086	972.09	0.05	0.14	0.966
0.70	5.012	4.496	7.15	160.61	91.04	409.3740	1210.80	0.07	0.17	0.866
0.60	3.981	3.572	8.95	251.65	114.79	409.9775	1620.18	0.09	0.21	0.766
0.50	3.162	2.837	10.80	366.44	148.28	420.6909	2030.15	0.11	0.26	0.666
0.40	2.512	2.254	12.80	514.72	230.34	519.0925	2450.84	0.14	0.30	0.566
0.30	1.995	1.790	15.40	745.06	1273.80	2280.1946	2969.94	0.17	0.36	0.466
0.20	1.585	1.422	25.35	2018.86	601.40	855.1379	5250.13	0.29	0.60	0.366
0.10	1.259	1.129	28.88	2620.26	968.82	1094.2476	6105.27	0.34	0.68	0.266
0.00	1.000	0.897	33.80	3589.08	1312.59	1177.6094	7199.51	0.40	0.80	0.166
-0.10	0.794	0.713	39.50	4901.67	1125.29	801.9265	8377.12	0.47	0.93	0.066
-0.20	0.631	0.566	43.80	6026.95	1271.72	719.8848	9179.05	0.51	1.03	-0.034
-0.30	0.501	0.450	48.20	7298.67	649.84	292.1970	9898.93	0.55	1.14	-0.134
-0.40	0.398	0.357	50.30	7948.51	1833.27	654.7864	10191.12	0.57	1.19	-0.234
-0.50	0.316	0.284	55.80	9781.78	2568.73	728.7700	10845.91	0.61	1.32	-0.334
-0.60	0.251	0.225	62.70	12350.51	3308.30	745.5525	11574.68	0.65	1.48	-0.434
-0.70	0.200	0.179	70.60	15658.81	4346.96	778.1411	12320.23	0.69	1.67	-0.534
-0.80	0.158	0.142	79.80	20005.77	1950.68	277.3699	13098.37	0.73	1.88	-0.634
-0.90	0.126	0.113	83.60	21956.46	2593.70	292.9497	13375.74	0.75	1.97	-0.734
-1.00	0.100	0.090	88.40	24550.16	3982.05	357.2566	13668.69	0.77	2.09	-0.834
-1.10	0.079	0.071	95.30	28532.21	5906.20	420.9023	14025.94	0.79	2.25	-0.934
-1.20	0.063	0.057	104.70	34438.41	3436.74	194.5455	14446.84	0.81	2.47	-1.034
-1.30	0.050	0.045	109.80	37875.15	5271.25	237.2917	14641.39	0.82	2.59	-1.134
-1.40	0.040	0.036	117.20	43152.40	12002.19	428.6824	14878.68	0.83	2.77	-1.234
-1.50	0.032	0.028	132.50	55154.59	12916.91	366.4668	15307.36	0.86	3.13	-1.334
-1.60	0.025	0.023	147.20	68071.50	15397.44	346.9963	15673.82	0.88	3.48	-1.434
-1.70	0.020	0.018	163.00	83468.94	15399.31	275.6628	16020.82	0.90	3.85	-1.534
-1.80	0.016	0.014	177.40	98868.25	12522.81	178.0655	16296.48	0.91	4.19	-1.634
-1.90	0.013	0.011	188.30	111391.06	19349.44	218.5479	16474.54	0.92	4.45	-1.734
-2.00	0.010		204.00	130740.50			16693.09	0.94	4.82	-1.834
-∞							17843.00	(1)		

PHOTOMETRIC PARAMETERS OF NGC 4293

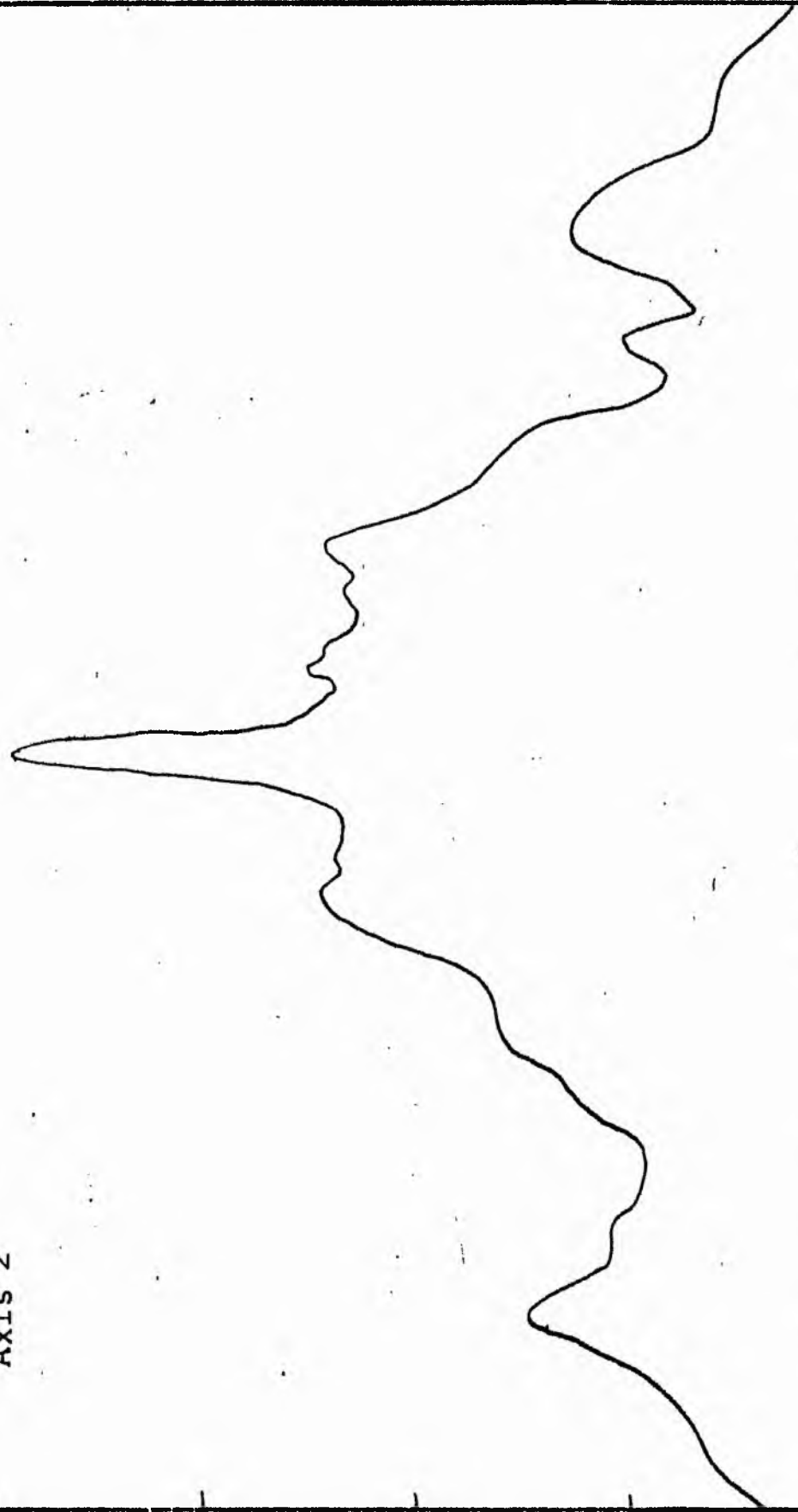
V-FILTER

Total luminosity	L_T	= 4.96
Total apparent magnitude	m_T	= ---
Apparent central surface brightness	μ_0	= ---
Major axis at threshold	$2a_m$	= 8.15
Minor axis at threshold	$2b_m$	= 5.77
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= ---
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= ---
Gradient of exponential component	$G(a)$	= -0.61
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.95
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.73
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.54
Axis ratio	b/a	= 0.45
Equivalent radius	r_1^*	= 0.36
Surface brightness	μ_1	= ---
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.19
Axis ratio	b/a	= 0.30
Equivalent radius	r_e^*	= 0.71
Surface brightness	μ_e	= ---
Mean surface brightness	μ_e'	= ---
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.19
Axis ratio	b/a	= 0.41
Equivalent radius	r_3^*	= 1.40
Surface brightness	μ_3	= ---
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.93 \\ = 1.98 \end{matrix}$

NGC 4303
B-Filter
Axis 1

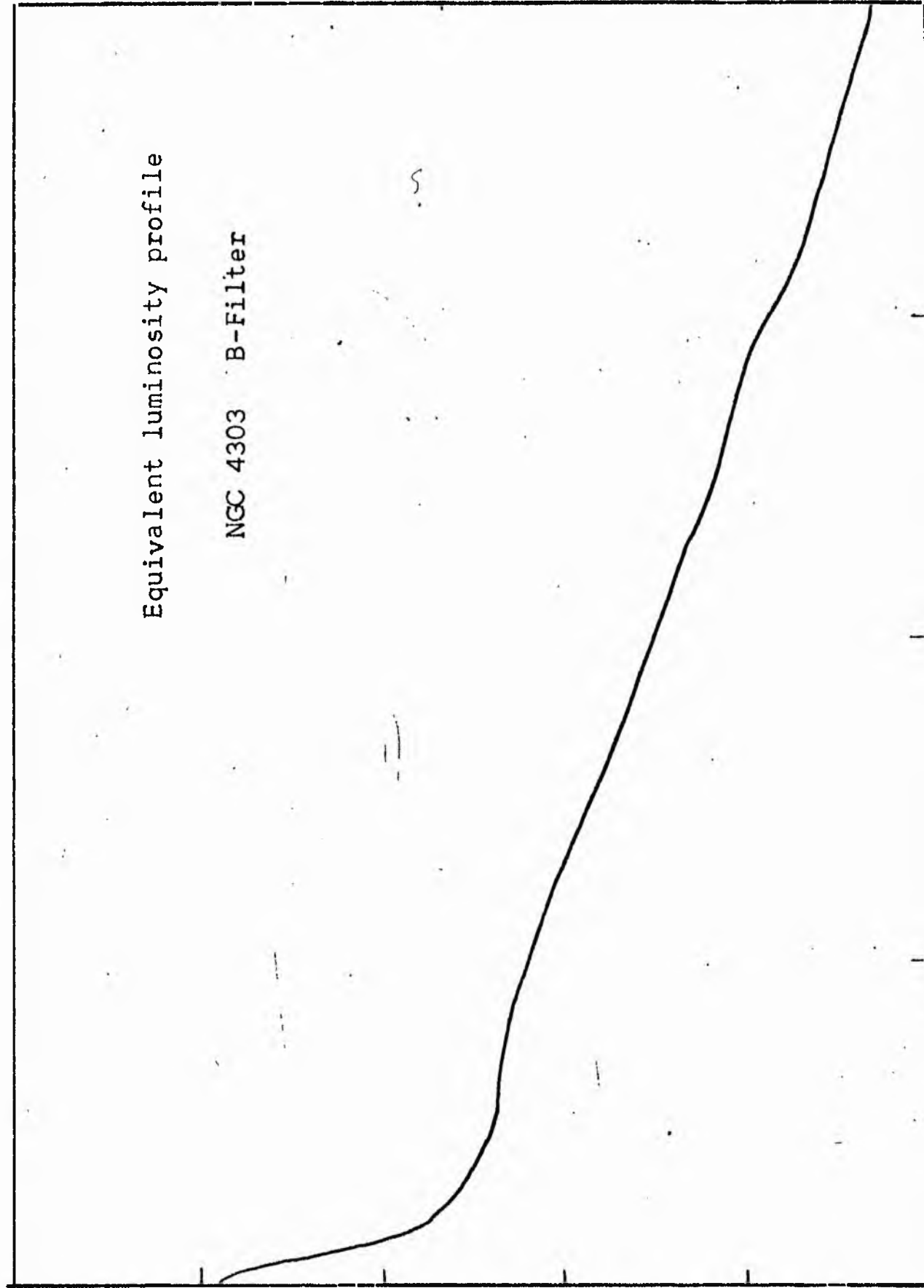


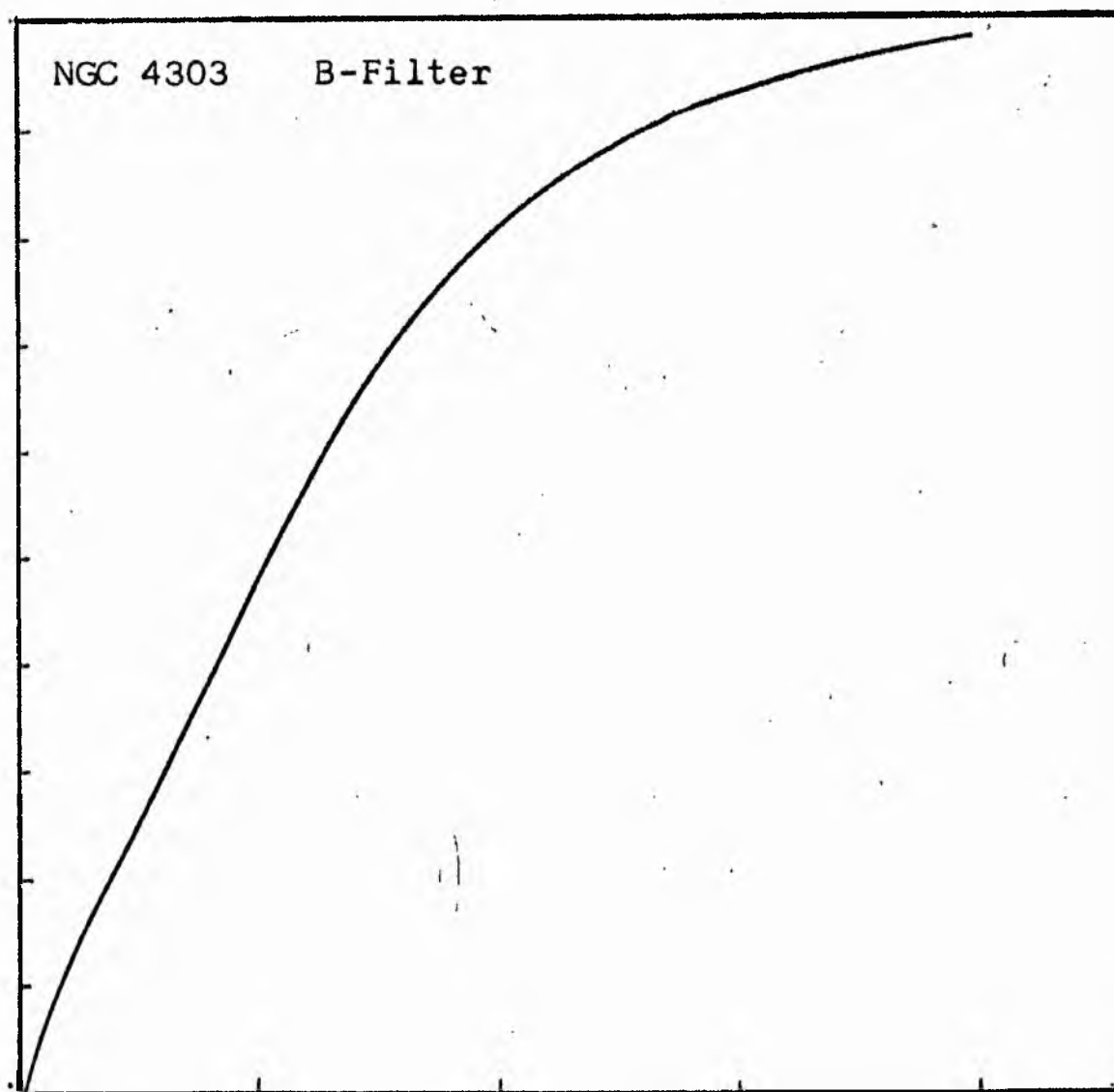
NGC 4303
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4303 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4303
B COLOUR

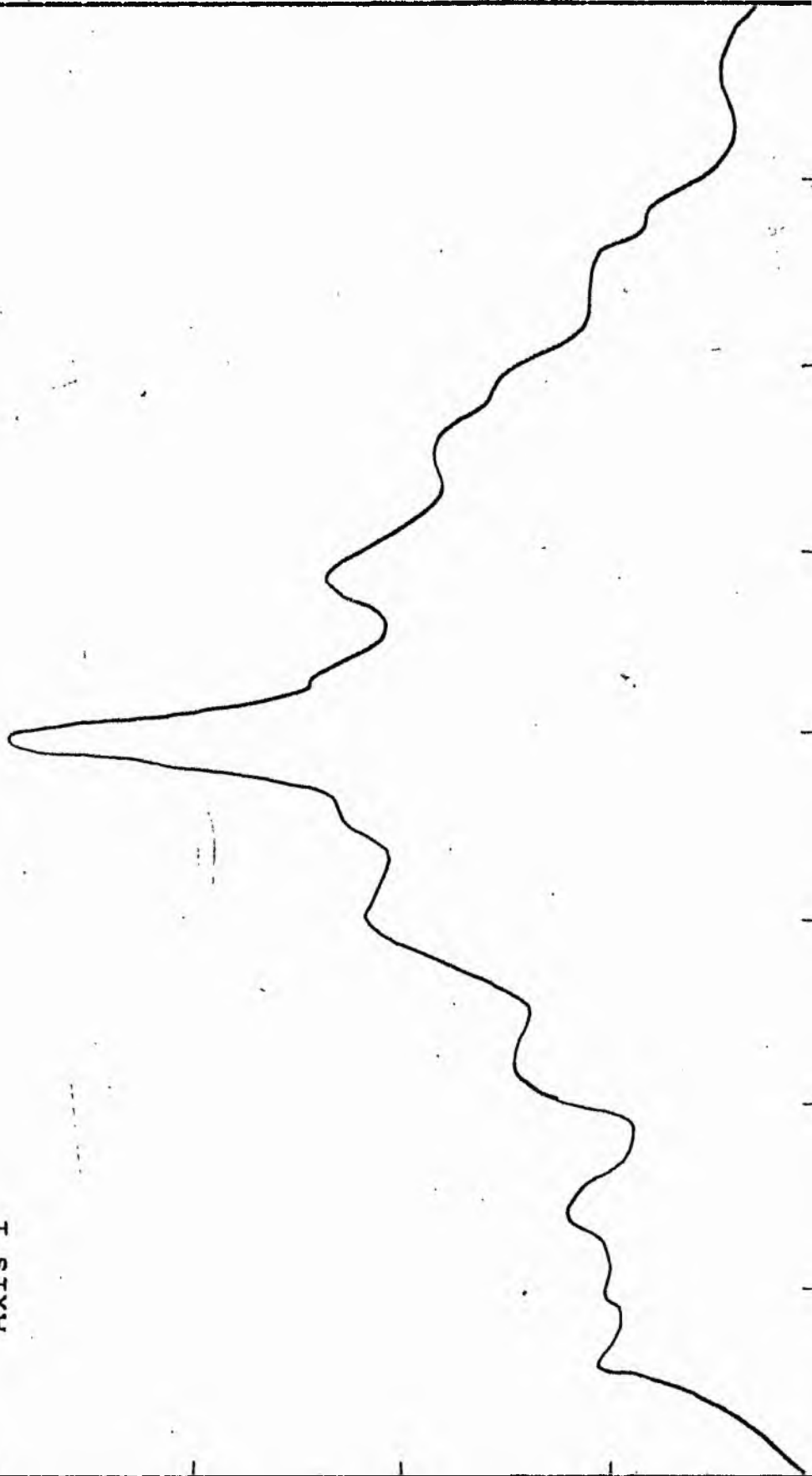
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.92	83.176		0.0	0.0			0.0	0.0	0.0	1.725	17.07
1.90	79.433	81.304	0.70	1.54	1.54	125.1585	125.16	0.00	0.01	1.705	17.12
1.80	63.096	71.264	1.40	6.16	4.62	329.1074	454.27	0.01	0.03	1.605	17.37
1.70	50.119	56.807	2.70	22.90	16.74	947.8679	1402.13	0.03	0.05	1.505	17.62
1.60	39.811	44.965	3.80	45.36	22.46	1010.0107	2412.14	0.05	0.07	1.405	17.87
1.50	31.623	35.717	4.55	65.04	19.67	702.6951	3114.84	0.06	0.09	1.305	18.12
1.40	25.119	28.371	4.80	72.38	7.34	208.3389	3323.18	0.07	0.09	1.205	18.37
1.30	19.952	22.536	5.50	95.03	22.65	510.4524	3833.63	0.08	0.11	1.105	18.62
1.20	15.849	17.901	6.17	119.60	24.56	439.7068	4273.34	0.09	0.12	1.005	18.87
1.10	12.589	14.219	7.46	174.83	55.24	785.4236	5058.76	0.10	0.14	0.905	19.12
1.00	10.000	11.295	7.80	191.13	16.30	184.0989	5242.86	0.11	0.15	0.805	19.37
0.90	7.943	8.972	8.20	211.24	20.11	180.3844	5423.24	0.11	0.16	0.705	19.62
0.80	6.310	7.126	9.50	283.53	72.29	515.1499	5938.39	0.12	0.18	0.605	19.87
0.70	5.012	5.661	11.70	430.05	146.52	829.4243	6767.81	0.14	0.23	0.505	20.12
0.60	3.981	4.496	14.70	678.87	248.81	1118.7769	7886.58	0.16	0.28	0.405	20.37
0.50	3.162	3.572	19.40	1182.37	503.50	1798.3311	9684.91	0.20	0.38	0.305	20.62
0.40	2.512	2.837	23.40	1720.21	537.84	1525.8865	11210.79	0.23	0.45	0.205	20.87
0.30	1.995	2.254	40.70	5204.02	3483.81	7850.9570	19061.75	0.39	0.79	0.105	21.12
0.20	1.585	1.790	51.30	8267.69	3063.68	5484.1758	24545.93	0.50	0.99	0.005	21.37
0.10	1.259	1.422	57.40	10350.79	2083.10	2961.9541	27507.88	0.56	1.11	-0.095	21.62
-0.00	1.000	1.129	64.60	13110.36	2759.57	3116.8049	30624.68	0.62	1.25	-0.195	21.87
-0.10	0.794	0.897	70.69	15698.77	2588.41	2322.2153	32946.90	0.67	1.37	-0.295	22.12
-0.20	0.631	0.713	76.64	18452.73	2753.96	1962.5796	34909.48	0.71	1.48	-0.395	22.37
-0.30	0.501	0.566	81.60	20918.46	2465.73	1395.7715	36305.25	0.73	1.58	-0.495	22.62
-0.40	0.398	0.455	87.68	24151.87	3233.40	1453.8804	37759.12	0.76	1.70	-0.595	22.87
-0.50	0.316	0.357	102.49	32999.90	8848.04	3160.2097	40919.33	0.83	1.98	-0.695	23.12
-0.60	0.251	0.284	107.50	36305.03	3305.13	937.6863	41857.02	0.85	2.08	-0.795	23.37
-0.70	0.200	0.225	114.50	41187.07	4882.04	1100.1960	42957.21	0.87	2.22	-0.895	23.62
-0.80	0.158	0.179	121.36	46270.14	5083.07	909.9043	43867.11	0.89	2.35	-0.995	23.87
-0.90	0.126	0.142	136.80	58792.50	12522.36	1780.5552	45647.67	0.92	2.65	-1.095	24.12
-1.00	0.100	0.113	143.50	64692.46	5899.96	666.3750	46314.04	0.94	2.78	-1.195	24.37
-1.10	0.079	0.090	146.50	67425.62	2733.16	245.2084	46559.25	0.94	2.84	-1.295	24.62
-1.20	0.063	0.071	150.20	70874.44	3448.81	245.7765	46805.02	0.95	2.91	-1.395	24.87
-1.30	0.050	0.057	159.60	80023.12	9148.69	517.8811	47322.90	0.96	3.09	-1.495	25.12
-1.40	0.040	0.045	168.00	88668.31	8645.19	388.7285	47711.63	0.96	3.25	-1.595	25.37
-1.50	0.032	0.036	176.40	97756.81	9088.50	324.6121	48036.24	0.97	3.41	-1.695	25.62
-1.60	0.025	0.028	186.00	111036.44	13279.62	376.7549	48412.99	0.98	3.64	-1.795	25.87
-1.70	0.020	0.023	194.30	118602.94	7566.50	170.5174	48583.51	0.98	3.76	-1.895	26.12
-1.80	0.016	0.018	200.40	126166.81	7563.87	135.4000	48718.91	0.99	3.88	-1.995	26.37
-1.90	0.013	0.014	206.50	133964.56	7797.75	110.8777	48829.78	0.99	4.00	-2.095	26.62
-2.00	0.010	0.011	212.30	141595.56	7631.00	86.1900	48915.97	0.99	4.11	-2.195	26.87
-∞							49455.00	111			∞

PHOTOMETRIC PARAMETERS OF NGC 4303

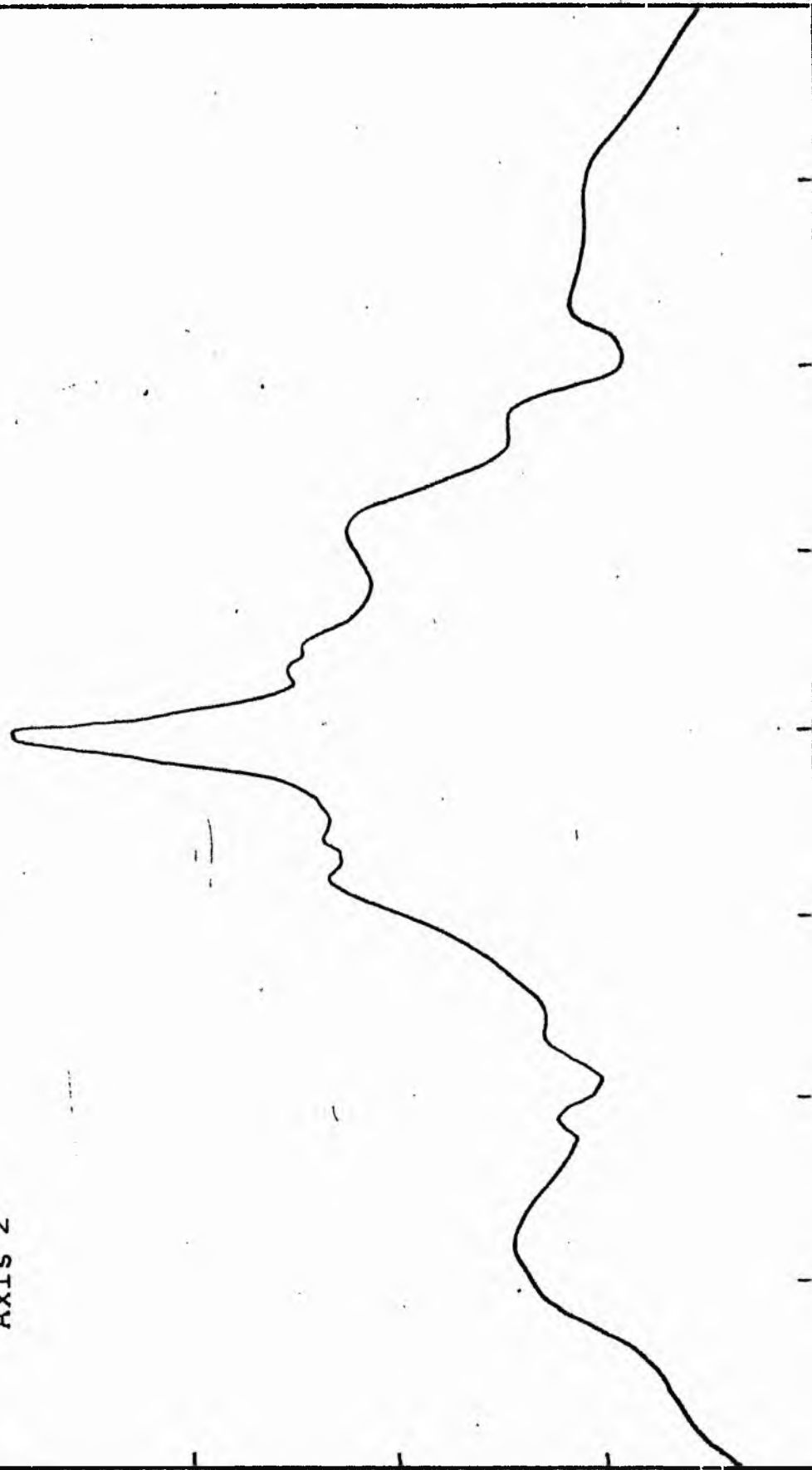
B-FILTER

Total luminosity	L_T	= 13.74
Total apparent magnitude	m_T	= 10.13
Apparent central surface brightness	μ_0	= 17.07
Major axis at threshold	$2a_m$	= 7.18
Minor axis at threshold	$2b_m$	= 7.25
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.46
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.95
Gradient of exponential component	$G(a)$	= -0.93
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.76
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.71
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.36
Axis ratio	b/a	= 1.13
Equivalent radius	r_1^*	= 0.44
Surface brightness	μ_1	= 20.90
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.77
Axis ratio	b/a	= 1.17
Equivalent radius	r_e^*	= 0.86
Surface brightness	μ_e	= 21.37
Mean surface brightness	μ_e'	= 11.78
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.53
Axis ratio	b/a	= 0.77
Equivalent radius	r_3^*	= 1.40
Surface brightness	μ_3	= 22.79
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.97 \\ = 1.64 \end{cases}$

NGC 4303
V-Filter
Axis 1

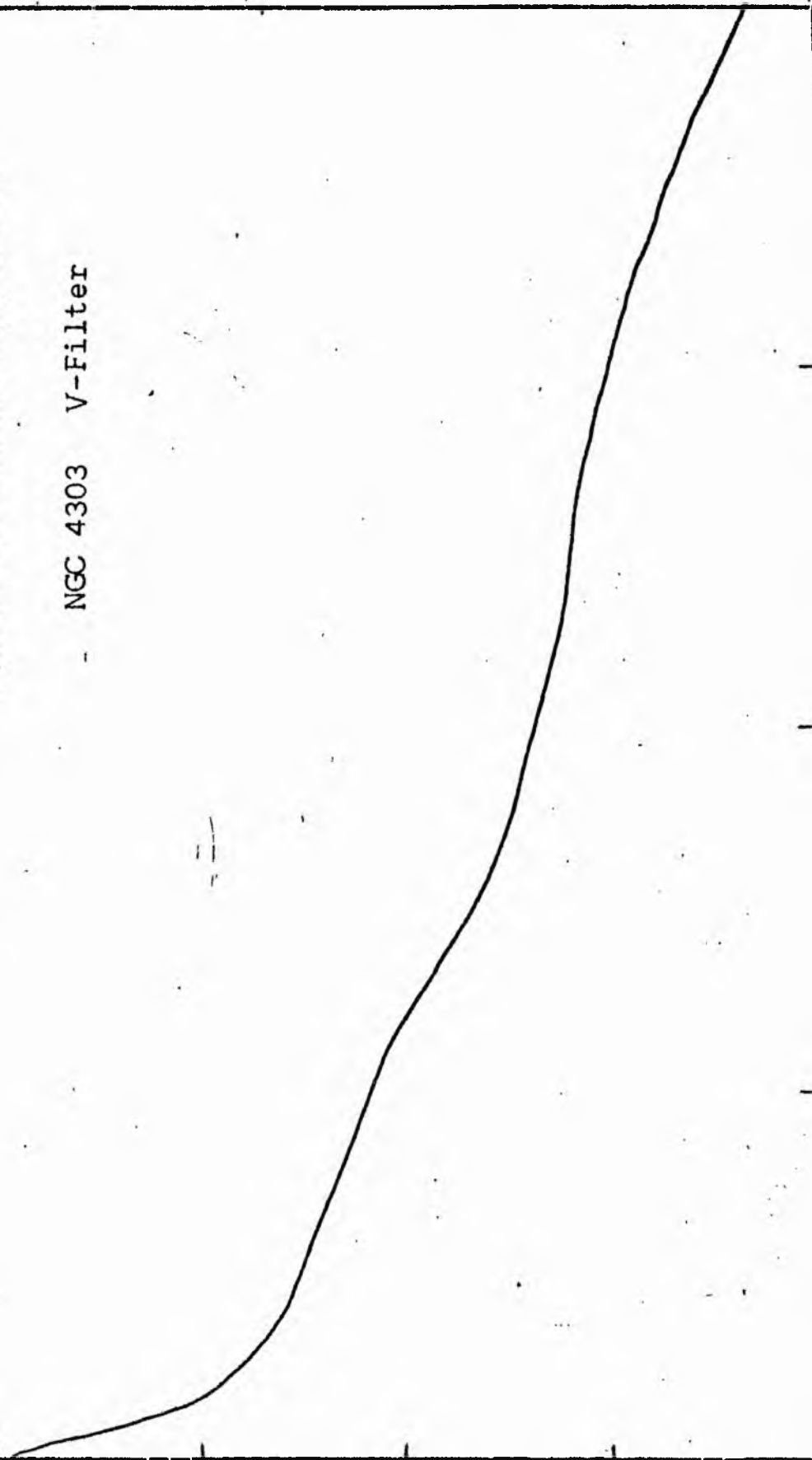


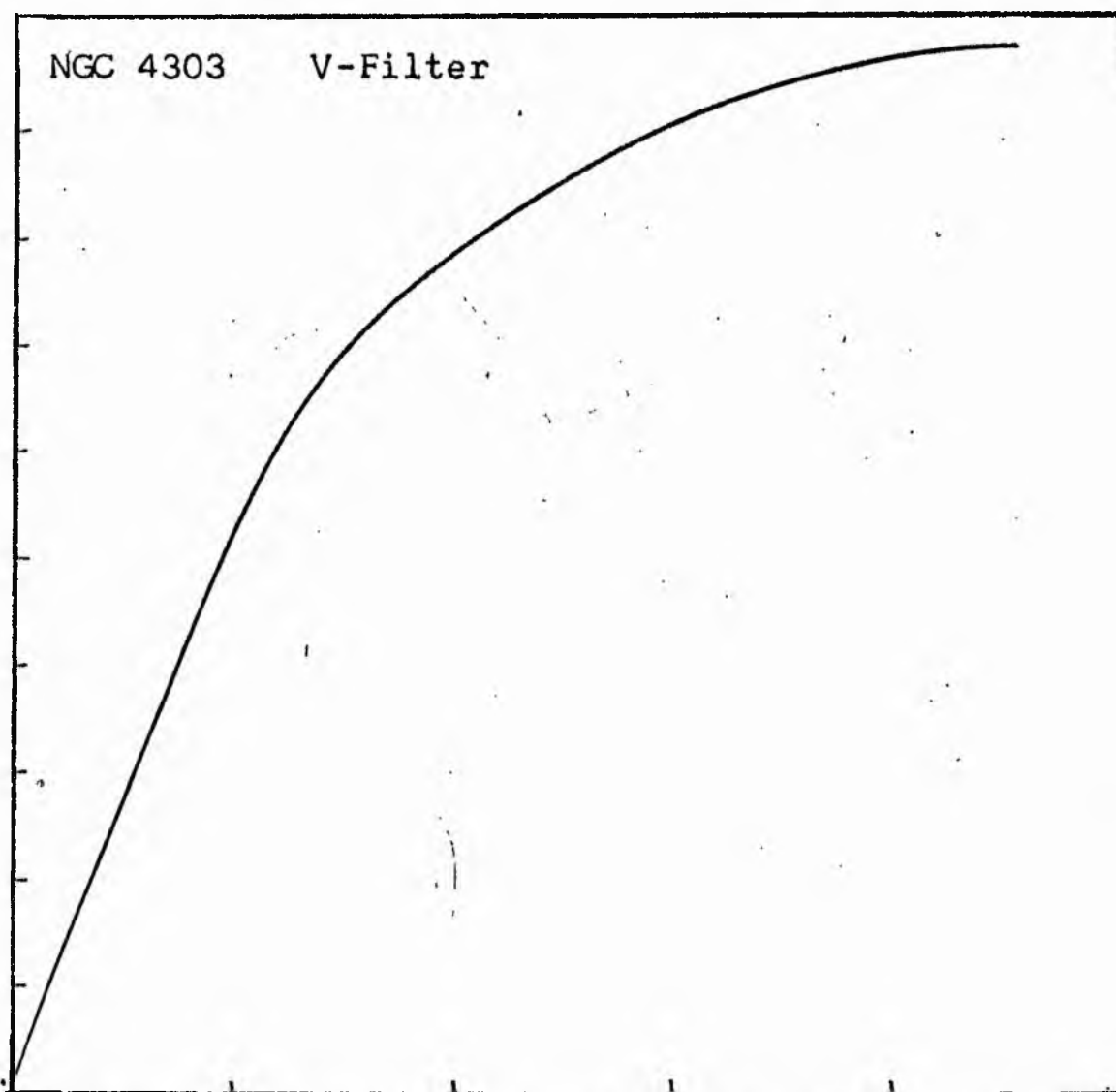
NGC 4303
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4303 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4303
V COLOUR

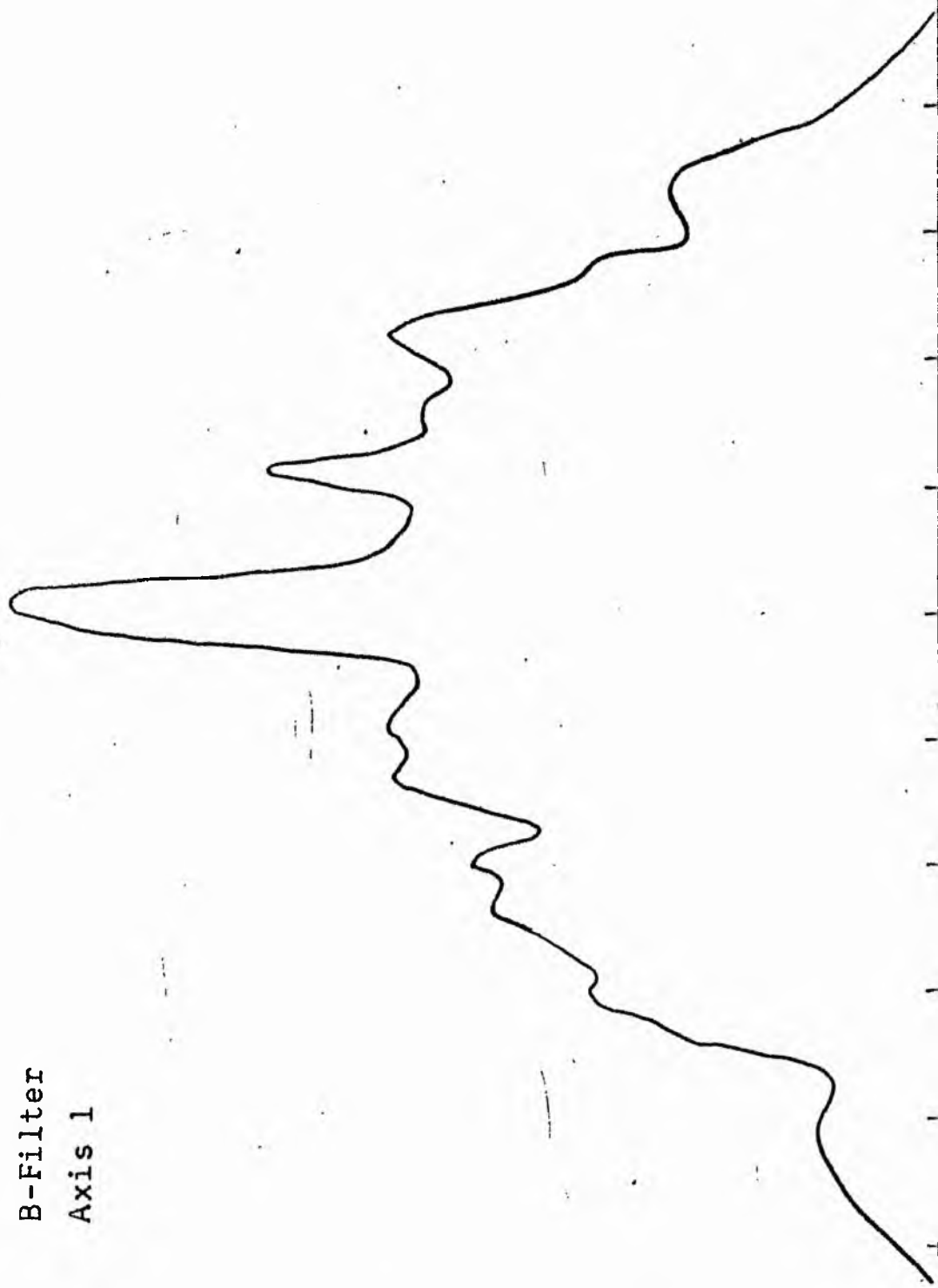
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.82	66.069		0.0	0.0			0.0	0.0	0.0	1.653	16.77
		64.582			2.01	129.8507					
1.80	63.096	56.607	0.80	2.01	7.07	400.1316	129.85	0.30	0.02	1.633	16.82
1.70	50.119	44.965	1.70	9.08	4.78	214.7156	529.98	0.01	0.04	1.533	17.07
1.60	39.811	35.717	2.10	13.85	31.51	1125.4358	744.70	0.02	0.04	1.433	17.32
1.50	31.623	28.371	3.80	45.36	7.45	211.2350	1870.13	0.04	0.08	1.333	17.57
1.40	25.119	22.536	4.10	52.81	19.57	441.0691	2081.37	0.04	0.09	1.233	17.82
1.30	19.952	17.901	4.80	72.38	29.69	531.4358	2522.44	0.05	0.10	1.133	18.07
1.20	15.849	14.219	5.70	102.07	74.64	1061.3655	3153.87	0.07	0.12	1.033	18.32
1.10	12.589	11.295	7.50	176.71	34.53	389.9556	4115.24	0.09	0.16	0.933	18.57
1.00	10.000	8.972	8.20	211.24	78.29	702.3691	4505.19	0.10	0.17	0.833	18.82
0.90	7.943	7.126	9.60	289.53	90.60	645.6738	5207.56	0.11	0.20	0.733	19.07
0.80	6.310	5.661	11.00	380.13	289.53	1638.9314	5853.23	0.13	0.23	0.633	19.32
0.70	5.012	4.496	14.60	669.66	452.55	2034.8420	7492.16	0.16	0.31	0.533	19.57
0.60	3.981	3.572	18.90	1122.21	810.00	2893.0181	9527.00	0.20	0.40	0.433	19.82
0.50	3.162	2.837	24.80	1932.20	1096.61	3111.1538	12420.02	0.27	0.52	0.333	20.07
0.40	2.512	2.254	31.05	3028.81	1460.01	3290.2251	15531.17	0.33	0.65	0.233	20.32
0.30	1.995	1.790	37.80	4488.83	2043.67	3658.3049	18821.39	0.40	0.79	0.133	20.57
0.20	1.585	1.422	45.60	6532.50	1735.19	2467.2666	22479.70	0.48	0.96	0.033	20.82
0.10	1.259	1.129	51.30	8267.69	1903.56	2149.9829	24946.96	0.53	1.08	-0.067	21.07
-0.00	1.000	0.897	56.90	10171.25	2416.76	2168.2163	27096.94	0.58	1.19	-0.167	21.32
-0.10	0.794	0.713	63.30	12588.01	2499.45	1781.2021	29265.16	0.63	1.33	-0.267	21.57
-0.20	0.631	0.566	69.30	15087.46	2349.16	1329.7869	31046.36	0.66	1.45	-0.367	21.82
-0.30	0.501	0.450	74.50	17436.62	1093.23	491.5627	32376.14	0.69	1.56	-0.467	22.07
-0.40	0.398	0.357	76.80	18529.85	1677.00	598.9644	32867.70	0.70	1.61	-0.567	22.32
-0.50	0.316	0.284	80.20	20206.85	6673.41	1893.2881	33466.66	0.71	1.68	-0.667	22.57
-0.60	0.251	0.225	92.50	26880.25	9357.26	2108.7144	35359.95	0.76	1.94	-0.767	22.82
-0.70	0.200	0.179	107.40	36237.51	6988.56	1250.9988	37468.66	0.80	2.25	-0.867	23.07
-0.80	0.158	0.142	117.30	43226.07	13606.13	1934.6558	38719.66	0.83	2.46	-0.967	23.32
-0.90	0.126	0.113	134.50	56832.20	20407.74	2304.9656	40654.31	0.87	2.82	-1.067	23.57
-1.00	0.100	0.090	156.80	77239.94	6639.19	595.6418	42959.27	0.92	3.29	-1.167	23.82
-1.10	0.079	0.071	163.40	83879.12	7126.62	507.8723	43554.91	0.93	3.43	-1.267	24.07
-1.20	0.063	0.057	170.20	91005.75	7306.00	413.5718	44062.79	0.94	3.57	-1.367	24.32
-1.30	0.050	0.045	176.90	98311.75	7242.25	325.6458	44476.36	0.95	3.71	-1.467	24.57
-1.40	0.040	0.036	183.30	105554.00	3717.62	132.7817	44802.00	0.96	3.84	-1.567	24.82
-1.50	0.032	0.028	186.50	109271.62	4378.75	124.2291	44934.78	0.96	3.91	-1.667	25.07
-1.60	0.025	0.023	190.20	113650.37	13273.06	299.1194	45059.01	0.96	3.99	-1.767	25.32
-1.70	0.020	0.018	201.00	126923.44	8994.44	161.0083	45358.12	0.97	4.21	-1.867	25.57
-1.80	0.016	0.014	208.00	135917.87	11607.81	165.0537	45519.13	0.97	4.36	-1.967	25.82
-1.90	0.013	0.011	216.70	147525.69	14072.25	158.9422	45684.18	0.98	4.54	-2.067	26.07
-2.00	0.010		226.80	161597.94			45843.12	0.98	4.75	-2.167	26.32
-∞							46813.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NGC 4303

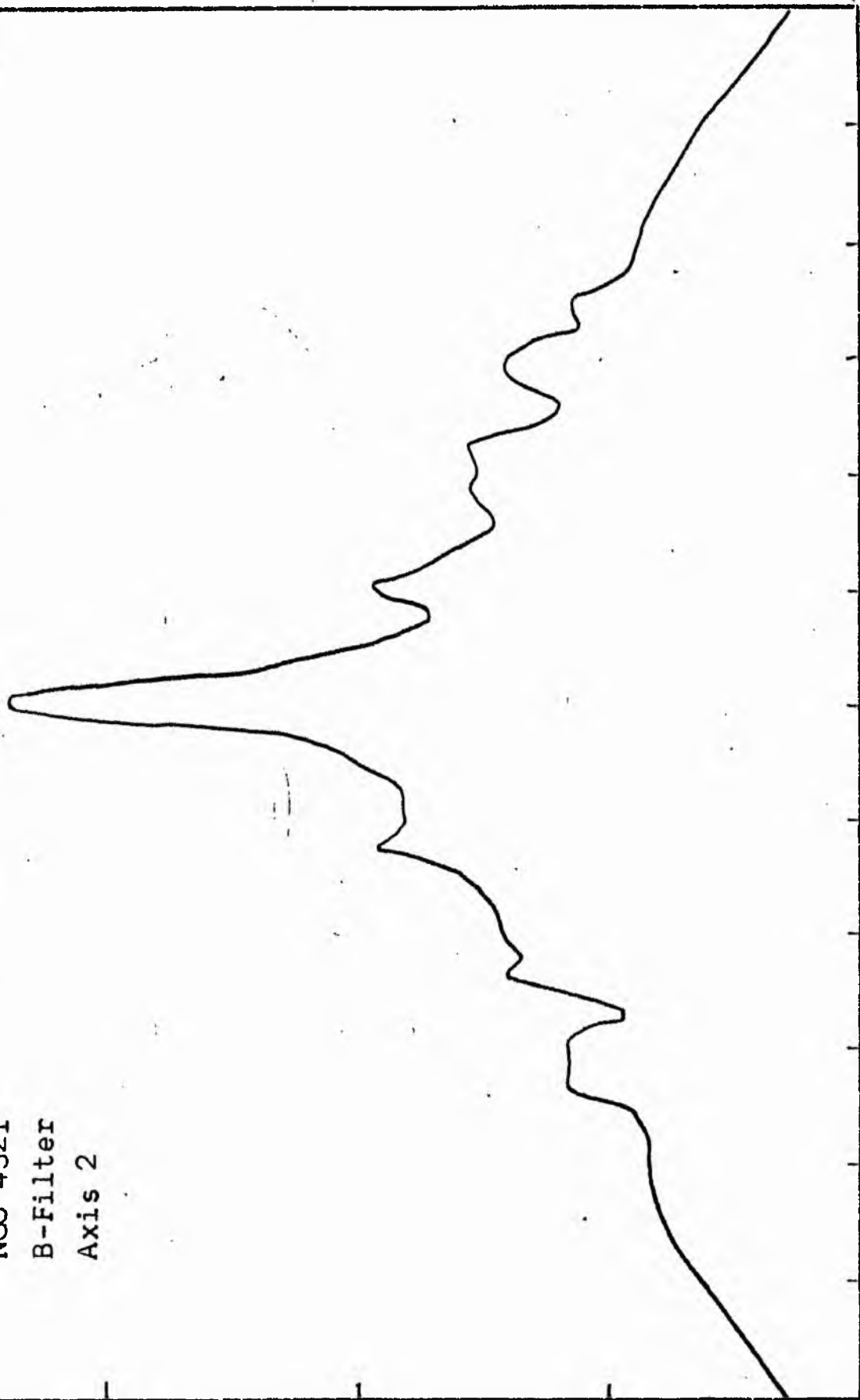
V-FILTER

Total luminosity	L_T	= 13.00
Total apparent magnitude	m_T	= 9.64
Apparent central surface brightness	μ_0	= 16.77
Major axis at threshold	$2a_m$	= 7.12
Minor axis at threshold	$2b_m$	= 7.23
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 5.80
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.96
Gradient of exponential component	$G(a)$	= -0.84
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.73
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.78
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.23
Axis ratio	b/a	= 1.27
Equivalent radius	r_1^*	= 0.39
Surface brightness	μ_1	= 19.99
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.67
Axis ratio	b/a	= 1.14
Equivalent radius	r_e^*	= 0.80
Surface brightness	μ_e	= 20.97
Mean surface brightness	μ_e'	= 11.14
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.48
Axis ratio	b/a	= 0.85
Equivalent radius	r_3^*	= 1.51
Surface brightness	μ_3	= 22.77
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.04 \\ = 1.90 \end{matrix}$

NGC 4321
B-Filter
Axis 1

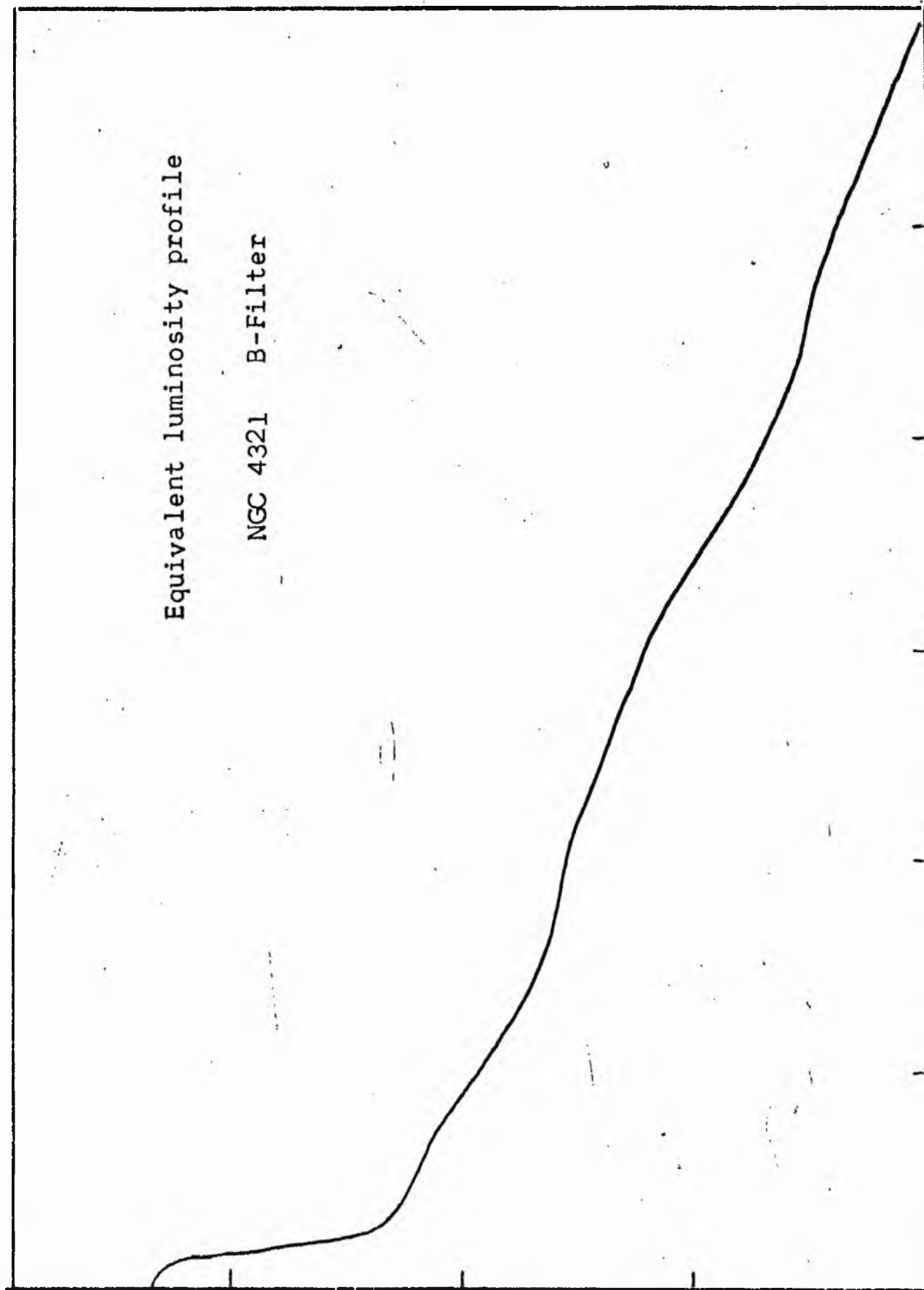


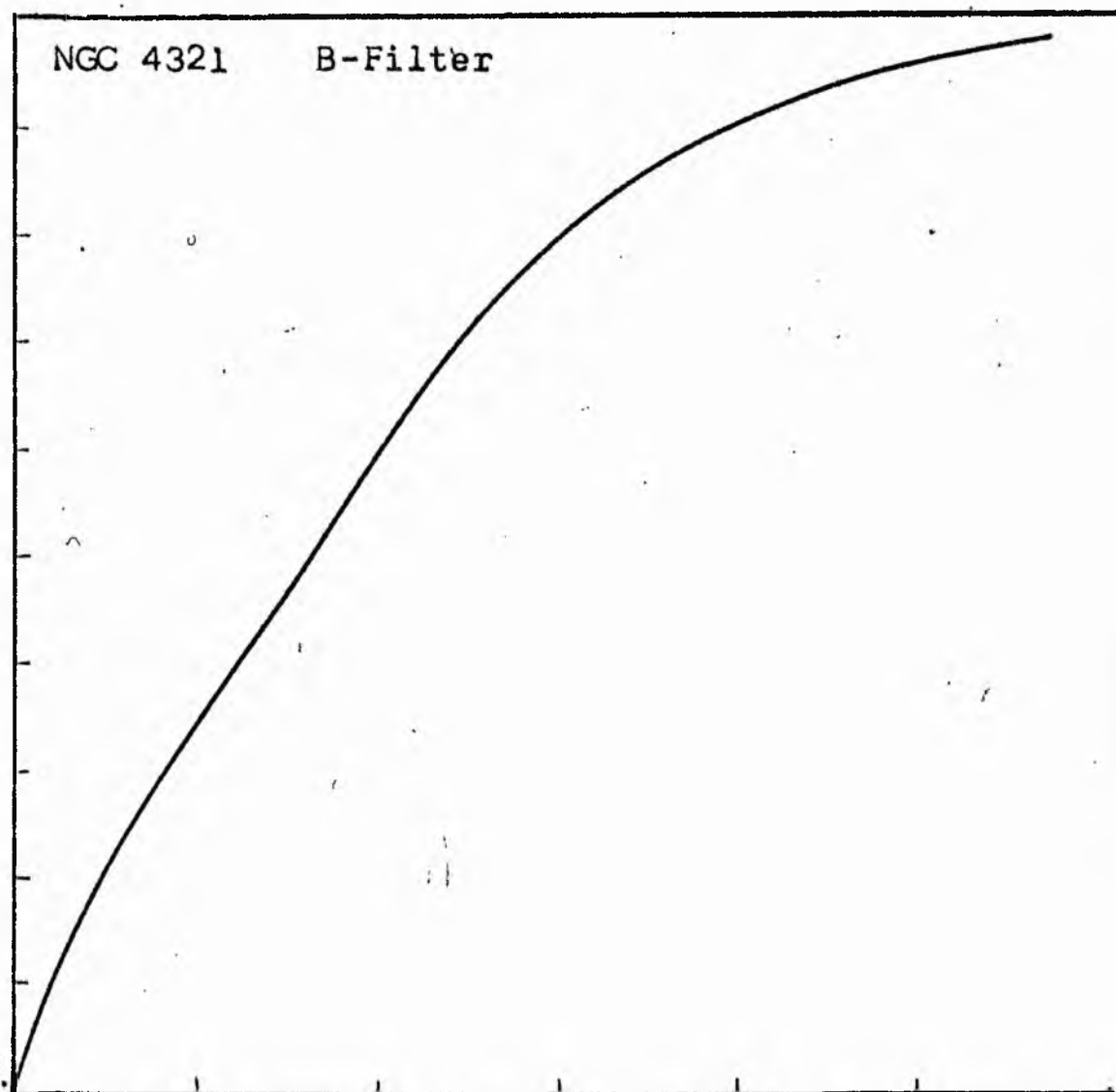
NGC 4321
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4321 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4321
B COLOUR

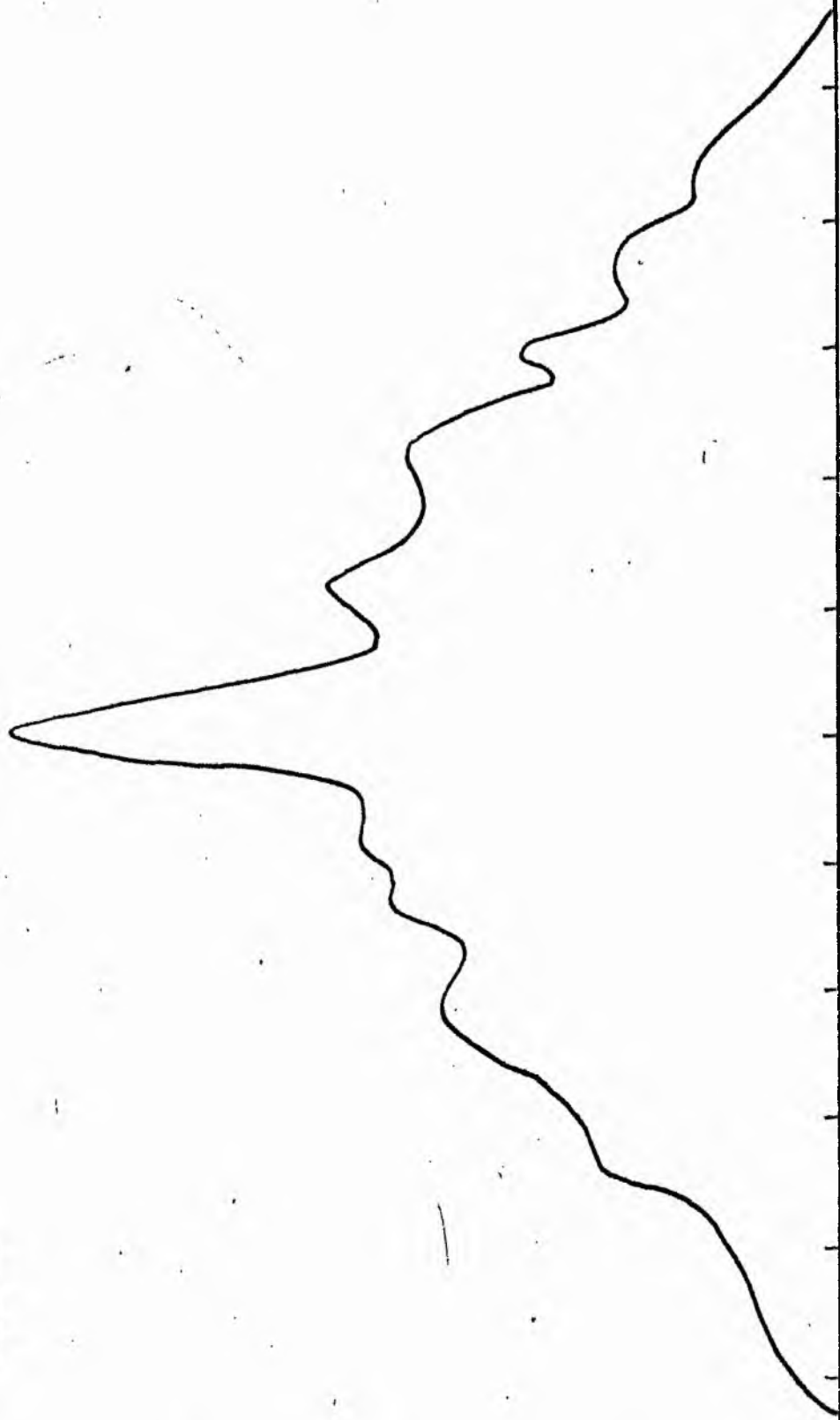
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.34	21.078		0.0	0.0	129.08	2499.7598	0.0	0.0	0.0	1.677	18.44
1.30	19.993	20.919	6.41	129.08	82.16	1470.6990	2699.76	0.06	0.08	1.637	18.54
1.20	15.849	17.901	6.20	211.24	15.74	223.7993	4170.46	0.09	0.10	1.537	18.79
1.10	12.589	14.219	6.50	226.98	21.87	246.9612	4394.25	0.09	0.10	1.437	19.04
1.00	10.000	11.295	6.90	248.85	71.83	642.6187	4641.21	0.10	0.11	1.337	19.29
0.90	7.943	8.972	10.10	320.47	92.78	376.1230	5283.83	0.11	0.12	1.237	19.54
0.80	6.310	7.126	10.90	373.25	20.83	117.9060	5659.95	0.12	0.13	1.137	19.79
0.70	5.012	5.661	11.20	394.08	43.35	194.9394	5777.86	0.12	0.14	1.037	20.04
0.60	3.981	4.496	11.80	437.44	69.27	247.4161	5972.79	0.12	0.15	0.937	20.29
0.50	3.162	3.572	12.70	506.71	49.01	139.0415	6220.21	0.13	0.16	0.837	20.54
0.40	2.512	2.837	13.30	555.72	218.66	492.7539	6359.25	0.13	0.16	0.737	20.79
0.30	1.995	2.254	15.70	774.37	1653.57	2960.0183	6852.00	0.14	0.19	0.637	21.04
0.20	1.585	1.790	27.80	2427.95	2374.95	3376.9502	9812.02	0.20	0.34	0.537	21.29
0.10	1.259	1.422	39.10	4802.89	2225.75	2513.8984	13188.96	0.27	0.48	0.437	21.54
-0.00	1.000	1.129	47.30	7028.65	1663.38	1492.3193	15702.86	0.33	0.58	0.337	21.79
-0.10	0.794	0.897	52.60	8692.03	2096.07	1493.7451	17195.18	0.36	0.65	0.237	22.04
-0.20	0.631	0.713	58.60	10788.10	6648.53	3763.5354	18688.92	0.39	0.72	0.137	22.29
-0.30	0.501	0.566	74.50	17436.62	11395.77	5124.0625	22452.46	0.47	0.92	0.037	22.54
-0.40	0.398	0.450	95.80	28832.40	9666.21	3452.4495	27576.52	0.57	1.18	-0.063	22.79
-0.50	0.316	0.357	110.70	38498.61	11615.08	3295.2871	31028.97	0.65	1.36	-0.163	23.04
-0.60	0.251	0.284	126.30	50113.70	11197.90	2523.5254	34324.25	0.71	1.56	-0.263	23.29
-0.70	0.200	0.225	139.70	61311.60	14272.28	2554.8433	36847.78	0.77	1.72	-0.363	23.54
-0.80	0.158	0.179	155.11	75583.87	7680.37	1092.0776	39402.62	0.82	1.91	-0.463	23.79
-0.90	0.126	0.142	162.80	83264.25	5879.69	664.0884	40494.69	0.84	2.01	-0.563	24.04
-1.00	0.100	0.113	168.45	89143.94	10226.56	917.4912	41158.78	0.86	2.08	-0.663	24.29
-1.10	0.079	0.090	177.85	99370.50	10276.44	732.3445	42076.27	0.88	2.19	-0.763	24.54
-1.20	0.063	0.071	186.82	109646.94	6575.25	372.2078	42808.61	0.89	2.30	-0.863	24.79
-1.30	0.050	0.057	192.34	116222.19	33202.12	1492.9304	43180.82	0.90	2.37	-0.963	25.04
-1.40	0.040	0.045	218.09	149424.31	27781.12	992.2568	44673.75	0.93	2.69	-1.063	25.29
-1.50	0.032	0.036	237.50	177205.44	11830.75	335.6504	45666.00	0.95	2.93	-1.163	25.54
-1.60	0.025	0.028	245.30	189036.19	24480.37	551.6882	46001.65	0.96	3.02	-1.263	25.79
-1.70	0.020	0.023	260.70	213516.56	22858.06	409.1812	46553.34	0.97	3.21	-1.363	26.04
-1.80	0.016	0.018	274.30	236374.62	17191.75	244.4539	46962.52	0.98	3.38	-1.463	26.29
-1.90	0.013	0.014	284.10	253566.37	24109.94	272.3159	47206.97	0.98	3.50	-1.563	26.54
-2.00	0.010	0.011	297.30	277676.31			47479.28	0.99	3.66	-1.663	26.79
-∞							48079.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4321

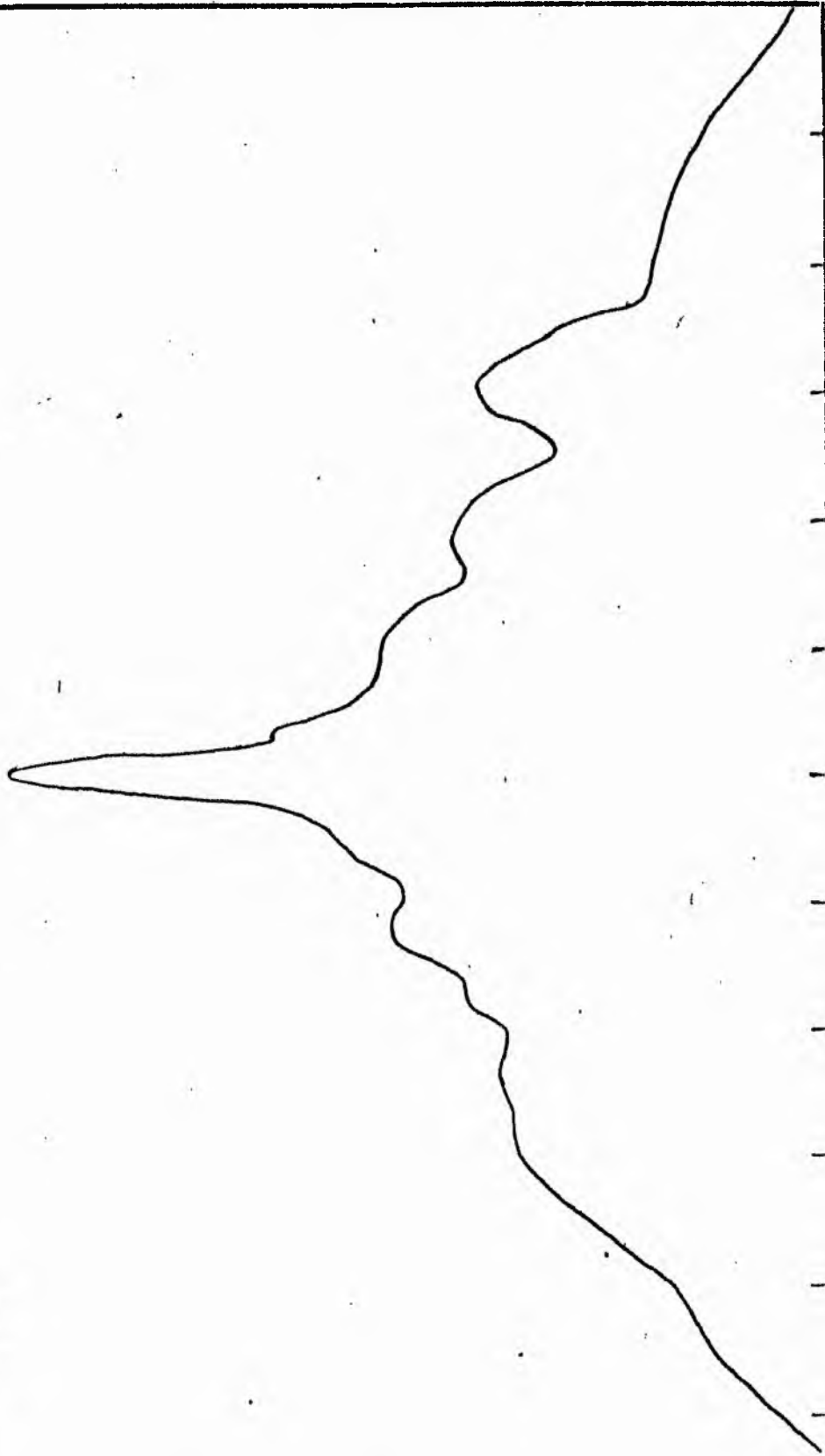
B-FILTER

Total luminosity	L_T	= 13.35
Total apparent magnitude	m_T	= 10.08
Apparent central surface brightness	μ_0	= 18.44
Major axis at threshold	$2a_m$	= 8.28
Minor axis at threshold	$2b_m$	= 11.50
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.96
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.90
Gradient of exponential component	$G(a)$	= -0.63
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.48
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.69
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.55
Axis ratio	b/a	= 0.48
Equivalent radius	r_1^*	= 0.58
Surface brightness	μ_1	= 21.47
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.68
Axis ratio	b/a	= 0.55
Equivalent radius	r_e^*	= 1.35
Surface brightness	μ_e	= 22.61
Mean surface brightness	μ_e'	= 12.73
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.25
Axis ratio	b/a	= 0.96
Equivalent radius	r_3^*	= 2.25
Surface brightness	μ_3	= 23.42
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.30 \\ = 1.67 \end{cases}$

NGC 4321
V-Filter
Axis 1

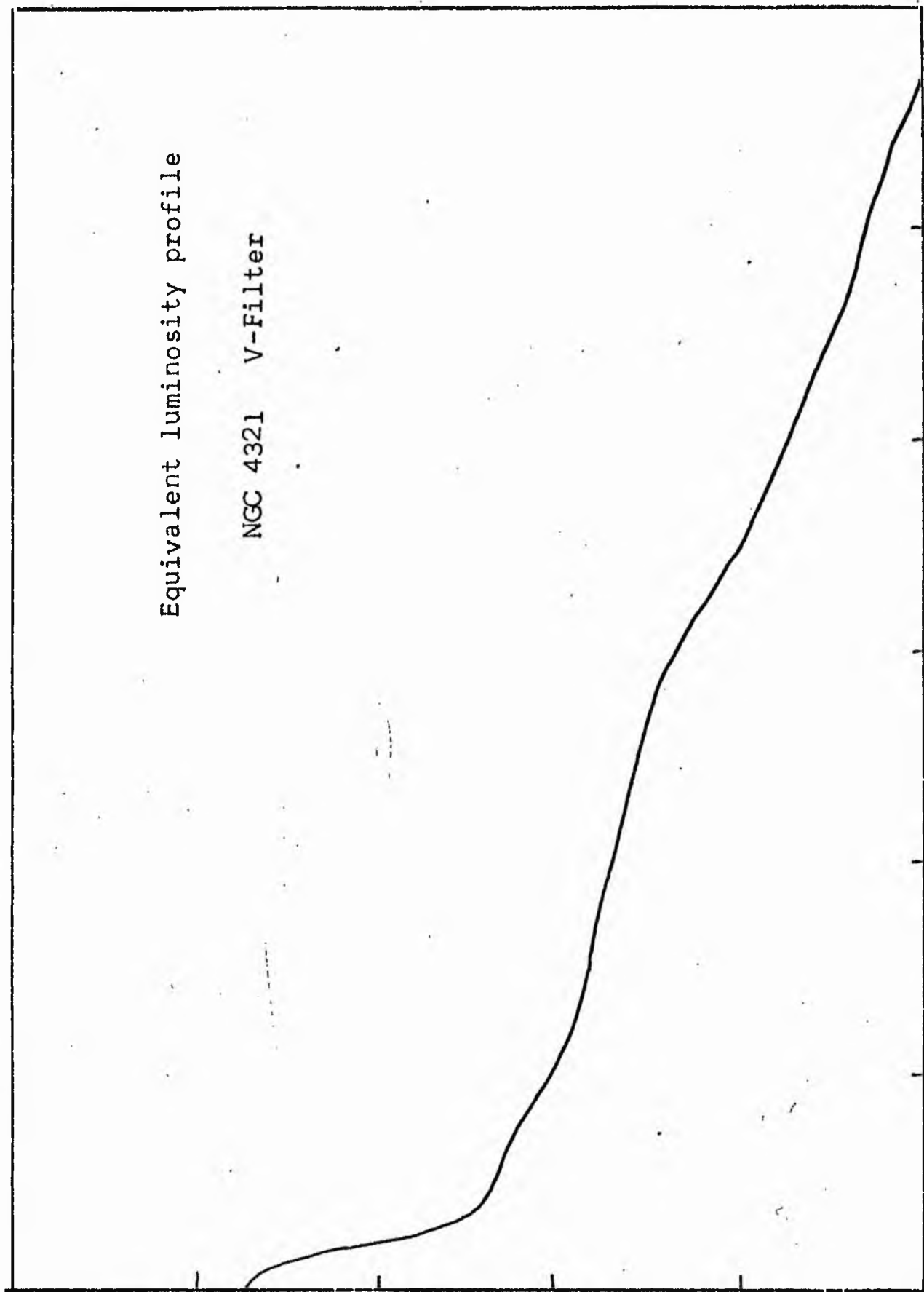


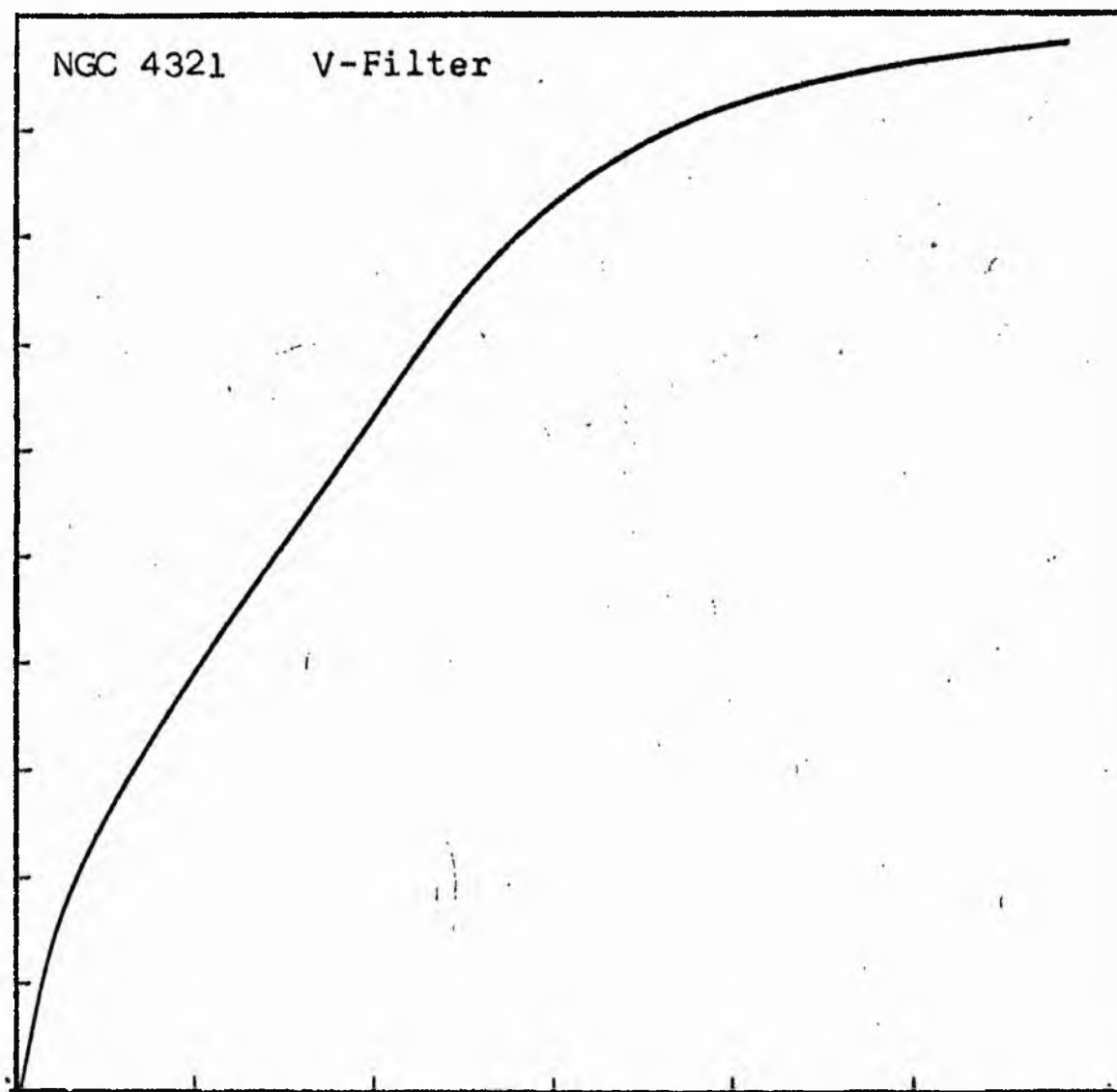
NGC 4321
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4321 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4321
V COLOUR

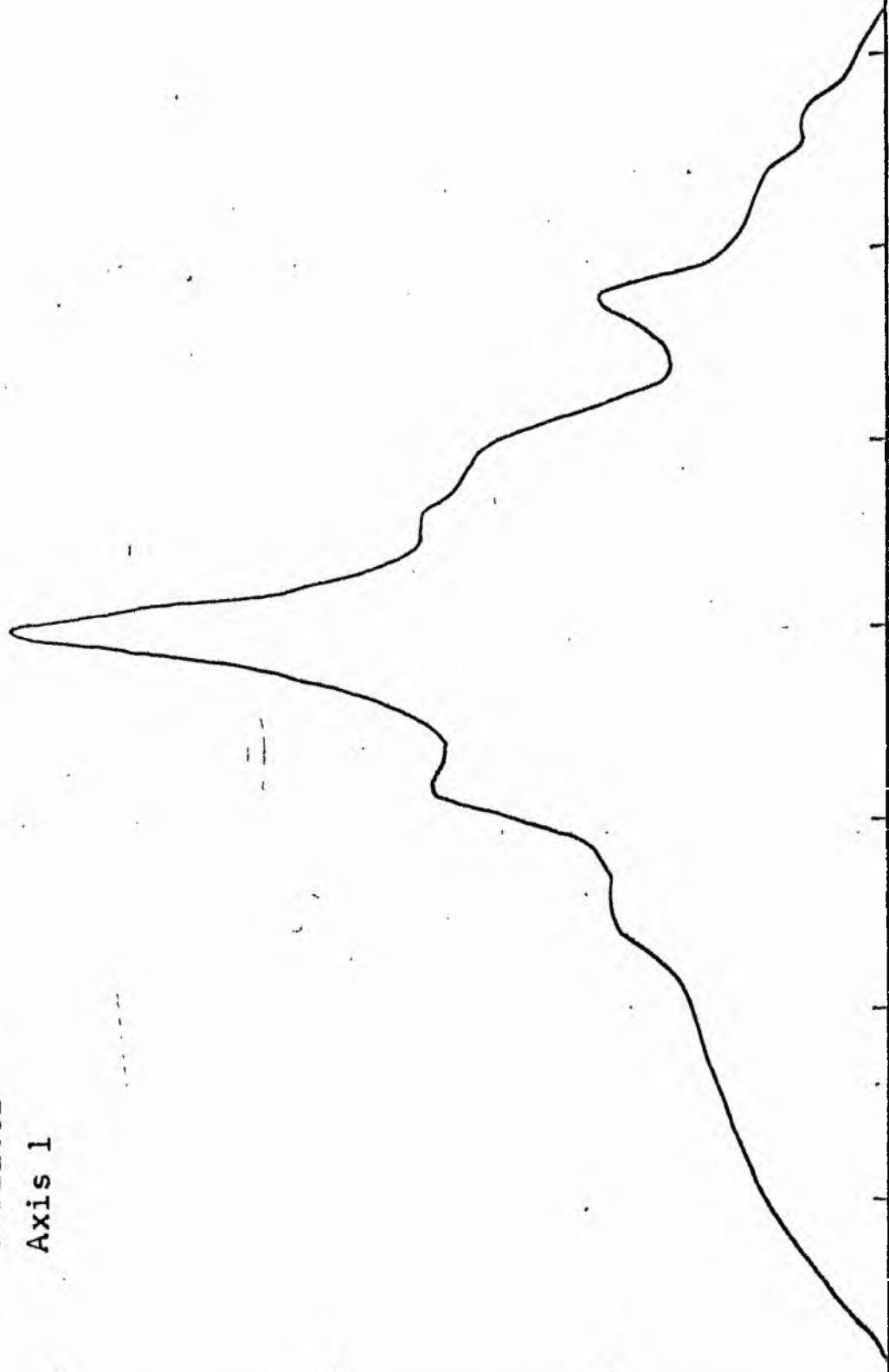
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.67	46.773		0.0	0.0			0.0	0.0	0.0	1.869	17.34
1.60	39.811	43.292	4.20	55.42	55.42	2399.1418	2399.14	0.04	0.06	1.799	17.52
1.50	31.623	35.717	6.10	116.90	61.48	2195.8962	4595.04	0.07	0.08	1.699	17.77
1.40	25.119	28.371	8.80	243.28	126.39	3585.6711	8180.70	0.13	0.12	1.599	18.02
1.30	19.953	22.536	9.20	265.90	22.62	509.7437	8690.45	0.14	0.13	1.499	18.27
1.20	15.849	17.901	9.60	289.53	23.62	422.8994	9113.34	0.15	0.13	1.399	18.52
1.10	12.589	14.219	10.10	320.47	30.94	440.0012	9553.34	0.16	0.14	1.299	18.77
1.00	10.000	11.295	11.20	394.08	73.61	831.3657	10384.71	0.17	0.15	1.199	19.02
0.90	7.943	8.972	11.80	437.44	43.35	388.9539	10773.66	0.18	0.16	1.099	19.27
0.80	6.310	7.126	12.30	475.29	37.86	269.7778	11043.44	0.18	0.17	0.999	19.52
0.70	5.012	5.661	13.40	564.10	88.81	502.7412	11546.18	0.19	0.18	0.899	19.77
0.60	3.981	4.496	14.60	669.66	105.56	474.6340	12020.81	0.20	0.20	0.799	20.02
0.50	3.162	3.572	15.10	716.31	46.65	166.6269	12187.43	0.20	0.21	0.699	20.27
0.40	2.512	2.837	15.90	794.23	77.91	221.0398	12408.47	0.20	0.22	0.599	20.52
0.30	1.995	2.254	29.70	2771.17	1976.94	4455.1562	16863.63	0.27	0.40	0.499	20.77
0.20	1.585	1.790	39.80	4976.40	2205.24	3947.5239	20811.15	0.34	0.54	0.399	21.02
0.10	1.259	1.422	42.30	5621.21	644.81	916.8599	21728.01	0.35	0.58	0.299	21.27
-0.00	1.000	1.129	47.50	7088.22	1467.00	1656.9170	23384.93	0.38	0.65	0.199	21.52
-0.10	0.794	0.897	53.50	8992.02	1903.80	1708.0166	25092.94	0.41	0.73	0.099	21.77
-0.20	0.631	0.713	73.60	17017.87	8025.85	5719.5312	30812.47	0.50	1.00	-0.001	22.02
-0.30	0.501	0.566	96.30	29134.15	12116.28	6858.6562	37671.13	0.61	1.31	-0.101	22.27
-0.40	0.398	0.450	112.83	39994.38	10860.23	4883.2500	42554.38	0.69	1.54	-0.201	22.52
-0.50	0.316	0.357	126.79	50503.30	10508.92	3753.4287	46307.80	0.75	1.73	-0.301	22.77
-0.60	0.251	0.284	137.50	59395.74	8892.44	2522.8484	48830.65	0.79	1.87	-0.401	23.02
-0.70	0.200	0.225	148.90	69652.87	10257.13	2311.5122	51142.16	0.83	2.03	-0.501	23.27
-0.80	0.158	0.179	156.30	76748.06	7095.19	1270.0884	52412.25	0.85	2.13	-0.601	23.52
-0.90	0.126	0.142	162.70	83162.00	6413.94	912.0002	53324.25	0.87	2.22	-0.701	23.77
-1.00	0.100	0.113	169.77	90546.50	7384.50	834.0491	54158.30	0.88	2.31	-0.801	24.02
-1.10	0.079	0.090	178.35	99930.00	9383.50	841.8523	55000.15	0.89	2.43	-0.901	24.27
-1.20	0.063	0.071	186.53	109306.81	9376.81	668.2310	55668.38	0.90	2.54	-1.001	24.52
-1.30	0.050	0.057	199.82	125437.56	16130.75	913.1177	56581.50	0.92	2.72	-1.101	24.77
-1.40	0.040	0.045	218.14	149492.87	24055.31	1081.6426	57663.14	0.94	2.97	-1.201	25.02
-1.50	0.032	0.036	232.54	169881.12	20388.25	728.2048	58391.34	0.95	3.17	-1.301	25.27
-1.60	0.025	0.028	243.08	185630.06	15748.94	446.8127	58838.15	0.96	3.31	-1.401	25.52
-1.70	0.020	0.023	249.63	195768.75	10138.69	228.4844	59066.64	0.96	3.40	-1.501	25.77
-1.80	0.016	0.018	262.40	216310.25	20541.50	367.7114	59434.35	0.97	3.57	-1.601	26.02
-1.90	0.013	0.014	275.10	237755.31	21445.06	304.9319	59739.28	0.97	3.75	-1.701	26.27
-2.00	0.010	0.011	289.30	262933.62	25178.31	284.3826	60023.66	0.98	3.94	-1.801	26.52
-2.00							61523.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4321

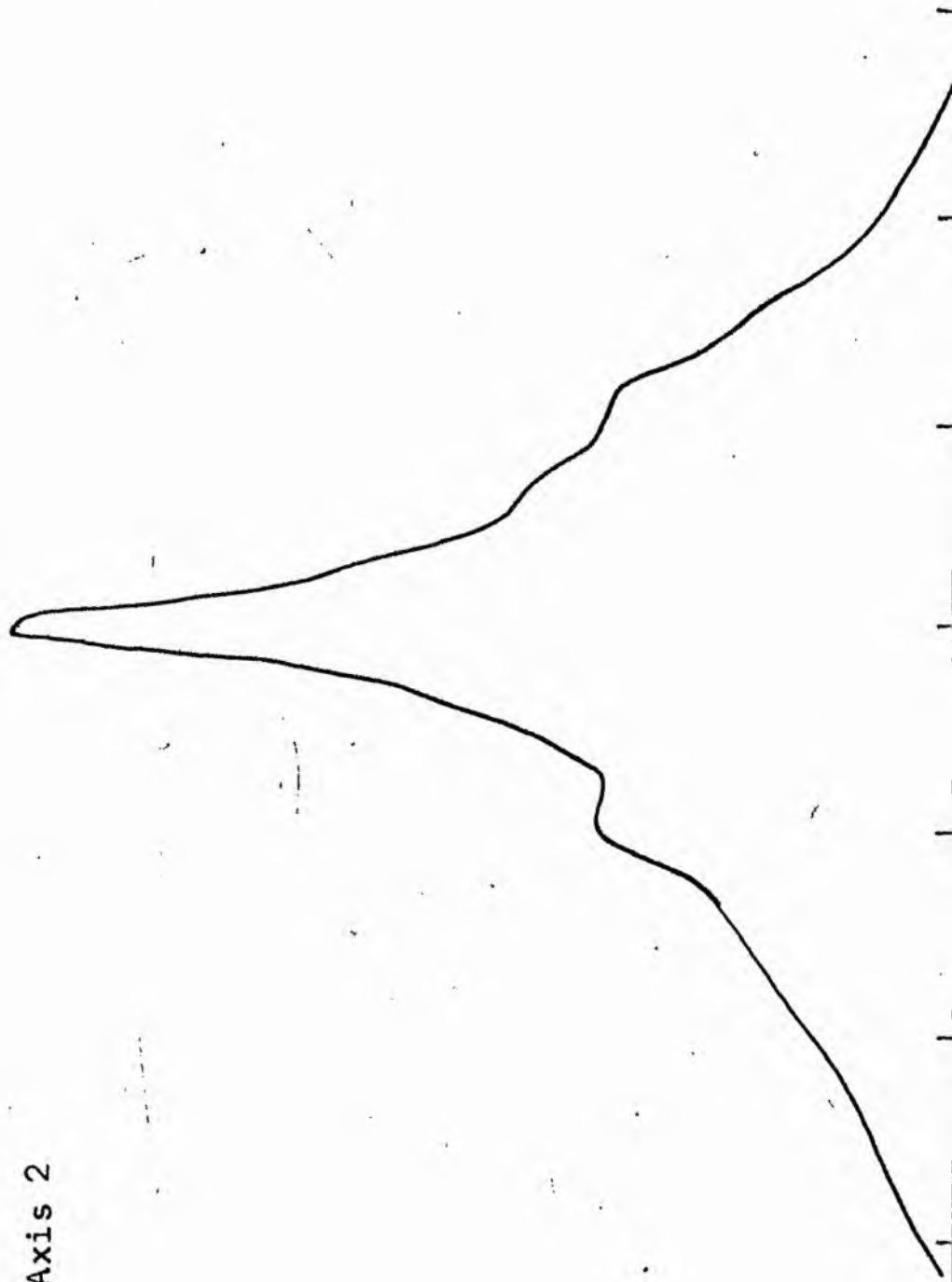
V-FILTER

Total luminosity	L_T	= 17.09
Total apparent magnitude	m_T	= 9.56
Apparent central surface brightness	μ_0	= 17.34
Major axis at threshold	$2a_m$	= 8.50
Minor axis at threshold	$2b_m$	= 10.06
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 7.03
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.94
Gradient of exponential component	$G(a)$	= -0.51
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.54
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.72
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.47
Axis ratio	b/a	= 0.61
Equivalent radius	r_1^*	= 0.42
Surface brightness	μ_1	= 20.69
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.62
Axis ratio	b/a	= 0.64
Equivalent radius	r_e^*	= 1.22
Surface brightness	μ_e	= 22.02
Mean surface brightness	μ_e'	= 11.99
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.26
Axis ratio	b/a	= 0.86
Equivalent radius	r_3^*	= 2.10
Surface brightness	μ_3	= 22.77
Concentration indices	$\{C_{21}$	= 2.89
	C_{32}	= 1.72

NGC 4340
B-Filter
Axis 1

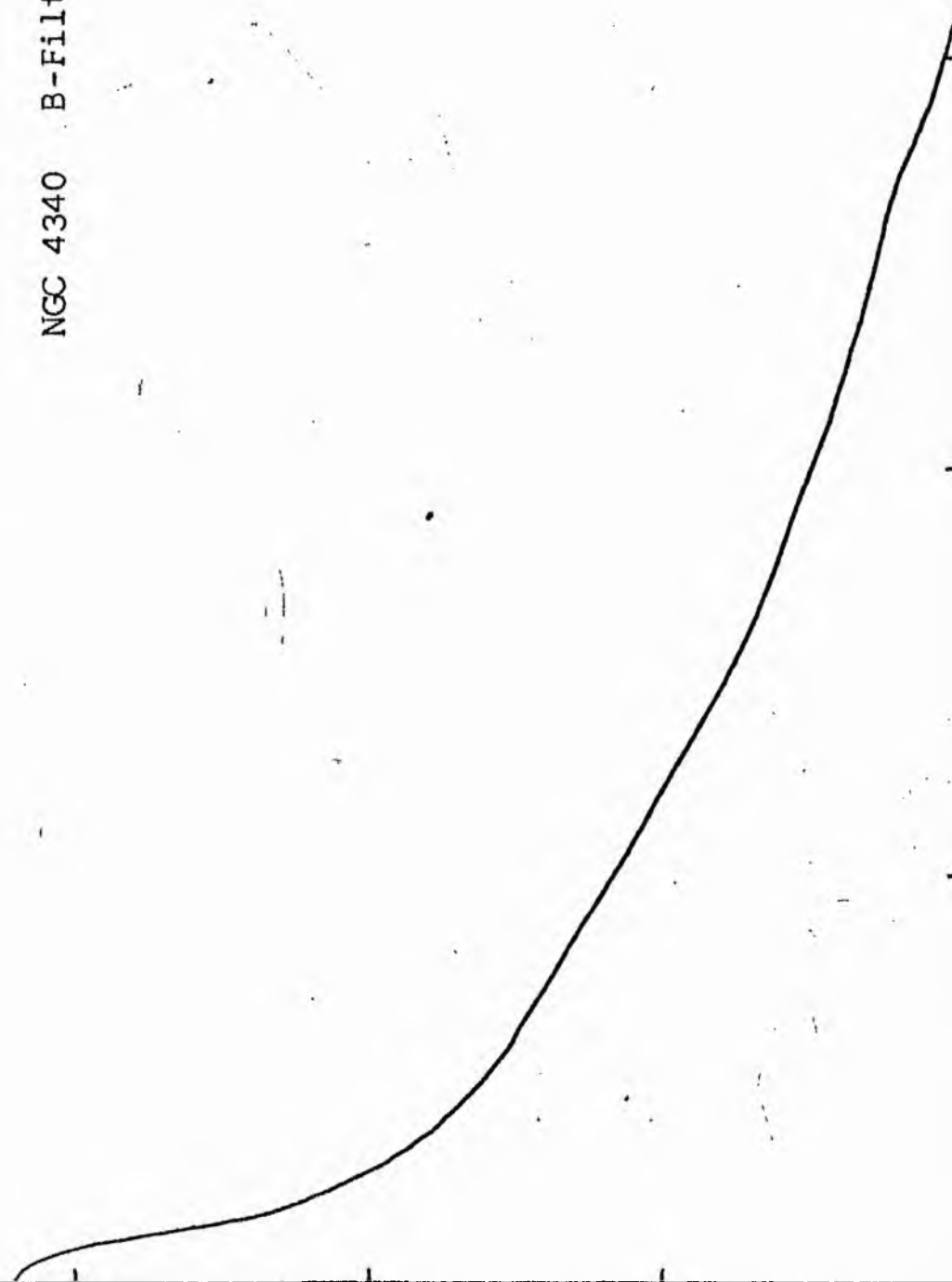


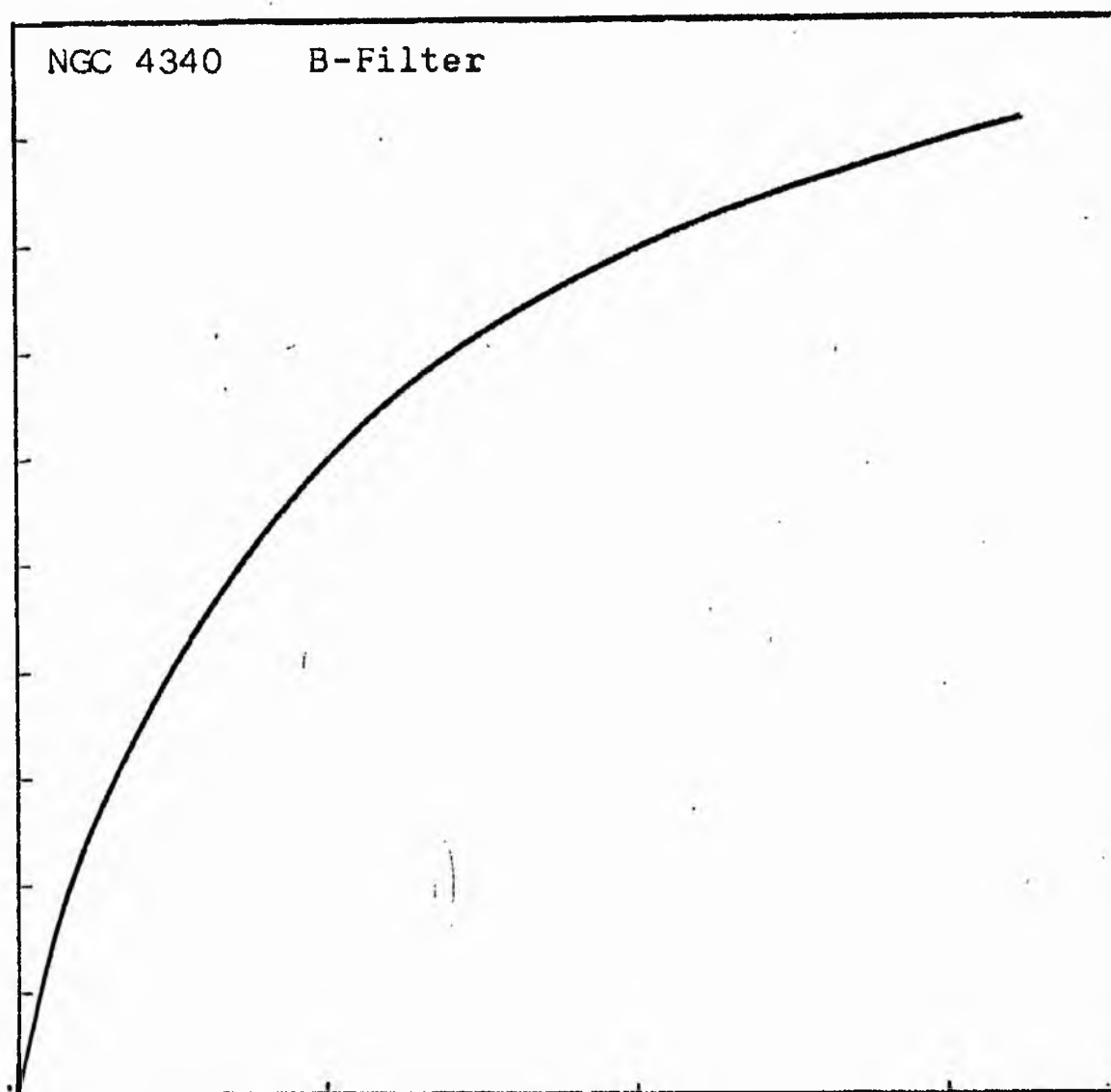
NGC 4340
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4340 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4340
B COLOR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.19	15.488		0.0	0.0			0.0	0.0	0.0	1.732	18.72
		14.039			26.42	370.9128					
1.10	12.589		2.90	26.42			370.91	0.05	0.08	1.642	18.95
		11.295			18.94	213.9627					
1.00	10.060		3.80	45.36			584.88	0.08	0.10	1.542	19.20
		8.972			12.25	163.7561					
0.90	7.943		4.50	63.62			748.63	0.10	0.12	1.442	19.45
		7.126			31.42	223.8829					
0.80	6.310		5.50	95.03			972.51	0.13	0.15	1.342	19.70
		5.661			14.33	81.0932					
0.70	5.012		5.90	109.36			1053.61	0.14	0.16	1.242	19.95
		4.456			58.06	261.0496					
0.60	3.981		7.30	167.42			1314.66	0.17	0.20	1.142	20.20
		3.572			18.85	67.3245					
0.50	3.162		7.70	186.27			1381.98	0.18	0.21	1.042	20.45
		2.837			24.98	70.8579					
0.40	2.512		8.20	211.24			1452.84	0.19	0.22	0.942	20.70
		2.254			48.91	110.2320					
0.30	1.995		9.10	260.16			1563.07	0.20	0.25	0.842	20.95
		1.790			207.44	371.3323					
0.20	1.565		12.20	467.59			1934.40	0.25	0.33	0.742	21.20
		1.422			96.51	137.2280					
0.10	1.259		13.40	564.10			2071.63	0.27	0.37	0.642	21.45
		1.129			133.36	150.6256					
-0.00	1.000		14.90	697.46			2222.26	0.29	0.41	0.542	21.70
		0.897			178.69	160.3174					
-0.10	0.794		16.70	876.16			2382.57	0.31	0.46	0.442	21.95
		0.713			210.71	150.1575					
-0.20	0.631		18.60	1086.86			2532.73	0.33	0.51	0.342	22.20
		0.566			475.42	269.1199					
-0.30	0.501		22.30	1562.28			2801.85	0.37	0.61	0.242	22.45
		0.450			1116.37	501.9707					
-0.40	0.398		29.20	2678.65			3303.82	0.43	0.80	0.142	22.70
		0.357			1038.99	371.0920					
-0.50	0.316		34.40	3717.63			3674.91	0.48	0.94	0.042	22.95
		0.284			1208.88	342.9700					
-0.60	0.251		39.60	4926.52			4017.88	0.53	1.08	-0.058	23.20
		0.225			1491.88	336.2068					
-0.70	0.200		45.20	6418.40			4354.09	0.57	1.24	-0.158	23.45
		0.179			1978.73	354.2080					
-0.80	0.158		51.70	8397.13			4708.30	0.62	1.41	-0.258	23.70
		0.142			1175.43	167.1349					
-0.90	0.126		55.20	9572.55			4875.43	0.64	1.51	-0.358	23.95
		0.113			1586.88	179.2321					
-1.00	0.100		59.60	11159.43			5054.66	0.66	1.63	-0.458	24.20
		0.090			2442.52	219.1343					
-1.10	0.079		65.80	13601.95			5273.79	0.69	1.80	-0.558	24.45
		0.071			3047.97	217.2121					
-1.20	0.063		72.80	16649.93			5491.00	0.72	1.99	-0.658	24.70
		0.057			3860.39	218.5267					
-1.30	0.050		80.80	20510.32			5709.53	0.75	2.21	-0.758	24.95
		0.045			4710.88	211.8244					
-1.40	0.040		89.60	25221.20			5921.35	0.77	2.45	-0.858	25.20
		0.036			6699.38	239.2815					
-1.50	0.032		100.00	31920.58			6160.63	0.81	2.76	-0.958	25.45
		0.028			9338.45	264.9414					
-1.60	0.025		114.60	41259.03			6425.57	0.84	3.13	-1.058	25.70
		0.023			9172.60	206.7132					
-1.70	0.020		126.70	50431.63			6632.29	0.87	3.46	-1.158	25.95
		0.018			9223.56	165.1106					
-1.80	0.016		137.80	59655.19			6797.39	0.89	3.77	-1.258	26.20
		0.014			8972.37	127.5804					
-1.90	0.013		147.80	68627.56			6924.97	0.91	4.04	-1.358	26.45
		0.011			7728.19	87.2881					
-2.00	0.010		155.90	76355.75			7012.26	0.92	4.26	-1.458	26.70
-∞							7642.00	1.11		∞	

PHOTOMETRIC PARAMETERS OF NGC 4340

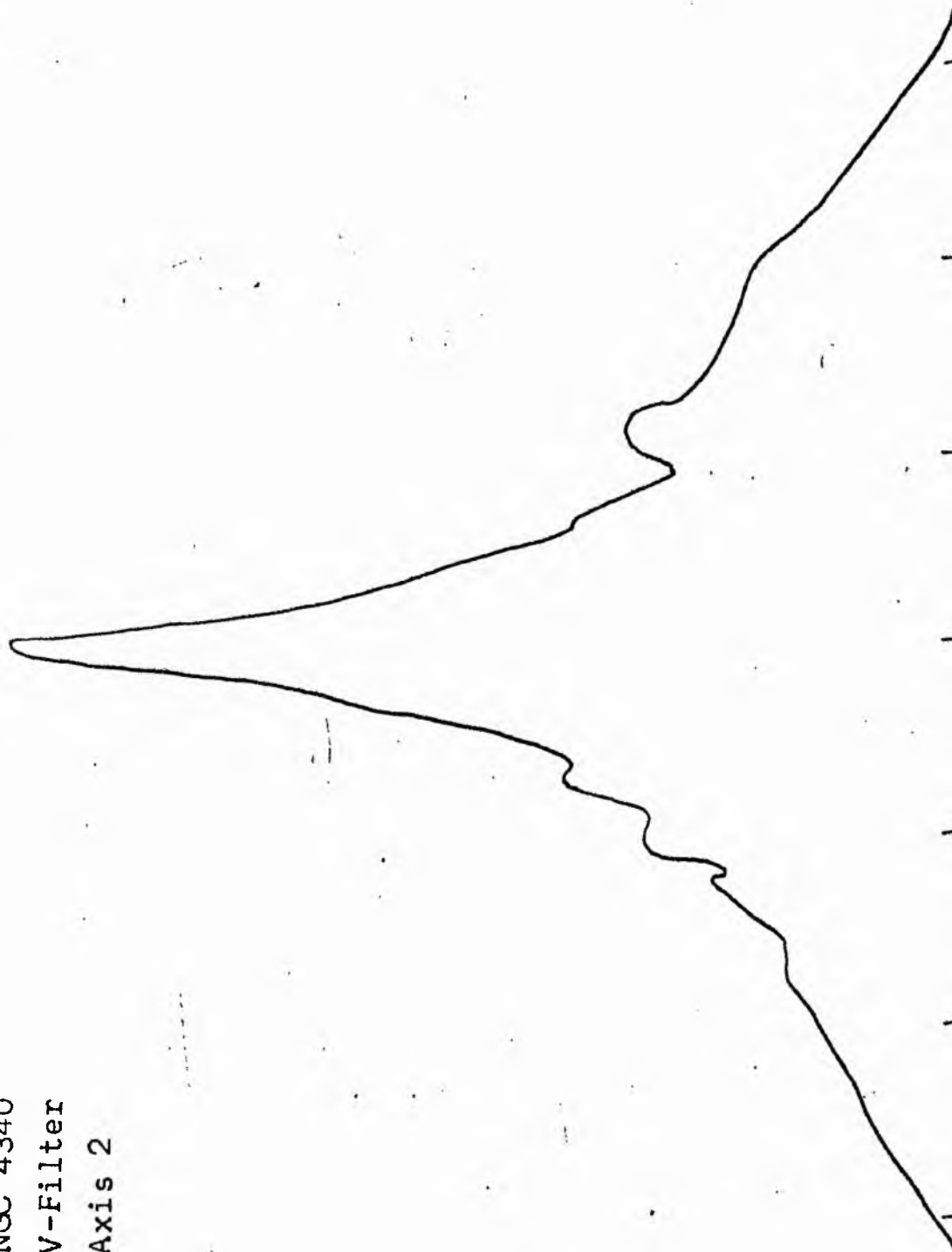
B-FILTER

Total luminosity	L_T	= 2.12
Total apparent magnitude	m_T	= 11.99
Apparent central surface brightness	μ_0	= 18.72
Major axis at threshold	$2a_m$	= 5.95
Minor axis at threshold	$2b_m$	= 4.70
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.05
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.76
Gradient of exponential component	$G(a)$	= -0.88
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.67
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.32
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 1.07
Equivalent radius	r_1^*	= 0.21
Surface brightness	μ_1	= 20.95
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.41
Axis ratio	b/a	= 0.96
Equivalent radius	r_e^*	= 0.61
Surface brightness	μ_e	= 23.05
Mean surface brightness	μ_e'	= 12.91
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.73
Axis ratio	b/a	= 0.71
Equivalent radius	r_3^*	= 1.31
Surface brightness	μ_3	= 24.95
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.90 \\ = 2.14 \end{matrix}$

NGC 4340
V-Filter
Axis 1

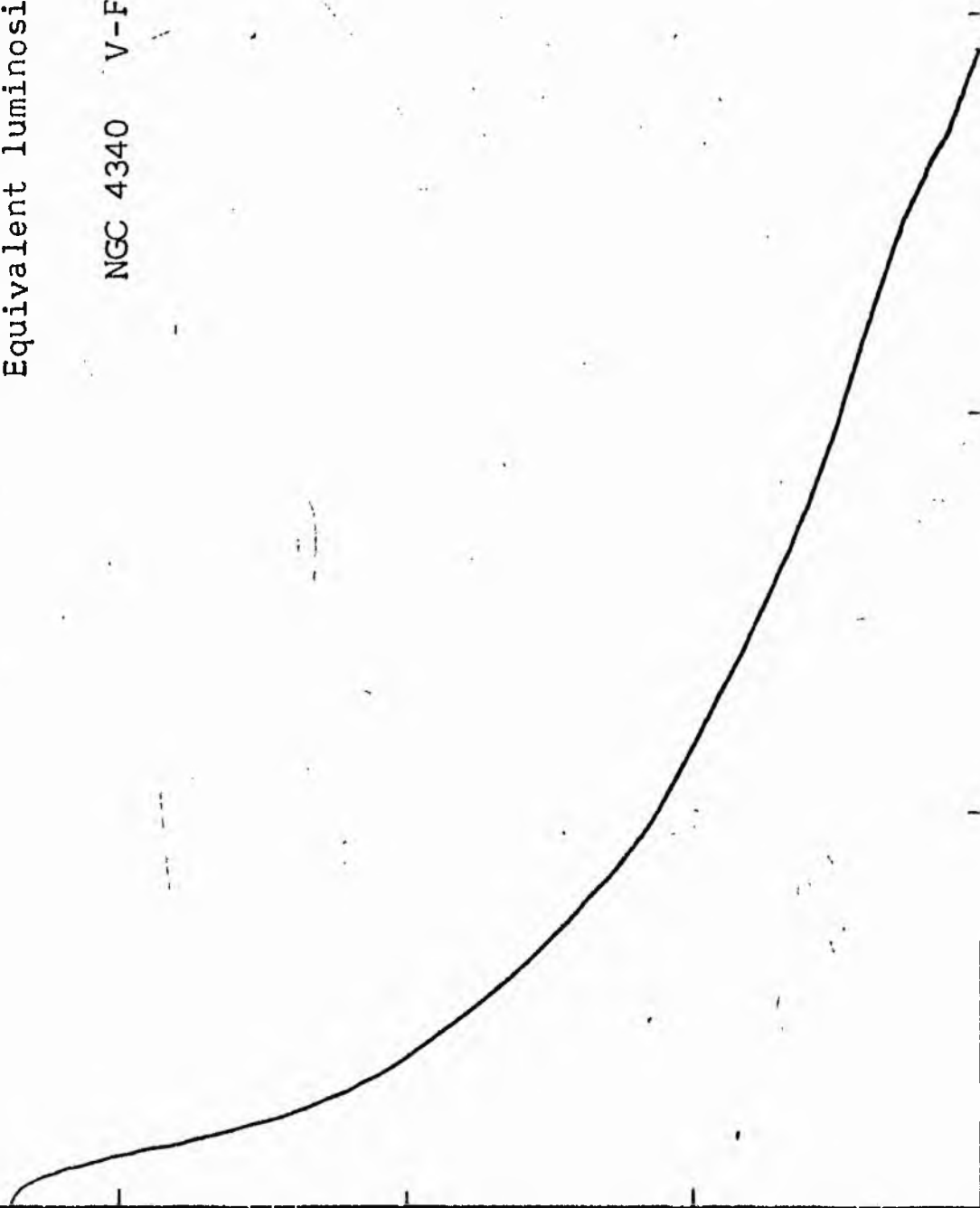


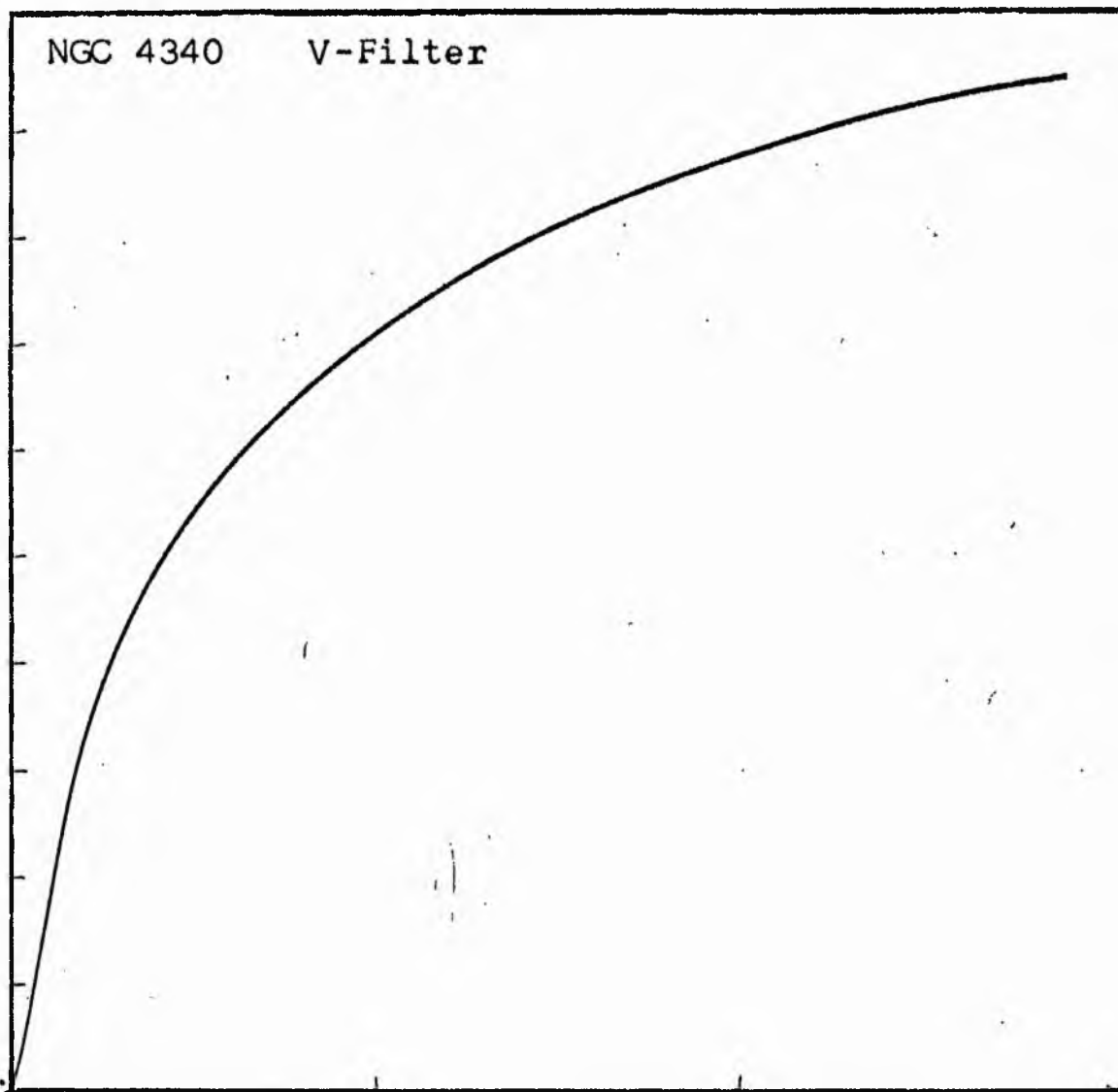
NGC 4340
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4340 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4340
v COLOUR

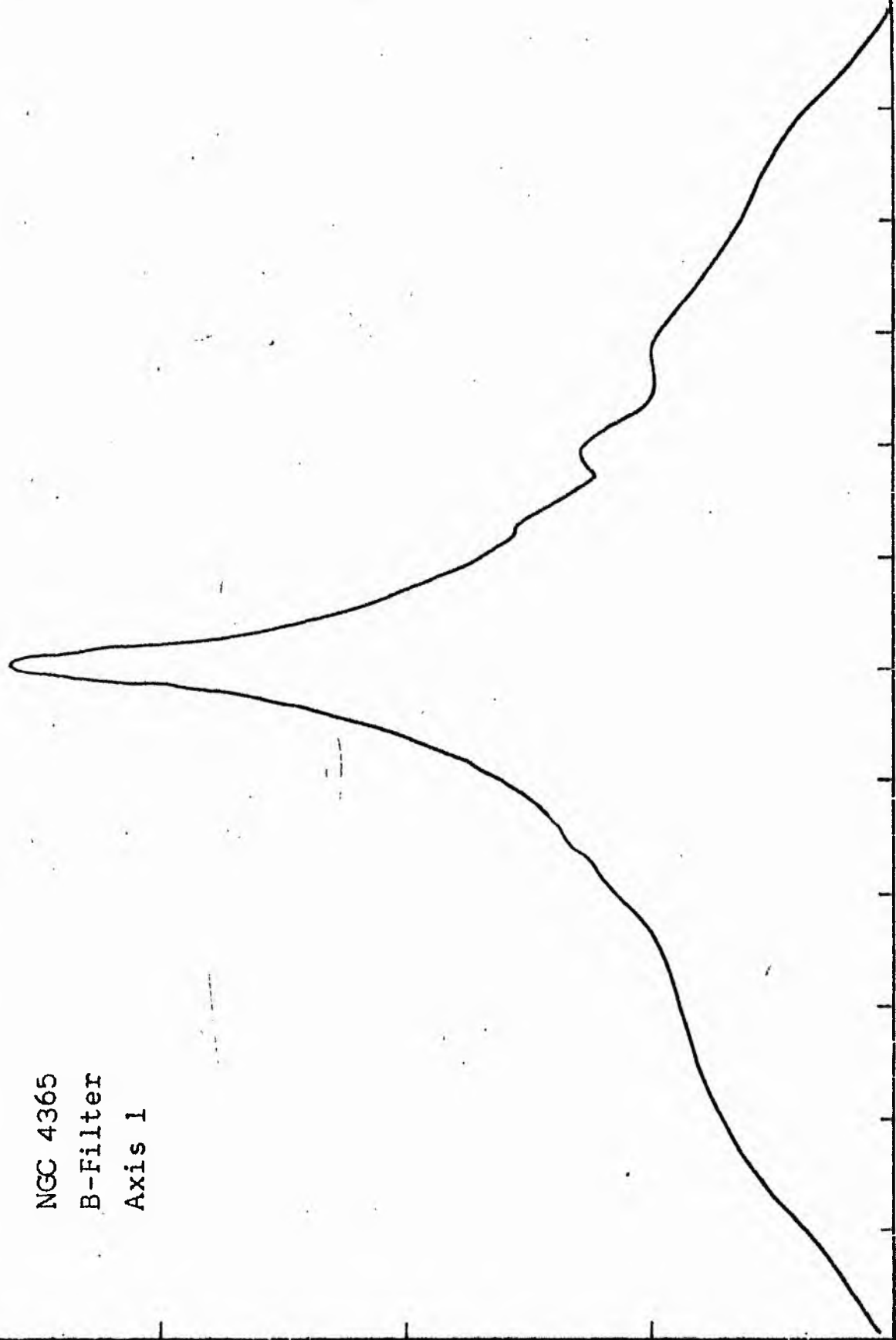
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.44	27.542		C.C	C.C			C.C	0.0	0.0	1.546	17.37
		26.331			9.08	239.0601					
1.40	25.119	22.536	1.70	9.08	36.29	817.7158	239.06	0.03	C.C8	1.506	17.47
1.30	19.953	17.901	3.80	45.36	21.11	377.9106	1056.78	0.12	0.19	1.406	17.72
1.20	15.849	14.219	4.60	66.48	28.56	406.0547	1434.69	0.16	0.23	1.306	17.97
1.10	12.589	11.295	5.50	95.03	10.65	120.2868	1840.74	0.20	0.27	1.206	18.22
1.00	10.000	8.972	5.80	105.68	52.68	472.6646	1961.03	0.22	0.29	1.106	18.47
0.90	7.943	7.126	7.10	158.37	9.05	64.4780	2433.69	0.27	0.35	1.006	18.72
0.80	6.310	5.661	7.30	167.42	43.83	248.6818	2498.17	0.27	0.36	0.906	18.97
0.70	5.012	4.446	8.20	211.24	43.23	194.3744	2746.25	0.30	0.40	0.806	19.22
0.60	3.781	3.577	9.00	254.47	53.44	190.8632	2940.63	0.32	0.44	0.706	19.47
0.50	3.162	2.837	9.90	307.91	72.23	204.9085	3131.49	0.34	0.49	0.606	19.72
0.40	2.512	2.254	11.00	380.13	118.63	267.3328	3336.40	0.37	0.54	0.506	19.97
0.30	1.935	1.790	12.60	498.76	108.23	193.7360	3603.73	0.40	0.62	0.406	20.22
0.20	1.585	1.422	13.90	606.99	177.28	252.0753	3797.47	0.42	0.68	0.306	20.47
0.10	1.259	1.129	15.80	784.27	211.11	238.4436	4049.54	0.44	0.78	0.206	20.72
-0.00	1.000	0.897	17.80	995.38	286.51	257.0498	4287.98	0.47	0.87	0.106	20.97
-0.10	0.794	0.713	20.20	1281.89	365.59	260.5320	4545.03	0.50	0.99	0.006	21.22
-0.20	0.631	0.566	22.90	1647.48	443.69	251.1573	4805.56	0.53	1.13	-0.094	21.47
-0.30	0.501	0.450	25.80	2091.17	514.59	231.3848	5056.72	0.55	1.27	-0.194	21.72
-0.40	0.398	0.357	28.80	2605.76	733.00	261.8020	5288.10	0.58	1.42	-0.294	21.97
-0.50	0.316	0.284	32.60	3338.76	846.63	240.1953	5549.90	0.61	1.60	-0.394	22.22
-0.60	0.251	0.225	36.50	4185.39	1044.23	235.3246	5790.09	0.63	1.79	-0.494	22.47
-0.70	0.200	0.179	40.80	5229.62	2190.70	392.1506	6025.42	0.66	2.01	-0.594	22.72
-0.80	0.158	0.142	48.60	7420.31	1774.53	252.3213	6417.57	0.70	2.39	-0.694	22.97
-0.90	0.126	0.113	54.10	9194.84	1564.59	221.8934	6669.89	0.73	2.66	-0.794	23.22
-1.00	0.100	0.090	59.60	11159.43	2442.52	219.1343	6891.78	0.76	2.93	-0.894	23.47
-1.10	0.079	0.071	65.80	13601.95	3369.71	240.1409	7110.91	0.78	3.23	-0.994	23.72
-1.20	0.063	0.057	73.50	16971.67	4410.80	249.6837	7351.05	0.81	3.61	-1.094	23.97
-1.30	0.050	0.045	82.50	21382.46	5964.75	268.2041	7600.73	0.83	4.05	-1.194	24.22
-1.40	0.040	0.036	93.30	27347.21	5852.62	249.0379	7868.94	0.86	4.59	-1.294	24.47
-1.50	0.032	0.028	102.80	33199.84	6067.67	172.1460	8077.97	0.89	5.05	-1.394	24.72
-1.60	0.025	0.023	111.80	39267.50	6273.52	141.3796	8250.12	0.90	5.49	-1.494	24.97
-1.70	0.020	0.018	120.40	45541.02	6253.02	111.9351	8391.50	0.92	5.92	-1.594	25.22
-1.80	0.016	0.014	128.40	51754.04	6227.42	88.5493	8503.43	0.93	6.31	-1.694	25.47
-1.90	0.013	0.011	135.90	58021.46	6541.38	68.2360	8591.98	0.94	6.68	-1.794	25.72
-2.00	0.010		142.80	64062.85			8660.21	0.95	7.02	-1.894	25.97
-∞							9120.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4340

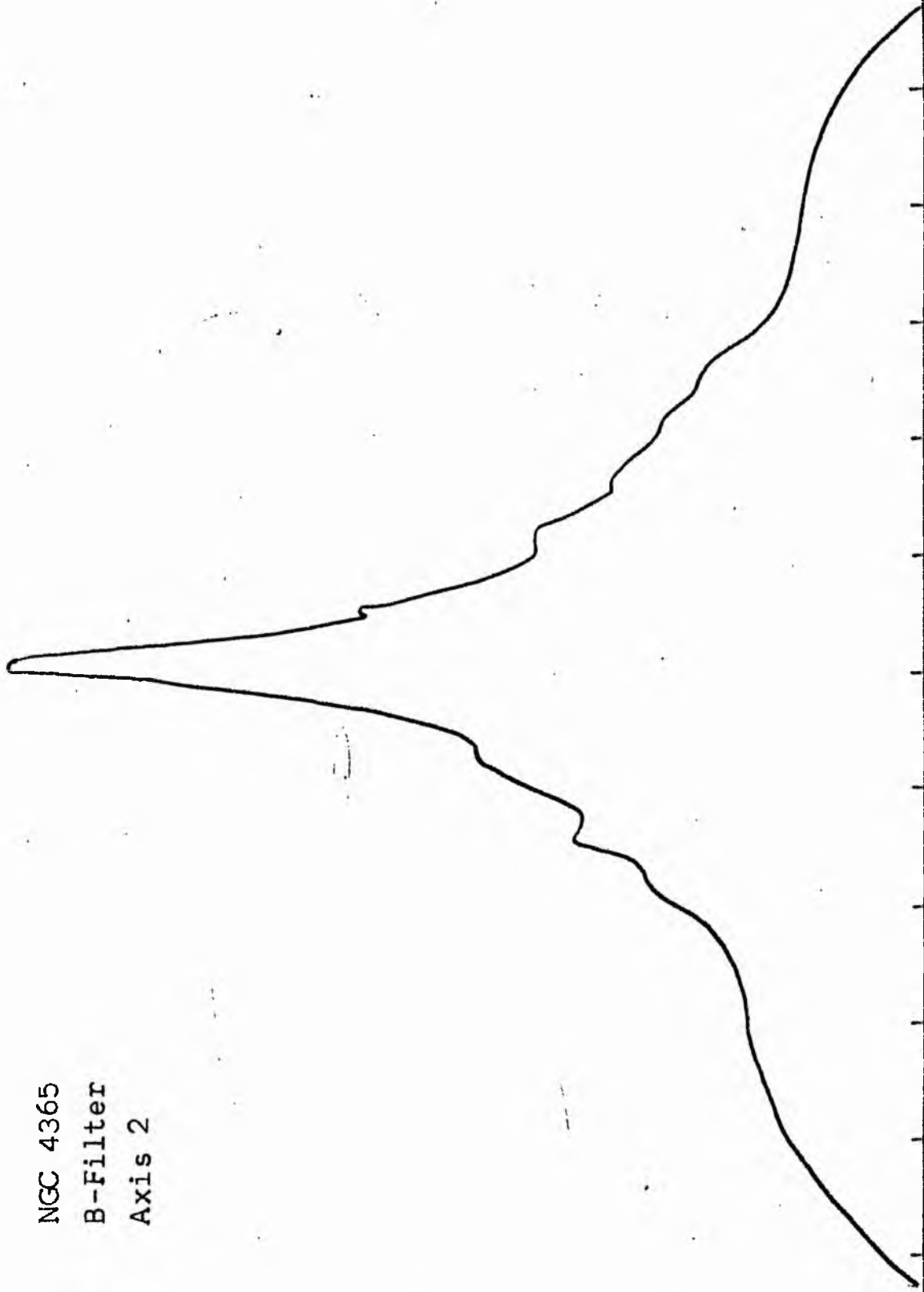
V-FILTER

Total luminosity	L_T	= 2.53
Total apparent magnitude	m_T	= 11.07
Apparent central surface brightness	μ_0	= 17.37
Major axis at threshold	$2a_m$	= 4.49
Minor axis at threshold	$2b_m$	= 5.08
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.93
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.91
Gradient of exponential component	$G(a)$	= -1.22
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.73
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.24
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.15
Axis ratio	b/a	= 1.11
Equivalent radius	r_1^*	= 0.12
Surface brightness	μ_1	= 18.62
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.36
Axis ratio	b/a	= 0.73
Equivalent radius	r_e^*	= 0.34
Surface brightness	μ_e	= 21.22
Mean surface brightness	μ_e'	= 12.38
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.05
Axis ratio	b/a	= 0.98
Equivalent radius	r_3^*	= 0.97
Surface brightness	μ_3	= 23.39
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.83 \\ = 2.85 \end{matrix}$

NGC 4365
B-Filter
Axis 1

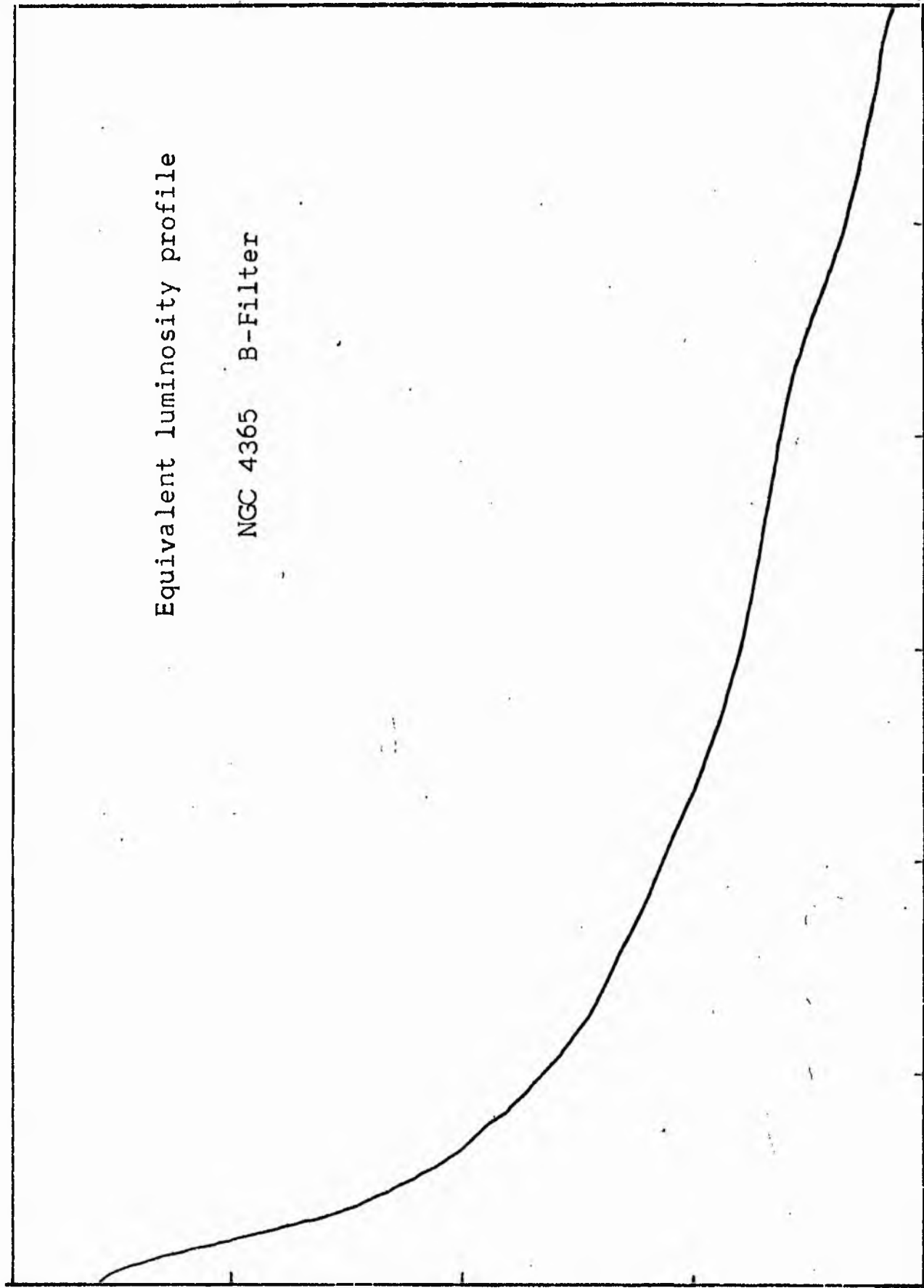


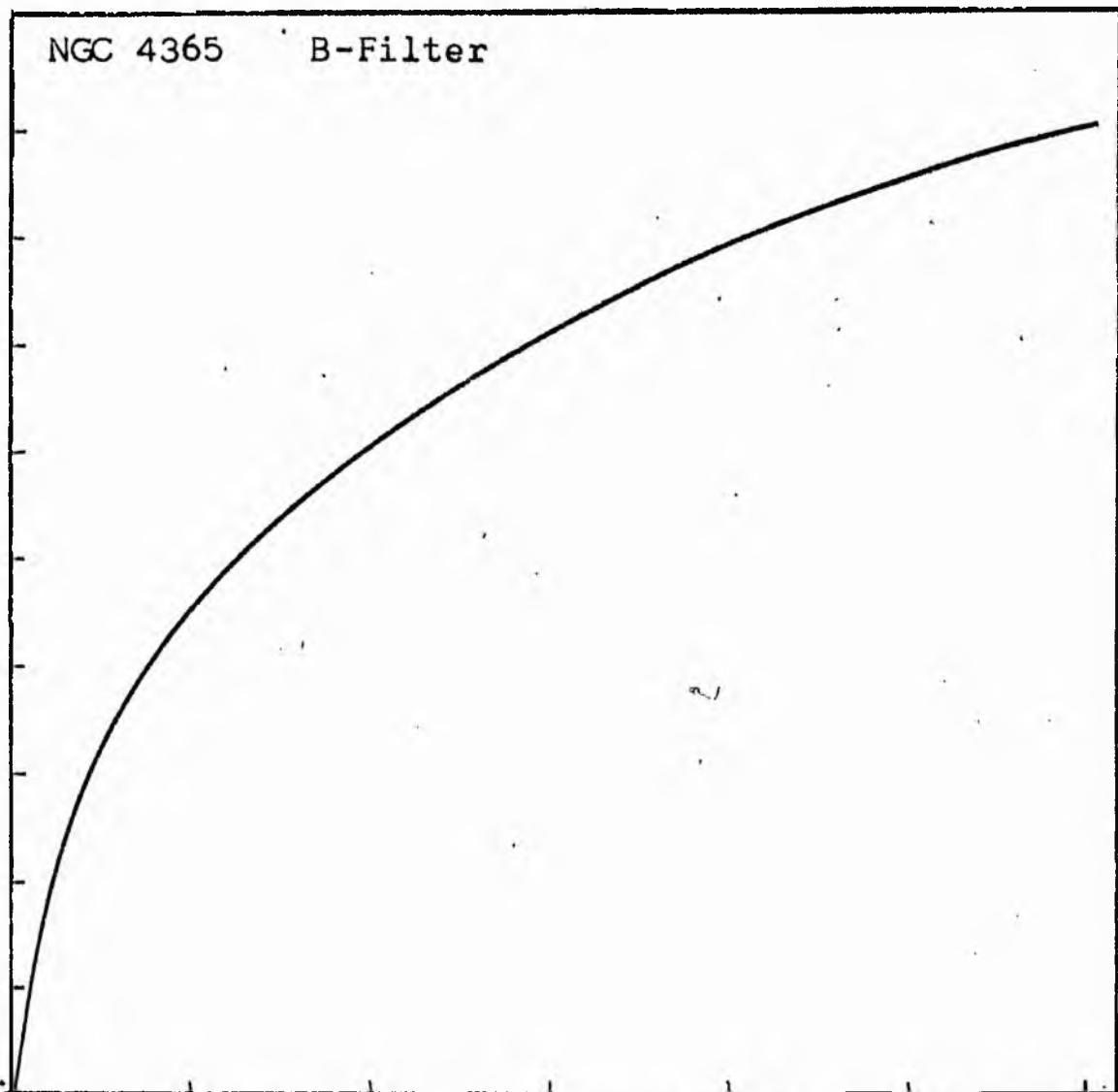
NGC 4365
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4365 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4365
B COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.67	46.773		0.0	0.0			0.0	0.0	0.0	2.244	17.61
1.60	39.811	43.292	2.50	19.63	19.63	850.0366	850.04	0.03	0.04	2.174	17.79
1.50	31.623	35.717	4.20	55.42	35.78	1278.0388	2128.08	0.06	0.07	2.074	18.04
1.40	25.119	28.371	5.45	93.31	37.90	1075.1223	3203.20	0.09	0.09	1.974	18.29
1.30	19.953	22.536	5.99	112.72	19.41	437.3618	3640.56	0.11	0.09	1.874	18.54
1.20	15.849	17.901	7.02	154.82	42.10	753.5876	4394.14	0.13	0.11	1.774	18.79
1.10	12.589	14.219	7.93	197.56	42.74	607.7178	5001.86	0.15	0.12	1.674	19.04
1.00	10.000	11.295	8.98	253.34	55.78	630.0161	5631.87	0.17	0.14	1.574	19.29
0.90	7.943	8.972	10.02	315.42	62.08	556.9353	6188.81	0.18	0.16	1.474	19.54
0.80	6.310	7.126	11.20	394.08	78.66	560.5920	6749.40	0.20	0.18	1.374	19.79
0.70	5.012	5.661	13.00	530.93	136.85	774.6533	7524.05	0.22	0.20	1.274	20.04
0.60	3.981	4.496	15.16	722.02	191.09	859.2217	8383.27	0.25	0.24	1.174	20.29
0.50	3.162	3.572	16.32	836.74	114.72	409.7397	8793.01	0.26	0.26	1.074	20.54
0.40	2.512	2.837	18.32	1054.39	217.65	617.4866	9410.49	0.28	0.29	0.974	20.79
0.30	1.995	2.254	20.80	1359.18	304.79	686.8606	10097.35	0.30	0.33	0.874	21.04
0.20	1.585	1.790	23.70	1764.60	405.42	725.7358	10823.09	0.32	0.37	0.774	21.29
0.10	1.259	1.422	26.90	2273.29	508.69	723.3020	11546.39	0.34	0.42	0.674	21.54
-0.00	1.000	1.129	30.50	2922.47	649.18	733.2205	12270.61	0.36	0.48	0.574	21.79
-0.10	0.794	0.897	34.50	3739.28	816.81	732.8125	13012.42	0.39	0.54	0.474	22.04
-0.20	0.631	0.713	38.69	4702.70	963.41	686.5669	13698.98	0.41	0.61	0.374	22.29
-0.30	0.501	0.566	44.20	6137.54	1434.84	812.2212	14511.20	0.43	0.70	0.274	22.54
-0.40	0.398	0.450	50.40	7980.14	1842.61	828.5186	15339.72	0.45	0.79	0.174	22.79
-0.50	0.316	0.357	57.70	10459.27	2479.12	885.4590	16225.18	0.48	0.91	0.074	23.04
-0.60	0.251	0.284	65.80	13601.95	3142.68	891.6016	17116.78	0.51	1.04	-0.026	23.29
-0.70	0.200	0.225	81.30	20764.95	7162.99	1614.2273	18731.00	0.56	1.28	-0.126	23.54
-0.80	0.158	0.179	92.40	26822.16	6057.21	1084.2834	19815.29	0.59	1.45	-0.226	23.79
-0.90	0.126	0.142	103.70	33783.71	6961.55	989.8655	20805.15	0.62	1.63	-0.326	24.04
-1.00	0.100	0.113	116.20	42419.16	8635.45	975.3391	21780.48	0.65	1.83	-0.426	24.29
-1.10	0.079	0.090	131.40	54242.61	11823.45	1060.7554	22841.24	0.68	2.07	-0.526	24.54
-1.20	0.063	0.071	148.70	69465.87	15223.27	1084.8738	23926.11	0.71	2.34	-0.626	24.79
-1.30	0.050	0.057	179.50	101222.87	31757.00	1797.6772	25723.79	0.76	2.83	-0.726	25.04
-1.40	0.040	0.045	209.60	138016.94	36794.06	1654.4382	27378.22	0.81	3.30	-0.826	25.29
-1.50	0.032	0.036	225.70	160034.25	22017.31	786.3899	28164.61	0.83	3.55	-0.926	25.54
-1.60	0.025	0.028	240.10	181106.56	21072.31	597.8420	28762.45	0.85	3.78	-1.026	25.79
-1.70	0.020	0.023	254.70	203801.69	22695.12	511.4548	29273.90	0.87	4.01	-1.126	26.04
-1.80	0.016	0.018	286.50	257869.00	54067.31	967.8542	30241.75	0.90	4.51	-1.226	26.29
-1.90	0.013	0.014	294.70	272841.06	14972.06	212.8911	30454.64	0.90	4.64	-1.326	26.54
-2.00	0.010	0.011	298.80	280485.50	7644.44	86.3420	30540.98	0.91	4.70	-1.426	26.79
-∞							33740.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4365

B-FILTER

Total luminosity	L_T	= 9.30
Total apparent magnitude	m_T	= 10.47
Apparent central surface brightness	μ_o	= 17.61
Major axis at threshold	$2a_m$	= 10.40
Minor axis at threshold	$2b_m$	= 9.37
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.43
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.75
Gradient of exponential component	$G(a)$	= -0.43
Equivalent gradient of exponential comp.,....	$G(r^*)$	= -0.38
Equivalent gradient of reduced exp. comp.,....	$G(\rho)$	= -0.28
Parameters at $k = \frac{1}{4}$:		

Semi-major axis	a_1	= 0.26
Axis ratio	b/a	= 0.69
Equivalent radius	r_1^*	= 0.25
Surface brightness	μ_1	= 20.29

Parameters at $k = \frac{1}{2}$ (effective) :

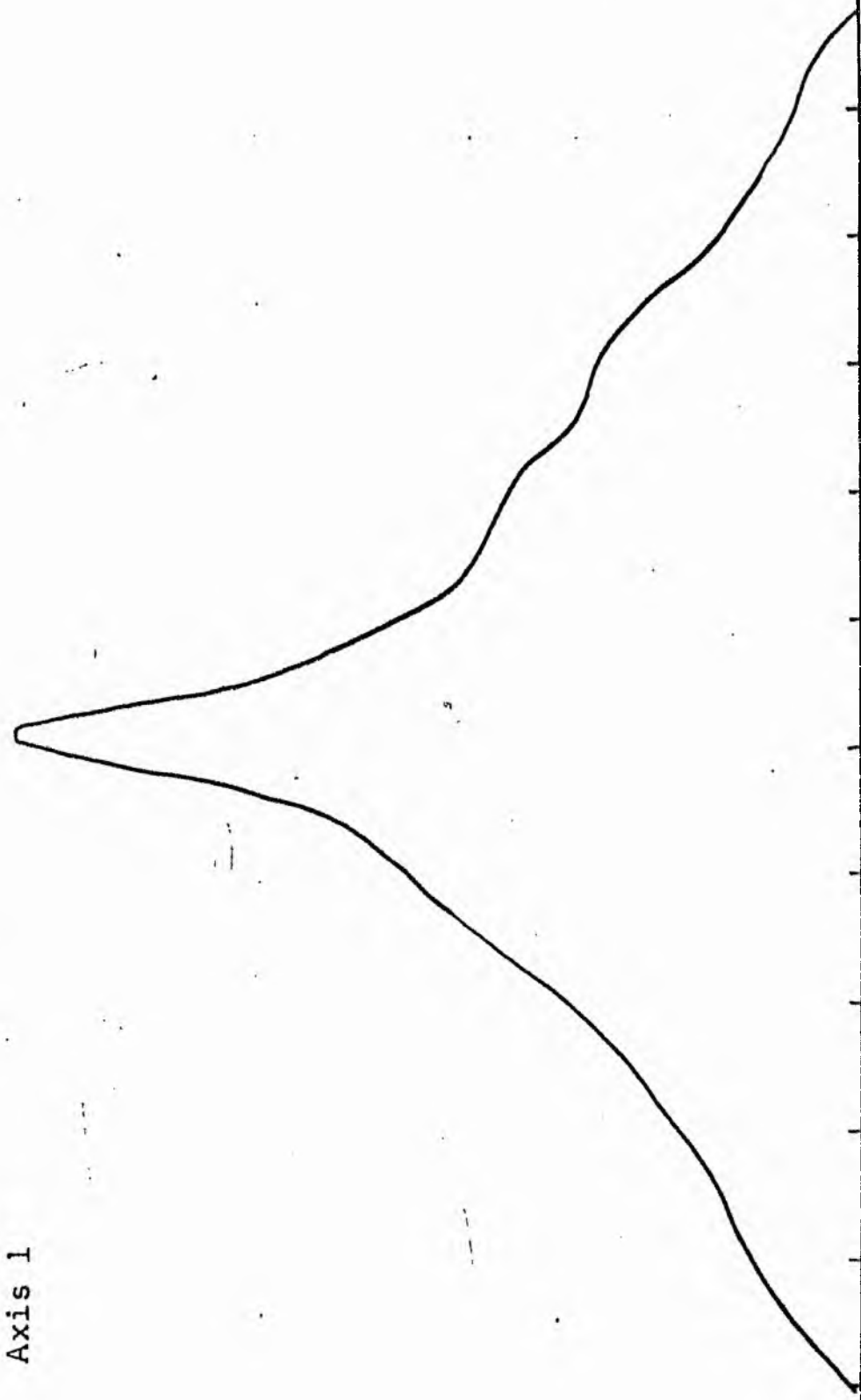
Semi-major axis	a_e	= 1.11
Axis ratio	b/a	= 0.95
Equivalent radius	r_e^*	= 1.06
Surface brightness	μ_e	= 23.17
Mean surface brightness	μ_e'	= 12.59

Parameters at $k = \frac{3}{4}$:

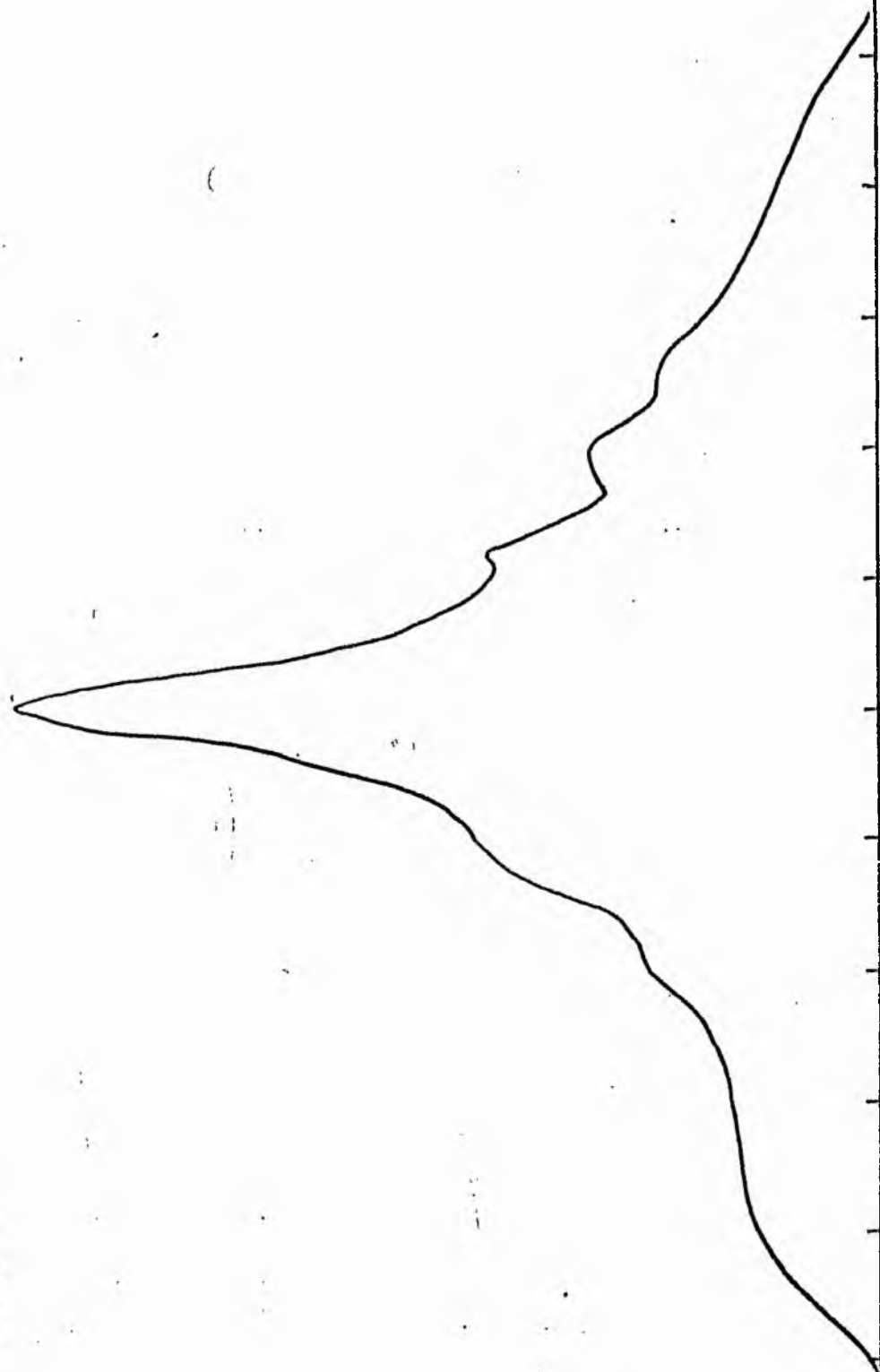
Semi-major axis	a_3	= 3.06
Axis ratio	b/a	= 0.83
Equivalent radius	r_3^*	= 2.87
Surface brightness	μ_3	= 24.99

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 4.15 \\ = 2.71 \end{cases}$
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NGC 4365
V-Filter
Axis 1

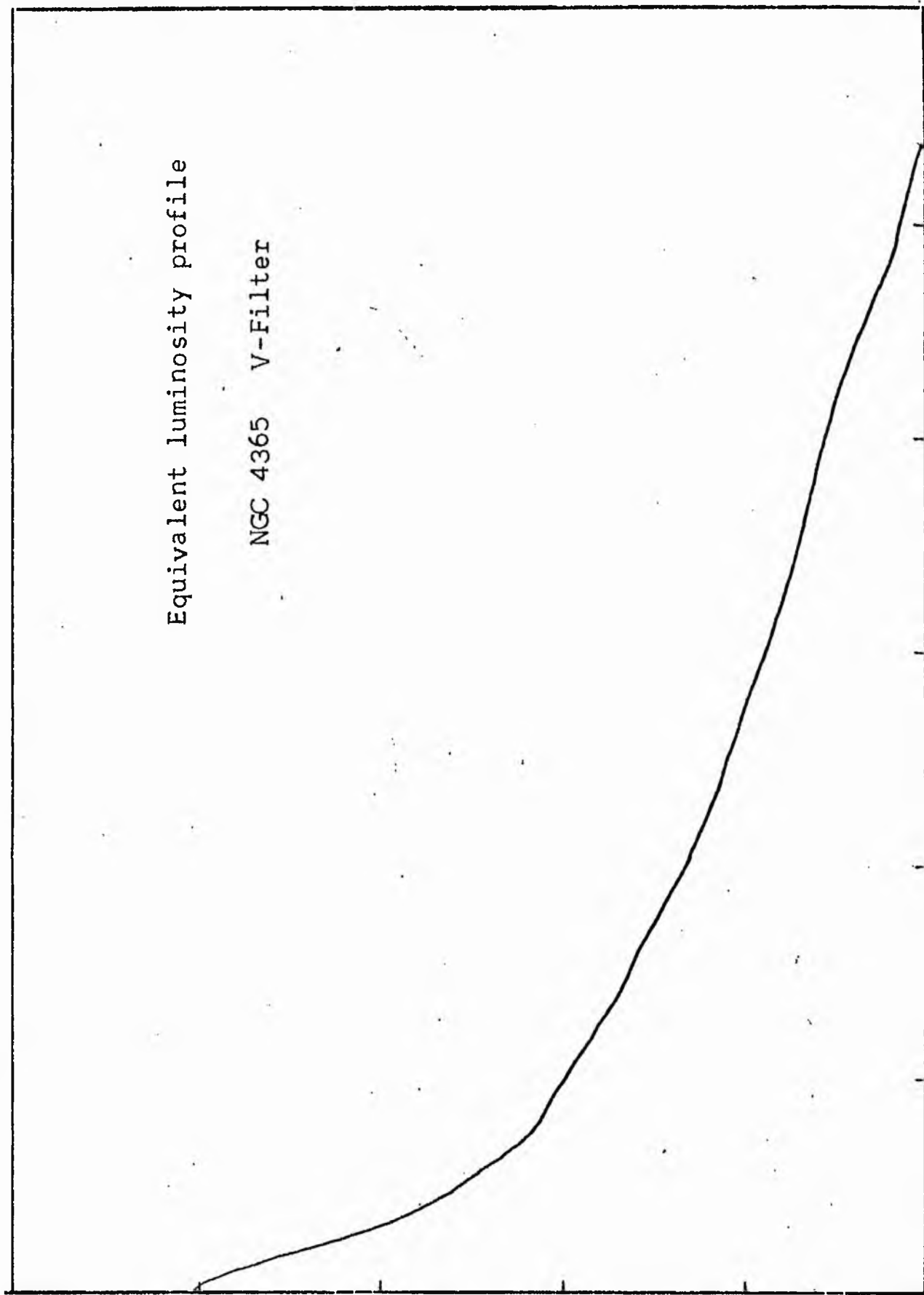


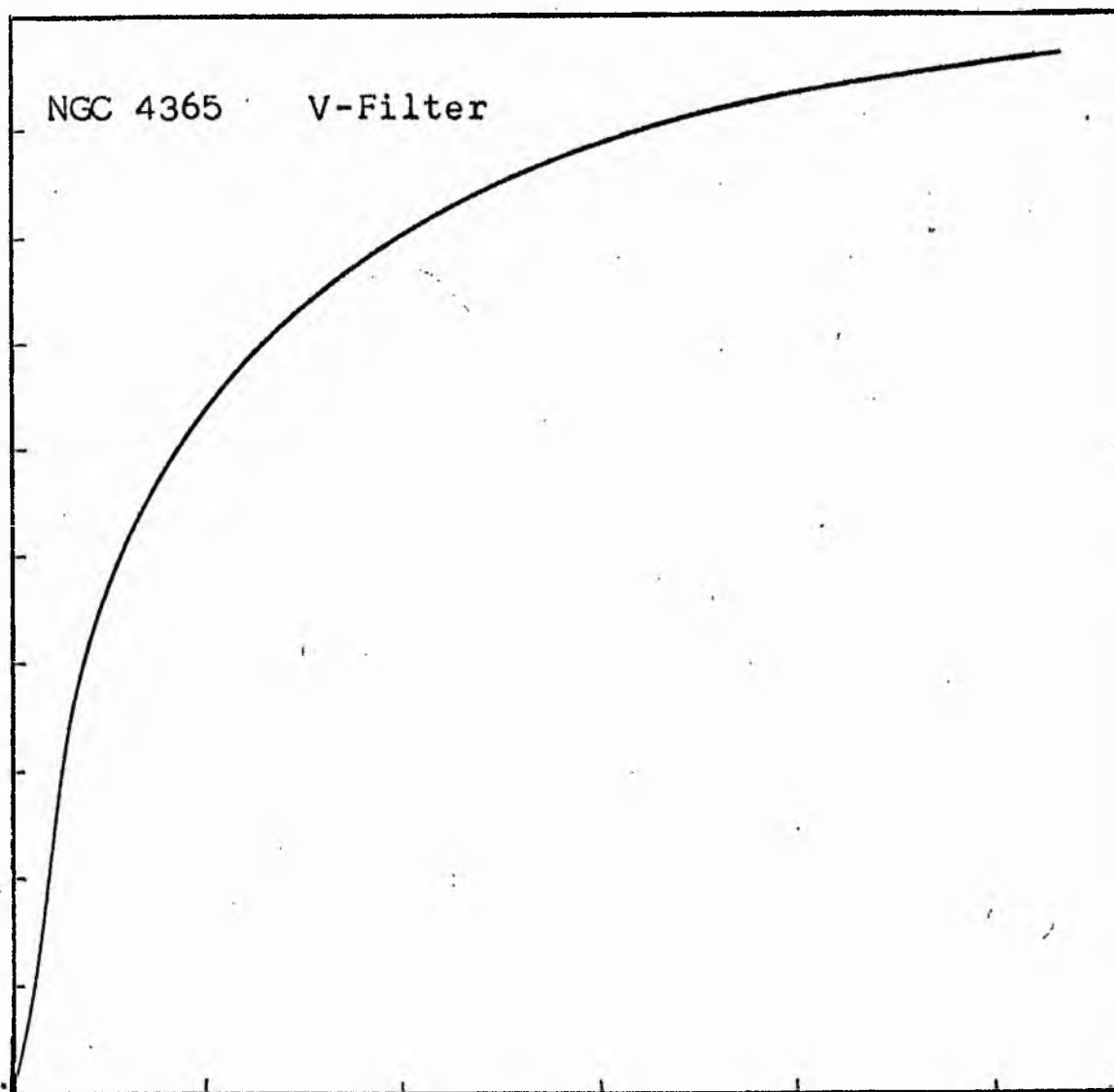
NGC 4365
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4365 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4365
V COLOUR

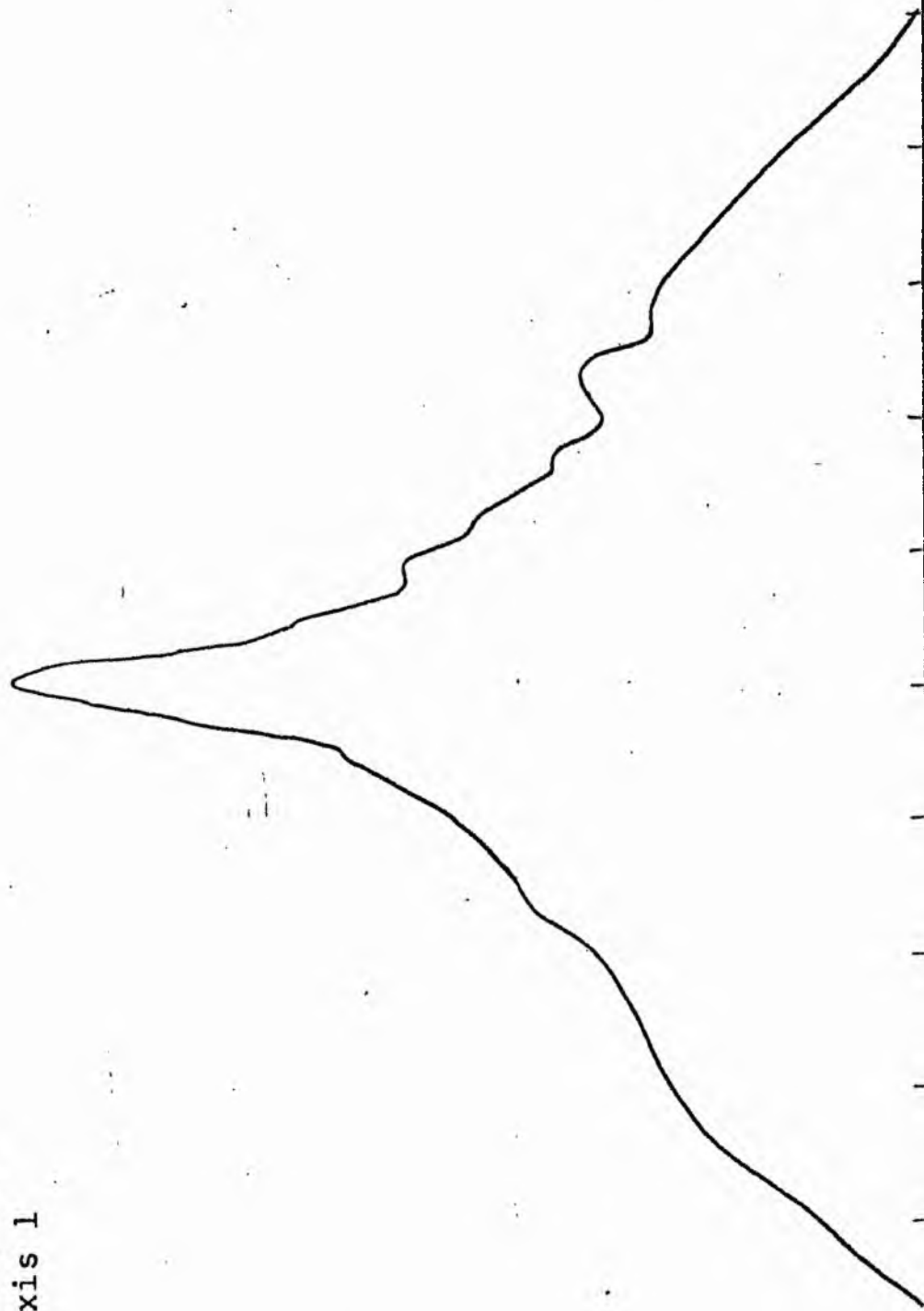
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.04	109.648		0.0	0.0	5.31	556.5396	0.0	0.0	0.0	1.582	16.96
2.00	100.000	104.824	1.30	5.31	19.32	1733.3889	556.54	0.01	0.05	1.542	17.06
1.90	79.433	89.716	2.80	24.63	16.08	1146.2812	2289.93	0.04	0.11	1.442	17.31
1.80	63.096	71.264	3.60	40.72	41.00	2320.7617	3436.21	0.06	0.14	1.342	17.56
1.70	50.119	56.607	5.10	81.71	72.23	3247.5791	5756.97	0.10	0.19	1.242	17.81
1.60	39.811	44.965	7.00	153.94	52.18	1863.7566	9004.55	0.16	0.26	1.142	18.06
1.50	31.623	35.717	8.10	206.12	140.24	3978.7231	10868.30	0.19	0.30	1.042	18.31
1.40	25.119	28.371	10.50	346.36	61.92	1395.4189	14847.02	0.26	0.39	0.942	18.56
1.30	19.952	22.536	11.40	408.28	82.59	1478.4600	16242.44	0.28	0.43	0.842	18.81
1.20	15.849	17.901	12.50	490.87	124.00	1763.1394	17720.90	0.31	0.47	0.742	19.06
1.10	12.589	14.219	13.99	614.87	139.89	1580.0449	19484.04	0.34	0.53	0.642	19.31
1.00	10.000	11.295	15.50	754.77	174.64	1566.7996	21064.08	0.37	0.58	0.542	19.56
0.90	7.943	8.972	17.20	929.41	192.80	1373.9587	22630.87	0.40	0.65	0.442	19.81
0.80	6.310	7.126	18.90	1122.21	250.07	1415.5669	24004.83	0.42	0.71	0.342	20.06
0.70	5.012	5.661	20.90	1372.28	289.62	1302.2788	25420.40	0.44	0.79	0.242	20.31
0.60	3.981	4.496	23.00	1661.90	380.92	1360.5051	26722.68	0.47	0.86	0.142	20.56
0.50	3.162	3.572	25.50	2042.82	473.25	1342.6299	28083.18	0.49	0.96	0.042	20.81
0.40	2.512	2.837	28.30	2516.07	640.89	1444.2749	29425.81	0.51	1.06	-0.058	21.06
0.30	1.995	2.254	31.70	3156.95	735.60	1316.7783	30870.08	0.54	1.19	-0.158	21.31
0.20	1.585	1.790	35.20	3892.56	548.90	780.4734	32186.86	0.56	1.32	-0.258	21.56
0.10	1.259	1.422	37.60	4441.45	1558.01	1759.7034	32967.33	0.58	1.41	-0.358	21.81
-0.00	1.000	1.129	43.70	5999.46	2300.49	2063.9041	34727.04	0.61	1.64	-0.458	22.06
-0.10	0.794	0.897	51.40	8299.96	3313.37	2361.2366	36790.94	0.64	1.93	-0.558	22.31
-0.20	0.631	0.713	60.80	11613.33	2828.05	1600.8716	39152.17	0.68	2.28	-0.658	22.56
-0.30	0.501	0.566	67.80	14441.39	2808.50	1262.8237	40753.04	0.71	2.55	-0.758	22.81
-0.40	0.398	0.450	74.10	17249.88	4654.09	1662.2771	42015.86	0.73	2.78	-0.858	23.06
-0.50	0.316	0.357	83.50	21903.97	2646.19	750.7415	43678.14	0.76	3.14	-0.958	23.31
-0.60	0.251	0.284	88.40	24550.16	4887.30	1101.3826	44428.87	0.78	3.32	-1.058	23.56
-0.70	0.200	0.225	96.80	29437.46	6800.05	1217.2544	45530.25	0.79	3.64	-1.158	23.81
-0.80	0.158	0.179	107.40	36237.51	9227.89	1312.1140	46747.51	0.82	4.03	-1.258	24.06
-0.90	0.126	0.142	120.30	45465.40	11535.93	1302.9333	48059.62	0.84	4.52	-1.358	24.31
-1.00	0.100	0.113	134.70	57001.33	6078.36	545.3262	49362.55	0.86	5.06	-1.458	24.56
-1.10	0.079	0.090	141.70	63079.69	11136.31	793.6191	49907.87	0.87	5.32	-1.558	24.81
-1.20	0.063	0.071	153.70	74216.00	13820.06	782.3142	50701.49	0.89	5.77	-1.658	25.06
-1.30	0.050	0.057	167.40	88036.06	19950.37	897.0632	51483.80	0.90	6.29	-1.758	25.31
-1.40	0.040	0.045	185.40	107986.44	24942.12	890.8528	52380.87	0.91	6.96	-1.858	25.56
-1.50	0.032	0.036	205.70	132928.56	10406.44	295.2397	53271.72	0.93	7.73	-1.958	25.81
-1.60	0.025	0.028	213.60	143335.00	14016.19	315.8665	53566.96	0.94	8.02	-2.058	26.06
-1.70	0.020	0.023	223.80	157351.19	18960.00	339.4006	53882.82	0.94	8.41	-2.158	26.31
-1.80	0.016	0.018	236.90	176311.19	31834.06	452.6545	54222.22	0.95	8.90	-2.258	26.56
-1.90	0.013	0.014	257.40	208145.25	9155.56	103.4095	54674.87	0.95	9.67	-2.358	26.81
-2.00	0.010	0.011	263.00	217300.81			54778.28	0.96	9.88	-2.458	27.06
							57278.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4365

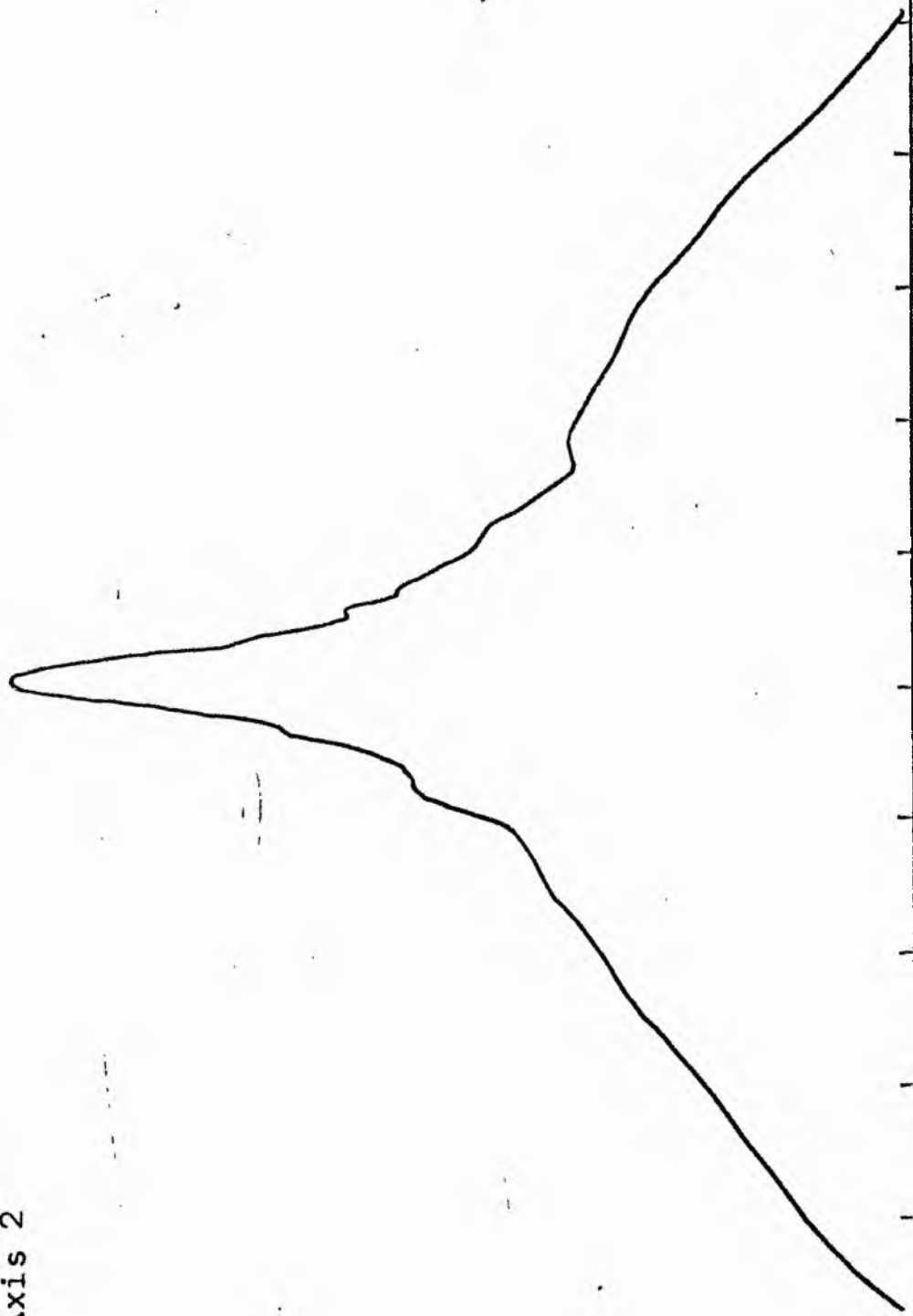
V-FILTER

Total luminosity	L_T	= 15.91
Total apparent magnitude	m_T	= 10.16
Apparent central surface brightness	μ_o	= 16.96
Major axis at threshold	$2a_m$	= 9.62
Minor axis at threshold	$2b_m$	= 8.83
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.37
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.88
Gradient of exponential component	$G(a)$	= -0.49
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.55
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.24
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 0.85
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 18.52
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.57
Axis ratio	b/a	= 0.74
Equivalent radius	r_e^*	= 0.44
Surface brightness	μ_e	= 20.93
Mean surface brightness	μ_e'	= 10.36
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.57
Axis ratio	b/a	= 0.77
Equivalent radius	r_3^*	= 1.32
Surface brightness	μ_3	= 23.24
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.62 \\ = 2.97 \end{cases}$

NGC 4374
B-Filter
Axis 1

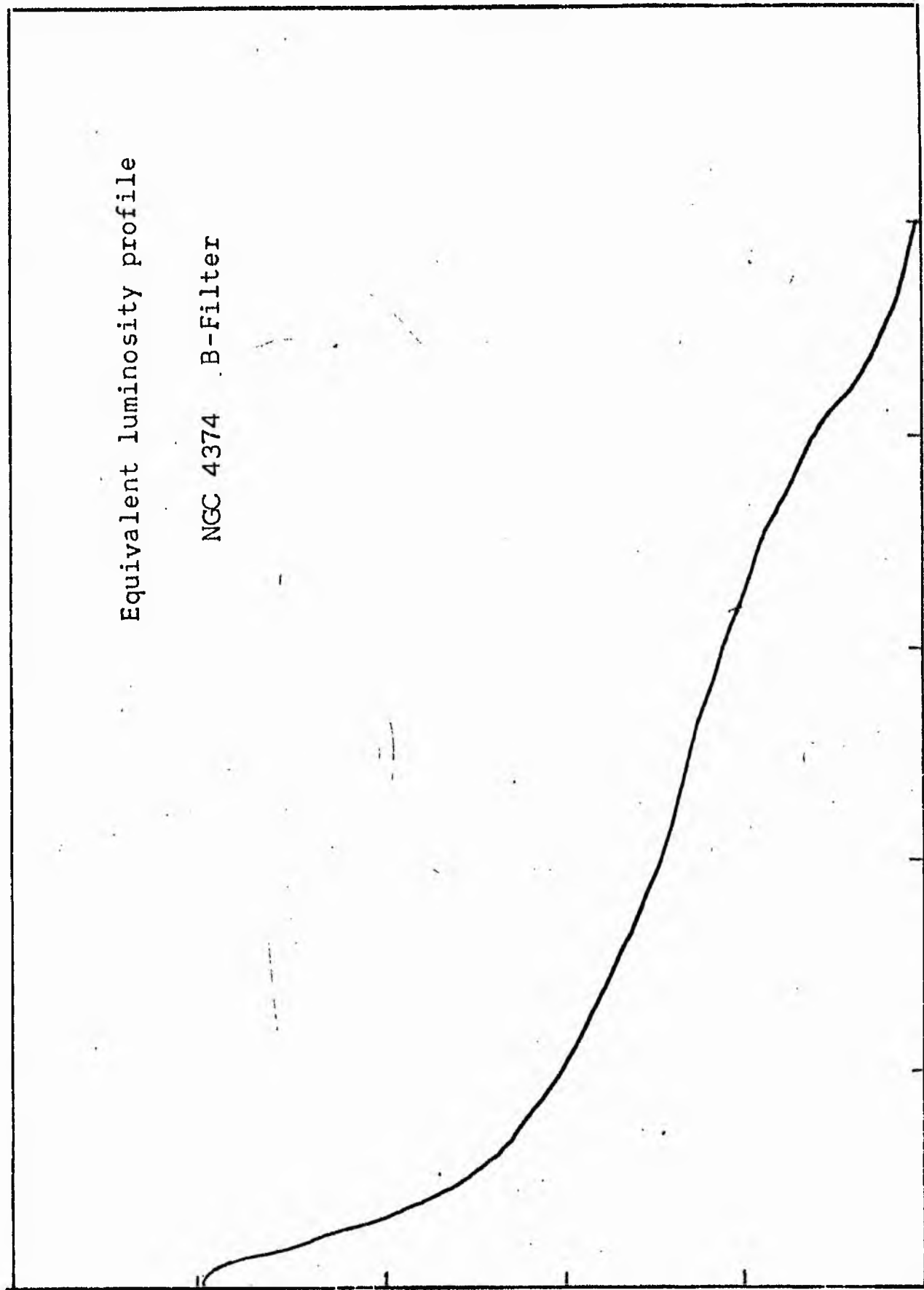


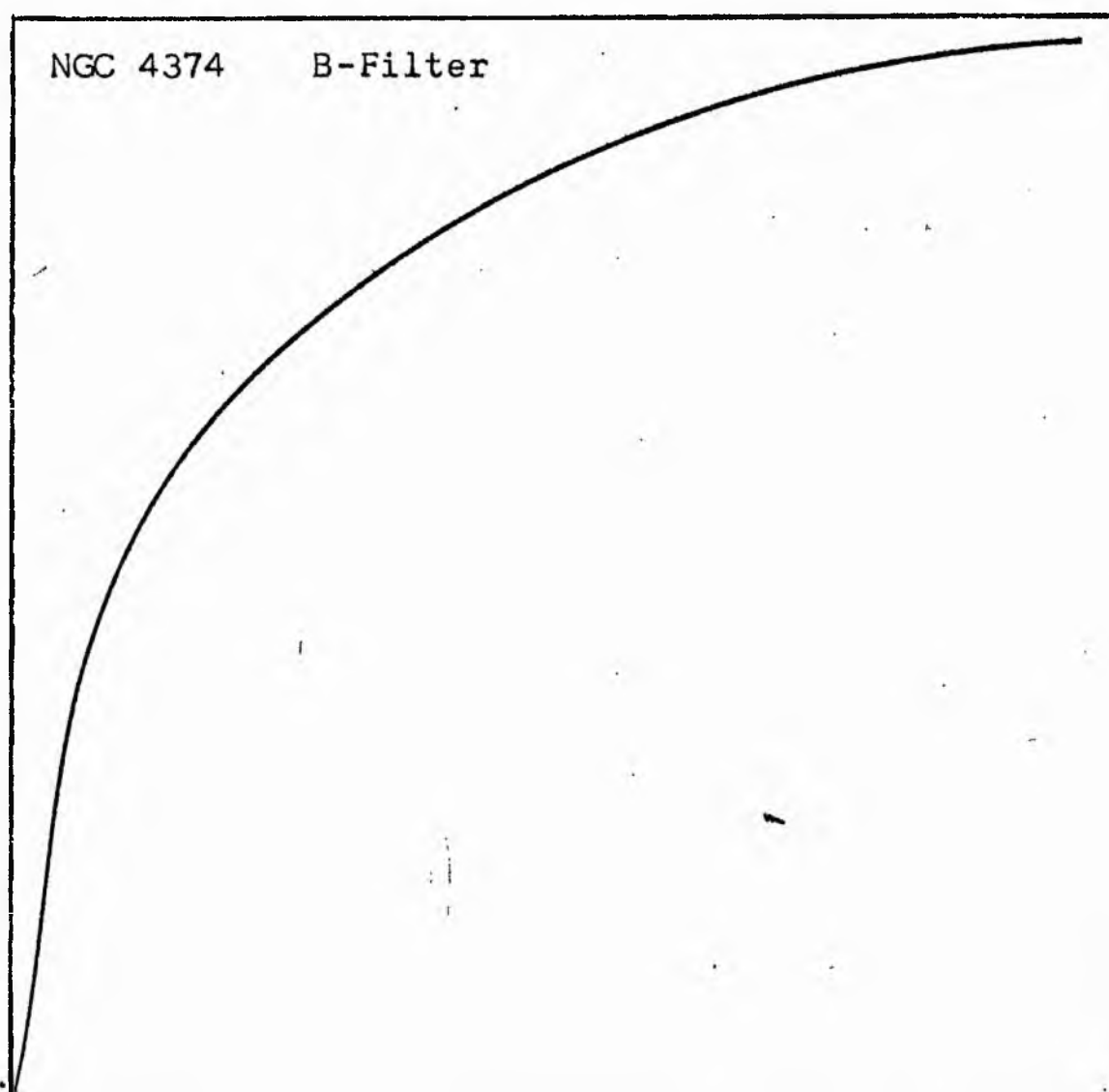
NGC 4374
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4374 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4374
B COLOUR

LOG I	I	I	R	AREA	ΔA	P	ΣP	KIRI	ρ	LOG J	μ
1.93	85.114		0.0	0.0			0.0	0.0	0.0	1.445	17.51
1.90	79.433	82.273	3.03	28.84	28.84	2372.9761	2372.98	0.04	0.12	1.415	17.59
1.80	63.096	71.264	6.63	138.09	109.25	7785.7500	10158.72	0.15	0.25	1.315	17.84
1.70	50.119	56.607	7.36	170.18	32.08	1816.1912	11974.91	0.18	0.28	1.215	18.09
1.60	39.811	44.965	8.35	219.04	48.86	2197.0085	14171.92	0.21	0.32	1.115	18.34
1.50	31.623	35.717	9.80	301.72	82.68	2953.0066	17124.92	0.25	0.38	1.015	18.59
1.40	25.119	28.371	10.80	366.44	64.72	1836.0586	18960.98	0.28	0.41	0.915	18.84
1.30	19.952	22.536	12.40	483.05	116.62	2628.0105	21588.99	0.32	0.48	0.815	19.09
1.20	15.849	17.901	13.50	572.56	89.50	1602.1780	23191.16	0.34	0.52	0.715	19.34
1.10	12.589	14.219	15.20	725.83	153.28	2179.4612	25370.62	0.38	0.58	0.615	19.59
1.00	10.000	11.295	16.90	847.27	171.44	1936.2830	27306.91	0.40	0.65	0.515	19.84
0.90	7.943	8.972	17.50	962.11	64.84	581.7498	27888.65	0.41	0.67	0.415	20.09
0.80	6.310	7.126	19.30	1170.21	208.10	1482.9788	29371.63	0.44	0.74	0.315	20.34
0.70	5.012	5.661	22.20	1548.30	378.09	2140.2551	31511.88	0.47	0.85	0.215	20.59
0.60	3.981	4.496	23.50	1734.94	186.64	839.2261	32351.11	0.48	0.90	0.115	20.84
0.50	3.162	3.572	25.70	2074.99	340.05	1214.5210	33565.62	0.50	0.99	0.015	21.09
0.40	2.512	2.837	28.20	2498.32	423.33	1201.0125	34766.64	0.52	1.08	-0.085	21.34
0.30	1.995	2.254	33.10	3441.96	943.64	2126.5452	36893.18	0.55	1.27	-0.185	21.59
0.20	1.585	1.790	39.10	4802.89	1360.94	2436.1636	39329.34	0.58	1.50	-0.285	21.84
0.10	1.259	1.422	45.90	6618.73	1815.84	2581.9404	41911.28	0.62	1.76	-0.385	22.09
-0.00	1.000	1.129	50.20	7916.94	1298.20	1466.2615	43377.54	0.64	1.93	-0.485	22.34
-0.10	0.794	0.897	56.70	10099.87	2182.93	1958.4355	45335.97	0.67	2.18	-0.585	22.59
-0.20	0.631	0.713	64.20	12948.50	2848.63	2030.0437	47366.01	0.70	2.46	-0.685	22.84
-0.30	0.501	0.566	72.50	16513.00	3564.49	2017.7454	49383.75	0.73	2.78	-0.785	23.09
-0.40	0.398	0.450	81.40	20816.06	4303.06	1934.8467	51318.60	0.76	3.13	-0.885	23.34
-0.50	0.316	0.357	90.70	25844.28	5028.22	1795.9045	53114.50	0.79	3.48	-0.985	23.59
-0.60	0.251	0.284	101.90	32621.06	6776.78	1922.6162	55037.11	0.82	3.91	-1.085	23.84
-0.70	0.200	0.225	114.60	41259.03	8637.97	1946.6191	56983.73	0.84	4.40	-1.185	24.09
-0.80	0.158	0.179	128.50	51874.77	10615.73	1900.2874	58884.02	0.87	4.93	-1.285	24.34
-0.90	0.126	0.142	144.60	65688.00	13813.23	1964.1042	60848.12	0.90	5.55	-1.385	24.59
-1.00	0.100	0.113	158.20	78625.37	12937.37	1461.2205	62309.34	0.92	6.07	-1.485	24.84
-1.10	0.079	0.090	169.80	90578.50	11953.12	1072.3875	63381.72	0.94	6.52	-1.585	25.09
-1.20	0.063	0.071	179.50	101222.87	10644.37	758.5615	64140.28	0.95	6.89	-1.685	25.34
-1.30	0.050	0.057	187.40	110328.81	9105.94	515.4612	64655.74	0.96	7.19	-1.785	25.59
-1.40	0.040	0.045	193.60	117749.87	7421.06	333.6860	64989.43	0.96	7.43	-1.885	25.84
-1.50	0.032	0.036	200.50	126242.81	8542.94	305.1262	65294.55	0.97	7.70	-1.985	26.09
-1.60	0.025	0.028	206.80	134354.06	8061.25	228.7049	65523.25	0.97	7.94	-2.085	26.34
-1.70	0.020	0.023	213.50	143200.87	8846.81	199.3703	65722.56	0.97	8.20	-2.185	26.59
-1.80	0.016	0.018	219.70	151638.62	8437.75	151.0432	65873.56	0.98	8.43	-2.285	26.84
-1.90	0.013	0.014	234.70	173051.75	21413.12	304.4773	66178.00	0.98	9.01	-2.385	27.09
-2.00	0.010	0.011	246.80	191355.12	18303.37	206.7315	66384.69	0.98	9.48	-2.485	27.34
-∞							67484.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4374

B-FILTER

Total luminosity	L_T	= 18.75
Total apparent magnitude	m_T	= 10.27
Apparent central surface brightness	μ_o	= 17.51
Major axis at threshold	$2a_m$	= 8.03
Minor axis at threshold	$2b_m$	= 7.70
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.85
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.95
Gradient of exponential component	$G(a)$	= -0.61
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.49
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.26

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 0.92
Equivalent radius	r_1^*	= 0.16
Surface brightness	μ_1	= 18.59

Parameters at $k = \frac{1}{2}$ (effective) :

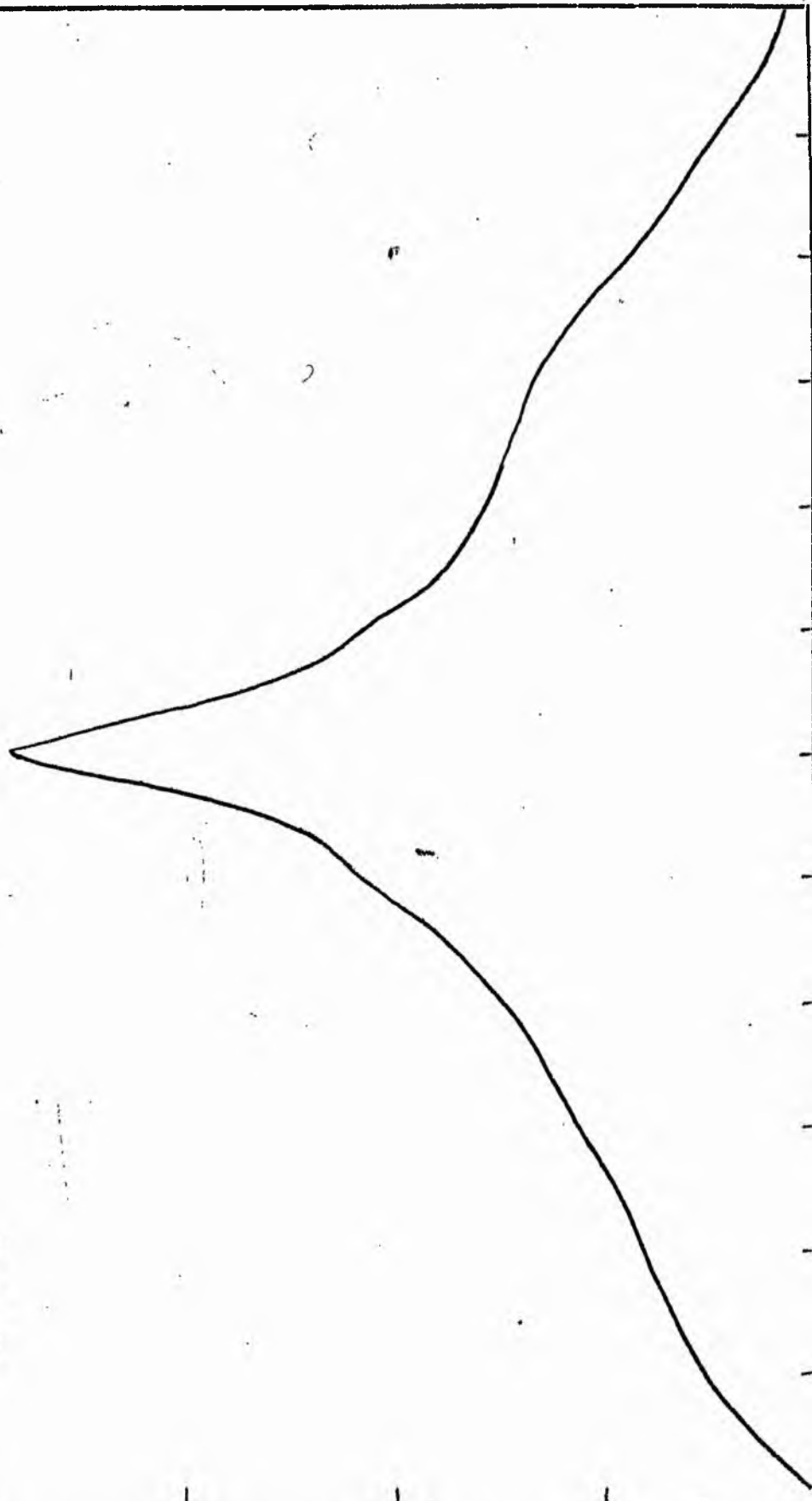
Semi-major axis	a_e	= 0.57
Axis ratio	b/a	= 0.79
Equivalent radius	r_e^*	= 0.43
Surface brightness	μ_e	= 21.09
Mean surface brightness	μ_e'	= 10.41

Parameters at $k = \frac{3}{4}$:

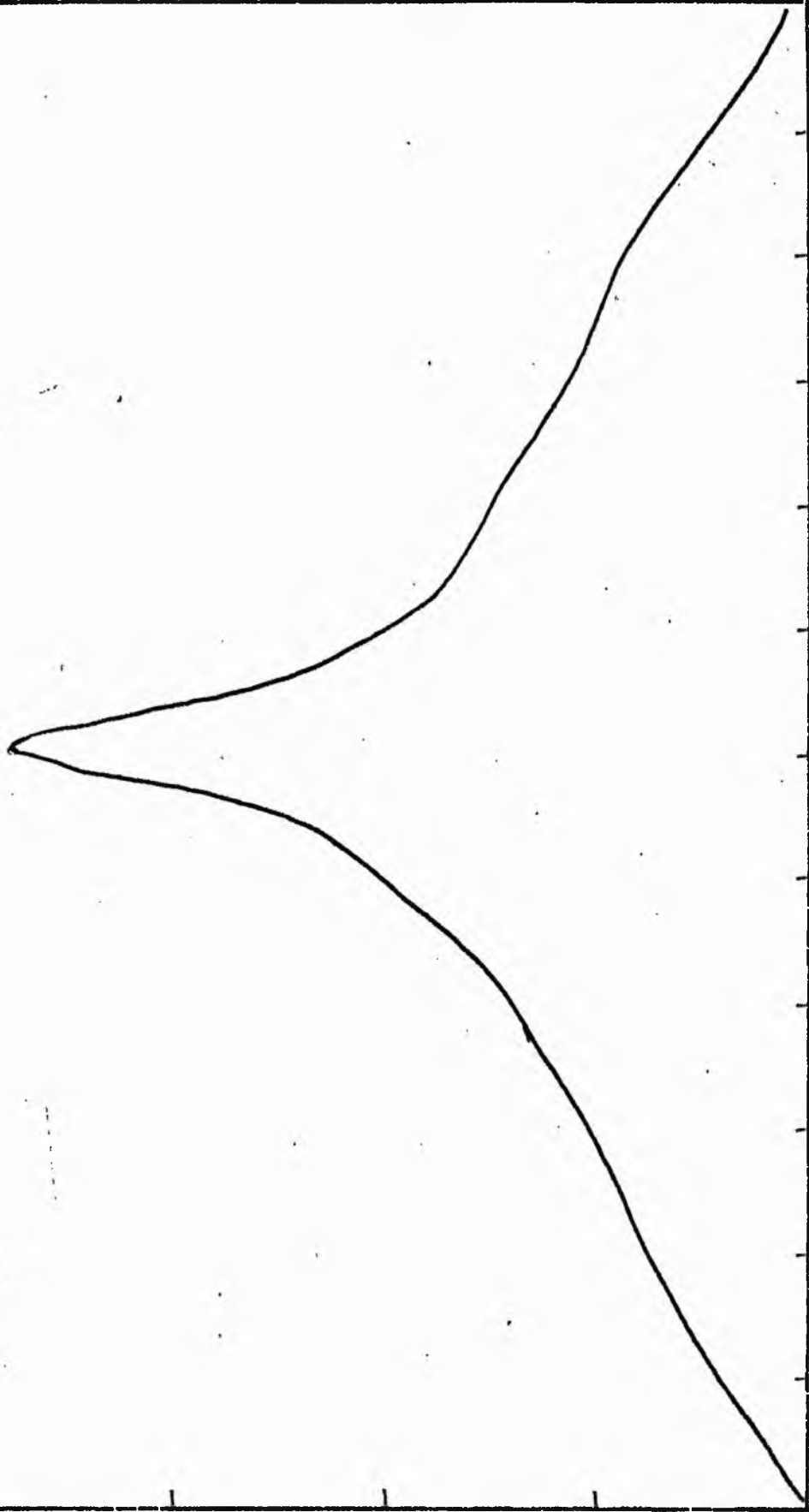
Semi-major axis	a_3	= 1.43
Axis ratio	b/a	= 0.95
Equivalent radius	r_3^*	= 1.30
Surface brightness	μ_3	= 23.26

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.69 \\ = 3.00 \end{matrix}$
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NGC 4374
V-Filter
Axis 1

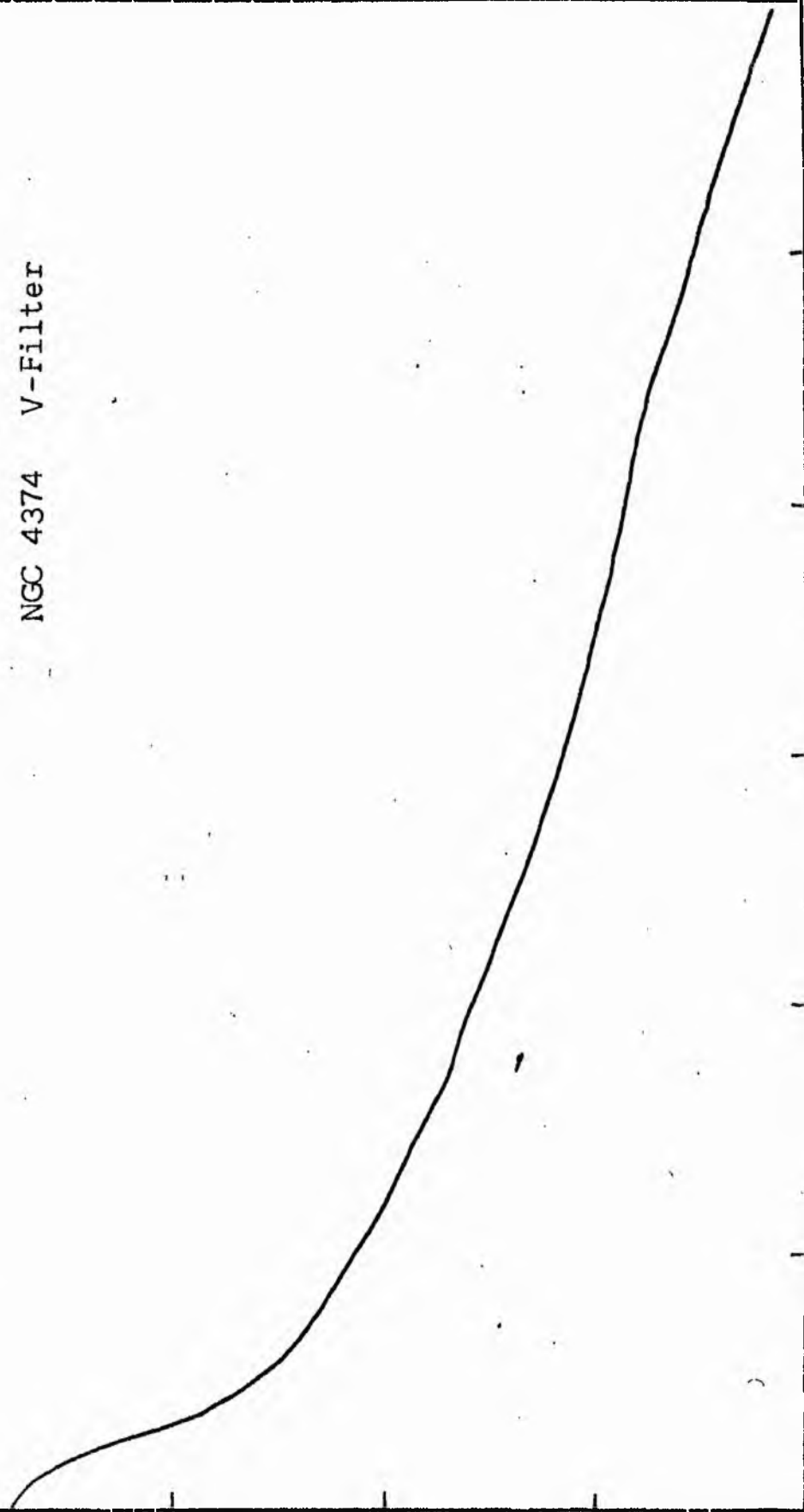


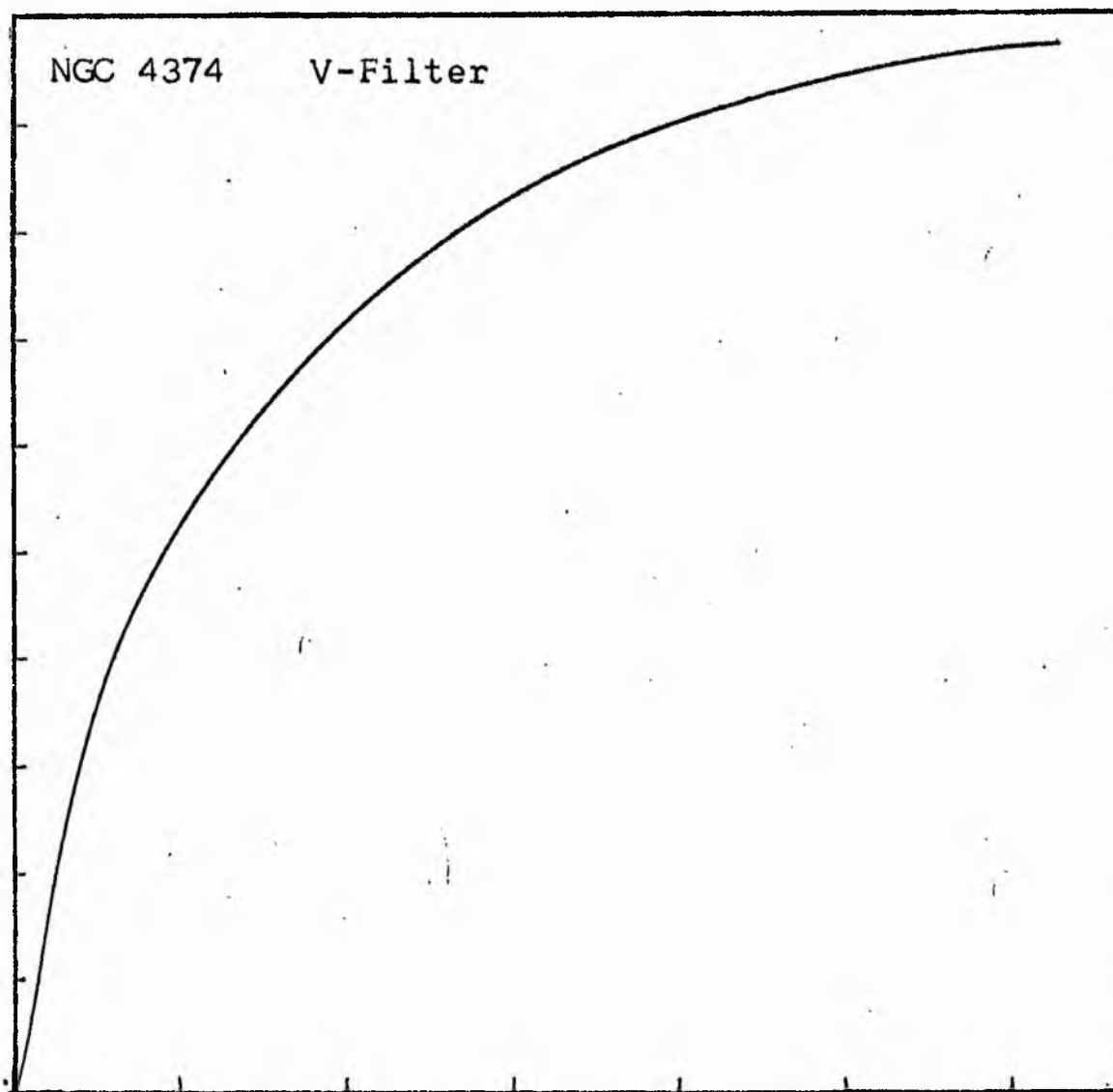
NGC 4374
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4374 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4374
V COLOUR

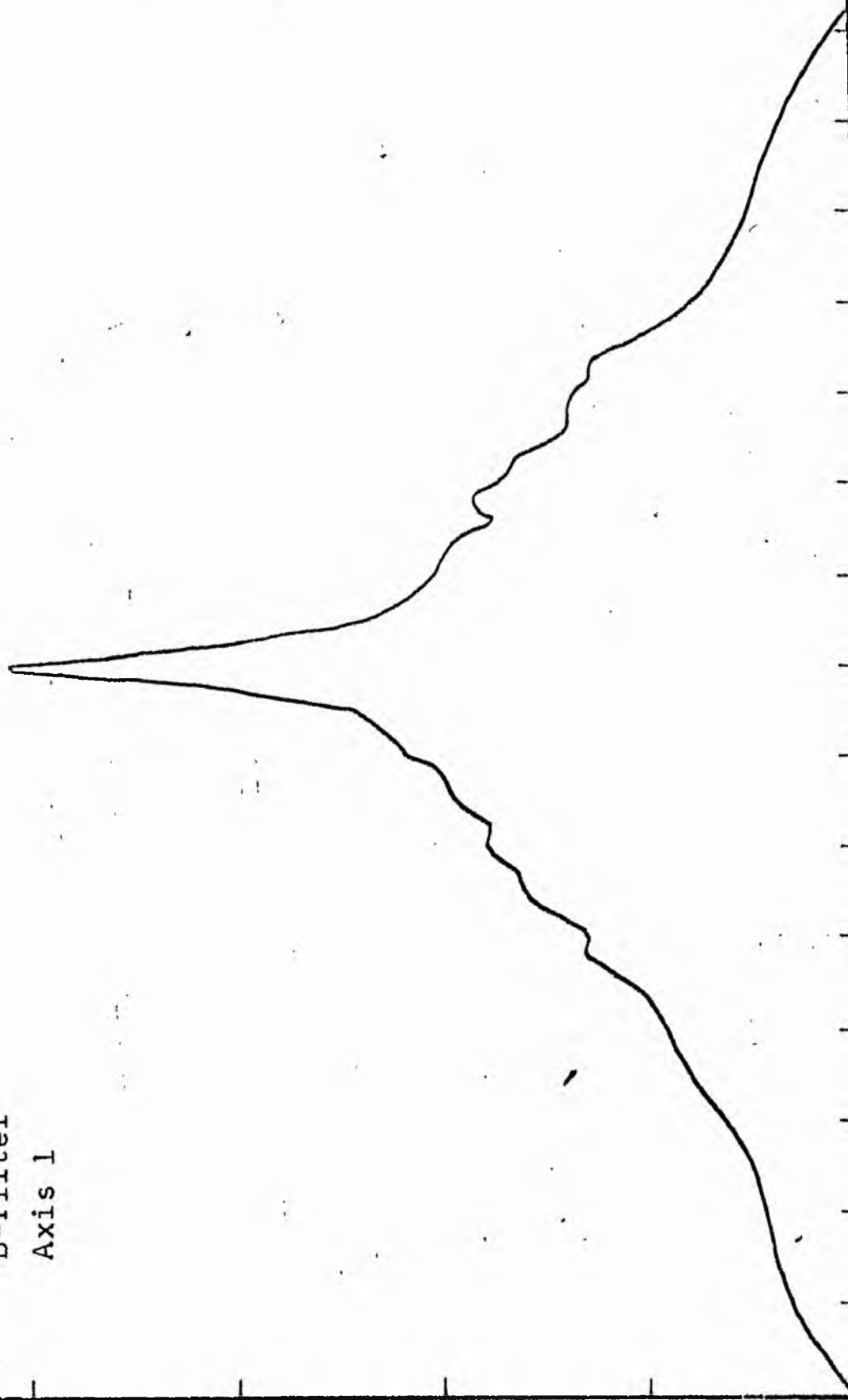
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.76	57.544		0.0	0.0			0.0	0.0	0.0	1.603	17.00
1.70	50.119	53.831	0.40	0.50	0.50	27.0585	27.06	0.00	0.01	1.543	17.23
1.60	39.811	44.965	2.10	13.85	13.35	600.3574	627.42	0.31	0.05	1.443	17.48
1.50	31.623	35.717	5.30	88.25	74.39	2657.0657	3284.48	0.04	0.12	1.343	17.73
1.40	25.119	28.371	8.20	211.24	122.99	3489.4138	6773.89	0.09	0.18	1.243	17.98
1.30	19.953	22.536	12.80	514.72	303.48	6839.0664	13612.96	0.19	0.28	1.143	18.23
1.20	15.849	17.901	15.30	735.42	220.70	3950.6301	17563.59	0.24	0.34	1.043	18.48
1.10	12.589	14.219	16.10	814.33	78.92	1122.0969	18685.68	0.26	0.35	0.943	18.73
1.00	10.000	11.295	17.23	932.65	118.32	1336.3945	20022.08	0.27	0.38	0.843	18.98
0.90	7.943	8.972	18.91	1123.39	190.74	1711.2588	21733.34	0.30	0.42	0.743	19.23
0.80	6.310	7.126	21.20	1411.96	288.56	2056.4023	23789.74	0.33	0.47	0.643	19.48
0.70	5.012	5.661	23.50	1734.94	322.99	1828.3347	25618.07	0.35	0.52	0.543	19.73
0.60	3.981	4.496	25.80	2091.17	356.22	1601.7356	27219.80	0.37	0.57	0.443	19.98
0.50	3.162	3.572	26.90	2273.29	182.12	650.4680	27870.27	0.38	0.59	0.343	20.23
0.40	2.512	2.837	30.60	2941.66	668.37	1896.2187	29766.49	0.41	0.67	0.243	20.48
0.30	1.995	2.254	37.10	4324.12	1382.46	3115.4543	32881.94	0.45	0.82	0.143	20.73
0.20	1.585	1.790	42.80	5754.89	1430.77	2561.1826	35443.12	0.48	0.94	0.043	20.98
0.10	1.259	1.422	49.60	7728.82	1973.93	2826.7283	38249.85	0.52	1.09	-0.057	21.23
-0.00	1.000	1.129	56.60	10064.28	2335.46	2637.8015	40887.65	0.56	1.25	-0.157	21.48
-0.10	0.794	0.897	66.50	13892.91	3828.63	3434.8909	44322.54	0.61	1.46	-0.257	21.73
-0.20	0.631	0.713	76.10	18193.61	4300.71	3064.8528	47387.39	0.65	1.67	-0.357	21.98
-0.30	0.501	0.566	84.20	22272.75	4079.14	2309.0752	49696.46	0.68	1.85	-0.457	22.23
-0.40	0.398	0.450	92.20	26706.16	4433.41	1993.4631	51689.93	0.71	2.03	-0.557	22.48
-0.50	0.316	0.357	102.50	33006.36	6300.20	2250.2151	53940.14	0.74	2.26	-0.657	22.73
-0.60	0.251	0.284	114.60	41259.03	8252.67	2341.3413	56281.48	0.77	2.52	-0.757	22.98
-0.70	0.200	0.225	126.80	50511.27	9252.23	2085.0518	58366.53	0.80	2.79	-0.857	23.23
-0.80	0.158	0.179	138.80	60524.15	10012.88	1792.3765	60158.91	0.82	3.05	-0.957	23.48
-0.90	0.126	0.142	151.90	72487.87	11963.73	1701.1270	61860.03	0.85	3.34	-1.057	23.73
-1.00	0.100	0.113	165.70	86257.06	13769.19	1555.1733	63415.20	0.87	3.65	-1.157	23.98
-1.10	0.079	0.090	182.40	104520.00	18262.94	1638.4822	65053.68	0.89	4.01	-1.257	24.23
-1.20	0.063	0.071	202.70	129079.50	24559.50	1750.2131	66803.87	0.91	4.46	-1.357	24.48
-1.30	0.050	0.057	216.80	147661.87	18582.37	1051.8975	67855.75	0.93	4.77	-1.457	24.73
-1.40	0.040	0.045	230.90	167493.37	19831.50	891.7197	68747.44	0.94	5.08	-1.557	24.98
-1.50	0.032	0.036	246.80	191355.12	23861.75	852.2673	69599.69	0.95	5.43	-1.657	25.23
-1.60	0.025	0.028	260.70	213516.56	22161.44	628.7417	70228.37	0.96	5.74	-1.757	25.48
-1.70	0.020	0.023	273.60	235169.81	21653.25	487.9753	70716.31	0.97	6.02	-1.857	25.73
-1.80	0.016	0.018	288.60	261662.87	26493.06	474.2500	71190.56	0.97	6.35	-1.957	25.98
-1.90	0.013	0.014	299.70	282177.87	20515.00	291.7073	71482.25	0.98	6.59	-2.057	26.23
-2.00	0.010	0.011	308.90	299768.00	17590.12	198.6760	71680.87	0.98	6.80	-2.157	26.48
							73130.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4374

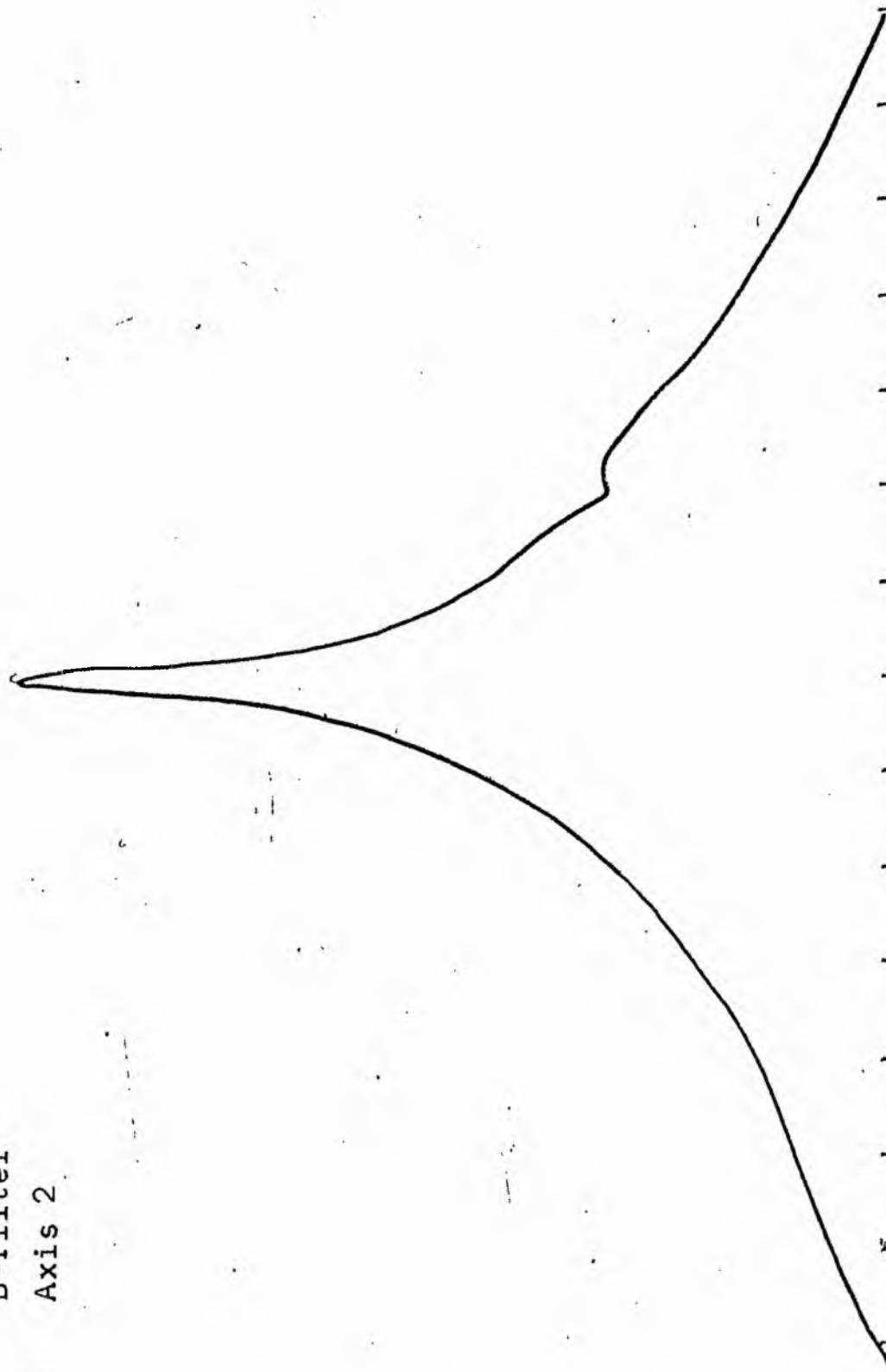
V-FILTER

Total luminosity	L_T	= 20.31
Total apparent magnitude	m_T	= 9.33
Apparent central surface brightness	μ_0	= 17.08
Major axis at threshold	$2a_m$	= 10.20
Minor axis at threshold	$2b_m$	= 10.12
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 7.72
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.94
Gradient of exponential component	$G(a)$	= -0.50
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.43
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.31
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 1.05
Equivalent radius	r_1^*	= 0.26
Surface brightness	μ_1	= 18.61
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.83
Axis ratio	b/a	= 0.90
Equivalent radius	r_e^*	= 0.76
Surface brightness	μ_e	= 21.10
Mean surface brightness	μ_e'	= 10.78
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.80
Axis ratio	b/a	= 0.88
Equivalent radius	r_3^*	= 1.78
Surface brightness	μ_3	= 22.81
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.88 \\ = 2.35 \end{matrix}$

NGC 4382
B-Filter
Axis 1

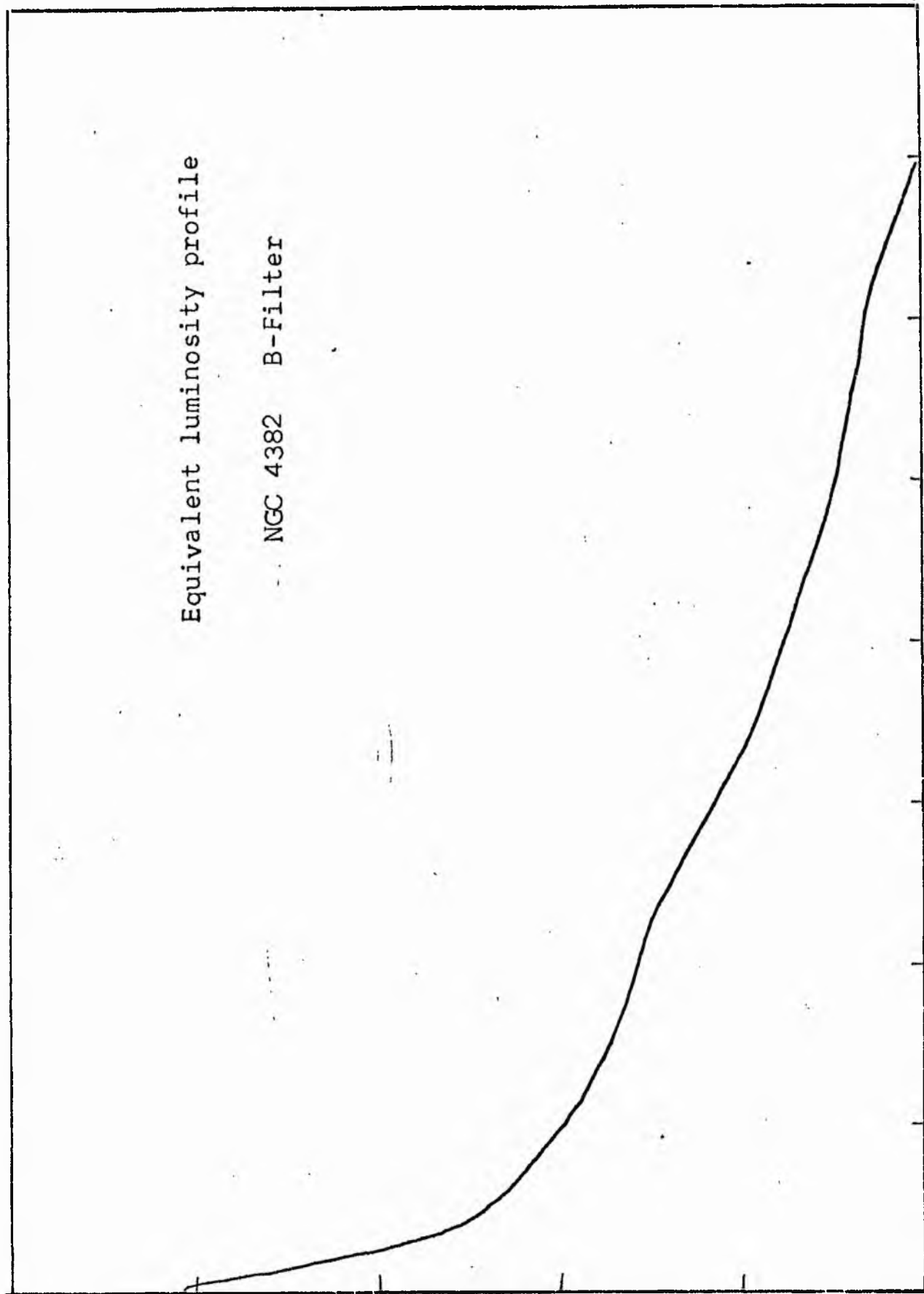


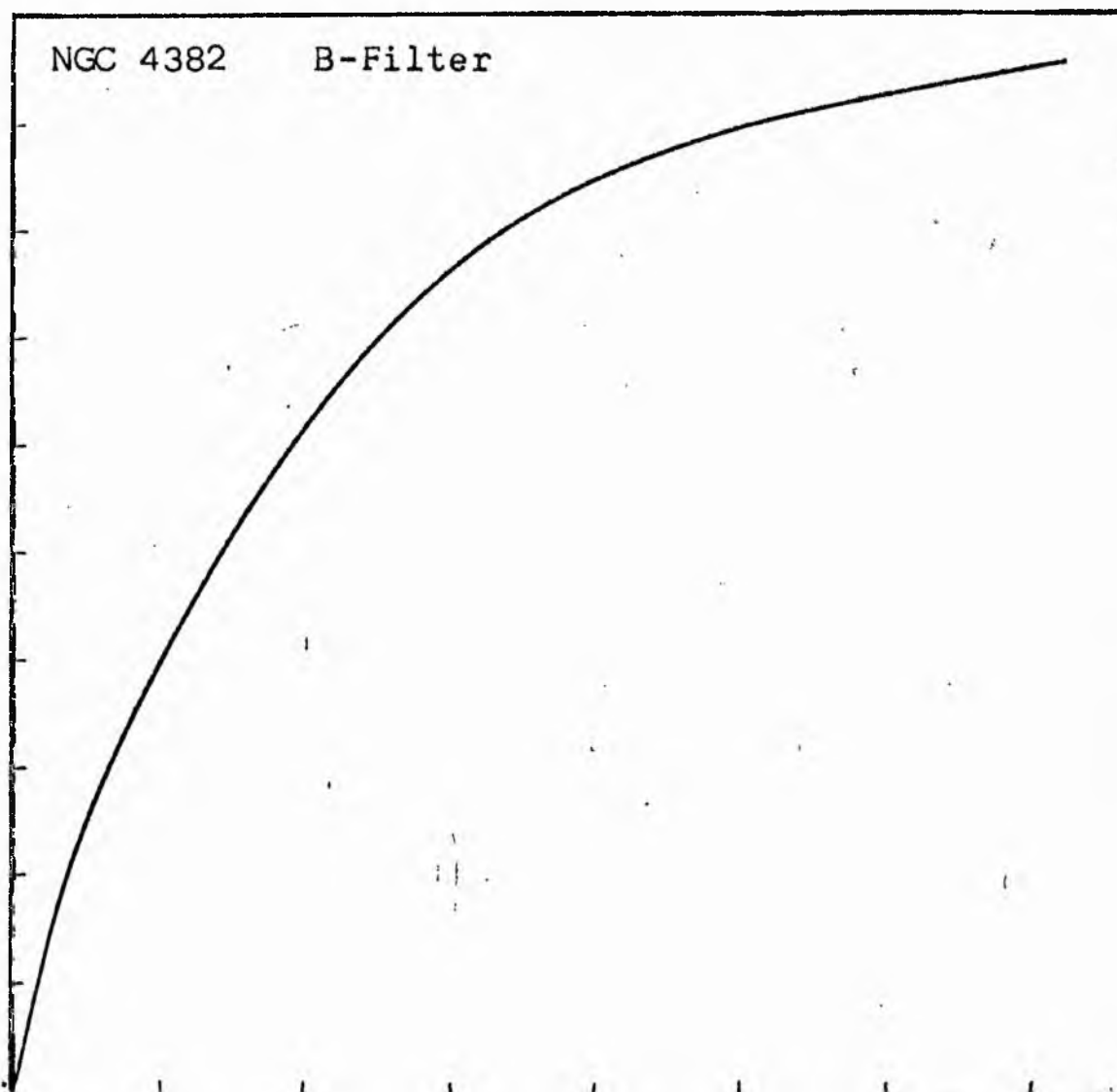
NGC 4382
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4382 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4302
B COLOUR

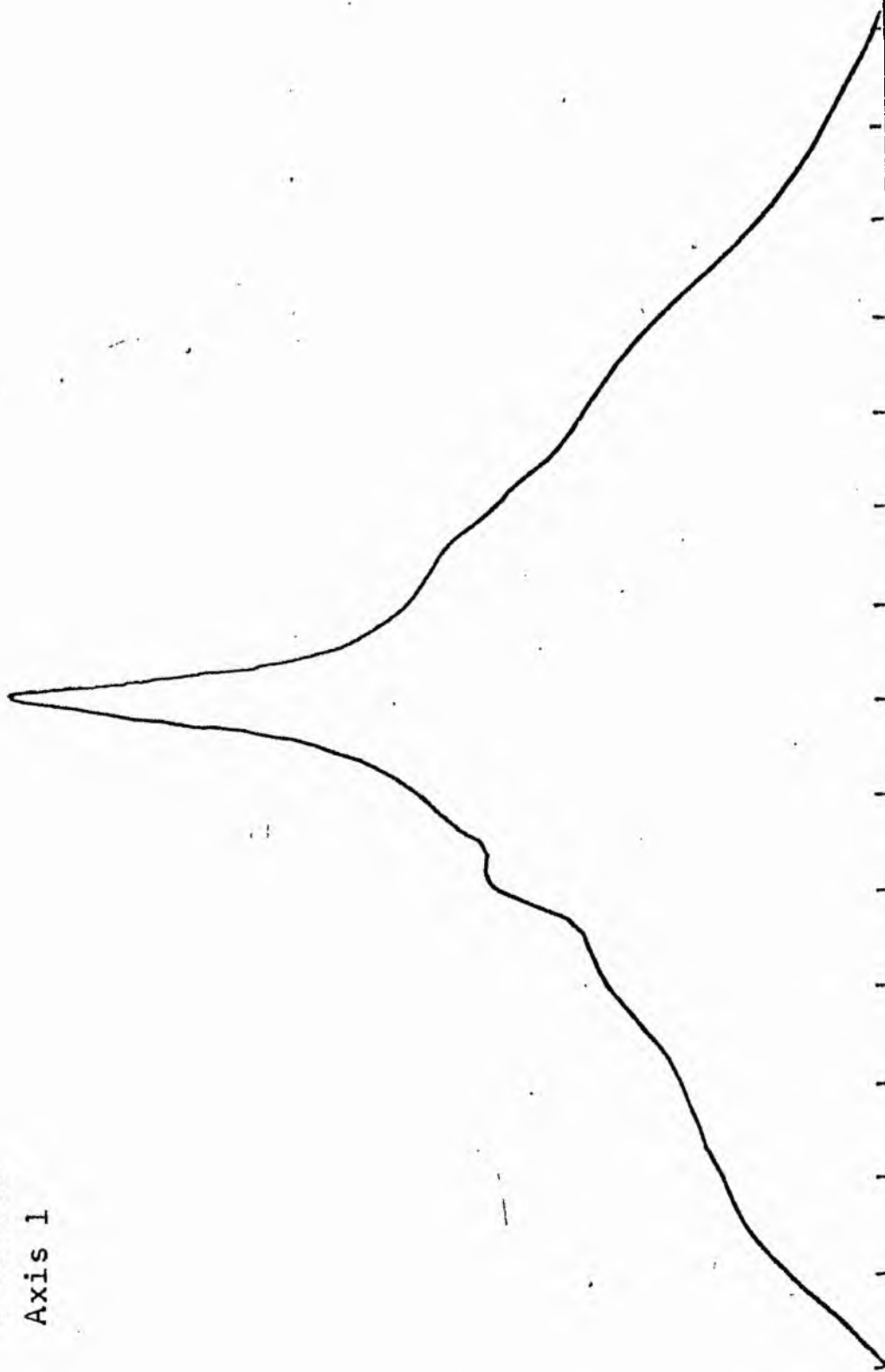
LOG I	I	T	n	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.09	123.027		0.0	0.0			0.0	0.0	0.0	2.337	16.79
2.00	100.000	111.513	0.90	2.54	2.54	283.7666	283.77	0.00	0.01	2.247	17.02
1.90	79.433	89.716	1.80	10.18	7.63	684.8999	968.67	0.02	0.03	2.147	17.27
1.80	63.096	71.264	2.80	24.63	14.45	1029.8608	1998.53	0.03	0.04	2.047	17.52
1.70	50.119	56.607	3.60	40.72	16.08	910.5227	2909.05	0.05	0.05	1.947	17.77
1.60	39.811	44.965	4.70	69.40	20.68	1289.7060	4198.75	0.07	0.07	1.847	18.02
1.50	31.623	35.717	5.66	100.64	31.24	1115.9648	5314.72	0.09	0.08	1.747	18.27
1.40	25.119	28.371	6.57	135.61	34.96	991.9441	6306.66	0.10	0.09	1.647	18.52
1.30	19.952	22.536	8.02	202.07	66.46	1497.7607	7804.42	0.13	0.11	1.547	18.77
1.20	15.849	17.901	8.47	225.38	23.31	417.3022	8221.72	0.13	0.12	1.447	19.02
1.10	12.589	14.219	8.93	250.53	25.15	357.5413	8579.26	0.14	0.13	1.347	19.27
1.00	10.000	11.295	9.87	306.04	55.52	627.0518	9206.31	0.15	0.14	1.247	19.52
0.90	7.943	0.972	12.10	459.96	153.92	1380.8696	10587.18	0.17	0.17	1.147	19.77
0.80	6.310	7.126	14.10	624.58	164.62	1173.1382	11760.31	0.19	0.20	1.047	20.02
0.70	5.012	5.661	16.11	815.34	190.76	1079.8442	12840.16	0.21	0.23	0.947	20.27
0.60	3.981	4.496	18.93	1125.77	310.43	1395.8291	14235.98	0.23	0.27	0.847	20.52
0.50	3.162	3.572	20.12	1271.76	145.99	521.4175	14757.40	0.24	0.29	0.747	20.77
0.40	2.512	2.837	23.84	1705.51	513.75	1457.5364	16214.93	0.27	0.34	0.647	21.02
0.30	1.995	2.254	28.51	2553.55	768.04	1730.8201	17945.75	0.29	0.41	0.547	21.27
0.20	1.565	1.790	34.00	3631.68	1078.13	1929.9285	19875.66	0.33	0.49	0.447	21.52
0.10	1.259	1.422	41.78	5483.86	1852.18	2633.6152	22509.29	0.37	0.60	0.347	21.77
-0.00	1.000	1.129	49.17	7595.39	2111.53	2384.8743	24894.16	0.41	0.70	0.247	22.02
-0.10	0.794	0.897	54.80	9434.32	1830.93	1649.8132	26543.97	0.44	0.76	0.147	22.27
-0.20	0.631	0.713	64.40	13029.30	3594.98	2561.9165	29105.89	0.48	0.92	0.047	22.52
-0.30	0.501	0.566	78.50	19359.28	6329.98	3583.2004	32609.09	0.54	1.12	-0.053	22.77
-0.40	0.398	0.450	91.50	26302.20	6942.92	3121.8423	35810.93	0.59	1.31	-0.153	23.02
-0.50	0.316	0.357	108.90	37256.80	10954.60	3912.6018	39723.53	0.65	1.56	-0.253	23.27
-0.60	0.251	0.284	124.10	48383.06	11126.26	3156.5908	42880.12	0.70	1.78	-0.353	23.52
-0.70	0.200	0.225	136.40	58449.20	10066.14	2268.4656	45148.58	0.74	1.95	-0.453	23.77
-0.80	0.158	0.179	144.00	65144.07	6694.87	1198.4265	46347.01	0.76	2.06	-0.553	24.02
-0.90	0.126	0.142	154.00	74506.00	9361.95	1331.1731	47678.18	0.78	2.21	-0.653	24.27
-1.00	0.100	0.113	163.90	84393.25	9887.25	1116.7219	48794.90	0.80	2.35	-0.753	24.52
-1.10	0.079	0.090	176.40	97756.81	13363.56	1198.9263	49993.82	0.82	2.53	-0.853	24.77
-1.20	0.063	0.071	191.70	115450.00	17693.19	1260.8887	51254.71	0.84	2.75	-0.953	25.02
-1.30	0.050	0.057	206.50	133964.56	18514.96	1048.0566	52302.77	0.86	2.96	-1.053	25.27
-1.40	0.040	0.045	219.80	151776.69	17012.12	800.9172	53103.68	0.87	3.15	-1.153	25.52
-1.50	0.032	0.036	239.70	100503.62	28726.94	1026.0342	54129.71	0.89	3.43	-1.253	25.77
-1.60	0.025	0.028	271.40	231403.06	50899.44	1444.0627	55573.77	0.91	3.89	-1.353	26.02
-1.70	0.020	0.023	300.50	283686.62	52283.56	1178.2537	56752.02	0.93	4.30	-1.453	26.27
-1.80	0.016	0.018	327.00	335927.37	52240.75	935.1553	57687.18	0.95	4.68	-1.553	26.52
-1.90	0.013	0.014	345.10	374144.44	38217.06	543.4158	58230.59	0.95	4.94	-1.653	26.77
-2.00	0.010	0.011	352.40	390140.87	15996.44	180.6753	58411.26	0.96	5.05	-1.753	27.02
-∞							61011.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NGC 4382

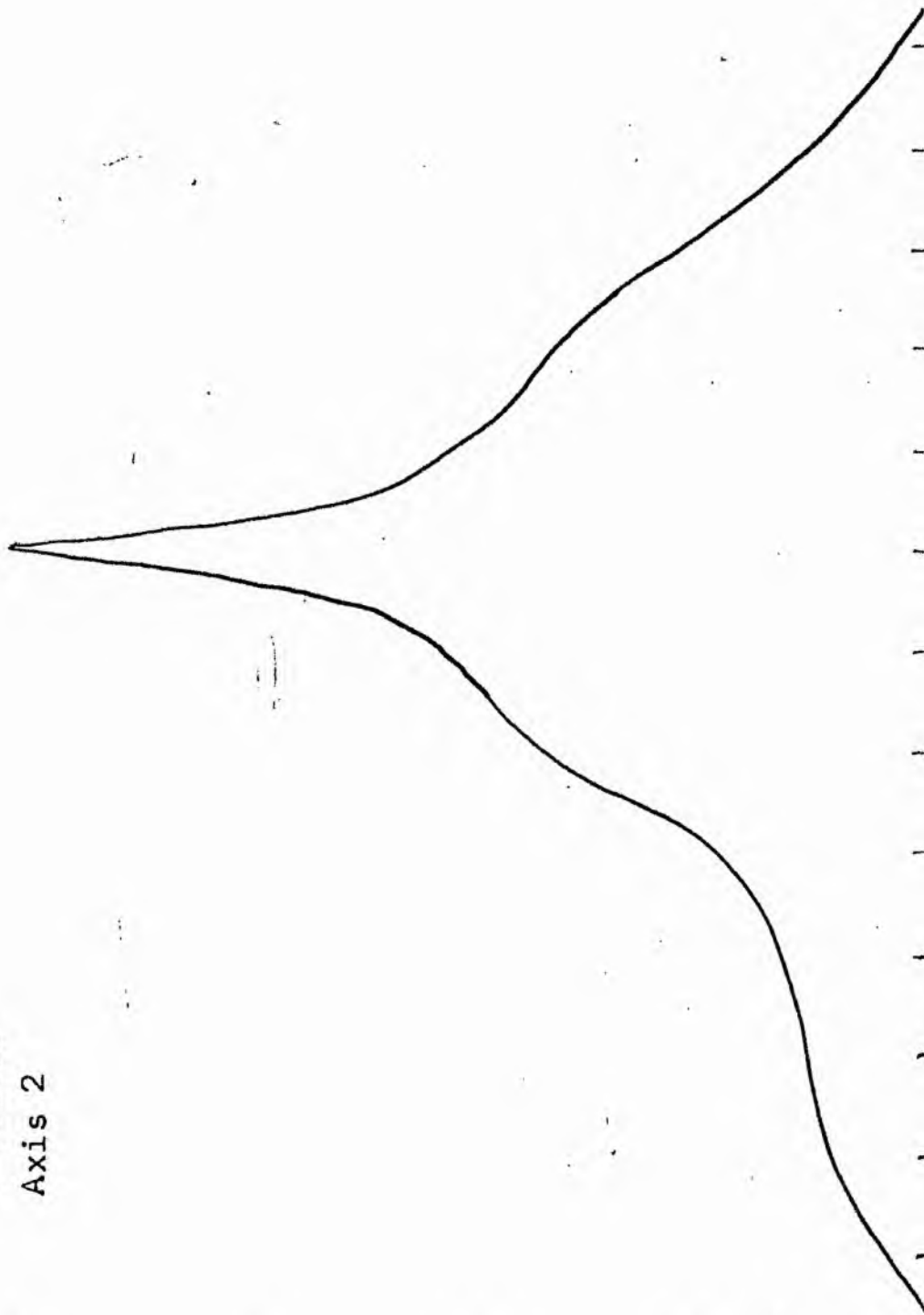
B - FILTER

Total luminosity	L_T	= 16.95
Total apparent magnitude	m_T	= 9.98
Apparent central surface brightness	μ_0	= 16.79
Major axis at threshold	$2a_m$	= 12.83
Minor axis at threshold	$2b_m$	= 10.64
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 6.85
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.84
Gradient of exponential component	$G(a)$	= -0.46
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.43
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.63
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.40
Axis ratio	b/a	= 0.96
Equivalent radius	r_1^*	= 0.36
Surface brightness	μ_1	= 20.84
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.57
Axis ratio	b/a	= 0.66
Equivalent radius	r_e^*	= 1.16
Surface brightness	μ_e	= 22.60
Mean surface brightness	μ_e'	= 12.30
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.55
Axis ratio	b/a	= 0.76
Equivalent radius	r_3^*	= 2.34
Surface brightness	μ_3	= 23.89
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.27 \\ = 2.01 \end{cases}$

NGC 4382
V-Filter
Axis 1

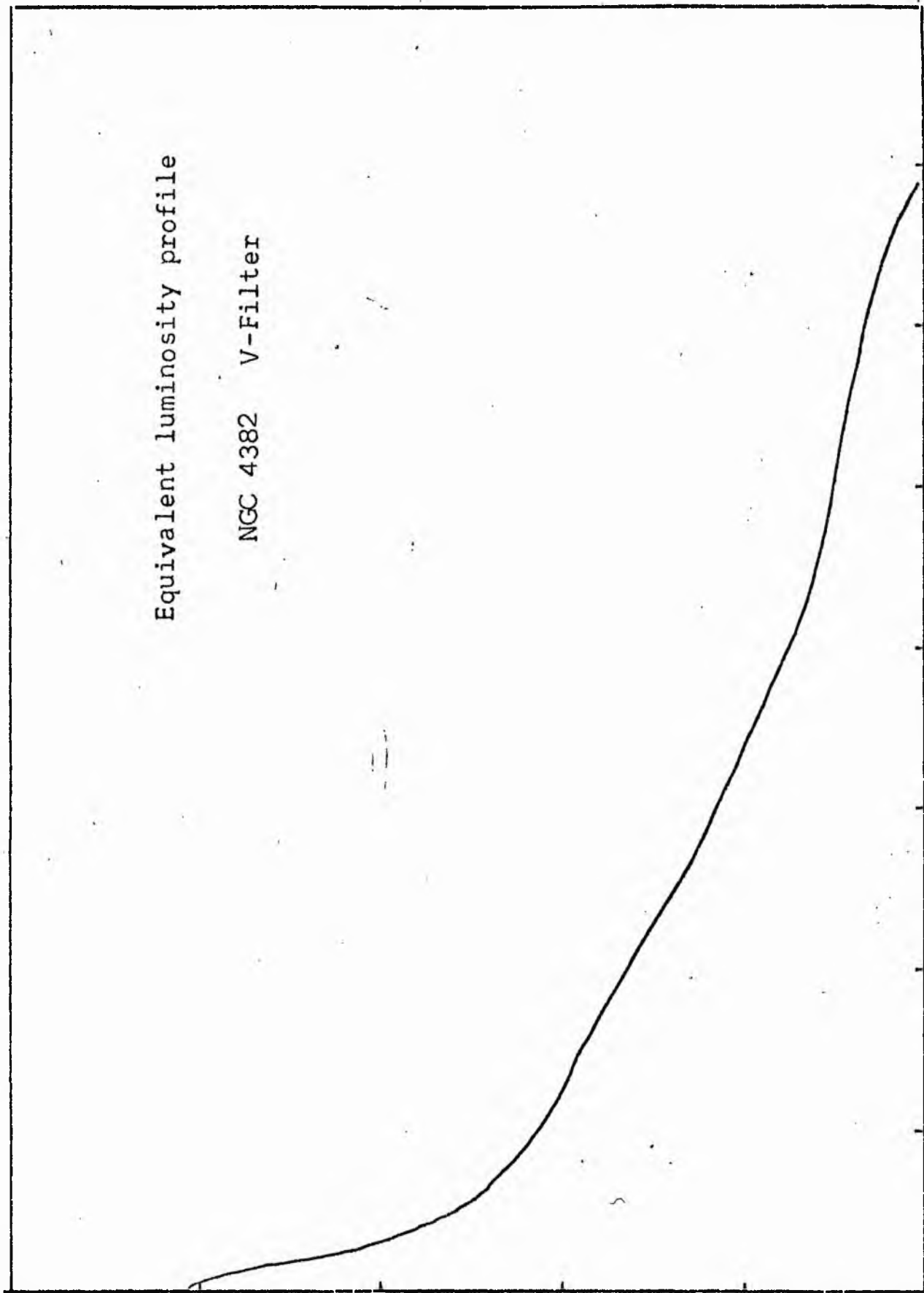


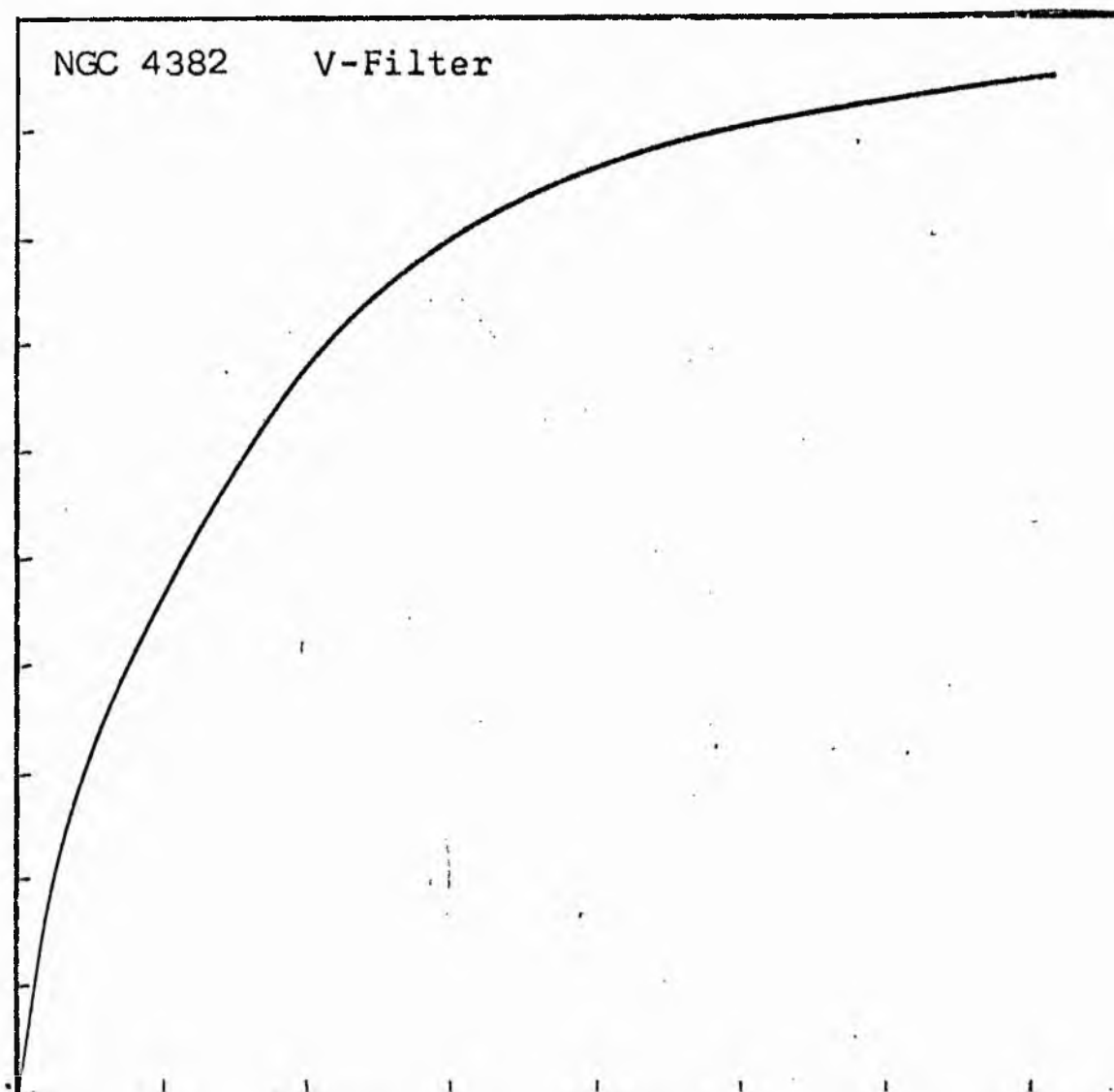
NGC 4382
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4382 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4382
V COLOUR

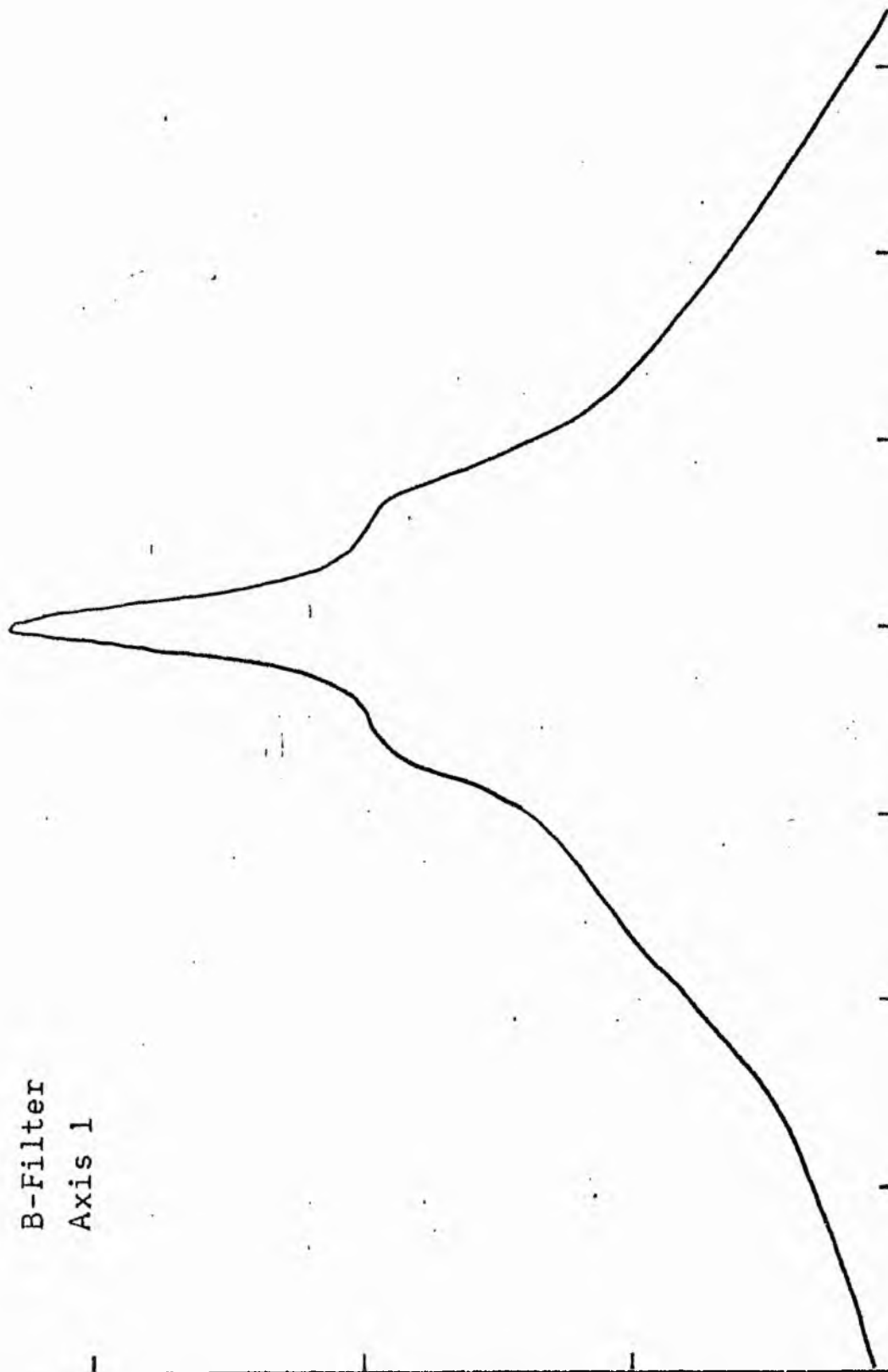
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.08	120.226		0.0	0.0			0.0	0.0	0.0	2.037	16.28
2.00	100.000	110.113	1.10	3.80	3.80	418.5754	418.58	0.31	0.02	1.957	16.48
1.90	79.433	89.716	2.30	16.62	12.82	1149.9556	1568.53	0.02	0.04	1.857	16.73
1.80	63.096	71.264	3.90	47.78	31.16	2220.9202	3749.45	0.05	0.07	1.757	16.98
1.70	50.119	56.607	5.00	78.54	30.76	1741.0186	5530.47	0.07	0.09	1.657	17.23
1.60	39.811	44.965	6.20	120.76	42.22	1898.5378	7429.00	0.09	0.11	1.557	17.48
1.50	31.623	35.717	7.86	194.09	73.32	2618.8665	10047.87	0.12	0.14	1.457	17.73
1.40	25.119	28.371	8.57	230.73	36.65	1039.7175	11087.58	0.13	0.15	1.357	17.98
1.30	19.952	22.536	10.55	349.67	118.93	2680.2275	13767.81	0.17	0.18	1.257	18.23
1.20	15.849	17.901	12.22	469.13	119.46	2138.4458	15906.25	0.19	0.21	1.157	18.48
1.10	12.589	14.219	13.54	575.95	106.82	1518.9312	17425.18	0.21	0.24	1.057	18.73
1.00	10.000	11.295	15.45	749.91	173.95	1964.7151	19389.90	0.24	0.27	0.957	18.98
0.90	7.943	8.972	17.34	944.60	194.69	1746.7041	21136.60	0.26	0.30	0.857	19.23
0.80	6.310	7.126	19.12	1148.48	203.89	1452.9614	22589.56	0.27	0.33	0.757	19.48
0.70	5.012	5.661	21.46	1446.80	298.32	1688.6709	24278.23	0.29	0.37	0.657	19.73
0.60	3.981	4.496	23.95	1802.02	355.22	1597.2422	25875.47	0.31	0.42	0.557	19.98
0.50	3.162	3.572	25.80	2091.17	289.14	1032.7141	26908.18	0.33	0.45	0.457	20.23
0.40	2.512	2.837	31.20	3058.15	966.98	2743.3921	29651.57	0.36	0.54	0.357	20.48
0.30	1.995	2.254	37.27	4363.84	1305.68	2942.4365	32594.01	0.40	0.65	0.257	20.73
0.20	1.585	1.790	42.80	5754.89	1391.05	2490.0779	35084.08	0.43	0.74	0.157	20.98
0.10	1.259	1.422	51.40	8299.96	2545.07	3618.8269	38702.91	0.47	0.89	0.057	21.23
-0.00	1.000	1.129	62.20	12154.32	3854.36	4353.3203	43056.23	0.52	1.08	-0.043	21.48
-0.10	0.794	0.897	72.60	16558.57	4404.26	3951.3135	47007.54	0.57	1.26	-0.143	21.73
-0.20	0.631	0.713	82.70	21486.26	4927.69	3511.6567	50519.20	0.61	1.44	-0.243	21.98
-0.30	0.501	0.566	94.50	28055.21	6568.95	3718.4707	54237.66	0.66	1.64	-0.343	22.23
-0.40	0.398	0.450	107.34	36197.03	8141.82	3660.9216	57898.58	0.70	1.86	-0.443	22.48
-0.50	0.316	0.357	121.12	46087.33	9890.30	3532.4707	61431.05	0.74	2.10	-0.543	22.73
-0.60	0.251	0.284	127.80	51311.12	5223.79	1482.0234	62913.07	0.76	2.22	-0.643	22.98
-0.70	0.200	0.225	133.50	55990.25	4679.13	1054.4707	63967.54	0.78	2.32	-0.743	23.23
-0.80	0.158	0.179	144.20	65325.14	9334.89	1671.0071	65638.50	0.80	2.50	-0.843	23.48
-0.90	0.126	0.142	153.60	74119.44	8794.30	1250.4612	66888.94	0.81	2.67	-0.943	23.73
-1.00	0.100	0.113	177.12	98556.44	24437.00	2760.0535	69648.94	0.84	3.07	-1.043	23.98
-1.10	0.079	0.090	188.23	111378.25	12751.81	1144.0425	70792.94	0.86	3.27	-1.143	24.23
-1.20	0.063	0.071	202.46	128774.00	17465.75	1244.6807	72037.56	0.87	3.51	-1.243	24.48
-1.30	0.050	0.057	218.70	150261.37	21487.37	1216.3391	73253.87	0.89	3.80	-1.343	24.73
-1.40	0.040	0.045	237.80	177653.37	27392.00	1231.6738	74485.50	0.90	4.13	-1.443	24.98
-1.50	0.032	0.036	260.30	212861.56	35208.19	1257.5237	75743.00	0.92	4.52	-1.543	25.23
-1.60	0.025	0.028	285.40	255892.31	43030.75	1220.8208	76963.81	0.93	4.95	-1.643	25.48
-1.70	0.020	0.023	300.10	282931.44	27039.12	619.3491	77573.12	0.94	5.21	-1.743	25.73
-1.80	0.016	0.018	322.70	327150.56	44219.12	791.5610	78364.62	0.95	5.60	-1.843	25.98
-1.90	0.013	0.014	327.90	337778.75	10628.19	151.1243	78515.69	0.95	5.69	-1.943	26.23
-2.00	0.010	0.011	351.40	387929.75	50151.00	566.4414	79082.12	0.96	6.10	-2.043	26.48
-∞							82482.00	(11)			∞

PHOTOMETRIC PARAMETERS OF NGC 4382

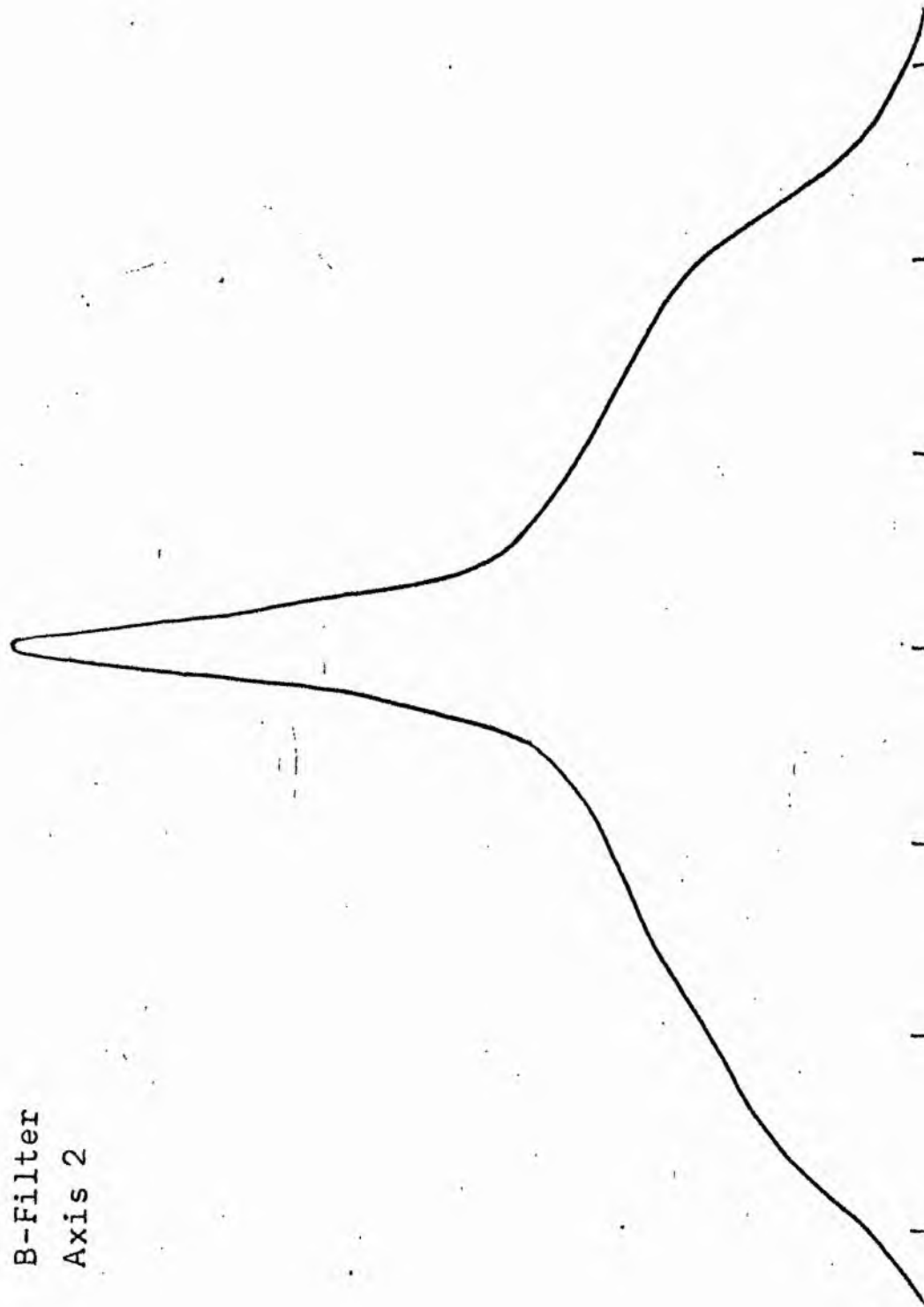
V-FILTER

Total luminosity	L_T	= 22.91
Total apparent magnitude	m_T	= 9.19
Apparent central surface brightness	μ_0	= 16.28
Major axis at threshold	$2a_m$	= 11.30
Minor axis at threshold	$2b_m$	= 10.62
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.72
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.90
Gradient of exponential component	$G(a)$	= -0.40
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.52
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.57
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.27
Axis ratio	b/a	= 1.05
Equivalent radius	r_1^*	= 0.28
Surface brightness	μ_1	= 19.11
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.05
Axis ratio	b/a	= 0.89
Equivalent radius	r_e^*	= 0.96
Surface brightness	μ_e	= 21.38
Mean surface brightness	μ_e'	= 11.05
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.45
Axis ratio	b/a	= 0.89
Equivalent radius	r_3^*	= 2.05
Surface brightness	μ_3	= 22.85
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.44 \\ = 2.14 \end{cases}$

NGC 4394
B-Filter
Axis 1

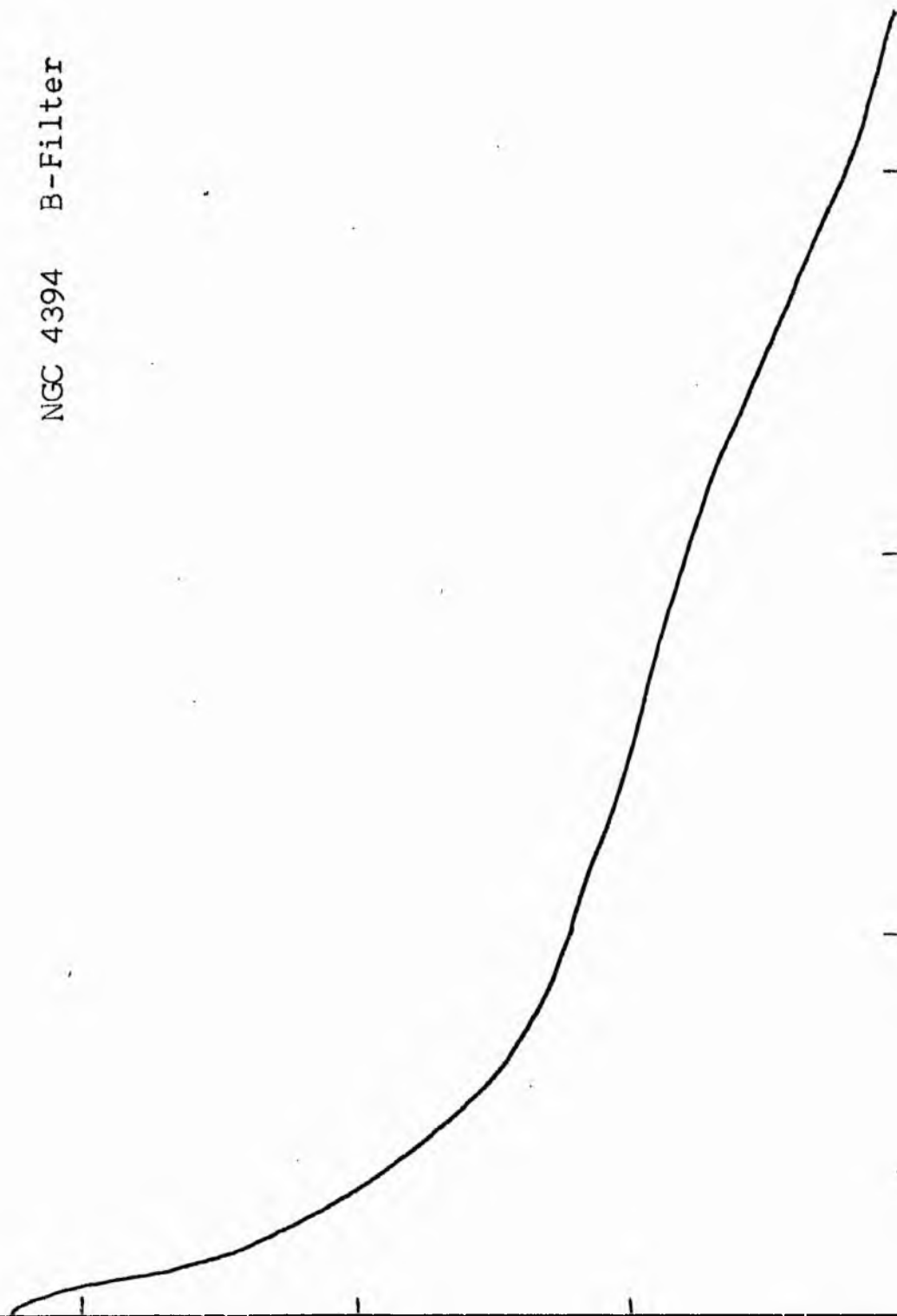


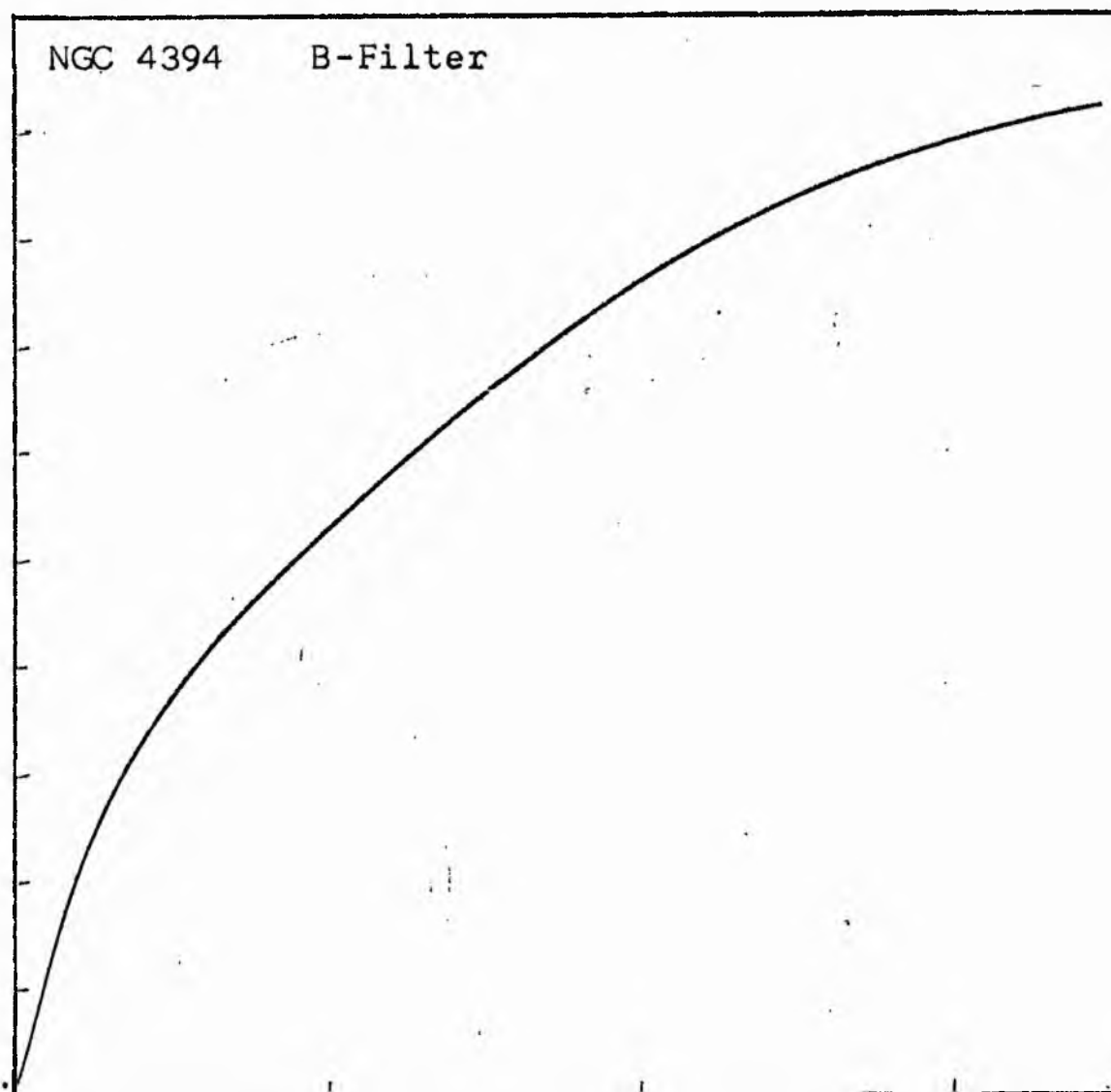
NGC 4394
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4394 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4394
B COLOUR

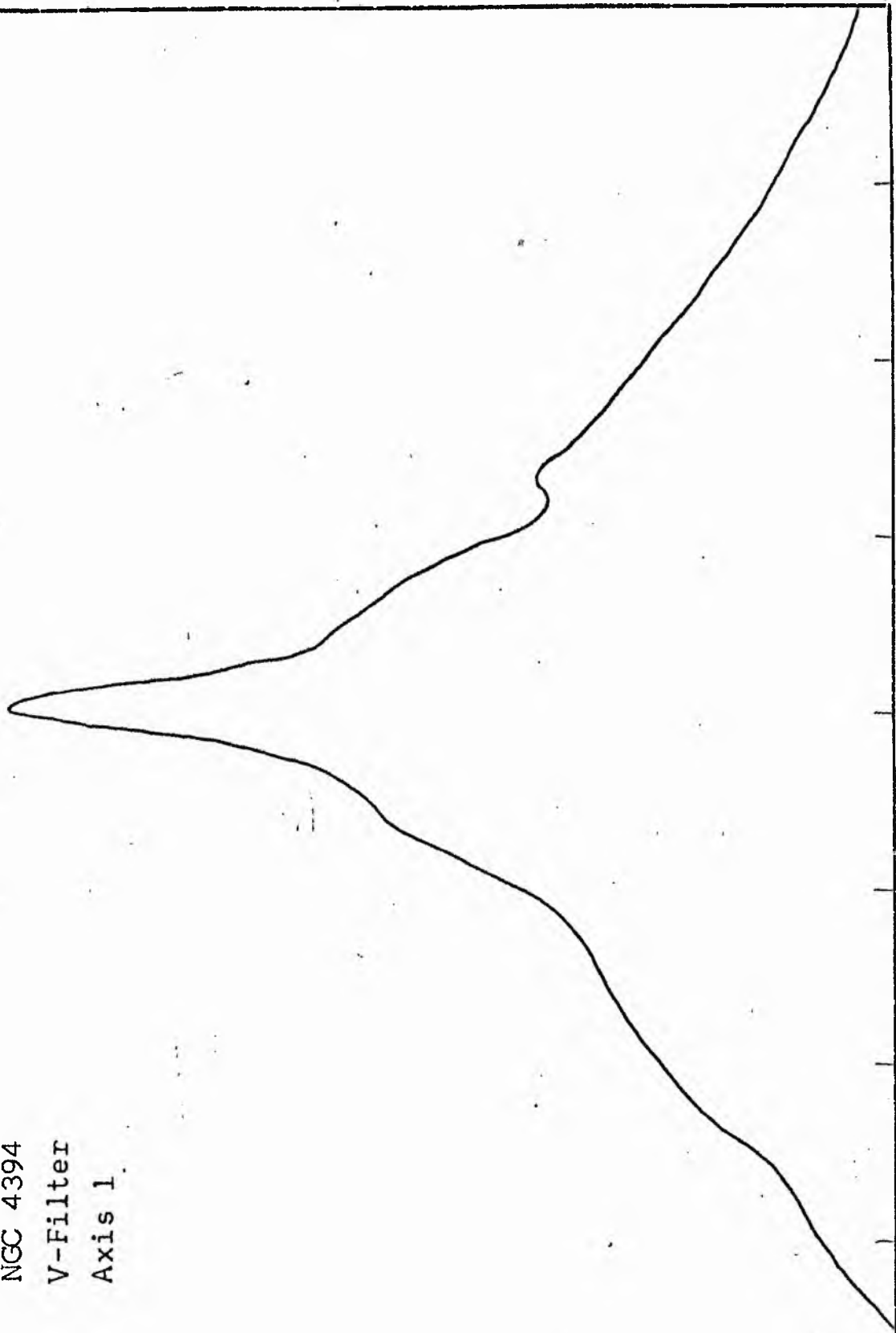
LOG I	I	T	R	AREA	ΔA	P	Σ P	K(I)	ρ	LOG J	μ
1.28	19.033		0.0	0.0			0.0	0.0	0.0	2.037	18.26
1.20	19.849	17.452	2.84	25.34	25.34	442.2065	442.21	0.05	0.06	1.957	18.46
1.10	12.589	14.219	3.40	40.72	15.38	218.6356	660.84	0.07	0.08	1.857	18.71
1.00	10.000	11.295	4.70	69.40	28.68	323.9600	984.80	0.10	0.10	1.757	18.96
0.90	7.943	8.972	5.30	88.25	18.85	169.1108	1153.91	0.12	0.11	1.657	19.21
0.80	6.310	7.126	5.80	103.68	17.44	124.2551	1278.17	0.13	0.12	1.557	19.46
0.70	5.012	5.661	6.20	120.76	15.08	85.3616	1363.53	0.14	0.13	1.457	19.71
0.60	3.981	4.496	7.10	158.37	37.60	169.0890	1532.62	0.16	0.15	1.357	19.96
0.50	3.162	3.572	7.94	198.06	39.69	141.7581	1674.38	0.17	0.17	1.257	20.21
0.40	2.512	2.837	9.01	255.03	56.98	161.6497	1836.03	0.19	0.19	1.157	20.46
0.30	1.995	2.254	10.48	345.04	90.01	202.8383	2038.86	0.21	0.22	1.057	20.71
0.20	1.585	1.790	12.75	510.71	165.66	296.5491	2335.41	0.24	0.27	0.957	20.96
0.10	1.259	1.422	15.03	709.69	198.98	282.9360	2618.35	0.27	0.32	0.857	21.21
-0.00	1.000	1.129	17.30	940.25	230.56	260.4055	2878.75	0.29	0.37	0.757	21.46
-0.10	0.794	0.897	19.37	1178.71	238.47	213.9454	3092.70	0.32	0.41	0.657	21.71
-0.20	0.631	0.713	21.99	1519.15	340.43	242.6066	3335.31	0.34	0.47	0.557	21.96
-0.30	0.501	0.566	24.97	1958.78	439.64	248.8656	3584.17	0.37	0.53	0.457	22.21
-0.40	0.398	0.450	26.88	2269.91	311.12	139.8958	3724.07	0.38	0.57	0.357	22.46
-0.50	0.316	0.357	31.02	3022.96	753.06	268.9678	3993.04	0.41	0.66	0.257	22.71
-0.60	0.251	0.284	36.20	4116.87	1093.90	310.3491	4303.38	0.44	0.77	0.157	22.96
-0.70	0.200	0.225	42.42	5653.16	1536.29	346.2141	4649.59	0.48	0.91	0.057	23.21
-0.80	0.158	0.179	52.00	8494.87	2841.71	508.6885	5158.28	0.53	1.11	-0.043	23.46
-0.90	0.126	0.142	63.80	12787.66	4292.79	610.3965	5768.68	0.59	1.36	-0.143	23.71
-1.00	0.100	0.113	77.40	18820.52	6032.86	681.3892	6450.06	0.66	1.65	-0.243	23.96
-1.10	0.079	0.090	88.50	24605.74	5785.22	519.0295	6969.09	0.71	1.89	-0.343	24.21
-1.20	0.063	0.071	99.30	30977.62	6371.89	454.0889	7423.18	0.76	2.12	-0.443	24.46
-1.30	0.050	0.057	109.70	37806.20	6828.57	386.5476	7809.72	0.80	2.34	-0.543	24.71
-1.40	0.040	0.045	116.50	42638.48	4832.29	217.2833	8027.00	0.82	2.49	-0.643	24.96
-1.50	0.032	0.036	125.30	49323.28	4684.79	238.7604	8265.76	0.85	2.68	-0.743	25.21
-1.60	0.025	0.028	132.40	55071.36	5748.09	163.0791	8428.84	0.86	2.83	-0.843	25.46
-1.70	0.020	0.023	140.20	61751.27	6679.91	150.5381	8579.37	0.88	3.00	-0.943	25.71
-1.80	0.016	0.018	149.70	70403.31	8652.04	154.8798	8734.25	0.89	3.20	-1.043	25.96
-1.90	0.013	0.014	163.50	83981.81	13578.50	193.0761	8927.33	0.91	3.49	-1.143	26.21
-2.00	0.010	0.011	174.80	95991.44	12009.62	135.6439	9062.97	0.93	3.73	-1.243	26.46
-∞							9772.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4394

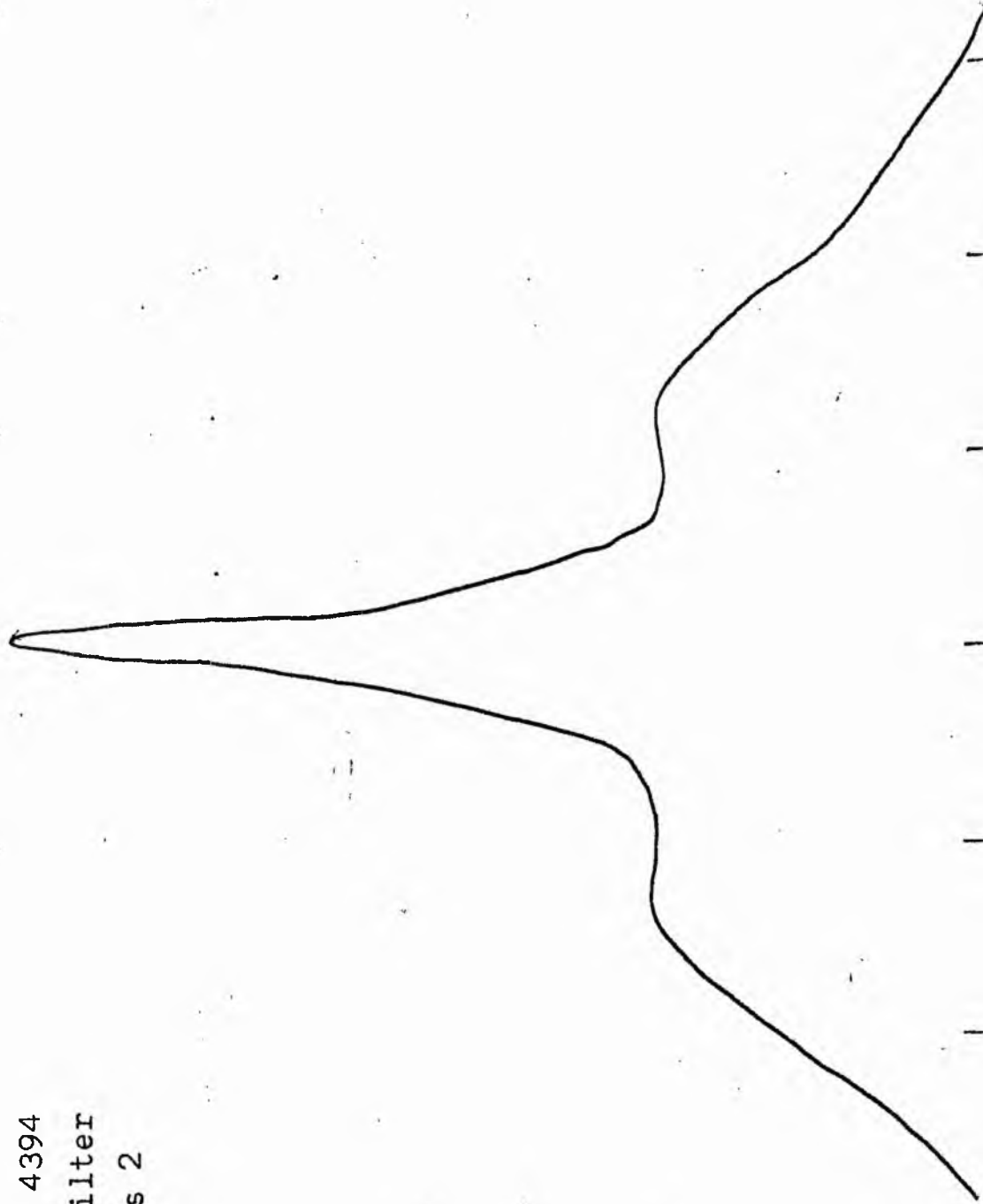
B-FILTER

Total luminosity	L_T	= 2.71
Total apparent magnitude	m_T	= 11.48
Apparent central surface brightness	μ_o	= 18.26
Major axis at threshold	$2a_m$	= 6.13
Minor axis at threshold	$2b_m$	= 5.56
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.03
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.84
Gradient of exponential component	$G(a)$	= -0.96
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.65
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.50
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.23
Axis ratio	b/a	= 0.94
Equivalent radius	r_1^*	= 0.23
Surface brightness	μ_1	= 21.04
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.08
Axis ratio	b/a	= 0.62
Equivalent radius	r_e^*	= 0.78
Surface brightness	μ_e	= 23.31
Mean surface brightness	μ_e'	= 12.93
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.56
Axis ratio	b/a	= 1.04
Equivalent radius	r_3^*	= 1.62
Surface brightness	μ_3	= 24.40
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.44 \\ = 2.07 \end{cases}$

NGC 4394
V-Filter
Axis 1.

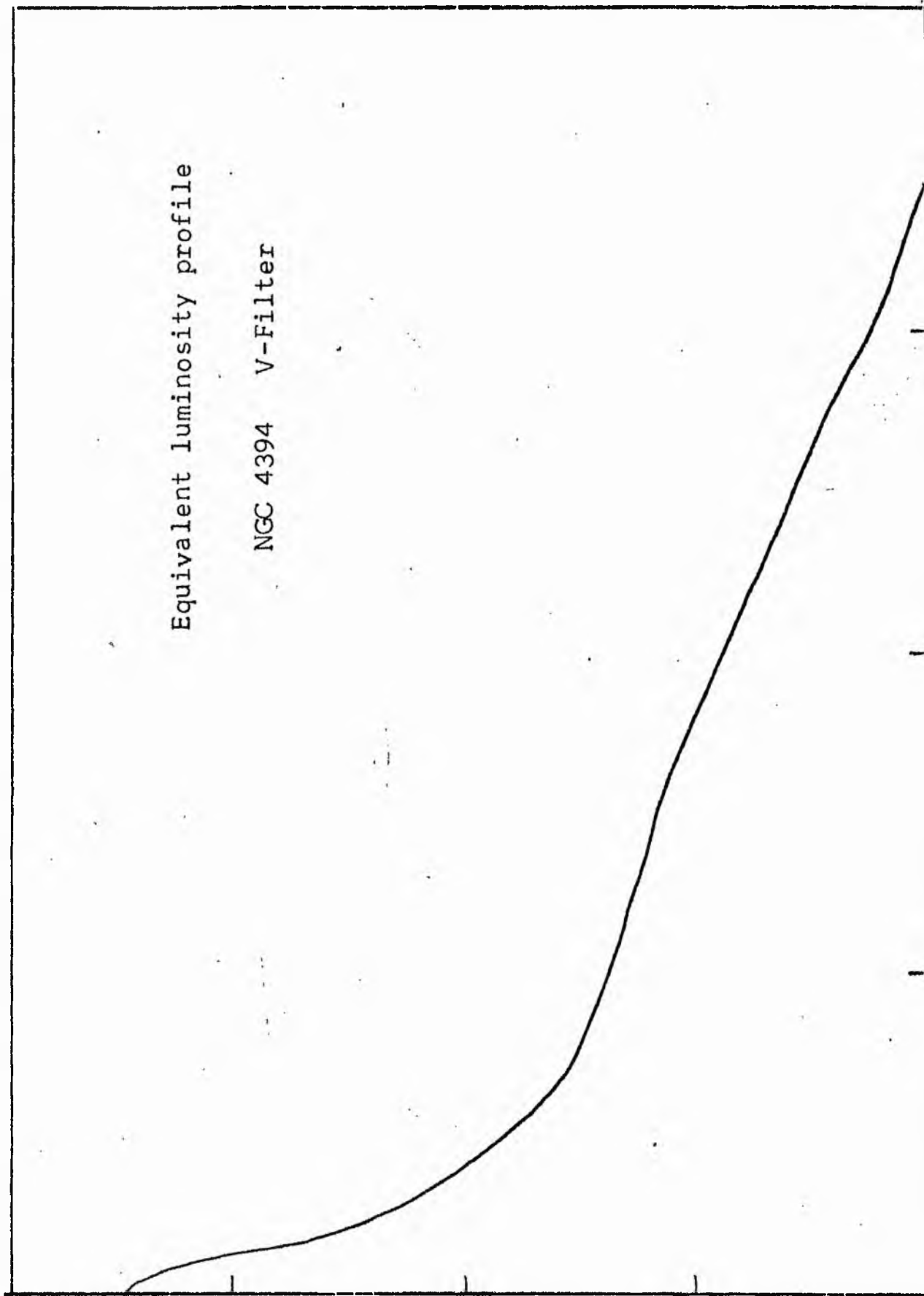


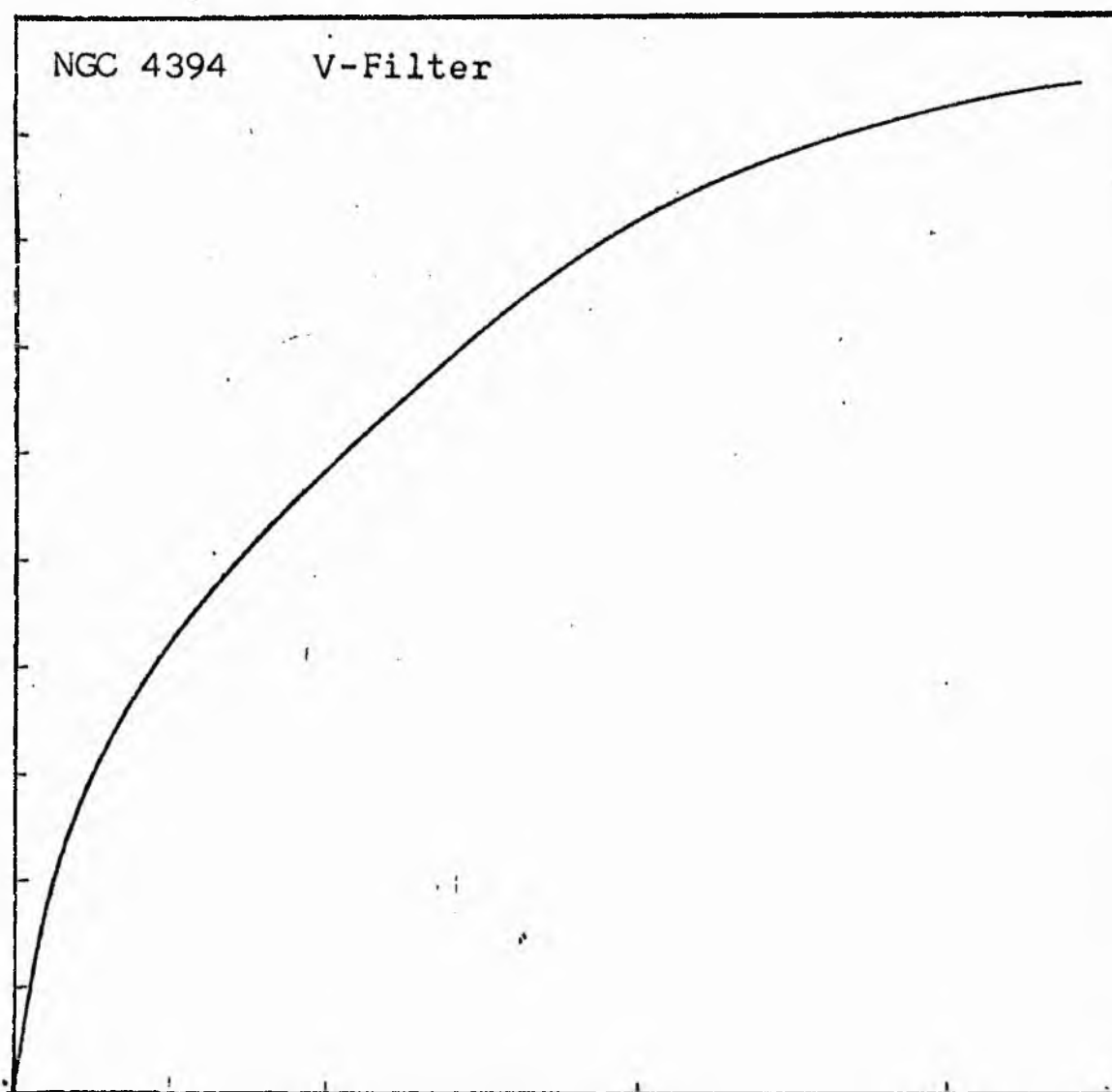
NGC 4394
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4394 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4394
V COLOUR

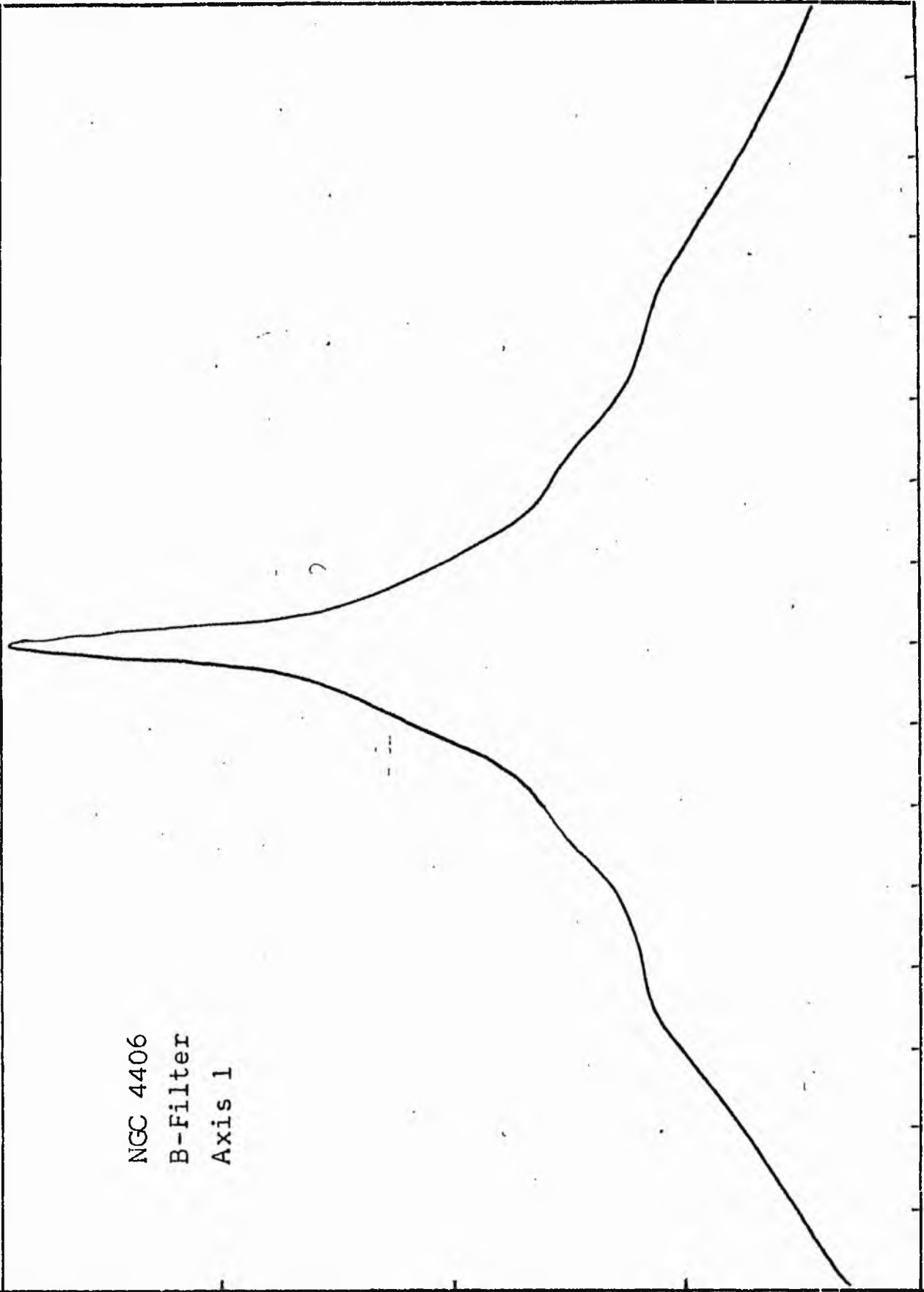
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.45	28.184		0.0	0.0			0.0	0.0	0.0	1.933	17.49
		26.651			17.65	470.2891					
1.40	25.119		2.37	17.65			470.29	0.04	0.06	1.883	17.62
		22.536			33.88	763.5972					
1.30	19.953		4.05	51.53			1233.89	0.10	0.11	1.783	17.87
		17.901			24.83	444.4058					
1.20	15.849		4.93	76.36			1678.29	0.13	0.13	1.683	18.12
		14.219			16.61	236.2483					
1.10	12.589		5.44	92.97			1914.54	0.15	0.15	1.583	18.37
		11.295			16.76	189.2840					
1.00	10.000		5.91	109.73			2103.82	0.16	0.16	1.483	18.62
		8.972			34.68	311.1721					
0.90	7.943		6.78	144.41			2415.00	0.19	0.18	1.383	18.87
		7.126			27.15	193.5172					
0.80	6.310		7.39	171.57			2608.51	0.20	0.20	1.283	19.12
		5.661			28.99	164.1075					
0.70	5.012		7.99	200.56			2772.62	0.21	0.22	1.183	19.37
		4.496			27.49	123.6068					
0.60	3.981		8.52	228.05			2896.23	0.22	0.23	1.083	19.62
		3.572			107.19	382.8337					
0.50	3.162		10.33	335.24			3279.06	0.25	0.28	0.983	19.87
		2.837			94.82	269.0015					
0.40	2.512		11.70	430.05			3548.06	0.27	0.32	0.883	20.12
		2.254			168.23	379.1226					
0.30	1.995		13.80	598.28			3927.19	0.30	0.37	0.783	20.37
		1.790			202.95	363.2944					
0.20	1.585		15.97	801.23			4290.48	0.33	0.43	0.683	20.62
		1.422			232.54	330.6448					
0.10	1.259		18.14	1033.77			4621.12	0.36	0.49	0.583	20.87
		1.129			243.05	274.5161					
-0.00	1.000		20.16	1276.82			4895.64	0.38	0.55	0.483	21.12
		0.897			352.01	315.8064					
-0.10	0.794		22.77	1628.83			5211.44	0.40	0.62	0.383	21.37
		0.713			446.16	317.9526					
-0.20	0.631		25.70	2074.99			5529.39	0.43	0.70	0.283	21.62
		0.566			368.70	208.7128					
-0.30	0.501		27.89	2443.69			5738.10	0.44	0.75	0.183	21.87
		0.450			594.88	267.4868					
-0.40	0.398		31.10	3038.58			6005.59	0.46	0.84	0.083	22.12
		0.357			1560.17	557.2400					
-0.50	0.316		38.26	4598.75			6562.82	0.51	1.04	-0.017	22.37
		0.284			2453.70	696.1331					
-0.60	0.251		47.38	7052.45			7258.96	0.56	1.28	-0.117	22.62
		0.225			4370.66	984.9597					
-0.70	0.200		60.30	11423.11			8243.91	0.64	1.63	-0.217	22.87
		0.179			3490.68	624.8574					
-0.80	0.158		68.90	14913.79			8868.77	0.69	1.86	-0.317	23.12
		0.142			3423.55	486.7969					
-0.90	0.126		76.40	18337.34			9355.57	0.72	2.07	-0.417	23.37
		0.113			5168.84	583.8013					
-1.00	0.100		86.50	23506.18			9939.37	0.77	2.34	-0.517	23.62
		0.090			5026.03	450.9177					
-1.10	0.079		95.30	28532.21			10390.28	0.80	2.58	-0.617	23.87
		0.071			5906.20	420.9016					
-1.20	0.063		104.70	34438.41			10811.18	0.84	2.83	-0.717	24.12
		0.057			4478.65	253.5246					
-1.30	0.050		111.30	38917.06			11064.70	0.86	3.01	-0.817	24.37
		0.045			5123.53	230.3791					
-1.40	0.040		118.40	44040.59			11295.08	0.87	3.20	-0.917	24.62
		0.036			4498.54	160.6742					
-1.50	0.032		124.30	48539.13			11455.75	0.89	3.36	-1.017	24.87
		0.028			8546.86	242.4833					
-1.60	0.025		134.80	57086.00			11698.23	0.91	3.65	-1.117	25.12
		0.023			7336.26	165.3295					
-1.70	0.020		143.20	64422.25			11863.56	0.92	3.87	-1.217	25.37
		0.018			7589.18	135.8537					
-1.80	0.016		151.40	72011.44			11999.41	0.93	4.10	-1.317	25.62
		0.014			10742.12	152.7450					
-1.90	0.013		162.30	82753.56			12152.16	0.94	4.39	-1.417	25.87
		0.011			8252.19	93.2065					
-2.00	0.010		170.20	91005.75			12245.36	0.95	4.60	-1.517	26.12
							12925.00	(11)			

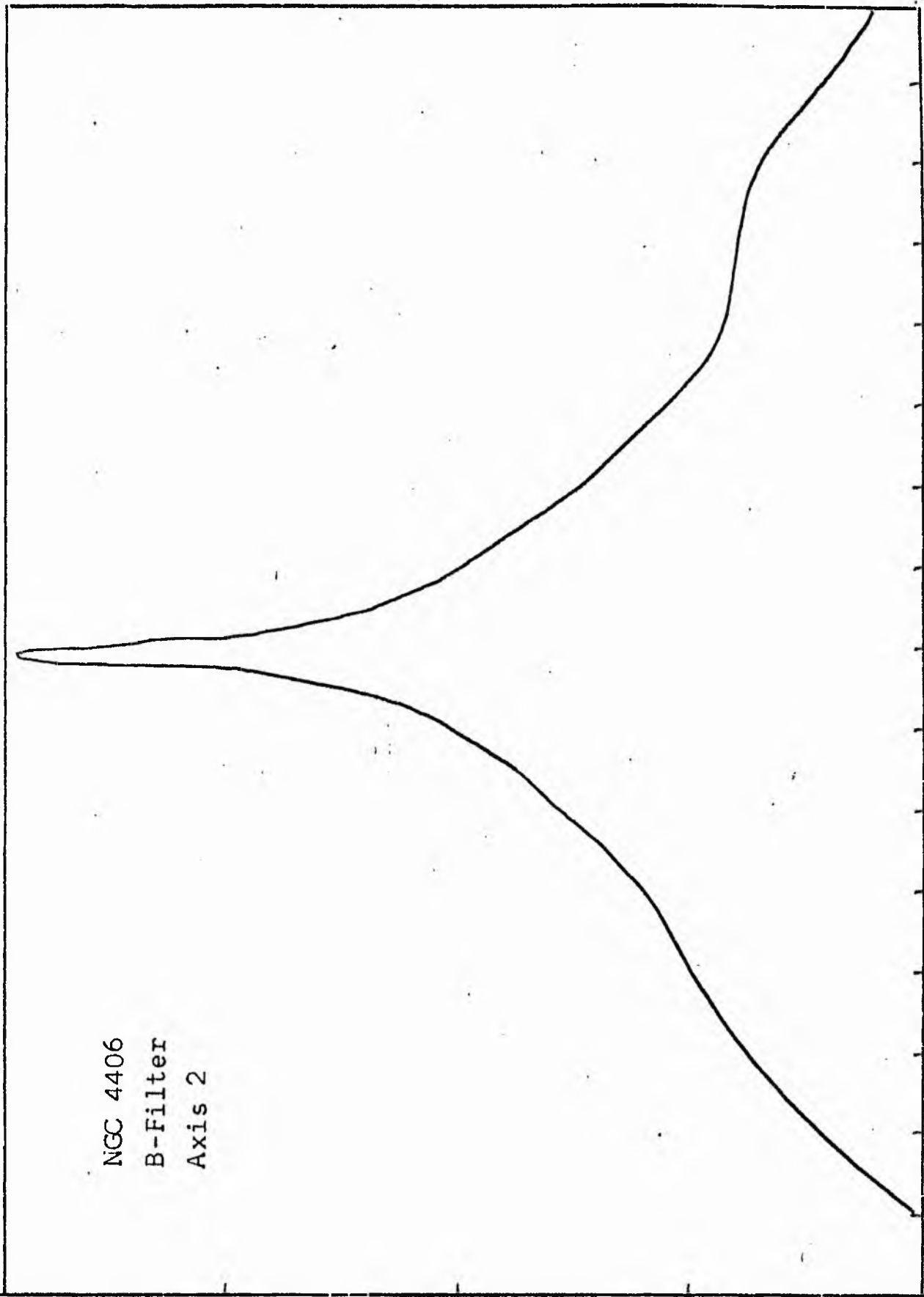
PHOTOMETRIC PARAMETERS OF NGC 4394

V-FILTER

Total luminosity	L_T	= 3.59
Total apparent magnitude	m_T	= 10.84
Apparent central surface brightness	μ_0	= 17.49
Major axis at threshold	$2a_m$	= 6.37
Minor axis at threshold	$2b_m$	= 4.82
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.10
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.90
Gradient of exponential component	$G(a)$	= -0.58
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.68
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.45
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 0.53
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 19.87
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.85
Axis ratio	b/a	= 0.49
Equivalent radius	r_e^*	= 0.62
Surface brightness	μ_e	= 22.32
Mean surface brightness	μ_e'	= 11.79
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.53
Axis ratio	b/a	= 0.78
Equivalent radius	r_3^*	= 1.37
Surface brightness	μ_3	= 23.52
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 3.66 \\ = 2.22 \end{matrix}$

NGC 4406
B-Filter
Axis 1

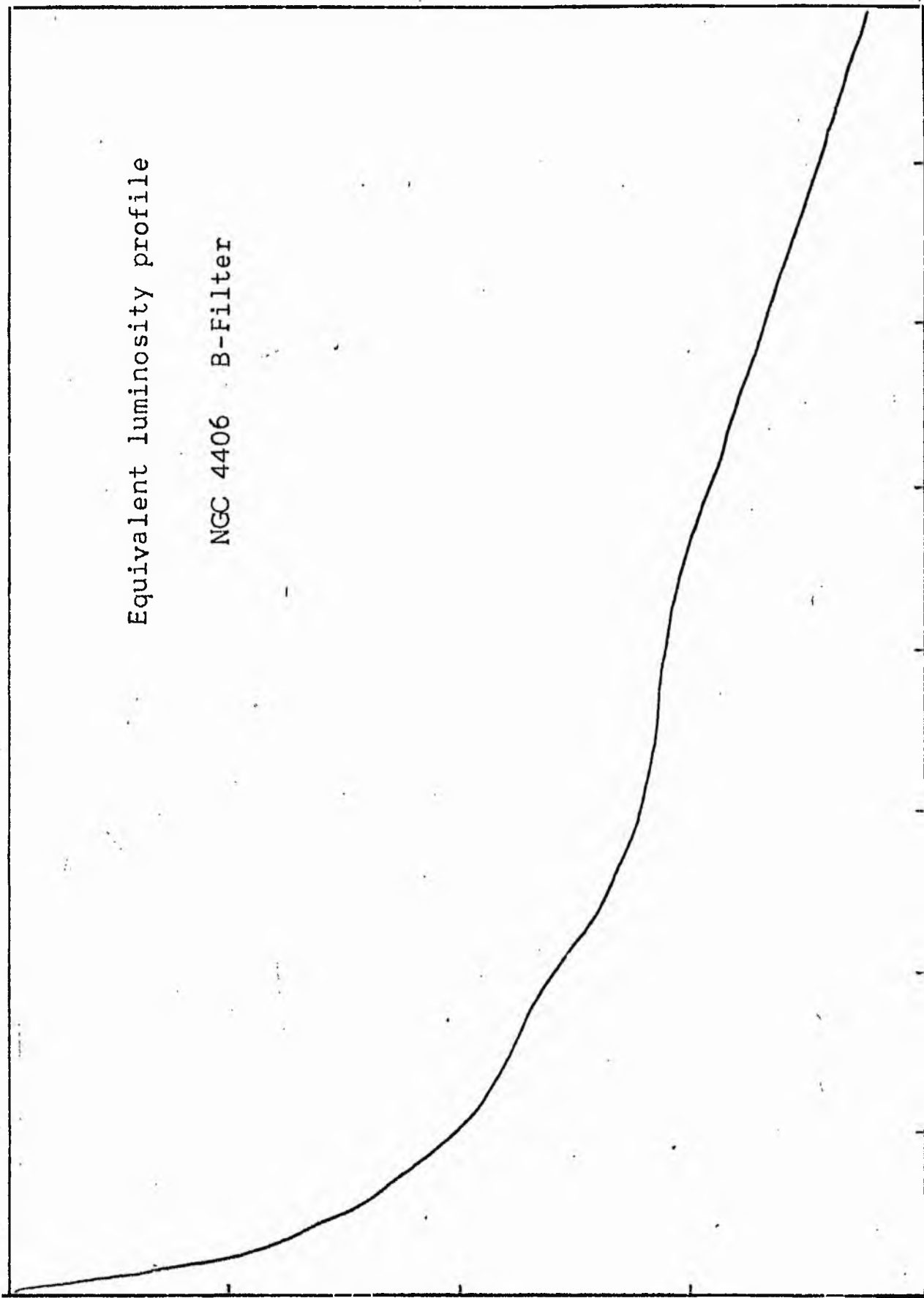


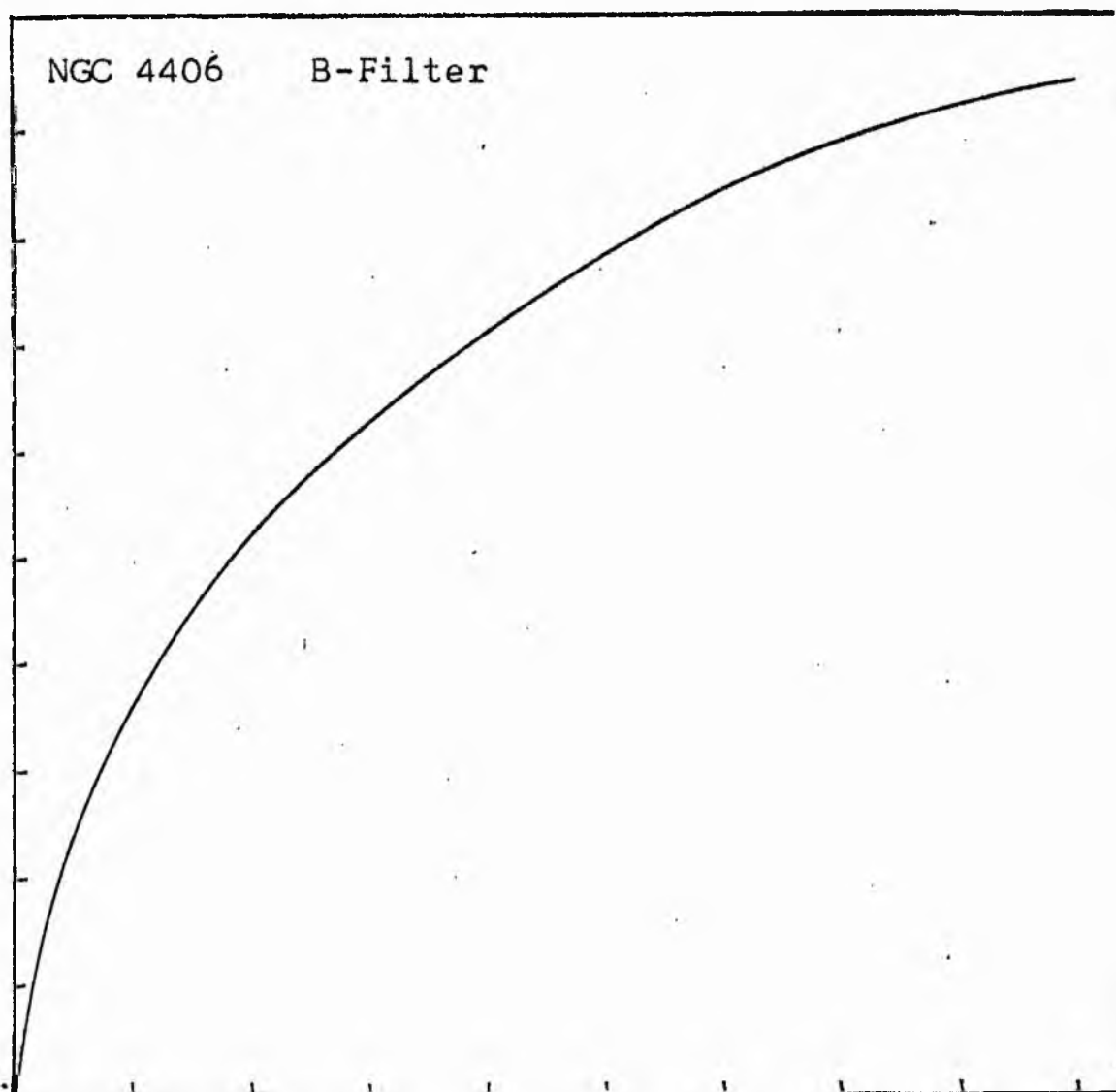


NGC 4406
B-Filter
Axis 2

Equivalent luminosity profile

NGC 4406 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4406
B COLOUR

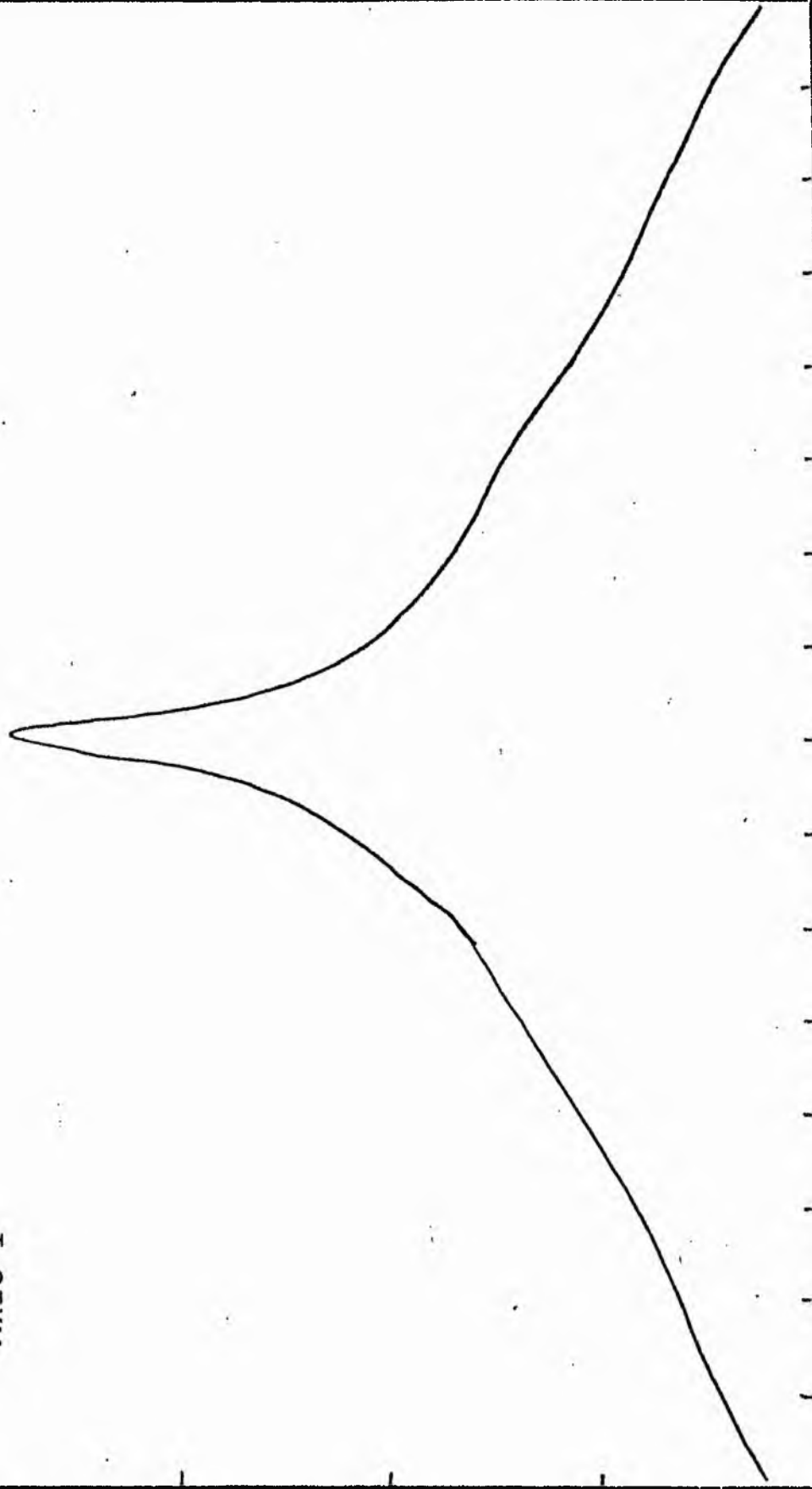
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	P	LOG J	μ
1.93	85.114		0.0	0.0			0.0	0.0	0.0	2.231	17.63
1.90	79.433	82.273	3.10	30.19	30.19	2483.8848	2483.88	0.03	0.03	2.201	17.71
1.80	63.096	71.264	4.50	63.62	33.43	2382.1160	4866.00	0.05	0.05	2.101	17.96
1.70	50.119	56.607	5.70	102.07	38.45	2176.7158	7042.71	0.08	0.06	2.001	18.21
1.60	39.811	44.965	6.20	120.76	18.69	840.4990	7883.21	0.09	0.07	1.901	18.46
1.50	31.623	35.717	6.60	136.85	16.08	574.4995	8457.71	0.09	0.07	1.801	18.71
1.40	25.119	28.371	7.70	186.27	49.42	1402.0012	9859.71	0.11	0.09	1.701	18.96
1.30	19.952	22.536	9.40	277.59	91.33	2058.0872	11917.79	0.13	0.10	1.601	19.21
1.20	15.849	17.901	11.70	430.05	152.46	2729.1584	14646.95	0.16	0.13	1.501	19.46
1.10	12.589	14.219	11.90	444.88	14.83	210.8443	14857.79	0.16	0.13	1.401	19.71
1.00	10.000	11.295	13.00	598.28	153.40	1732.6230	16590.41	0.18	0.15	1.301	19.96
0.90	7.943	8.972	15.70	774.37	176.04	1574.7708	18170.18	0.20	0.17	1.201	20.21
0.80	6.310	7.126	16.30	834.69	60.32	429.8403	186.0.02	0.20	0.18	1.101	20.46
0.70	5.012	5.661	18.40	1063.62	228.93	1275.8884	19845.91	0.21	0.20	1.001	20.71
0.60	3.981	4.496	22.10	1534.38	470.77	2116.7747	22012.68	0.24	0.25	0.901	20.96
0.50	3.162	3.572	26.00	2123.72	509.33	2104.8901	24117.57	0.26	0.29	0.801	21.21
0.40	2.512	2.837	29.70	2771.17	647.45	1836.8535	25954.42	0.28	0.33	0.701	21.46
0.30	1.995	2.254	34.30	3696.05	924.88	2084.2781	28038.70	0.30	0.38	0.601	21.71
0.20	1.585	1.790	39.50	4901.67	1205.62	2158.1350	30146.83	0.33	0.44	0.501	21.96
0.10	1.259	1.422	46.30	6734.59	1832.93	2606.2349	32803.07	0.35	0.52	0.401	22.21
-0.00	1.000	1.129	53.40	8958.44	2223.84	2511.7305	35314.89	0.38	0.59	0.301	22.46
-0.10	0.794	0.897	61.90	12037.36	3078.92	2762.2747	38077.07	0.41	0.69	0.201	22.71
-0.20	0.631	0.713	73.80	17110.49	5073.13	3615.3066	41692.37	0.45	0.82	0.101	22.96
-0.30	0.501	0.566	89.60	25221.20	8110.71	4591.2148	46283.59	0.50	1.00	0.001	23.21
-0.40	0.398	0.450	103.60	33718.58	8497.38	3820.7983	50104.39	0.54	1.15	-0.099	23.46
-0.50	0.316	0.357	110.20	38151.62	4433.04	1583.3284	51687.71	0.56	1.23	-0.199	23.71
-0.60	0.251	0.284	121.40	46300.66	8149.04	2311.9336	53999.65	0.58	1.35	-0.299	23.96
-0.70	0.200	0.225	135.60	57765.59	11464.93	2583.6902	56583.34	0.61	1.51	-0.399	24.21
-0.80	0.158	0.179	153.60	74119.44	16353.85	2927.4487	59510.78	0.64	1.71	-0.499	24.46
-0.90	0.126	0.142	189.90	113292.12	39172.69	5569.9648	65040.75	0.70	2.12	-0.599	24.71
-1.00	0.100	0.113	236.70	176013.62	62721.50	7084.1211	72164.81	0.78	2.64	-0.699	24.96
-1.10	0.079	0.090	250.30	196821.00	20807.37	1866.7559	74031.56	0.80	2.79	-0.799	25.21
-1.20	0.063	0.071	273.50	234998.19	38177.19	2720.6624	76752.19	0.83	3.05	-0.899	25.46
-1.30	0.050	0.057	298.60	280110.06	45111.87	2553.6548	79305.81	0.86	3.33	-0.999	25.71
-1.40	0.040	0.045	320.70	323107.87	42997.81	1933.3850	81239.19	0.88	3.57	-1.099	25.96
-1.50	0.032	0.036	344.60	373061.00	49953.12	1784.1658	83023.31	0.90	3.84	-1.199	26.21
-1.60	0.025	0.028	371.80	434278.25	61217.25	1736.7883	84760.06	0.91	4.14	-1.299	26.46
-1.70	0.020	0.023	386.90	470269.75	35991.50	811.0984	85571.12	0.92	4.31	-1.399	26.71
-1.80	0.016	0.018	405.20	515808.69	45538.94	815.1870	86386.25	0.93	4.51	-1.499	26.96
-1.90	0.013	0.014	419.70	553385.31	37576.62	534.3093	86920.50	0.94	4.67	-1.599	27.21
-2.00	0.010	0.011	434.60	593374.62	39989.31	451.6680	87372.12	0.94	4.84	-1.699	27.46
-∞							92672.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4406

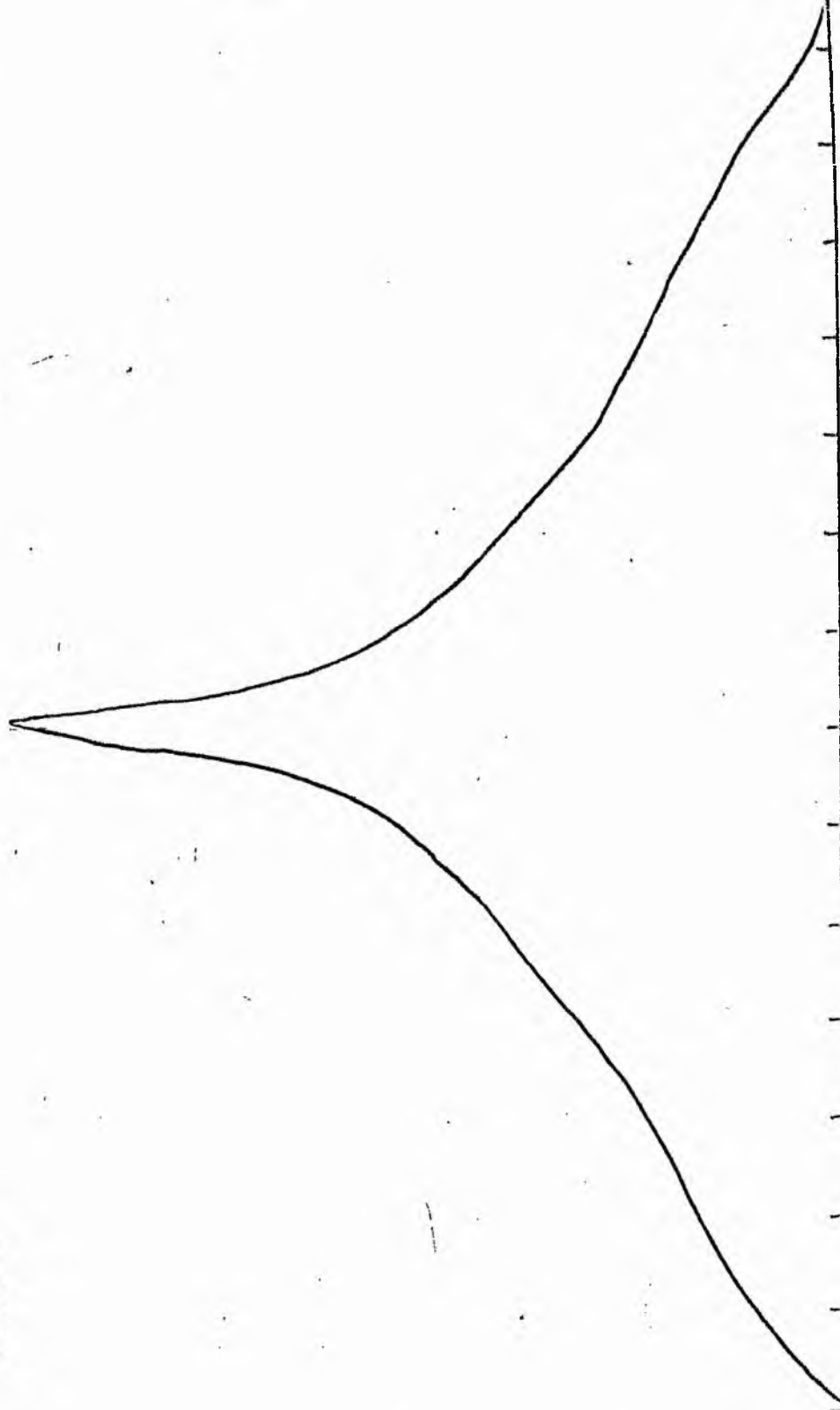
B-FILTER

Total luminosity	L_T	= 25.74
Total apparent magnitude	m_T	= 10.04
Apparent central surface brightness	μ_0	= 17.63
Major axis at threshold	$2a_m$	= 16.03
Minor axis at threshold	$2b_m$	= 12.42
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 8.10
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.78
Gradient of exponential component	$G(a)$	= -0.21
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.26
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.44
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.45
Axis ratio	b/a	= 0.81
Equivalent radius	r_1^*	= 0.40
Surface brightness	μ_1	= 21.08
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.32
Axis ratio	b/a	= 1.04
Equivalent radius	r_e^*	= 1.50
Surface brightness	μ_e	= 23.21
Mean surface brightness	μ_e'	= 12.92
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 3.86
Axis ratio	b/a	= 0.75
Equivalent radius	r_3^*	= 3.64
Surface brightness	μ_3	= 24.87
Concentration indices	$\{C_{21}$	= 3.71
	C_{32}	= 2.43

NGC 4406
V-Filter
Axis 1

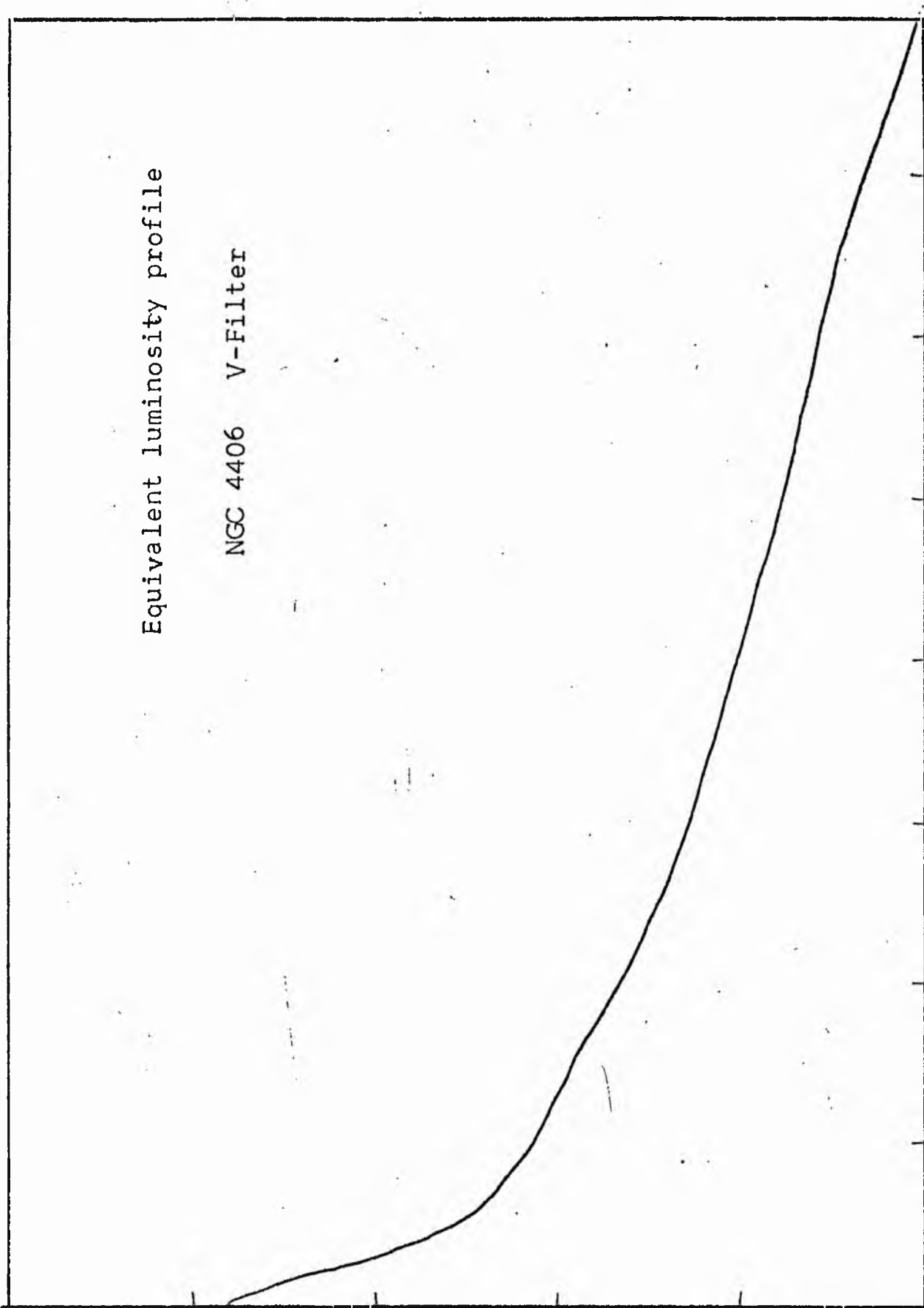


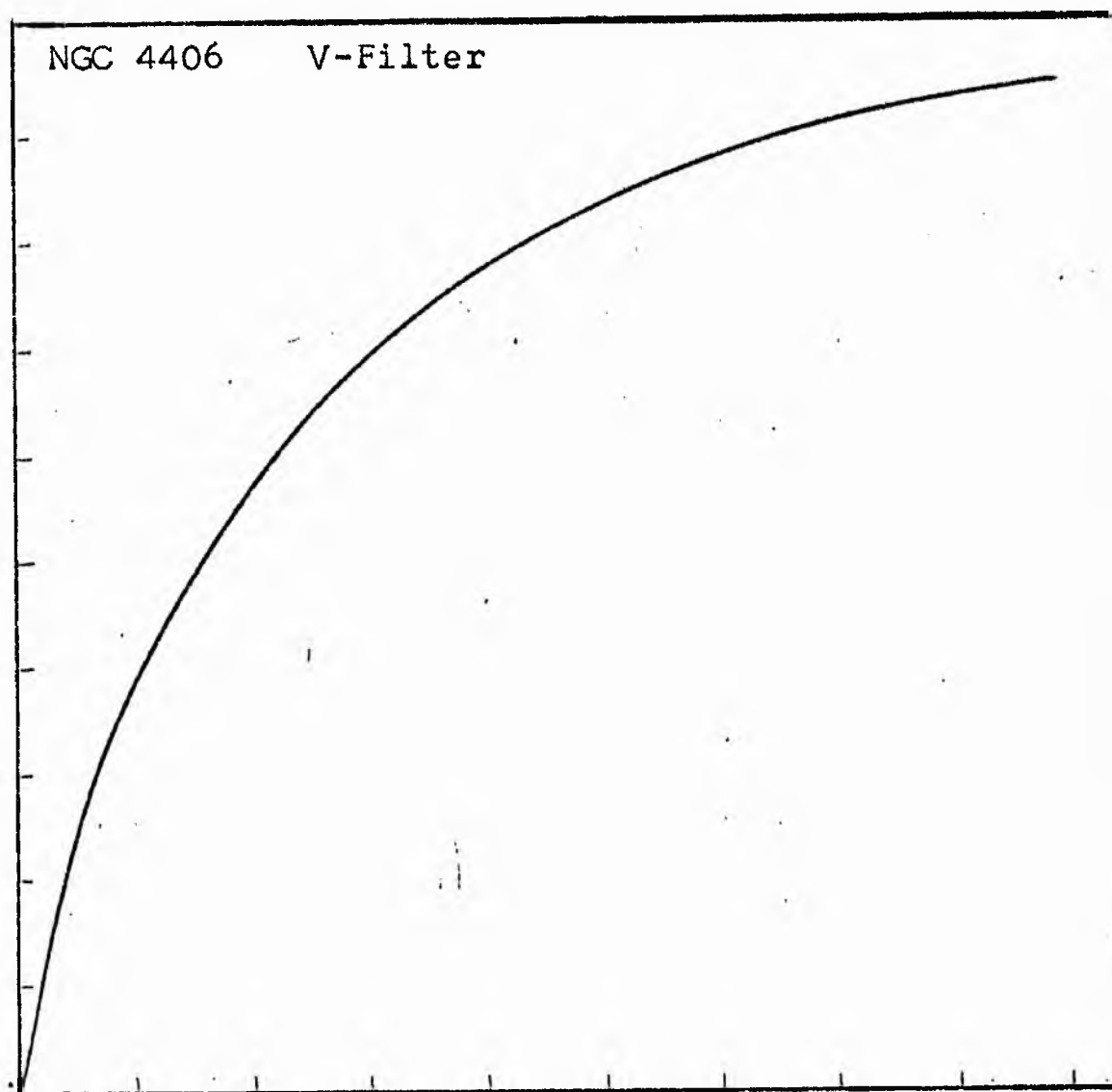
NGC 4406
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4406 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4406
V COLOUR

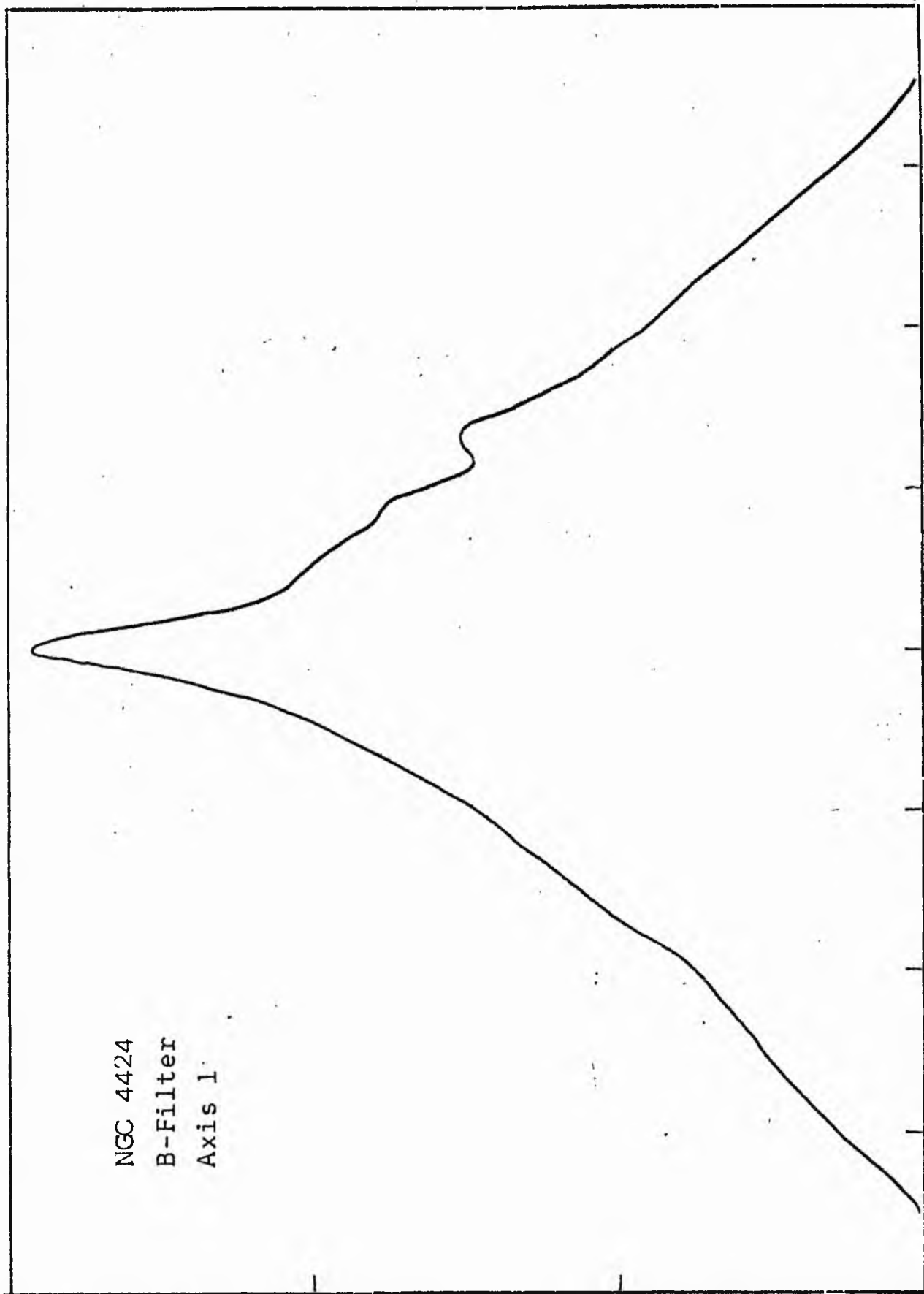
LOG I	I	T	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
1.84	69.183		0.0	0.0			0.0	0.0	0.0	1.942	17.03
		66.139			8.04	531.9231					
1.80	63.096	56.607	1.60	8.04	22.15	1253.7454	531.92	0.01	0.02	1.902	17.13
1.70	50.119	44.965	3.10	30.17	25.23	1134.3193	1785.67	0.02	0.04	1.802	17.38
1.60	39.811	35.717	4.20	55.42	46.65	1666.2744	2919.99	0.03	0.06	1.702	17.63
1.50	31.623	28.371	5.70	102.07	135.72	3850.3762	4586.26	0.05	0.08	1.602	17.88
1.40	25.119	22.536	8.70	237.79	237.50	5352.3008	8436.64	0.09	0.12	1.502	18.13
1.30	19.952	17.901	12.30	475.29	72.10	1290.6296	13788.94	0.15	0.17	1.402	18.38
1.20	15.849	14.219	13.20	547.39	150.07	2133.8950	15079.57	0.16	0.18	1.302	18.63
1.10	12.589	11.295	14.90	697.46	41.60	537.5652	17213.46	0.19	0.20	1.202	18.88
1.00	10.000	8.972	15.40	745.06	131.10	1176.1577	17751.02	0.19	0.21	1.102	19.13
0.90	7.943	7.126	16.70	876.16	294.05	2095.5210	18927.18	0.21	0.23	1.002	19.38
0.80	6.310	5.661	19.30	1170.21	295.53	1672.8997	21022.70	0.23	0.26	0.902	19.63
0.70	5.012	4.496	21.60	1465.74	374.10	1682.1211	22695.60	0.25	0.29	0.802	19.88
0.60	3.981	3.572	24.20	1839.84	588.10	2100.5012	24377.72	0.26	0.33	0.702	20.13
0.50	3.162	2.837	27.80	2427.95	689.30	1955.5854	26478.22	0.29	0.38	0.602	20.38
0.40	2.512	2.254	31.50	3117.25	1137.22	2562.7896	28433.80	0.31	0.43	0.502	20.63
0.30	1.995	1.790	36.80	4254.46	1313.72	2351.6487	30996.59	0.34	0.50	0.402	20.88
0.20	1.585	1.422	42.10	5568.19	3156.93	4488.8281	33348.24	0.36	0.57	0.302	21.13
0.10	1.259	1.129	52.70	8725.11	2850.05	3219.0029	37837.07	0.41	0.71	0.202	21.38
-0.00	1.000	0.897	60.70	11575.16	5442.71	4882.9648	41056.07	0.45	0.82	0.102	21.63
-0.10	0.794	0.713	73.60	17017.87	8655.71	6168.3867	45939.03	0.50	1.00	0.002	21.88
-0.20	0.631	0.566	90.40	25673.59	6247.00	3536.2256	52107.42	0.57	1.22	-0.098	22.13
-0.30	0.501	0.450	100.80	31920.58	6369.64	2864.0715	55643.64	0.60	1.36	-0.198	22.38
-0.40	0.398	0.357	110.40	38290.22	4494.77	1605.3762	58507.71	0.64	1.49	-0.298	22.63
-0.50	0.316	0.284	116.70	42785.00	8526.12	2418.9163	60113.09	0.65	1.58	-0.398	22.88
-0.60	0.251	0.225	127.80	51311.12	9300.28	2095.8745	62532.00	0.68	1.73	-0.498	23.13
-0.70	0.200	0.179	138.90	60611.40	18212.85	3260.2214	64627.87	0.70	1.88	-0.598	23.38
-0.80	0.158	0.142	158.40	78824.25	27075.50	3849.8662	67888.06	0.74	2.14	-0.698	23.63
-0.90	0.126	0.113	183.60	105899.75	20771.25	2346.0227	71737.87	0.78	2.49	-0.798	23.88
-1.00	0.100	0.090	200.80	126671.00	34784.50	3120.7288	74083.87	0.80	2.72	-0.898	24.13
-1.10	0.079	0.071	226.70	161455.50	31453.44	2241.5005	77204.56	0.84	3.07	-0.998	24.38
-1.20	0.063	0.057	247.80	192908.94	37471.87	2121.1763	79446.06	0.86	3.35	-1.098	24.63
-1.30	0.050	0.045	270.80	230380.81	29291.44	1317.0815	81567.19	0.89	3.67	-1.198	24.88
-1.40	0.040	0.036	287.50	259672.25	33917.25	1211.4155	82884.25	0.90	3.89	-1.298	25.13
-1.50	0.032	0.028	305.70	293589.50	43777.50	1242.0068	84095.62	0.91	4.14	-1.398	25.38
-1.60	0.025	0.023	327.70	337367.00	40690.69	916.9985	85337.62	0.93	4.44	-1.498	25.63
-1.70	0.020	0.018	346.90	378057.69	46926.56	840.0266	86254.56	0.94	4.70	-1.598	25.88
-1.80	0.016	0.014	367.80	424984.25	13040.00	185.4183	87094.56	0.95	4.98	-1.698	26.13
-1.90	0.013	0.011	373.40	438024.25	1691.69	809.7375	87279.94	0.95	5.05	-1.798	26.38
-2.00	0.010		402.80	509715.94			88789.62	0.96	5.45	-1.898	26.63
-∞							92089.00	(11)			∞

PHOTOMETRIC PARAMETERS OF NGC 4406

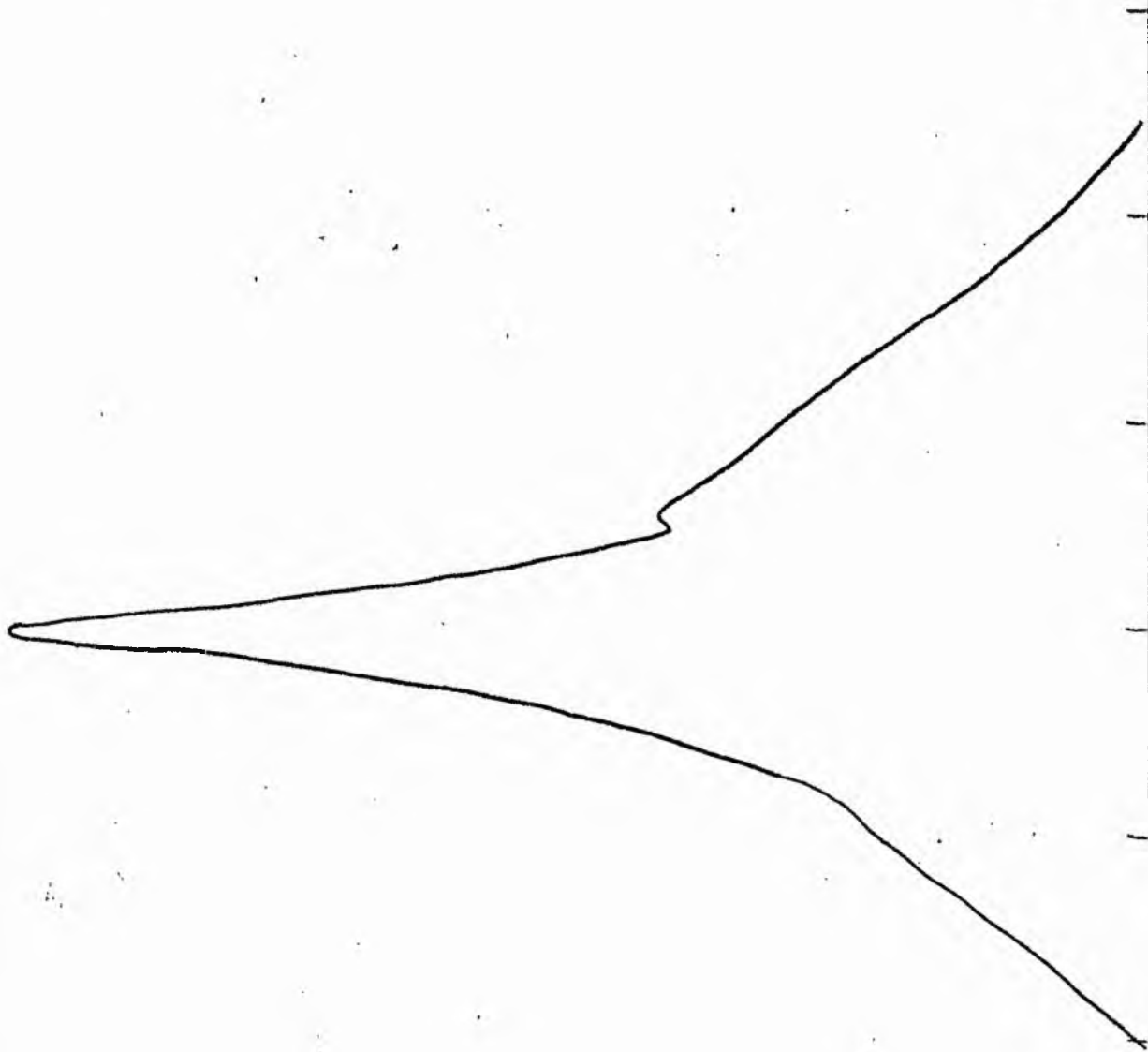
V-FILTER

Total luminosity	L_T	= 25.58
Total apparent magnitude	m_T	= 9.21
Apparent central surface brightness	μ_o	= 17.03
Major axis at threshold	$2a_m$	= 14.58
Minor axis at threshold	$2b_m$	= 11.70
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 9.81
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.28
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.31
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.34
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.53
Axis ratio	b/a	= 0.66
Equivalent radius	r_1^*	= 0.37
Surface brightness	μ_1	= 19.88
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.37
Axis ratio	b/a	= 0.69
Equivalent radius	r_e^*	= 1.23
Surface brightness	μ_e	= 21.88
Mean surface brightness	μ_e'	= 11.65
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 3.08
Axis ratio	b/a	= 0.76
Equivalent radius	r_3^*	= 2.76
Surface brightness	μ_3	= 23.69
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.35 \\ = 2.24 \end{cases}$

NGC 4424
B-Filter
Axis 1

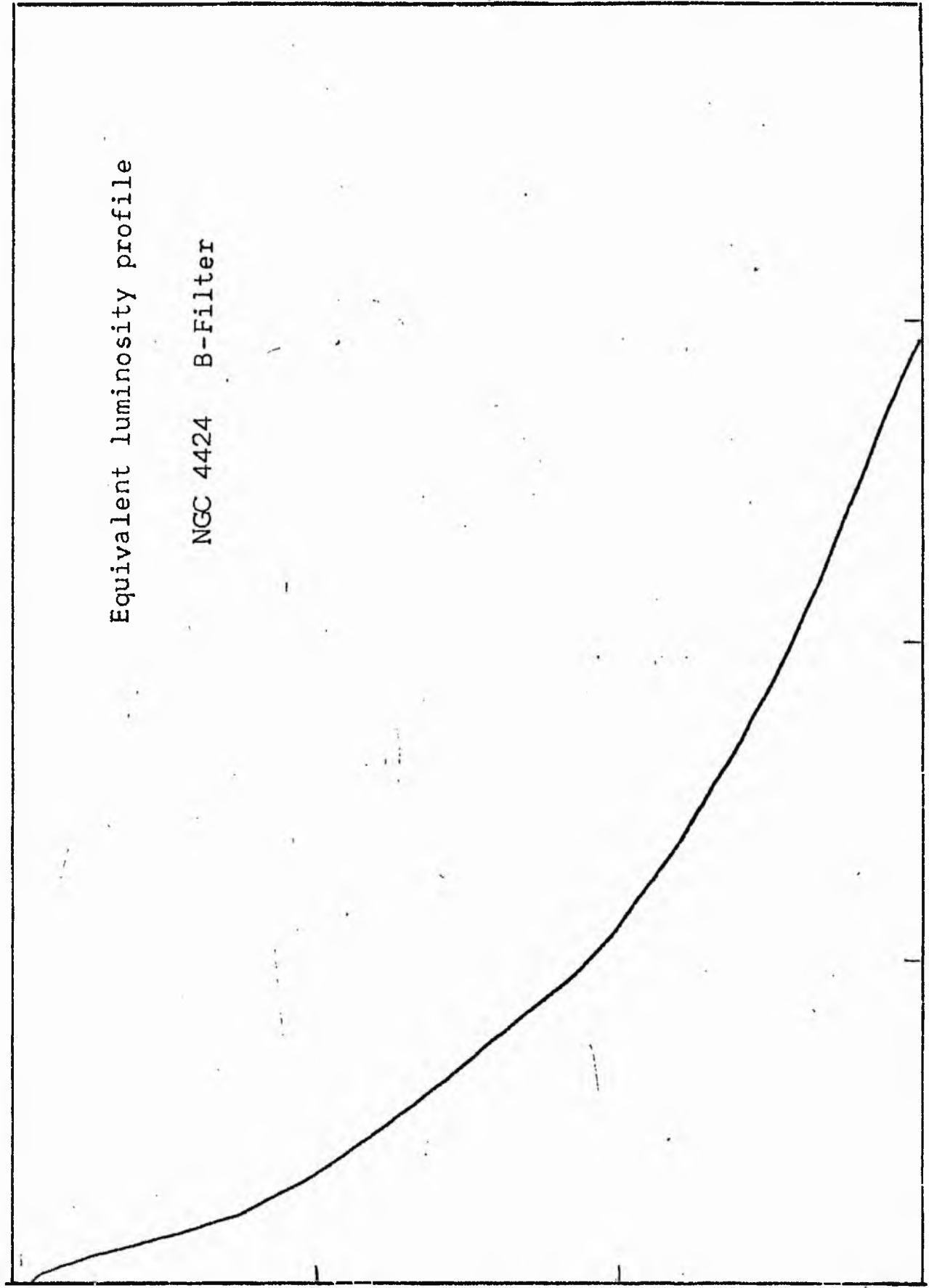


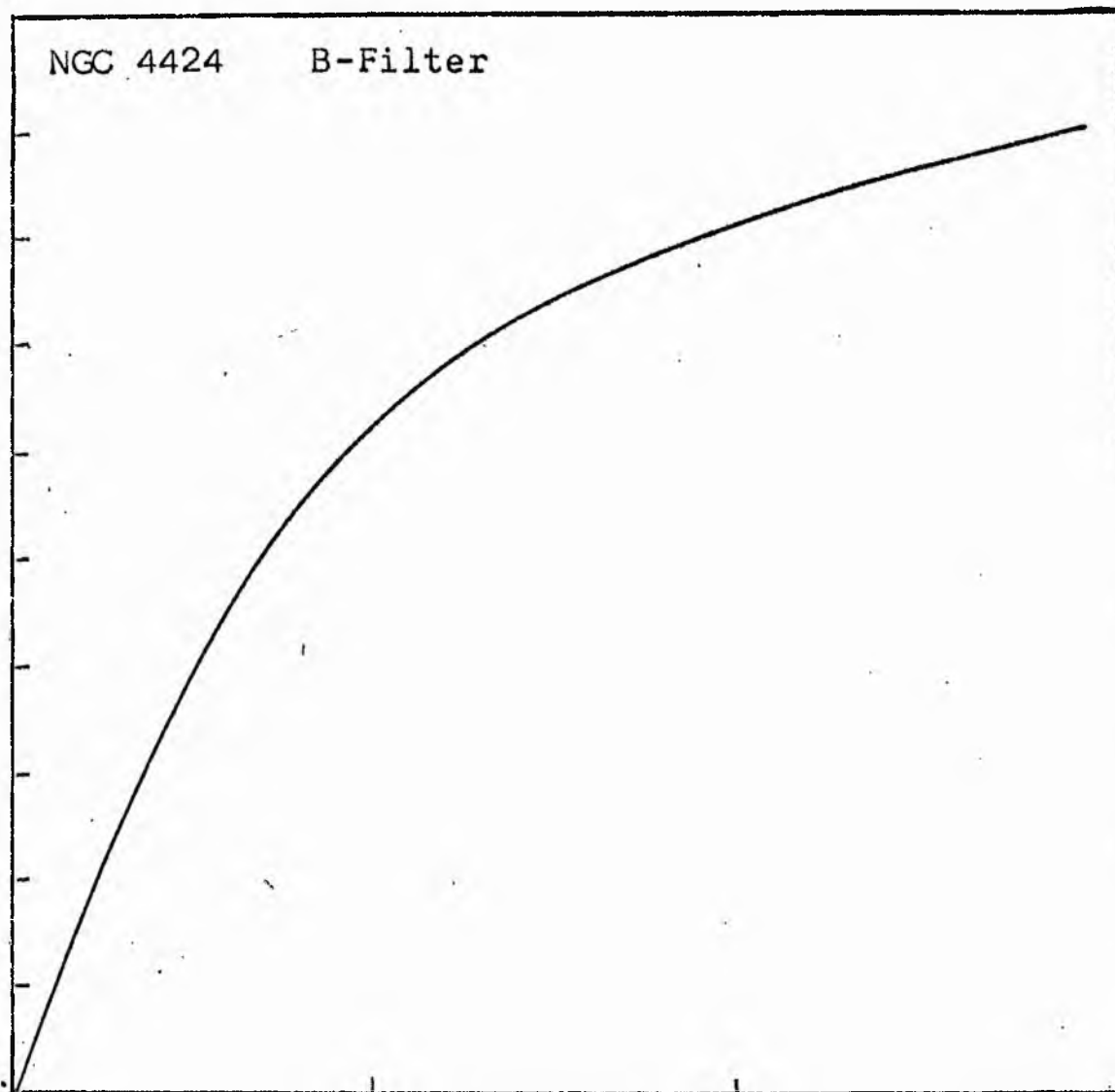
NGC 4424
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4424 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4424
B COLOUR

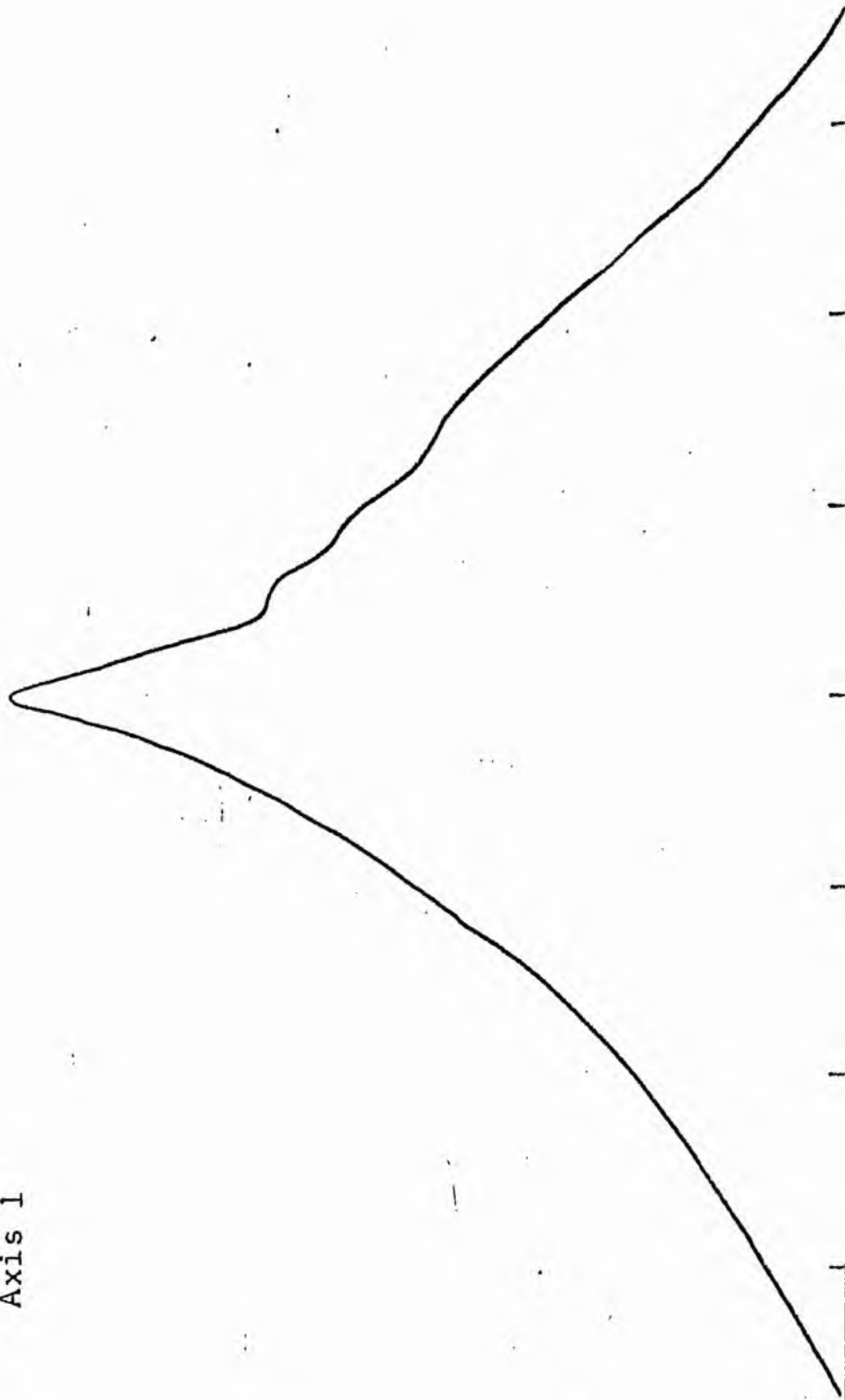
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.88	7.588	6.948	0.0	0.0	49.51	344.0090	0.0	0.0	0.0	1.352	19.65
0.80	6.310	5.661	3.97	49.51	21.67	122.6486	344.01	0.05	0.12	1.272	19.65
0.70	5.012	4.496	4.76	71.18	31.61	142.1195	466.66	0.07	0.14	1.172	20.10
0.60	3.981	3.572	5.72	102.79	42.05	150.1976	608.78	0.09	0.17	1.072	20.35
0.50	3.162	2.837	6.79	144.84	61.28	173.8551	758.97	0.11	0.20	0.972	20.60
0.40	2.512	2.254	8.10	206.12	106.78	240.6456	932.83	0.13	0.24	0.872	20.65
0.30	1.995	1.790	9.98	312.90	117.15	209.7053	1173.47	0.17	0.30	0.772	21.10
0.20	1.585	1.422	11.70	430.05	218.68	310.9382	1383.18	0.20	0.35	0.672	21.35
0.10	1.259	1.129	14.37	648.73	281.76	318.2366	1694.12	0.24	0.43	0.572	21.60
0.00	1.000	0.897	17.21	930.49	352.68	316.4075	2012.35	0.29	0.51	0.472	21.85
-0.10	0.794	0.713	20.21	1263.16	400.49	285.4031	2328.76	0.33	0.60	0.372	22.10
-0.20	0.631	0.566	23.15	1683.65	552.61	312.8145	2614.17	0.38	0.69	0.272	22.35
-0.30	0.501	0.450	26.68	2236.25	699.64	314.5918	2926.98	0.42	0.79	0.172	22.60
-0.40	0.398	0.357	30.57	2935.90	868.70	310.2705	3241.57	0.47	0.91	0.072	22.85
-0.50	0.316	0.284	34.80	3804.59	684.24	194.1237	3551.84	0.51	1.03	-0.028	23.10
-0.60	0.251	0.225	37.80	4488.83	843.84	190.1650	3745.97	0.54	1.12	-0.128	23.35
-0.70	0.200	0.179	41.20	5332.66	944.52	169.0766	3936.13	0.57	1.22	-0.228	23.60
-0.80	0.158	0.142	44.70	6277.18	1639.75	233.1584	4105.20	0.59	1.32	-0.328	23.85
-0.90	0.126	0.113	50.20	7916.94	2578.61	291.2451	4338.36	0.62	1.49	-0.428	24.10
-1.00	0.100	0.090	57.80	10495.55	2052.72	184.1633	4629.60	0.67	1.71	-0.528	24.35
-1.10	0.079	0.071	63.20	12548.27	3199.39	228.0032	4813.76	0.69	1.87	-0.628	24.60
-1.20	0.063	0.057	70.80	15747.66	2637.72	149.3150	5041.76	0.72	2.10	-0.728	24.85
-1.30	0.050	0.045	76.50	18385.39	4152.67	186.7248	5191.07	0.75	2.27	-0.828	25.10
-1.40	0.040	0.036	84.70	22538.05	4926.54	175.9612	5377.80	0.77	2.51	-0.928	25.35
-1.50	0.032	0.028	93.50	27464.59	4709.84	133.6232	5553.76	0.80	2.77	-1.028	25.60
-1.60	0.025	0.023	101.20	32174.43	9156.64	206.3539	5687.38	0.82	3.00	-1.128	25.85
-1.70	0.020	0.018	114.70	41331.07	7442.65	133.2309	5893.73	0.85	3.40	-1.228	26.10
-1.80	0.016	0.014	124.60	48773.72	9162.38	130.2825	6026.96	0.87	3.69	-1.328	26.35
-1.90	0.013	0.011	135.80	57936.10	9213.65	104.0663	6157.24	0.88	4.02	-1.428	26.60
-2.00	0.010		146.20	67149.75			6261.30	0.90	4.33	-1.528	26.85
-∞							6961.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4424

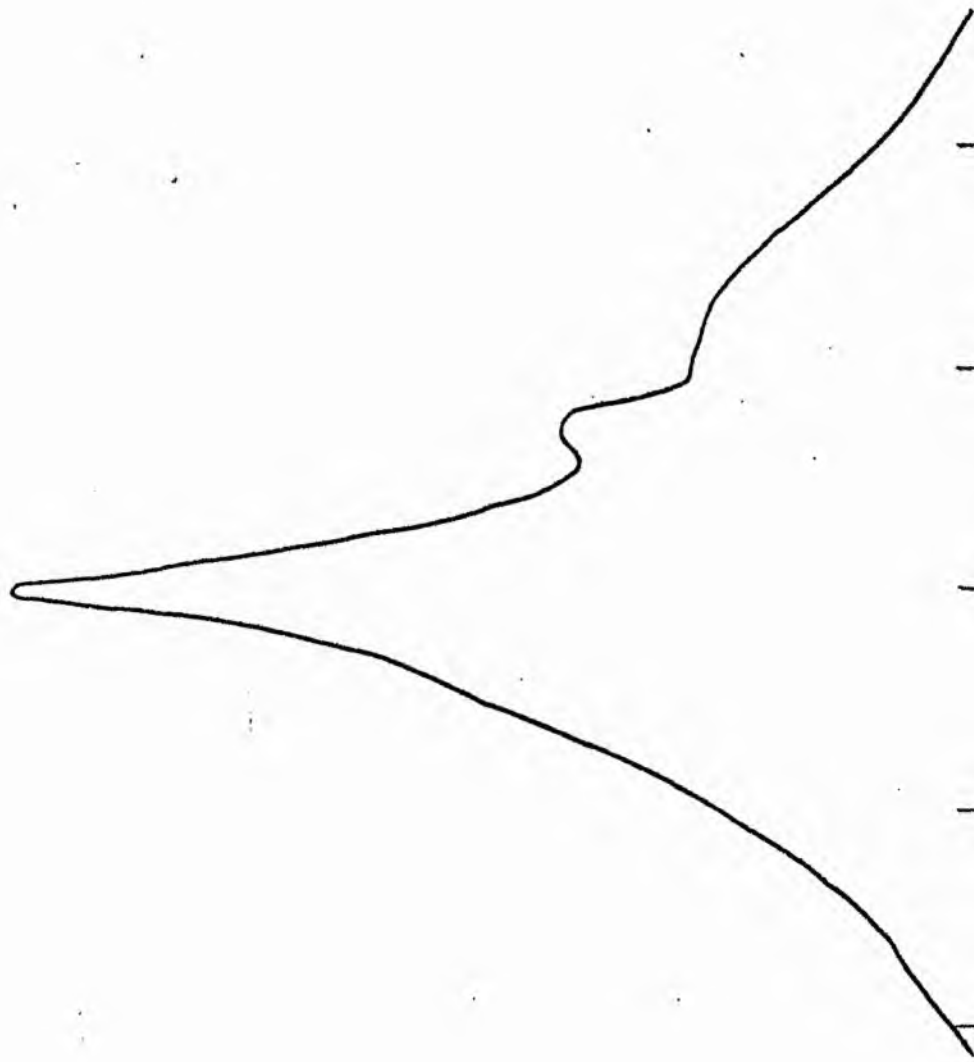
B-FILTER

Total luminosity	L_T	= 1.93
Total apparent magnitude	m_T	= 12.24
Apparent central surface brightness	μ_o	= 19.65
Major axis at threshold	$2a_m$	= 5.94
Minor axis at threshold	$2b_m$	= 3.85
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.68
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.74
Gradient of exponential component	$G(a)$	= -0.89
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.67
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.96
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.38
Axis ratio	b/a	= 0.48
Equivalent radius	r_1^*	= 0.25
Surface brightness	μ_1	= 21.65
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.11
Axis ratio	b/a	= 0.32
Equivalent radius	r_e^*	= 0.56
Surface brightness	μ_e	= 23.04
Mean surface brightness	μ_e'	= 12.94
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.87
Axis ratio	b/a	= 0.51
Equivalent radius	r_3^*	= 1.30
Surface brightness	μ_3	= 25.10
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.29 \\ = 2.30 \end{cases}$

NGC 4424
V-Filter
Axis 1

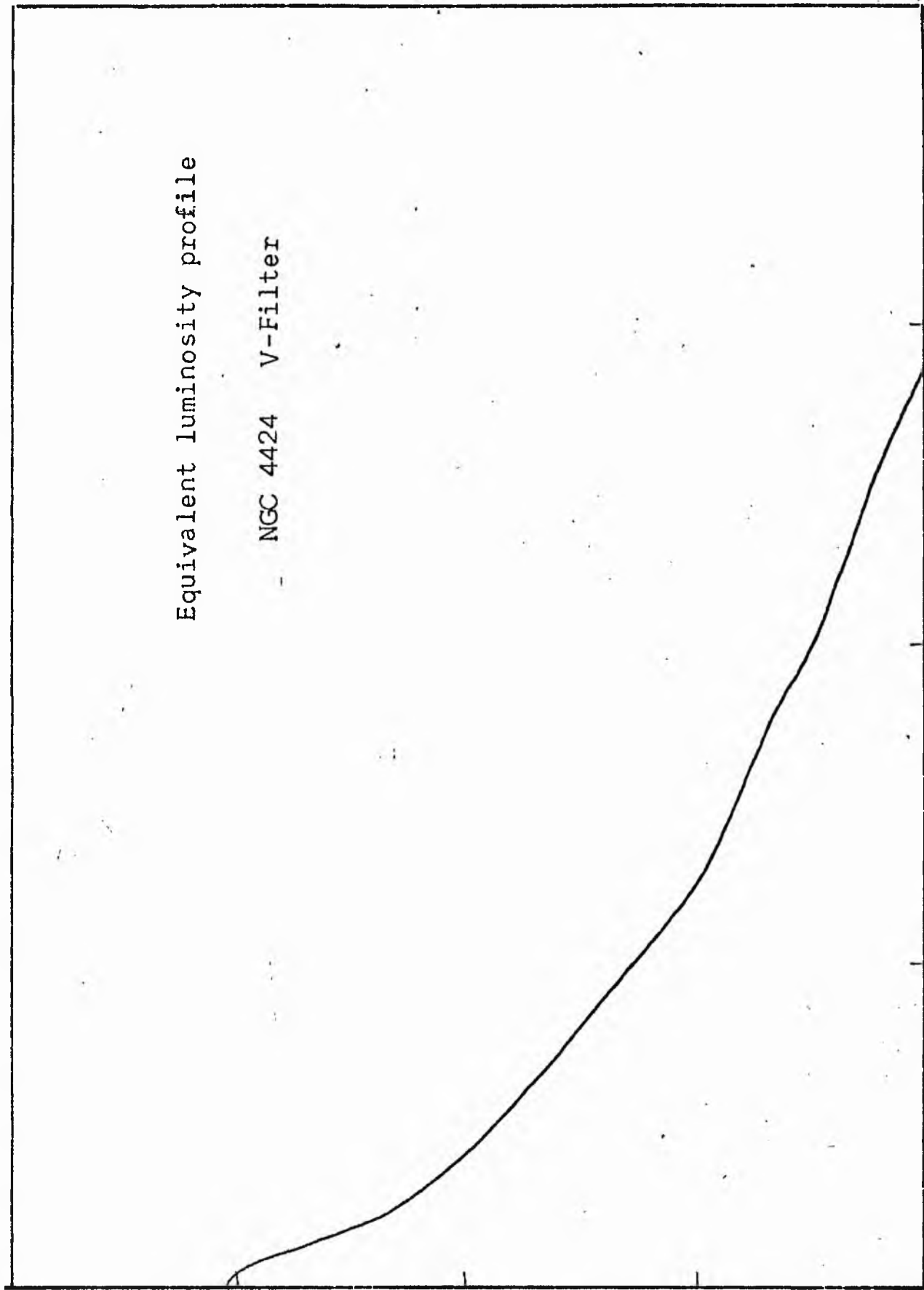


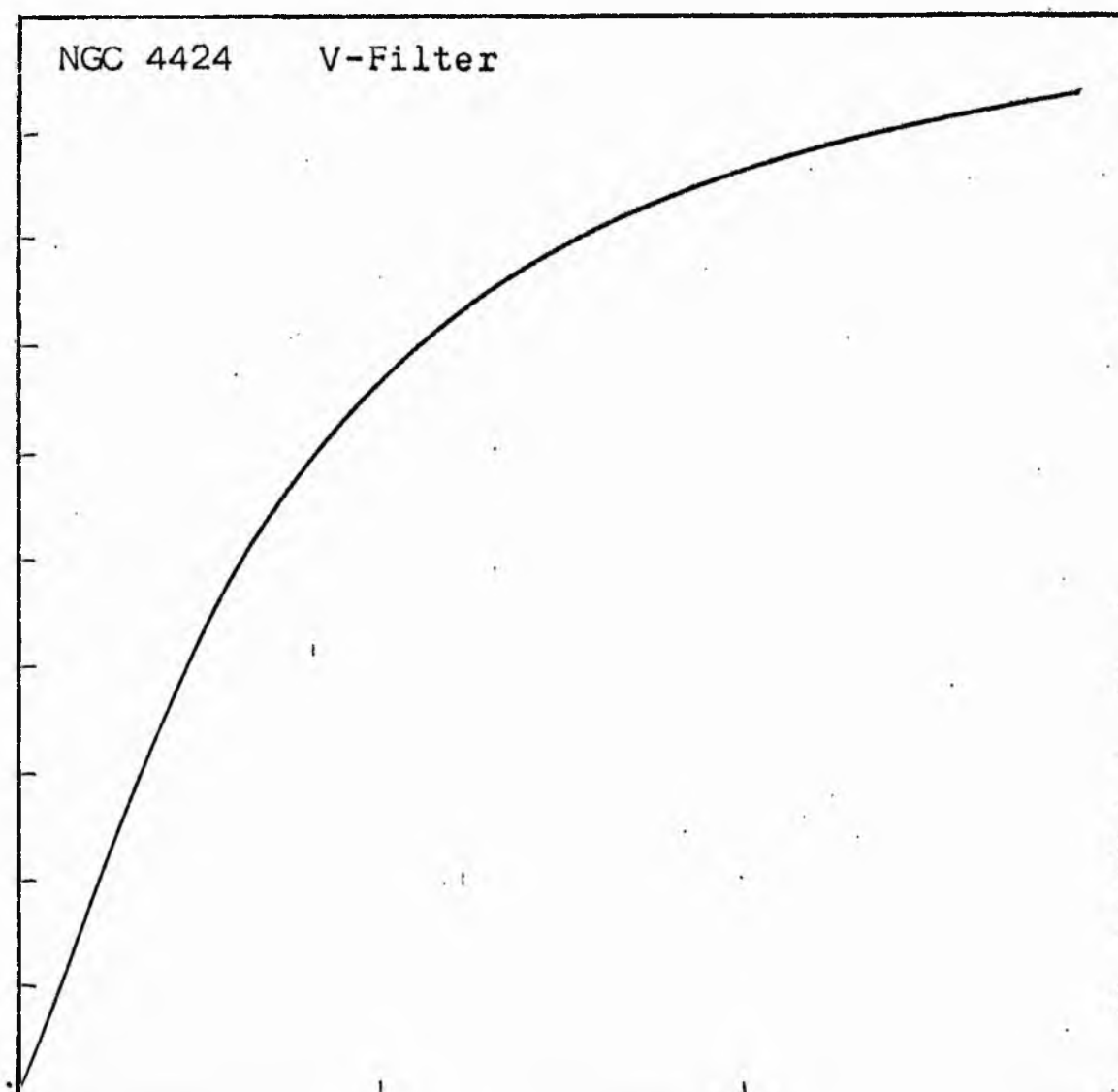
NGC 4424
V-Filter
Axis 2



Equivalent luminosity profile

- NGC 4424 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4424
V COLOUR

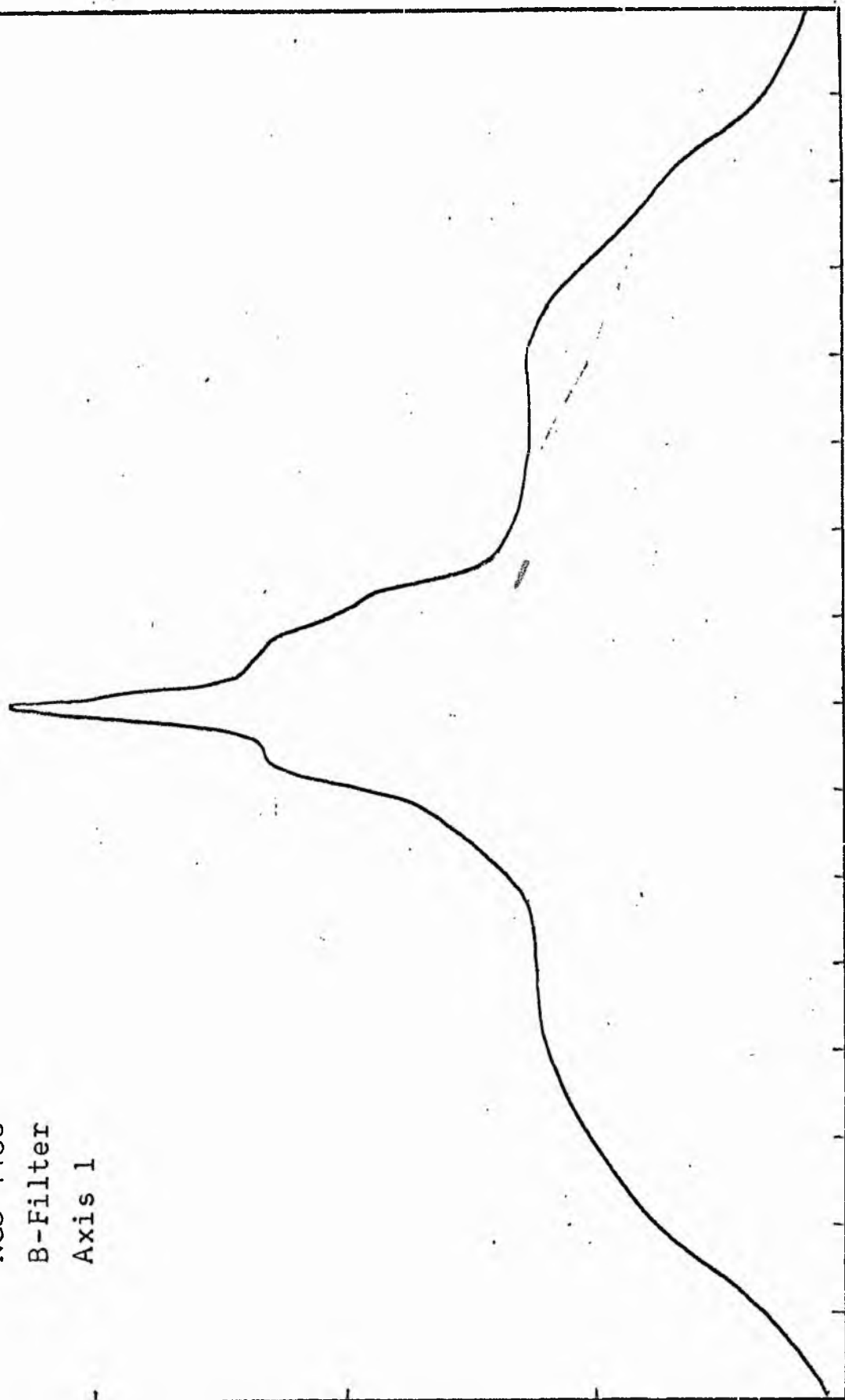
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.03	10.715		0.0	0.0			0.0	0.0	0.0	1.324	18.91
		10.358			28.46	294.8091					
1.00	10.000	8.972	3.01	28.46			294.81	0.03	0.10	1.294	18.99
					34.31	307.8047					
0.90	7.943	7.126	4.47	62.77			602.61	0.07	0.14	1.194	19.24
					20.87	148.7640					
0.80	6.310	5.661	5.16	83.65			751.38	0.08	0.17	1.094	19.49
					54.86	310.5740					
0.70	5.012	4.496	6.64	138.51			1061.95	0.12	0.21	0.994	19.74
					59.05	265.5039					
0.60	3.981	3.572	7.93	197.56			1327.46	0.15	0.25	0.894	19.99
					72.41	258.6152					
0.50	3.162	2.837	9.27	269.97			1586.07	0.18	0.30	0.794	20.24
					69.83	198.1089					
0.40	2.512	2.254	10.40	339.79			1784.18	0.20	0.33	0.694	20.49
					130.10	293.1958					
0.30	1.995	1.790	12.23	469.90			2077.38	0.23	0.39	0.594	20.74
					237.90	425.8667					
0.20	1.585	1.422	15.01	707.80			2503.24	0.28	0.48	0.494	20.99
					292.06	415.2817					
0.10	1.259	1.129	17.84	999.86			2918.52	0.32	0.57	0.394	21.24
					359.32	405.8347					
0.00	1.000	0.897	20.80	1359.18			3324.36	0.37	0.67	0.294	21.49
					375.77	337.1255					
-0.10	0.794	0.713	23.50	1734.94			3661.48	0.41	0.75	0.194	21.74
					585.91	417.5469					
-0.20	0.631	0.566	27.18	2320.86			4079.03	0.45	0.87	0.094	21.99
					802.33	454.1753					
-0.30	0.501	0.450	31.53	3123.19			4533.20	0.50	1.01	-0.006	22.24
					1050.74	472.4631					
-0.40	0.398	0.357	36.45	4173.93			5005.66	0.56	1.17	-0.106	22.49
					1352.02	482.8965					
-0.50	0.316	0.284	41.94	5525.94			5488.56	0.61	1.34	-0.206	22.74
					1410.88	400.2786					
-0.60	0.251	0.225	46.99	6936.82			5888.84	0.65	1.50	-0.306	22.99
					854.45	192.5569					
-0.70	0.200	0.179	49.80	7791.27			6081.39	0.67	1.59	-0.406	23.24
					801.89	143.5449					
-0.80	0.158	0.142	52.30	8593.16			6224.93	0.69	1.67	-0.506	23.49
					1506.71	214.2411					
-0.90	0.126	0.113	56.70	10099.87			6439.17	0.71	1.81	-0.606	23.74
					1743.80	196.9566					
-1.00	0.100	0.090	61.40	11643.68			6636.12	0.74	1.96	-0.706	23.99
					4669.32	418.9158					
-1.10	0.079	0.071	72.50	16513.00			7055.04	0.78	2.32	-0.806	24.24
					3044.07	216.9343					
-1.20	0.063	0.057	78.90	19557.07			7271.97	0.81	2.52	-0.906	24.49
					3949.12	223.5498					
-1.30	0.050	0.045	86.50	23506.18			7495.52	0.83	2.77	-1.006	24.74
					3257.94	146.4934					
-1.40	0.040	0.036	92.30	26764.12			7642.01	0.85	2.95	-1.106	24.99
					2856.09	102.0111					
-1.50	0.032	0.028	97.10	29620.21			7744.02	0.86	3.11	-1.206	25.24
					4686.77	132.9686					
-1.60	0.025	0.023	104.50	34306.98			7876.98	0.87	3.34	-1.306	25.49
					8112.18	182.8160					
-1.70	0.020	0.018	116.20	42419.16			8059.80	0.89	3.72	-1.406	25.74
					9860.09	176.5054					
-1.80	0.016	0.014	129.00	52279.25			8236.30	0.91	4.13	-1.506	25.99
					6685.31	95.0604					
-1.90	0.013	0.011	137.00	58964.56			8331.36	0.92	4.38	-1.606	26.24
					7269.75	82.1104					
-2.00	0.010		145.20	66234.31			8413.47	0.93	4.64	-1.706	26.49
							9013.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4424

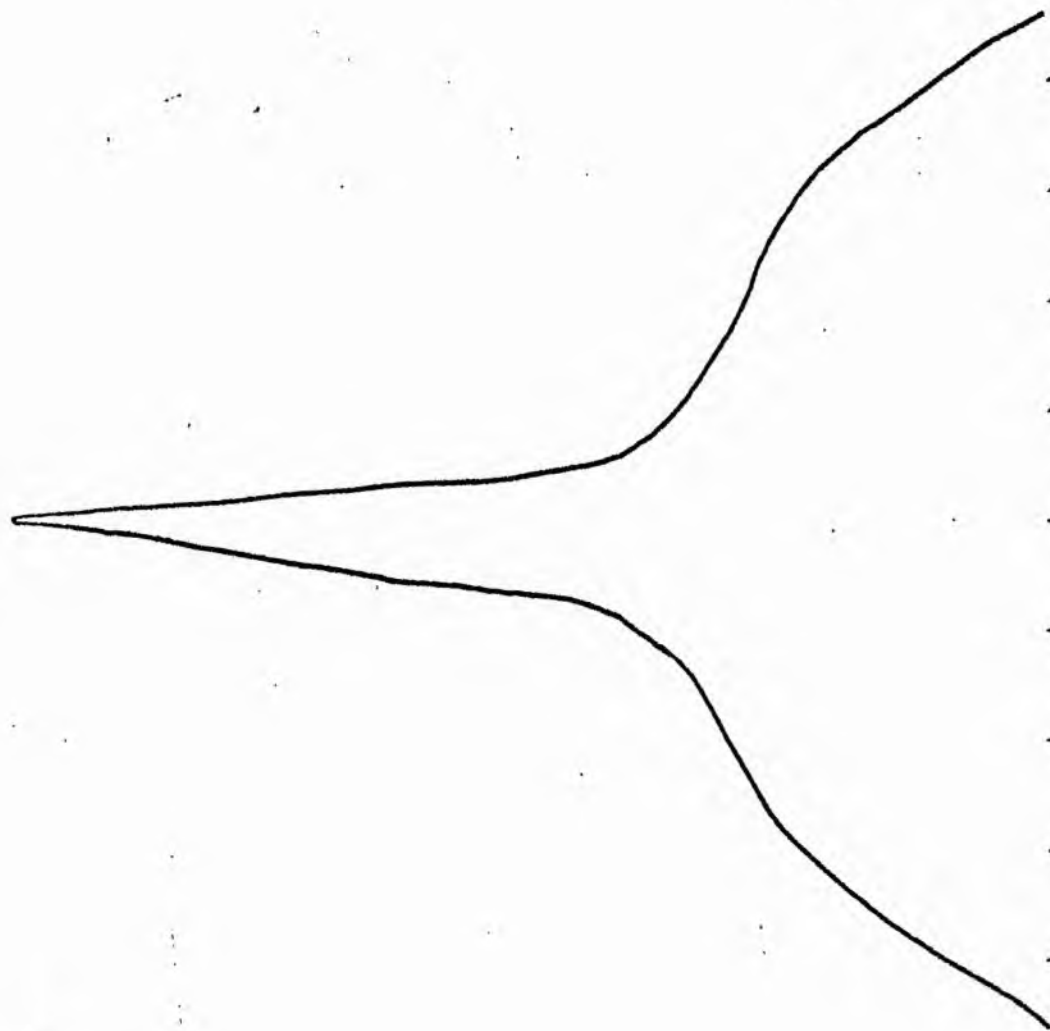
V-FILTER

Total luminosity	L_T	= 2.50
Total apparent magnitude	m_T	= 11.60
Apparent central surface brightness	μ_0	= 18.91
Major axis at threshold	$2a_m$	= 5.76
Minor axis at threshold	$2b_m$	= 3.81
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.11
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.85
Gradient of exponential component	$G(a)$	= -0.69
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.83
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.77
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.31
Axis ratio	b/a	= 0.72
Equivalent radius	r_1^*	= 0.22
Surface brightness	μ_1	= 20.84
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.88
Axis ratio	b/a	= 0.44
Equivalent radius	r_e^*	= 0.52
Surface brightness	μ_e	= 22.24
Mean surface brightness	μ_e'	= 12.15
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.76
Axis ratio	b/a	= 0.45
Equivalent radius	r_3^*	= 1.07
Surface brightness	μ_3	= 24.06
Concentration indices	$\{C_{21}$	= 2.34
	C_{32}	= 2.06

NGC 4438
B-Filter
Axis 1

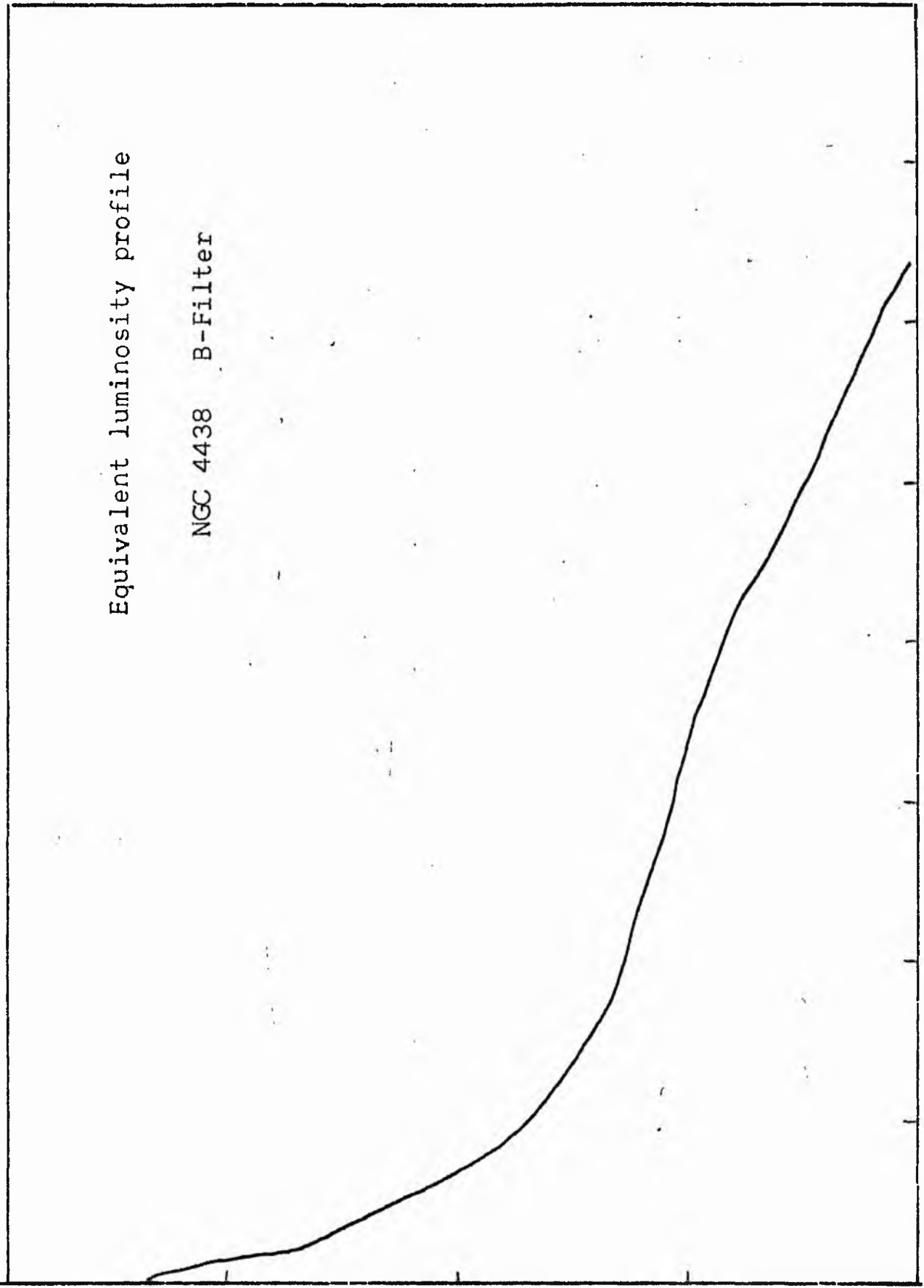


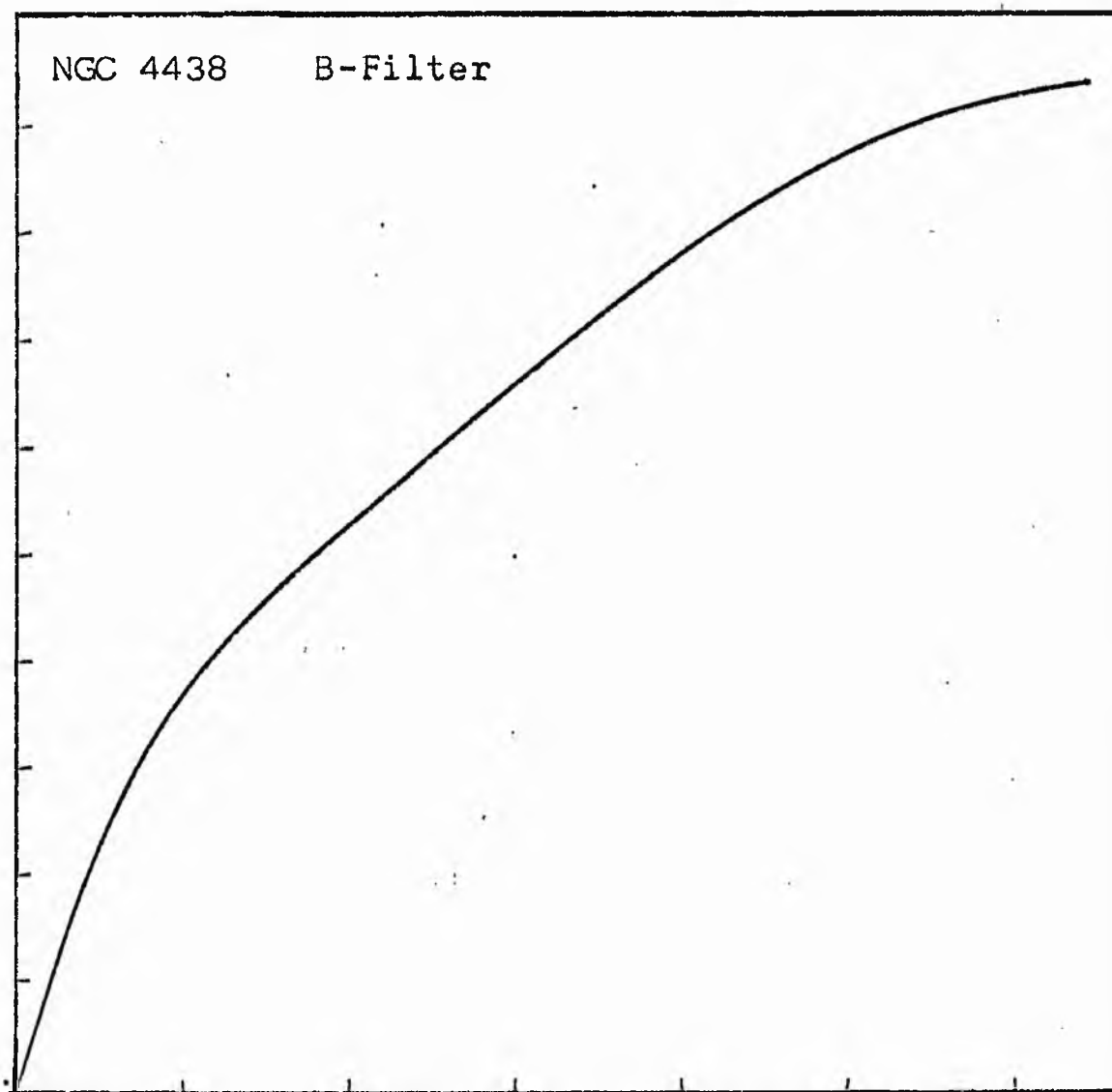
NGC 4438
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4438 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4438
B COLOUR

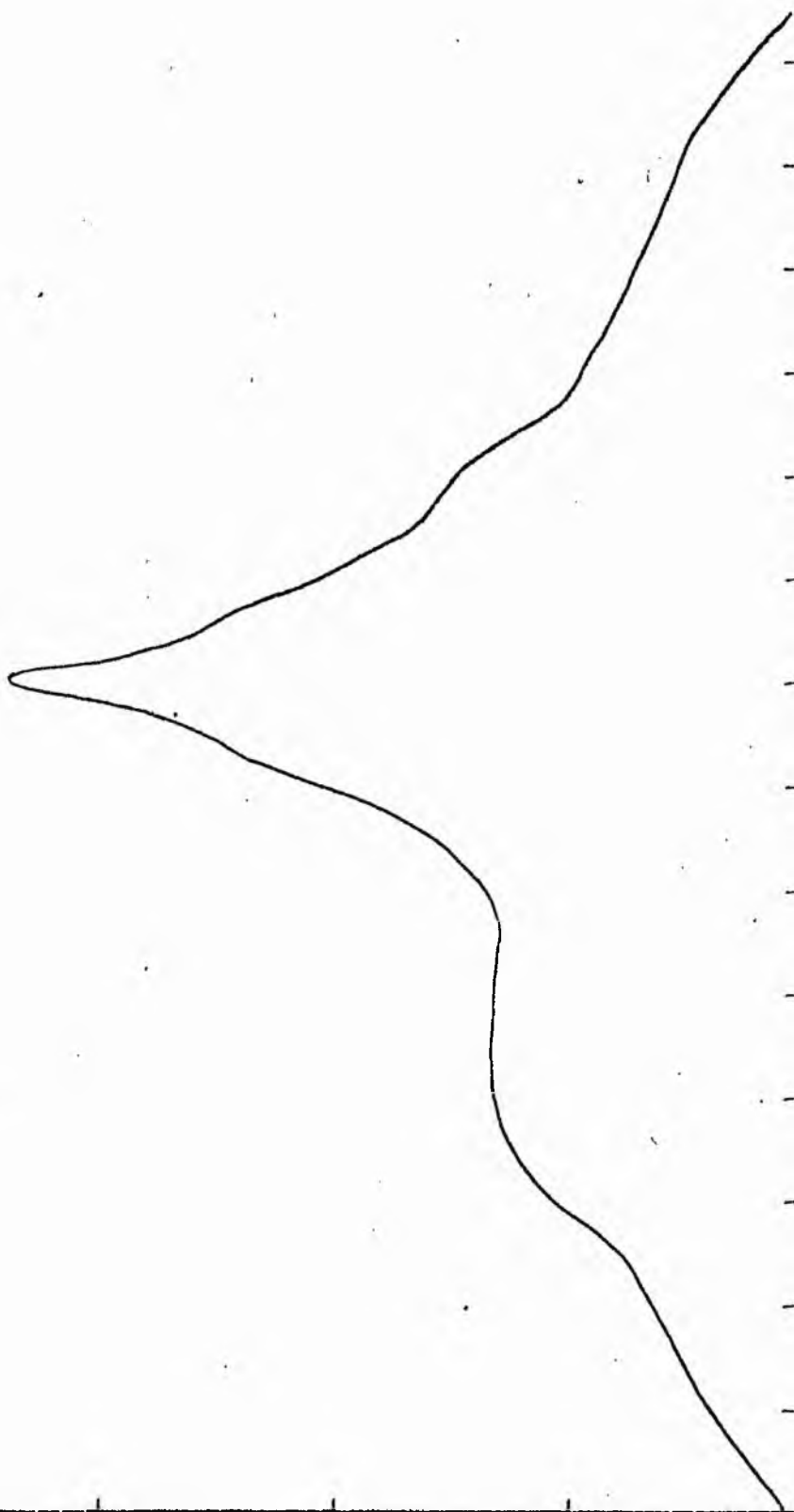
LOG I	I	T	R	AREA	ΔA	P	ΣP	R(I)	P	LOG J	M
1.35	22.387		0.0	0.0			0.0	0.0	0.0	2.028	18.97
1.30	19.953	21.170	2.60	21.24	21.24	449.5876	449.59	0.01	0.03	1.978	19.10
1.20	15.849	17.901	4.70	69.40	48.16	862.1099	1311.70	0.03	0.05	1.878	19.35
1.10	12.589	14.219	6.10	116.90	47.50	675.4167	1987.11	0.05	0.07	1.778	19.60
1.00	10.000	11.295	6.90	149.57	32.67	349.0229	2336.14	0.06	0.08	1.678	19.85
0.90	7.943	8.972	8.20	211.24	61.67	553.2742	2909.41	0.07	0.09	1.578	20.10
0.80	6.310	7.126	9.90	307.91	96.67	688.8848	3598.30	0.09	0.11	1.478	20.35
0.70	5.012	5.661	11.60	422.73	114.83	649.9927	4248.29	0.11	0.13	1.378	20.60
0.60	3.981	4.496	13.00	530.93	108.20	484.5007	4734.79	0.12	0.15	1.278	20.85
0.50	3.162	3.572	18.50	1075.21	544.28	1943.9873	6678.77	0.17	0.21	1.178	21.10
0.40	2.512	2.837	19.90	1244.10	168.89	479.1548	7157.92	0.18	0.22	1.078	21.35
0.30	1.995	2.254	24.70	1916.65	672.55	1515.6440	8673.56	0.22	0.28	0.978	21.60
0.20	1.585	1.790	30.23	2870.95	954.30	1708.2639	10381.82	0.27	0.34	0.878	21.85
0.10	1.259	1.422	32.23	3263.40	392.45	558.6227	10939.84	0.28	0.36	0.778	22.10
-0.00	1.000	1.129	35.40	3936.92	673.92	760.7095	11700.55	0.30	0.40	0.678	22.35
-0.10	0.794	0.897	37.50	4417.86	480.95	431.4866	12132.04	0.31	0.42	0.578	22.60
-0.20	0.631	0.713	42.60	5701.23	1283.37	914.5796	13046.61	0.33	0.48	0.478	22.85
-0.30	0.501	0.566	48.10	7268.41	1567.18	887.1365	13933.75	0.36	0.54	0.378	23.10
-0.40	0.398	0.450	56.70	10099.87	2831.46	1273.1538	15206.90	0.39	0.63	0.278	23.35
-0.50	0.316	0.357	68.50	14741.14	4641.27	1657.7051	16864.60	0.43	0.76	0.178	23.60
-0.60	0.251	0.284	77.00	18626.50	3885.37	1102.3081	17966.91	0.46	0.86	0.078	23.85
-0.70	0.200	0.225	93.20	27288.62	8662.12	1952.0696	19918.97	0.51	1.04	-0.022	24.10
-0.80	0.158	0.179	110.70	38498.61	11209.99	2006.6709	21925.64	0.56	1.24	-0.122	24.35
-0.90	0.126	0.142	132.70	55321.22	16822.61	2392.0178	24317.66	0.62	1.48	-0.222	24.60
-1.00	0.100	0.113	167.40	88036.06	32714.84	3695.0176	28012.67	0.72	1.87	-0.322	24.85
-1.10	0.079	0.090	184.50	106940.56	18904.50	1696.0454	29708.71	0.76	2.06	-0.422	25.10
-1.20	0.063	0.071	208.40	136441.12	29500.56	2102.3411	31811.05	0.82	2.33	-0.522	25.35
-1.30	0.050	0.057	220.80	153160.87	16719.75	946.4614	32757.52	0.84	2.46	-0.622	25.60
-1.40	0.040	0.045	237.40	177056.25	23895.37	1074.4534	33831.97	0.87	2.65	-0.722	25.85
-1.50	0.032	0.036	251.30	198396.87	21340.62	762.2219	34594.19	0.89	2.80	-0.822	26.10
-1.60	0.025	0.028	261.30	214500.25	16103.37	456.8691	35051.05	0.90	2.92	-0.922	26.35
-1.70	0.020	0.023	274.10	236030.00	21529.75	485.1931	35536.25	0.91	3.06	-1.022	26.60
-1.80	0.016	0.018	293.50	270623.87	34593.87	619.2634	36155.51	0.93	3.28	-1.122	26.85
-1.90	0.013	0.014	310.70	303271.87	32648.00	464.2300	36619.73	0.94	3.47	-1.222	27.10
-2.00	0.010	0.011	318.20	318090.00	14816.12	167.3673	36787.10	0.94	3.55	-1.322	27.35
-∞							38987.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4438

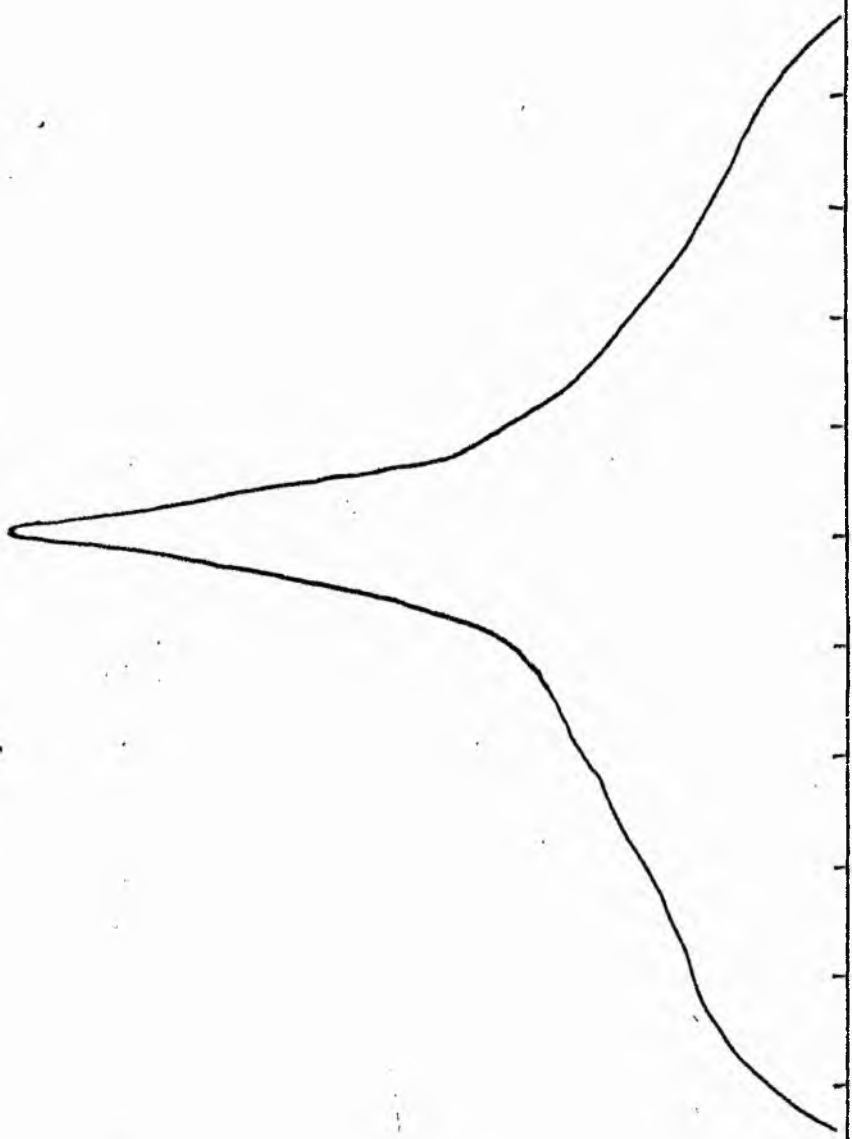
B-FILTER

Total luminosity	L_T	= 10.83
Total apparent magnitude	m_T	= 10.86
Apparent central surface brightness	μ_0	= 18.97
Major axis at threshold	$2a_m$	= 14.15
Minor axis at threshold	$2b_m$	= 7.85
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.70
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.74
Gradient of exponential component	$G(a)$	= -0.40
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.33
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.53
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.68
Axis ratio	b/a	= 0.54
Equivalent radius	r_1^*	= 0.47
Surface brightness	μ_1	= 21.75
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.97
Axis ratio	b/a	= 0.42
Equivalent radius	r_e^*	= 1.50
Surface brightness	μ_e	= 24.05
Mean surface brightness	μ_e'	= 13.74
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 4.43
Axis ratio	b/a	= 0.47
Equivalent radius	r_3^*	= 2.99
Surface brightness	μ_3	= 25.04
Concentration indices	$\{C_{21}$	= 3.19
	C_{32}	= 2.00

NGC 4438
V-Filter
Axis 1

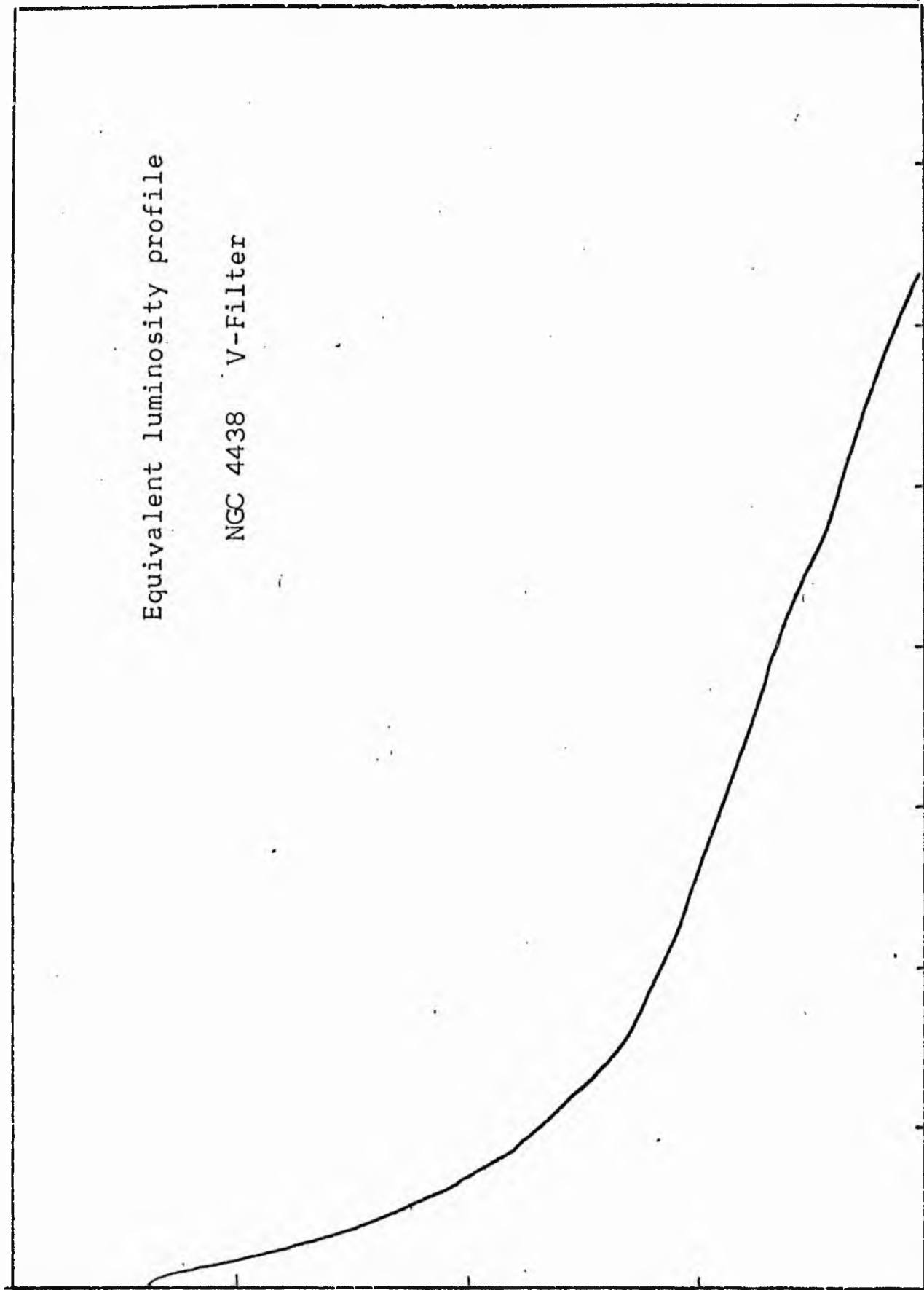


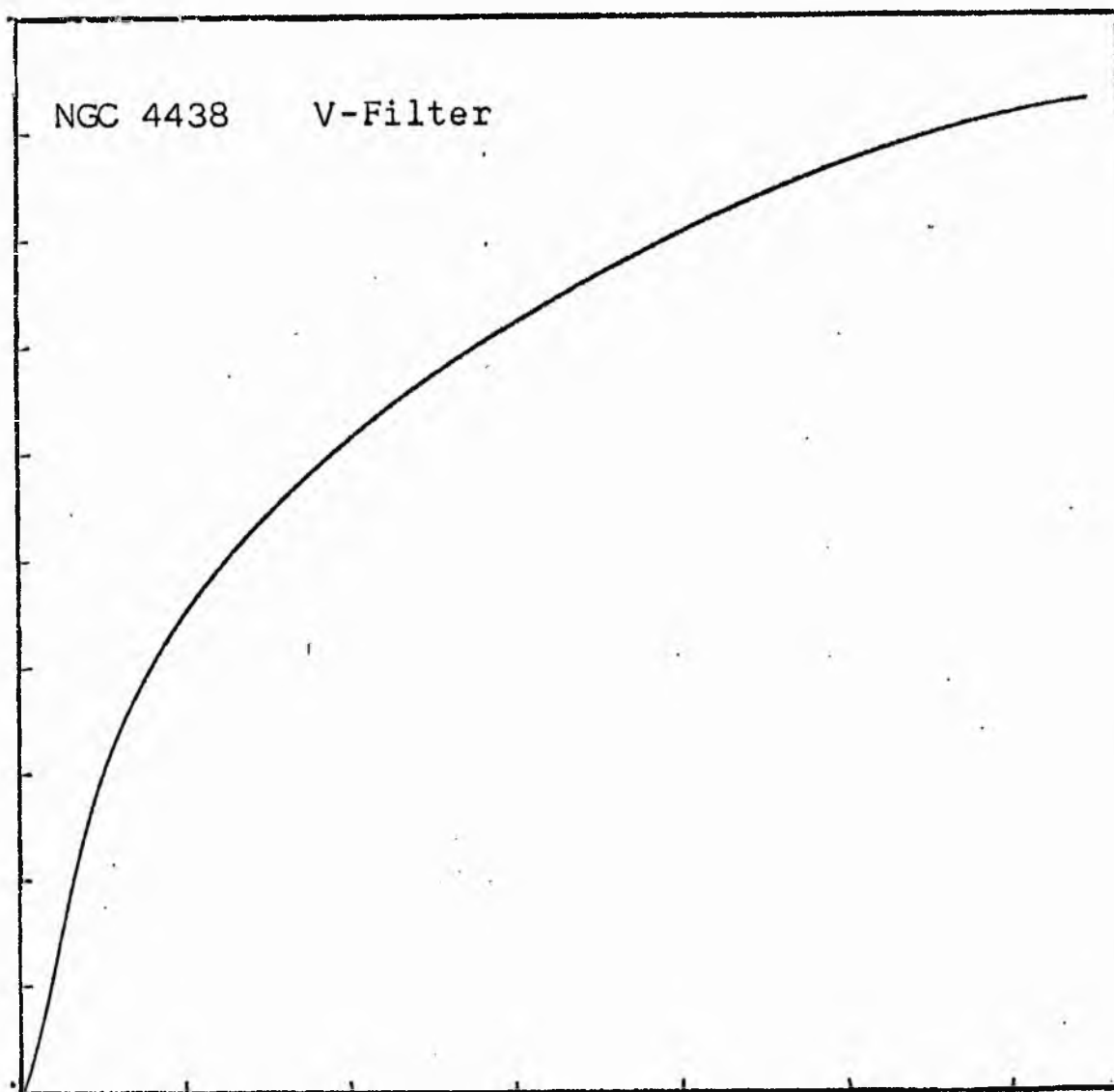
NGC 4438
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4438 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4438
V COLOUR

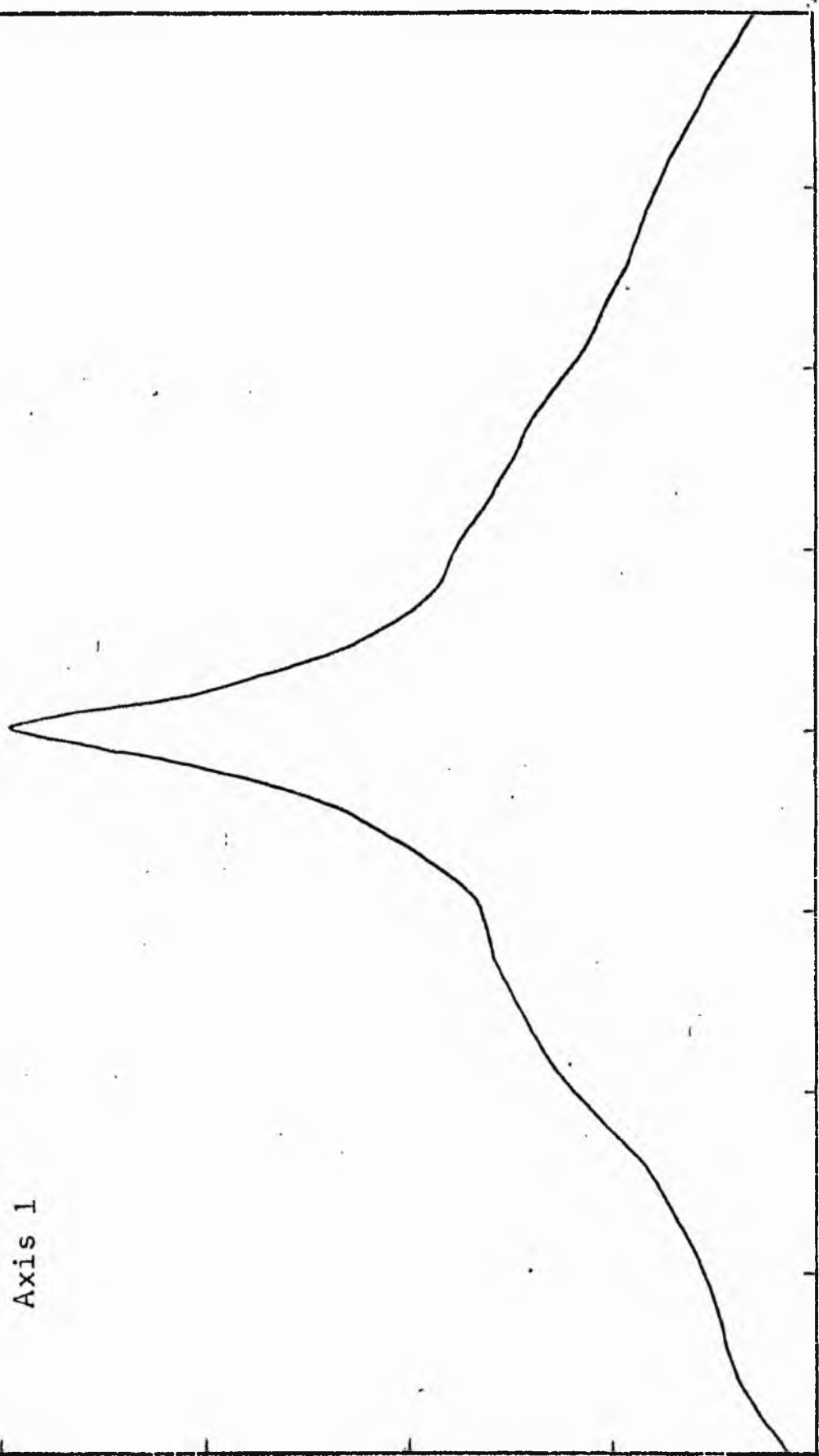
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ_i	LOG J	μ
1.37	23.442		0.0	0.0	66.48	1442.3586	0.0	0.0	0.0	1.812	18.09
1.30	19.953	21.697	4.60	66.48	42.88	767.6321	1442.36	0.04	0.08	1.742	18.27
1.20	15.849	17.901	5.90	109.36	40.21	571.7815	2209.99	0.06	0.10	1.642	18.52
1.10	12.589	14.219	6.90	149.57	118.65	1340.1045	2781.77	0.08	0.11	1.542	18.77
1.00	10.000	11.295	9.24	268.22	93.48	838.6565	4121.87	0.11	0.15	1.442	19.02
0.90	7.943	8.972	10.73	361.70	151.41	1079.0132	4960.53	0.13	0.18	1.342	19.27
0.80	6.310	7.126	12.78	513.11	230.01	1302.0461	6039.54	0.16	0.21	1.242	19.52
0.70	5.012	5.661	15.38	743.13	277.01	1245.5720	7341.59	0.20	0.25	1.142	19.77
0.60	3.981	4.496	18.02	1020.14	201.56	719.8948	8587.16	0.23	0.30	1.042	20.02
0.50	3.162	3.572	19.72	1221.70	392.86	1114.5703	9307.05	0.25	0.32	0.942	20.27
0.40	2.512	2.837	22.67	1614.55	402.71	907.5339	10421.62	0.28	0.37	0.842	20.52
0.30	1.995	2.254	25.34	2017.26	579.46	1037.2722	11329.15	0.31	0.42	0.742	20.77
0.20	1.585	1.790	28.75	2596.72	721.58	1026.0242	12366.42	0.34	0.47	0.642	21.02
0.10	1.259	1.422	32.50	3318.31	574.25	648.5925	13392.45	0.36	0.53	0.542	21.27
-0.00	1.000	1.129	35.20	3892.56	836.91	750.8477	14041.04	0.38	0.58	0.442	21.52
-0.10	0.794	0.897	38.80	4729.47	1160.66	827.1365	14791.88	0.40	0.64	0.342	21.77
-0.20	0.631	0.713	43.30	5890.14	2185.29	1237.0298	15619.02	0.42	0.71	0.242	22.02
-0.30	0.501	0.566	50.70	8075.43	2639.16	1186.6863	16856.04	0.46	0.83	0.142	22.27
-0.40	0.398	0.450	58.40	10714.59	2233.92	797.8809	18042.73	0.49	0.96	0.042	22.52
-0.50	0.316	0.357	64.20	12948.50	3022.37	857.4695	18840.61	0.51	1.05	-0.058	22.77
-0.60	0.251	0.284	71.30	15970.87	5359.78	1207.8633	19698.07	0.53	1.17	-0.158	23.02
-0.70	0.200	0.225	82.40	21330.65	6843.42	1225.0232	20905.94	0.57	1.35	-0.258	23.27
-0.80	0.158	0.179	94.70	28174.07	11445.45	1627.4363	22130.96	0.60	1.55	-0.358	23.52
-0.90	0.126	0.142	112.30	39619.52	12174.52	1375.0657	23758.39	0.65	1.84	-0.458	23.77
-1.00	0.100	0.113	128.40	51794.04	16277.46	1460.3562	25133.45	0.68	2.10	-0.558	24.02
-1.10	0.079	0.090	147.20	68071.50	20913.75	1490.4067	26593.81	0.72	2.41	-0.658	24.27
-1.20	0.063	0.071	168.30	88985.25	18768.31	1062.4250	28084.21	0.76	2.76	-0.758	24.52
-1.30	0.050	0.057	185.20	107753.56	26990.56	1213.6282	29146.64	0.79	3.03	-0.858	24.77
-1.40	0.040	0.045	207.10	134744.12	32894.37	1174.8867	30360.26	0.82	3.39	-0.958	25.02
-1.50	0.032	0.036	231.00	167638.50	37124.50	1053.2603	31535.15	0.86	3.79	-1.058	25.27
-1.60	0.025	0.028	255.30	204763.00	27664.56	623.4470	32588.41	0.88	4.18	-1.158	25.52
-1.70	0.020	0.023	272.00	232427.56	28148.69	503.8884	33211.85	0.90	4.46	-1.258	25.77
-1.80	0.016	0.018	288.00	260576.25	29947.81	425.8354	33715.74	0.92	4.72	-1.358	26.02
-1.90	0.013	0.014	304.10	290524.06	25171.44	284.3054	34141.57	0.93	4.98	-1.458	26.27
-2.00	0.010	0.011	317.00	315695.50			34425.87	0.93	5.19	-1.558	26.52
-∞							36825.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4438

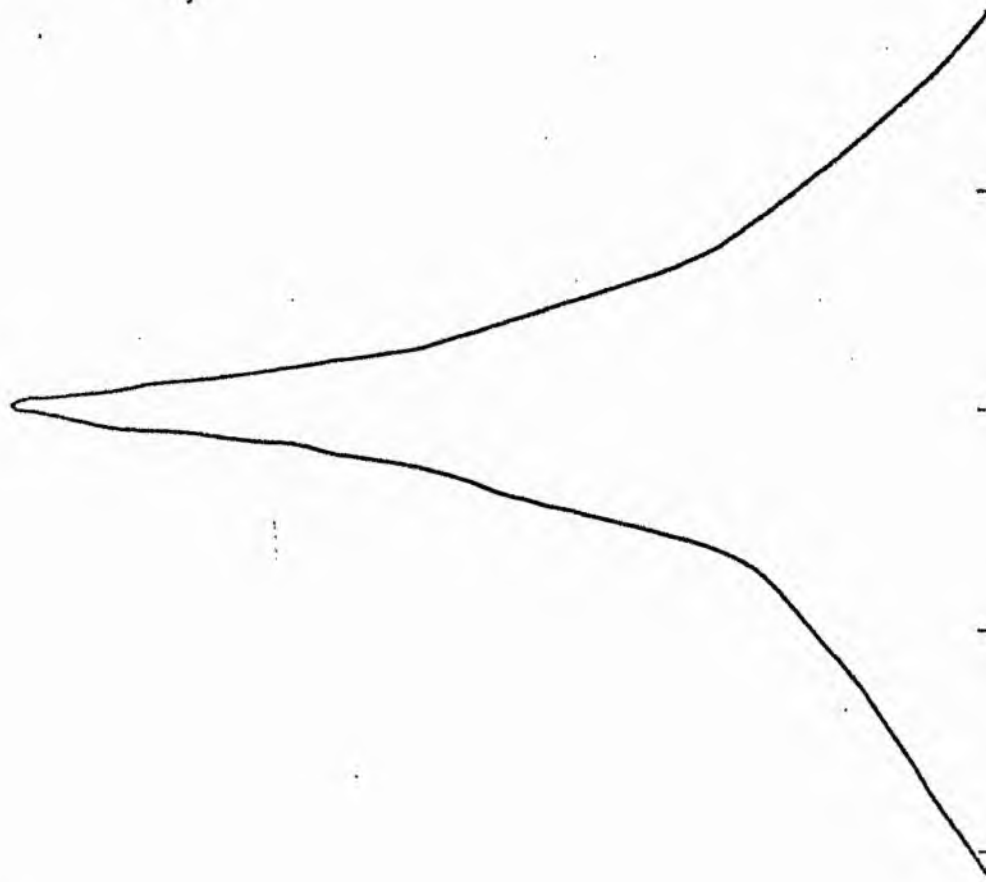
V-FILTER

Total luminosity	L_T	= 10.23
Total apparent magnitude	m_T	= 10.10
Apparent central surface brightness	μ_0	= 18.09
Major axis at threshold	$2a_m$	= 12.68
Minor axis at threshold	$2b_m$	= 8.60
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.58
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.82
Gradient of exponential component	$G(a)$	= -0.31
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.32
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.33
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.57
Axis ratio	b/a	= 0.56
Equivalent radius	r_1^*	= 0.32
Surface brightness	μ_1	= 20.27
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.38
Axis ratio	b/a	= 0.43
Equivalent radius	r_e^*	= 1.02
Surface brightness	μ_e	= 22.65
Mean surface brightness	μ_e'	= 12.14
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 3.75
Axis ratio	b/a	= 0.59
Equivalent radius	r_3^*	= 2.69
Surface brightness	μ_3	= 24.46
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 3.13 \\ = 2.64 \end{matrix}$

NGC 4442
B-Filter
Axis 1

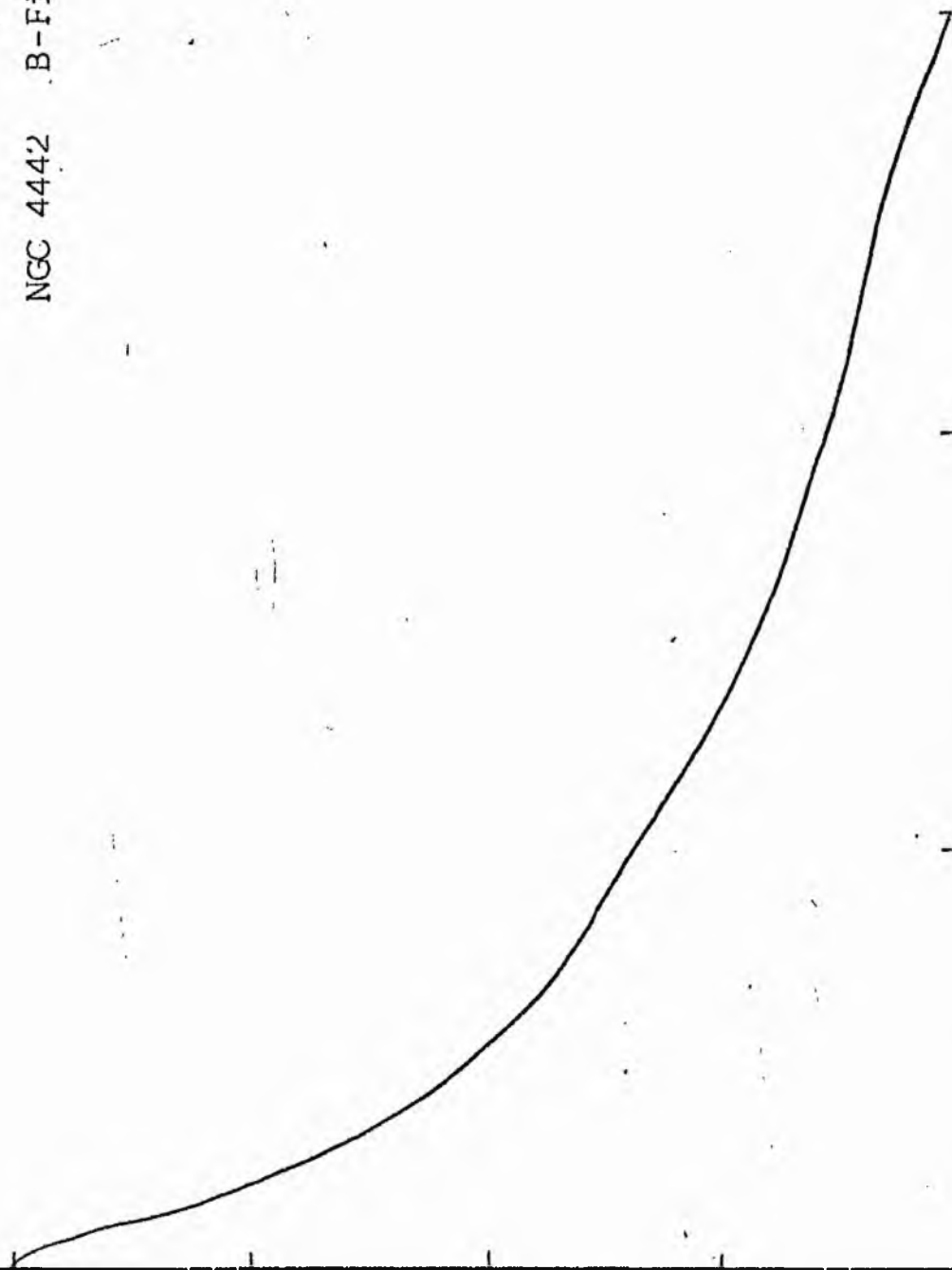


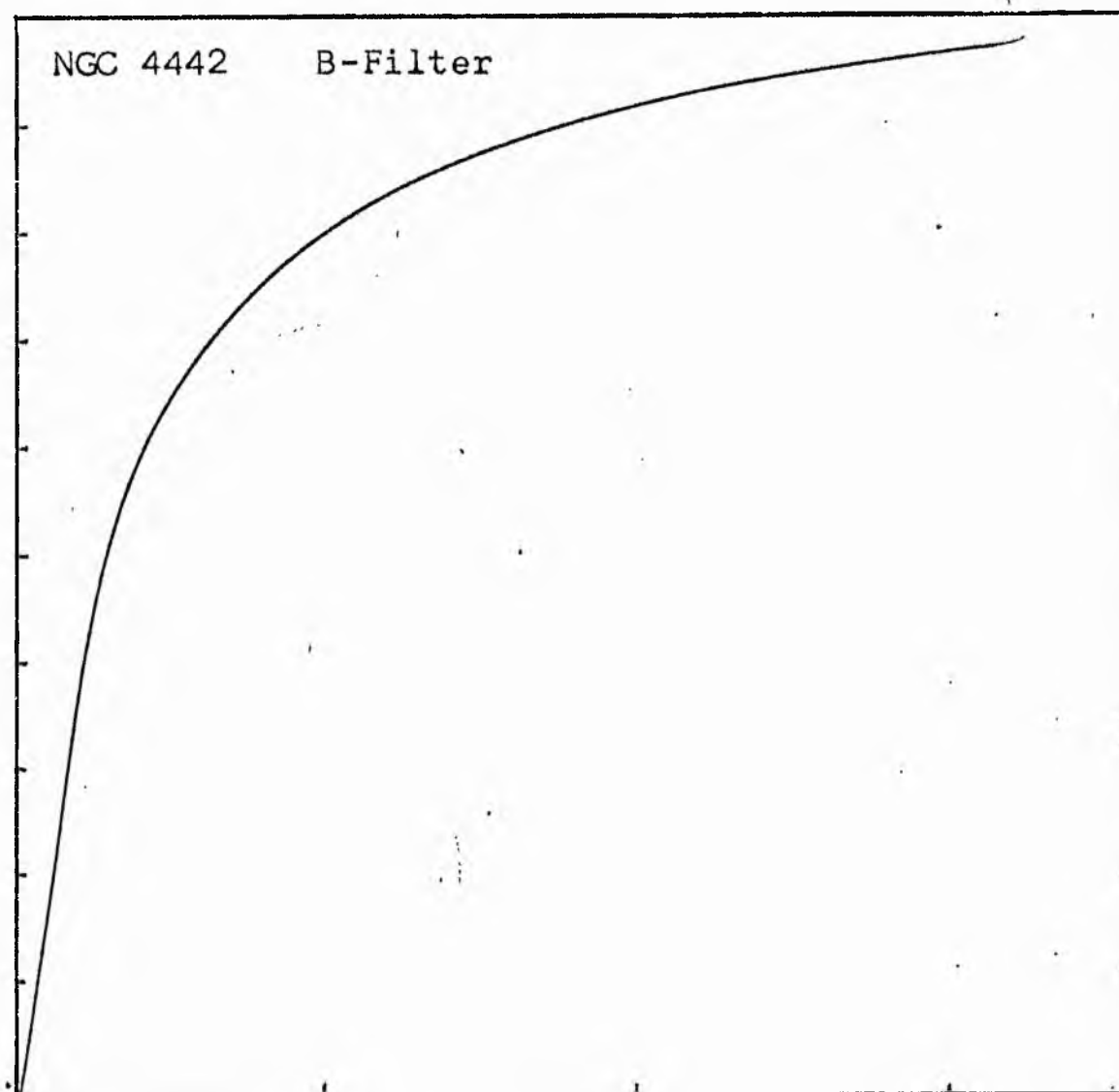
NGC 4442
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4442 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4442
B COLOUR

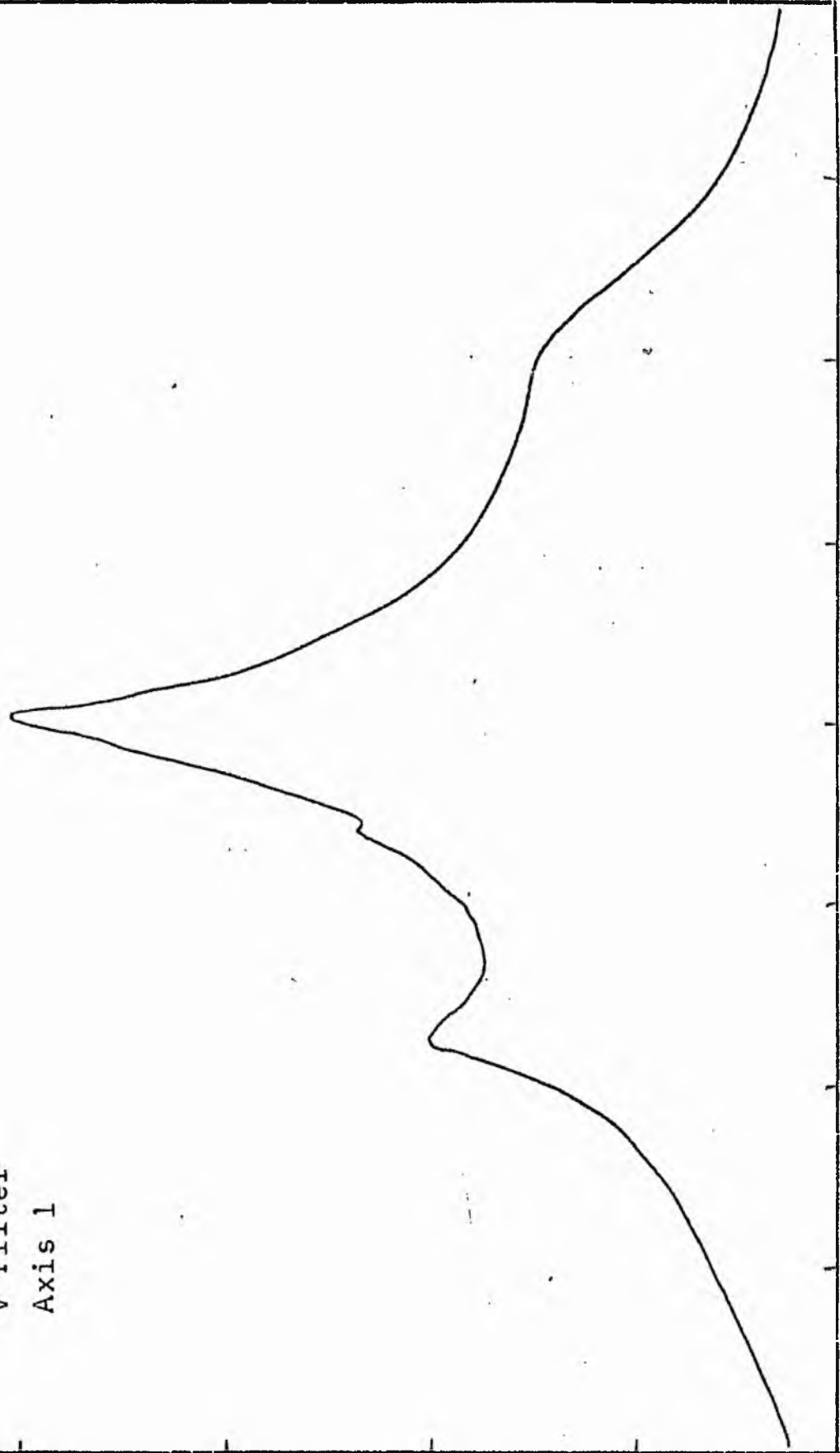
LOG I	I	T	R	AREB	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.94	87.096		0.0	0.0			0.0	0.0	0.0	1.326	17.17
1.90	79.433	83.263	0.00	2.01	2.01	167.4131	167.41	0.01	0.06	1.286	17.27
1.80	63.096	71.244	1.50	7.07	9.06	360.4317	927.88	0.03	0.11	1.186	17.52
1.70	50.119	56.607	2.80	24.63	17.56	994.1040	1521.97	0.09	0.21	1.086	17.77
1.60	39.811	44.965	3.33	34.84	10.21	458.9409	1980.91	0.12	0.24	0.986	18.02
1.50	31.623	35.717	4.50	63.62	28.78	1027.9399	3008.85	0.18	0.33	0.886	18.27
1.40	25.119	28.371	5.10	81.71	18.10	913.3813	3522.23	0.21	0.37	0.786	18.52
1.30	19.952	22.936	5.90	109.36	27.65	623.0200	4145.25	0.24	0.43	0.686	18.77
1.20	15.849	17.901	6.80	145.27	35.91	642.7830	4788.03	0.28	0.50	0.586	19.02
1.10	12.589	14.219	8.10	206.12	60.85	865.2629	5653.29	0.33	0.59	0.486	19.27
1.00	10.000	11.299	9.23	267.64	61.52	694.8560	6348.15	0.37	0.68	0.386	19.52
0.90	7.943	8.972	10.70	359.68	92.04	825.7388	7173.89	0.42	0.78	0.286	19.77
0.80	6.310	7.126	11.90	444.88	85.20	607.1660	7781.05	0.46	0.87	0.186	20.02
0.70	5.012	5.661	12.70	506.71	61.83	349.9792	8131.03	0.48	0.93	0.086	20.27
0.60	3.981	4.496	13.80	598.28	91.58	411.7717	8542.80	0.50	1.01	-0.014	20.52
0.50	3.162	3.572	15.10	716.31	118.03	421.5605	8964.36	0.53	1.11	-0.114	20.77
0.40	2.512	2.837	16.50	855.30	138.98	394.3064	9358.66	0.55	1.21	-0.214	21.02
0.30	1.995	2.254	18.20	1040.62	185.32	417.6333	9776.29	0.58	1.33	-0.314	21.27
0.20	1.585	1.790	20.90	1372.28	331.66	593.6882	10369.98	0.61	1.53	-0.414	21.52
0.10	1.259	1.422	23.50	1734.94	362.67	515.6753	10885.65	0.64	1.72	-0.514	21.77
-0.00	1.000	1.129	25.10	1979.23	244.29	275.9131	11161.56	0.66	1.84	-0.614	22.02
-0.10	0.794	0.897	27.60	2393.14	413.90	371.3372	11532.90	0.68	2.02	-0.714	22.27
-0.20	0.631	0.713	30.60	2941.66	548.52	390.8977	11923.79	0.70	2.24	-0.814	22.52
-0.30	0.501	0.566	33.00	3421.19	479.53	271.4492	12195.24	0.72	2.42	-0.914	22.77
-0.40	0.398	0.450	38.80	4729.47	1308.28	588.2595	12783.50	0.75	2.84	-1.014	23.02
-0.50	0.316	0.357	42.30	5621.21	891.74	318.4990	13102.00	0.77	3.10	-1.114	23.27
-0.60	0.251	0.284	47.60	7118.09	1496.87	424.6729	13526.67	0.80	3.49	-1.214	23.52
-0.70	0.200	0.225	52.30	8593.16	1475.07	332.4158	13859.08	0.82	3.83	-1.314	23.77
-0.80	0.158	0.179	56.90	10171.25	1578.09	282.4885	14141.57	0.83	4.17	-1.414	24.02
-0.90	0.126	0.142	60.30	11423.11	1251.86	178.0019	14319.57	0.84	4.42	-1.514	24.27
-1.00	0.100	0.113	65.70	13560.65	2137.54	241.4258	14561.00	0.86	4.82	-1.614	24.52
-1.10	0.079	0.090	74.60	17483.45	3922.80	351.9385	14912.93	0.88	5.47	-1.714	24.77
-1.20	0.063	0.071	81.20	20713.89	3230.44	230.2144	15143.14	0.89	5.95	-1.814	25.02
-1.30	0.050	0.057	88.60	24661.37	3947.47	223.4552	15366.60	0.91	6.49	-1.914	25.27
-1.40	0.040	0.045	96.30	29134.15	4472.78	201.1174	15567.71	0.92	7.06	-2.014	25.52
-1.50	0.032	0.036	109.70	37806.20	8672.05	309.7378	15877.45	0.94	8.04	-2.114	25.77
-1.60	0.025	0.028	119.20	44637.75	6831.55	193.8172	16071.27	0.95	8.74	-2.214	26.02
-1.70	0.020	0.023	126.40	50193.09	5555.34	125.1943	16196.46	0.95	9.27	-2.314	26.27
-1.80	0.016	0.018	134.70	57001.33	6808.24	121.8735	16318.33	0.96	9.87	-2.414	26.52
-1.90	0.013	0.014	142.30	63615.01	6613.68	94.0412	16412.37	0.97	10.43	-2.514	26.77
-2.00	0.010	0.011	150.30	70968.81	7353.80	83.0591	16495.43	0.97	11.02	-2.614	27.02
-∞							16975.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4442

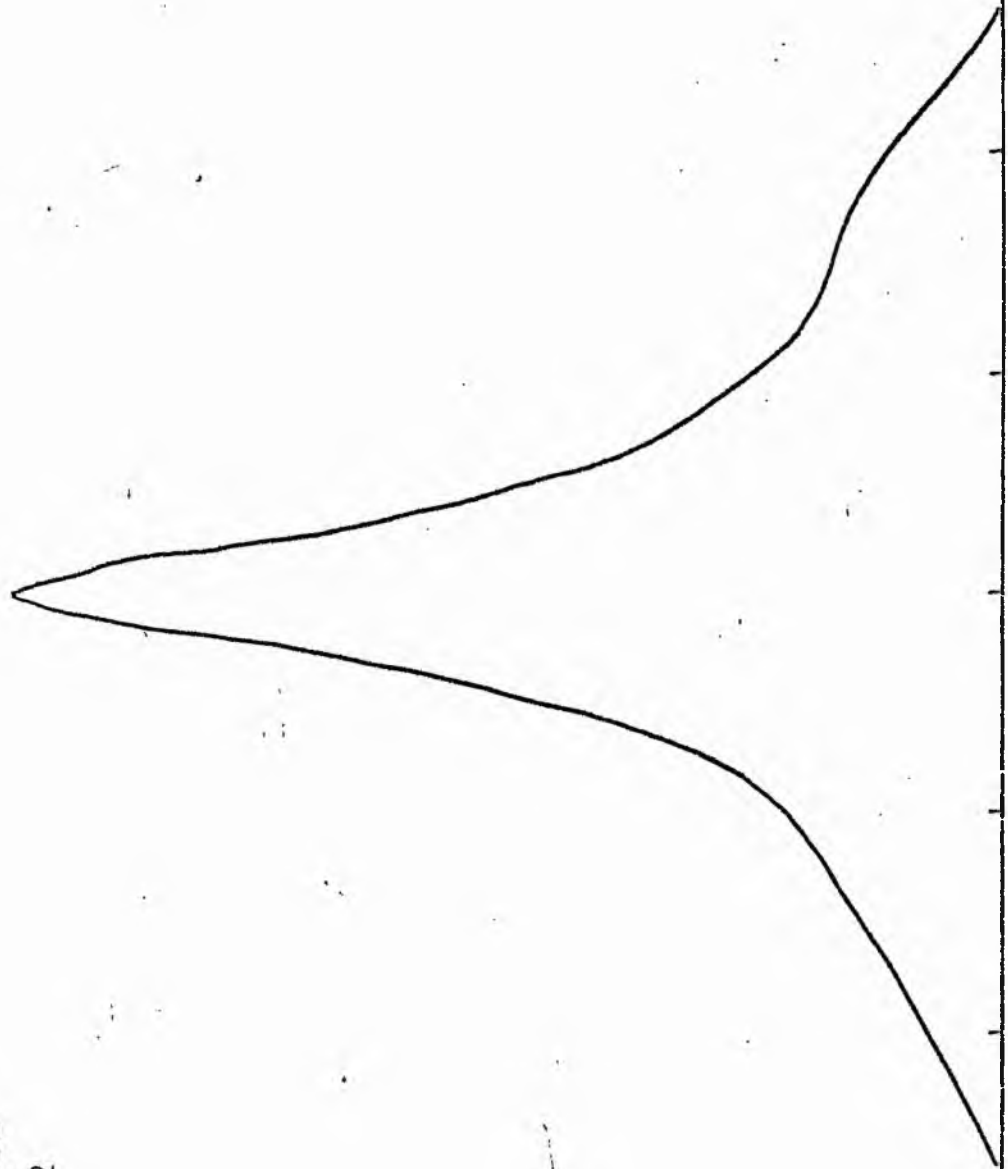
B-FILTER

Total luminosity	L_T	= 4.72
Total apparent magnitude	m_T	= 11.45
Apparent central surface brightness	μ_o	= 17.17
Major axis at threshold	$2a_m$	= 9.47
Minor axis at threshold	$2b_m$	= 3.28
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 4.43
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.51
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.81
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -1.41
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.15
Axis ratio	b/a	= 0.55
Equivalent radius	r_1^*	= 0.10
Surface brightness	μ_1	= 18.83
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.28
Axis ratio	b/a	= 0.59
Equivalent radius	r_e^*	= 0.23
Surface brightness	μ_e	= 20.52
Mean surface brightness	μ_e'	= 10.25
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.08
Axis ratio	b/a	= 0.41
Equivalent radius	r_3^*	= 0.64
Surface brightness	μ_3	= 23.02
Concentration indices	$\left\{ \begin{matrix} C_{21} \\ C_{32} \end{matrix} \right.$	= 2.26
		= 2.80

NGC 4442
V-Filter
Axis 1

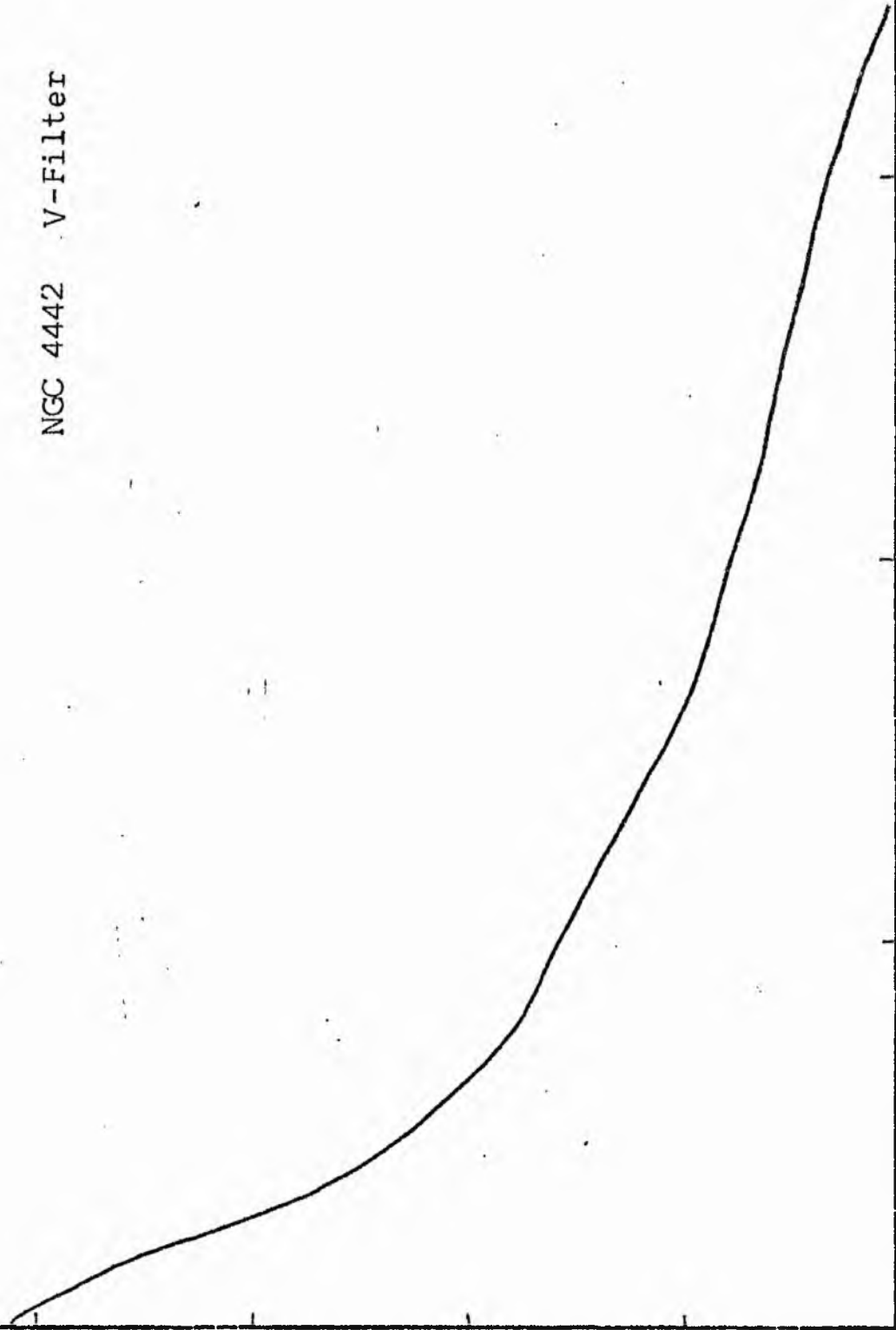


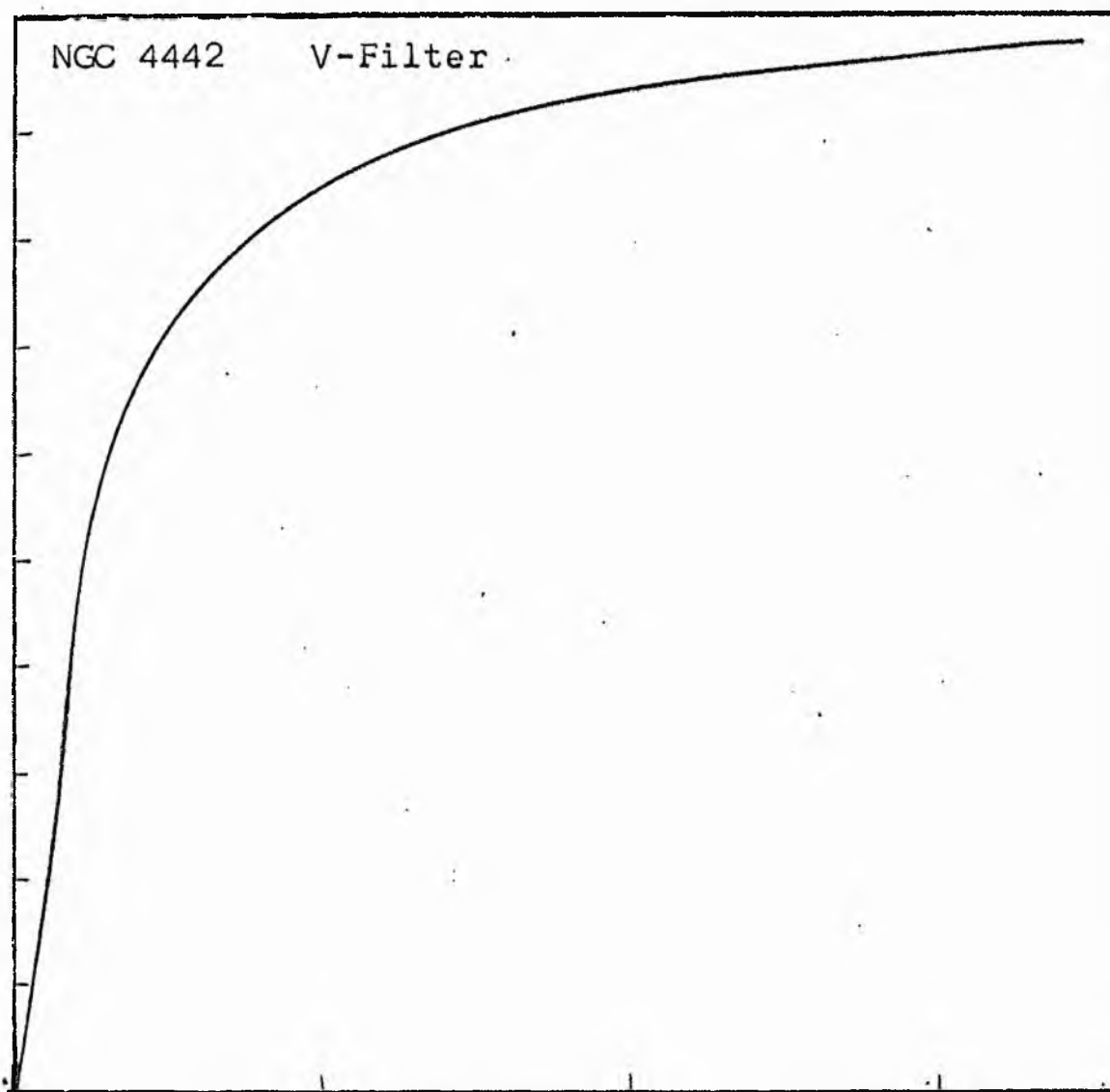
NGC 4442
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4442 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4442
V COLOUR

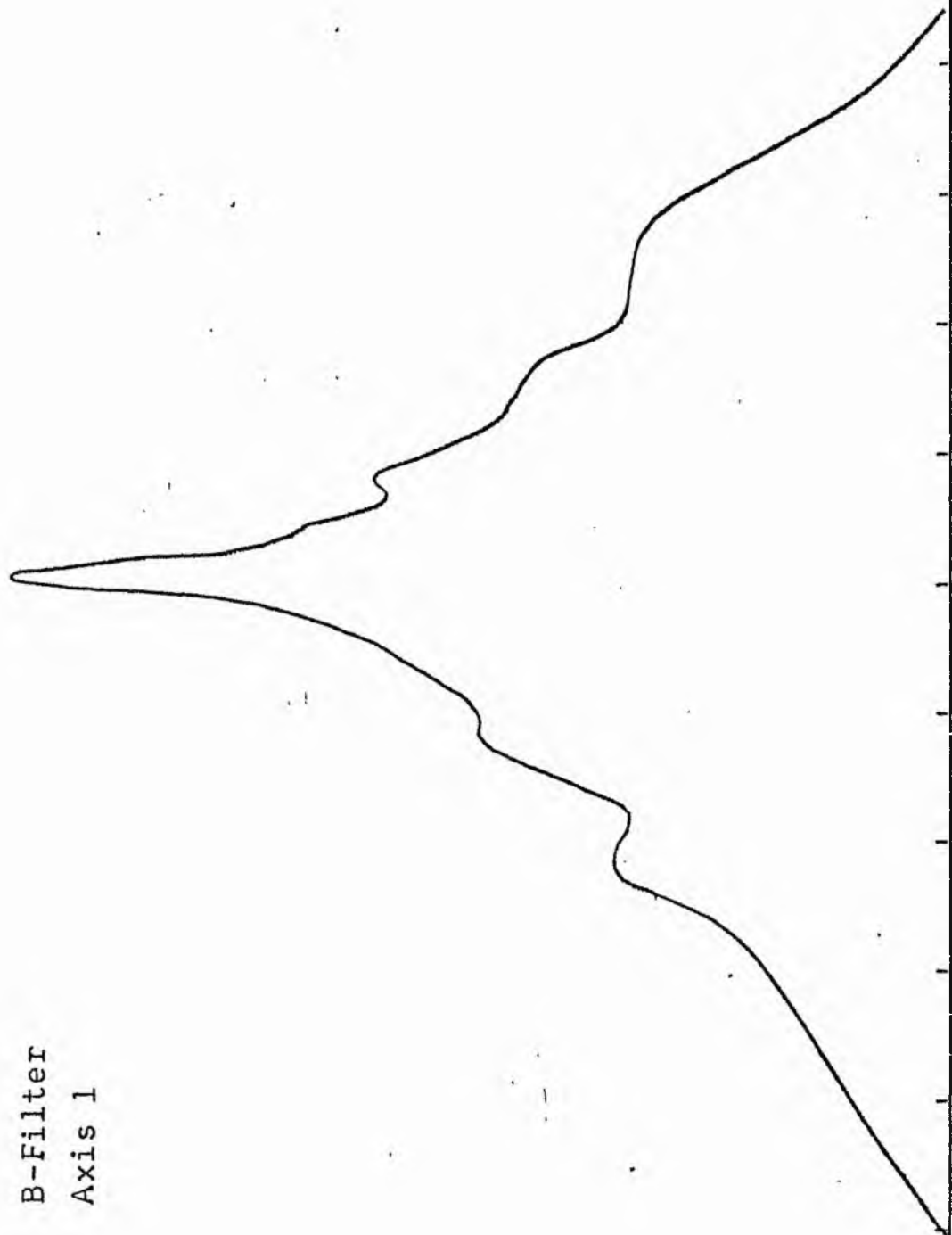
LOG I	I	T	N	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.05	112.202		0.0	0.0			0.0	0.0	0.0	0.863	16.63
		106.101			5.31	563.3193					
2.00	100.000		1.30	5.31			563.32	0.02	0.11	0.813	16.76
		89.716			24.88	2232.2678					
1.90	79.433		3.10	30.19			2795.59	0.08	0.25	0.713	17.01
		71.264			25.23	1797.7783					
1.80	63.096		4.20	55.42			4593.36	0.13	0.34	0.613	17.26
		56.607			39.62	2242.5193					
1.70	50.119		5.50	95.03			6835.88	0.20	0.45	0.513	17.51
		44.965			37.70	1695.1250					
1.60	39.811		6.50	132.73			8531.00	0.25	0.53	0.413	17.76
		35.717			25.64	915.6077					
1.50	31.623		7.10	158.37			9446.61	0.28	0.58	0.313	18.01
		28.371			137.22	3893.1541					
1.40	25.119		9.70	295.59			13339.76	0.39	0.79	0.213	18.26
		22.536			112.69	2539.5132					
1.30	19.952		11.40	408.28			15879.27	0.47	0.93	0.113	18.51
		17.901			59.31	1061.7454					
1.20	15.849		12.20	467.59			16941.02	0.50	0.99	0.013	18.76
		14.219			71.53	1017.1428					
1.10	12.589		13.10	539.13			17958.16	0.53	1.06	-0.087	19.01
		11.295			40.23	454.4060					
1.00	10.000		13.58	579.36			18412.56	0.54	1.10	-0.187	19.26
		8.972			105.99	950.8630					
0.90	7.943		14.77	685.35			19363.42	0.57	1.20	-0.287	19.51
		7.126			166.84	1188.9761					
0.80	6.310		16.47	852.19			20552.39	0.60	1.34	-0.387	19.76
		5.661			171.35	969.9448					
0.70	5.012		18.05	1023.54			21522.33	0.63	1.47	-0.487	20.01
		4.496			182.10	818.8230					
0.60	3.981		19.59	1205.64			22341.15	0.65	1.59	-0.587	20.26
		3.572			171.89	613.9395					
0.50	3.162		20.94	1377.53			22955.09	0.67	1.70	-0.687	20.51
		2.837			229.91	652.2559					
0.40	2.512		22.62	1607.44			23607.34	0.69	1.84	-0.787	20.76
		2.254			278.30	627.1670					
0.30	1.995		24.50	1885.74			24234.51	0.71	1.99	-0.887	21.01
		1.790			332.11	594.5012					
0.20	1.585		26.57	2217.85			24829.01	0.73	2.16	-0.987	21.26
		1.422			438.82	623.9595					
0.10	1.259		29.08	2656.67			25452.96	0.75	2.36	-1.087	21.51
		1.129			540.24	610.1770					
-0.00	1.000		31.90	3196.91			26063.14	0.76	2.59	-1.187	21.76
		0.897			245.04	219.8433					
-0.10	0.794		33.10	3441.96			26282.98	0.77	2.69	-1.287	22.01
		0.713			1236.46	881.1477					
-0.20	0.631		38.59	4678.42			27164.12	0.80	3.14	-1.387	22.26
		0.566			1925.90	1090.1919					
-0.30	0.501		45.85	6604.32			28254.32	0.83	3.73	-1.487	22.51
		0.450			1215.14	546.3809					
-0.40	0.398		49.89	7819.46			28800.70	0.84	4.05	-1.587	22.76
		0.357			1239.91	442.8540					
-0.50	0.316		53.70	9059.37			29243.55	0.86	4.36	-1.687	23.01
		0.284			1436.18	407.4521					
-0.60	0.251		57.80	10495.55			29651.00	0.87	4.70	-1.787	23.26
		0.225			1697.87	382.6262					
-0.70	0.200		62.30	12193.43			30033.62	0.88	5.06	-1.887	23.51
		0.179			2078.07	371.9878					
-0.80	0.158		67.40	14271.49			30405.61	0.89	5.48	-1.987	23.76
		0.142			1699.38	241.6353					
-0.90	0.126		71.30	15970.87			30847.24	0.90	5.79	-2.087	24.01
		0.113			3388.41	382.7058					
-1.00	0.100		78.50	19359.28			31029.94	0.91	6.38	-2.187	24.26
		0.090			3499.22	313.9360					
-1.10	0.079		85.30	22858.50			31343.87	0.92	6.93	-2.287	24.51
		0.071			5256.11	374.5715					
-1.20	0.063		94.60	28114.61			31718.45	0.93	7.69	-2.387	24.76
		0.057			7384.40	418.0098					
-1.30	0.050		106.30	35499.00			32136.45	0.94	8.64	-2.487	25.01
		0.045			5472.51	246.0698					
-1.40	0.040		114.20	40971.51			32382.52	0.95	9.28	-2.587	25.26
		0.036			8351.77	298.2983					
-1.50	0.032		125.30	49323.28			32680.82	0.96	10.18	-2.687	25.51
		0.028			10331.91	293.1255					
-1.60	0.025		137.80	59655.19			32973.94	0.97	11.23	-2.787	25.76
		0.023			7311.00	164.7594					
-1.70	0.020		146.00	66966.19			33138.70	0.97	11.86	-2.887	26.01
		0.018			8705.44	155.8350					
-1.80	0.016		155.20	75671.62			33294.53	0.98	12.61	-2.987	26.26
		0.014			10164.50	144.6021					
-1.90	0.013		165.30	85841.12			33439.13	0.98	13.43	-3.087	26.51
		0.011			7099.75	80.1897					
-2.00	0.010		172.00	92940.87			33519.32	0.98	13.98	-3.187	26.76
							34139.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4442

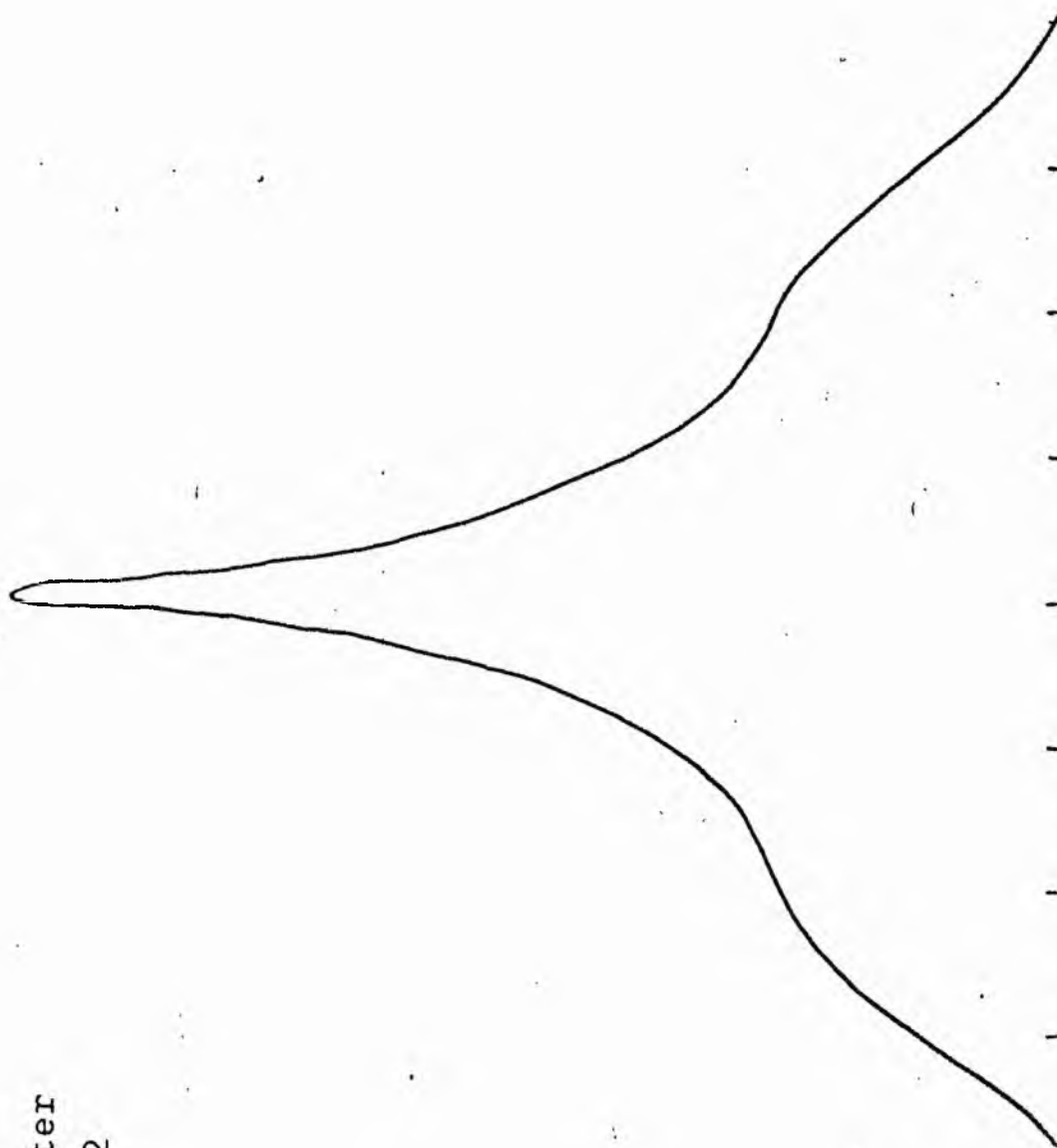
V-FILTER

Total luminosity	L_T	= 9.48
Total apparent magnitude	m_T	= 10.43
Apparent central surface brightness	μ_0	= 16.63
Major axis at threshold	$2a_m$	= 7.41
Minor axis at threshold	$2b_m$	= 4.48
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 4.95
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.94
Gradient of exponential component	$G(a)$	= -0.62
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.89
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.99
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.12
Axis ratio	b/a	= 0.65
Equivalent radius	r_1^*	= 0.11
Surface brightness	μ_1	= 17.76
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.23
Axis ratio	b/a	= 0.61
Equivalent radius	r_e^*	= 0.21
Surface brightness	μ_e	= 18.76
Mean surface brightness	μ_e'	= 9.03
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.63
Axis ratio	b/a	= 0.55
Equivalent radius	r_3^*	= 0.50
Surface brightness	μ_3	= 21.51
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.89 \\ = 2.42 \end{cases}$

NGC 4450
B-Filter
Axis 1

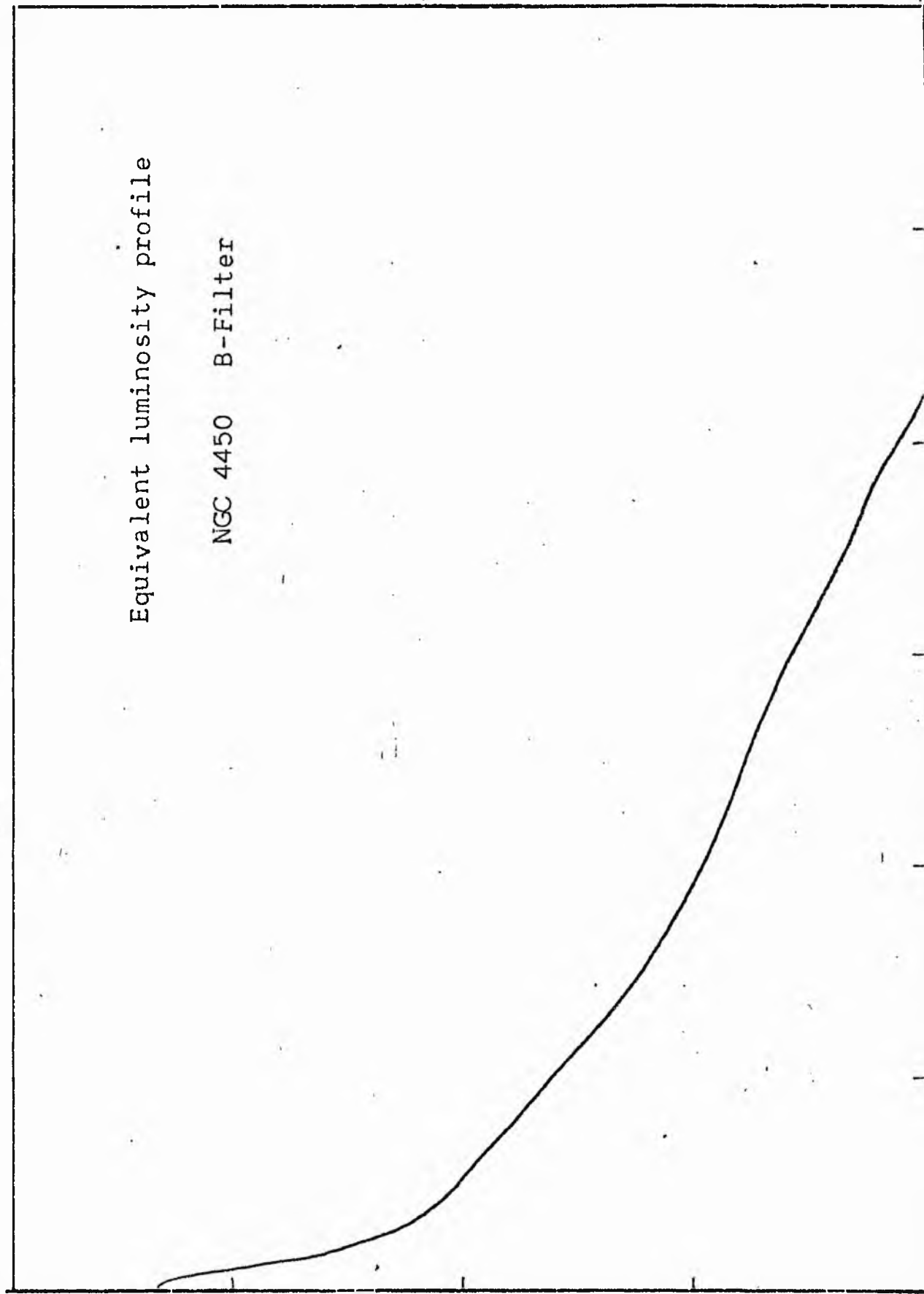


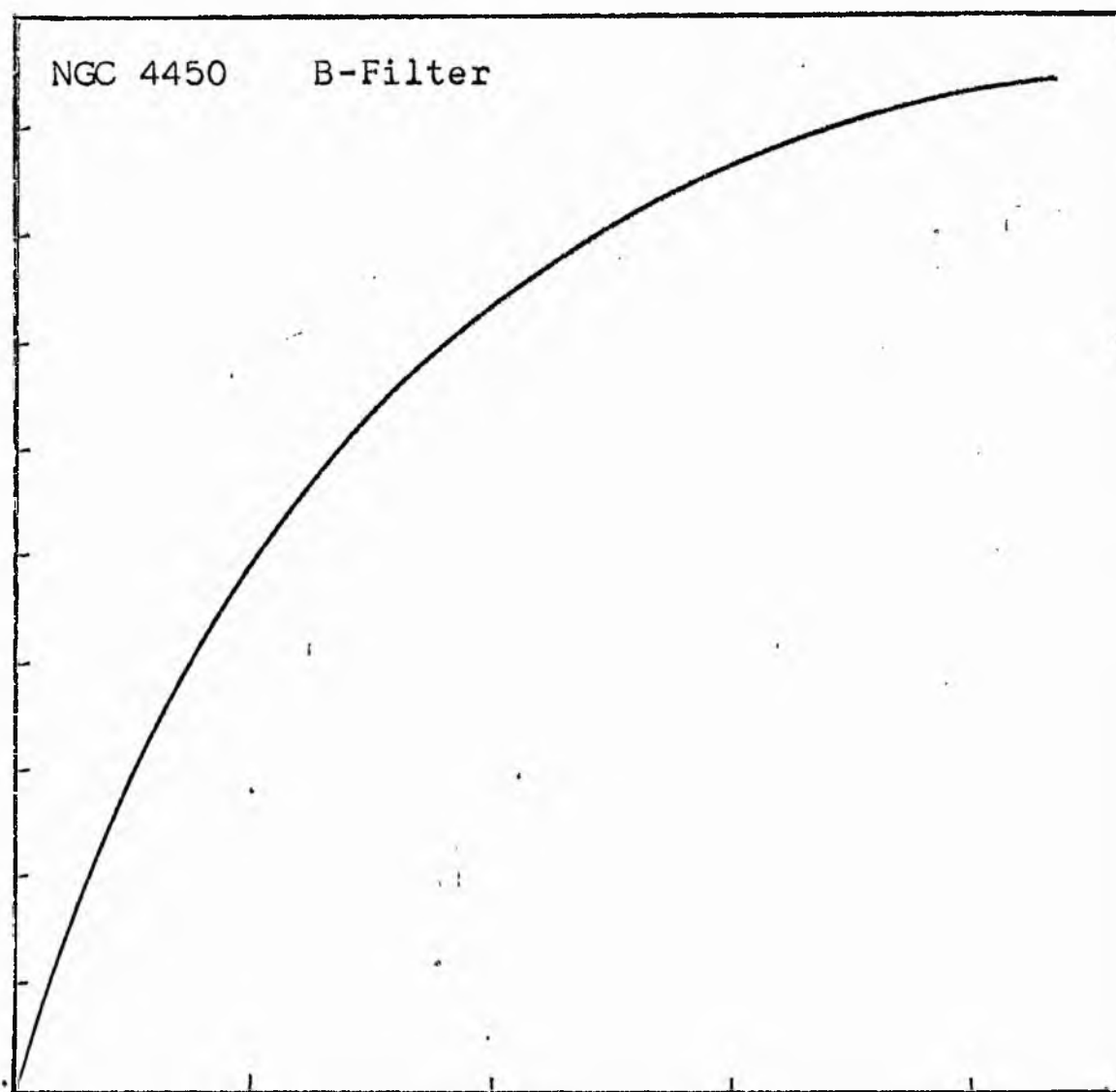
NGC 4450
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4450 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4450
B COLOUR

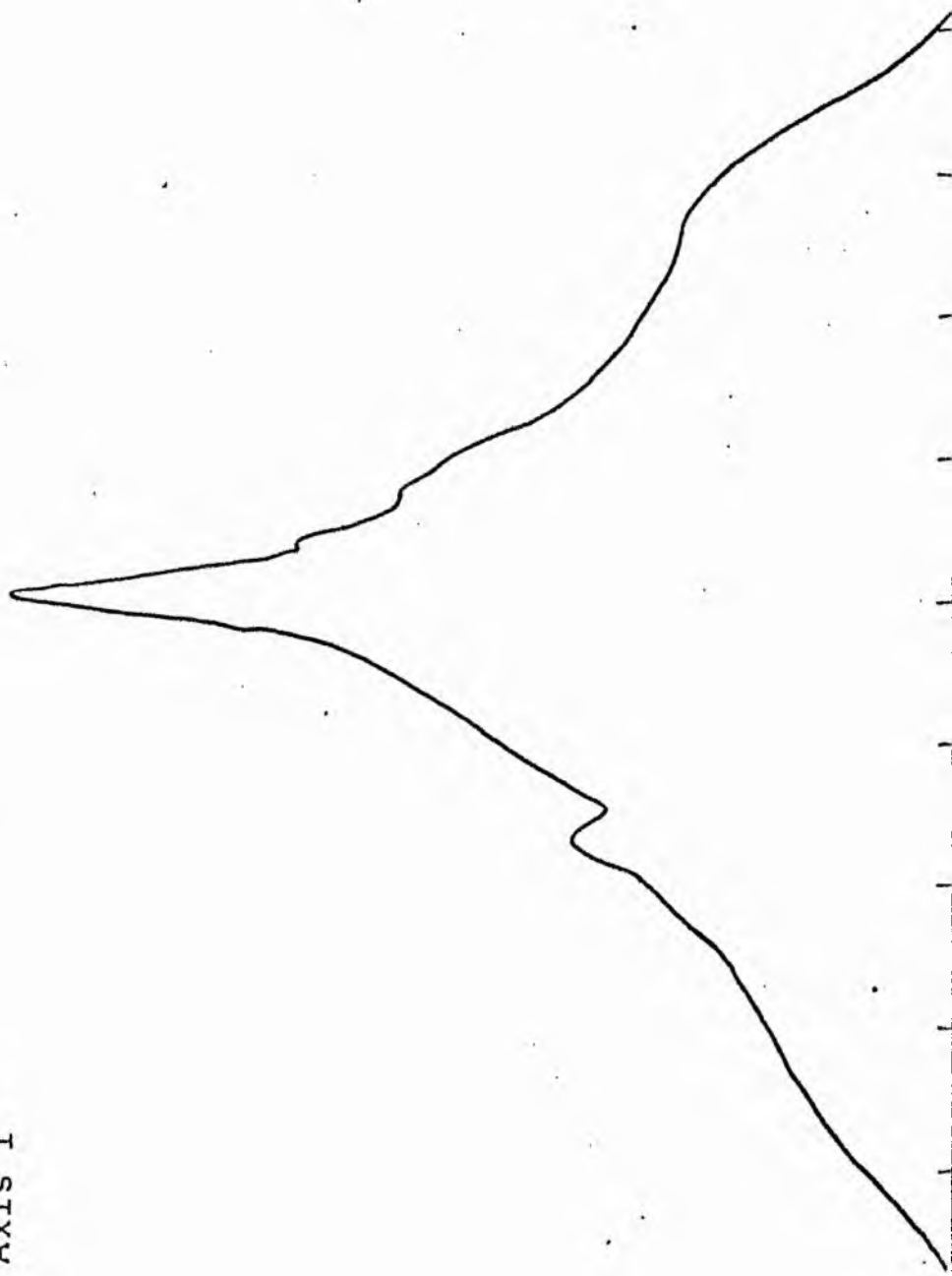
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.31	20.417		0.0	0.0			0.0	0.0	0.0	1.709	18.17
1.30	19.953	20.185	2.40	18.10	18.10	365.2583	365.26	0.02	0.05	1.699	18.20
1.20	15.849	17.901	3.81	45.60	27.51	492.4146	857.67	0.05	0.08	1.599	18.45
1.10	12.589	14.219	4.09	52.55	6.95	98.8100	956.48	0.05	0.08	1.499	18.70
1.00	10.000	11.295	5.16	83.65	31.09	351.1926	1307.68	0.07	0.10	1.399	18.95
0.90	7.943	8.972	6.25	122.72	34.07	350.5356	1658.21	0.09	0.13	1.299	19.20
0.80	6.310	7.126	6.70	141.03	18.31	130.4671	1788.68	0.10	0.14	1.199	19.45
0.70	5.012	5.661	7.66	184.33	43.31	245.1500	2033.84	0.11	0.16	1.099	19.70
0.60	3.981	4.496	9.47	201.74	97.41	437.9814	2471.82	0.14	0.19	0.999	19.95
0.50	3.162	3.572	10.43	341.76	60.02	214.3608	2686.18	0.15	0.21	0.899	20.20
0.40	2.512	2.837	11.50	415.48	73.72	209.1427	2895.32	0.16	0.23	0.799	20.45
0.30	1.995	2.254	15.01	767.80	292.33	658.7751	3554.10	0.20	0.30	0.699	20.70
0.20	1.585	1.790	17.77	992.03	284.23	508.7874	4062.88	0.23	0.36	0.599	20.95
0.10	1.259	1.422	18.99	1132.92	140.89	200.3344	4263.21	0.24	0.39	0.499	21.20
-0.00	1.000	1.129	25.60	2058.87	925.95	1045.8254	5309.04	0.30	0.52	0.399	21.45
-0.10	0.794	0.897	32.98	3417.05	1358.18	1218.5024	6527.54	0.37	0.67	0.299	21.70
-0.20	0.631	0.713	37.36	4384.93	967.89	689.7544	7217.29	0.41	0.76	0.199	21.95
-0.30	0.501	0.566	43.27	5881.98	1497.04	847.4319	8064.72	0.46	0.88	0.099	22.20
-0.40	0.398	0.450	49.32	7641.80	1759.82	791.2981	8856.02	0.50	1.00	-0.001	22.45
-0.50	0.316	0.357	55.43	9652.49	2010.69	718.1519	9574.17	0.54	1.13	-0.101	22.70
-0.60	0.251	0.284	61.60	11998.49	2346.00	665.5769	10239.74	0.58	1.25	-0.201	22.95
-0.70	0.200	0.225	68.70	14027.34	2828.85	637.5005	10877.24	0.61	1.39	-0.301	23.20
-0.80	0.158	0.179	75.19	17761.09	2933.76	525.1643	11402.41	0.64	1.53	-0.401	23.45
-0.90	0.126	0.142	82.11	21180.77	3419.67	486.2454	11888.65	0.67	1.67	-0.501	23.70
-1.00	0.100	0.113	96.70	29376.68	8195.91	925.6973	12814.34	0.72	1.96	-0.601	23.95
-1.10	0.079	0.090	108.70	37120.08	7743.40	694.7104	13509.05	0.76	2.21	-0.701	24.20
-1.20	0.063	0.071	125.70	49638.70	12518.62	892.1326	14401.18	0.81	2.55	-0.801	24.45
-1.30	0.050	0.057	136.80	58792.50	9153.80	518.1729	14919.35	0.84	2.78	-0.901	24.70
-1.40	0.040	0.045	147.50	68349.25	9556.75	429.7180	15349.07	0.87	2.99	-1.001	24.95
-1.50	0.032	0.036	153.60	74119.44	5770.19	206.0935	15555.16	0.88	3.12	-1.101	25.20
-1.60	0.025	0.028	164.20	84702.44	10583.00	300.2505	15855.41	0.90	3.33	-1.201	25.45
-1.70	0.020	0.023	177.20	98645.50	13943.06	314.2200	16169.62	0.91	3.60	-1.301	25.70
-1.80	0.016	0.018	188.50	111627.81	12982.31	232.3959	16402.02	0.93	3.83	-1.401	25.95
-1.90	0.013	0.014	196.30	121057.12	9429.31	134.0778	16536.09	0.93	3.98	-1.501	26.20
-2.00	0.010	0.011	210.30	138940.31	17883.19	201.9864	16738.08	0.95	4.27	-1.601	26.45
- ∞							17698.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4450

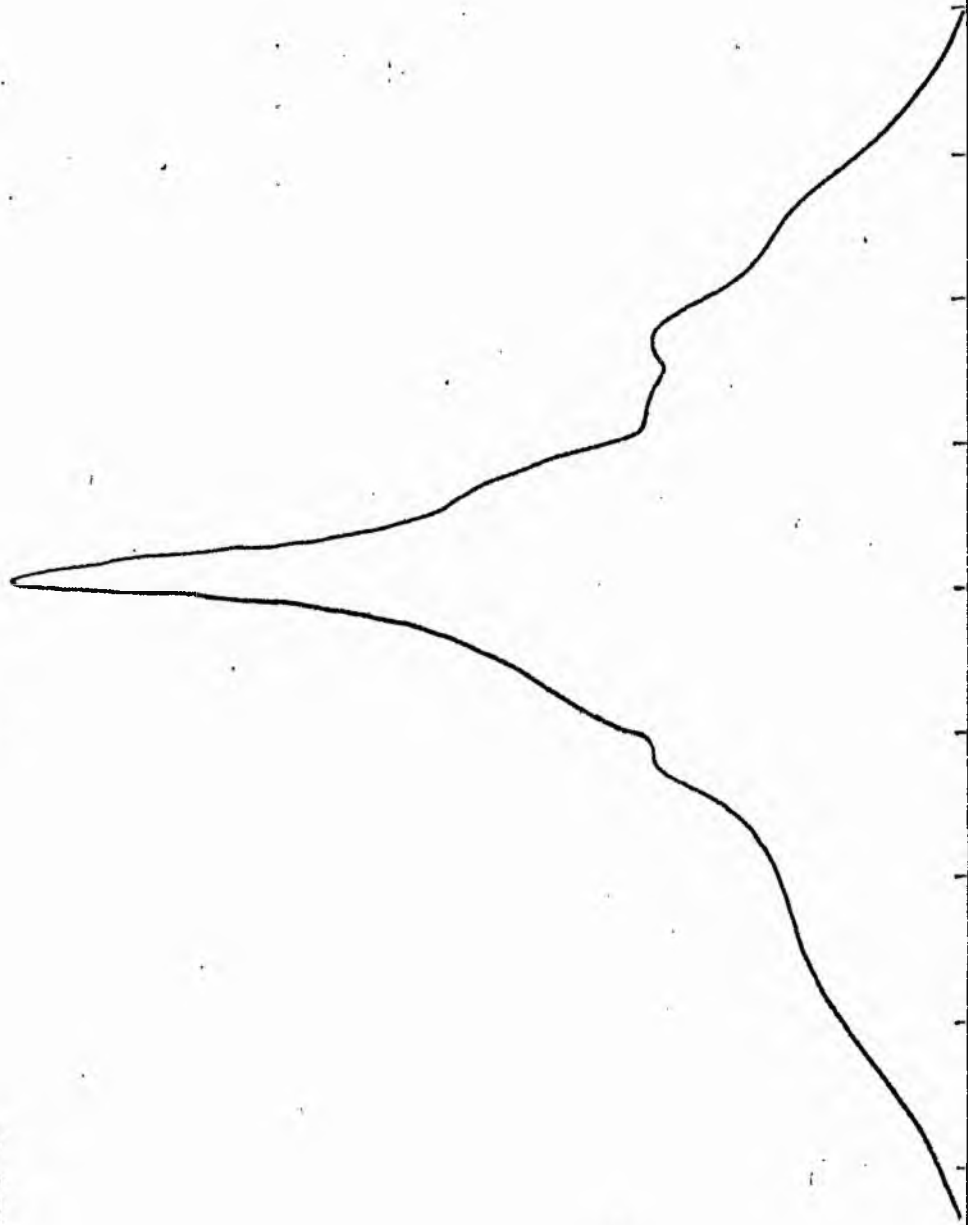
B-FILTER

Total luminosity	L_T	= 4.92
Total apparent magnitude	m_T	= 10.86
Apparent central surface brightness	μ_o	= 18.17
Major axis at threshold	$2a_m$	= 7.70
Minor axis at threshold	$2b_m$	= 6.05
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 5.28
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.87
Gradient of exponential component	$G(a)$	= -0.78
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.63
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.44
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.43
Axis ratio	b/a	= 0.81
Equivalent radius	r_1^*	= 0.33
Surface brightness	μ_1	= 21.25
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.94
Axis ratio	b/a	= 0.65
Equivalent radius	r_e^*	= 0.82
Surface brightness	μ_e	= 22.45
Mean surface brightness	μ_e'	= 12.41
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.15
Axis ratio	b/a	= 0.75
Equivalent radius	r_3^*	= 1.74
Surface brightness	μ_3	= 24.14
Concentration indices	$\left\{ \begin{array}{l} C_{21} \\ C_{32} \end{array} \right.$	= 2.47
		= 2.12

NGC 4450
V-Filter
Axis 1

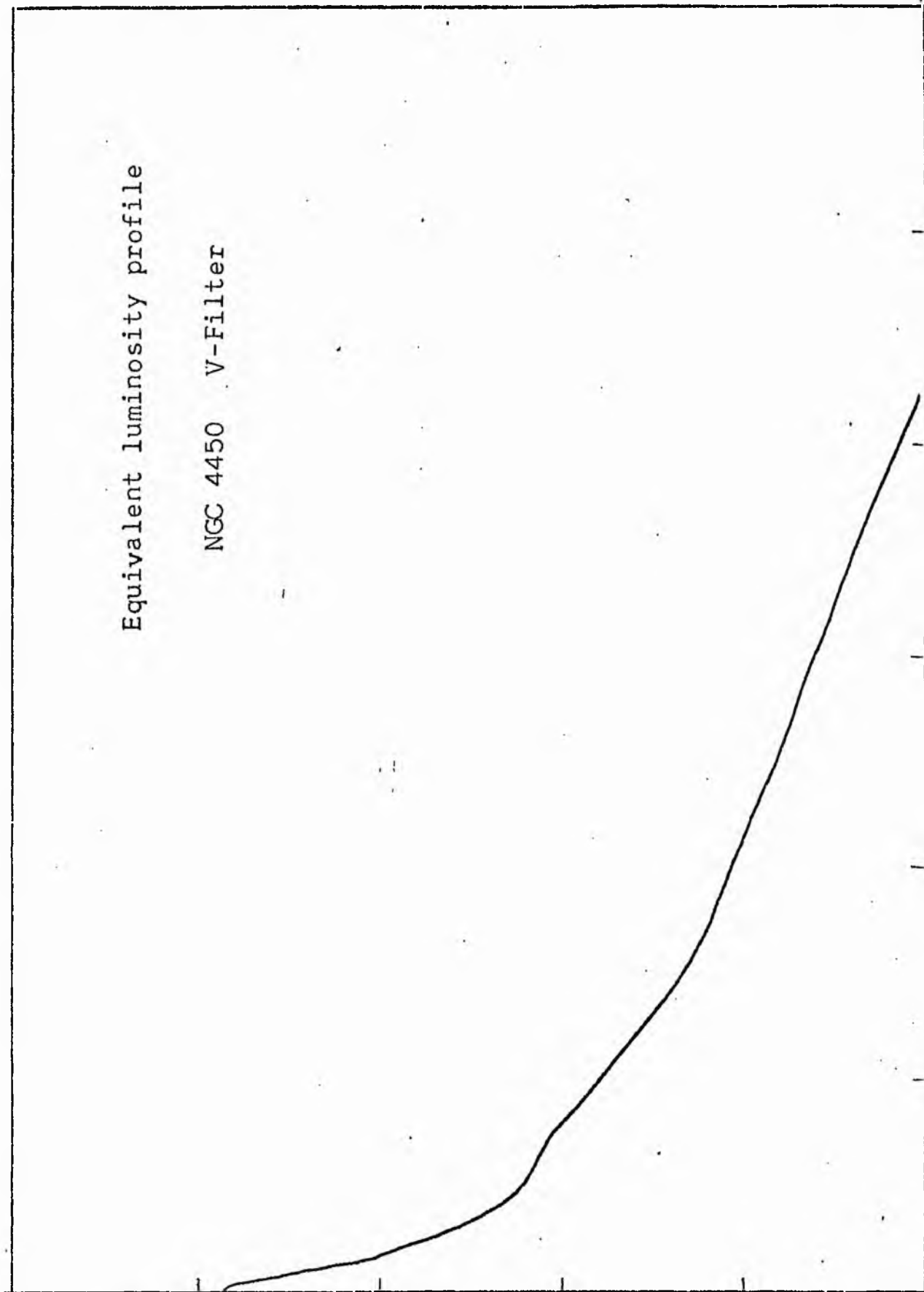


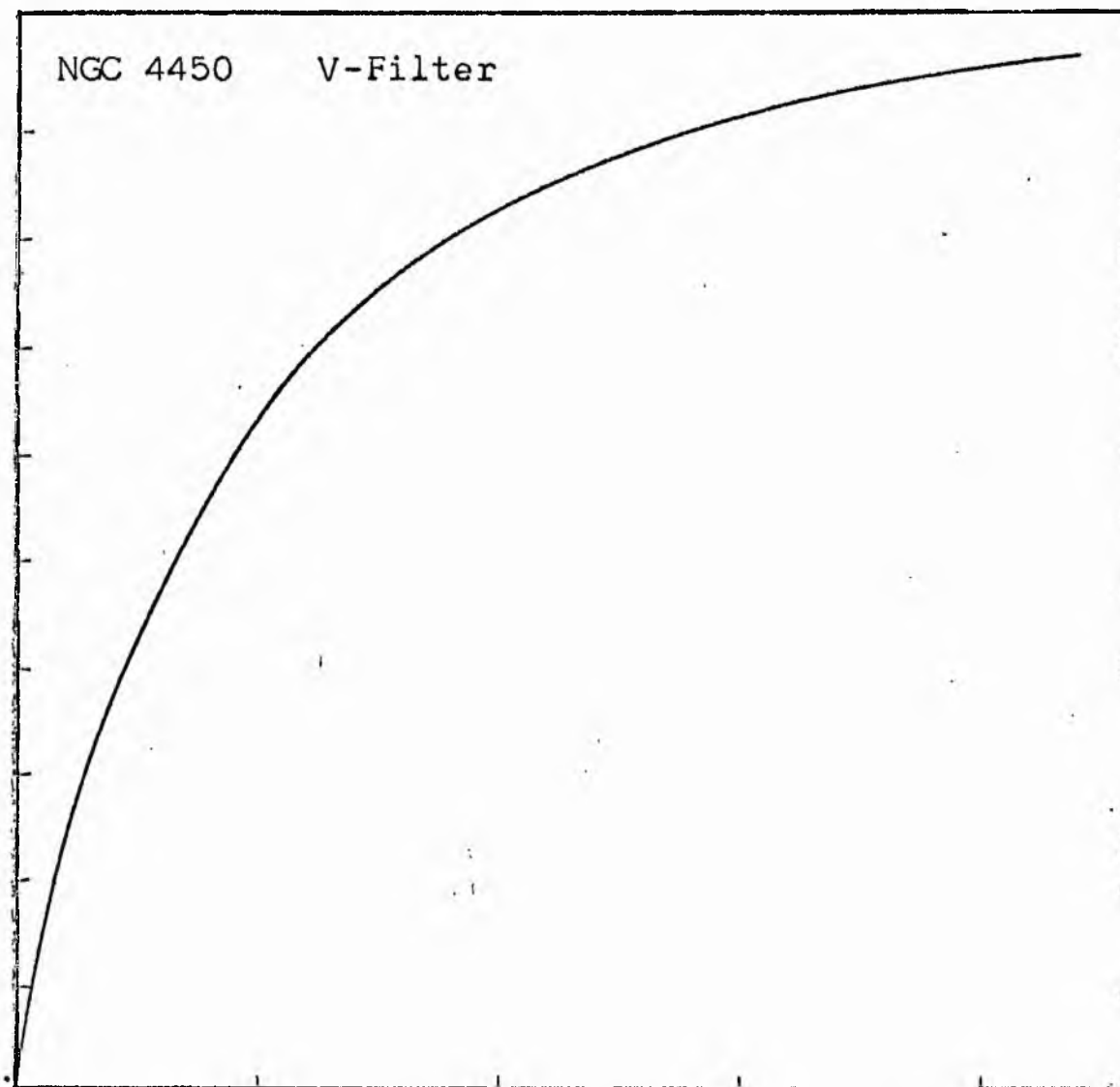
NGC 4450
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4450 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4450
V COLOUR

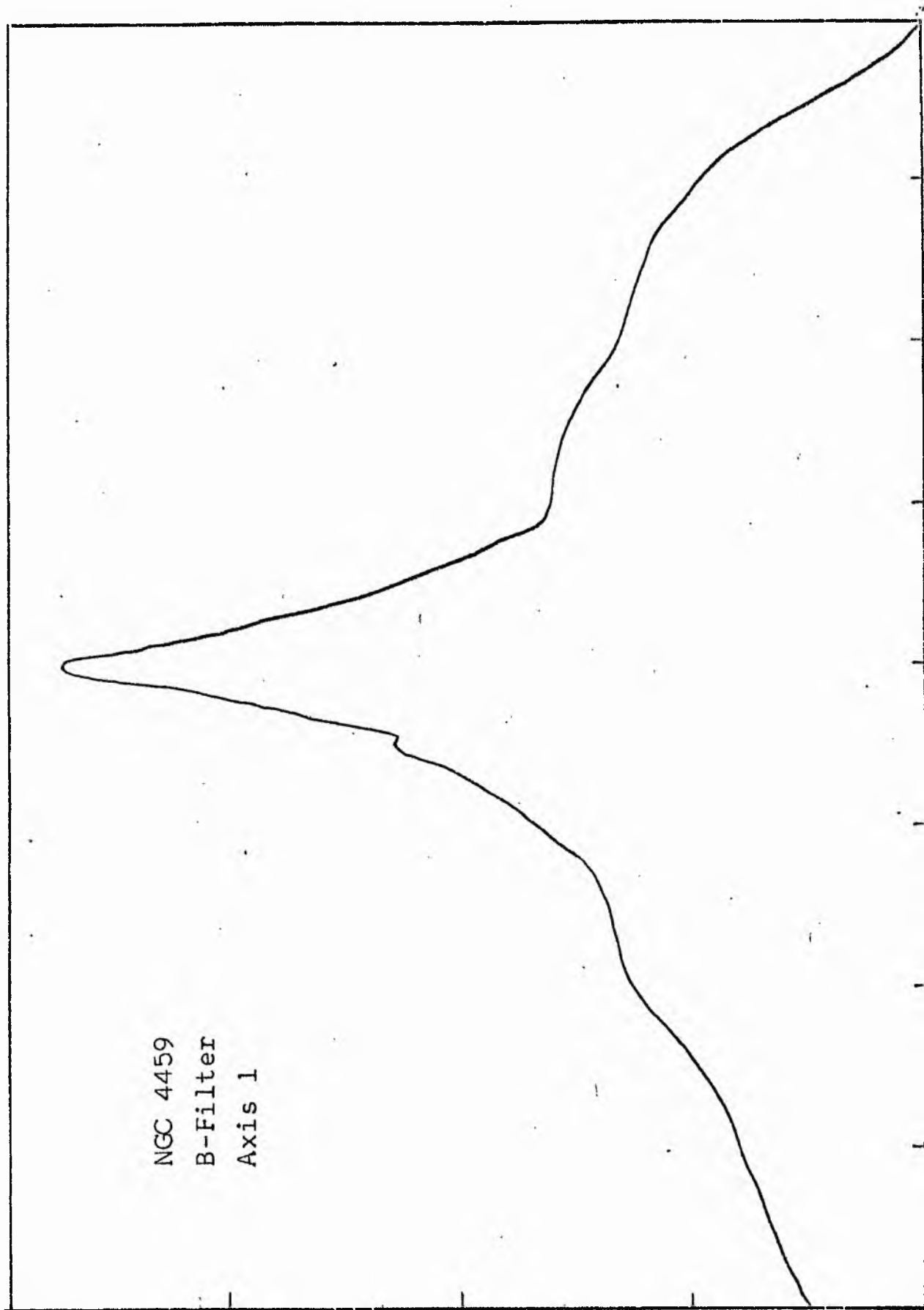
LOG I	I	\bar{I}	H	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.83	67.608		0.0	0.0						1.803	16.50
1.80	63.096	65.352	1.20	4.52	4.52	295.6448	295.64	0.01	0.03	1.773	16.58
1.70	50.119	56.607	2.70	22.90	18.38	1040.3425	1335.99	0.05	0.07	1.673	16.83
1.60	39.811	44.965	3.50	38.48	15.58	700.6516	2036.64	0.08	0.10	1.573	17.08
1.50	31.623	35.717	4.20	55.42	16.93	604.7939	2641.43	0.10	0.12	1.473	17.33
1.40	25.119	28.371	4.60	66.48	11.06	313.7344	2955.17	0.11	0.13	1.373	17.58
1.30	19.952	22.536	5.50	95.03	28.56	643.5515	3598.72	0.14	0.15	1.273	17.83
1.20	15.849	17.901	6.40	128.68	33.65	602.2930	4201.01	0.16	0.18	1.173	18.08
1.10	12.589	14.219	7.00	153.94	25.26	359.1492	4560.16	0.17	0.19	1.073	18.33
1.00	10.000	11.295	8.50	226.98	73.04	824.9751	5385.13	0.20	0.23	0.973	18.58
0.90	7.943	8.972	9.60	289.53	62.55	561.1621	5946.29	0.23	0.27	0.873	18.83
0.80	6.310	7.126	10.23	328.78	39.25	279.6924	6225.98	0.24	0.28	0.773	19.08
0.70	5.012	5.661	11.43	410.43	81.66	462.2295	6688.21	0.25	0.32	0.673	19.33
0.60	3.981	4.496	13.19	546.56	136.13	612.0957	7300.30	0.28	0.36	0.573	19.58
0.50	3.162	3.572	15.45	749.91	203.34	726.2727	8026.57	0.31	0.43	0.473	19.83
0.40	2.512	2.837	17.48	959.91	210.01	595.8049	8622.38	0.33	0.48	0.373	20.08
0.30	1.995	2.254	19.05	1140.09	180.18	406.0371	9028.41	0.34	0.53	0.273	20.33
0.20	1.585	1.790	23.50	1734.94	594.85	1064.8284	10093.24	0.38	0.65	0.173	20.58
0.10	1.259	1.422	29.89	2806.74	1071.79	1523.9792	11617.22	0.44	0.83	0.073	20.83
-0.00	1.000	1.129	38.32	4613.18	1806.44	2040.2957	13657.51	0.52	1.06	-0.027	21.08
-0.10	0.794	0.897	45.25	6432.61	1819.43	1632.3118	15289.82	0.58	1.25	-0.127	21.33
-0.20	0.631	0.713	51.50	8332.29	1899.68	1353.7864	16643.61	0.63	1.42	-0.227	21.58
-0.30	0.501	0.566	54.70	9399.93	1067.64	604.3550	17247.96	0.66	1.51	-0.327	21.83
-0.40	0.398	0.450	58.20	10641.32	1241.40	558.1875	17806.14	0.68	1.61	-0.427	22.08
-0.50	0.316	0.357	62.30	12193.43	1552.10	554.3564	18360.50	0.70	1.72	-0.527	22.33
-0.60	0.251	0.284	66.78	14010.14	1816.71	515.4136	18875.91	0.72	1.84	-0.627	22.58
-0.70	0.200	0.225	74.24	17315.12	3304.98	744.7971	19620.71	0.75	2.05	-0.727	22.83
-0.80	0.158	0.179	82.50	21382.46	4067.34	728.0815	20348.79	0.77	2.28	-0.827	23.08
-0.90	0.126	0.142	91.55	26330.94	4948.48	703.6240	21052.41	0.80	2.53	-0.927	23.33
-1.00	0.100	0.113	101.40	32301.72	5970.78	674.3738	21726.78	0.83	2.80	-1.027	23.58
-1.10	0.079	0.090	108.70	37120.08	4818.36	432.2839	22159.06	0.84	3.00	-1.127	23.83
-1.20	0.063	0.071	119.80	45088.25	7968.17	567.8445	22726.90	0.86	3.31	-1.227	24.08
-1.30	0.050	0.057	134.60	56916.73	1828.48	669.5764	23396.48	0.89	3.72	-1.327	24.33
-1.40	0.040	0.045	147.80	68627.56	11710.83	526.5742	23923.05	0.91	4.08	-1.427	24.58
-1.50	0.032	0.036	156.00	76453.75	7826.19	279.5264	24202.57	0.92	4.31	-1.527	24.83
-1.60	0.025	0.028	164.10	84599.31	8145.56	231.0969	24433.67	0.93	4.53	-1.627	25.08
-1.70	0.020	0.023	173.20	94242.25	9642.94	217.3117	24650.98	0.94	4.78	-1.727	25.33
-1.80	0.016	0.018	187.40	110328.81	16086.56	287.9636	24938.94	0.95	5.17	-1.827	25.58
-1.90	0.013	0.014	197.80	122914.25	12585.44	178.9548	25117.89	0.95	5.46	-1.927	25.83
-2.00	0.010	0.011	211.00	139866.81	16952.56	191.4745	25309.36	0.96	5.83	-2.027	26.08
							26309.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4450

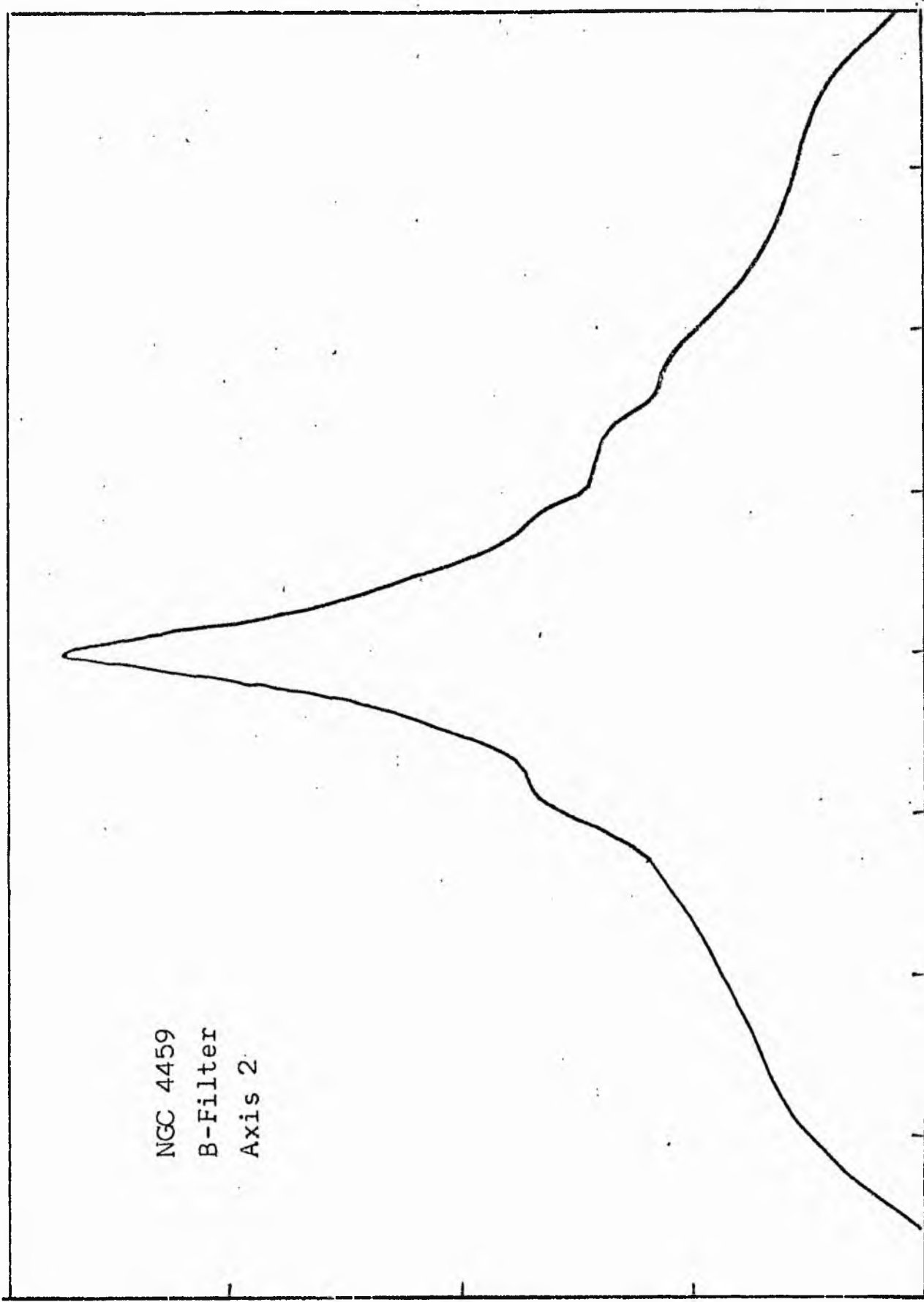
V-FILTER

Total luminosity	L_T	= 7.31
Total apparent magnitude	m_T	= 10.03
Apparent central surface brightness	μ_o	= 16.50
Major axis at threshold	$2a_m$	= 7.61
Minor axis at threshold	$2b_m$	= 6.32
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.20
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.93
Gradient of exponential component	$G(a)$	= -0.71
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.68
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.37
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.32
Axis ratio	b/a	= 0.53
Equivalent radius	r_1^*	= 0.19
Surface brightness	μ_1	= 19.33
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.76
Axis ratio	b/a	= 0.51
Equivalent radius	r_e^*	= 0.60
Surface brightness	μ_e	= 21.00
Mean surface brightness	μ_e'	= 10.93
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.61
Axis ratio	b/a	= 0.63
Equivalent radius	r_3^*	= 1.26
Surface brightness	μ_3	= 22.83
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.25 \\ = 2.08 \end{cases}$

NGC 4459
B-Filter
Axis 1

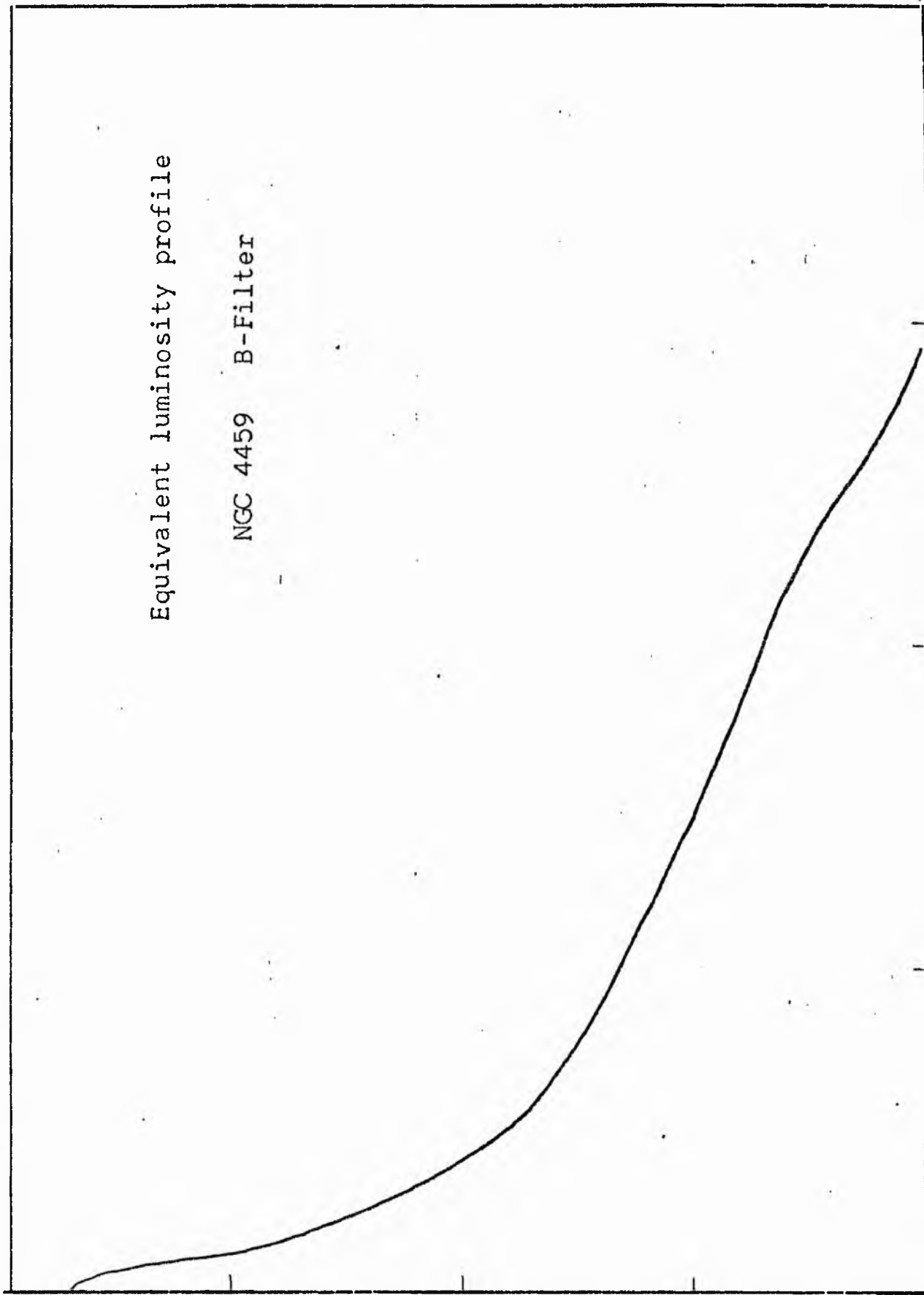


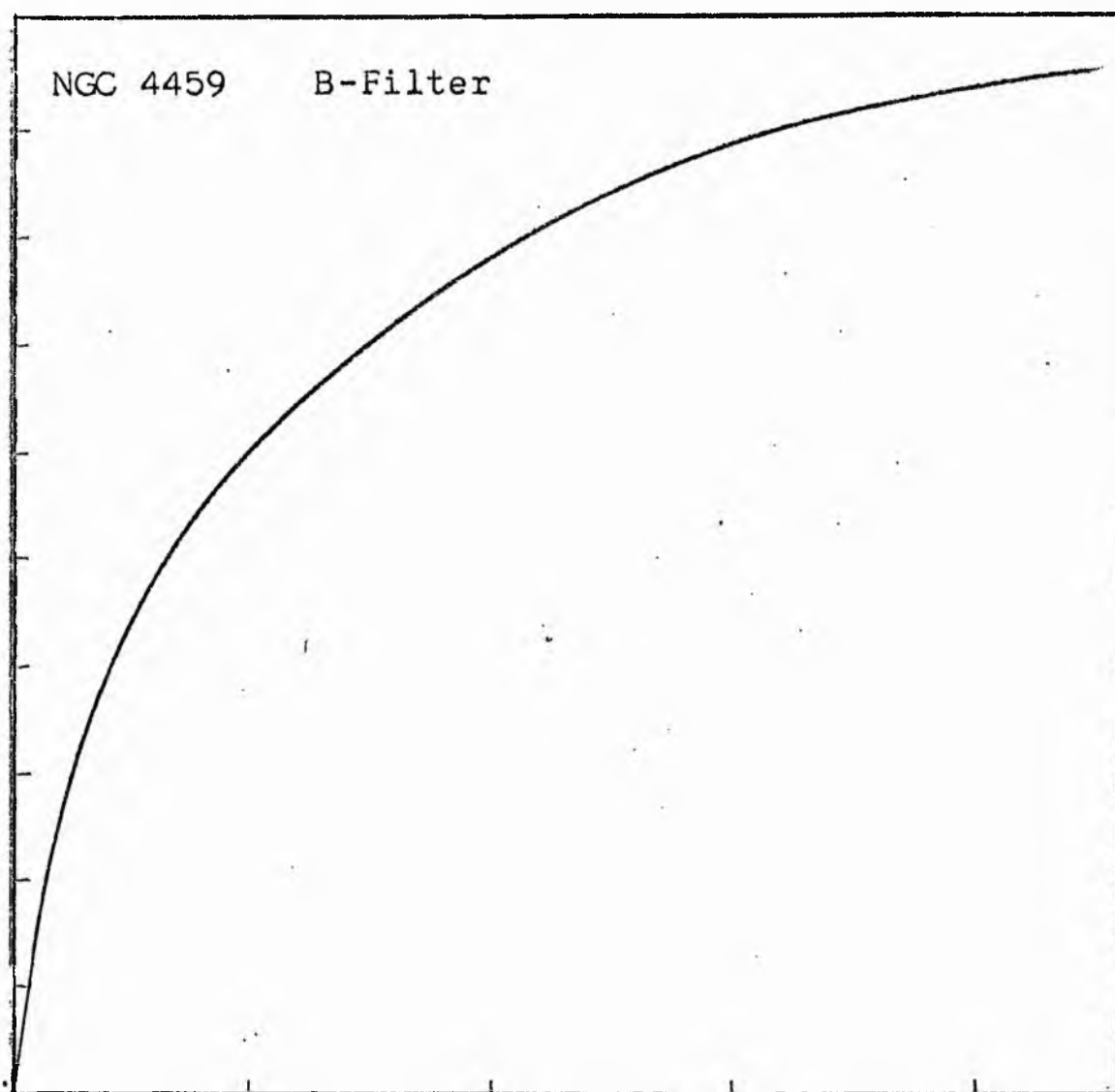
NGC 4459
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4459 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4459
B COLOUR

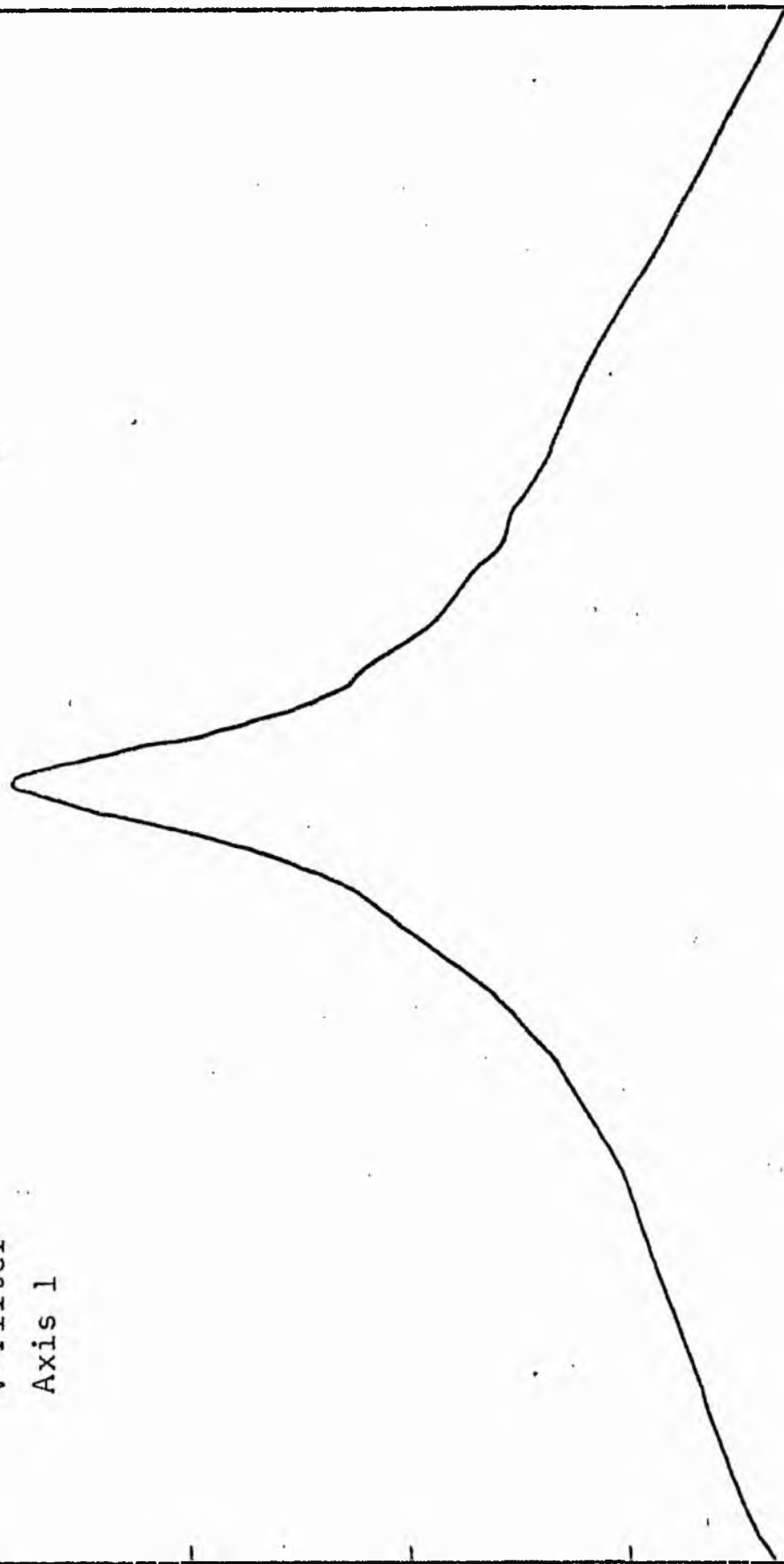
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.68	47.863		0.0	0.0			0.0	0.0	0.0	1.722	18.32
		43.837			23.24	1018.8870					
1.60	39.811		2.72	23.24			1018.89	0.04	0.09	1.642	18.52
		35.717			42.94	1533.8381					
1.50	31.623		4.59	64.19			2552.73	0.09	0.14	1.542	18.77
		28.371			20.05	818.3777					
1.40	25.119		5.50	95.03			3371.10	0.12	0.17	1.442	19.02
		22.536			26.90	606.2336					
1.30	19.953		6.23	121.93			3977.34	0.14	0.20	1.342	19.27
		17.901			43.65	781.3918					
1.20	15.849		7.26	165.59			4758.73	0.17	0.23	1.242	19.52
		14.219			66.23	941.6731					
1.10	12.589		8.59	231.81			5700.43	0.20	0.27	1.142	19.77
		11.295			78.59	887.6223					
1.00	10.000		9.94	310.40			6588.02	0.23	0.31	1.042	20.02
		8.972			71.12	638.0181					
0.90	7.943		11.02	381.52			7226.04	0.26	0.35	0.942	20.27
		7.126			100.76	718.0266					
0.80	6.310		12.39	482.27			7944.06	0.28	0.39	0.842	20.52
		5.661			172.79	978.1245					
0.70	5.012		14.44	655.66			8922.18	0.32	0.45	0.742	20.77
		4.496			147.17	661.7578					
0.60	3.981		15.98	802.24			9583.94	0.34	0.50	0.642	21.02
		3.572			195.38	697.8318					
0.50	3.162		17.82	997.62			10281.77	0.36	0.56	0.542	21.27
		2.837			251.49	713.4875					
0.40	2.512		19.94	1249.11			10995.25	0.39	0.63	0.442	21.52
		2.254			378.29	852.5063					
0.30	1.995		22.76	1627.40			11847.75	0.42	0.71	0.342	21.77
		1.790			545.61	976.6729					
0.20	1.585		26.30	2173.01			12824.43	0.45	0.83	0.242	22.02
		1.422			560.97	797.6375					
0.10	1.259		29.50	2733.97			13622.06	0.48	0.93	0.142	22.27
		1.129			296.80	335.2190					
-0.00	1.000		31.06	3030.77			13957.28	0.49	0.98	0.042	22.52
		0.897			765.45	690.3230					
-0.10	0.794		34.78	3800.22			14647.60	0.52	1.09	-0.058	22.77
		0.713			1091.52	777.8599					
-0.20	0.631		39.46	4891.74			15425.46	0.55	1.24	-0.158	23.02
		0.566			1472.81	833.7122					
-0.30	0.501		45.01	6364.55			16259.17	0.58	1.41	-0.258	23.27
		0.450			1974.21	887.6926					
-0.40	0.398		51.52	8338.76			17146.86	0.61	1.62	-0.358	23.52
		0.357			2582.30	922.3083					
-0.50	0.316		58.96	10921.05			18069.17	0.64	1.85	-0.458	23.77
		0.284			4384.89	1244.0254					
-0.60	0.251		69.80	15305.95			19313.19	0.68	2.19	-0.558	24.02
		0.225			5407.95	1218.7158					
-0.70	0.200		81.20	20713.09			20531.91	0.73	2.55	-0.658	24.27
		0.179			4902.92	877.6575					
-0.80	0.158		90.30	25616.82			21409.56	0.76	2.84	-0.758	24.52
		0.142			6684.91	950.5293					
-0.90	0.126		101.40	32301.72			22360.09	0.79	3.18	-0.858	24.77
		0.113			3464.96	391.3525					
-1.00	0.100		106.70	35766.68			22751.44	0.81	3.35	-0.958	25.02
		0.090			8273.91	742.3042					
-1.10	0.079		118.40	44040.59			23493.74	0.83	3.72	-1.058	25.27
		0.071			11698.30	833.6699					
-1.20	0.063		133.20	55738.89			24327.41	0.86	4.18	-1.158	25.52
		0.057			12610.36	713.8379					
-1.30	0.050		147.50	68349.25			25041.25	0.89	4.63	-1.258	25.77
		0.045			15941.00	716.7842					
-1.40	0.040		163.80	84290.25			25758.03	0.91	5.14	-1.358	26.02
		0.036			9083.37	324.4297					
-1.50	0.032		172.40	93373.62			26082.46	0.92	5.41	-1.458	26.27
		0.028			8753.50	248.3454					
-1.60	0.025		180.30	102127.12			26330.80	0.93	5.66	-1.558	26.52
		0.023			12121.56	273.1702					
-1.70	0.020		190.70	114248.69			26603.97	0.94	5.99	-1.658	26.77
		0.018			10662.12	190.8618					
-1.80	0.016		199.40	124910.81			26794.83	0.95	6.26	-1.758	27.02
		0.014			16684.75	237.2441					
-1.90	0.013		212.30	141595.56			27032.07	0.96	6.67	-1.858	27.27
		0.011			12816.50	144.7591					
-2.00	0.010		221.70	154412.06			27176.83	0.96	6.96	-1.958	27.52
-∞							28226.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4459

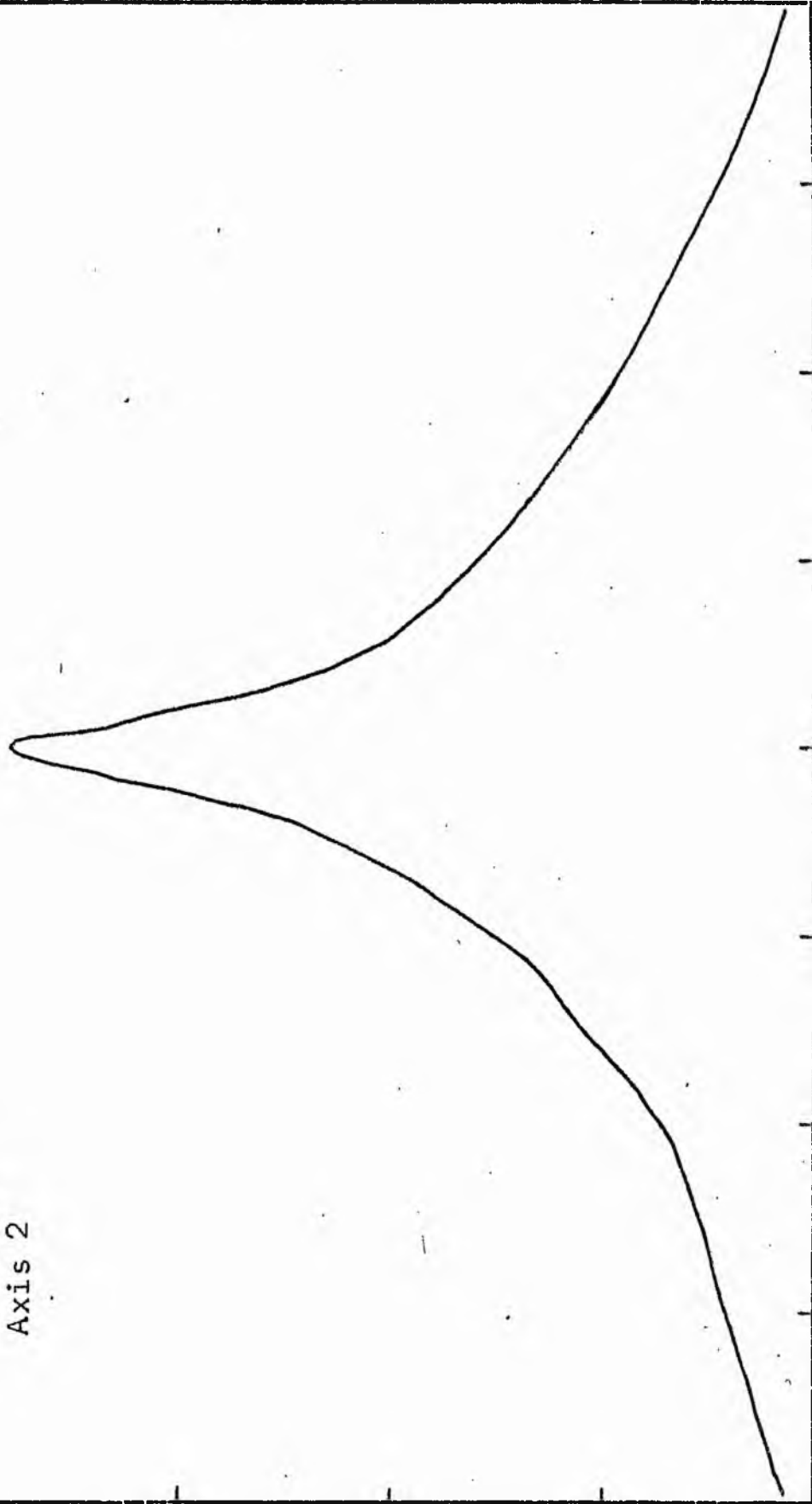
B-FILTER

Total luminosity	L_T	= 7.84
Total apparent magnitude	m_T	= 11.39
Apparent central surface brightness	μ_0	= 18.32
Major axis at threshold	$2a_m$	= 7.97
Minor axis at threshold	$2b_m$	= 6.18
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.43
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.81
Gradient of exponential component	$G(a)$	= -0.69
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.55
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.30
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 0.69
Equivalent radius	r_1^*	= 0.18
Surface brightness	μ_1	= 20.19
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.58
Axis ratio	b/a	= 0.69
Equivalent radius	r_e^*	= 0.53
Surface brightness	μ_e	= 22.60
Mean surface brightness	μ_e'	= 11.99
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.93
Axis ratio	b/a	= 0.57
Equivalent radius	r_3^*	= 1.46
Surface brightness	μ_3	= 24.44
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.97 \\ = 2.75 \end{cases}$

NGC 4459
V-Filter
Axis 1



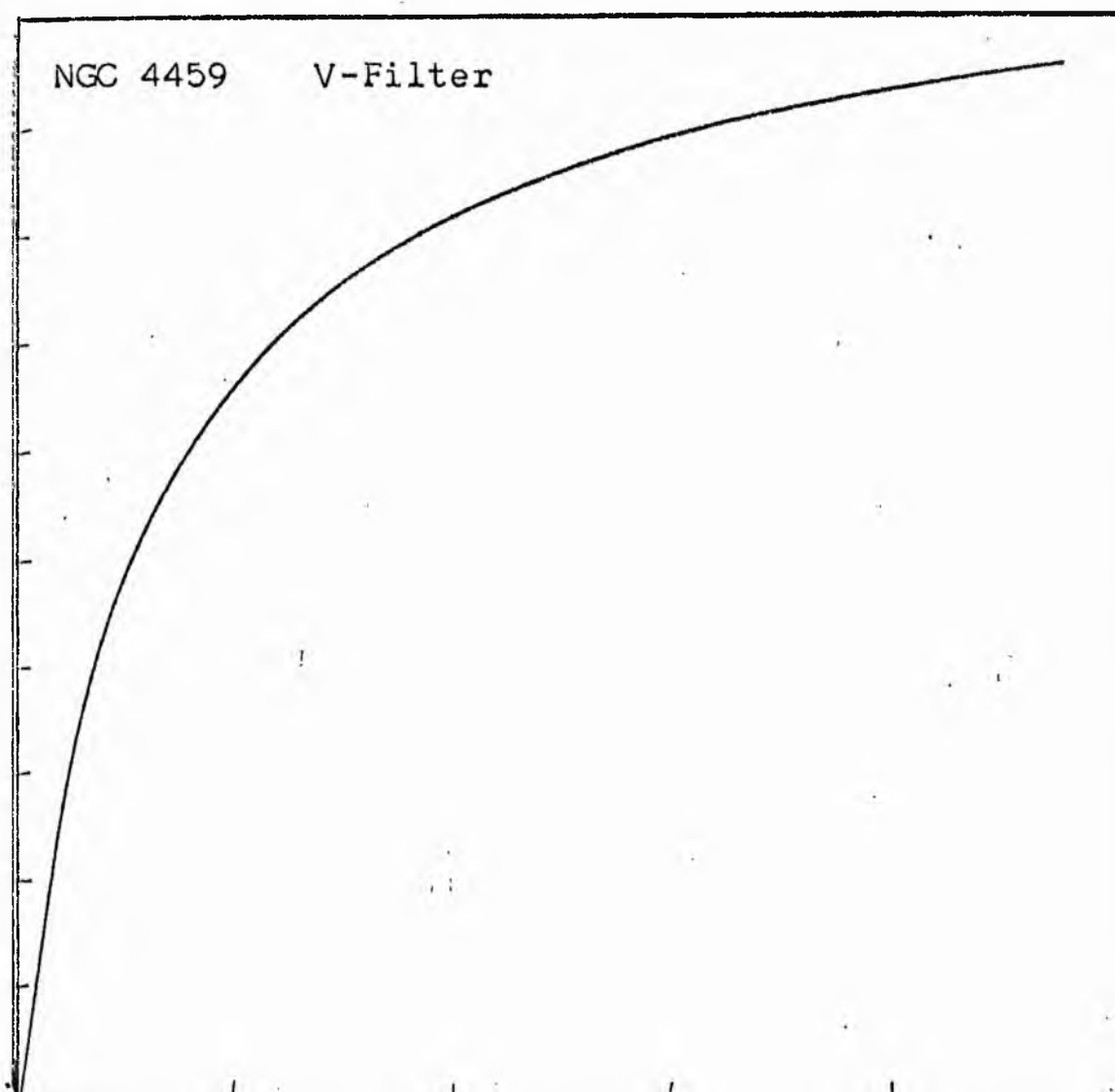
NGC 4459
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4459 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4459
V COLOUR

LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LUG J	μ
1.78	60.256		0.0	0.0			0.0	0.0	0.0	1.535	17.30
		55.187			5.31	293.0049					
1.70	50.119		1.30	5.31			293.00	0.01	0.05	1.455	17.50
		44.965			21.11	949.2708					
1.60	39.811		2.90	26.42			1242.28	0.04	0.11	1.355	17.75
		35.717			29.00	1035.6716					
1.50	31.623		4.20	55.42			2277.95	0.08	0.16	1.255	18.00
		28.371			39.62	1123.9216					
1.40	25.119		5.50	95.03			3401.87	0.12	0.20	1.155	18.25
		22.536			48.53	1093.6560					
1.30	19.953		6.76	143.56			4495.52	0.15	0.25	1.055	18.50
		17.901			68.71	1229.9409					
1.20	15.849		8.22	212.27			5725.46	0.20	0.31	0.955	18.75
		14.219			84.54	1202.0774					
1.10	12.589		9.72	296.81			6927.54	0.24	0.36	0.855	19.00
		11.295			68.27	771.0435					
1.00	10.000		10.78	365.08			7698.58	0.26	0.40	0.755	19.25
		8.972			81.30	729.3723					
0.90	7.943		11.92	446.38			8427.95	0.29	0.44	0.655	19.50
		7.126			155.38	1107.3020					
0.80	6.310		13.84	601.76			9535.25	0.33	0.51	0.555	19.75
		5.661			177.55	1005.0735					
0.70	5.012		15.75	779.31			10540.32	0.36	0.58	0.455	20.00
		4.496			177.31	797.2671					
0.60	3.981		17.45	956.62			11337.59	0.39	0.65	0.355	20.25
		3.572			169.15	634.1472					
0.50	3.162		18.93	1125.77			11941.73	0.41	0.70	0.255	20.50
		2.837			318.33	903.1304					
0.40	2.512		21.44	1444.10			12844.86	0.44	0.80	0.155	20.75
		2.254			385.11	867.8638					
0.30	1.995		24.13	1829.21			13712.72	0.47	0.90	0.055	21.00
		1.790			769.32	1377.1282					
0.20	1.585		28.76	2598.53			15089.85	0.52	1.07	-0.045	21.25
		1.422			592.38	842.3013					
0.10	1.259		31.87	3190.91			15932.15	0.55	1.18	-0.145	21.50
		1.129			513.77	580.2825					
-0.00	1.000		34.34	3704.68			16512.43	0.57	1.27	-0.245	21.75
		0.897			968.89	869.2522					
-0.10	0.794		38.57	4673.57			17381.68	0.60	1.43	-0.345	22.00
		0.713			1262.91	899.9966					
-0.20	0.631		43.47	5936.48			18281.68	0.63	1.61	-0.445	22.25
		0.566			1621.88	918.0984					
-0.30	0.501		49.05	7558.36			19199.77	0.66	1.82	-0.545	22.50
		0.450			2052.39	922.8457					
-0.40	0.398		55.31	9610.75			20122.62	0.69	2.05	-0.645	22.75
		0.357			1876.86	670.3501					
-0.50	0.316		60.47	11487.61			20792.96	0.71	2.24	-0.745	23.00
		0.284			1879.71	533.2871					
-0.60	0.251		65.23	13367.32			21326.25	0.73	2.42	-0.845	23.25
		0.225			2114.57	476.5332					
-0.70	0.200		70.20	15481.89			21802.78	0.75	2.60	-0.945	23.50
		0.179			3047.96	545.6064					
-0.80	0.158		76.80	18529.85			22348.39	0.77	2.85	-1.045	23.75
		0.142			4922.00	699.8616					
-0.90	0.126		86.40	23451.86			23048.25	0.79	3.21	-1.145	24.00
		0.113			7152.56	807.8521					
-1.00	0.100		98.70	30604.41			23856.10	0.82	3.66	-1.245	24.25
		0.090			5633.10	575.3804					
-1.10	0.079		107.40	36237.51			24361.48	0.84	3.99	-1.345	24.50
		0.071			7654.43	545.4068					
-1.20	0.063		118.20	43891.94			24906.96	0.85	4.39	-1.445	24.75
		0.057			12014.45	680.1050					
-1.30	0.050		133.40	55906.39			25587.06	0.88	4.95	-1.545	25.00
		0.045			8066.78	362.7212					
-1.40	0.040		142.70	63973.16			25949.78	0.89	5.30	-1.645	25.25
		0.036			12774.90	456.2795					
-1.50	0.032		156.30	76748.06			26406.06	0.91	5.80	-1.745	25.50
		0.028			21008.75	596.0388					
-1.60	0.025		176.40	97756.81			27002.09	0.93	6.55	-1.845	25.75
		0.023			11280.56	254.2176					
-1.70	0.020		186.30	109037.37			27256.31	0.94	6.91	-1.945	26.00
		0.018			21831.31	390.8003					
-1.80	0.016		204.10	130868.69			27647.11	0.95	7.57	-2.045	26.25
		0.014			21599.31	377.1252					
-1.90	0.013		220.30	152468.00			27954.23	0.96	8.17	-2.145	26.50
		0.011			16625.06	187.7758					
-2.00	0.010		232.00	169093.06			28142.00	0.97	8.61	-2.245	26.75
-∞							29142.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NCC 4459
V-FILTER

Total luminosity	L_T	= 8.09
Total apparent magnitude	m_T	= 10.59
Apparent central surface brightness	μ_0	= 17.30
Major axis at threshold	$2a_m$	= 7.71
Minor axis at threshold	$2b_m$	= 6.99
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.37
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.88
Gradient of exponential component	$G(a)$	= -0.65
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.56
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.21
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.23
Axis ratio	b/a	= 0.92
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 19.12
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.42
Axis ratio	b/a	= 0.97
Equivalent radius	r_e^*	= 0.45
Surface brightness	μ_e	= 21.15
Mean surface brightness	μ_e'	= 10.84
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.43
Axis ratio	b/a	= 0.69
Equivalent radius	r_3^*	= 1.18
Surface brightness	μ_3	= 23.50
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.64 \\ = 2.63 \end{cases}$

PHOTOGRAPHIC SURFACE PHOTOMETRY OF GALAXIES

IN THE VIRGO CLUSTER

by

Christopher W. Fraser

VOLUME THREE



Th 5841

PHOTOMETRIC DATA FOR GALAXIES

NGC 4661 to NGC 4762

EXPLANATORY NOTES FOR USE WITH THE PHOTOMETRIC DATA

The data for each galaxy occupy twelve pages, divided equally between the B and the V colours. The galaxies are listed in increasing order of NGC identification numbers and each page may be described as follows:-

1. Luminosity profile of major axis. The y-axis denotes $\log I$, with the value of $\log I = -2.00$ at the origin. Each division on the y-axis represents a value of $+1.00$ in $\log I$. For the x-axis, which denotes distance r , the centre of the galaxy is plotted at the centre of the axis. Each division represents 50 arc sec.
2. Luminosity profile of minor axis. The units for the x and y axes are as before.
3. Equivalent luminosity profile. The y-axis denotes increasing $\log I$, each division representing an integer value from a minimum at the origin of $\log I = -2.00$. The distance r is given on the x-axis, from zero at the origin and increasing by 50 arc sec per division.
4. Relative integrated luminosity curve $k(r)$ is plotted as ordinate, each positive division in y

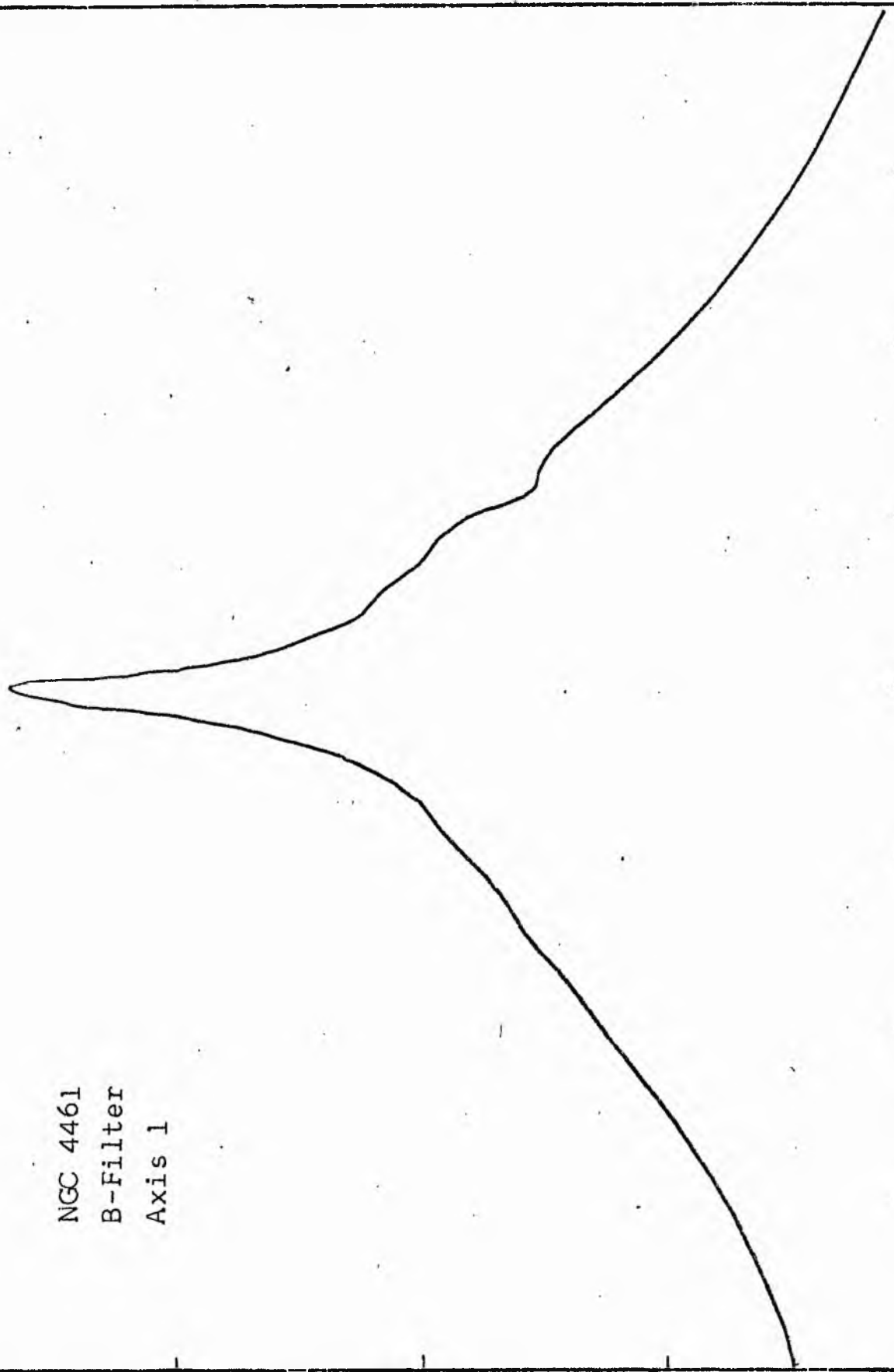
representing a value of 0.1, and r^* , as in case 3, denotes increasing distance r^* , in 50 arc sec steps.

5. Table of mean luminosity distribution. The areas in column 5 are expressed in square seconds of arc, and in the last column, μ is in mag sec^{-2} .

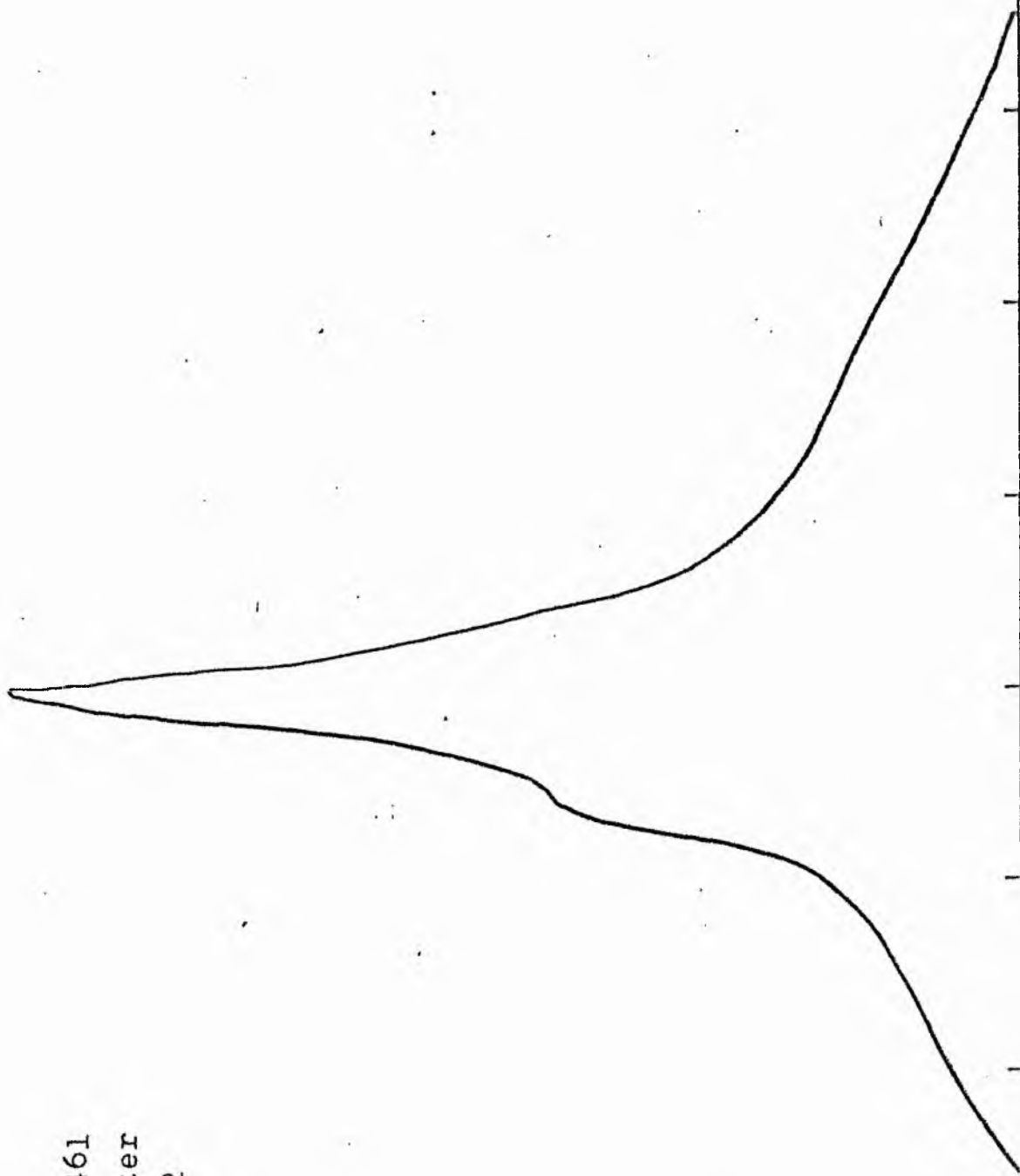
6. Table of photometric parameters. The units are as follows:-

- a. The total luminosity is expressed in sky units per square minute.
- b. The apparent central surface brightness is in mag. per square second, as are the surface brightnesses at $k = \frac{1}{4}, \frac{1}{2}$ and $\frac{3}{4}$.
- c. The major axis and minor axis at threshold together with the major axis at $\mu = 25.0 \text{ mag. sec}^{-2}$, and the semi-major axis measurements at $k = \frac{1}{4}, \frac{1}{2}$ and $\frac{3}{4}$, are expressed in minutes of arc.
- d. The gradient of the exponential component $G(a) = d \log I/da$ is expressed per minute of arc, as is the equivalent gradient of the exponential component $G(r^*) = d \log I/dr^*$.
- e. The mean surface brightness is in magnitudes per square minute.

NGC 4461
B-Filter
Axis 1

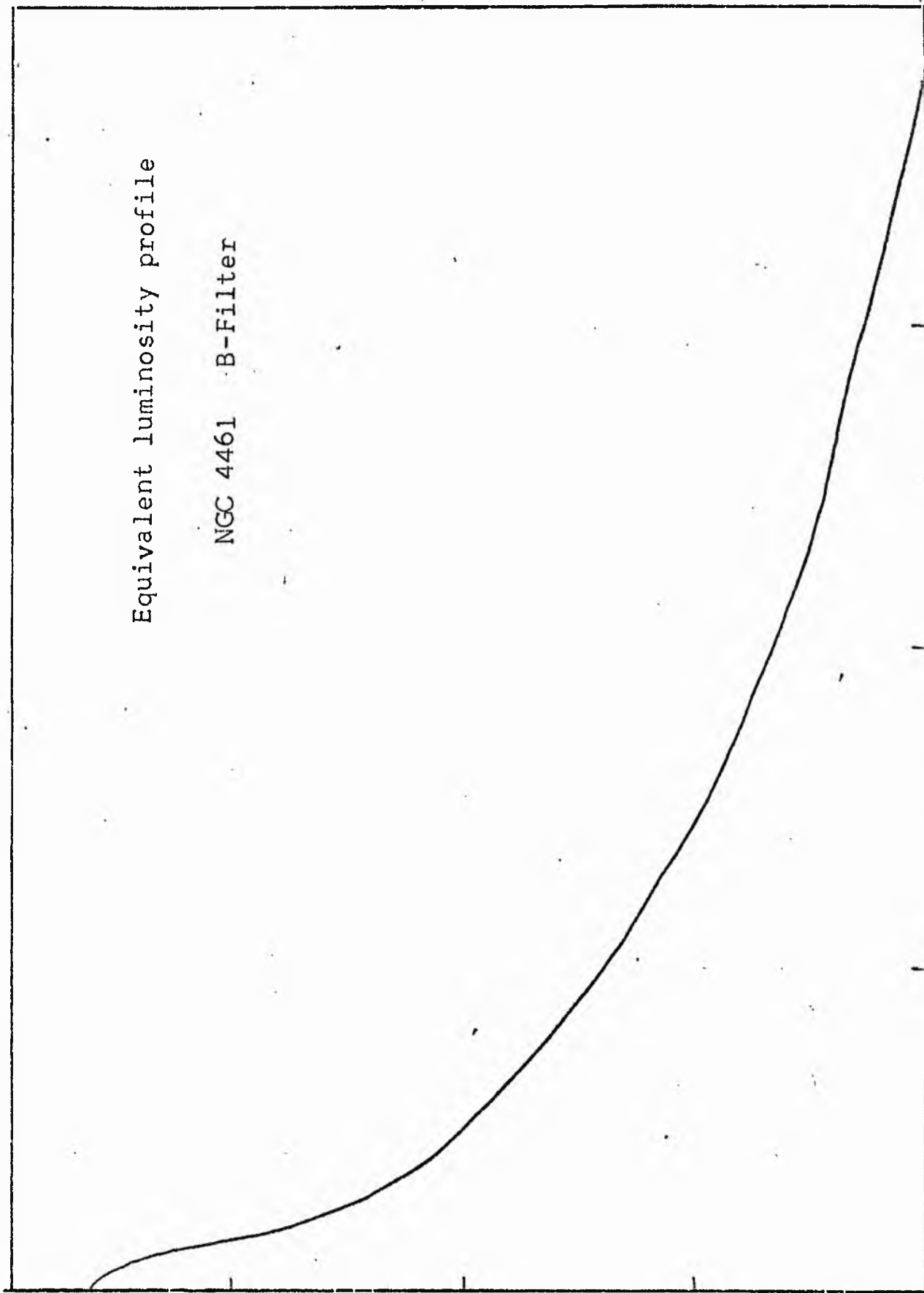


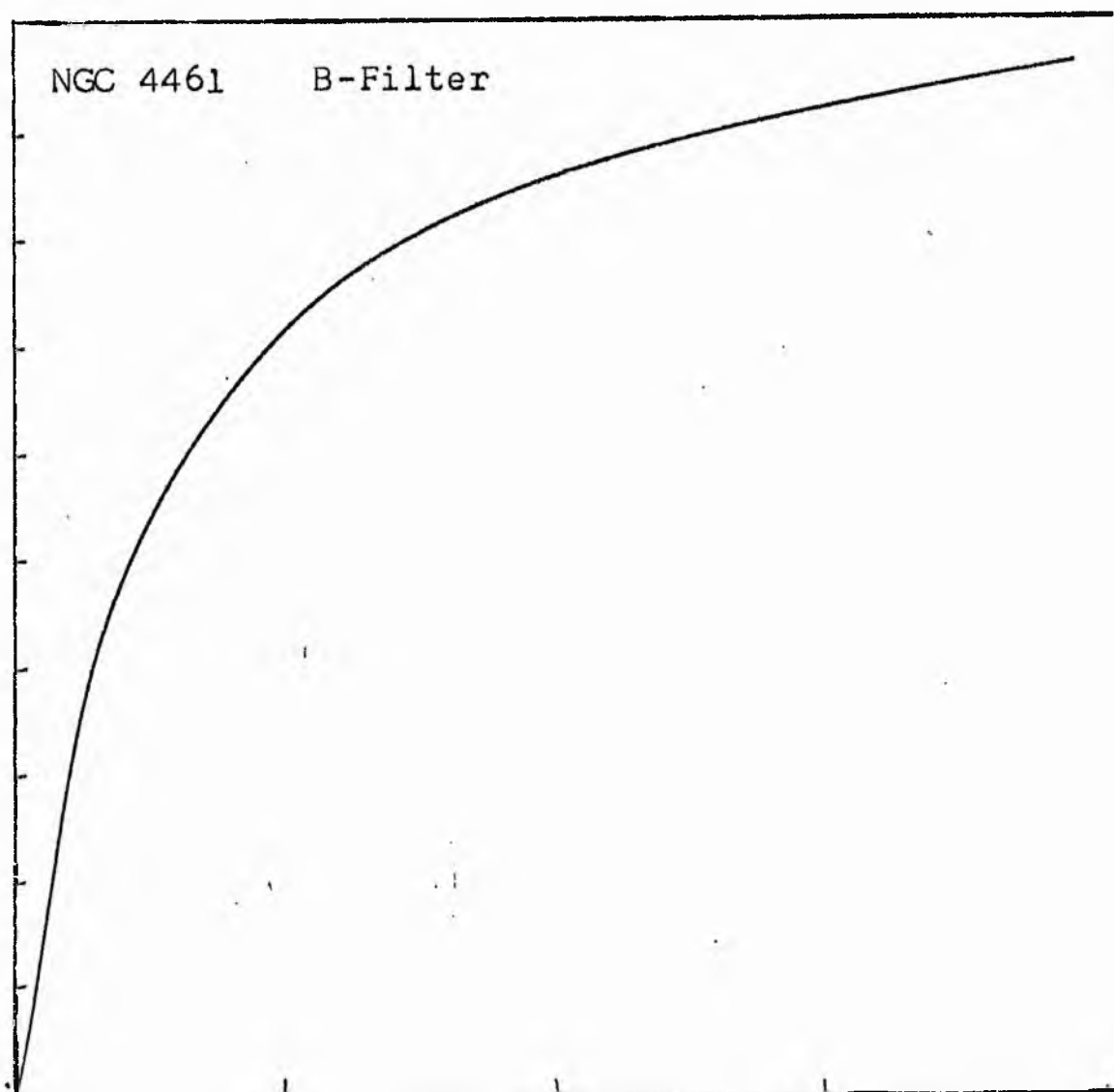
NGC 4461
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4461 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4461
B COLOUR

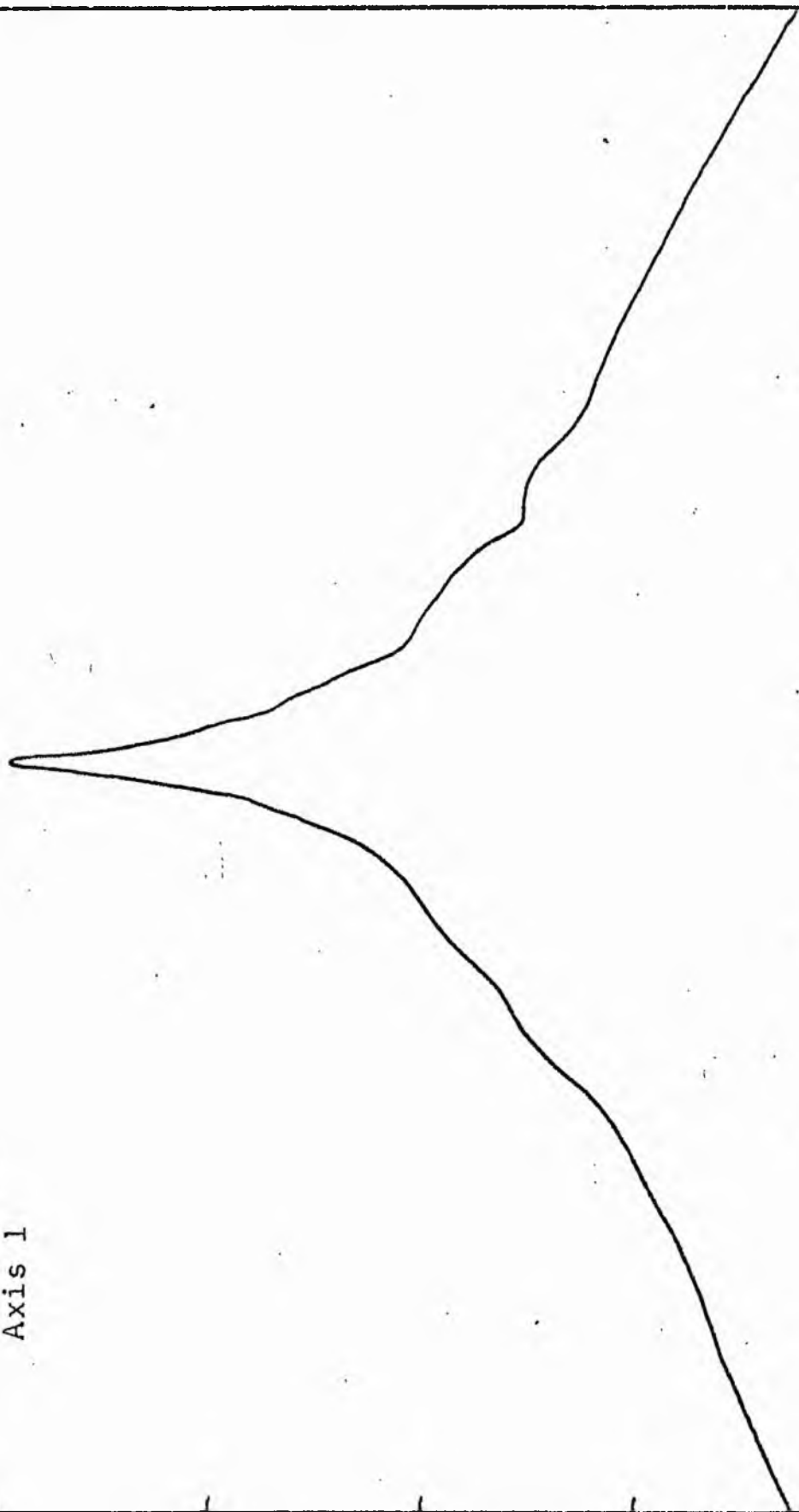
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.65	44.668		0.0	0.0			0.0	0.0	0.0	1.481	18.35
1.60	39.811	42.239	2.55	20.43	20.43	862.8760	862.88	0.05	0.12	1.431	18.48
1.50	31.623	35.717	3.10	30.19	9.76	348.6838	1211.56	0.07	0.15	1.331	18.73
1.40	25.119	28.371	5.01	78.85	48.66	1380.6208	2592.18	0.15	0.24	1.231	18.98
1.30	19.953	22.536	5.76	104.23	25.38	571.8699	3164.05	0.19	0.27	1.131	19.23
1.20	15.849	17.901	6.87	148.27	44.04	788.3992	3952.45	0.23	0.32	1.031	19.49
1.10	12.589	14.219	7.70	186.27	37.99	540.2031	4492.65	0.26	0.36	0.931	19.73
1.00	10.000	11.295	8.33	217.99	31.73	358.3379	4850.99	0.27	0.39	0.831	19.98
0.90	7.943	8.972	8.94	251.09	33.10	296.9209	5147.91	0.30	0.42	0.731	20.23
0.80	6.310	7.126	10.26	330.71	79.62	567.4055	5715.31	0.34	0.48	0.631	20.48
0.70	5.012	5.661	11.10	387.08	56.37	319.0798	6034.39	0.35	0.52	0.531	20.73
0.60	3.981	4.496	12.20	467.59	80.52	362.0493	6396.43	0.38	0.57	0.431	20.99
0.50	3.162	3.572	13.32	557.39	89.79	320.7136	6717.14	0.39	0.63	0.331	21.23
0.40	2.512	2.837	15.10	716.31	158.93	450.8828	7168.03	0.42	0.71	0.231	21.48
0.30	1.995	2.254	17.60	973.14	256.82	578.7683	7746.79	0.46	0.83	0.131	21.73
0.20	1.585	1.790	20.33	1298.45	325.31	582.3242	8329.12	0.49	0.96	0.031	21.98
0.10	1.259	1.422	23.35	1712.87	414.42	589.2625	8918.38	0.52	1.10	-0.069	22.23
-0.00	1.000	1.129	26.73	2244.64	531.78	600.6711	9519.00	0.56	1.26	-0.169	22.48
-0.10	0.794	0.897	30.24	2872.85	628.21	563.6013	10082.60	0.59	1.43	-0.269	22.73
-0.20	0.631	0.713	33.97	3625.27	752.42	536.2046	10618.83	0.62	1.60	-0.369	22.98
-0.30	0.501	0.566	37.92	4517.37	892.10	504.9424	11123.79	0.65	1.79	-0.469	23.23
-0.40	0.398	0.450	42.09	5565.54	1048.17	471.3035	11595.09	0.68	1.98	-0.569	23.48
-0.50	0.316	0.357	46.48	6787.06	1221.52	436.2852	12031.38	0.71	2.19	-0.669	23.73
-0.60	0.251	0.284	51.08	8196.93	1409.87	399.9902	12431.37	0.73	2.41	-0.769	23.98
-0.70	0.200	0.225	55.00	9503.32	1306.38	294.4019	12725.77	0.75	2.59	-0.869	24.23
-0.80	0.158	0.179	58.70	10824.95	1321.64	236.5823	12962.35	0.76	2.77	-0.969	24.48
-0.90	0.126	0.142	65.80	13601.95	2777.00	394.8625	13357.21	0.78	3.10	-1.069	24.73
-1.00	0.100	0.113	70.20	15481.89	1879.94	212.3313	13569.54	0.80	3.31	-1.169	24.98
-1.10	0.079	0.090	75.80	18050.45	2568.56	230.4415	13799.98	0.81	3.57	-1.269	25.23
-1.20	0.063	0.071	86.90	23724.07	5673.62	404.3259	14204.30	0.83	4.10	-1.369	25.48
-1.30	0.050	0.057	96.30	29134.15	5410.08	306.2498	14510.55	0.85	4.54	-1.469	25.73
-1.40	0.040	0.045	106.70	35766.68	6632.53	298.2305	14808.78	0.87	5.03	-1.569	25.98
-1.50	0.032	0.036	119.20	44637.75	8871.07	316.8469	15125.62	0.89	5.62	-1.669	26.23
-1.60	0.025	0.028	136.40	50449.20	13811.45	391.8442	15517.46	0.91	6.43	-1.769	26.48
-1.70	0.020	0.023	148.70	69465.87	11016.68	248.2706	15765.73	0.93	7.01	-1.869	26.73
-1.80	0.016	0.018	161.20	81635.62	12169.75	217.8496	15983.58	0.94	7.60	-1.969	26.98
-1.90	0.013	0.014	177.40	98868.25	17232.62	245.0345	16228.61	0.95	8.36	-2.069	27.23
-2.00	0.010	0.011	190.30	113769.87	14901.62	168.3100	16396.92	0.96	8.97	-2.169	27.48
-∞							17016.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4461

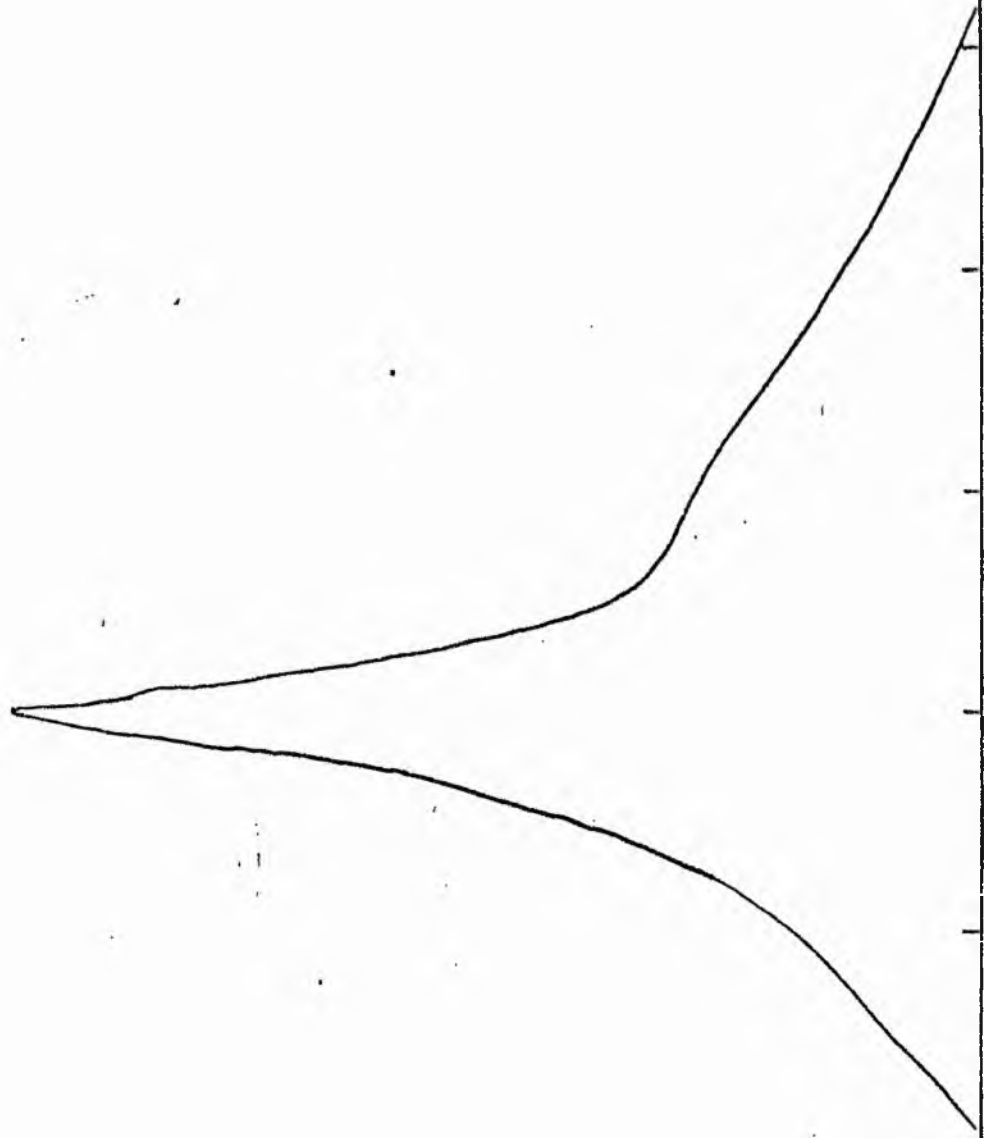
B-FILTER

Total luminosity	L_T	= 4.73
Total apparent magnitude	m_T	= 11.90
Apparent central surface brightness	μ_0	= 18.35
Major axis at threshold	$2a_m$	= 7.53
Minor axis at threshold	$2b_m$	= 5.22
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.85
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.80
Gradient of exponential component	$G(a)$	= -0.85
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.65
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.53
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.13
Axis ratio	b/a	= 0.88
Equivalent radius	r_1^*	= 0.12
Surface brightness	μ_1	= 19.65
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.46
Axis ratio	b/a	= 0.72
Equivalent radius	r_e^*	= 0.35
Surface brightness	μ_e	= 22.06
Mean surface brightness	μ_e'	= 11.60
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.42
Axis ratio	b/a	= 0.39
Equivalent radius	r_3^*	= 0.93
Surface brightness	μ_3	= 24.23
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.90 \\ = 2.62 \end{cases}$

NGC 4461
V-Filter
Axis 1

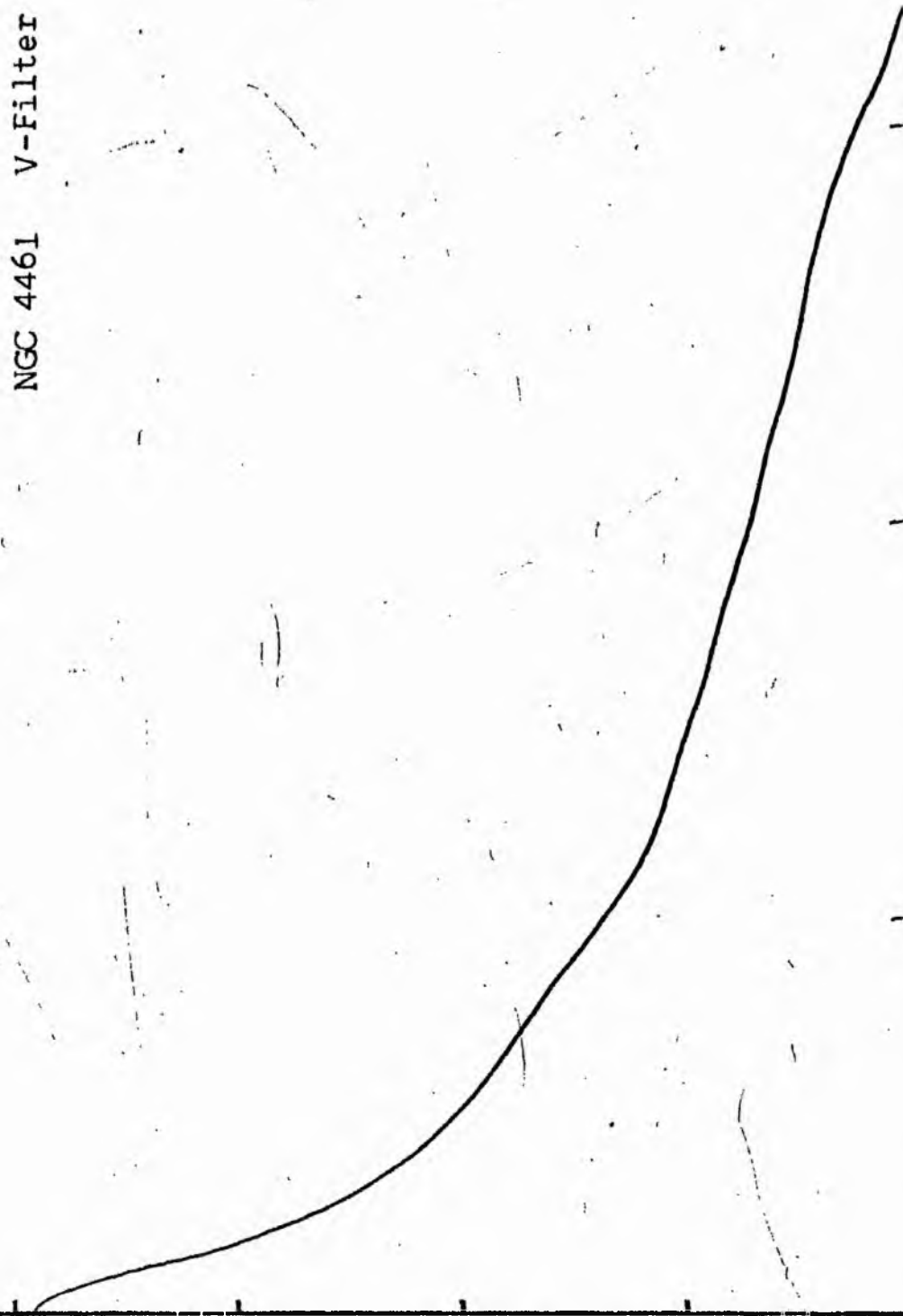


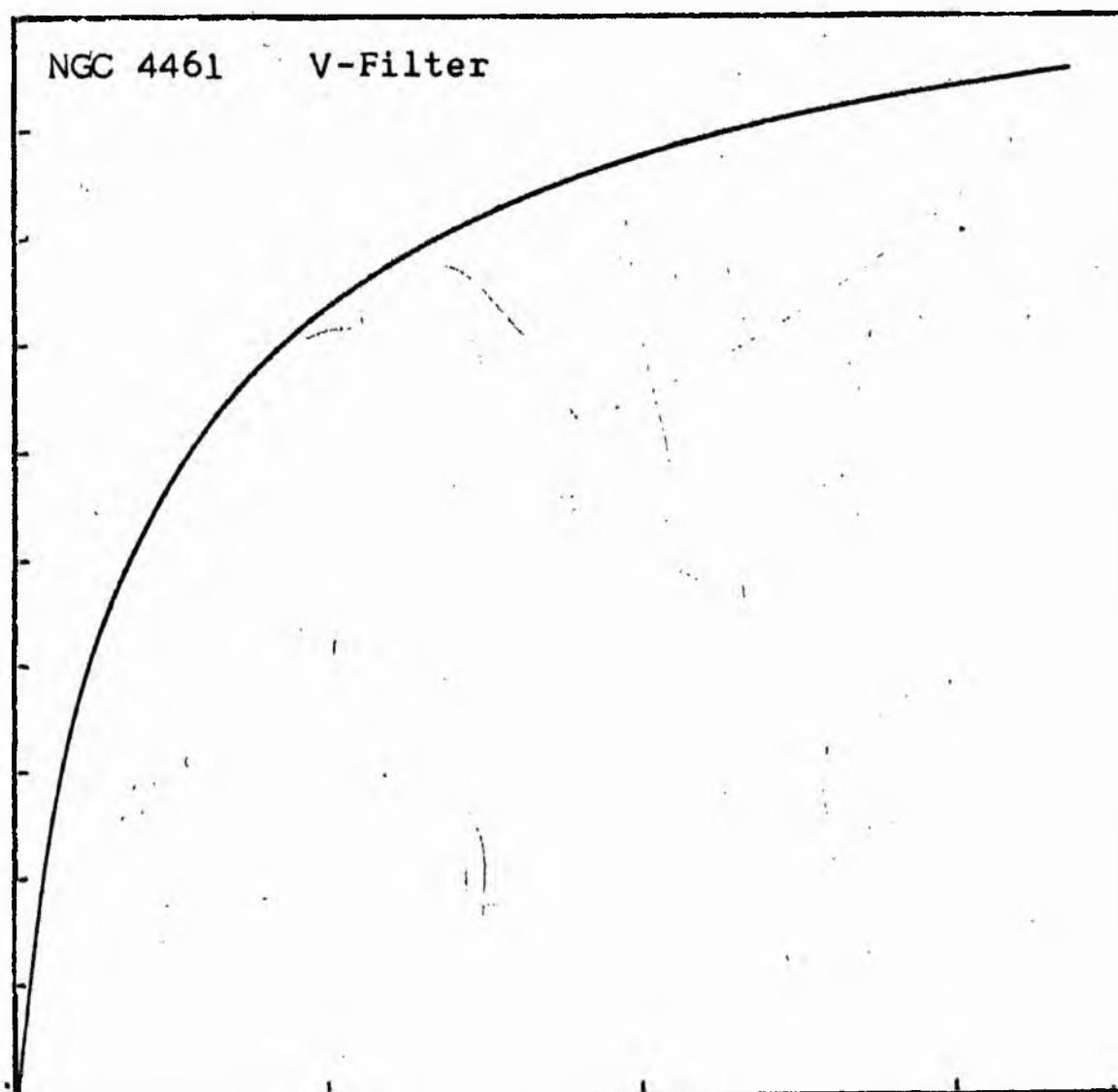
NGC 4461
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4461 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4461
V COLOUR

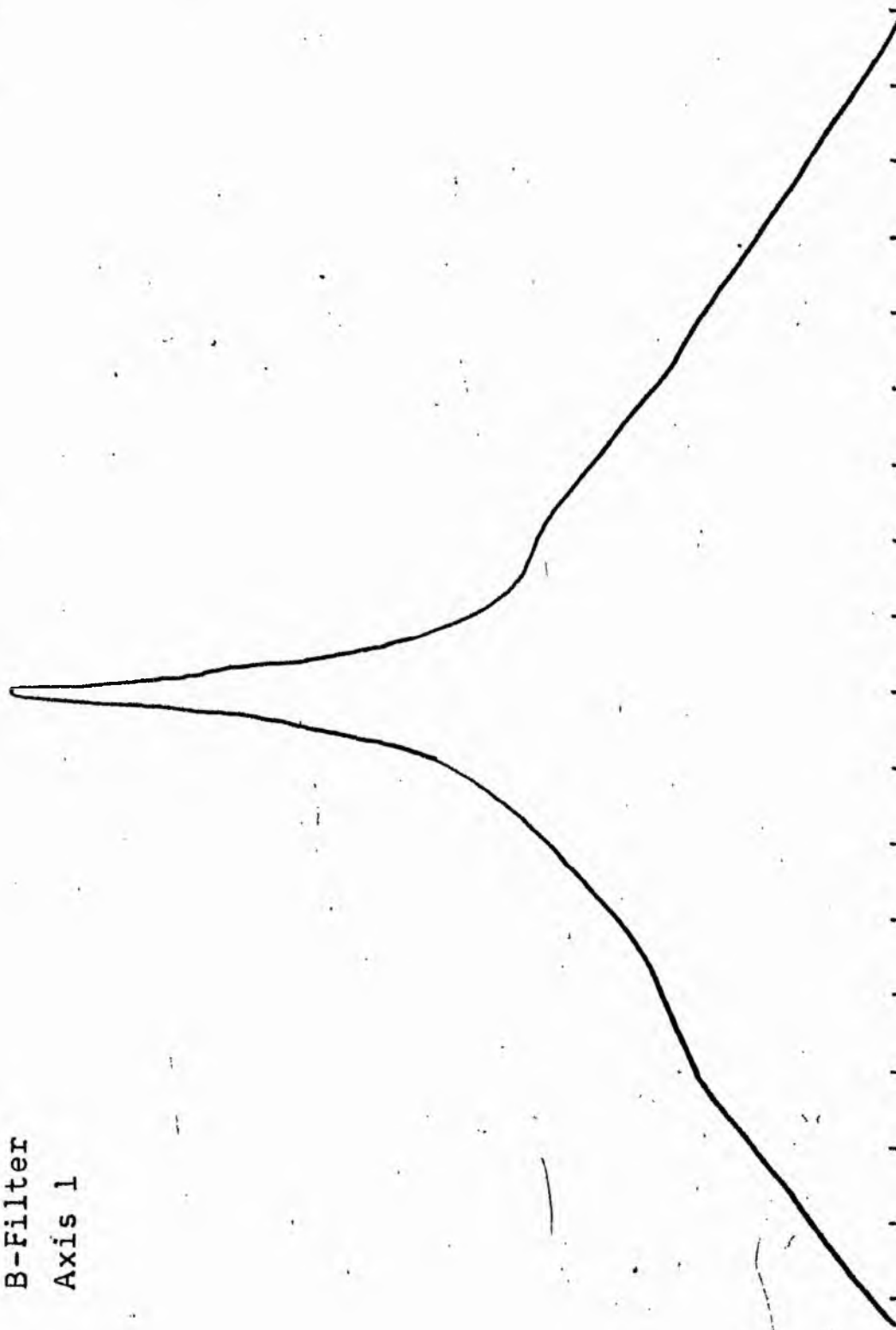
LOG I	I	I	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
1.86	72.444		0.0	0.0			0.0	0.0	0.0	1.616	17.11
		67.770			2.01	136.2587					
1.80	63.096		0.80	2.01			136.26	0.01	0.04	1.556	17.26
		56.607			9.33	528.1736					
1.70	50.119		1.90	11.34			664.43	0.04	0.10	1.456	17.51
		44.965			20.83	936.5549					
1.60	39.811		3.20	32.17			1600.99	0.09	0.18	1.356	17.76
		35.717			23.25	830.3306					
1.50	31.623		4.20	55.42			2431.32	0.14	0.23	1.256	18.01
		28.371			26.30	746.0095					
1.40	25.119		5.10	81.71			3177.33	0.19	0.28	1.156	18.26
		22.536			28.76	648.1462					
1.30	19.952		5.93	110.47			3825.47	0.23	0.33	1.056	18.51
		17.901			34.79	622.8245					
1.20	15.849		6.80	145.27			4448.30	0.26	0.38	0.956	18.76
		14.219			41.00	582.9475					
1.10	12.589		7.70	186.27			5031.24	0.30	0.42	0.856	19.01
		11.295			30.16	340.6345					
1.00	10.000		8.30	216.42			5371.87	0.32	0.46	0.756	19.26
		8.972			61.17	548.7617					
0.90	7.943		9.40	277.59			5920.64	0.35	0.52	0.656	19.51
		7.126			70.09	499.4878					
0.80	6.310		10.52	347.68			6420.12	0.38	0.58	0.556	19.76
		5.661			53.47	392.6692					
0.70	5.012		11.30	401.15			6722.79	0.40	0.62	0.456	20.01
		4.496			81.12	364.7610					
0.60	3.981		12.39	482.27			7087.55	0.42	0.68	0.356	20.26
		3.572			86.05	307.3386					
0.50	3.162		13.45	568.32			7394.88	0.44	0.74	0.256	20.51
		2.837			136.65	387.6924					
0.40	2.512		14.98	704.97			7782.57	0.46	0.83	0.156	20.76
		2.254			189.11	426.1755					
0.30	1.995		16.87	894.09			8208.75	0.48	0.93	0.056	21.01
		1.790			260.41	466.1562					
0.20	1.585		19.17	1154.50			8674.90	0.51	1.06	-0.044	21.26
		1.422			348.11	494.9817					
0.10	1.259		21.87	1502.61			9169.88	0.54	1.21	-0.144	21.51
		1.129			471.89	532.9800					
-0.00	1.000		25.07	1974.50			9702.86	0.57	1.38	-0.244	21.76
		0.897			580.83	521.0999					
-0.10	0.794		28.52	2555.34			10223.96	0.60	1.57	-0.344	22.01
		0.713			718.20	511.8130					
-0.20	0.631		32.28	3273.53			10735.77	0.63	1.78	-0.444	22.26
		0.566			884.37	500.6145					
-0.30	0.501		36.38	4157.91			11236.38	0.66	2.01	-0.544	22.51
		0.450			1069.15	480.7363					
-0.40	0.398		40.79	5227.05			11717.12	0.69	2.25	-0.644	22.76
		0.357			1285.41	459.1023					
-0.50	0.316		45.53	6512.46			12176.22	0.72	2.51	-0.744	23.01
		0.284			938.42	266.2361					
-0.60	0.251		48.70	7450.88			12442.45	0.73	2.69	-0.844	23.26
		0.225			1440.57	324.6421					
-0.70	0.200		53.20	8891.46			12767.09	0.75	2.93	-0.944	23.51
		0.179			611.86	109.5269					
-0.80	0.158		55.00	9503.32			12876.62	0.76	3.03	-1.044	23.76
		0.142			4473.27	636.0547					
-0.90	0.126		66.70	13976.59			13512.67	0.80	3.68	-1.144	24.01
		0.113			2856.81	322.6646					
-1.00	0.100		73.20	16833.40			13835.34	0.82	4.04	-1.244	24.26
		0.090			4342.21	389.5657					
-1.10	0.079		82.10	21175.61			14224.90	0.84	4.53	-1.344	24.51
		0.071			4554.82	324.5950					
-1.20	0.063		90.50	25730.43			14549.49	0.86	4.99	-1.444	24.76
		0.057			8053.28	455.8730					
-1.30	0.050		103.70	33783.71			15005.36	0.89	5.72	-1.544	25.01
		0.045			5906.41	265.5801					
-1.40	0.040		112.40	39690.12			15270.94	0.90	6.20	-1.644	25.26
		0.036			5247.71	187.4316					
-1.50	0.032		119.60	44937.84			15458.37	0.91	6.60	-1.744	25.51
		0.028			7991.83	226.7354					
-1.60	0.025		129.80	52929.67			15685.11	0.93	7.16	-1.844	25.76
		0.023			6379.70	143.7719					
-1.70	0.020		137.40	59309.37			15828.87	0.93	7.58	-1.944	26.01
		0.018			8762.13	156.8498					
-1.80	0.016		147.20	68071.50			15985.72	0.94	8.12	-2.044	26.26
		0.014			10752.75	152.8954					
-1.90	0.013		158.40	78824.25			16138.62	0.95	8.74	-2.144	26.51
		0.011			9001.56	101.6702					
-2.00	0.010		167.20	87825.81			16240.29	0.96	9.22	-2.244	26.76
-∞							16940.00	(1)		∞	

PHOTOMETRIC PARAMETERS OF NGC 4461

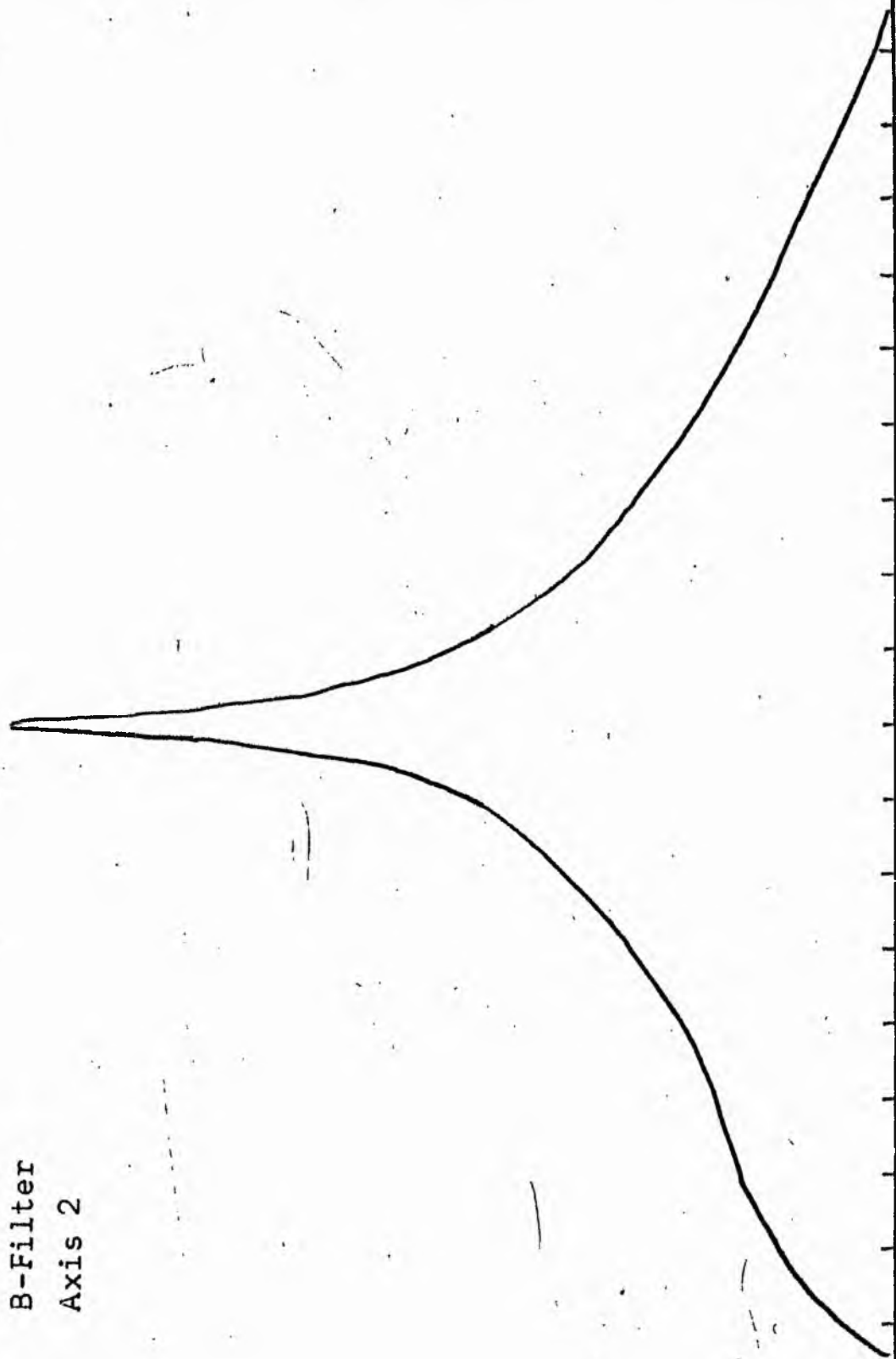
V-FILTER

Total luminosity	L_T	= 4.71
Total apparent magnitude	m_T	= 10.62
Apparent central surface brightness	μ_0	= 17.11
Major axis at threshold	$2a_m$	= 7.18
Minor axis at threshold	$2b_m$	= 4.32
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.43
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.68
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.77
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.48
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.15
Axis ratio	b/a	= 0.67
Equivalent radius	r_1^*	= 0.11
Surface brightness	μ_1	= 18.68
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.41
Axis ratio	b/a	= 0.63
Equivalent radius	r_e^*	= 0.30
Surface brightness	μ_e	= 21.18
Mean surface brightness	μ_e'	= 9.97
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.32
Axis ratio	b/a	= 0.37
Equivalent radius	r_3^*	= 0.87
Surface brightness	μ_3	= 23.51
Concentration indices	$\left\{ \begin{matrix} C_{21} \\ C_{32} \end{matrix} \right.$	= 2.79
		= 2.88

NGC 4472
B-Filter
Axis 1

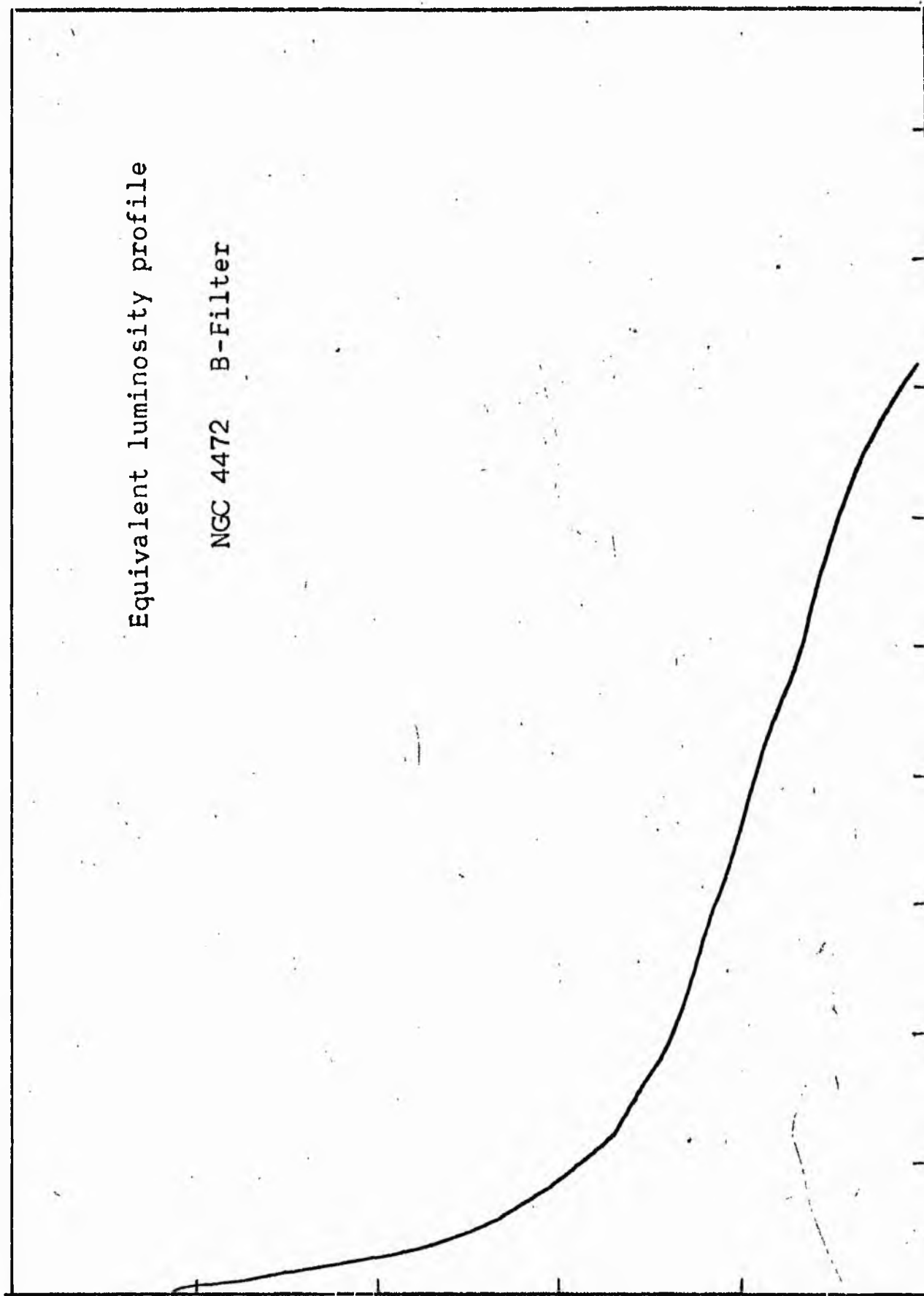


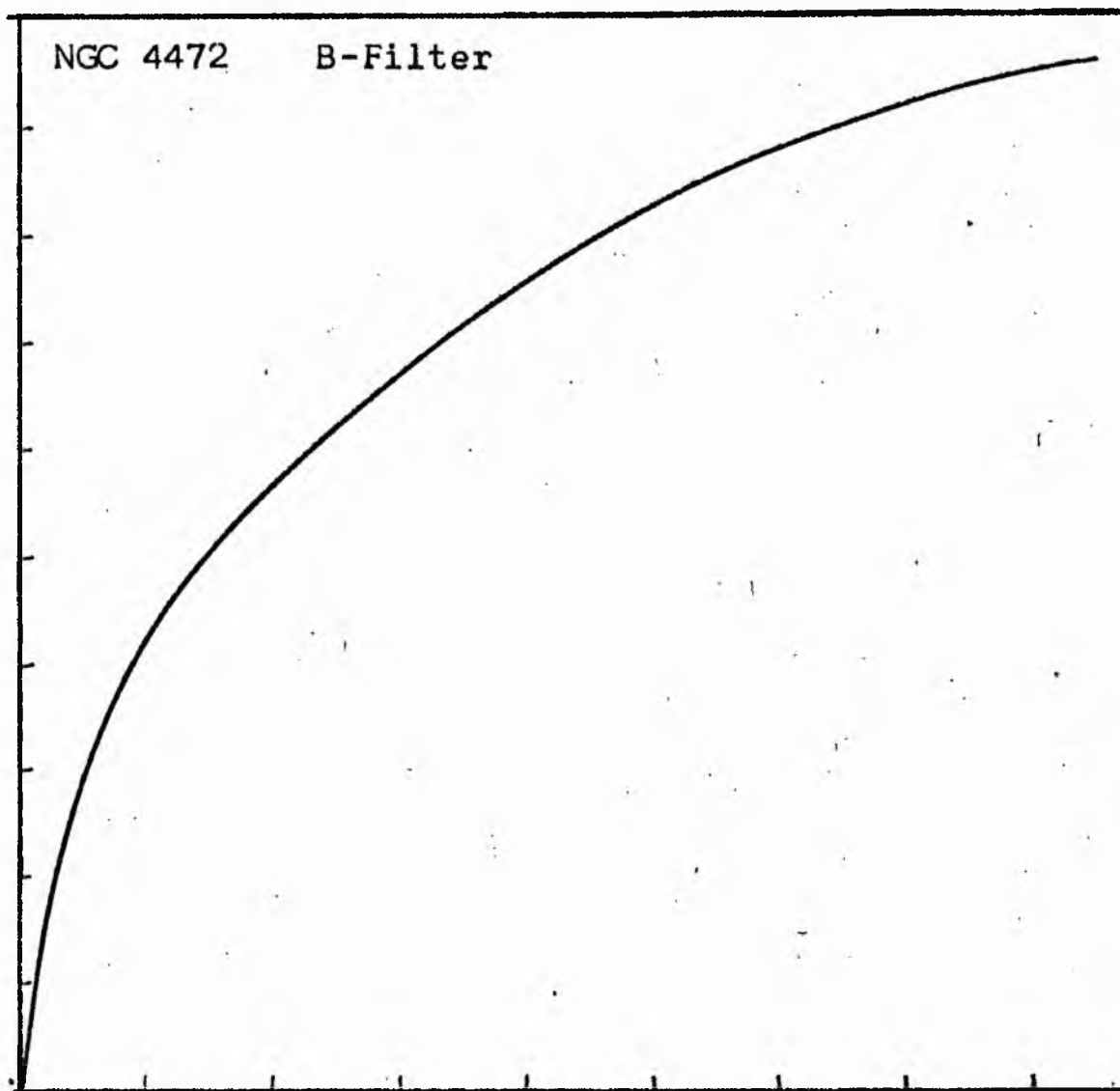
NGC 4472
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4472 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4472
B COLOUR

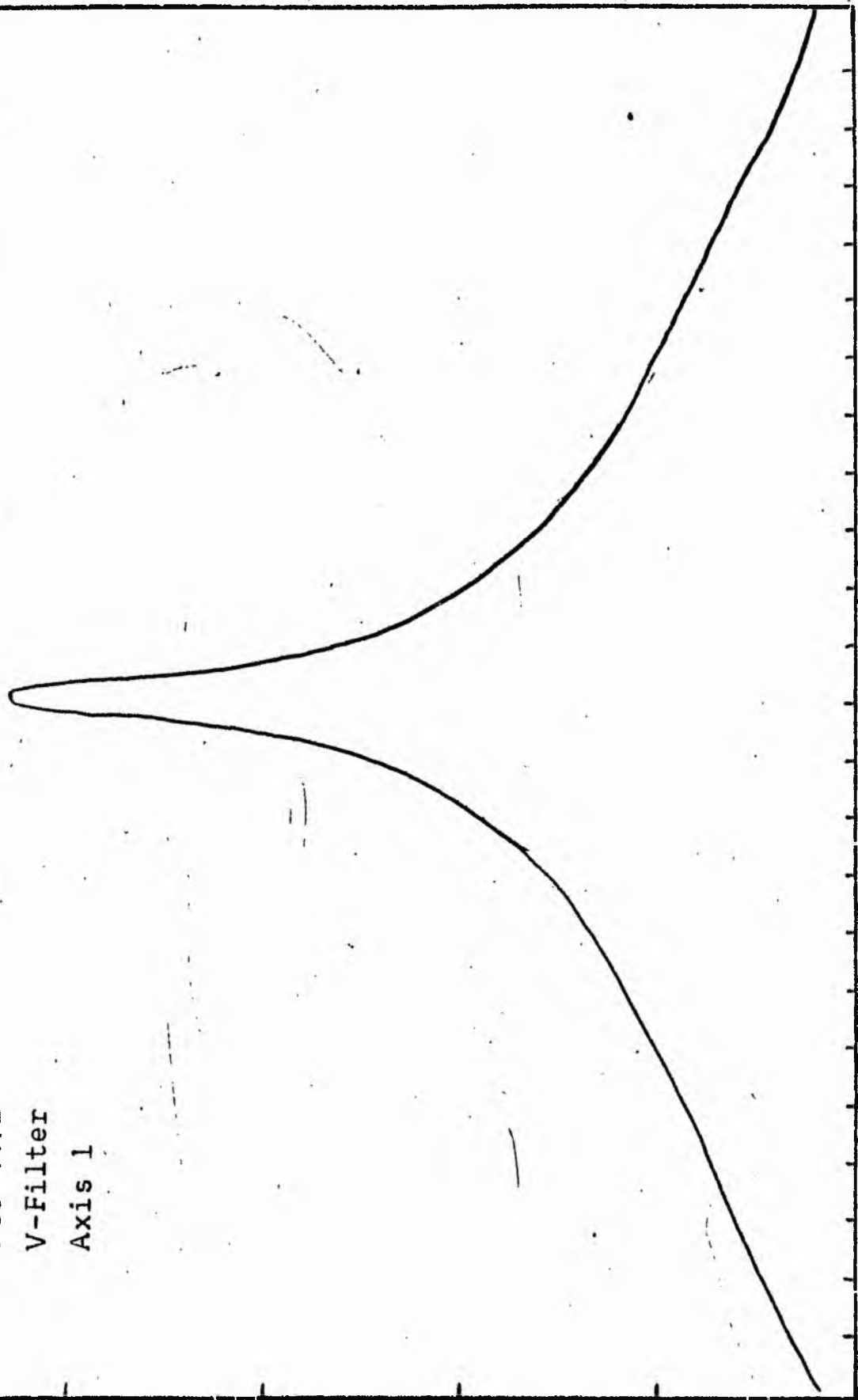
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.02	104.713		0.0	0.0			0.0	0.0	0.0	2.339	16.69
2.00	100.000	102.356	0.80	2.01	2.01	205.7995	205.80	0.00	0.01	2.319	16.74
1.90	79.433	89.716	1.90	11.34	9.33	817.1006	1047.90	0.01	0.03	2.219	16.47
1.80	63.096	71.264	3.80	43.36	34.02	2424.6511	3467.55	0.04	0.05	2.119	17.24
1.70	50.119	56.607	5.20	84.95	39.58	2240.7383	5708.29	0.06	0.07	2.019	17.49
1.60	39.811	44.965	6.30	124.69	39.74	1786.9429	7495.23	0.08	0.08	1.919	17.74
1.50	31.623	35.717	7.76	169.18	64.49	2303.3396	9798.97	0.11	0.10	1.819	17.99
1.40	25.119	28.371	9.04	256.74	67.56	1916.6294	11715.20	0.13	0.12	1.719	18.24
1.30	19.952	22.536	10.27	331.35	74.62	1681.5374	13346.73	0.15	0.14	1.619	18.49
1.20	15.849	17.901	12.05	456.17	124.81	2234.2554	15630.90	0.18	0.16	1.519	18.74
1.10	12.589	14.219	14.84	691.06	235.69	3351.2979	18982.28	0.21	0.20	1.419	18.99
1.00	10.000	11.295	15.08	714.42	22.56	254.7969	19237.08	0.22	0.20	1.319	19.24
0.90	7.943	8.972	16.95	902.59	188.17	1688.1655	20925.24	0.23	0.23	1.219	19.49
0.80	6.310	7.126	20.34	1299.73	397.14	2830.1540	23755.39	0.27	0.27	1.119	19.74
0.70	5.012	5.661	21.25	1418.63	118.90	673.0564	24428.45	0.27	0.28	1.019	19.99
0.60	3.981	4.496	23.79	1778.03	359.40	1616.0271	26044.47	0.29	0.32	0.919	20.24
0.50	3.162	3.572	27.80	2427.95	649.92	2321.2791	28365.75	0.32	0.37	0.819	20.49
0.40	2.512	2.837	31.39	3045.51	667.57	1893.9258	30259.68	0.34	0.42	0.719	20.74
0.30	1.995	2.254	34.55	3750.12	654.61	1475.2056	31734.88	0.36	0.46	0.619	20.99
0.20	1.585	1.790	39.11	4805.35	1055.23	1888.9260	33623.80	0.38	0.52	0.519	21.24
0.10	1.259	1.422	44.80	6305.30	1499.95	2132.7103	35756.57	0.40	0.60	0.419	21.49
-0.00	1.000	1.129	51.35	8283.82	1978.52	2234.6479	37991.22	0.43	0.68	0.319	21.74
-0.10	0.794	0.897	58.78	10854.48	2570.66	2306.2908	40297.51	0.45	0.78	0.219	21.99
-0.20	0.631	0.713	67.09	14140.52	3286.04	2341.7532	42639.26	0.48	0.89	0.119	22.24
-0.30	0.501	0.566	73.72	17073.40	2932.89	1660.2136	44299.47	0.50	0.98	0.019	22.49
-0.40	0.398	0.450	86.03	23251.43	6178.03	2777.9138	47077.38	0.53	1.14	-0.081	22.74
-0.50	0.316	0.357	101.40	32301.72	9050.29	3232.4492	50309.83	0.56	1.35	-0.181	22.99
-0.60	0.251	0.284	123.70	48071.67	15769.95	4474.0352	54783.86	0.61	1.65	-0.281	23.24
-0.70	0.200	0.225	140.80	62280.93	14209.25	3202.1411	57986.00	0.65	1.87	-0.381	23.49
-0.80	0.158	0.179	164.05	84547.75	22266.82	3985.9104	61971.91	0.69	2.18	-0.481	23.74
-0.90	0.126	0.142	186.70	109506.12	24958.37	3548.8323	65520.75	0.73	2.40	-0.581	23.99
-1.00	0.100	0.113	217.90	149164.06	39657.94	4479.1914	69999.94	0.78	2.90	-0.681	24.24
-1.10	0.079	0.090	246.25	190503.25	41339.19	3708.7896	73708.69	0.83	3.28	-0.781	24.49
-1.20	0.063	0.071	266.43	223005.56	32502.31	2316.2476	76024.87	0.85	3.55	-0.881	24.74
-1.30	0.050	0.057	283.90	253209.62	30204.06	1709.7659	77734.62	0.87	3.78	-0.981	24.99
-1.40	0.040	0.045	298.35	279641.44	26431.01	1188.4993	78923.06	0.88	3.97	-1.081	25.24
-1.50	0.032	0.036	332.50	347322.69	67681.25	2417.3577	81340.37	0.91	4.43	-1.181	25.49
-1.60	0.025	0.028	363.27	414579.87	67257.19	1908.1467	83248.50	0.93	4.83	-1.281	25.74
-1.70	0.020	0.023	383.62	462329.94	47750.06	1076.0876	84324.56	0.94	5.11	-1.381	25.99
-1.80	0.016	0.018	397.93	497465.50	35135.56	628.9573	84953.50	0.95	5.30	-1.481	26.24
-1.90	0.013	0.014	409.73	527406.25	29940.75	425.7332	85379.19	0.96	5.45	-1.581	26.49
-2.00	0.010	0.011	418.50	550225.62	22819.37	257.7383	85636.87	0.96	5.57	-1.681	26.74
-∞							89276.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4472

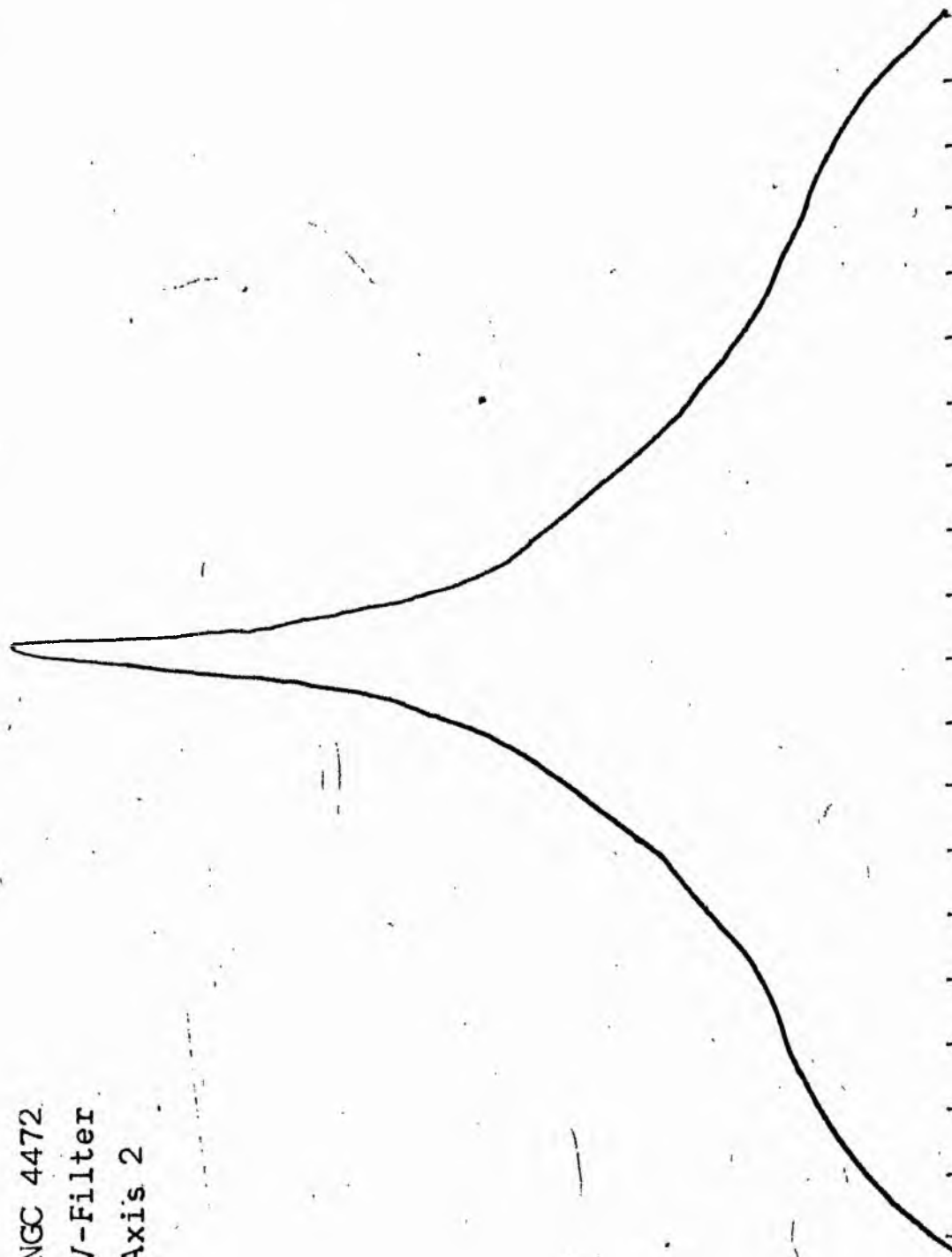
B-FILTER

Total luminosity	L_T	=	24.80
Total apparent magnitude	m_T	=	9.36
Apparent central surface brightness	μ_0	=	16.69
Major axis at threshold	$2a_m$	=	13.70
Minor axis at threshold	$2b_m$	=	13.71
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	=	10.08
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	=	0.87
Gradient of exponential component	$G(a)$	=	-0.28
Equivalent gradient of exponential comp....	$G(r^*)$	=	-0.27
Equivalent gradient of reduced exp. comp....	$G(\rho)$	=	-0.32
Parameters at $k = \frac{1}{4}$:			
Semi-major axis	a_1	=	0.38
Axis ratio	b/a	=	0.97
Equivalent radius	r_1^*	=	0.31
Surface brightness	μ_1	=	19.61
Parameters at $k = \frac{1}{2}$ (effective) :			
Semi-major axis	a_e	=	1.33
Axis ratio	b/a	=	0.98
Equivalent radius	r_e^*	=	1.25
Surface brightness	μ_e	=	22.49
Mean surface brightness	μ_e'	=	11.84
Parameters at $k = \frac{3}{4}$:			
Semi-major axis	a_3	=	3.80
Axis ratio	b/a	=	0.81
Equivalent radius	r_3^*	=	3.27
Surface brightness	μ_3	=	24.09
Concentration indices	$\begin{Bmatrix} C_{21} \\ C_{32} \end{Bmatrix}$	=	$\begin{Bmatrix} 4.04 \\ 2.61 \end{Bmatrix}$

NGC 4472
V-Filter
Axis 1

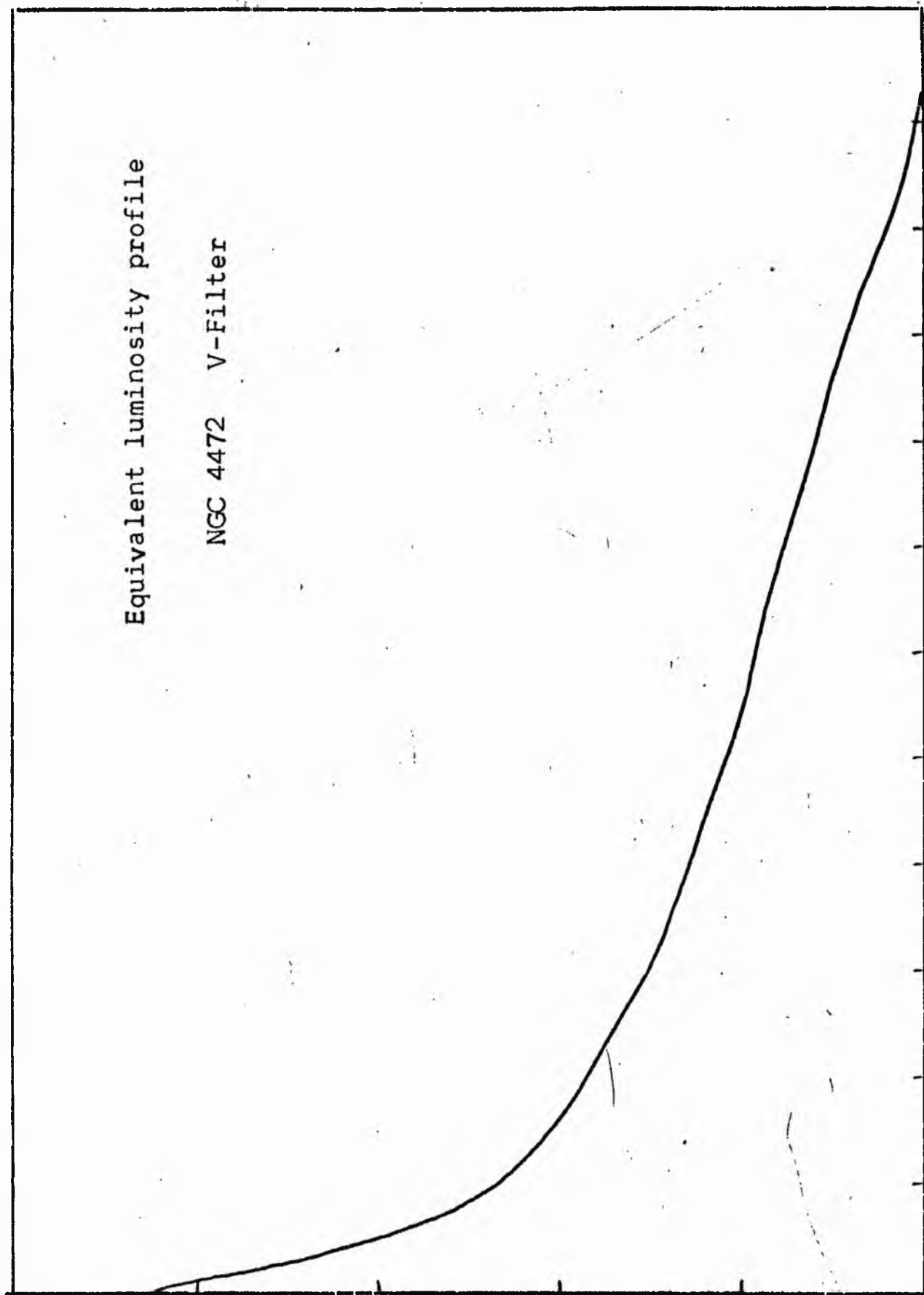


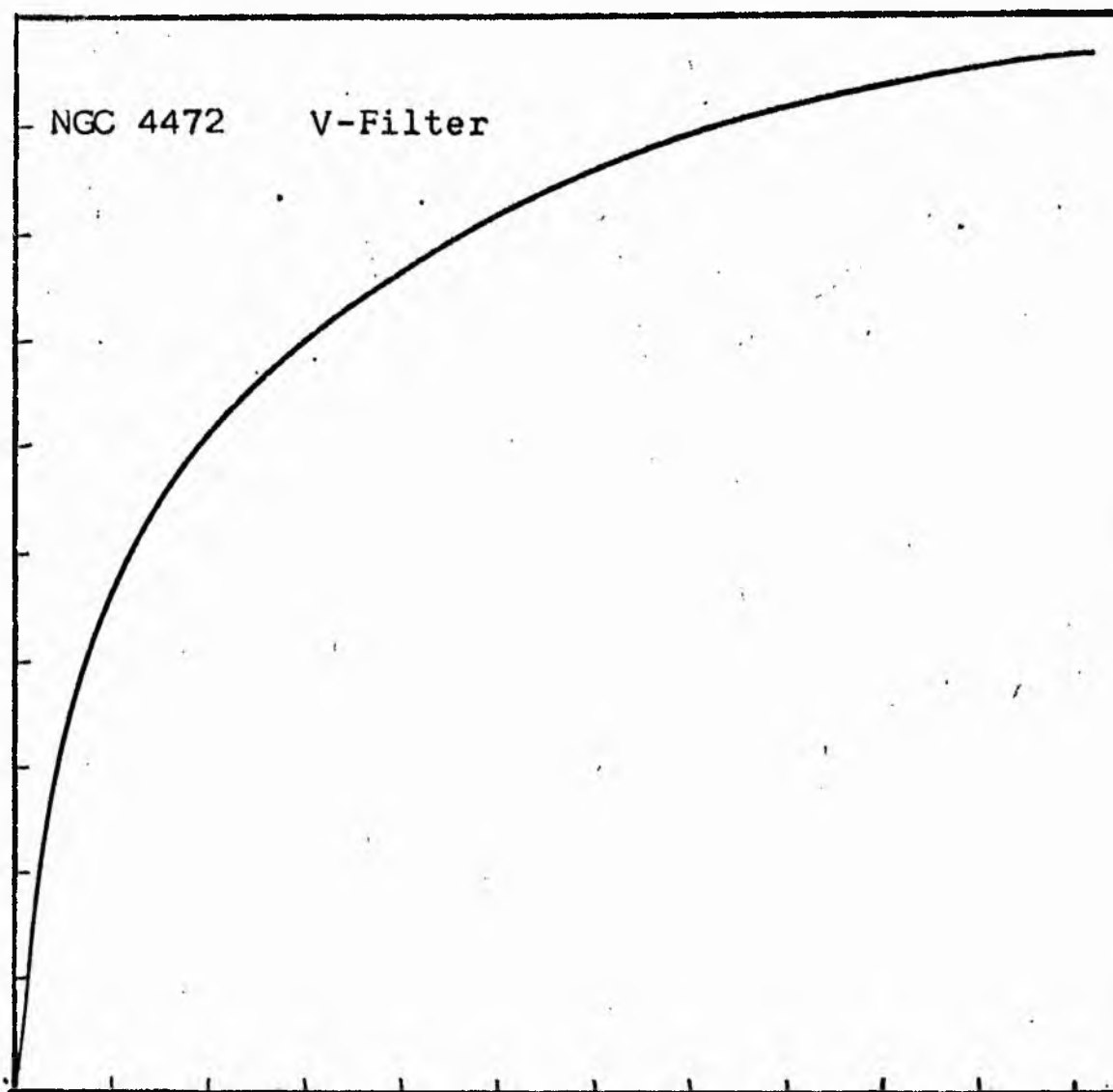
NGC 4472
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4472 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4472
V COLOUR

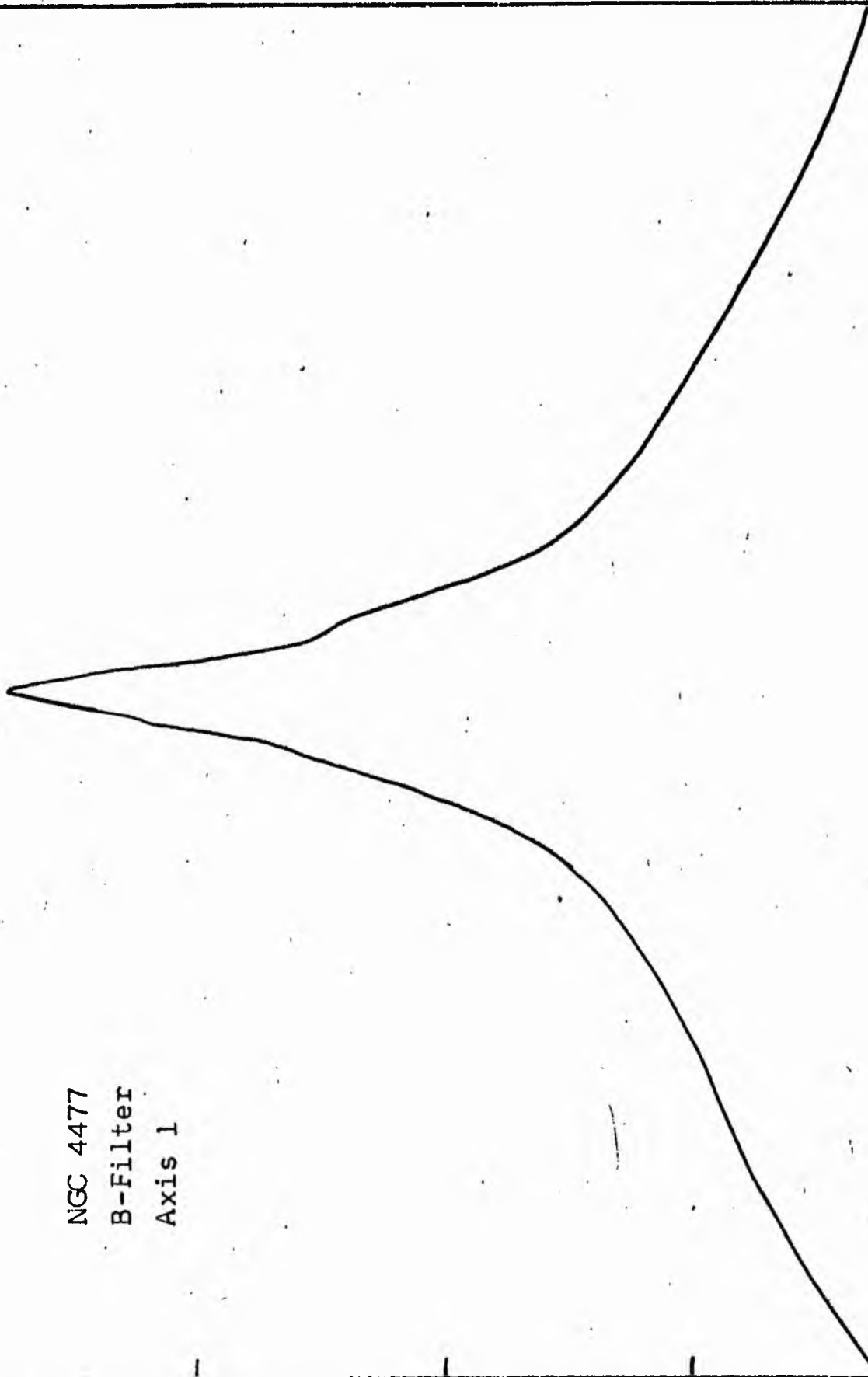
LOG I	I	I	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
2.24	173.780		0.0	0.0			0.0	0.0	0.0	2.000	16.05
2.20	158.489	166.135	1.20	4.52	4.52	751.5742	751.57	0.00	0.02	1.960	16.15
2.10	125.892	142.191	3.00	28.27	23.75	3377.0891	4128.66	0.02	0.05	1.860	16.40
2.00	100.000	112.946	4.20	55.42	27.14	3065.7312	7194.39	0.04	0.07	1.760	16.65
1.90	79.433	89.716	6.10	116.90	61.48	5515.8359	12710.23	0.07	0.11	1.660	16.90
1.80	63.095	71.264	8.30	216.42	99.53	7092.5937	19802.82	0.10	0.15	1.560	17.15
1.70	50.118	56.607	10.40	339.79	123.37	6983.6055	26786.43	0.14	0.19	1.460	17.40
1.60	39.810	44.964	11.50	415.48	75.68	3402.9622	30189.39	0.16	0.21	1.360	17.65
1.50	31.623	35.717	15.10	716.31	300.84	10744.9102	40934.30	0.21	0.27	1.260	17.90
1.40	25.119	28.371	17.25	934.82	218.51	6199.1328	47133.43	0.24	0.31	1.160	18.15
1.30	19.952	22.536	20.32	1297.17	362.35	8165.7422	55299.17	0.29	0.36	1.060	18.40
1.20	15.849	17.901	22.33	1566.49	269.32	4820.9414	60120.11	0.31	0.40	0.960	18.65
1.10	12.589	14.219	24.50	1885.74	319.25	4539.4609	64659.57	0.33	0.44	0.860	18.90
1.00	10.000	11.295	26.80	2256.42	370.67	4186.5820	68846.12	0.36	0.48	0.760	19.15
0.90	7.943	8.972	28.58	2566.10	309.69	2778.3657	71624.44	0.37	0.51	0.660	19.40
0.80	6.310	7.126	31.53	3123.19	557.08	3969.9746	75594.37	0.39	0.56	0.560	19.65
0.70	5.012	5.661	33.79	3586.96	463.77	2625.2510	78219.62	0.40	0.60	0.460	19.90
0.60	3.981	4.496	37.93	4519.76	932.80	4194.2734	82413.87	0.43	0.68	0.360	20.15
0.50	3.162	3.572	43.47	5936.48	1416.72	5060.0078	87473.87	0.45	0.77	0.260	20.40
0.40	2.512	2.837	48.61	7423.36	1486.89	4218.3750	91692.25	0.47	0.87	0.160	20.65
0.30	1.995	2.254	51.72	8403.62	980.26	2209.0710	93901.31	0.49	0.92	0.060	20.90
0.20	1.585	1.790	59.20	11010.15	2606.52	4665.8320	98567.12	0.51	1.06	-0.040	21.15
0.10	1.259	1.422	68.86	14896.48	3886.33	5525.9531	104093.06	0.54	1.23	-0.140	21.40
-0.00	1.000	1.129	81.78	21010.86	6114.39	6905.9023	110998.94	0.57	1.46	-0.240	21.65
-0.10	0.794	0.897	93.09	27224.25	6213.38	5574.3711	116573.25	0.60	1.66	-0.340	21.90
-0.20	0.631	0.713	104.57	34352.95	7128.70	5080.1680	121653.37	0.63	1.86	-0.440	22.15
-0.30	0.501	0.566	116.21	42426.46	8073.51	4570.1445	126223.50	0.65	2.07	-0.540	22.40
-0.40	0.398	0.450	125.75	49678.20	7251.74	3260.6963	129484.19	0.67	2.24	-0.640	22.65
-0.50	0.316	0.357	139.25	60917.25	11239.05	4014.1892	133498.37	0.69	2.48	-0.740	22.90
-0.60	0.251	0.284	162.37	82824.94	21907.68	6215.3359	139713.69	0.72	2.89	-0.840	23.15
-0.70	0.200	0.225	191.88	115666.94	32842.00	7401.1289	147114.81	0.76	3.42	-0.940	23.40
-0.80	0.158	0.179	218.08	149410.62	33743.69	6040.3320	153155.12	0.79	3.89	-1.040	23.65
-0.90	0.126	0.142	235.08	173612.56	24201.94	3441.2686	156596.37	0.81	4.19	-1.140	23.90
-1.00	0.100	0.113	265.44	221351.31	47738.75	5391.8711	161988.19	0.84	4.73	-1.240	24.15
-1.10	0.079	0.090	302.31	287113.87	65762.56	5899.9453	167888.12	0.87	5.39	-1.340	24.40
-1.20	0.063	0.071	328.16	338314.56	51200.69	3648.7642	171536.87	0.89	5.85	-1.440	24.65
-1.30	0.050	0.057	361.57	410708.87	72394.31	4098.0273	175634.87	0.91	6.45	-1.540	24.90
-1.40	0.040	0.045	397.71	496915.75	86206.87	3876.2585	179511.12	0.93	7.09	-1.640	25.15
-1.50	0.032	0.036	428.77	577561.44	80645.69	2880.3987	182391.50	0.94	7.64	-1.740	25.40
-1.60	0.025	0.028	450.19	636709.56	59148.12	1678.0820	184069.56	0.95	8.03	-1.840	25.65
-1.70	0.020	0.023	467.65	687054.94	50345.37	1134.5728	185204.12	0.96	8.34	-1.940	25.90
-1.80	0.016	0.018	491.27	758210.69	71155.75	1273.7478	186477.81	0.96	8.76	-2.040	26.15
-1.90	0.013	0.014	518.40	844266.75	86056.06	1223.6448	187701.44	0.97	9.24	-2.140	26.40
-2.00	0.010	0.011	542.30	923908.12	79811.37	899.5247	188600.94	0.98	9.67	-2.240	26.65
-	-	-	-	-	-	-	193350.00	(1)	-	-	00

PHOTOMETRIC PARAMETERS OF NGC 4472

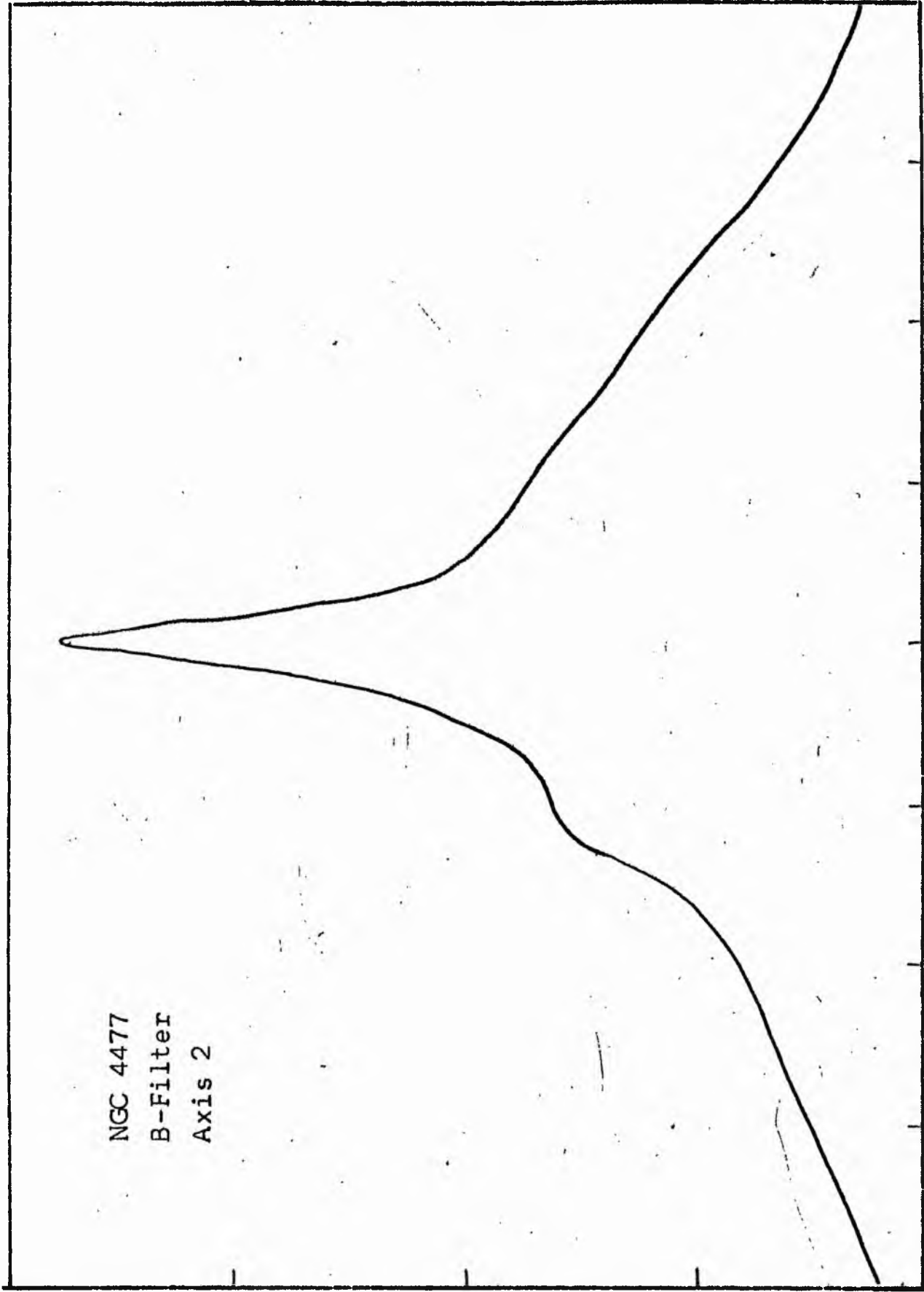
V-FILTER

Total luminosity	L_T	= 53.7
Total apparent magnitude	m_T	= 8.43
Apparent central surface brightness	μ_0	= 16.05
Major axis at threshold	$2a_m$	= 22.03
Minor axis at threshold	$2b_m$	= 15.98
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 13.83
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.92
Gradient of exponential component	$G(a)$	= -0.21
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.23
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.21
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.38
Axis ratio	b/a	= 0.90
Equivalent radius	r_1^*	= 0.29
Surface brightness	μ_1	= 18.20
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.17
Axis ratio	b/a	= 0.75
Equivalent radius	r_e^*	= 0.93
Surface brightness	μ_e	= 21.02
Mean surface brightness	μ_e'	= 10.27
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 3.53
Axis ratio	b/a	= 0.74
Equivalent radius	r_3^*	= 3.05
Surface brightness	μ_3	= 23.34
Concentration indices	$\{C_{21}$	= 3.17
	C_{32}	= 3.27

NGC 4477
B-Filter
Axis 1

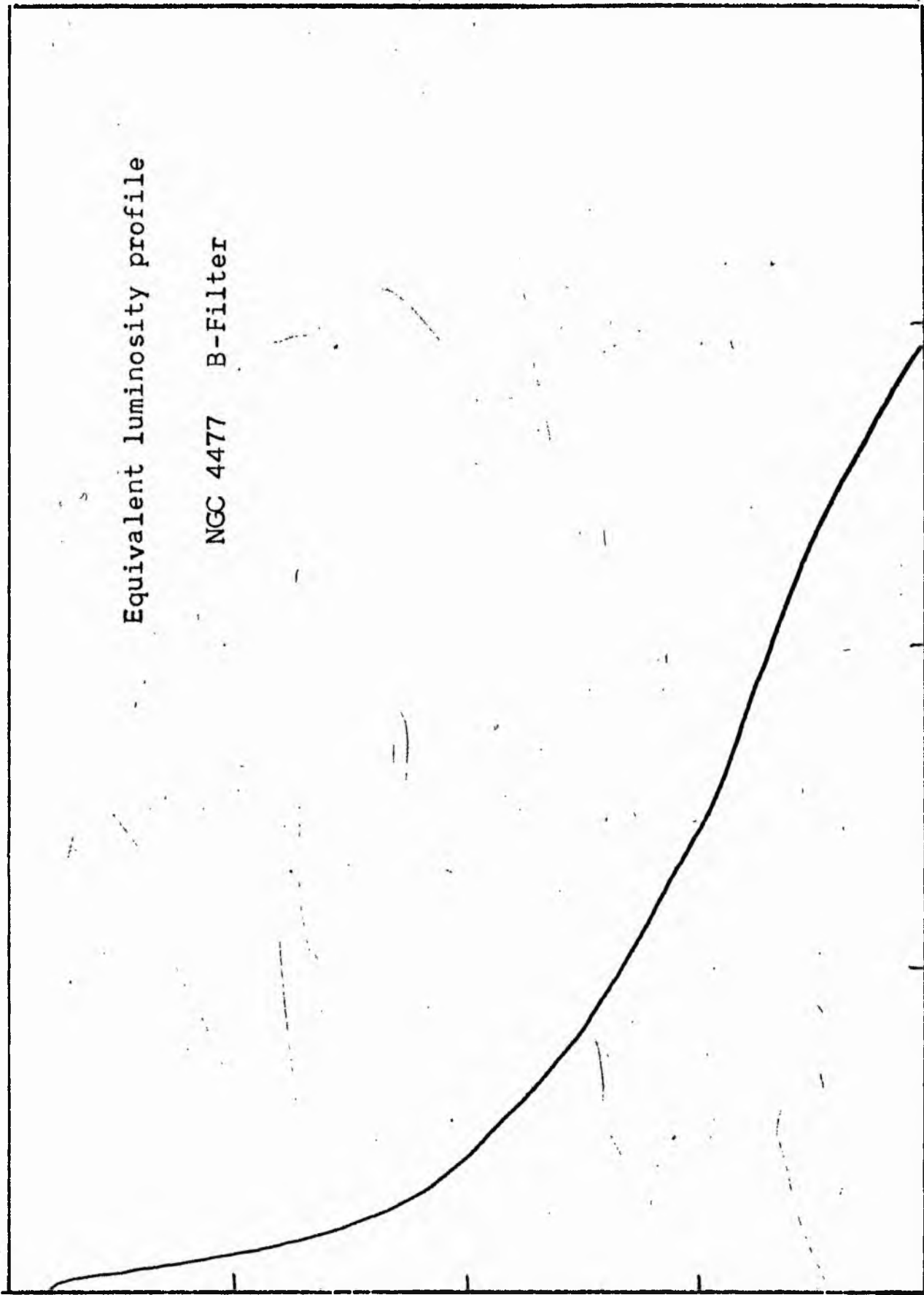


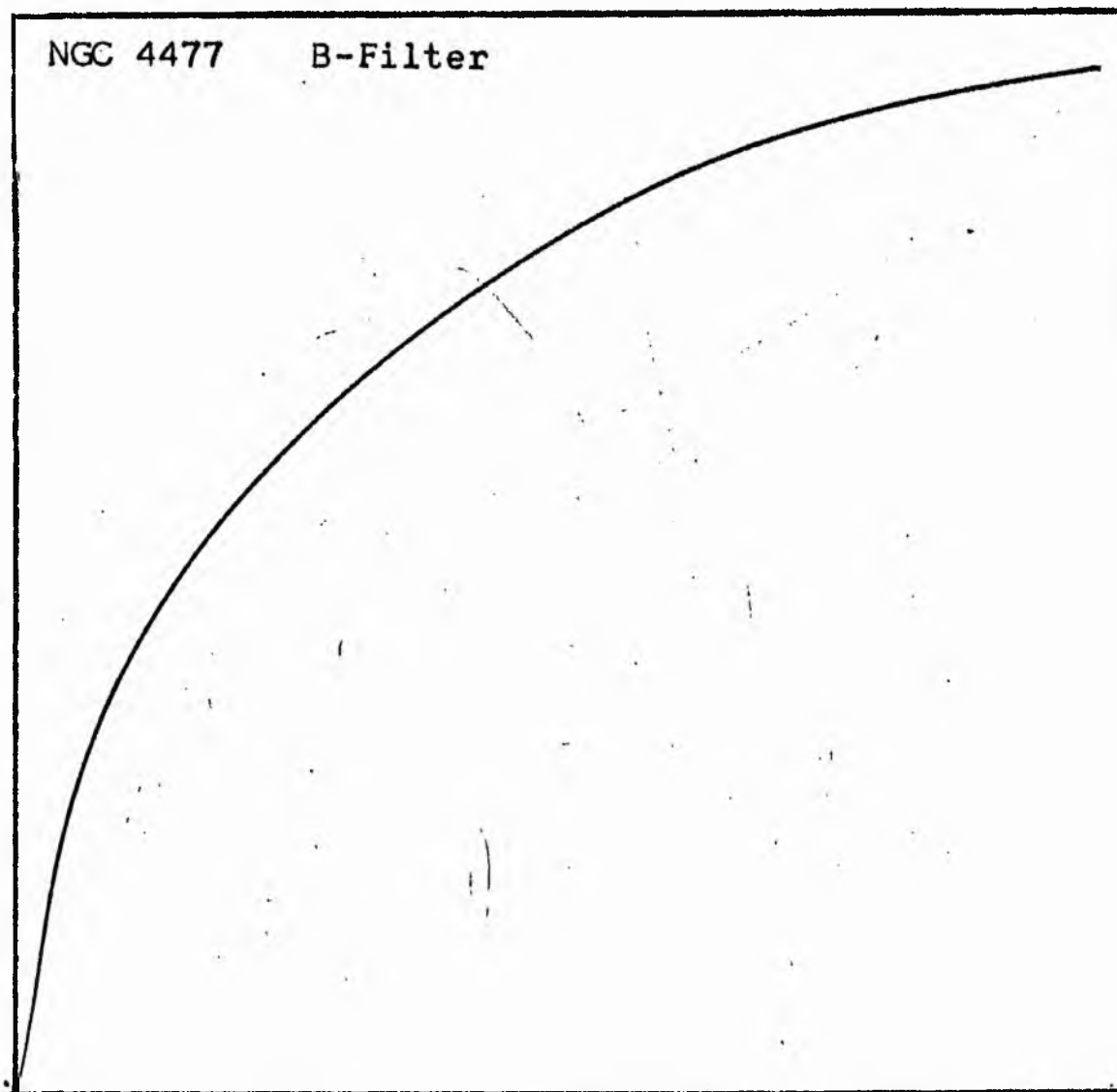
NGC 4477
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4477 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4477 B COLOUR

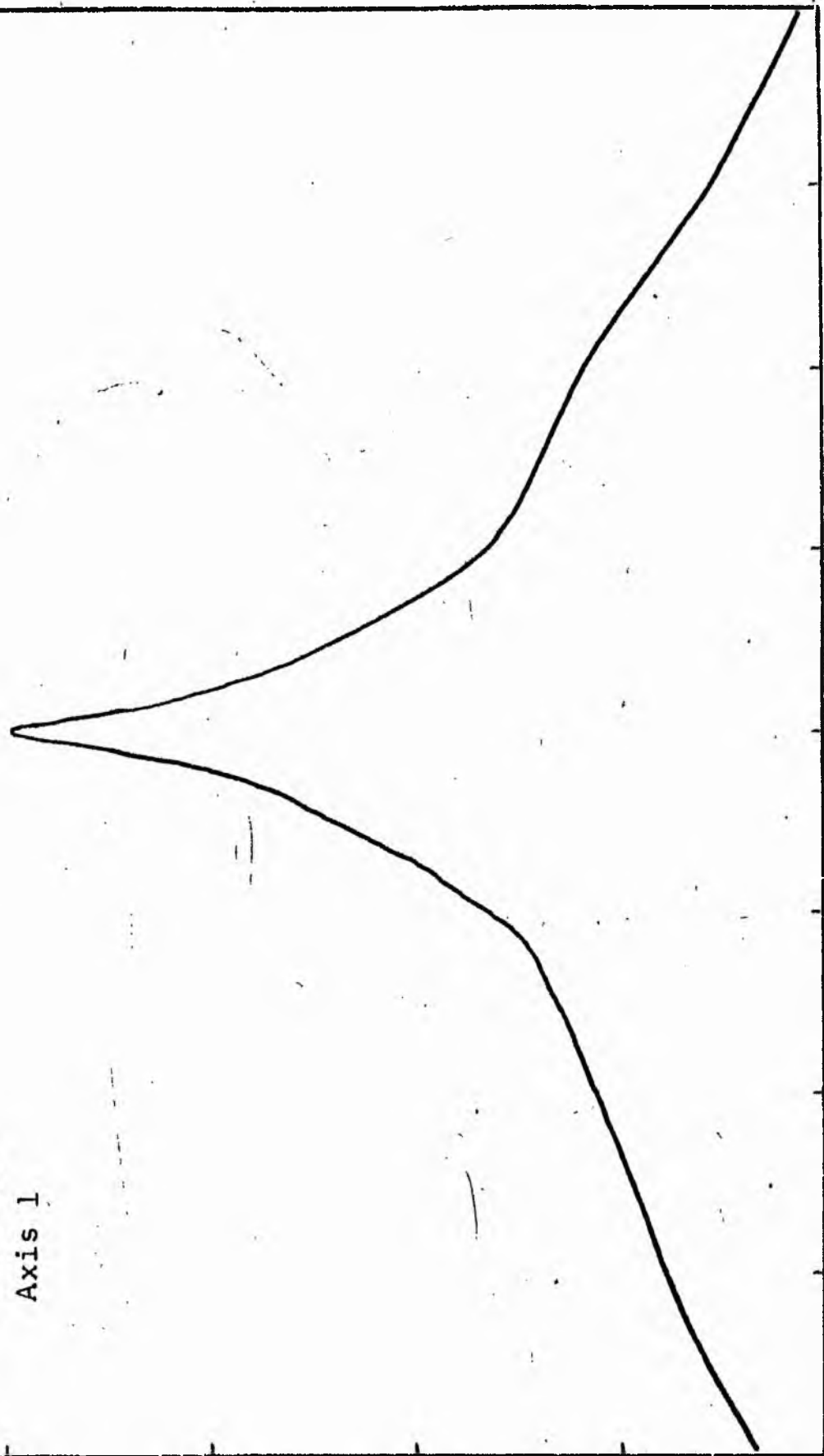
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.77	-98.884		0.0	0.0	13.85	755.0859	0.0	0.0	0.0	1.942	17.81
1.70	50.119	34.501	2.10	13.85	4.24	190.7020	755.09	0.03	0.05	1.872	17.99
1.60	39.811	44.965	2.40	18.10	39.99	1428.3972	945.79	0.04	0.06	1.772	18.24
1.50	31.623	35.717	4.30	58.09	11.31	320.8660	2374.19	0.09	0.11	1.672	18.49
1.40	25.119	28.371	4.70	69.40	36.29	817.7146	2695.05	0.11	0.12	1.572	18.74
1.30	19.953	22.536	5.80	105.68	52.68	943.0896	3512.77	0.14	0.15	1.472	18.99
1.20	15.849	17.901	7.10	158.37	47.75	678.9888	4455.85	0.17	0.19	1.372	19.24
1.10	12.589	14.219	8.10	206.12	42.73	482.5676	5134.84	0.20	0.21	1.272	19.49
1.00	10.000	11.295	8.90	248.85	65.31	585.9663	5617.41	0.22	0.23	1.172	19.74
0.90	7.943	8.972	10.00	314.16	86.99	619.9292	6203.37	0.24	0.26	1.072	19.99
0.80	6.310	7.126	11.30	401.15	89.72	507.8989	6823.30	0.27	0.30	0.972	20.24
0.70	5.012	5.661	12.50	490.87	98.77	444.1216	7331.20	0.29	0.33	0.872	20.49
0.60	3.981	4.496	13.70	589.65	107.82	385.0933	7775.32	0.30	0.36	0.772	20.74
0.50	3.162	3.572	14.90	697.46	157.83	447.7847	8160.41	0.32	0.39	0.672	20.99
0.40	2.512	2.837	16.50	855.30	231.57	521.8467	8608.19	0.34	0.43	0.572	21.24
0.30	1.945	2.254	18.60	1088.86	298.58	534.4761	9130.03	0.36	0.49	0.472	21.49
0.20	1.585	1.790	21.00	1385.44	484.94	689.5298	9664.50	0.38	0.55	0.372	21.74
0.10	1.259	1.422	24.40	1870.38	717.32	810.1819	10354.03	0.40	0.64	0.272	21.99
-0.00	1.000	1.129	28.70	2587.70	1173.29	1052.6274	11164.21	0.44	0.75	0.172	22.24
-0.10	0.794	0.897	34.60	3760.99	1165.53	830.6016	12216.84	0.48	0.91	0.072	22.49
-0.20	0.631	0.713	39.60	4926.52	1463.51	828.4495	13047.44	0.51	1.04	-0.028	22.74
-0.30	0.501	0.566	45.10	6390.03	2335.09	1049.9600	13875.89	0.54	1.18	-0.128	22.99
-0.40	0.398	0.450	52.70	8725.11	2509.34	896.2505	14925.84	0.58	1.38	-0.228	23.24
-0.50	0.316	0.357	59.80	11234.45	2326.20	659.9580	15822.09	0.62	1.57	-0.328	23.49
-0.60	0.251	0.284	65.70	13560.65	5503.81	1240.3203	16482.05	0.64	1.72	-0.428	23.74
-0.70	0.200	0.225	77.90	19064.46	4659.61	834.1028	17722.37	0.69	2.04	-0.528	23.99
-0.80	0.158	0.179	86.90	23724.07	4212.50	598.9763	18556.47	0.72	2.28	-0.628	24.24
-0.90	0.126	0.142	94.30	27936.57	6778.70	765.6257	19155.45	0.75	2.47	-0.728	24.49
-1.00	0.100	0.113	105.12	34715.26	8216.51	737.1543	19921.07	0.78	2.75	-0.828	24.74
-1.10	0.079	0.090	116.90	42931.77	11890.30	847.3528	20658.22	0.81	3.06	-0.928	24.99
-1.20	0.063	0.071	132.10	54822.07	15882.55	899.0681	21505.57	0.84	3.46	-1.028	25.24
-1.30	0.050	0.057	150.02	70704.62	12079.56	543.1553	22404.64	0.87	3.93	-1.128	25.49
-1.40	0.040	0.045	162.33	82784.19	8542.69	305.1182	22947.79	0.89	4.25	-1.228	25.74
-1.50	0.032	0.036	170.50	91326.87	9941.12	282.0393	23252.91	0.91	4.46	-1.328	25.99
-1.60	0.025	0.028	179.54	101268.00	8003.62	180.3688	23534.95	0.92	4.70	-1.428	26.24
-1.70	0.020	0.023	186.50	109271.62	20700.94	370.5657	23715.32	0.92	4.88	-1.528	26.49
-1.80	0.016	0.018	203.40	129972.56	14842.56	211.0497	24085.88	0.94	5.33	-1.628	26.74
-1.90	0.013	0.014	214.70	144815.12	12958.19	146.3594	24296.93	0.95	5.62	-1.728	26.99
-2.00	0.010	0.011	224.10	157773.31			24443.29	0.95	5.87	-1.828	27.24
-∞							25643.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4477

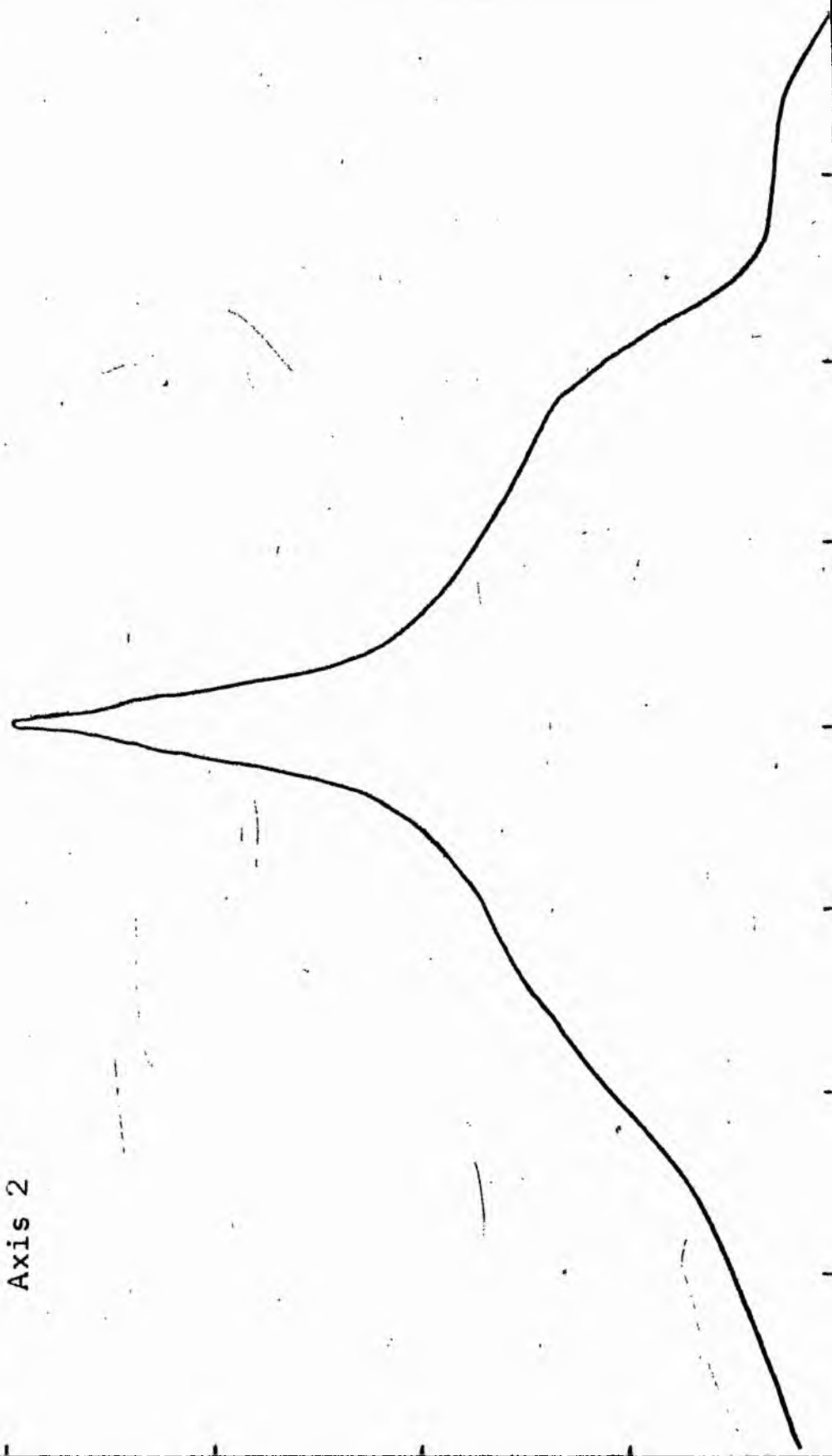
3-FILTER

Total luminosity	L_T	= 7.12
Total apparent magnitude	m_T	= 11.22
Apparent central surface brightness	μ_0	= 17.81
Major axis at threshold	$2a_m$	= 7.44
Minor axis at threshold	$2b_m$	= 7.21
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.67
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.81
Gradient of exponential component	$G(a)$	= -0.62
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.53
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.40
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.18
Axis ratio	b/a	= 0.81
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 20.07
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.62
Axis ratio	b/a	= 0.86
Equivalent radius	r_e^*	= 0.64
Surface brightness	μ_e	= 22.67
Mean surface brightness	μ_e'	= 12.21
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.46
Axis ratio	b/a	= 0.97
Equivalent radius	r_3^*	= 1.59
Surface brightness	μ_3	= 24.49
Concentration indices	$\left\{ \begin{array}{l} C_{21} \\ C_{32} \end{array} \right.$	$\begin{array}{l} = 3.67 \\ = 2.50 \end{array}$

NGC 4477
V-Filter
Axis 1

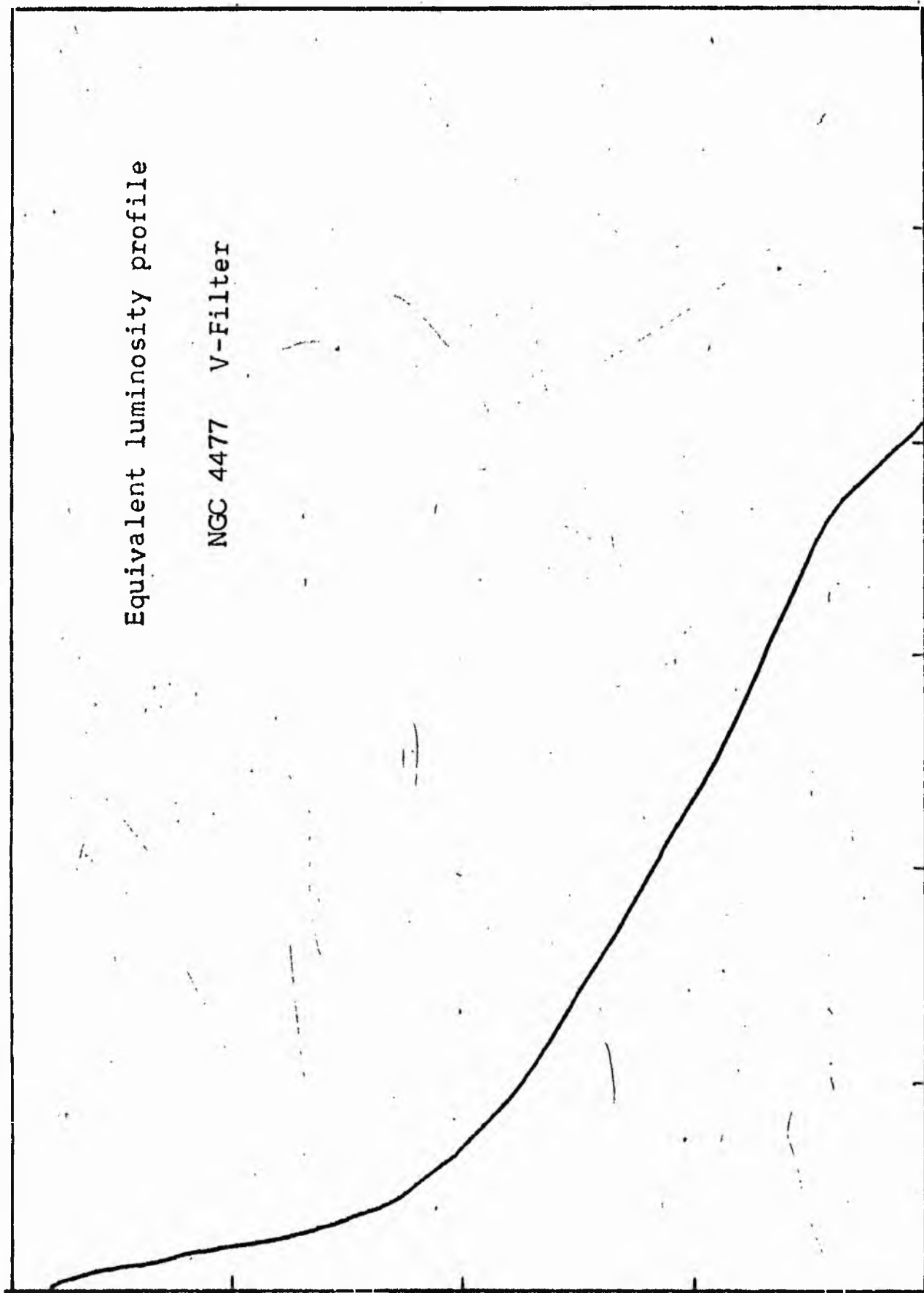


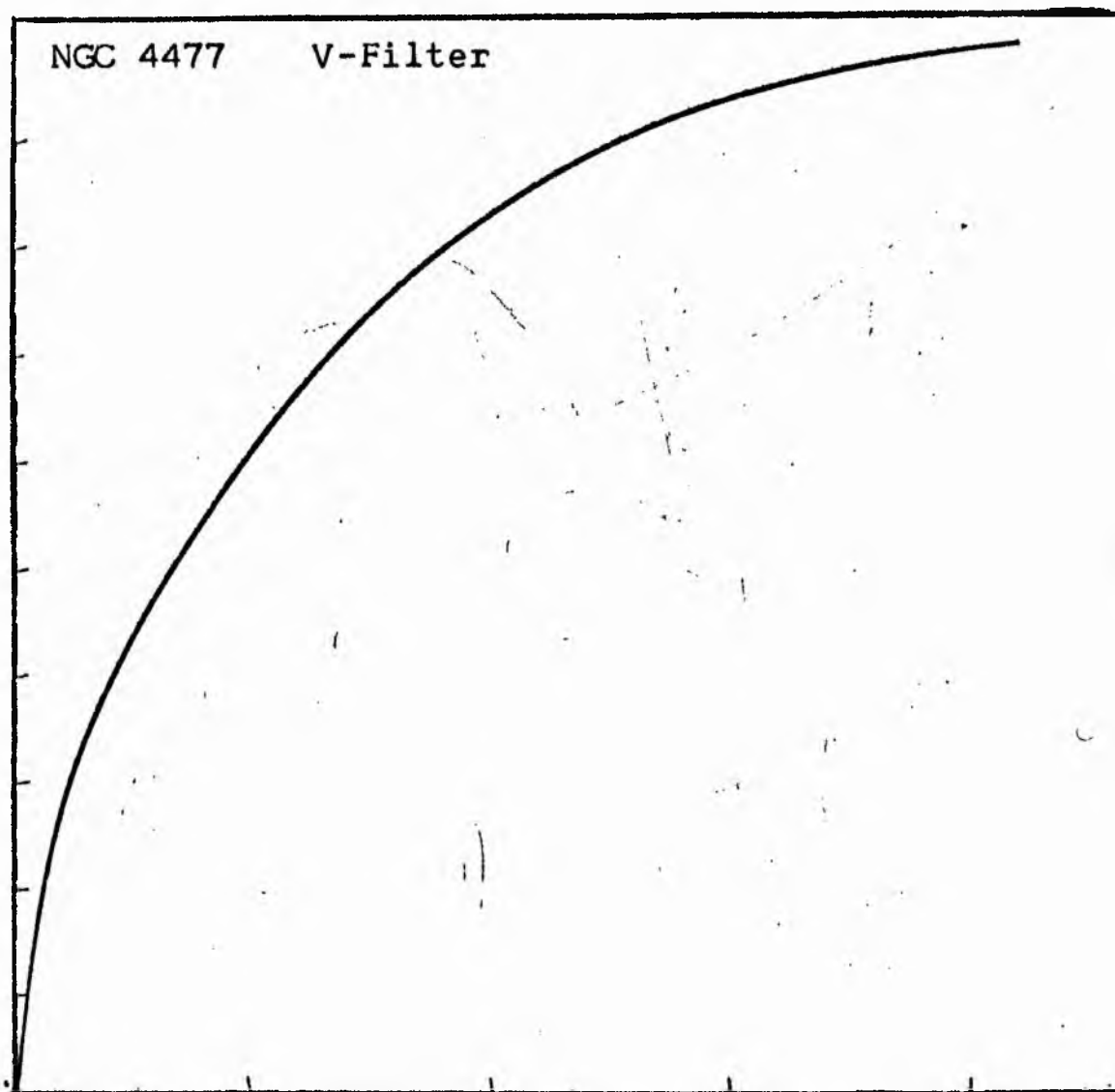
NGC 4477
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4477 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4477
V COLOUR

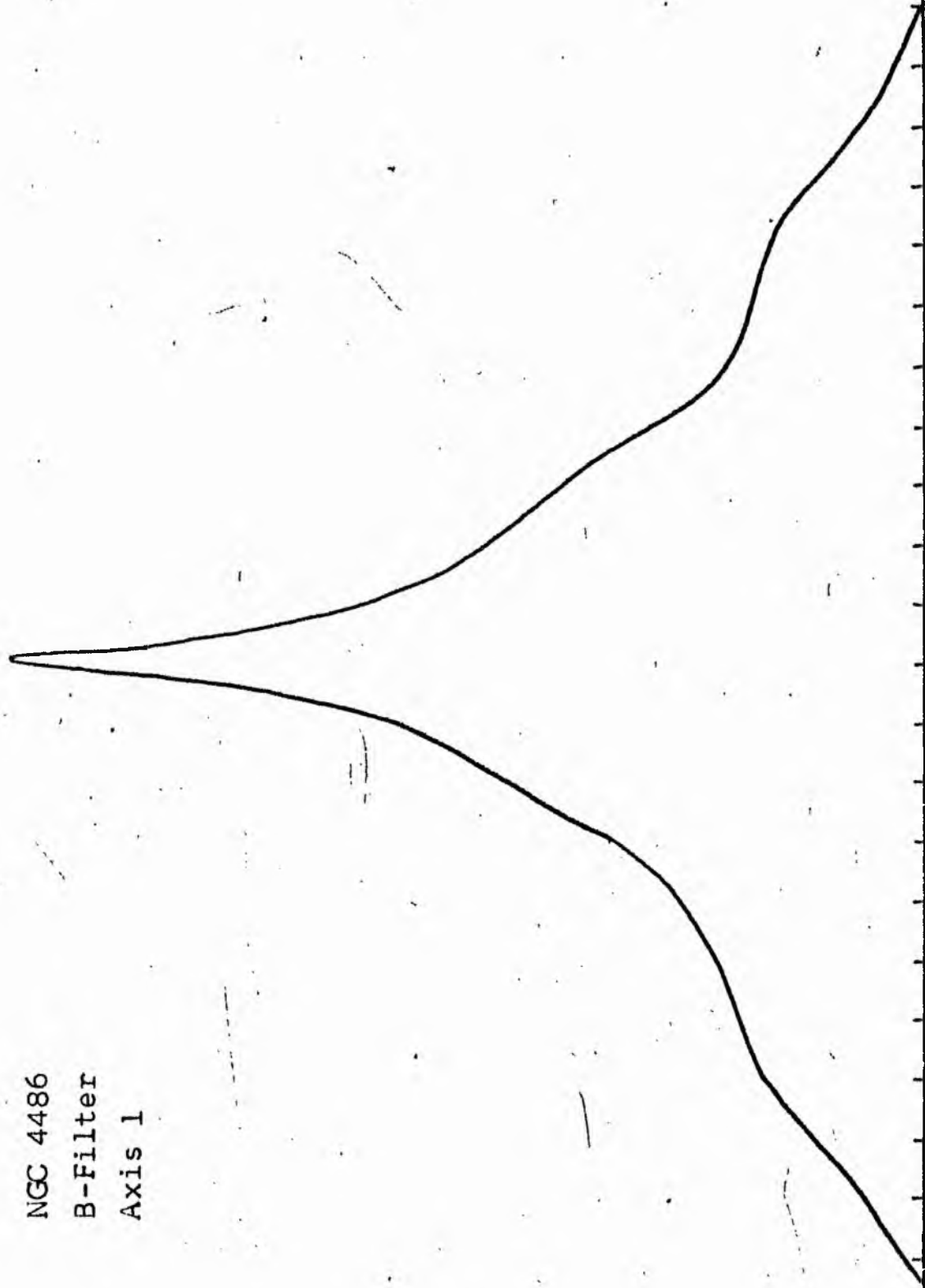
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	ΔA
1.88	75.858		0.0	0.0			0.0	0.0	0.0	1.910	16.77
1.80	63.096	69.477	0.80	2.01	2.01	139.6909	139.69	0.00	0.02	1.830	16.97
1.70	50.119	56.607	1.80	10.18	8.17	462.3740	602.06	0.02	0.05	1.730	17.22
1.60	39.811	44.965	3.20	32.17	21.99	988.8218	1590.89	0.05	0.09	1.630	17.47
1.50	31.623	35.717	4.60	66.48	34.31	1225.2996	2816.19	0.10	0.13	1.530	17.72
1.40	25.119	28.371	5.50	95.03	28.56	810.1843	3626.37	0.13	0.16	1.430	17.97
1.30	19.952	22.536	6.10	116.90	21.87	492.7512	4119.12	0.14	0.17	1.330	18.22
1.20	15.849	17.901	8.30	216.42	99.53	1781.5137	5900.69	0.20	0.24	1.230	18.47
1.10	12.589	14.219	9.10	260.16	43.73	621.8064	6522.50	0.23	0.26	1.130	18.72
1.00	10.000	11.295	10.10	320.47	60.32	681.2698	7203.77	0.25	0.29	1.030	18.97
0.90	7.943	8.972	11.40	408.28	87.81	787.7717	7991.54	0.28	0.33	0.930	19.22
0.80	6.310	7.126	12.40	483.05	74.77	532.8386	8524.37	0.29	0.35	0.830	19.47
0.70	5.012	5.661	13.60	581.07	98.02	554.8457	9079.21	0.31	0.39	0.730	19.72
0.60	3.981	4.496	14.90	697.46	116.40	523.3669	9602.58	0.33	0.42	0.630	19.97
0.50	3.162	3.572	16.20	824.48	127.01	453.6479	10056.22	0.35	0.46	0.530	20.22
0.40	2.512	2.837	17.95	1012.23	187.75	332.6560	10588.87	0.37	0.51	0.430	20.47
0.30	1.995	2.254	20.19	1280.62	268.40	604.8481	11193.72	0.39	0.58	0.330	20.72
0.20	1.585	1.790	22.76	1627.40	346.77	620.7490	11814.47	0.41	0.65	0.230	20.97
0.10	1.259	1.422	27.10	2307.22	679.82	966.6294	12781.10	0.44	0.77	0.130	21.22
-0.00	1.000	1.129	33.30	3483.68	1176.46	1328.7612	14109.86	0.49	0.95	0.030	21.47
-0.10	0.794	0.897	39.17	4820.11	1336.43	1198.9893	15308.84	0.53	1.12	-0.070	21.72
-0.20	0.631	0.713	45.40	6475.32	1655.21	1179.5659	16488.41	0.57	1.29	-0.170	21.97
-0.30	0.501	0.566	51.20	8235.49	1760.17	996.3772	17484.78	0.60	1.46	-0.270	22.22
-0.40	0.398	0.450	61.96	12060.70	3825.21	1719.9819	19204.76	0.66	1.77	-0.370	22.47
-0.50	0.316	0.357	71.69	16146.07	4085.37	1459.1506	20663.91	0.71	2.04	-0.470	22.72
-0.60	0.251	0.284	80.40	20307.75	4161.68	1180.6951	21844.60	0.75	2.29	-0.570	22.97
-0.70	0.200	0.225	88.12	24394.89	4087.14	921.0618	22765.66	0.79	2.51	-0.670	23.22
-0.80	0.158	0.179	93.95	27729.59	3334.70	596.9333	23362.59	0.81	2.68	-0.770	23.47
-0.90	0.126	0.142	100.66	31831.97	4102.38	583.3179	23945.91	0.83	2.87	-0.870	23.72
-1.00	0.100	0.113	107.11	36042.07	4210.10	475.5127	24421.42	0.84	3.05	-0.970	23.97
-1.10	0.079	0.090	120.45	45578.86	9536.79	855.6028	25277.02	0.87	3.43	-1.070	24.22
-1.20	0.063	0.071	132.70	55321.22	9742.36	694.2805	25971.30	0.90	3.78	-1.170	24.47
-1.30	0.050	0.057	146.80	67702.06	12380.84	700.8442	26672.14	0.92	4.19	-1.270	24.72
-1.40	0.040	0.045	155.94	76394.94	8692.87	390.8726	27063.01	0.93	4.45	-1.370	24.97
-1.50	0.032	0.036	174.50	95662.25	19267.31	688.1667	27751.18	0.96	4.98	-1.470	25.22
-1.60	0.025	0.028	180.33	102161.12	6498.87	184.3789	27935.55	0.96	5.14	-1.570	25.47
-1.70	0.020	0.023	184.70	107172.56	5011.44	112.9369	28048.49	0.97	5.27	-1.670	25.72
-1.80	0.016	0.018	190.50	114009.19	6836.62	122.3816	28170.87	0.97	5.43	-1.770	25.97
-1.90	0.013	0.014	195.60	120195.31	6186.12	87.9617	28258.83	0.98	5.58	-1.870	26.22
-2.00	0.010	0.011	202.50	128824.94	8629.62	97.4692	28356.30	0.98	5.77	-1.970	26.47
							28956.00	(1)			00

PHOTOMETRIC PARAMETERS OF NGC 4477

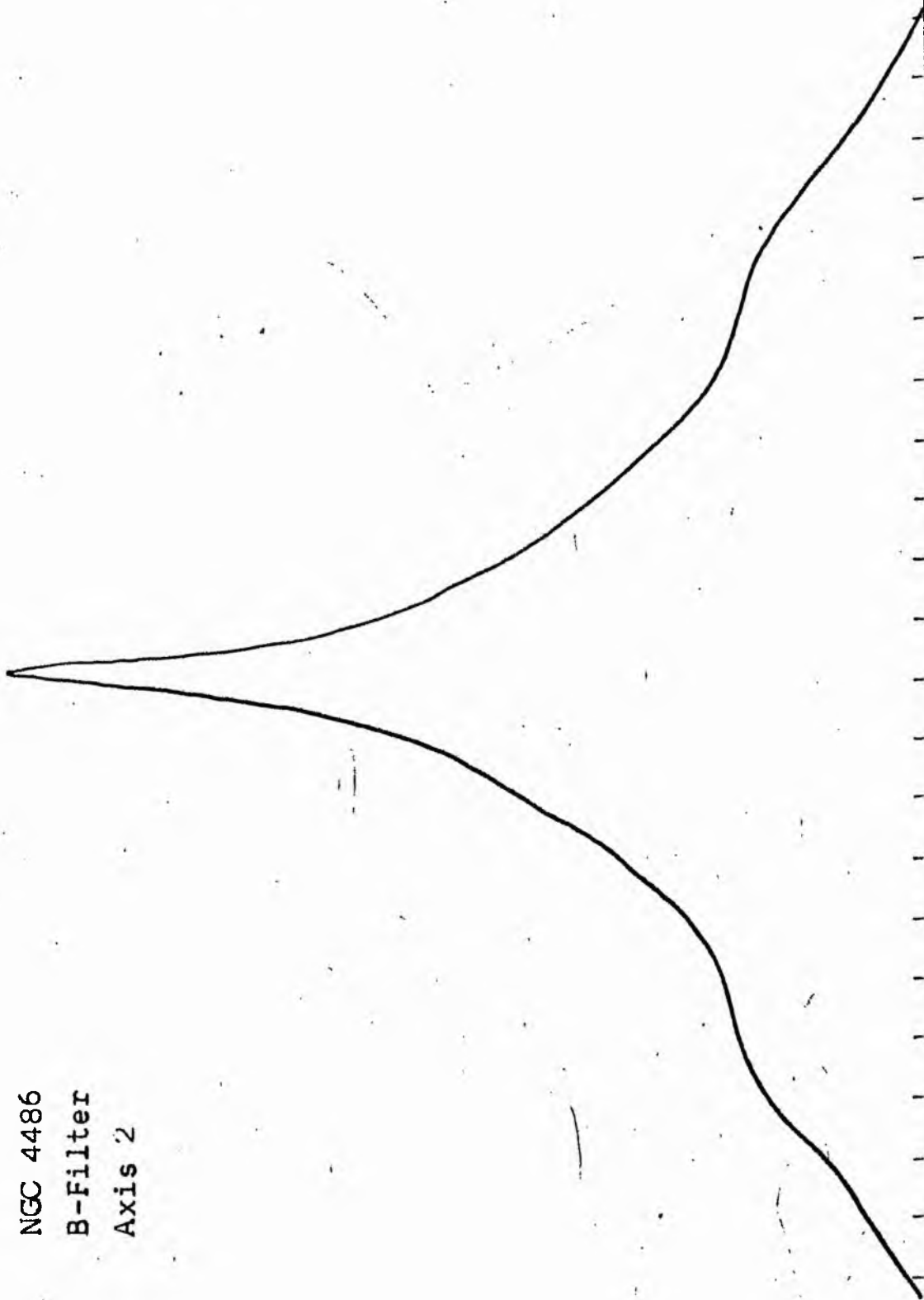
V-FILTER

Total luminosity	L_T	= 8.04
Total apparent magnitude	m_T	= 10.32
Apparent central surface brightness	μ_o	= 16.77
Major axis at threshold	$2a_m$	= 6.86
Minor axis at threshold	$2b_m$	= 6.52
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.20
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.93
Gradient of exponential component	$G(a)$	= -0.73
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.65
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.38
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.21
Axis ratio	b/a	= 0.99
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 18.97
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.61
Axis ratio	b/a	= 0.92
Equivalent radius	r_e^*	= 0.58
Surface brightness	μ_e	= 21.53
Mean surface brightness	μ_e'	= 11.12
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.26
Axis ratio	b/a	= 1.03
Equivalent radius	r_3^*	= 1.32
Surface brightness	μ_3	= 22.97
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.45 \\ = 2.26 \end{cases}$

NGC 4486
B-Filter
Axis 1

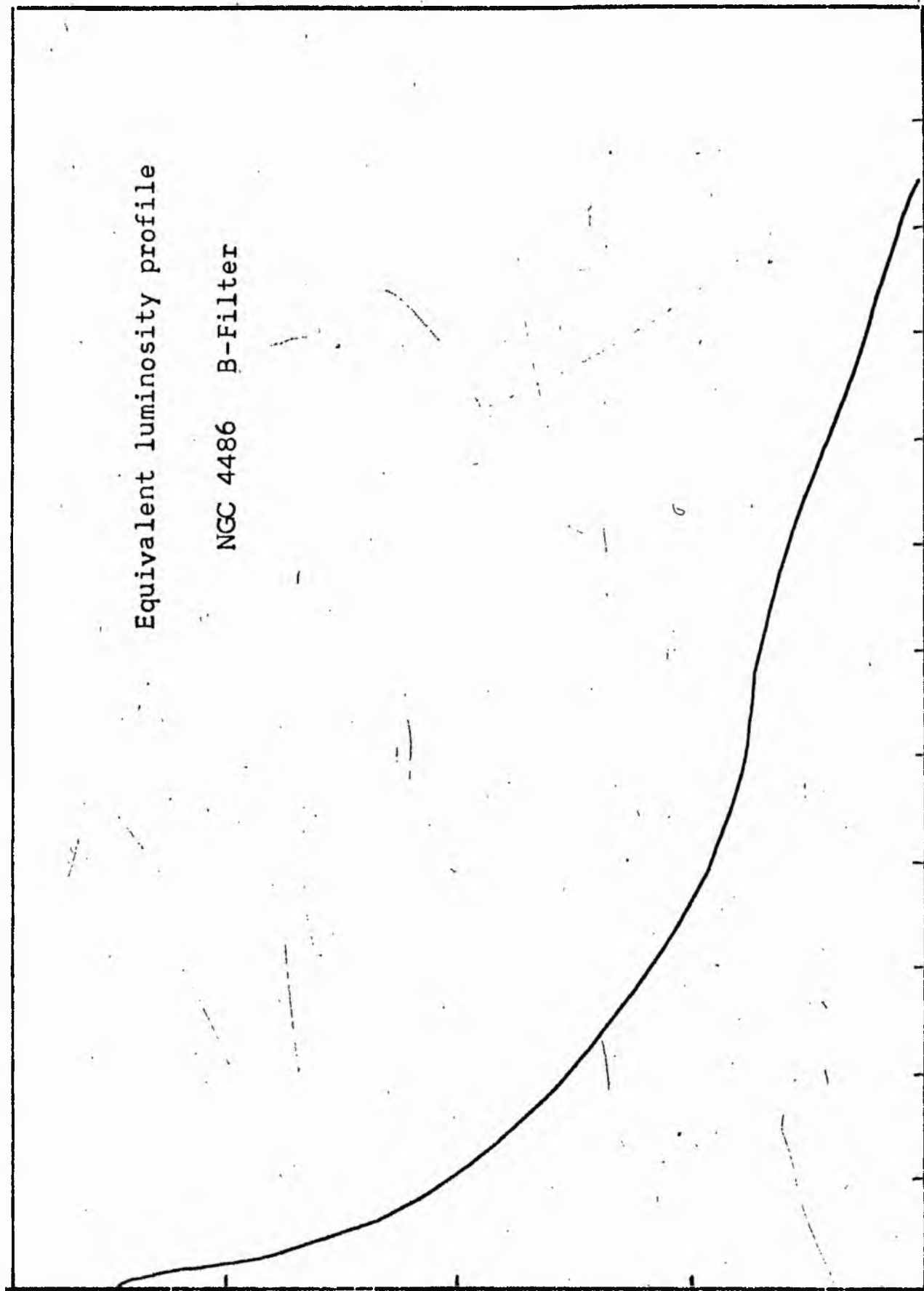


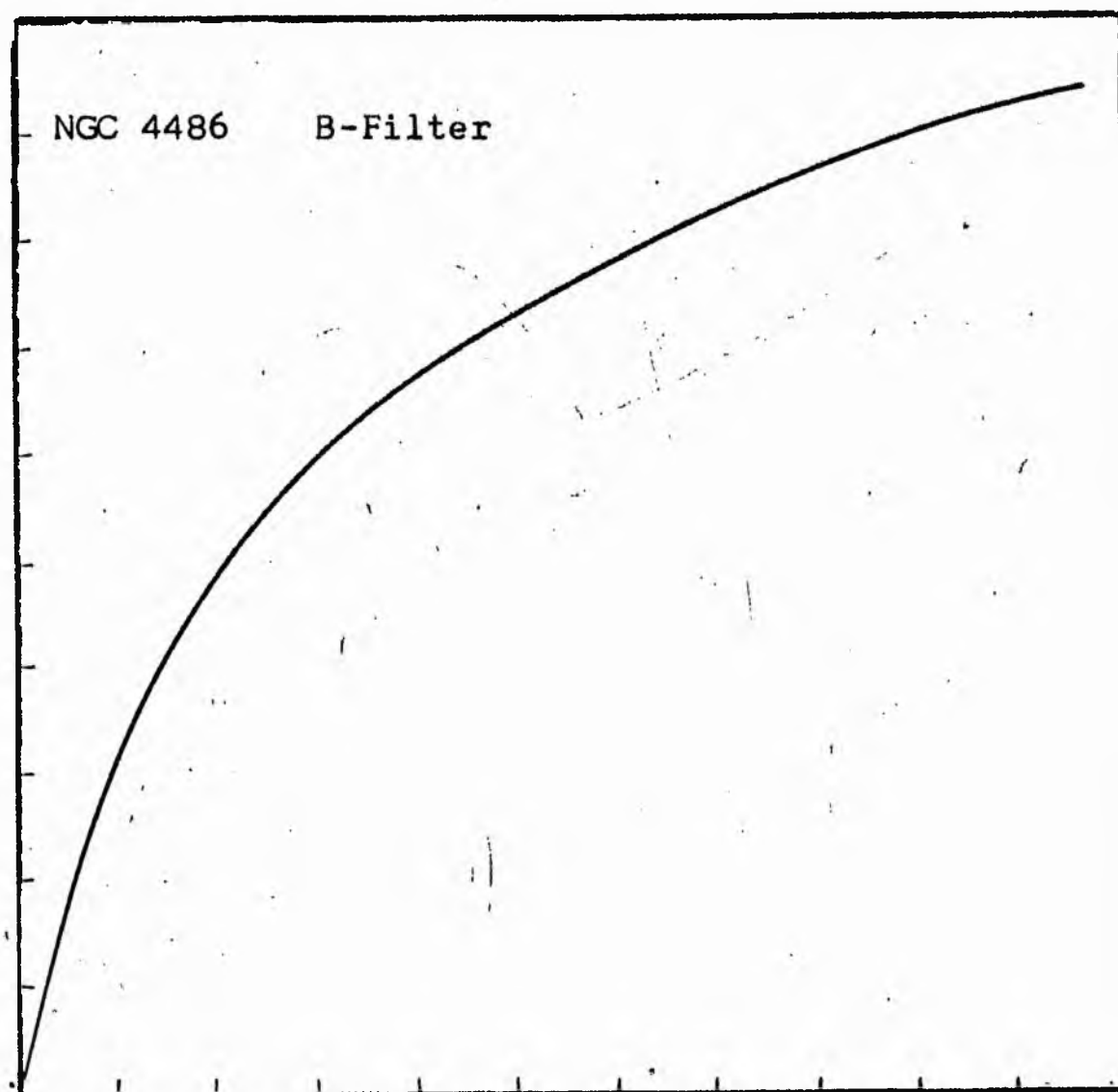
NGC 4486
B-Filter
Axis '2



Equivalent luminosity profile

NGC 4486 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

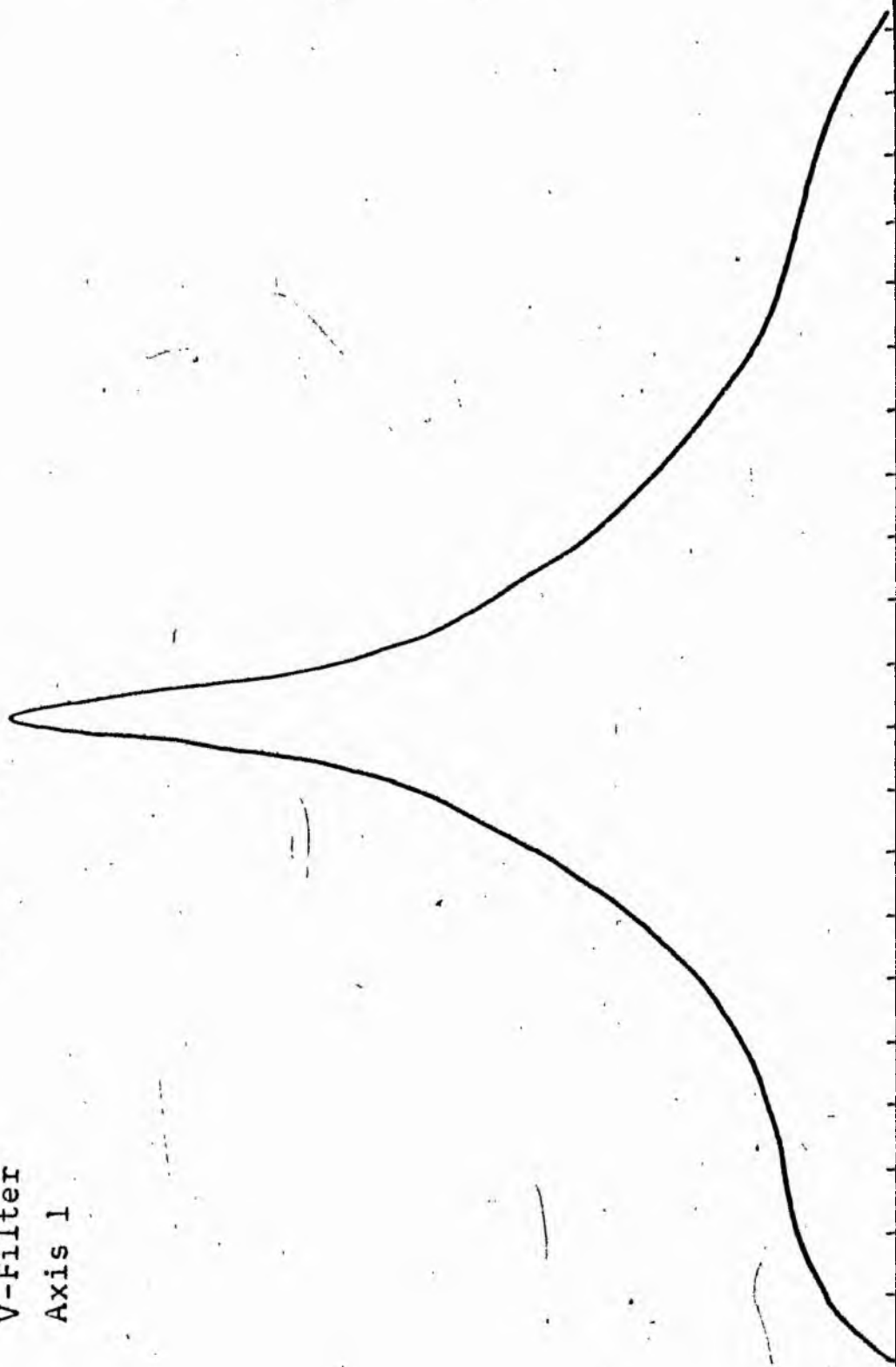
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	u
1.43	26.915		0.0	0.0			0.0	0.0	0.0	1.833	18.43
1.40	25.119	26.017	4.33	58.90	58.90	1532.4402	1532.44	0.02	0.04	1.803	18.51
1.30	19.953	22.536	7.19	162.41	103.51	2332.5947	3865.03	0.04	0.07	1.703	18.76
1.20	15.849	17.901	9.32	272.89	116.48	1977.6392	5842.67	0.06	0.09	1.603	19.01
1.10	12.589	14.219	11.69	429.32	156.43	2224.3042	8066.97	0.09	0.11	1.503	19.26
1.00	10.000	11.295	14.10	624.50	195.26	2205.4043	10272.37	0.11	0.14	1.403	19.51
0.90	7.943	8.972	16.05	809.20	164.70	1657.0632	11929.44	0.13	0.16	1.303	19.76
0.80	6.310	7.126	18.60	1086.06	277.58	1978.1682	13907.61	0.15	0.18	1.203	20.01
0.70	5.012	5.661	20.78	1356.57	269.70	1526.6997	15434.30	0.17	0.20	1.103	20.26
0.60	3.981	4.496	23.90	1794.51	437.94	1969.1885	17403.49	0.19	0.23	1.003	20.51
0.50	3.162	3.572	27.26	2334.54	546.03	1928.8147	19332.30	0.21	0.27	0.903	20.76
0.40	2.512	2.837	31.17	3052.27	717.73	2036.2593	21368.56	0.23	0.31	0.803	21.01
0.30	1.995	2.254	35.99	4069.24	1016.97	2291.8044	23660.36	0.26	0.35	0.703	21.26
0.20	1.585	1.790	42.46	5663.82	1594.58	2854.4138	26514.77	0.29	0.42	0.603	21.51
0.10	1.259	1.422	48.87	7502.99	1839.17	2615.1267	29129.90	0.32	0.48	0.503	21.76
-0.00	1.000	1.129	54.83	9444.66	1941.66	2193.0305	31322.93	0.34	0.54	0.403	22.01
-0.10	0.794	0.897	64.37	13017.18	3572.52	3205.1267	34528.05	0.38	0.63	0.303	22.26
-0.20	0.631	0.713	77.23	18737.93	5726.75	4676.8423	38604.89	0.42	0.76	0.203	22.51
-0.30	0.501	0.566	89.36	25086.26	6340.33	3593.6016	42198.49	0.46	0.88	0.103	22.76
-0.40	0.398	0.450	101.39	32295.35	7209.09	3241.5413	45440.03	0.50	1.00	0.003	23.01
-0.50	0.316	0.357	114.88	41460.89	9165.54	3273.6248	48713.65	0.53	1.13	-0.097	23.26
-0.60	0.251	0.284	129.65	52807.40	11346.51	3219.0918	51932.74	0.57	1.27	-0.197	23.51
-0.70	0.200	0.225	139.33	60987.25	8179.85	1843.3857	53776.12	0.59	1.37	-0.297	23.76
-0.80	0.158	0.179	151.62	72220.87	11233.62	2010.9021	55787.02	0.61	1.49	-0.397	24.01
-0.90	0.126	0.142	166.98	87594.87	15374.00	2186.0396	57973.06	0.64	1.64	-0.497	24.26
-1.00	0.100	0.113	185.06	107590.75	19995.87	2258.4583	60231.52	0.66	1.82	-0.597	24.51
-1.10	0.079	0.090	206.69	134211.19	26620.44	2388.2922	62619.81	0.69	2.03	-0.697	24.76
-1.20	0.063	0.071	237.49	177190.50	42979.31	3062.8967	65682.69	0.72	2.33	-0.797	25.01
-1.30	0.050	0.057	309.96	301828.94	124638.44	7055.4570	72738.12	0.80	3.04	-0.897	25.26
-1.40	0.040	0.045	350.30	385504.62	83675.69	3762.4695	76500.56	0.84	3.44	-0.997	25.51
-1.50	0.032	0.036	376.05	444263.44	58758.81	2098.6855	78599.19	0.86	3.69	-1.097	25.76
-1.60	0.025	0.028	394.67	489348.06	45084.62	1279.0972	79878.25	0.88	3.88	-1.197	26.01
-1.70	0.020	0.023	424.90	567182.75	77834.69	1754.0781	81632.31	0.90	4.17	-1.297	26.26
-1.80	0.016	0.018	460.50	666206.94	99024.19	1772.6282	83404.94	0.92	4.52	-1.397	26.51
-1.90	0.013	0.014	494.56	768400.25	102193.31	1453.1125	84858.00	0.93	4.86	-1.497	26.76
-2.00	0.010	0.011	514.75	832420.19	64019.94	723.0901	85581.06	0.94	5.06	-1.597	27.01
-∞							91099.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4486

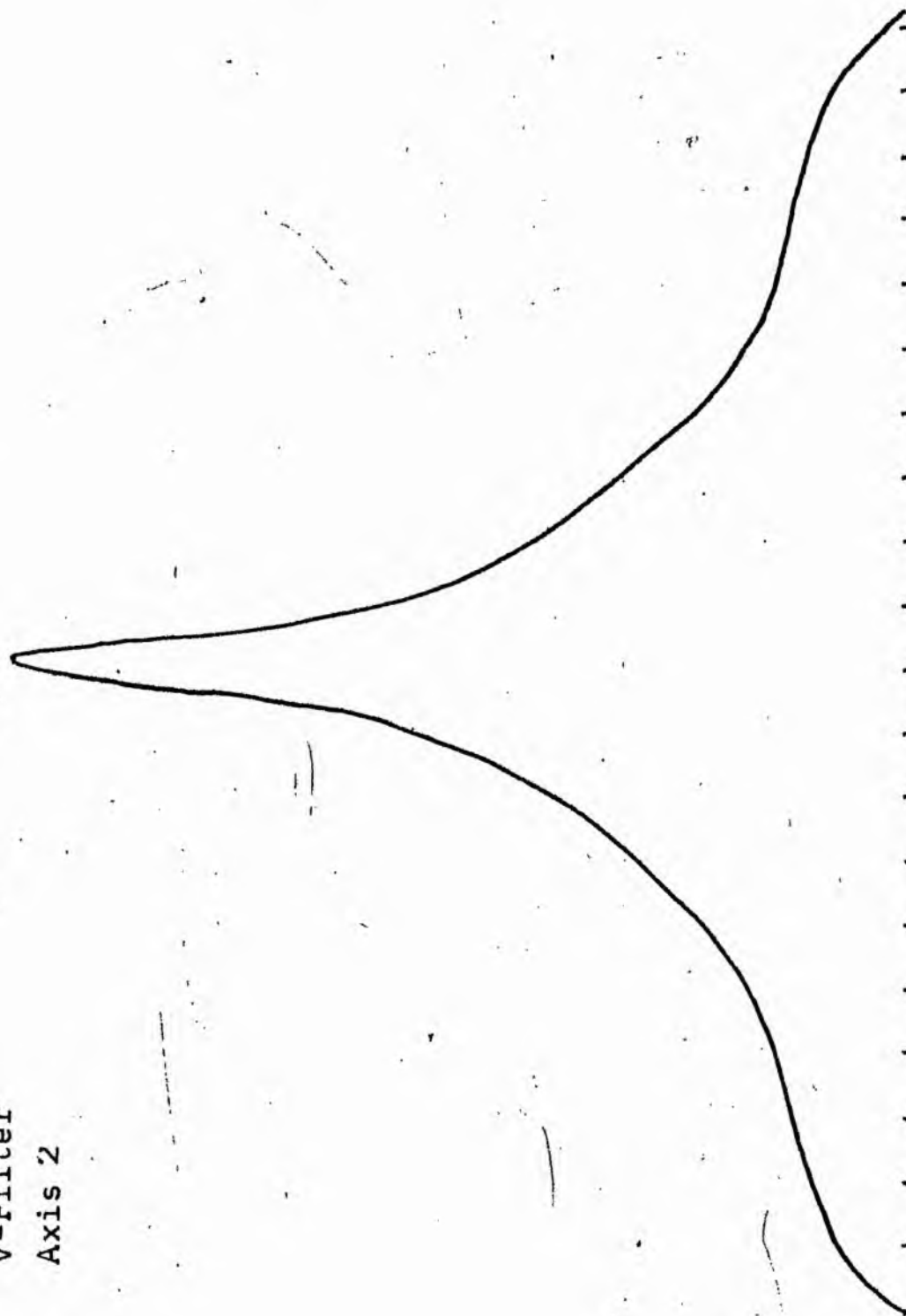
B-FILTER

Total luminosity	L_T	= 25.31
Total apparent magnitude	m_T	= 9.62
Apparent central surface brightness	μ_o	= 18.43
Major axis at threshold	$2a_m$	= 17.68
Minor axis at threshold	$2b_m$	= 17.20
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 7.82
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.72
Gradient of exponential component	$G(a)$	= -0.22
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.21
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.29
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.65
Axis ratio	b/a	= 0.90
Equivalent radius	r_1^*	= 0.56
Surface brightness	μ_1	= 21.18
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.82
Axis ratio	b/a	= 0.84
Equivalent radius	r_e^*	= 1.70
Surface brightness	μ_e	= 23.01
Mean surface brightness	μ_e'	= 12.77
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 4.36
Axis ratio	b/a	= 0.99
Equivalent radius	r_3^*	= 4.41
Surface brightness	μ_3	= 25.10
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.99 \\ = 2.60 \end{cases}$

NGC 4486
V-Filter
Axis 1

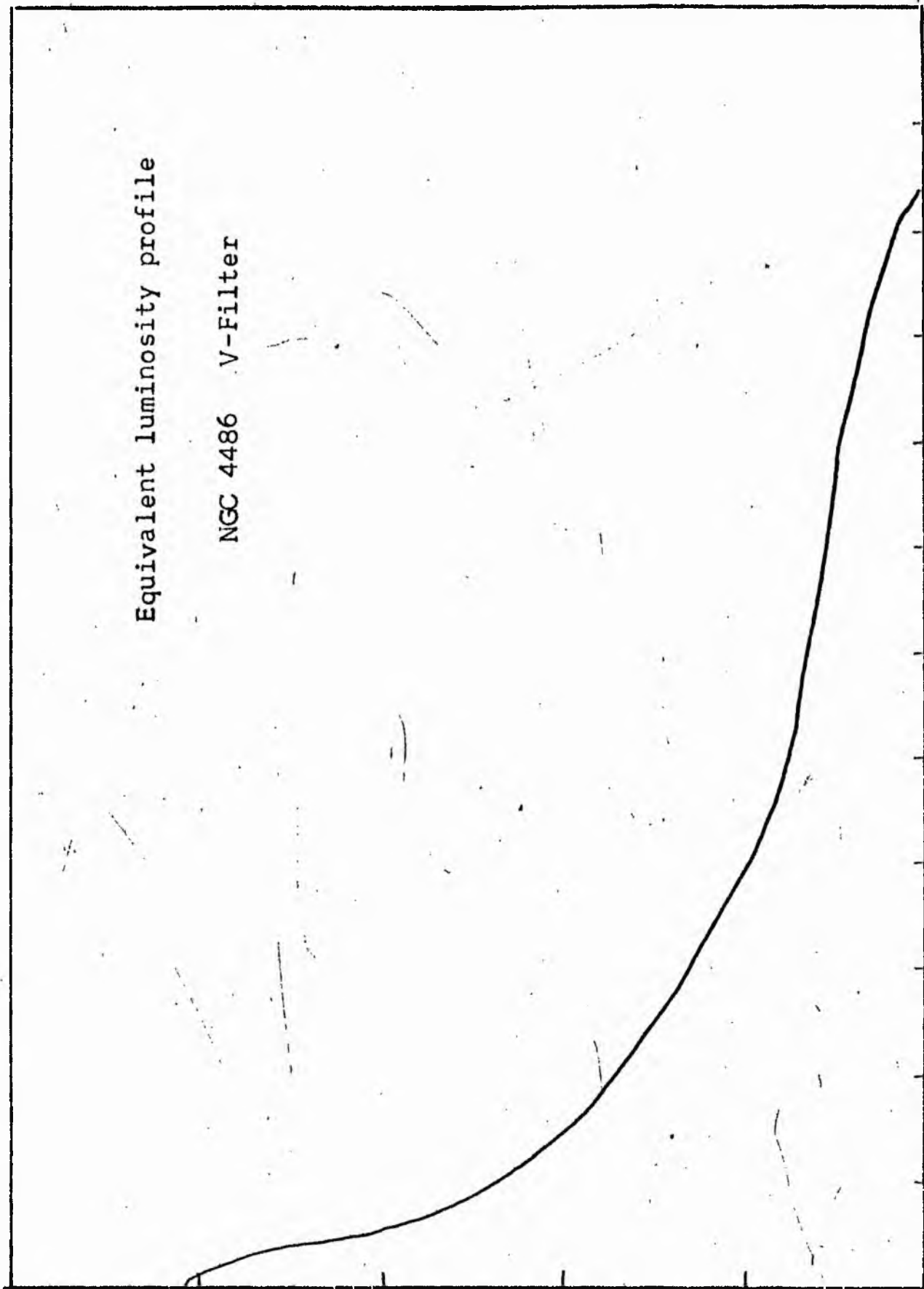


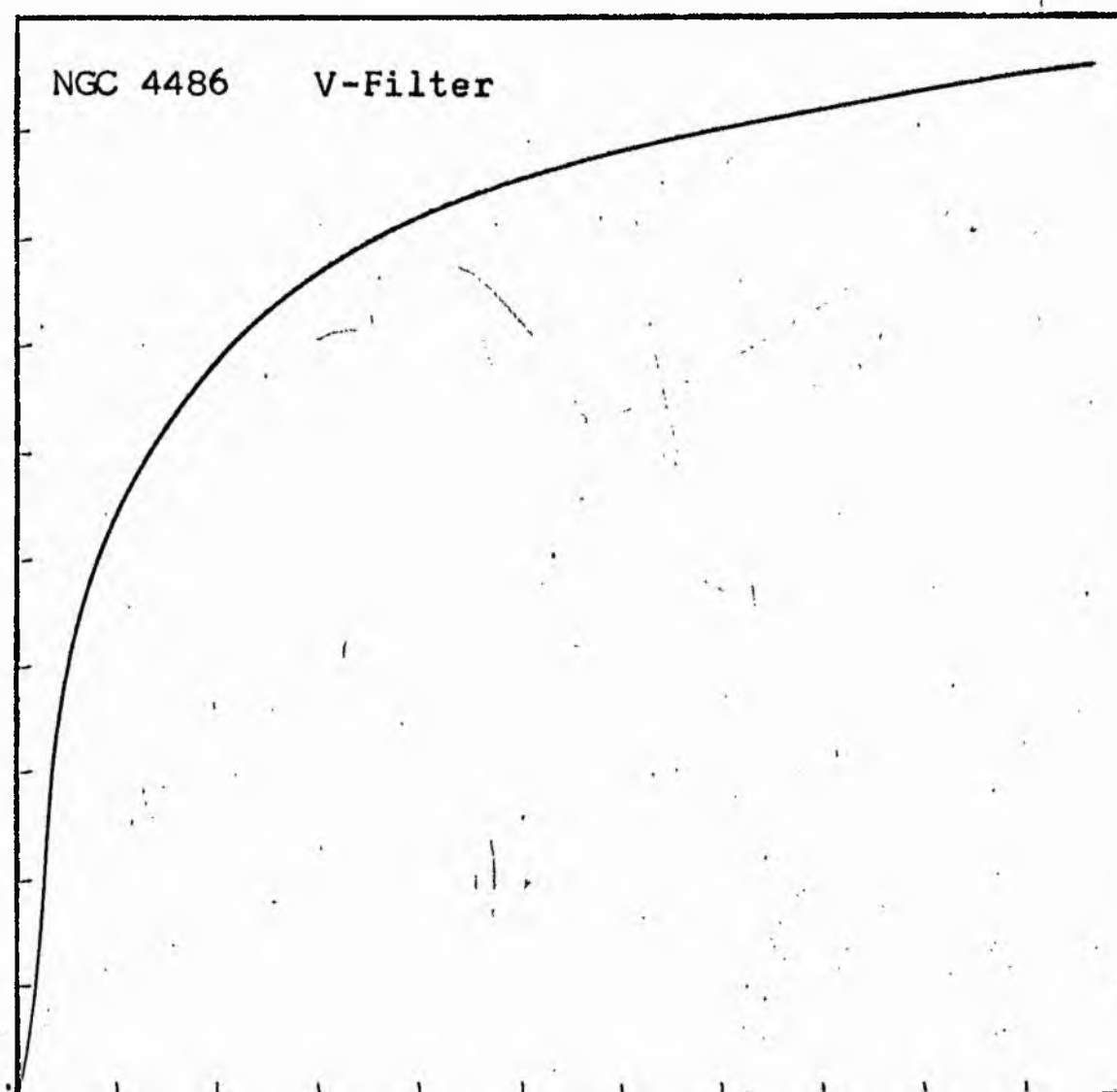
NGC 4486
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4486 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

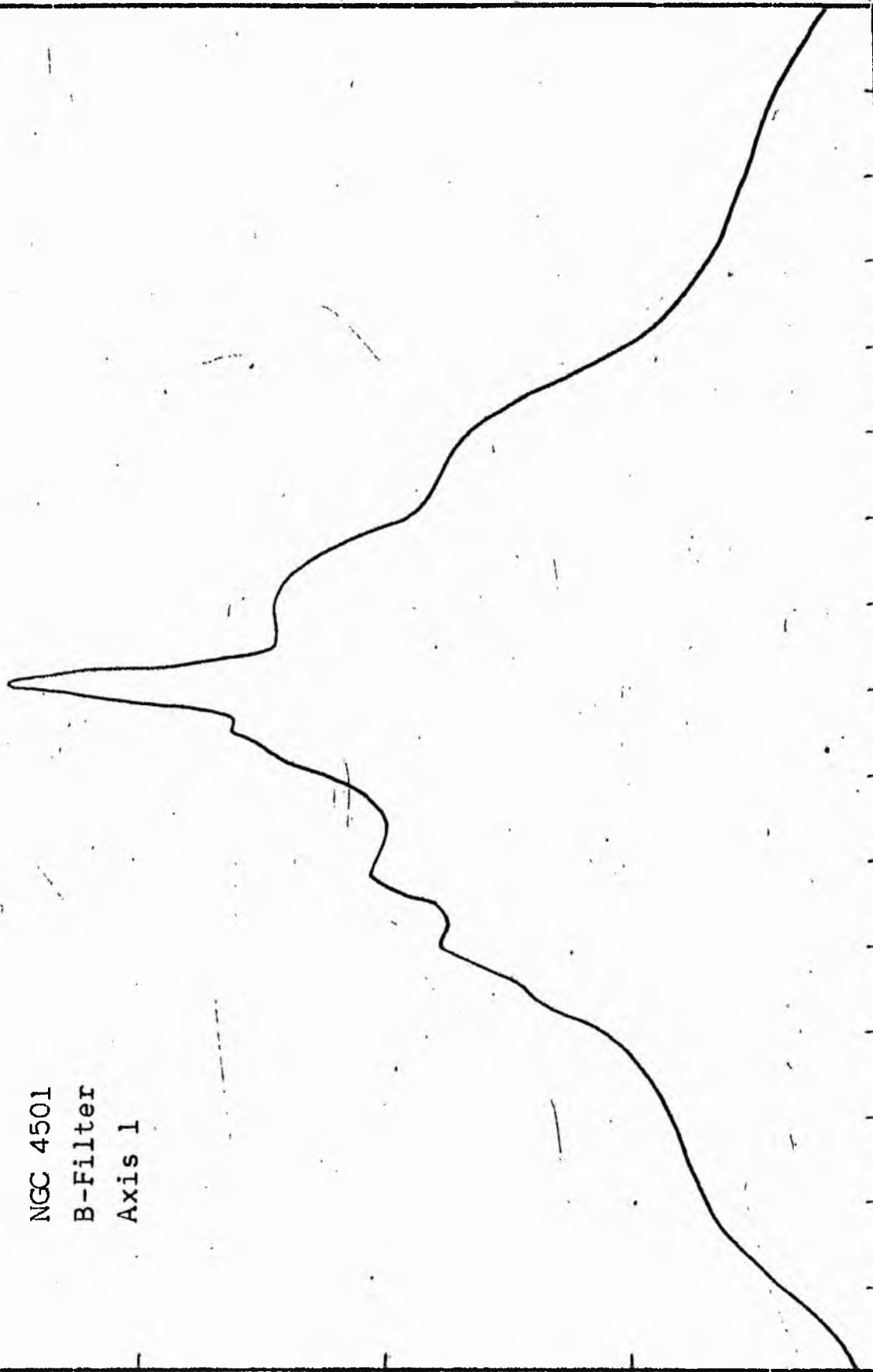
MEAN LUMINOSITY DISTRIBUTION IN NGC 4486
V COLOUR

LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	P	LOG J	μ
2.02	104.713		0.0	0.0			0.0	0.0	0.0	1.438	16.67
		102.356			21.24	2173.7571					
2.00	100.000	89.716	2.60	21.24	42.38	3802.1877	2173.76	0.01	0.07	1.418	16.72
1.90	79.433	71.264	4.50	63.62	73.23	5218.7070	5975.94	0.04	0.12	1.318	16.97
1.80	63.096	56.607	6.60	136.85	202.95	11488.2187	11194.65	0.07	0.18	1.218	17.22
1.70	50.119	44.965	10.40	339.79	248.13	11151.0859	22682.87	0.15	0.28	1.118	17.47
1.60	39.811	35.717	13.68	587.93	216.32	7726.3125	33839.95	0.22	0.37	1.018	17.72
1.50	31.623	28.371	16.00	804.25	161.17	4572.3867	41566.27	0.27	0.44	0.918	17.97
1.40	25.119	22.536	17.53	965.41	155.61	3506.7007	46138.65	0.30	0.48	0.818	18.22
1.30	19.952	17.901	18.89	1121.02	305.63	5470.9102	49645.35	0.33	0.52	0.718	18.47
1.20	15.849	14.219	21.31	1426.65	206.48	2935.8789	55116.26	0.36	0.58	0.618	18.72
1.10	12.589	11.295	22.80	1633.12	314.69	3554.3174	58052.14	0.38	0.62	0.518	18.97
1.00	10.000	8.972	24.90	1947.82	445.32	3995.2180	61606.46	0.41	0.68	0.418	19.22
0.90	7.943	7.126	27.60	2393.14	491.12	3499.9358	65601.62	0.43	0.75	0.318	19.47
0.80	6.310	5.661	30.30	2884.26	532.79	3015.9253	69101.50	0.46	0.83	0.218	19.72
0.70	5.012	4.496	32.98	3417.05	677.11	3044.5596	72117.37	0.48	0.90	0.118	19.97
0.60	3.981	3.572	36.10	4094.15	708.74	2531.3748	75161.87	0.50	0.98	0.018	20.22
0.50	3.162	2.837	39.10	4802.89	1133.58	3216.0413	77693.19	0.51	1.07	-0.082	20.47
0.40	2.512	2.254	43.47	5936.48	1662.00	3745.4224	80939.19	0.53	1.19	-0.182	20.72
0.30	1.995	1.790	49.18	7598.48	1994.89	3570.9905	84654.56	0.56	1.34	-0.282	20.97
0.20	1.585	1.422	55.26	9593.37	2443.98	3475.0928	88225.50	0.58	1.51	-0.382	21.22
0.10	1.259	1.129	61.90	12037.36	2963.15	3346.7415	91700.56	0.61	1.69	-0.482	21.47
-0.00	1.000	0.897	69.10	15000.50	2907.36	2608.3596	95047.25	0.63	1.88	-0.582	21.72
-0.10	0.794	0.713	75.50	17907.86	4470.82	3186.0730	97655.56	0.65	2.06	-0.682	21.97
-0.20	0.631	0.566	84.40	22378.68	6333.45	3585.1660	100841.62	0.67	2.30	-0.782	22.22
-0.30	0.501	0.450	95.60	28712.13	7660.47	3444.4849	104426.75	0.69	2.61	-0.882	22.47
-0.40	0.398	0.357	107.60	36372.60	9928.06	3545.9565	107871.19	0.71	2.94	-0.982	22.72
-0.50	0.316	0.284	121.40	46300.66	9687.59	2749.0000	111417.12	0.74	3.31	-1.082	22.97
-0.60	0.251	0.225	133.50	55990.25	10701.00	2411.5339	114166.12	0.75	3.64	-1.182	23.22
-0.70	0.200	0.179	145.70	66691.25	14035.37	2512.4260	116577.62	0.77	3.97	-1.282	23.47
-0.80	0.158	0.142	160.30	80726.62	12322.31	1752.1099	119090.00	0.79	4.37	-1.382	23.72
-0.90	0.126	0.113	172.10	93048.94	21799.62	2462.1733	120842.06	0.80	4.69	-1.482	23.97
-1.00	0.100	0.090	191.20	114848.56	23827.62	2137.7209	123304.19	0.81	5.22	-1.582	24.22
-1.10	0.079	0.071	210.10	138676.19	26936.50	1919.6050	125441.87	0.83	5.73	-1.682	24.47
-1.20	0.063	0.057	229.60	165612.69	43341.87	2453.4602	127361.44	0.84	6.26	-1.782	24.72
-1.30	0.050	0.045	257.90	208954.56	54524.75	2451.6907	129814.87	0.86	7.04	-1.882	24.97
-1.40	0.040	0.036	289.60	263479.31	138709.56	4954.2617	132266.56	0.87	7.90	-1.982	25.22
-1.50	0.032	0.028	357.80	402188.87	131854.00	3740.8164	137220.81	0.91	9.76	-2.082	25.47
-1.60	0.025	0.023	412.30	534042.87	116059.12	2615.4895	140961.62	0.93	11.25	-2.182	25.72
-1.70	0.020	0.018	454.90	653102.03	39453.25	706.2478	143577.06	0.95	12.41	-2.282	25.97
-1.80	0.016	0.014	468.50	689555.25	75865.00	1078.7390	144283.25	0.95	12.78	-2.382	26.22
-1.90	0.013	0.011	493.60	765420.25	66837.94	754.9158	145361.94	0.96	13.46	-2.482	26.47
-2.00	0.010		514.70	832258.19			146116.81	0.97	14.04	-2.582	26.72
-∞							151316.00	(1)			∞

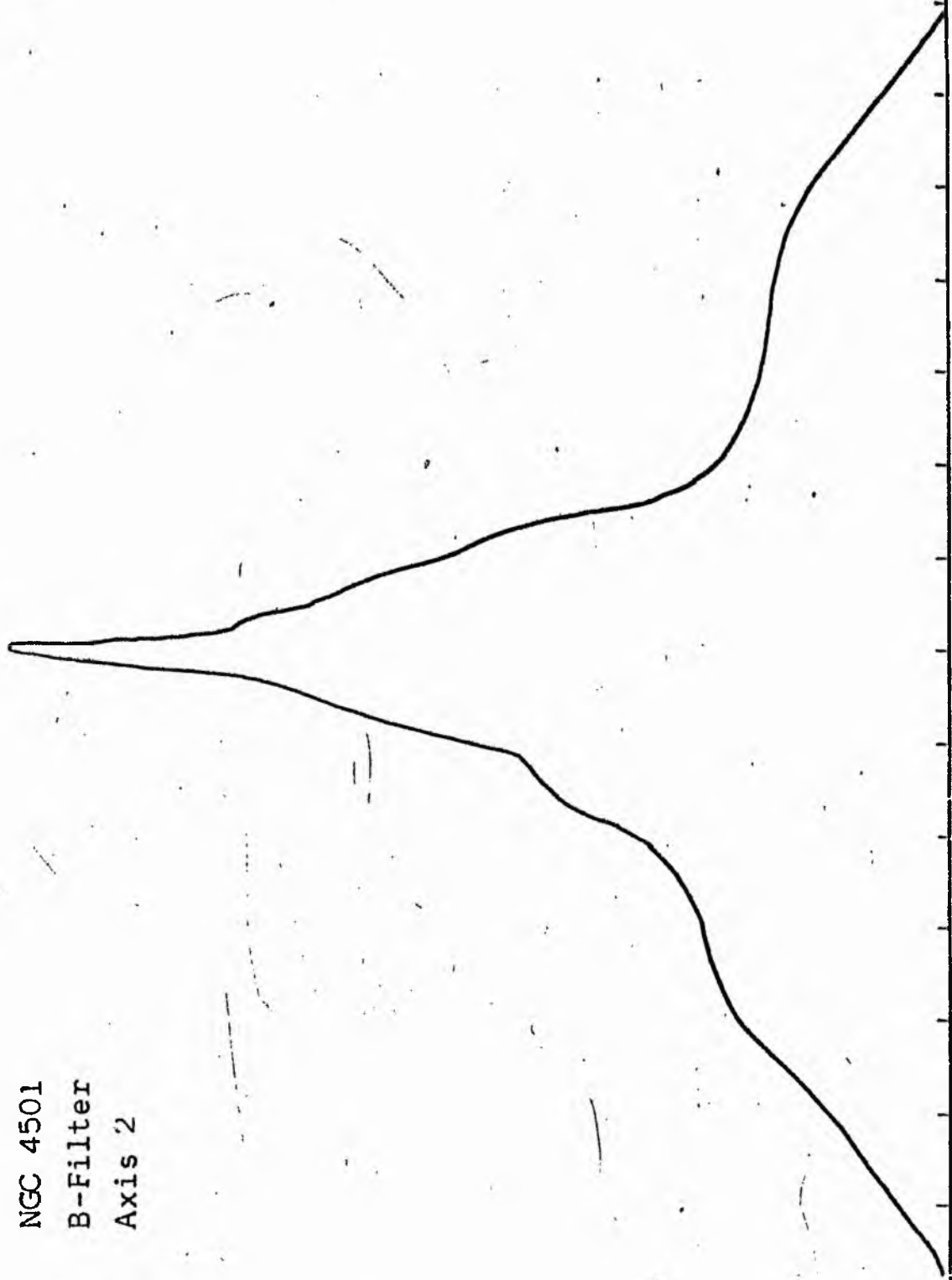
PHOTOMETRIC PARAMETERS OF NGC 4486
V-FILTER

Total luminosity	L_T	= 42.03
Total apparent magnitude	m_T	= 8.77
Apparent central surface brightness	μ_0	= 16.67
Major axis at threshold	$2a_m$	= 17.33
Minor axis at threshold	$2b_m$	= 17.21
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 9.03
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.86
Gradient of exponential component	$G(a)$	= -0.23
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.25
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.21
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.31
Axis ratio	b/a	= 0.95
Equivalent radius	r_1^*	= 0.25
Surface brightness	μ_1	= 17.87
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.70
Axis ratio	b/a	= 0.98
Equivalent radius	r_e^*	= 0.61
Surface brightness	μ_e	= 20.22
Mean surface brightness	μ_e'	= 9.67
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.37
Axis ratio	b/a	= 0.96
Equivalent radius	r_3^*	= 2.17
Surface brightness	μ_3	= 23.22
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.47 \\ = 3.56 \end{cases}$

NGC 4501
B-Filter
Axis 1

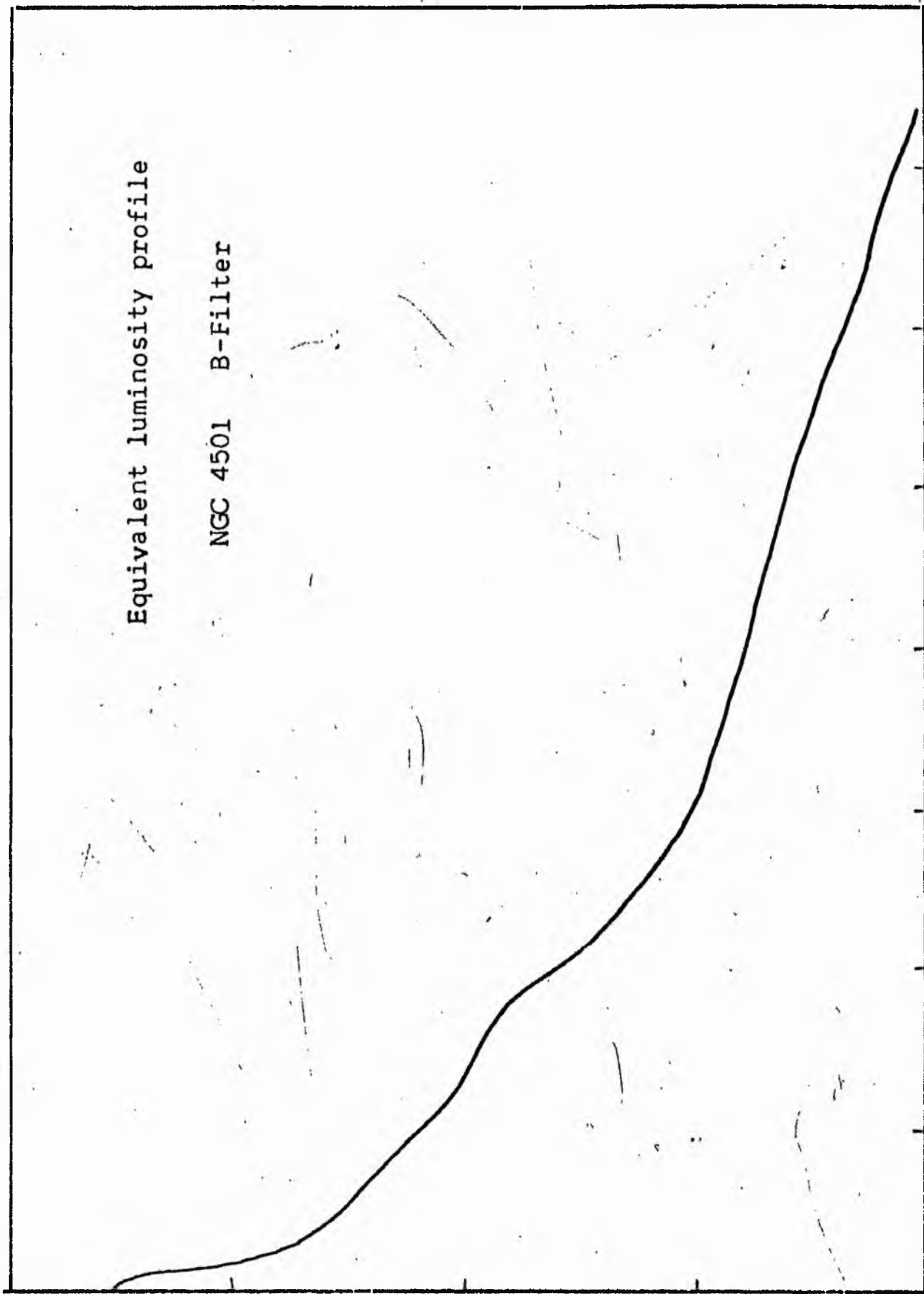


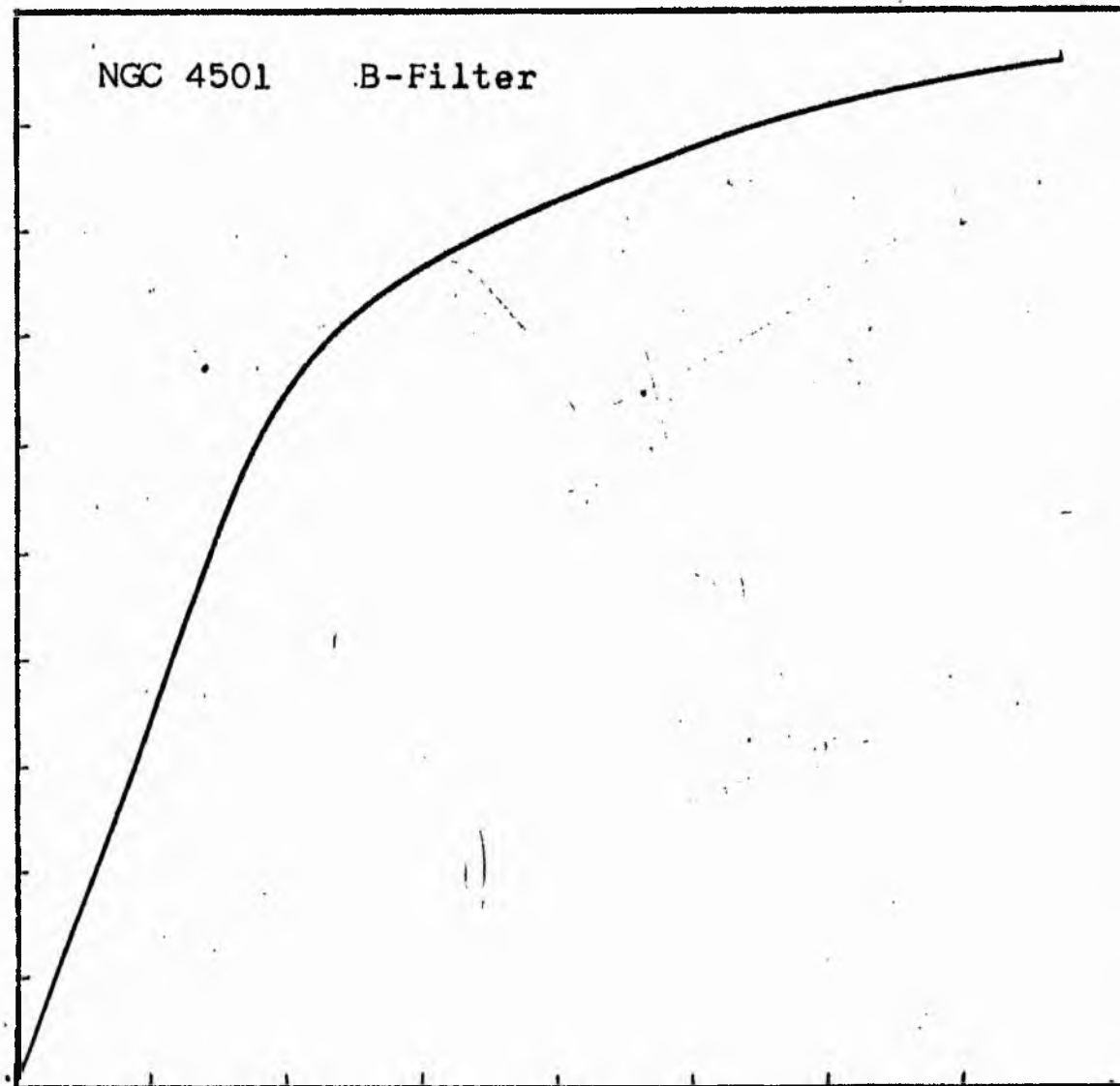
NGC 4501
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4501 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

B COLOUR

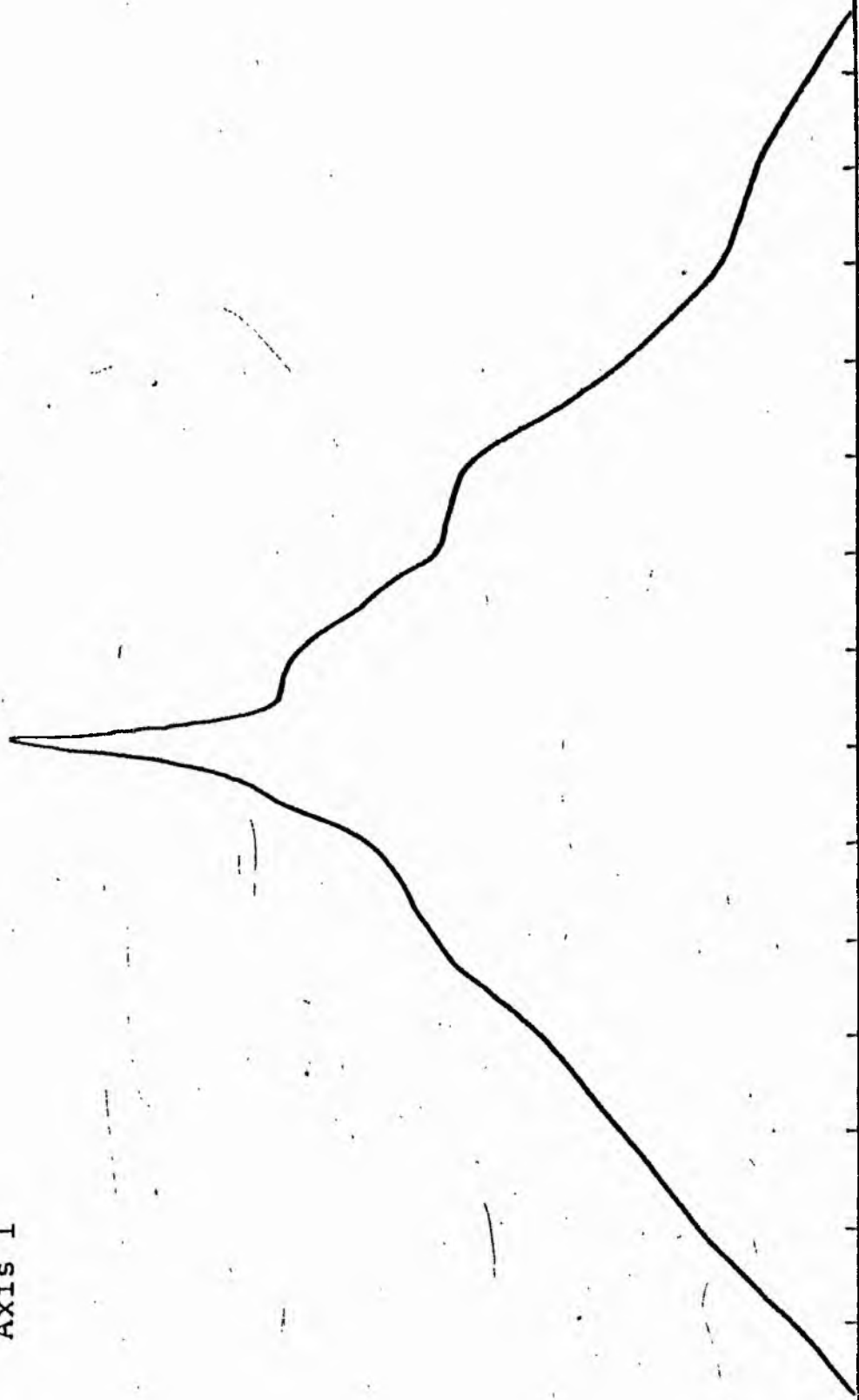
LOG I	I	T	R	AREA	ΔA	P	Σ P	KIR)	P	LOG J	JL
1.52	33.113	32.360	0.0	0.0	24.45	791.5390	0.0	0.0	0.0	1.549	18.37
1.50	31.623	28.371	2.79	24.45	13.37	379.4055	791.54	0.01	0.04	1.529	18.42
1.40	25.114	22.536	3.47	37.83	25.79	501.1080	1170.95	0.02	0.05	1.429	18.67
1.30	19.453	17.901	4.50	63.62	42.07	753.0048	1752.13	0.03	0.07	1.329	18.92
1.20	15.849	14.219	5.80	105.68	61.73	677.7737	2505.14	0.04	0.09	1.229	19.17
1.10	12.509	11.295	7.30	167.42	59.56	672.7576	3382.92	0.05	0.11	1.129	19.42
1.00	10.000	8.972	8.50	226.98	119.38	1071.0344	4055.67	0.06	0.13	1.029	19.67
0.90	7.943	7.126	10.50	346.36	13.32	94.9259	5126.71	0.08	0.16	0.929	19.92
0.80	6.310	5.661	10.70	359.68	319.19	1806.8162	5221.63	0.08	0.16	0.829	20.17
0.70	5.012	4.496	14.70	678.87	396.34	1702.1344	7028.45	0.11	0.22	0.729	20.42
0.60	3.981	3.572	18.50	1075.21	473.09	1089.7251	8810.58	0.13	0.27	0.629	20.67
0.50	3.162	2.837	22.20	1548.30	2799.16	7941.4023	10500.30	0.16	0.33	0.529	20.92
0.40	2.512	2.254	37.20	4347.46	1353.77	3050.8201	18441.71	0.28	0.55	0.429	21.17
0.30	1.995	1.790	42.60	5701.23	2027.59	3629.5256	21442.52	0.33	0.63	0.329	21.42
0.20	1.585	1.422	49.60	7728.82	1263.21	1746.1597	25122.05	0.38	0.73	0.229	21.67
0.10	1.259	1.129	53.50	8992.02	3900.07	4414.9648	26918.26	0.41	0.74	0.129	21.92
-0.00	1.000	0.897	64.06	12892.09	7511.76	6715.2578	31323.17	0.48	0.95	0.029	22.17
-0.10	0.794	0.713	80.59	20403.85	3648.97	2600.4065	36062.43	0.58	1.14	-0.071	22.42
-0.20	0.631	0.566	87.50	24052.82	3824.53	2164.9543	40662.83	0.62	1.30	-0.171	22.67
-0.30	0.501	0.450	94.20	27677.35	2479.50	1114.8994	42827.79	0.65	1.34	-0.271	22.92
-0.40	0.398	0.357	98.30	30356.86	3426.85	1223.9570	43942.68	0.67	1.46	-0.371	23.17
-0.50	0.316	0.284	103.70	33783.71	3336.37	946.5537	45166.64	0.69	1.54	-0.471	23.42
-0.60	0.251	0.225	108.70	37120.08	4572.12	1030.3604	46113.14	0.70	1.61	-0.571	23.67
-0.70	0.200	0.179	115.20	41642.20	6426.11	1150.3218	47143.55	0.72	1.71	-0.671	23.92
-0.80	0.158	0.142	123.76	46118.32	8515.67	1211.4175	48293.87	0.74	1.83	-0.771	24.17
-0.90	0.126	0.113	134.27	56637.98	9596.33	1083.8689	49505.29	0.76	1.99	-0.871	24.42
-1.00	0.100	0.090	145.20	66234.31	24643.12	2210.8948	50589.15	0.77	2.15	-0.971	24.67
-1.10	0.079	0.071	170.08	90877.44	27964.75	1993.2485	52800.05	0.81	2.52	-1.071	24.92
-1.20	0.063	0.057	194.50	118847.19	37002.94	2094.6399	54793.29	0.84	2.88	-1.171	25.17
-1.30	0.050	0.045	222.73	155850.12	35970.50	1617.4162	56887.93	0.87	3.30	-1.271	25.42
-1.40	0.040	0.036	247.10	191820.62	39582.44	1413.7639	58505.34	0.89	3.66	-1.371	25.67
-1.50	0.032	0.028	271.40	231403.06	43106.94	1222.9883	59919.10	0.91	4.02	-1.471	25.92
-1.60	0.025	0.023	295.60	274510.00	46588.00	914.6890	61142.09	0.93	4.38	-1.571	26.17
-1.70	0.020	0.018	316.70	315048.00	24951.12	446.6489	62056.78	0.95	4.69	-1.671	26.42
-1.80	0.016	0.014	329.00	340049.12	56092.75	797.5969	62503.43	0.95	4.87	-1.771	26.67
-1.90	0.013	0.011	355.10	396141.87	46587.25	526.1921	63301.02	0.97	5.26	-1.871	26.92
-2.00	0.010		375.40	442729.12			63827.21	0.97	5.56	-1.971	27.17
-∞							65527.00	(11			∞

PHOTOMETRIC PARAMETERS OF NGC 4501

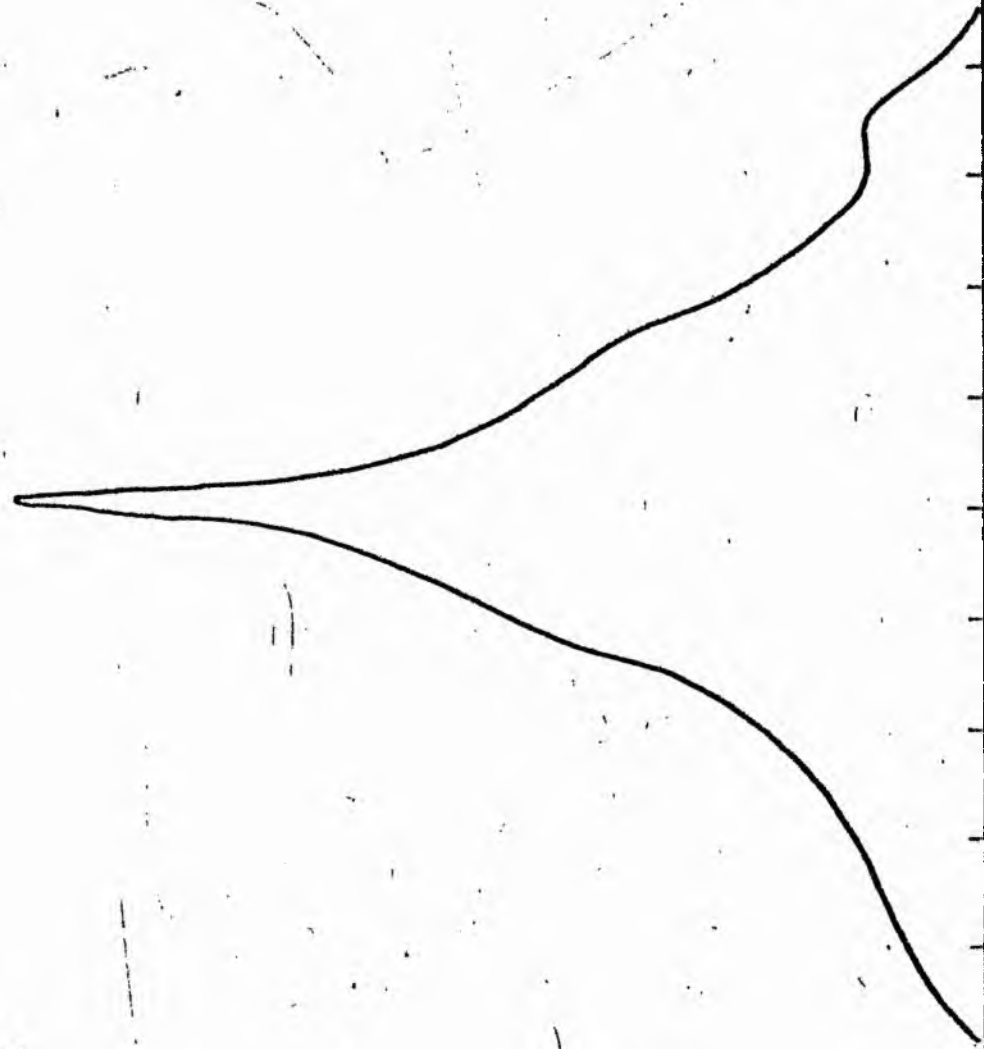
B-FILTER

Total luminosity	L_T	= 18.20
Total apparent magnitude	m_T	= 10.13
Apparent central surface brightness	μ_0	= 18.37
Major axis at threshold	$2a_m$	= 13.30
Minor axis at threshold	$2b_m$	= 11.40
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 7.55
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.40
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.35
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.89
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.68
Axis ratio	b/a	= 0.56
Equivalent radius	r_1^*	= 0.56
Surface brightness	μ_1	= 21.12
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.56
Axis ratio	b/a	= 0.46
Equivalent radius	r_e^*	= 1.13
Surface brightness	μ_e	= 22.22
Mean surface brightness	μ_e^r	= 12.39
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.91
Axis ratio	b/a	= 0.53
Equivalent radius	r_3^*	= 2.18
Surface brightness	μ_3	= 24.49
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.02 \\ = 1.94 \end{matrix}$

NGC 4501
V-Filter
Axis 1

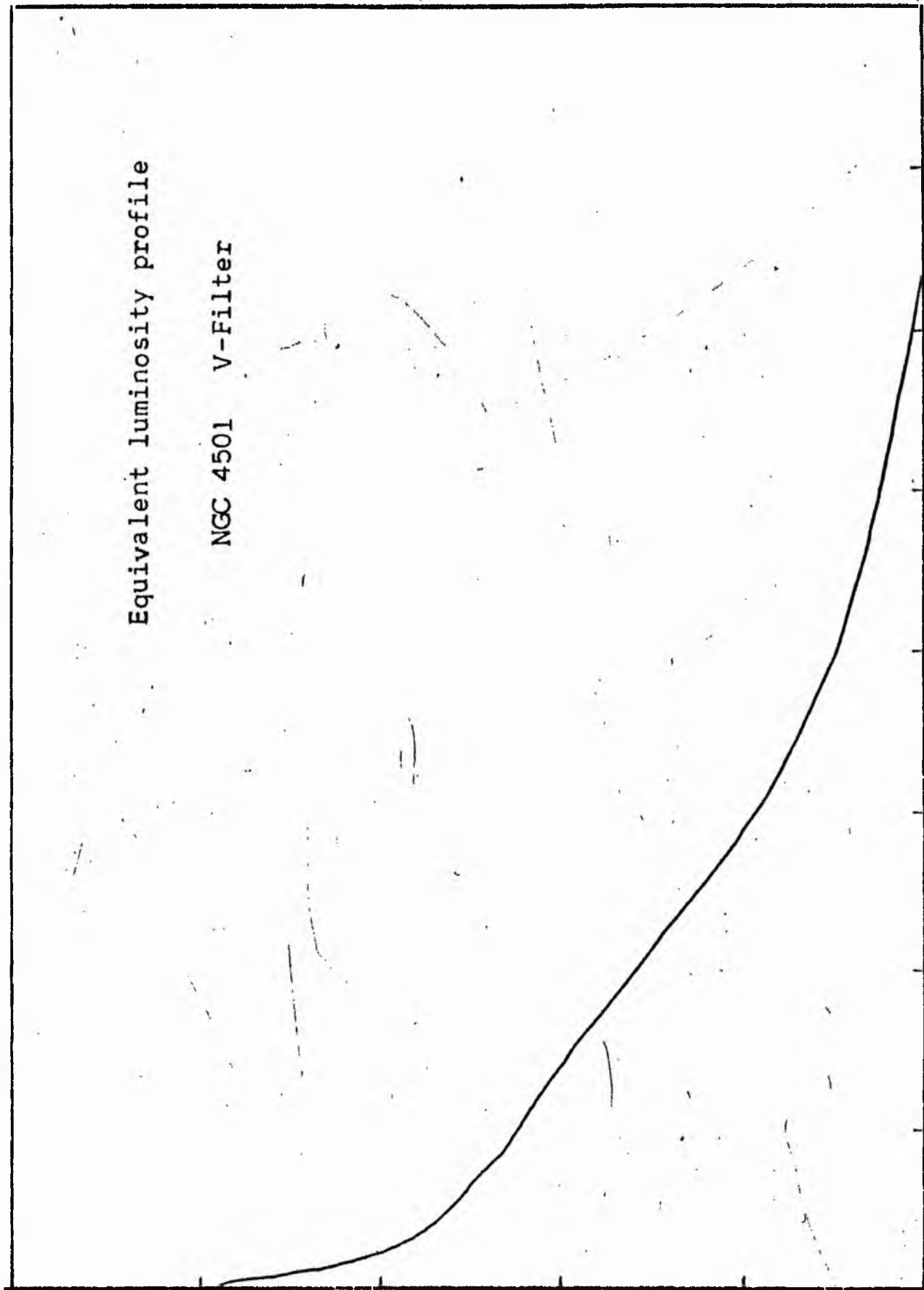


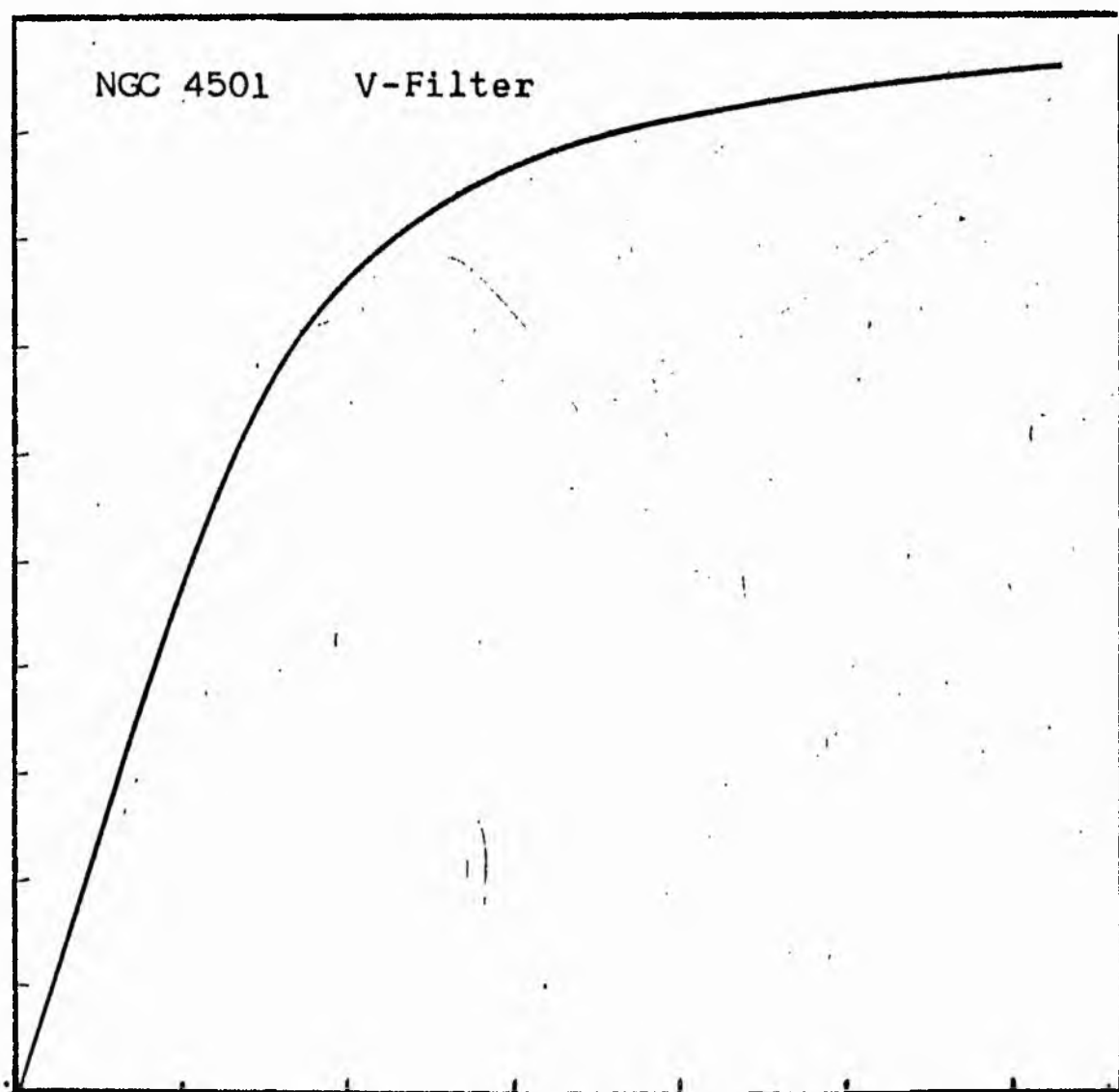
NGC 4501
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4501 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4501
V COLOUR

LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.91	81.283		0.0	0.0			0.0	0.0	0.0	1.739	16.80
1.90	79.433	80.358	0.60	1.13	1.13	90.8825	90.88	0.00	0.01	1.729	16.83
1.80	61.096	71.264	1.40	6.16	5.03	358.2124	449.09	0.01	0.03	1.629	17.08
1.70	50.119	56.607	2.10	13.85	7.70	435.6990	884.79	0.01	0.04	1.529	17.33
1.60	39.811	44.965	2.90	26.42	12.57	565.0410	1449.83	0.02	0.05	1.429	17.58
1.50	31.623	35.717	3.60	40.72	14.29	510.5420	1960.38	0.03	0.07	1.329	17.83
1.40	25.119	28.371	4.20	55.42	14.70	417.1230	2377.50	0.04	0.08	1.229	18.08
1.30	19.952	22.536	5.10	81.71	26.30	592.5759	2970.08	0.05	0.09	1.129	18.33
1.20	15.849	17.901	7.20	162.86	81.15	1452.5913	4422.66	0.07	0.13	1.029	18.58
1.10	12.589	14.219	8.20	211.24	48.38	687.9214	5110.58	0.08	0.15	0.929	18.83
1.00	10.000	11.295	10.30	333.29	122.05	1378.5061	6489.09	0.11	0.19	0.829	19.08
0.90	7.943	8.972	11.20	394.08	60.79	545.3794	7034.46	0.12	0.21	0.729	19.33
0.80	6.310	7.126	13.13	541.60	147.52	1051.2766	8085.74	0.13	0.24	0.629	19.58
0.70	5.012	5.661	17.28	938.07	396.47	2244.3059	10330.04	0.17	0.32	0.529	19.83
0.60	3.981	4.496	22.10	1534.38	596.31	2681.2683	13011.31	0.22	0.41	0.429	20.08
0.50	3.162	3.572	30.40	2903.33	1368.95	4889.3984	17900.71	0.30	0.56	0.329	20.33
0.40	2.512	2.837	34.20	3674.53	771.20	2187.9385	20088.64	0.33	0.63	0.229	20.58
0.30	1.995	2.254	37.20	4347.46	672.93	1516.4768	21605.12	0.36	0.69	0.129	20.83
0.20	1.585	1.790	51.80	8429.64	4082.18	7307.3711	28912.49	0.48	0.96	0.029	21.08
0.10	1.259	1.422	58.60	10788.10	2358.46	3353.4873	32265.98	0.54	1.08	-0.071	21.33
-0.00	1.000	1.129	64.70	13150.98	2362.88	2668.7644	34934.74	0.58	1.20	-0.171	21.58
-0.10	0.794	0.897	71.54	16078.57	2927.60	2626.5166	37561.25	0.62	1.32	-0.271	21.83
-0.20	0.631	0.713	82.30	21278.90	5200.33	3705.9507	41267.20	0.69	1.52	-0.371	22.08
-0.30	0.501	0.566	88.40	24550.16	3271.26	1851.7549	43118.96	0.72	1.63	-0.471	22.33
-0.40	0.398	0.450	95.70	28772.24	4222.08	1898.4343	45017.39	0.75	1.77	-0.571	22.58
-0.50	0.316	0.357	100.50	31730.87	2958.63	1056.7190	46074.11	0.77	1.86	-0.671	22.83
-0.60	0.251	0.284	107.14	36062.27	4331.40	1228.8455	47302.95	0.79	1.98	-0.771	23.08
-0.70	0.200	0.225	114.54	41215.83	5153.56	1161.3862	48464.34	0.81	2.12	-0.871	23.33
-0.80	0.158	0.179	122.71	47305.28	6089.45	1090.0520	49554.39	0.82	2.27	-0.971	23.58
-0.90	0.126	0.142	131.67	54465.76	7160.48	1018.1484	50572.54	0.84	2.43	-1.071	23.83
-1.00	0.100	0.113	141.34	62759.58	8293.82	936.7510	51509.29	0.86	2.61	-1.171	24.08
-1.10	0.079	0.090	147.50	68349.25	5589.67	501.4834	52010.77	0.86	2.72	-1.271	24.33
-1.20	0.063	0.071	154.30	74776.56	6447.31	459.4617	52470.23	0.87	2.85	-1.371	24.58
-1.30	0.050	0.057	166.20	86778.44	11981.87	678.2598	53148.48	0.88	3.07	-1.471	24.83
-1.40	0.040	0.045	179.50	101222.87	14444.44	649.4902	53797.97	0.89	3.32	-1.571	25.08
-1.50	0.032	0.036	193.40	117506.69	16283.81	581.6055	54379.58	0.90	3.57	-1.671	25.33
-1.60	0.025	0.028	215.30	145625.62	28118.94	797.7593	55177.34	0.92	3.98	-1.771	25.58
-1.70	0.020	0.023	243.70	186578.19	40952.56	922.9001	56100.23	0.93	4.50	-1.871	25.83
-1.80	0.016	0.018	258.10	209278.75	22700.56	406.3599	56506.59	0.94	4.77	-1.971	26.08
-1.90	0.013	0.014	279.60	245547.31	36318.56	516.4207	57023.01	0.95	5.17	-2.071	26.33
-2.00	0.010	0.011	304.10	290524.06	44926.75	507.4351	57530.45	0.96	5.62	-2.171	26.58
-∞							60130.00	(11)			∞

PHOTOMETRIC PARAMETERS OF NGC 4501

V-FILTER

Total luminosity	L_T	= 16.70
Total apparent magnitude	m_T	= 9.63
Apparent central surface brightness	μ_0	= 16.80
Major axis at threshold	$2a_m$	= 12.38
Minor axis at threshold	$2b_m$	= 8.26
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.37
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.88
Gradient of exponential component	$G(a)$	= -0.48
Equivalent gradient of exponential comp.	$G(r^*)$	= -0.64
Equivalent gradient of reduced exp. comp.	$G(\rho)$	= -0.76
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.53
Axis ratio	b/a	= 0.31
Equivalent radius	r_1^*	= 0.43
Surface brightness	μ_1	= 20.18
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.21
Axis ratio	b/a	= 0.26
Equivalent radius	r_e^*	= 0.90
Surface brightness	μ_e	= 21.16
Mean surface brightness	μ_e'	= 11.38
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.47
Axis ratio	b/a	= 0.46
Equivalent radius	r_3^*	= 1.60
Surface brightness	μ_3	= 22.58
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.11 \\ = 1.77 \end{cases}$

NGC 4519
B-Filter
Axis 1

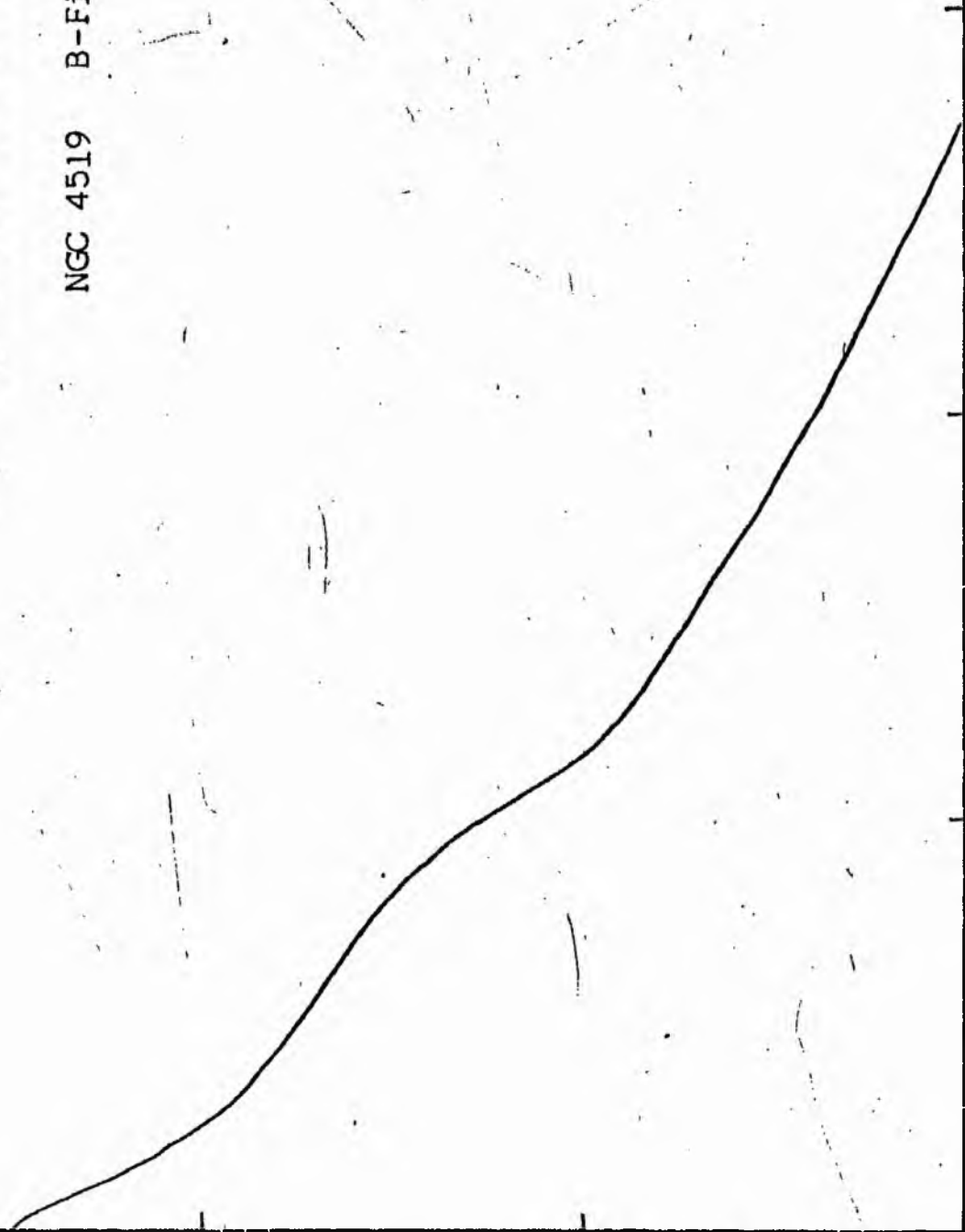


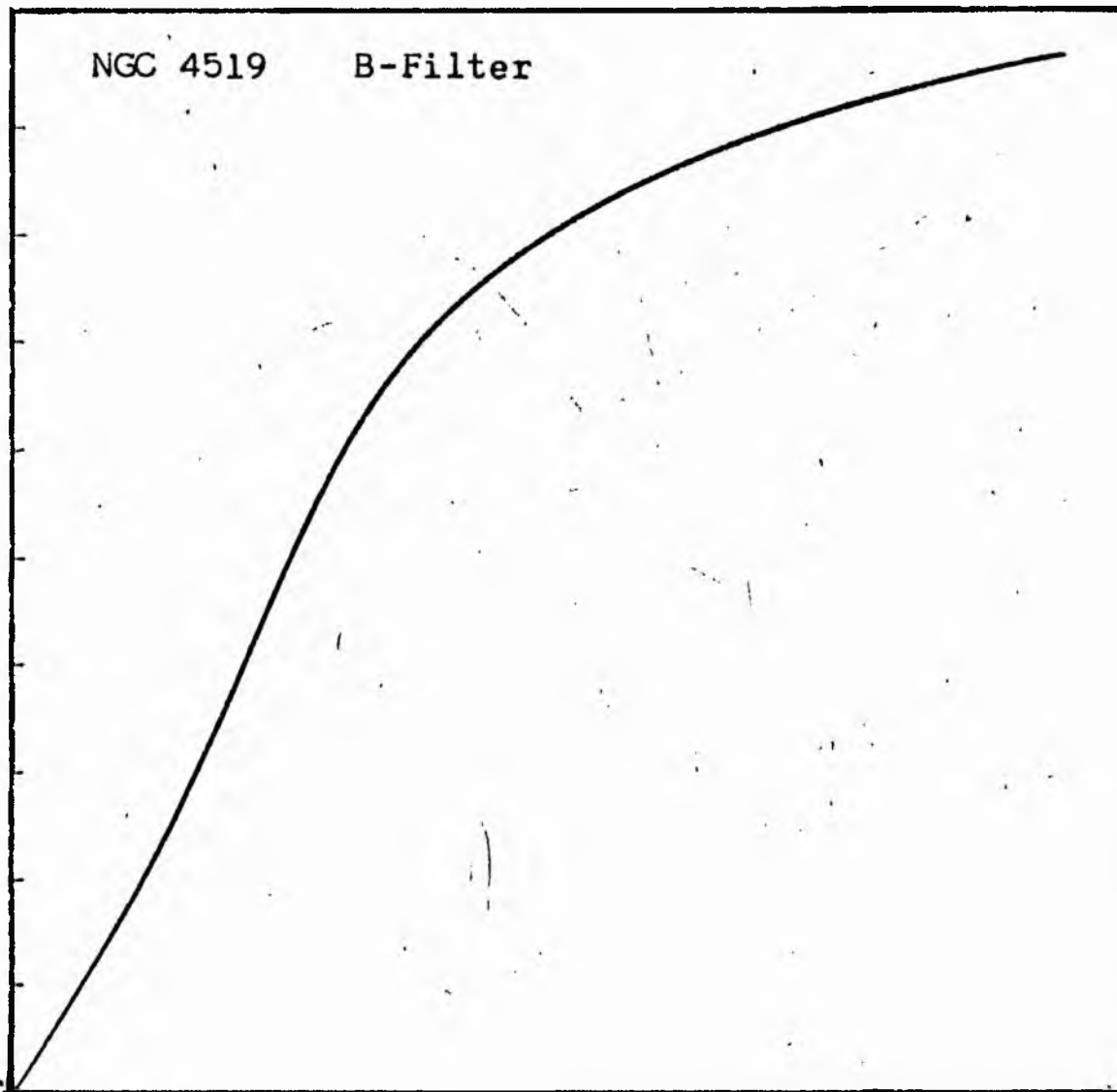
NGC 4519
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4519 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4519
B COLOR

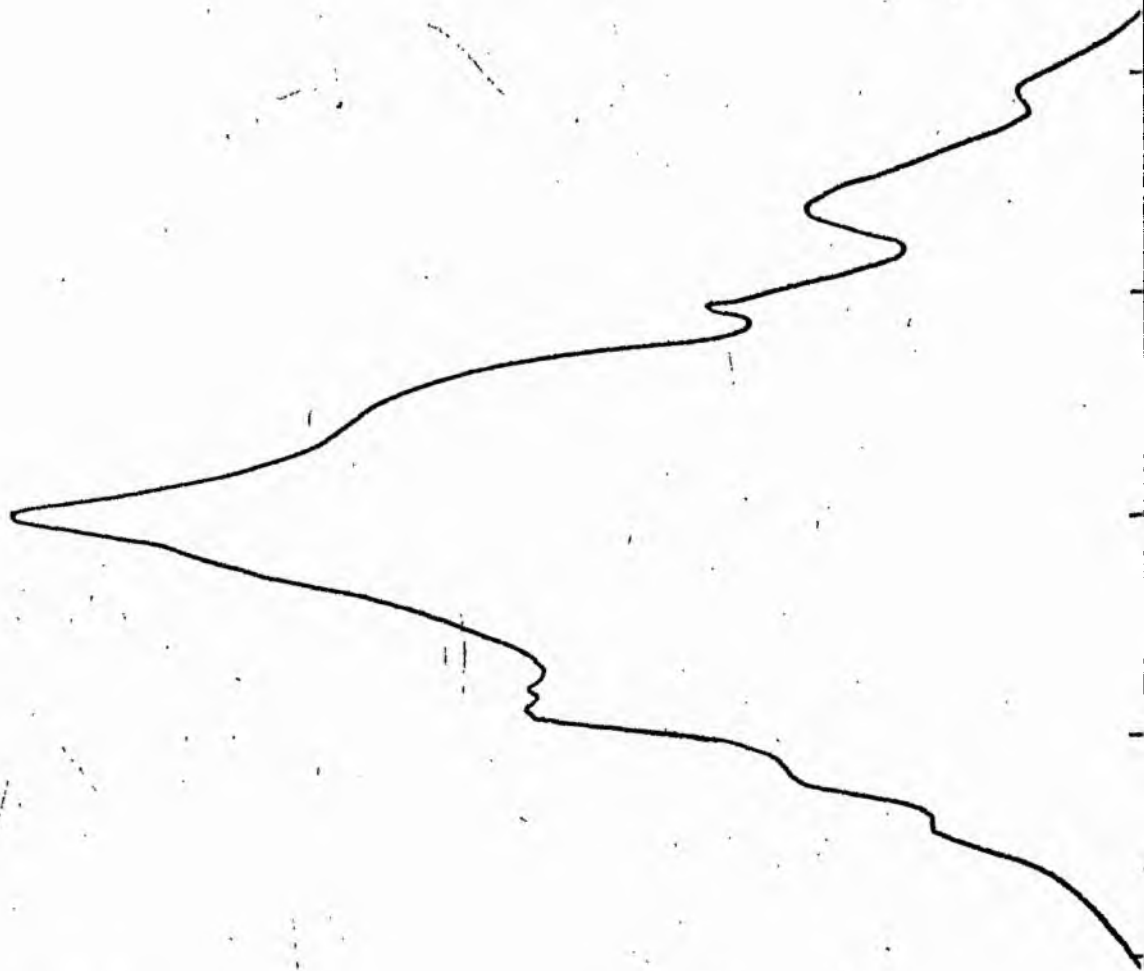
LOG I	I	I	R	AREA	A A	P	EP	K(R)	Q	LOG J	M
0.46	2.084		0.0	0.0			0.0	0.0	0.0	0.919	20.53
0.40	2.512	2.698	3.79	45.13	45.13	121.7484	121.75	0.02	0.10	0.859	20.68
0.30	1.995	2.254	5.54	96.42	51.29	115.5456	237.34	0.04	0.14	0.759	20.73
0.20	1.585	1.790	7.20	162.86	66.44	110.9322	356.28	0.06	0.19	0.659	21.18
0.10	1.254	1.422	8.99	253.90	91.04	129.4558	485.73	0.08	0.24	0.559	21.43
0.00	1.000	1.129		400.44	146.54	165.5073	651.24	0.11	0.30	0.459	21.68
-0.10	0.794	0.897	15.33	738.30	337.86	303.1174	954.36	0.16	0.40	0.359	21.73
-0.20	0.631	0.713	20.50	1320.25	581.95	414.7239	1369.08	0.23	0.54	0.259	22.18
-0.30	0.501	0.566	28.24	2505.41	1185.16	670.8840	2039.96	0.35	0.74	0.159	22.43
-0.40	0.398	0.450	34.67	3776.22	1270.81	571.4167	2611.38	0.45	0.91	0.059	22.68
-0.50	0.316	0.357	40.57	5170.82	1394.60	446.1055	3109.49	0.53	1.06	-0.041	22.93
-0.60	0.251	0.284	45.90	6618.73	1447.91	410.7854	3520.27	0.60	1.20	-0.141	23.18
-0.70	0.200	0.225	49.78	7785.02	1166.28	262.8303	3783.10	0.65	1.30	-0.241	23.43
-0.80	0.158	0.179	52.34	8606.31	821.30	147.0187	3930.12	0.67	1.37	-0.341	23.68
-0.90	0.126	0.142	55.31	9610.75	1004.43	142.8215	4072.94	0.70	1.45	-0.441	23.93
-1.00	0.100	0.113	58.30	10677.92	1067.18	120.5337	4193.47	0.72	1.53	-0.541	24.18
-1.10	0.079	0.090	63.58	12699.62	2021.70	181.3799	4374.85	0.75	1.66	-0.641	24.43
-1.20	0.063	0.071	69.51	15179.04	2479.41	176.6941	4551.54	0.78	1.82	-0.741	24.68
-1.30	0.050	0.057	76.09	18188.84	3009.80	170.3772	4721.92	0.81	1.99	-0.841	24.93
-1.40	0.040	0.045	83.33	21814.85	3626.01	163.0433	4884.96	0.84	2.18	-0.941	25.18
-1.50	0.032	0.036	91.75	26446.12	4631.27	165.4150	5050.37	0.87	2.40	-1.041	25.43
-1.60	0.025	0.028	99.70	31277.84	4831.72	137.0808	5187.45	0.89	2.61	-1.141	25.68
-1.70	0.020	0.023	107.97	36623.17	5345.32	120.4619	5307.91	0.91	2.82	-1.241	25.93
-1.80	0.016	0.018	116.34	42521.43	5898.27	105.5846	5413.50	0.93	3.04	-1.341	26.18
-1.90	0.013	0.014	124.88	48993.18	6471.74	92.0233	5505.52	0.94	3.27	-1.441	26.43
-2.00	0.010	0.011	135.70	57850.82	8857.65	100.0451	5605.56	0.96	3.55	-1.541	26.68
-∞							5835.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4519

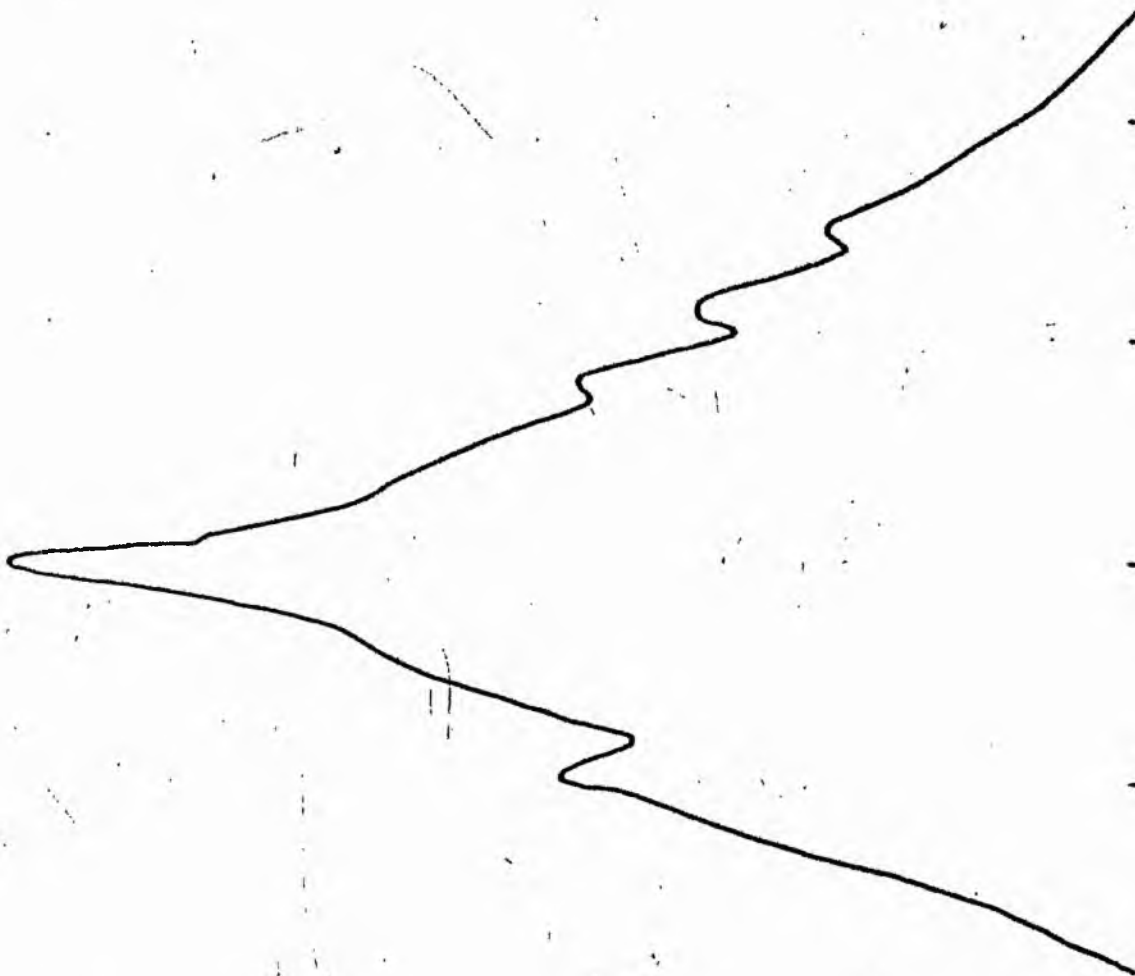
B-FILTER

Total luminosity	L_T	= 1.62
Total apparent magnitude	m_T	= 12.26
Apparent central surface brightness	μ_0	= 20.53
Major axis at threshold	$2a_m$	= 4.35
Minor axis at threshold	$2b_m$	= 4.26
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 2.47
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.82
Gradient of exponential component	$G(a)$	= -1.21
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.96
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.81
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.35
Axis ratio	b/a	= 0.41
Equivalent radius	r_1^*	= 0.36
Surface brightness	μ_1	= 22.22
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.65
Axis ratio	b/a	= 0.82
Equivalent radius	r_e^*	= 0.64
Surface brightness	μ_e	= 22.83
Mean surface brightness	μ_e'	= 13.26
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.95
Axis ratio	b/a	= 1.12
Equivalent radius	r_3^*	= 1.06
Surface brightness	μ_3	= 24.43
Concentration indices	$\left\{ \begin{array}{l} C_{21} \\ C_{32} \end{array} \right.$	= 1.77
		= 1.66

NGC 4519
V-Filter
Axis 1

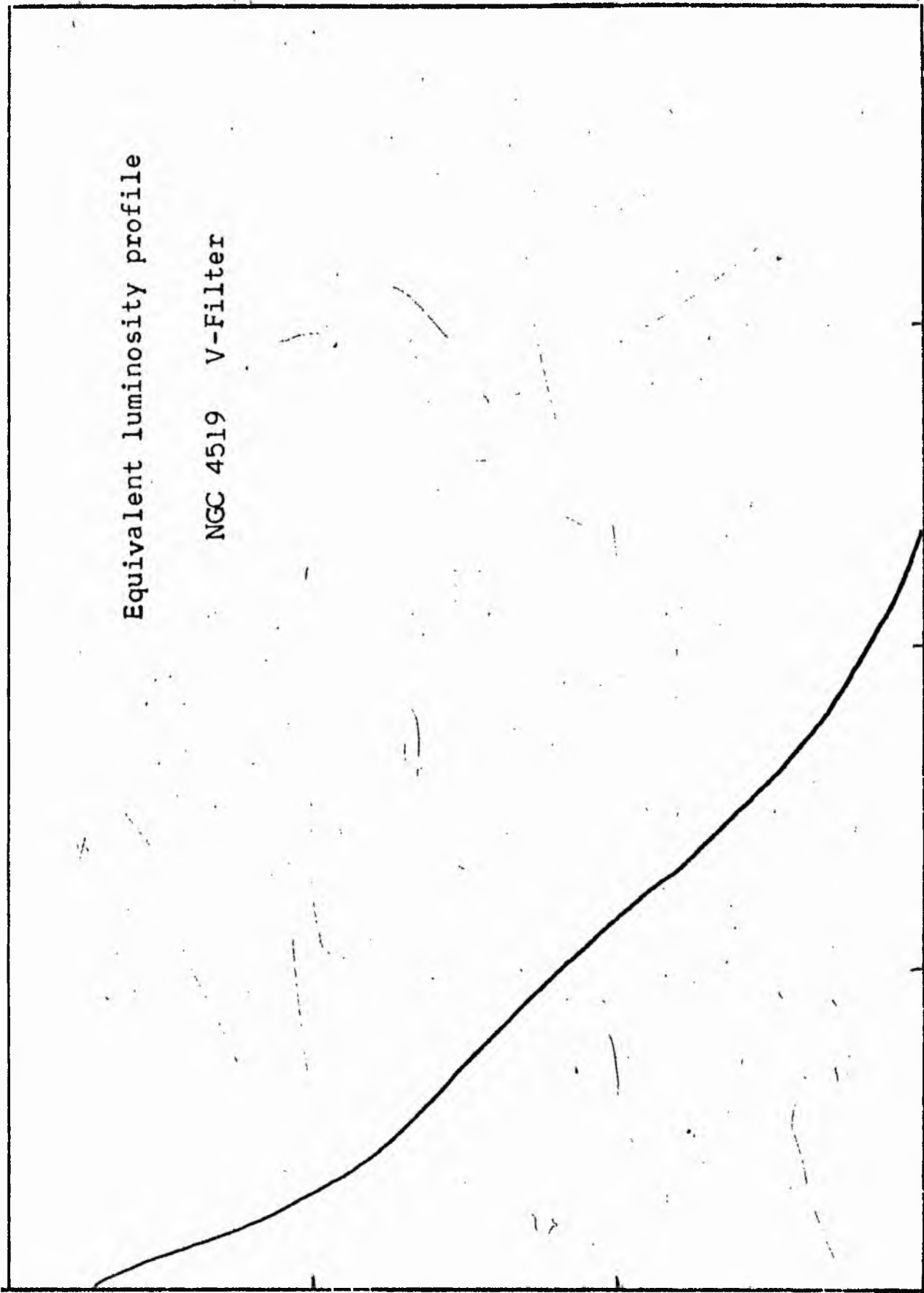


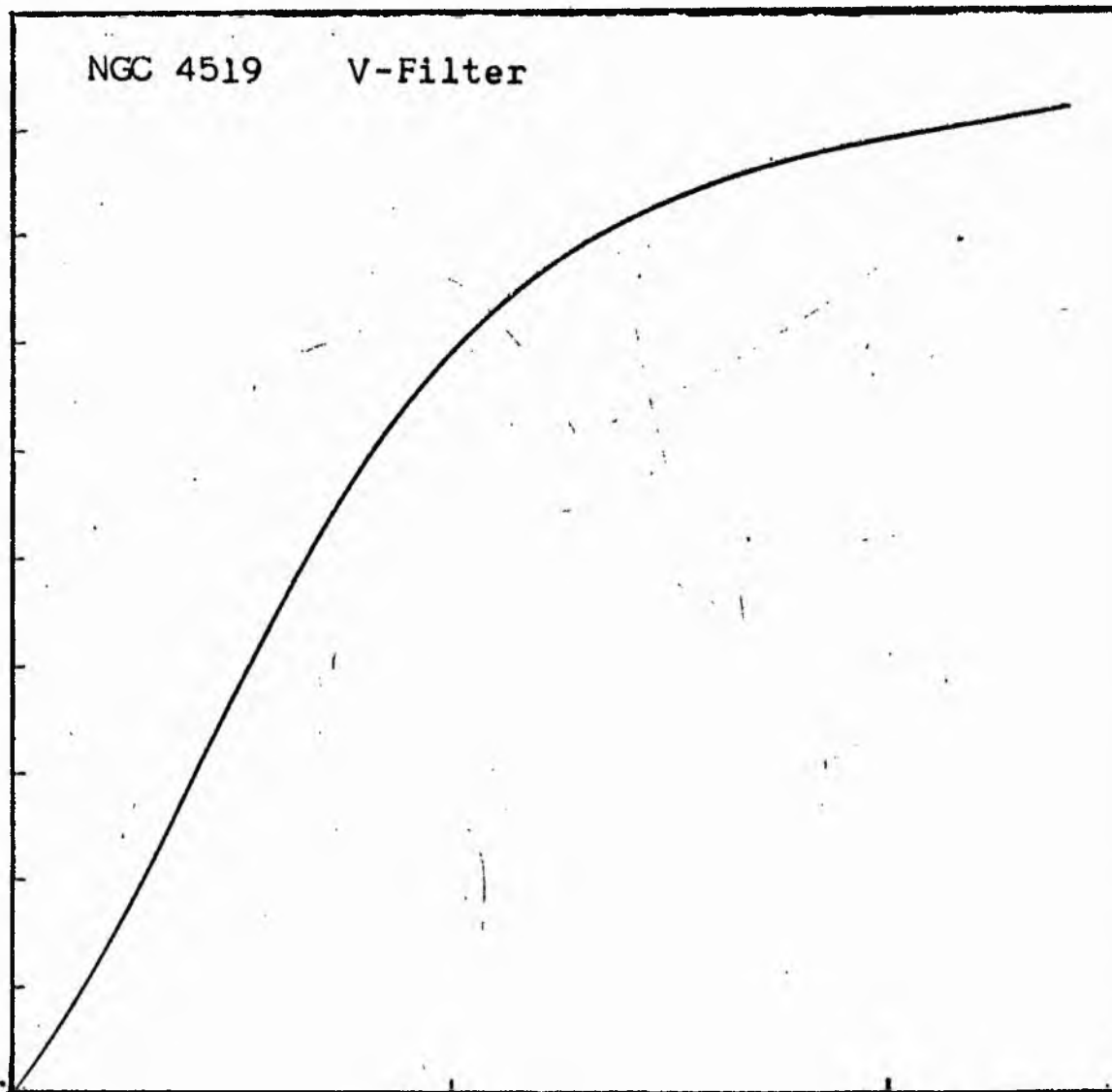
NGC 4519
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4519 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4519
V COLOUR

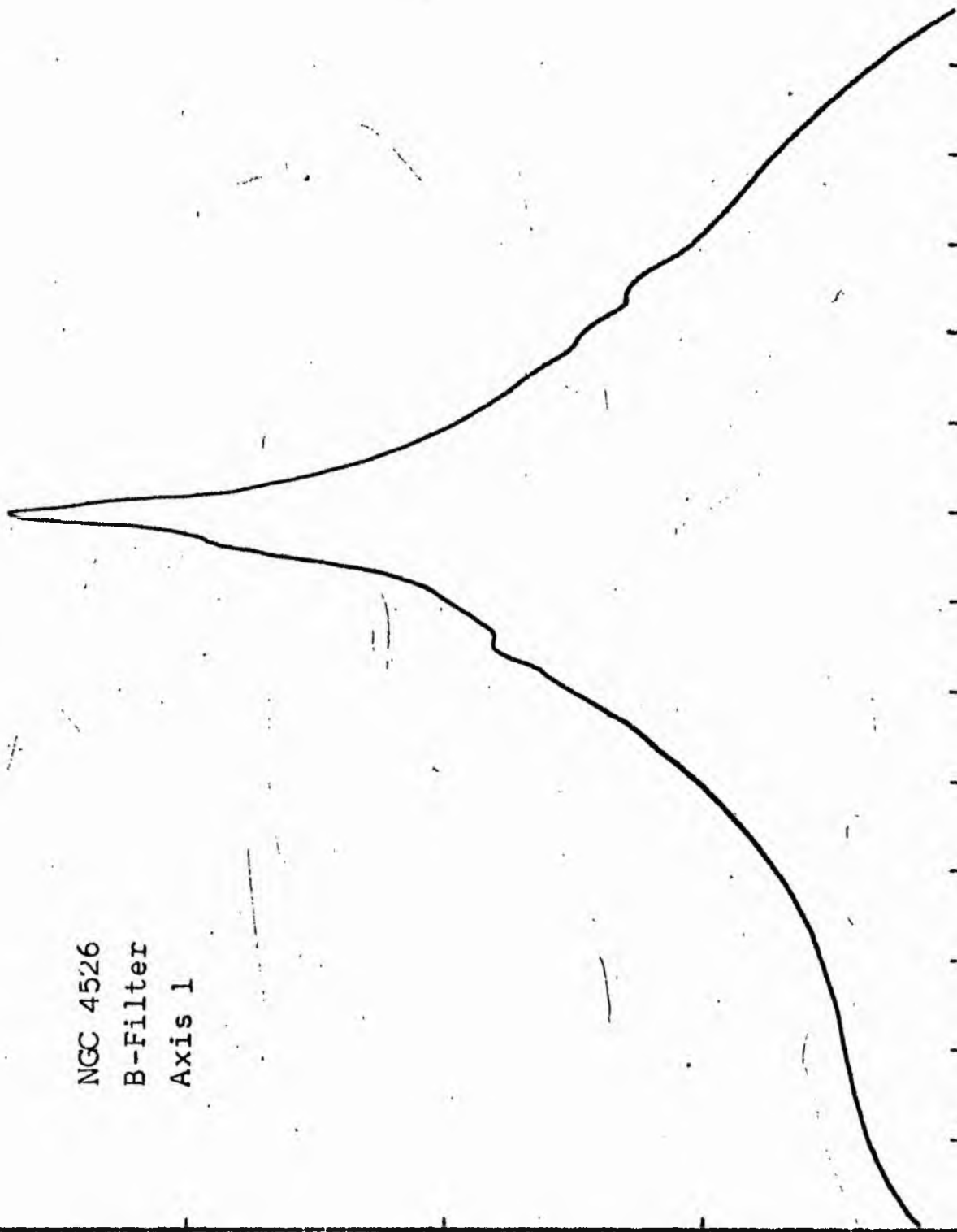
LOG I	I	T	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
0.69	4.898		0.0	0.0			0.0	0.0	0.0	1.123	19.35
		4.439			32.17	142.8159					
0.60	3.981	3.572	3.20	32.17	21.28	97.4240	142.82	0.03	0.10	1.033	19.50
0.50	3.162	2.837	4.35	59.45	59.76	169.5520	243.24	0.04	0.13	0.933	19.83
0.40	2.512	2.254	6.16	119.21	62.73	141.3591	409.79	0.07	0.18	0.833	20.08
0.30	1.995	1.790	7.61	181.94	77.08	137.9733	551.15	0.10	0.23	0.733	20.33
0.20	1.585	1.422	9.08	259.01	97.31	138.3720	689.12	0.12	0.27	0.633	20.58
0.10	1.259	1.129	10.65	356.33	278.93	315.0398	827.50	0.15	0.32	0.533	20.83
0.00	1.000	0.897	14.22	635.26	388.20	348.3513	1142.54	0.20	0.43	0.433	21.08
-0.10	0.794	0.713	18.05	1023.54	524.76	373.9700	1499.89	0.27	0.54	0.333	21.33
-0.20	0.631	0.566	22.20	1548.30	777.68	440.2241	1864.86	0.33	0.67	0.233	21.58
-0.30	0.501	0.450	27.21	2325.90	850.92	382.6128	2305.08	0.41	0.82	0.133	21.83
-0.40	0.398	0.357	31.80	3176.90	987.86	352.8333	2687.69	0.48	0.95	0.033	22.08
-0.50	0.316	0.284	36.41	4164.77	1118.82	317.4199	3040.53	0.54	1.09	-0.067	22.33
-0.60	0.251	0.225	41.01	5283.59	1254.64	282.7429	3357.95	0.60	1.23	-0.167	22.58
-0.70	0.200	0.179	45.62	6538.23	1391.32	249.0583	3640.69	0.65	1.37	-0.267	22.83
-0.80	0.158	0.142	50.24	7929.55	1521.99	216.4138	3889.75	0.69	1.51	-0.367	23.08
-0.90	0.126	0.113	54.85	9451.55	1248.36	140.9982	4106.16	0.73	1.65	-0.467	23.33
-1.00	0.100	0.090	58.36	10699.91	1407.55	126.2806	4247.16	0.76	1.75	-0.567	23.58
-1.10	0.079	0.071	62.08	12107.46	1585.60	112.9970	4373.43	0.78	1.86	-0.667	23.83
-1.20	0.063	0.057	66.02	13693.06	1788.83	101.2609	4486.43	0.80	1.98	-0.767	24.08
-1.30	0.050	0.045	70.20	15481.89	2001.56	90.0001	4587.69	0.82	2.11	-0.867	24.33
-1.40	0.040	0.036	74.60	17483.45	2242.52	80.0959	4677.69	0.83	2.24	-0.967	24.58
-1.50	0.032	0.028	79.24	19725.97	2493.90	70.7545	4757.78	0.85	2.38	-1.067	24.83
-1.60	0.025	0.023	84.10	22219.87	5303.49	119.5191	4828.54	0.86	2.52	-1.167	25.08
-1.70	0.020	0.018	93.60	27523.36	3616.71	64.7426	4948.05	0.88	2.81	-1.267	25.33
-1.80	0.016	0.014	99.56	31140.07	5980.01	85.0312	5012.79	0.89	2.99	-1.367	25.58
-1.90	0.013	0.011	108.70	37120.08	3279.42	37.0403	5097.82	0.91	3.26	-1.467	25.83
-2.00	0.010		113.40	40399.50			5134.86	0.91	3.40	-1.567	26.08
-∞							5614.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4519

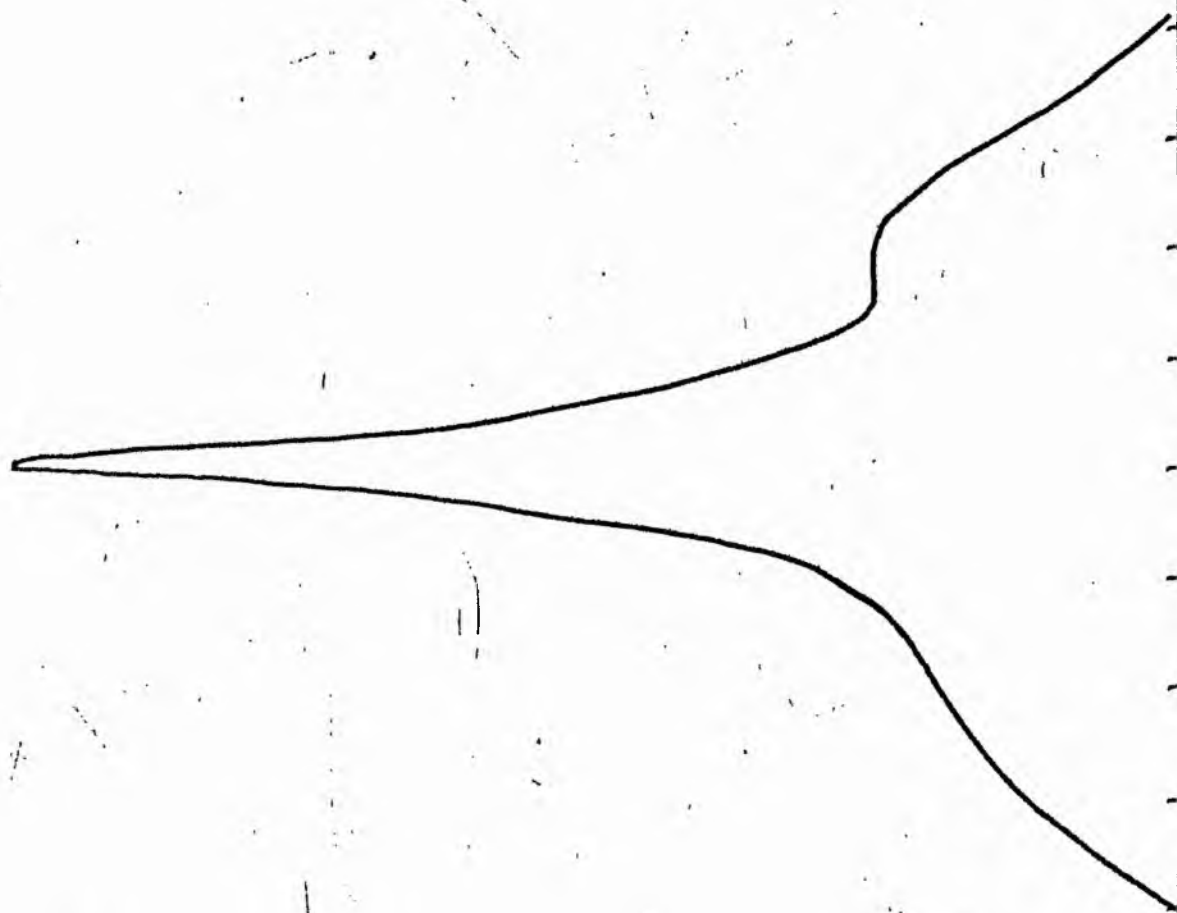
V-FILTER

Total luminosity	L_T	= 1.56
Total apparent magnitude	m_T	= 11.71
Apparent central surface brightness	μ_o	= 19.35
Major axis at threshold	$2a_m$	= 3.71
Minor axis at threshold	$2b_m$	= 3.64
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 2.66
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.85
Gradient of exponential component	$G(a)$	= -1.06
Equivalent gradient of exponential comp....	$G(r^*)$	= -1.17
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.76
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.33
Axis ratio	b/a	= 0.65
Equivalent radius	r_1^*	= 0.28
Surface brightness	μ_1	= 21.15
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.53
Axis ratio	b/a	= 0.91
Equivalent radius	r_e^*	= 0.56
Surface brightness	μ_e	= 22.16
Mean surface brightness	μ_e'	= 12.41
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.78
Axis ratio	b/a	= 1.19
Equivalent radius	r_3^*	= 0.96
Surface brightness	μ_3	= 23.50
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.95 \\ = 1.72 \end{matrix}$

NGC 4526
B-Filter
Axis 1

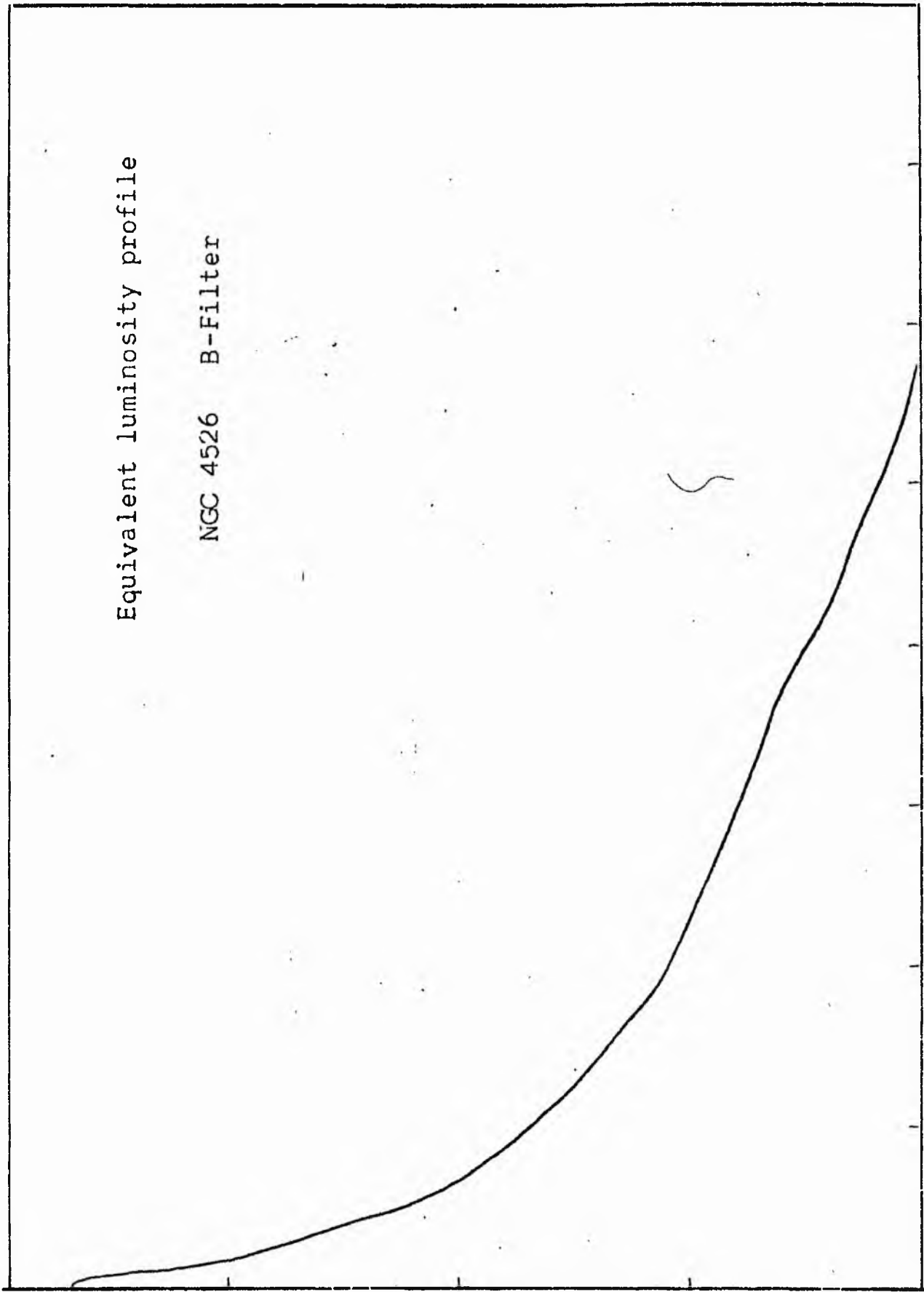


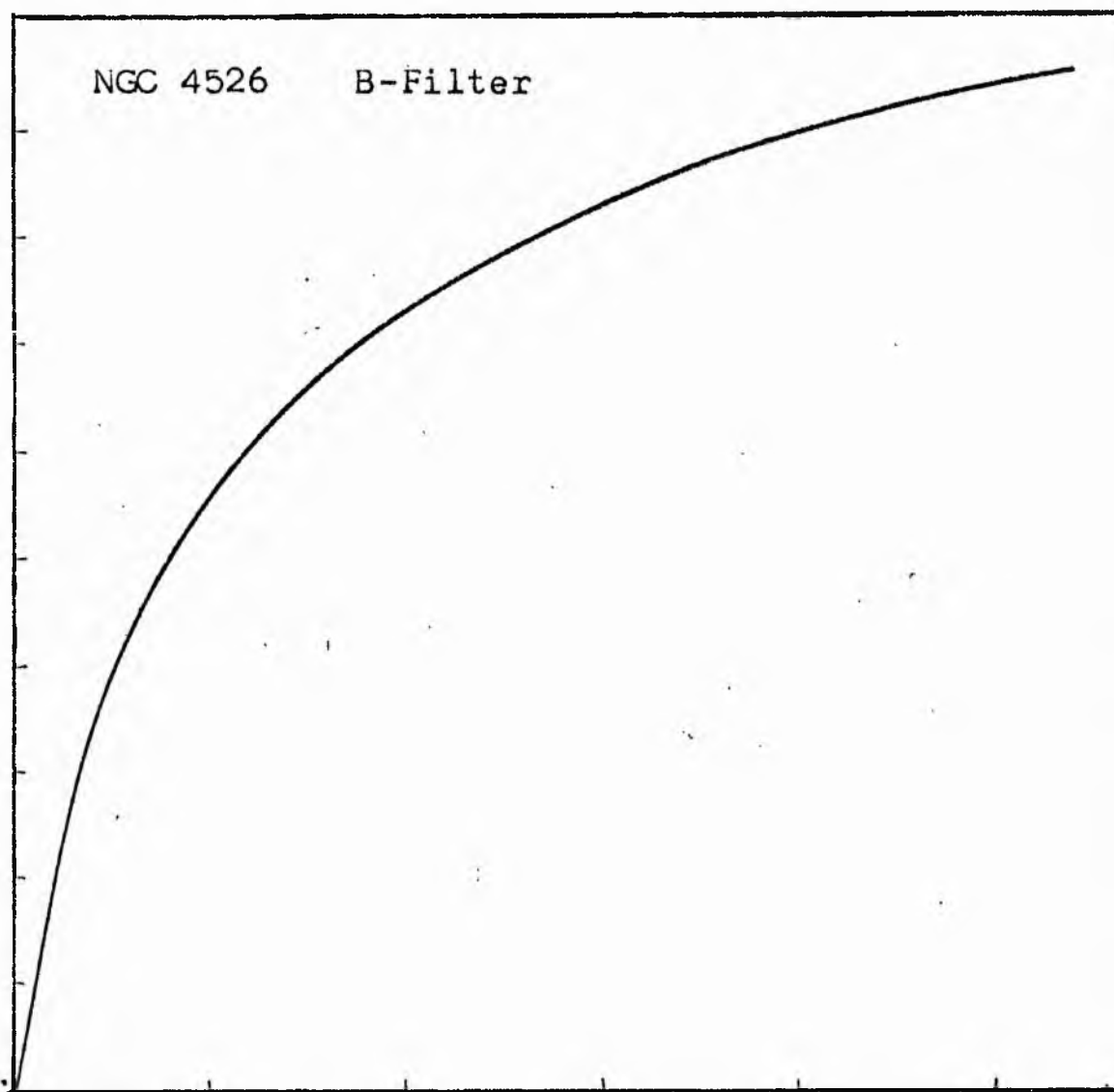
NGC 4526
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4526 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4324
BY COLOUR

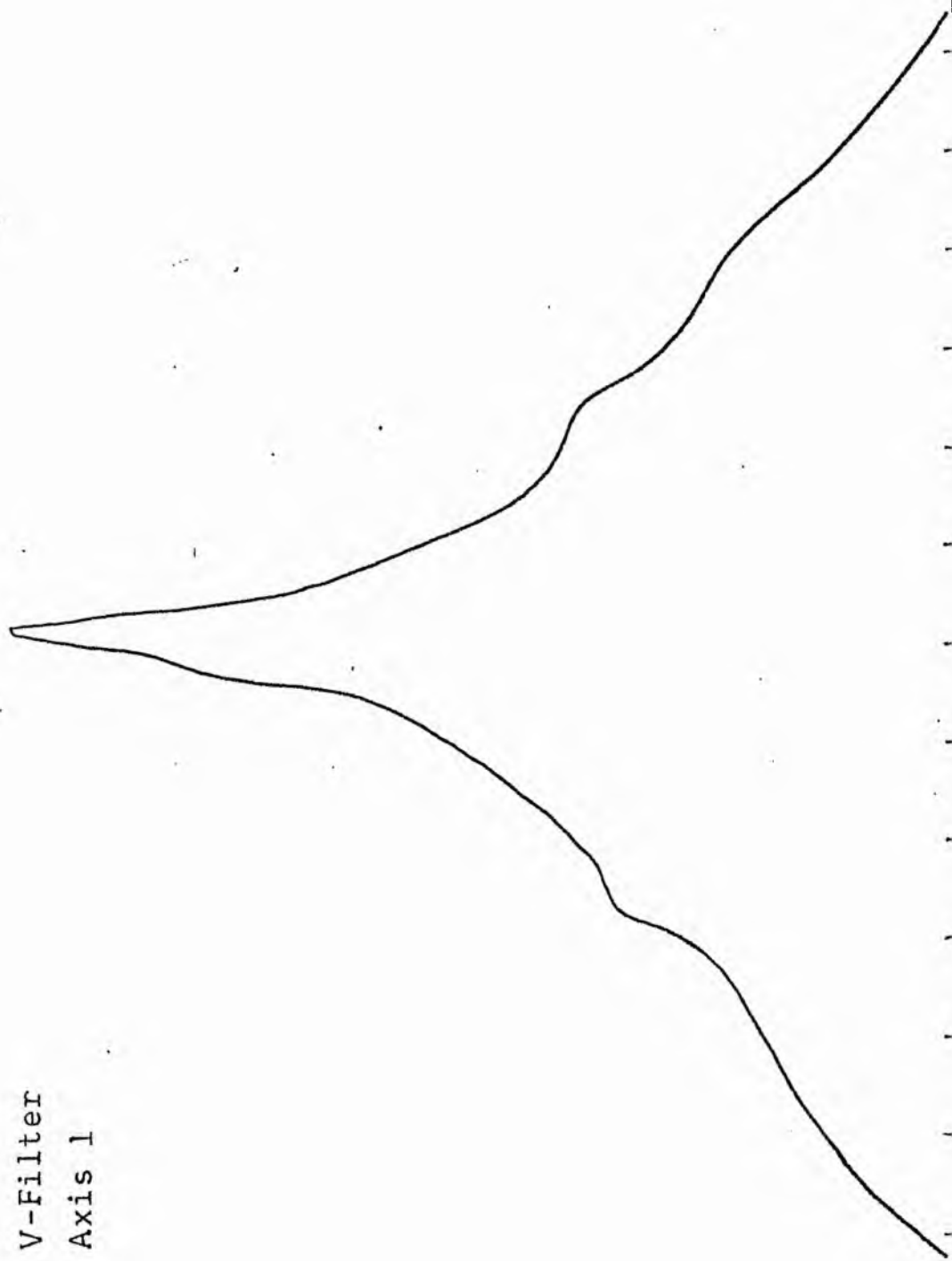
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	Δ
1.66	45.709	42.760	0.0	0.0	31.97	1366.9917	0.0	0.0	0.0	1.748	17.68
1.60	39.811	35.717	3.19	31.97	26.39	942.5388	1366.99	0.04	0.08	1.688	17.83
1.50	31.623	28.371	4.31	58.36	40.16	1139.4209	2309.93	0.07	0.11	1.588	18.08
1.40	25.119	22.534	5.60	98.52	22.24	501.2488	3448.95	0.11	0.15	1.488	18.33
1.30	19.953	17.901	6.20	120.76	51.27	917.7822	3950.20	0.12	0.16	1.388	18.58
1.20	15.849	14.219	7.40	172.03	60.32	857.6702	4867.98	0.15	0.19	1.288	18.83
1.10	12.589	11.295	8.60	232.35	69.37	783.4624	5725.65	0.17	0.23	1.188	17.08
1.00	10.000	8.972	9.80	301.72	57.96	520.0142	6509.11	0.20	0.26	1.088	19.33
0.90	7.943	7.126	10.70	359.68	208.64	1486.8552	7029.12	0.21	0.28	0.988	19.58
0.80	6.310	5.661	13.45	568.32	240.96	1363.9939	8515.97	0.26	0.35	0.888	19.83
0.70	5.012	4.496	16.05	809.28	189.46	851.8843	9879.96	0.30	0.42	0.788	20.08
0.60	3.981	3.572	17.83	998.74	147.34	526.2651	10731.85	0.33	0.47	0.688	20.33
0.50	3.162	2.837	19.10	1146.08	306.12	868.4775	11258.11	0.34	0.50	0.588	20.58
0.40	2.512	2.254	21.50	1452.20	312.40	704.0110	12126.59	0.37	0.56	0.488	20.83
0.30	1.995	1.790	23.70	1764.60	491.81	880.3823	12830.59	0.39	0.62	0.388	21.08
0.20	1.585	1.422	26.80	2256.42	774.35	1101.0535	13710.97	0.42	0.70	0.288	21.33
0.10	1.259	1.129	31.06	3030.77	686.87	775.7842	14812.02	0.45	0.81	0.188	21.58
-0.00	1.000	0.897	34.40	3717.63	987.50	885.9409	15587.80	0.48	0.90	0.088	21.83
-0.10	0.794	0.713	38.70	4705.13	1319.07	940.0239	16473.74	0.50	1.01	-0.012	22.08
-0.20	0.631	0.566	43.79	6024.20	1814.07	1026.8921	17413.77	0.53	1.15	-0.112	22.33
-0.30	0.501	0.450	49.95	7838.28	2172.73	976.9558	18440.66	0.56	1.31	-0.212	22.58
-0.40	0.398	0.357	56.45	10011.00	2636.74	941.7529	19417.61	0.59	1.48	-0.312	22.83
-0.50	0.316	0.284	63.45	12647.74	3287.31	932.6338	20359.36	0.62	1.66	-0.412	23.08
-0.60	0.251	0.225	71.22	15935.05	3910.59	881.2778	21291.99	0.65	1.87	-0.512	23.33
-0.70	0.200	0.179	79.48	19845.65	4433.25	256.5625	22173.27	0.68	2.08	-0.612	23.58
-0.80	0.158	0.142	82.30	21278.90	8037.04	1142.7903	22429.83	0.69	2.16	-0.712	23.83
-0.90	0.126	0.113	96.60	29315.95	8835.68	997.9539	23572.62	0.72	2.53	-0.812	24.08
-1.00	0.100	0.090	110.20	38151.62	17169.59	1540.3916	24570.57	0.75	2.89	-0.912	24.33
-1.10	0.079	0.071	132.70	55321.22	14144.66	1008.0076	26110.96	0.80	3.48	-1.012	24.58
-1.20	0.063	0.057	148.70	69465.87	16271.44	921.0815	27118.97	0.83	3.90	-1.112	24.83
-1.30	0.050	0.045	165.20	85737.31	15485.56	696.3054	28040.05	0.86	4.33	-1.212	25.08
-1.40	0.040	0.036	179.50	101222.87	12786.31	456.6873	28736.35	0.88	4.71	-1.312	25.33
-1.50	0.032	0.028	190.50	114009.19	14815.75	420.3372	29193.04	0.89	4.99	-1.412	25.58
-1.60	0.025	0.023	202.50	128824.94	32203.50	725.7346	29613.37	0.90	5.31	-1.512	25.83
-1.70	0.020	0.018	226.40	161028.44	22196.31	397.3342	30339.11	0.93	5.94	-1.612	26.08
-1.80	0.016	0.014	241.50	183224.75	27028.31	384.3215	30736.44	0.94	6.33	-1.712	26.33
-1.90	0.013	0.011	258.70	210253.06	8868.87	100.1717	31120.76	0.95	6.78	-1.812	26.58
-2.00	0.010		264.10	219121.94			31220.93	0.95	6.92	-1.912	26.83
-∞							32730.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4526

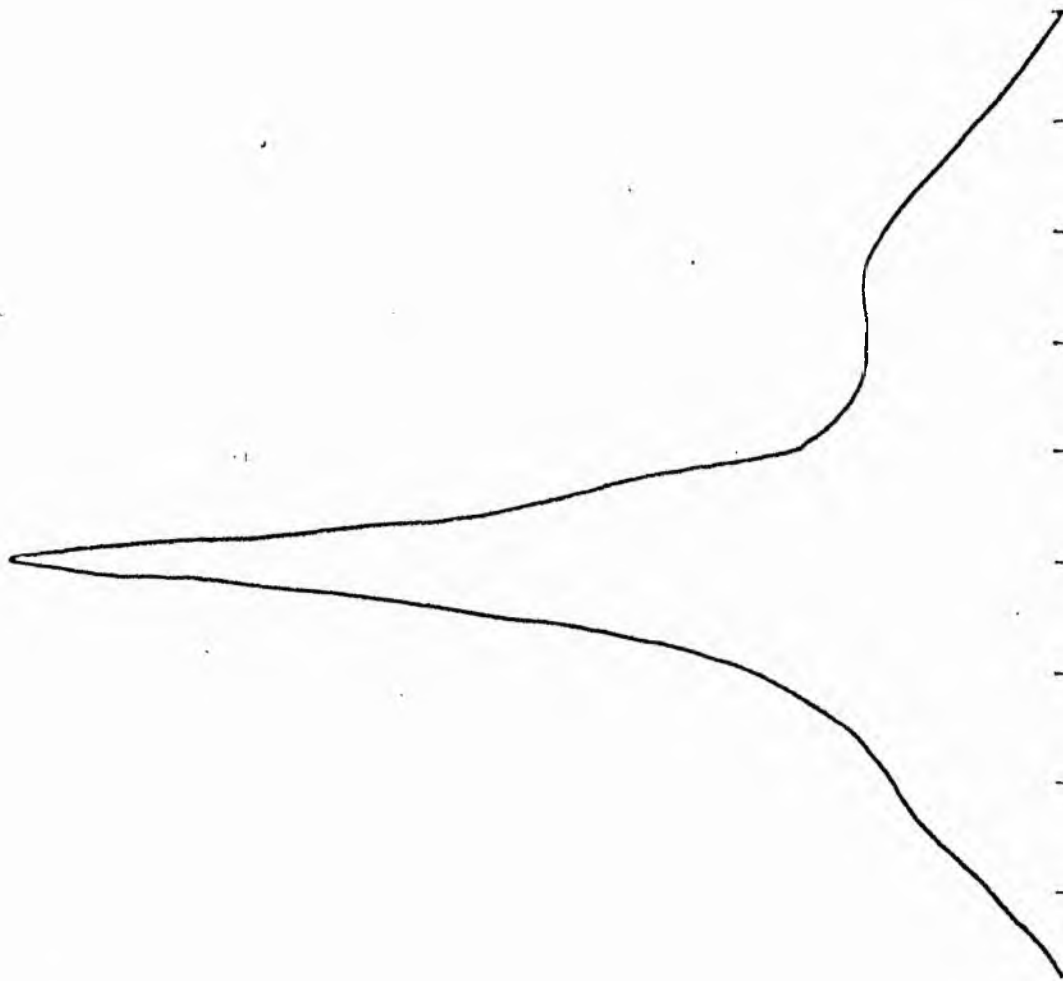
B-FILTER

Total luminosity	L_T	= 9.09
Total apparent magnitude	m_T	= 10.54
Apparent central surface brightness	μ_0	= 17.68
Major axis at threshold	$2a_m$	= 12.38
Minor axis at threshold	$2b_m$	= 6.58
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.54
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.85
Gradient of exponential component	$G(a)$	= -0.48
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.44
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.25
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 0.59
Equivalent radius	r_1^*	= 0.21
Surface brightness	μ_1	= 19.78
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.88
Axis ratio	b/a	= 0.43
Equivalent radius	r_e^*	= 0.64
Surface brightness	μ_e	= 22.08
Mean surface brightness	μ_e'	= 11.54
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.43
Axis ratio	b/a	= 0.58
Equivalent radius	r_3^*	= 1.83
Surface brightness	μ_3	= 24.33
Concentration indices	$\{C_{21}$	= 2.97
	$\{C_{32}$	= 2.88

NGC 4526
V-Filter
Axis 1

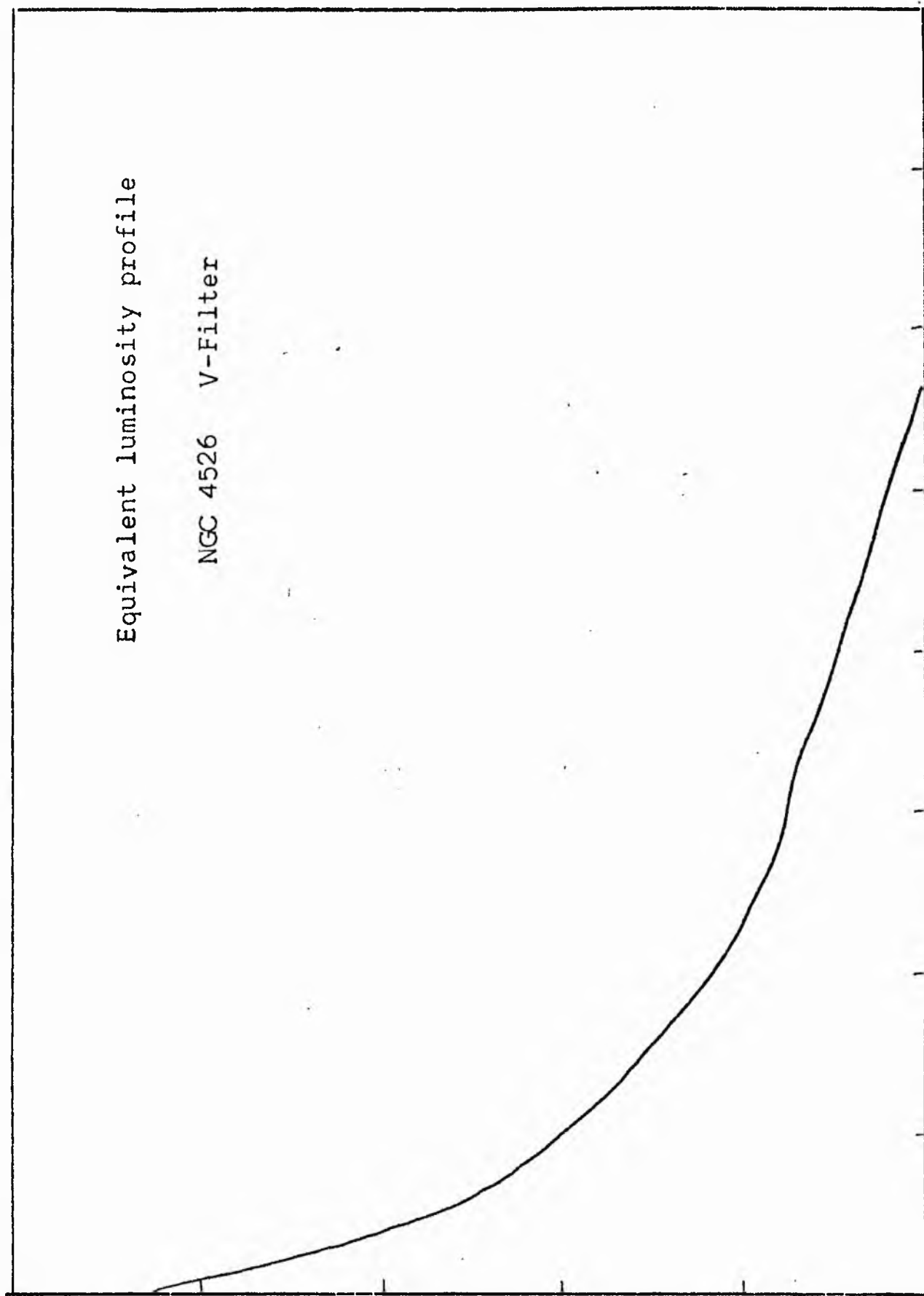


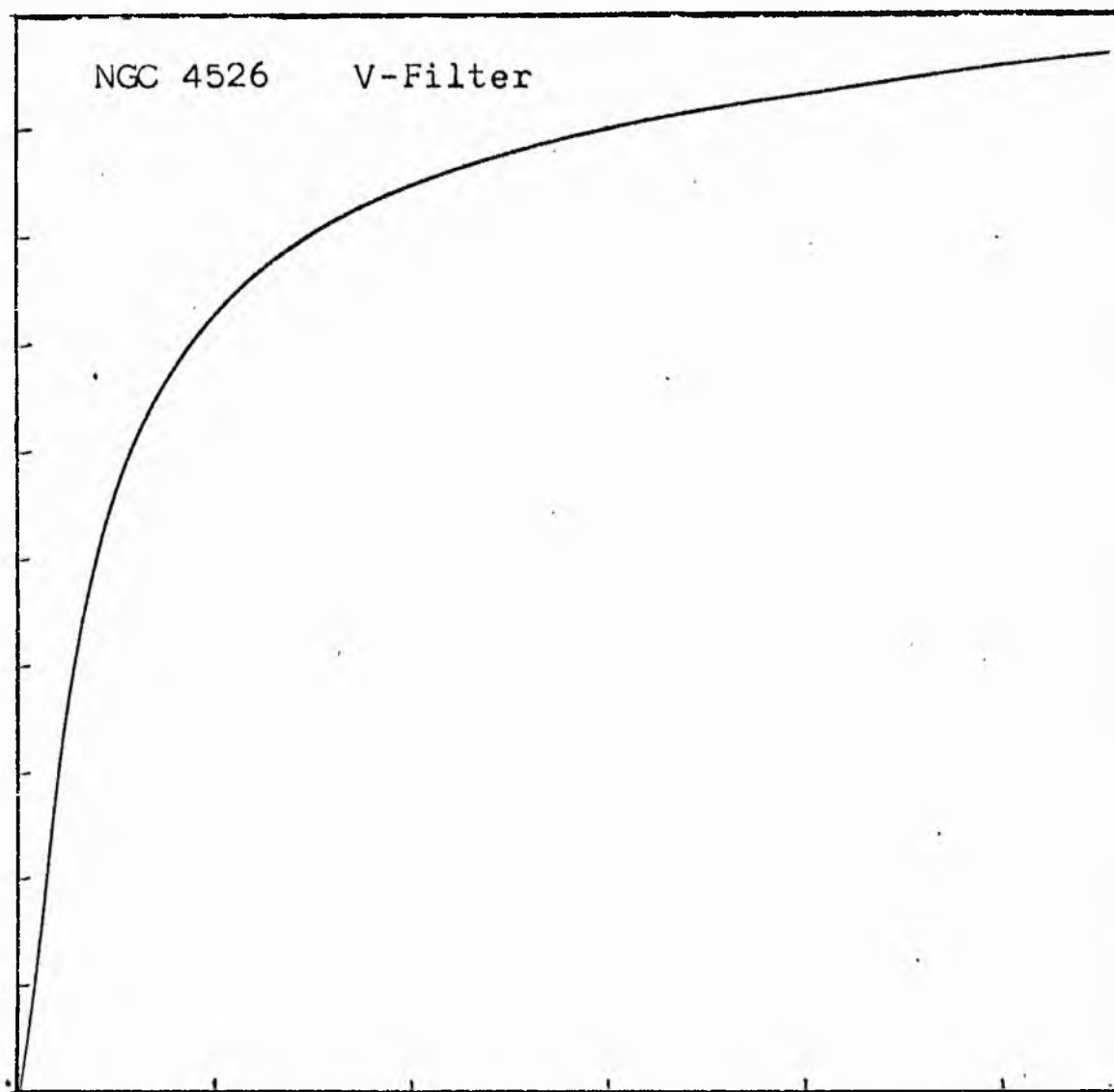
NGC 4526
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4526 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

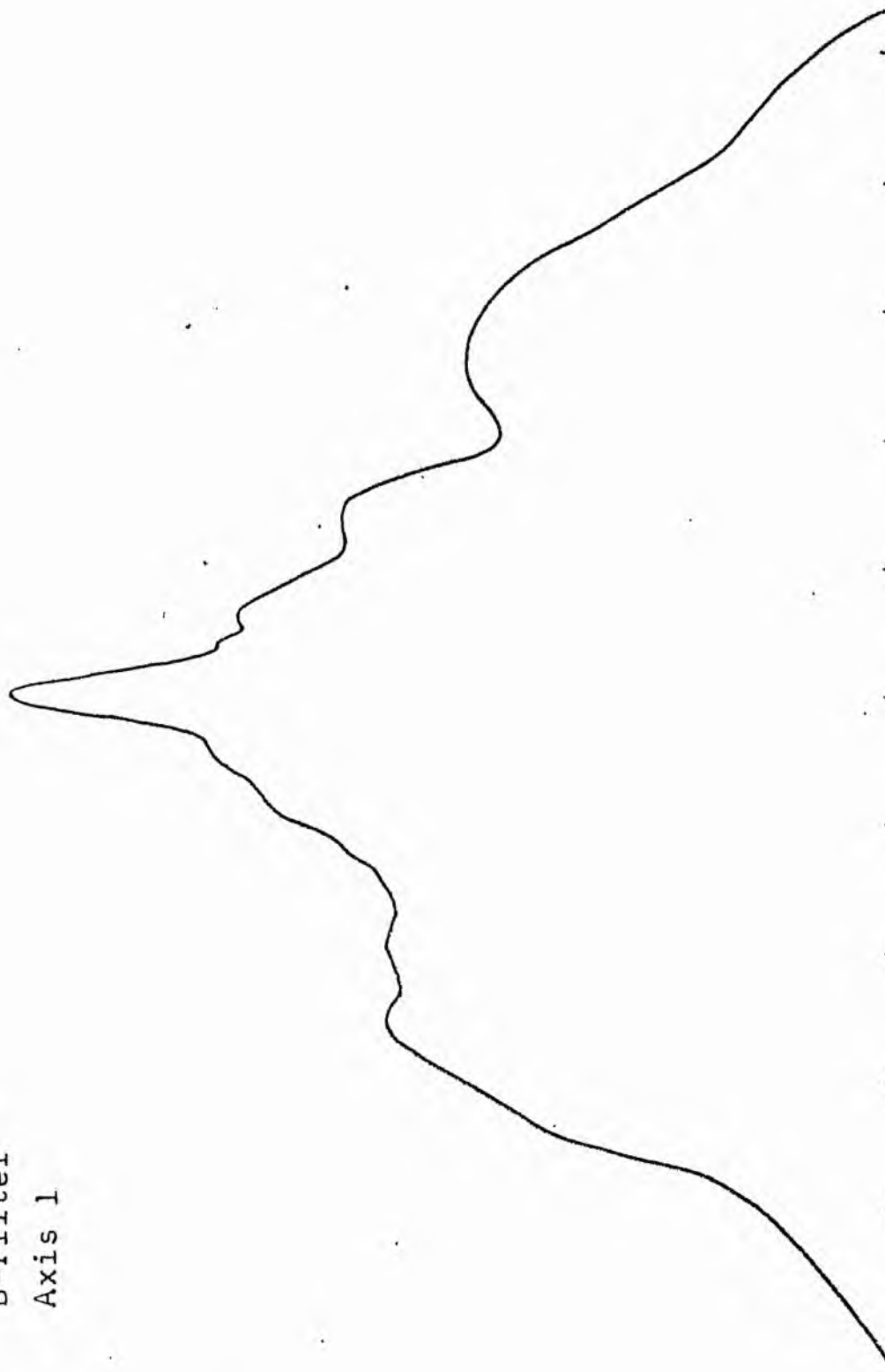
MEAN LUMINOSITY DISTRIBUTION IN NGC 4526
V COLOUR

LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.11	128.825		0.0	0.0	0.13	16.0043	0.0	0.0	0.0	1.186	16.38
2.10	125.892	127.359	0.20	0.13	6.94	784.1753	16.00	0.00	0.01	1.176	16.41
2.00	100.000	112.946	1.50	7.07	40.72	3652.7949	800.18	0.01	0.08	1.076	16.66
1.90	79.433	89.716	3.90	47.78	61.58	4388.0937	4452.97	0.07	0.20	0.976	16.91
1.80	63.095	71.264	5.90	109.36	49.01	2774.2422	8841.07	0.14	0.31	0.876	17.16
1.70	50.118	56.607	7.10	158.37	68.61	3085.1208	11615.31	0.18	0.37	0.776	17.41
1.60	39.810	44.964	8.50	226.98	103.73	3704.7900	14700.43	0.23	0.44	0.676	17.66
1.50	31.623	35.717	10.26	330.71	56.37	1599.1851	18455.21	0.29	0.53	0.576	17.91
1.40	25.119	28.371	11.10	387.08	111.68	2516.8513	20004.40	0.31	0.58	0.476	18.16
1.30	19.952	22.536	12.60	498.76	208.10	3725.1074	22521.25	0.35	0.65	0.376	18.41
1.20	15.849	17.901	15.00	706.86	237.74	3380.4265	26246.35	0.41	0.78	0.276	18.66
1.10	12.589	14.219	17.34	944.60	73.28	827.6233	29626.78	0.47	0.90	0.176	18.91
1.00	10.000	11.295	18.00	1017.88	206.30	1850.8223	30454.40	0.48	0.93	0.076	19.16
0.90	7.943	8.972	19.74	1224.18	301.89	2151.3674	32305.22	0.51	1.02	-0.024	19.41
0.80	6.310	7.126	22.04	1526.06	268.44	1519.5693	34456.59	0.54	1.14	-0.124	19.66
0.70	5.012	5.661	23.90	1794.51	229.13	1030.2656	35976.15	0.56	1.24	-0.224	19.91
0.60	3.981	4.496	25.38	2023.64	483.55	1727.0662	37006.42	0.58	1.32	-0.324	20.16
0.50	3.162	3.572	28.25	2507.19	550.96	1563.1135	38733.48	0.61	1.46	-0.424	20.41
0.40	2.512	2.837	31.20	3058.15	599.21	1350.3486	40296.59	0.63	1.62	-0.524	20.66
0.30	1.995	2.254	34.12	3657.36	550.99	986.3008	41646.94	0.65	1.77	-0.624	20.91
0.20	1.585	1.790	36.60	4208.35	944.65	1343.1917	42633.24	0.67	1.90	-0.724	21.16
0.10	1.259	1.422	40.50	5153.00	1288.14	1454.8926	43976.43	0.69	2.10	-0.824	21.41
-0.00	1.000	1.129	45.28	6441.14	1955.99	1754.8296	45431.32	0.71	2.35	-0.924	21.66
-0.10	0.794	0.897	51.70	8397.13	1831.40	1305.1230	47186.15	0.74	2.68	-1.024	21.91
-0.20	0.631	0.713	57.06	10228.53	1992.31	1127.7817	48491.27	0.76	2.96	-1.124	22.16
-0.30	0.501	0.566	62.37	12220.84	2046.42	920.1577	49619.05	0.78	3.23	-1.224	22.41
-0.40	0.398	0.450	67.39	14267.26	2483.45	887.0010	50539.21	0.79	3.49	-1.324	22.66
-0.50	0.316	0.357	73.02	16750.71	2485.45	705.1375	51426.21	0.81	3.78	-1.424	22.91
-0.60	0.251	0.284	78.25	19236.17	2762.33	622.5068	52131.35	0.82	4.06	-1.524	23.16
-0.70	0.200	0.225	83.68	21998.50	2551.66	456.7632	52753.85	0.83	4.34	-1.624	23.41
-0.80	0.158	0.179	88.40	24550.16	10217.97	1452.8909	53210.61	0.84	4.58	-1.724	23.66
-0.90	0.126	0.142	105.20	34768.13	3522.09	397.8042	54663.50	0.86	5.45	-1.824	23.91
-1.00	0.100	0.113	110.40	38290.22	16117.62	1446.0071	55061.30	0.86	5.72	-1.924	24.16
-1.10	0.079	0.090	131.60	54407.85	13941.40	993.5195	56507.31	0.89	6.82	-2.024	24.41
-1.20	0.063	0.071	147.50	68349.25	17180.56	972.5405	57500.83	0.90	7.64	-2.124	24.66
-1.30	0.050	0.057	165.00	85529.81	20139.37	915.5591	58473.37	0.92	8.55	-2.224	24.91
-1.40	0.040	0.045	183.40	105669.19	25199.50	970.0432	59378.93	0.93	9.51	-2.324	25.16
-1.50	0.032	0.036	204.10	130868.69	12466.31	353.6797	60278.97	0.95	10.58	-2.424	25.41
-1.60	0.025	0.028	213.60	143335.00	20120.81	453.4382	60632.65	0.95	11.07	-2.524	25.66
-1.70	0.020	0.023	228.10	163455.81	26506.25	474.4841	61086.09	0.96	11.82	-2.624	25.91
-1.80	0.016	0.018	245.90	189962.06	21919.69	311.6794	61560.57	0.97	12.74	-2.724	26.16
-1.90	0.013	0.014	259.70	211881.75	24320.62	274.6938	61872.24	0.97	13.46	-2.824	26.41
-2.00	0.010	0.011	274.20	236202.37			62146.93	0.98	14.21	-2.924	26.66
-∞							63706.00	111			∞

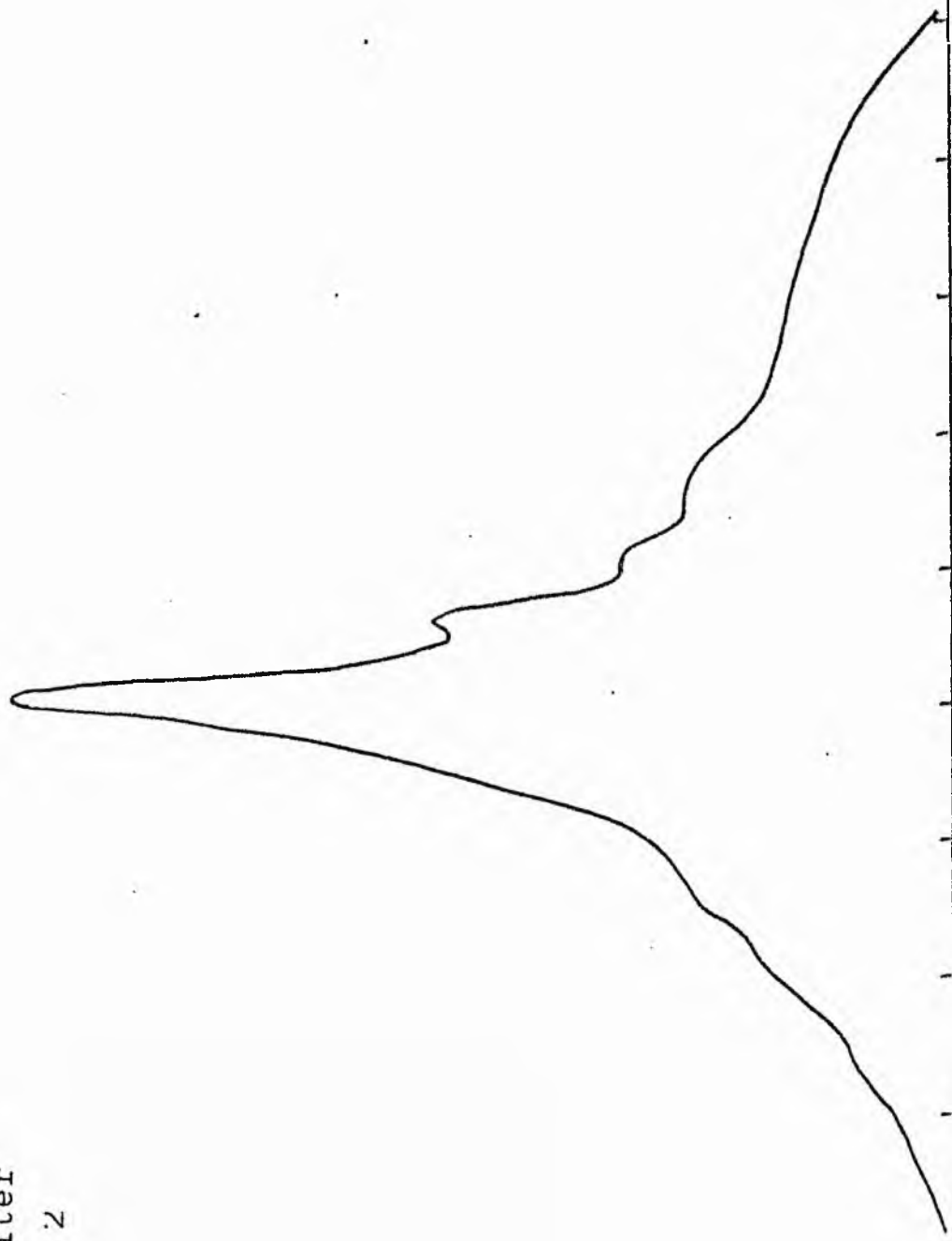
PHOTOMETRIC PARAMETERS OF NGC 4526
V-FILTER

Total luminosity	L_T	= 17.70
Total apparent magnitude	m_T	= 9.65
Apparent central surface brightness	μ_o	= 16.38
Major axis at threshold	$2a_m$	= 10.38
Minor axis at threshold	$2b_m$	= 7.98
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 7.81
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.92
Gradient of exponential component	$G(a)$	= -0.52
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.42
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.35
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.23
Axis ratio	b/a	= 0.57
Equivalent radius	r_1^*	= 0.15
Surface brightness	μ_1	= 17.75
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.47
Axis ratio	b/a	= 0.52
Equivalent radius	r_e^*	= 0.32
Surface brightness	μ_e	= 19.34
Mean surface brightness	μ_e'	= 9.15
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.40
Axis ratio	b/a	= 0.44
Equivalent radius	r_3^*	= 0.90
Surface brightness	μ_3	= 22.04
Concentration indices	$\begin{Bmatrix} C_{21} \\ C_{32} \end{Bmatrix}$	$\begin{matrix} = 2.13 \\ = 2.80 \end{matrix}$

NGC 4527
B-Filter
Axis 1

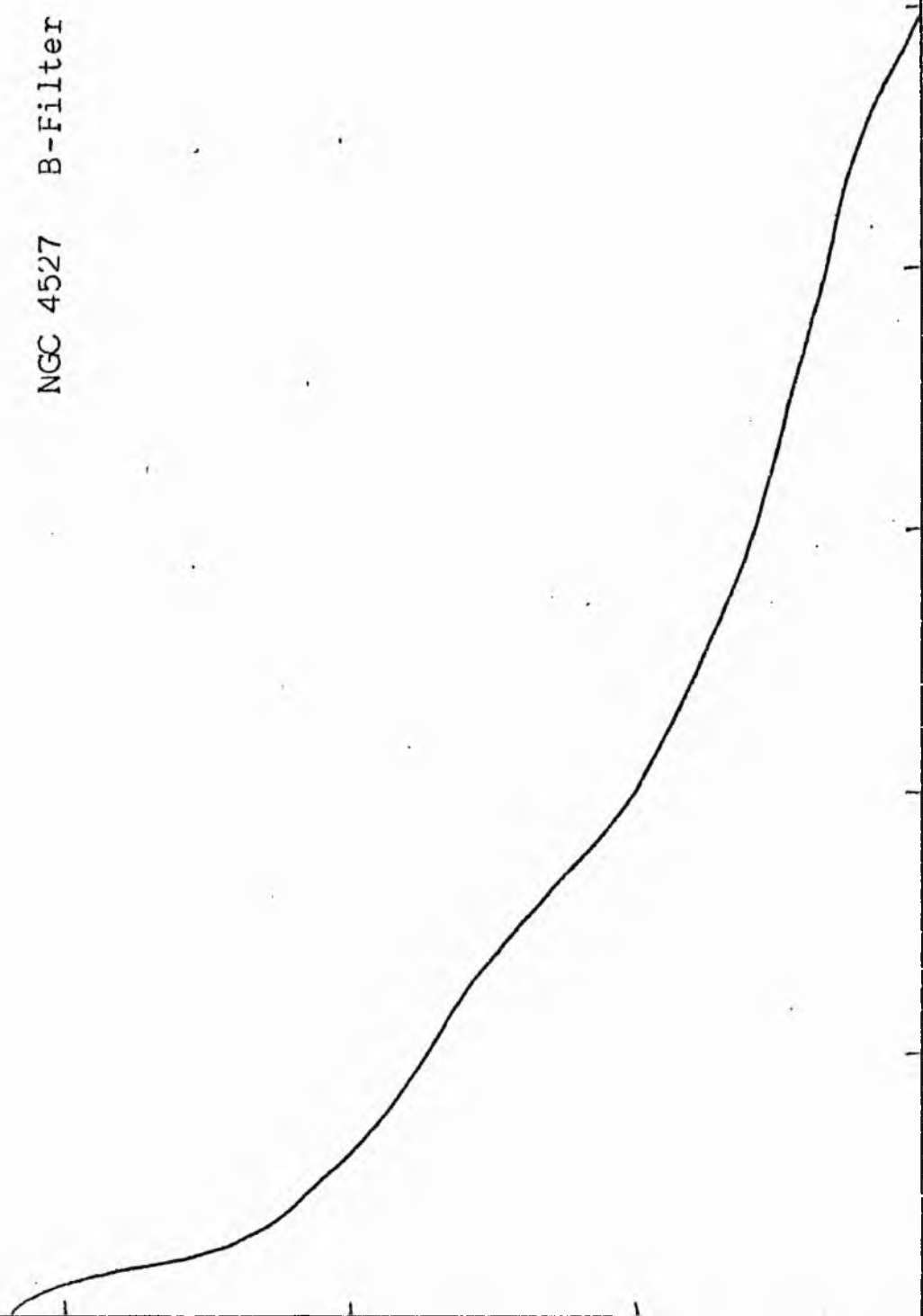


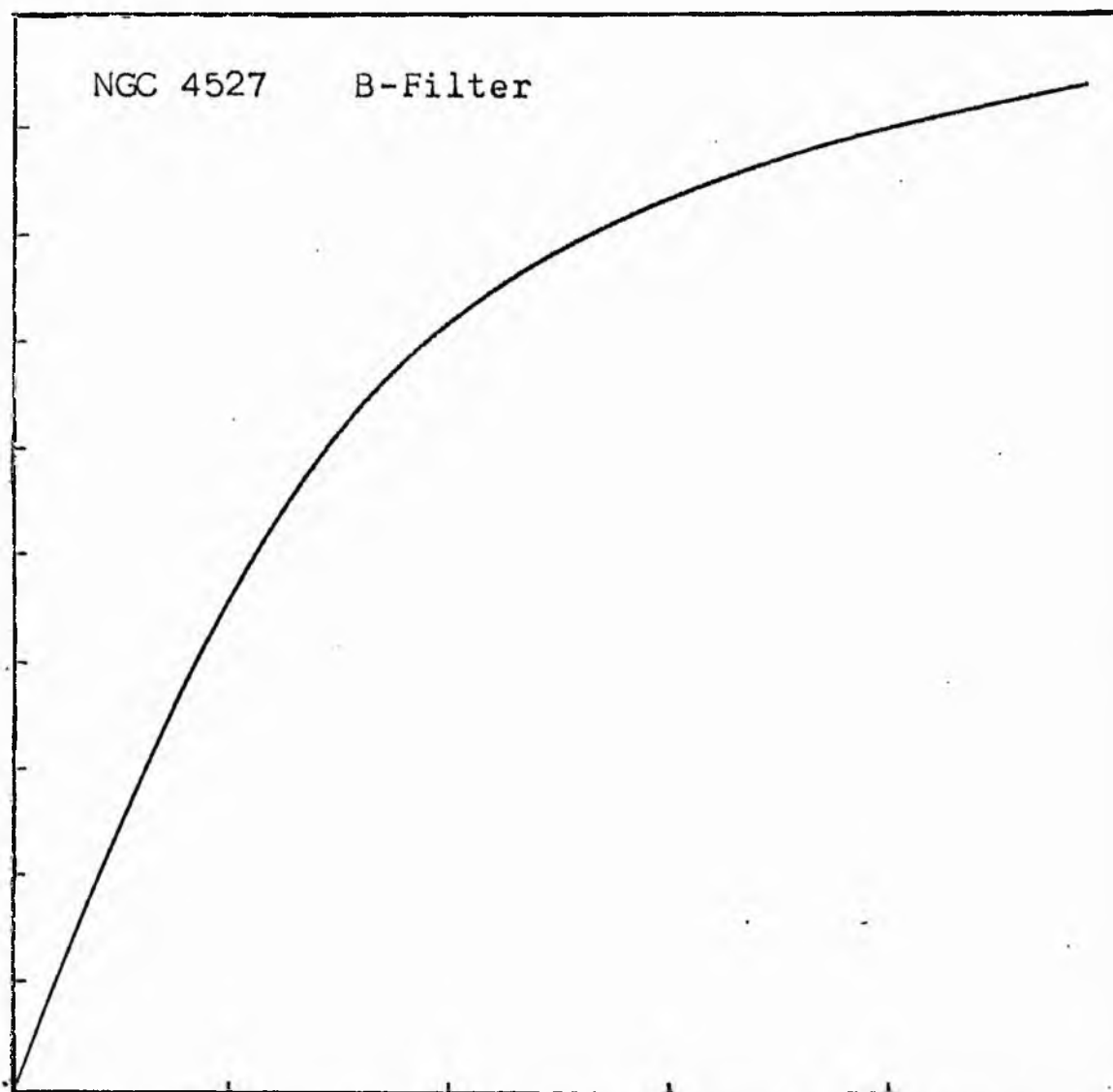
NGC 4527
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4527 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

B COLOUR

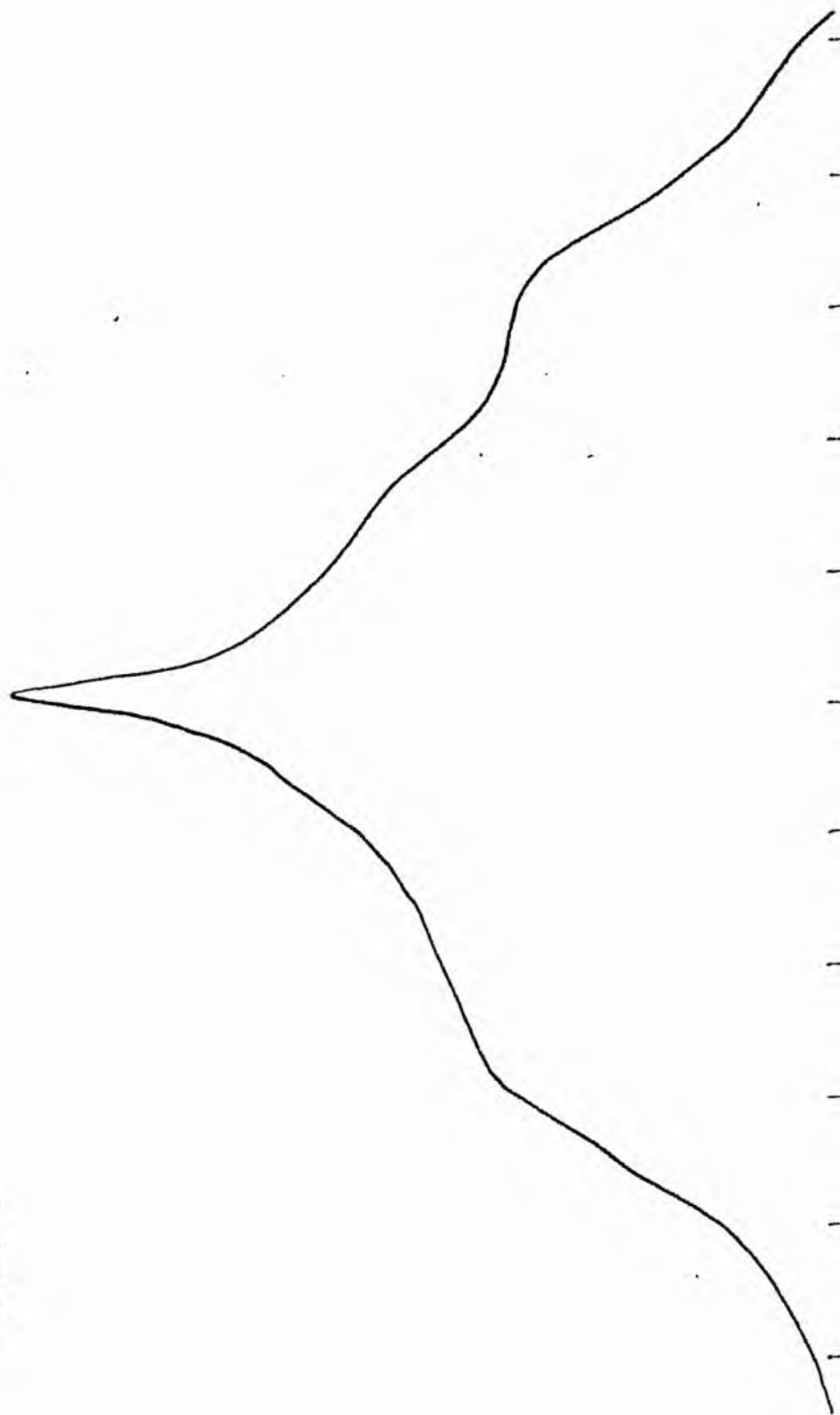
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(M)	Q	LOG J	AL
1.17	14.791		0.0	0.0							
		13.690									
1.10	12.589		3.19	31.97	31.97	437.6616	C.C	0.0	0.0	1.490	19.28
		11.295					437.66	0.02	0.06	1.420	19.46
1.00	10.000		6.10	116.90	84.93	959.2446					
		8.972					1346.91	0.06	0.11	1.320	19.71
0.90	7.943		7.20	162.86	45.96	412.3496					
		7.126					1819.26	0.08	0.13	1.220	19.76
0.80	6.310		8.40	221.67	58.81	419.1084					
		5.661					2228.36	0.10	0.16	1.120	20.21
0.70	5.012		8.70	237.79	16.12	91.2303					
		4.496					2319.59	0.10	0.16	1.020	20.46
0.60	3.981		10.10	320.47	82.69	371.7469					
		3.572					2641.34	0.12	0.19	0.920	20.71
0.50	3.162		11.70	430.05	109.58	391.3794					
		2.837					3082.77	0.14	0.22	0.820	20.76
0.40	2.512		13.40	564.10	134.05	340.3152					
		2.254					3463.09	0.15	0.25	0.720	21.21
0.30	1.995		16.18	822.44	254.34	582.1868					
		1.790					4045.27	0.18	0.36	0.620	21.46
0.20	1.585		22.42	1579.14	750.70	1354.5457					
		1.422					5349.82	0.24	0.42	0.520	21.71
0.10	1.259		27.01	2291.92	712.78	1013.5020					
		1.129					6413.32	0.29	0.50	0.420	21.76
-0.00	1.000		29.77	2704.24	492.33	556.0637					
		0.897					6969.38	0.31	0.56	0.320	22.21
-0.10	0.794		35.75	4015.15	1230.91	1134.3259					
		0.713					6013.76	0.36	0.67	0.220	22.46
-0.20	0.631		44.33	6173.69	2158.54	1538.2661					
		0.566					4611.97	0.43	0.83	0.120	22.71
-0.30	0.501		52.11	8530.84	2357.15	1334.3157					
		0.450					10946.28	0.49	0.97	0.020	22.96
-0.40	0.398		60.00	11329.73	2770.89	1249.5217					
		0.357					12145.80	0.54	1.12	-0.080	23.21
-0.50	0.316		67.99	14522.44	3212.71	1147.4746					
		0.284					13343.27	0.60	1.77	-0.180	23.46
-0.60	0.251		74.21	17301.13	2778.69	788.3379					
		0.225					14131.61	0.63	1.38	-0.280	23.71
-0.70	0.200		80.20	20206.85	2905.71	654.8247					
		0.179					14786.43	0.66	1.50	-0.380	23.96
-0.80	0.158		85.70	23073.39	2866.54	513.1335					
		0.142					15299.57	0.68	1.60	-0.480	24.21
-0.90	0.126		92.30	26764.12	3690.73	524.7891					
		0.113					15824.36	0.71	1.72	-0.580	24.46
-1.00	0.100		98.70	30604.41	3840.29	433.7463					
		0.090					16258.10	0.73	1.84	-0.680	24.71
-1.10	0.079		114.30	41043.29	10438.87	936.5391					
		0.071					17194.64	0.77	2.13	-0.780	24.96
-1.20	0.063		124.00	48305.13	7261.84	517.5112					
		0.057					17712.15	0.79	2.31	-0.880	25.21
-1.30	0.050		135.00	57255.53	8950.40	506.6587					
		0.045					18218.80	0.81	2.52	-0.980	25.46
-1.40	0.040		143.00	64242.43	6986.90	314.1653					
		0.036					18532.97	0.83	2.67	-1.080	25.71
-1.50	0.032		164.70	85219.06	20976.63	749.2212					
		0.028					19282.19	0.86	3.07	-1.180	25.96
-1.60	0.025		180.30	102127.12	16908.06	479.6990					
		0.023					19761.88	0.88	3.36	-1.280	26.21
-1.70	0.020		202.10	128316.44	26189.31	590.2009					
		0.018					20352.08	0.91	3.77	-1.380	26.46
-1.80	0.016		217.80	149027.19	20710.75	370.7422					
		0.014					20722.82	0.93	4.06	-1.480	26.71
-1.90	0.013		229.40	165324.25	16297.06	231.7321					
		0.011					20954.55	0.94	4.28	-1.580	26.96
-2.00	0.010		246.30	190580.56	25256.31	285.2642					
							21239.82	0.95	4.59	-1.680	27.21
-∞							22389.00	(1)		∞	

PHOTOMETRIC PARAMETERS OF NGC 4527

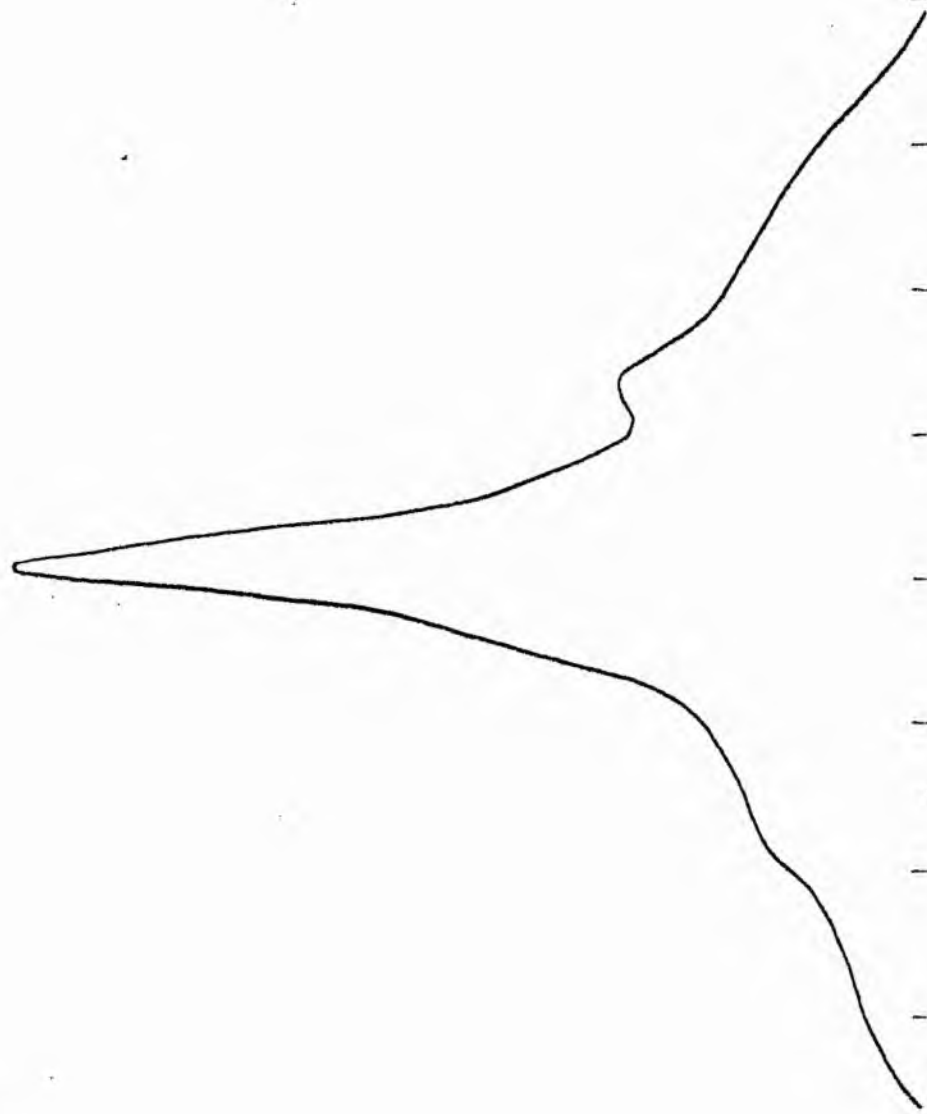
B-FILTER

Total luminosity	L_T	= 6.22
Total apparent magnitude	m_T	= 11.33
Apparent central surface brightness	μ_0	= 19.28
Major axis at threshold	$2a_m$	= 8.77
Minor axis at threshold	$2b_m$	= 7.65
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.05
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.77
Gradient of exponential component	$G(a)$	= -0.53
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.59
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.76
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.80
Axis ratio	b/a	= 0.31
Equivalent radius	r_1^*	= 0.39
Surface brightness	μ_1	= 21.76
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.81
Axis ratio	b/a	= 0.28
Equivalent radius	r_e^*	= 0.89
Surface brightness	μ_e	= 23.01
Mean surface brightness	μ_e'	= 13.07
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.88
Axis ratio	b/a	= 0.39
Equivalent radius	r_3^*	= 1.79
Surface brightness	μ_3	= 24.83
Concentration indices	$\{C_{21}$	= 2.30
	C_{32}	= 2.00

NGC 4527
V-Filter
Axis 1

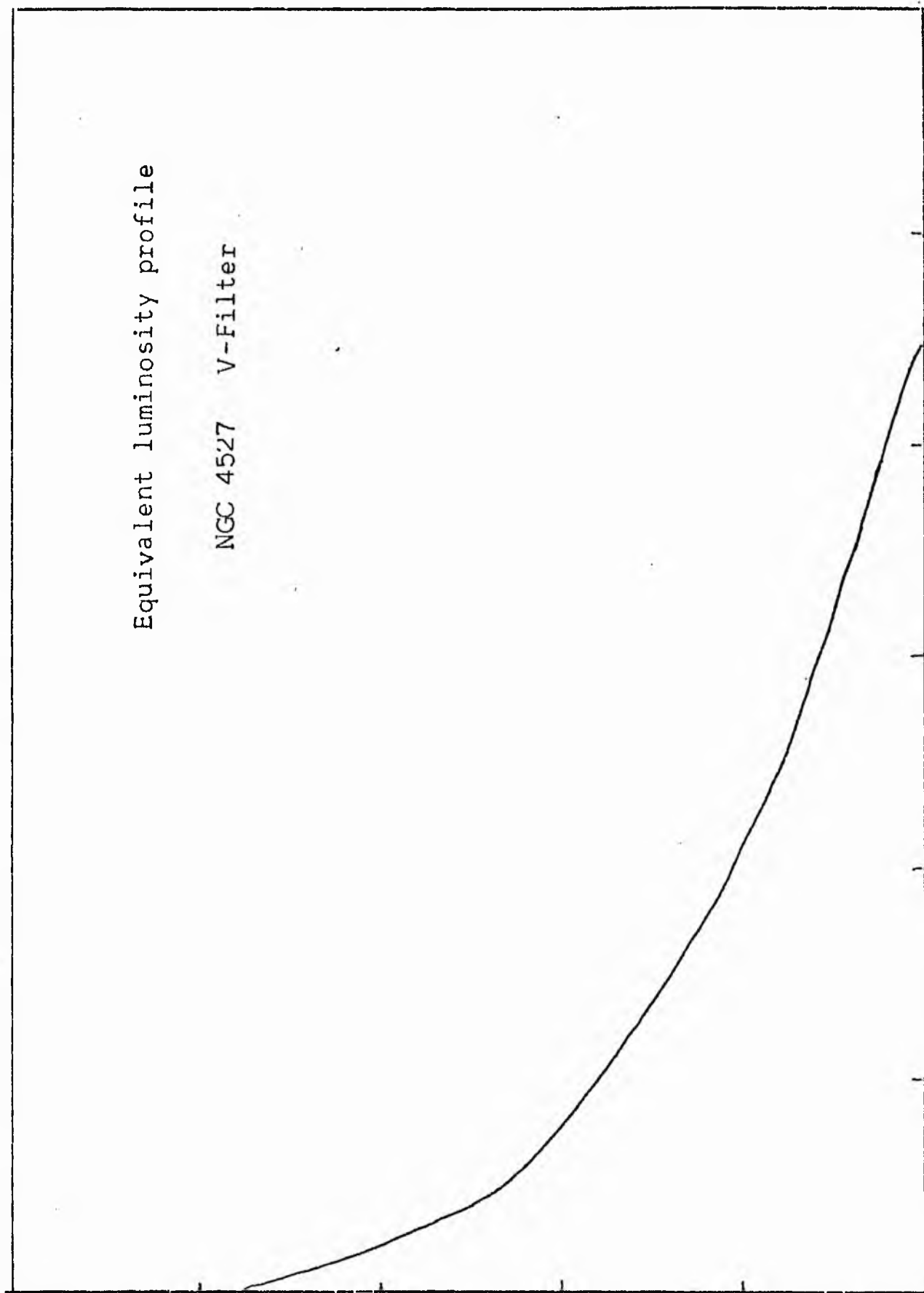


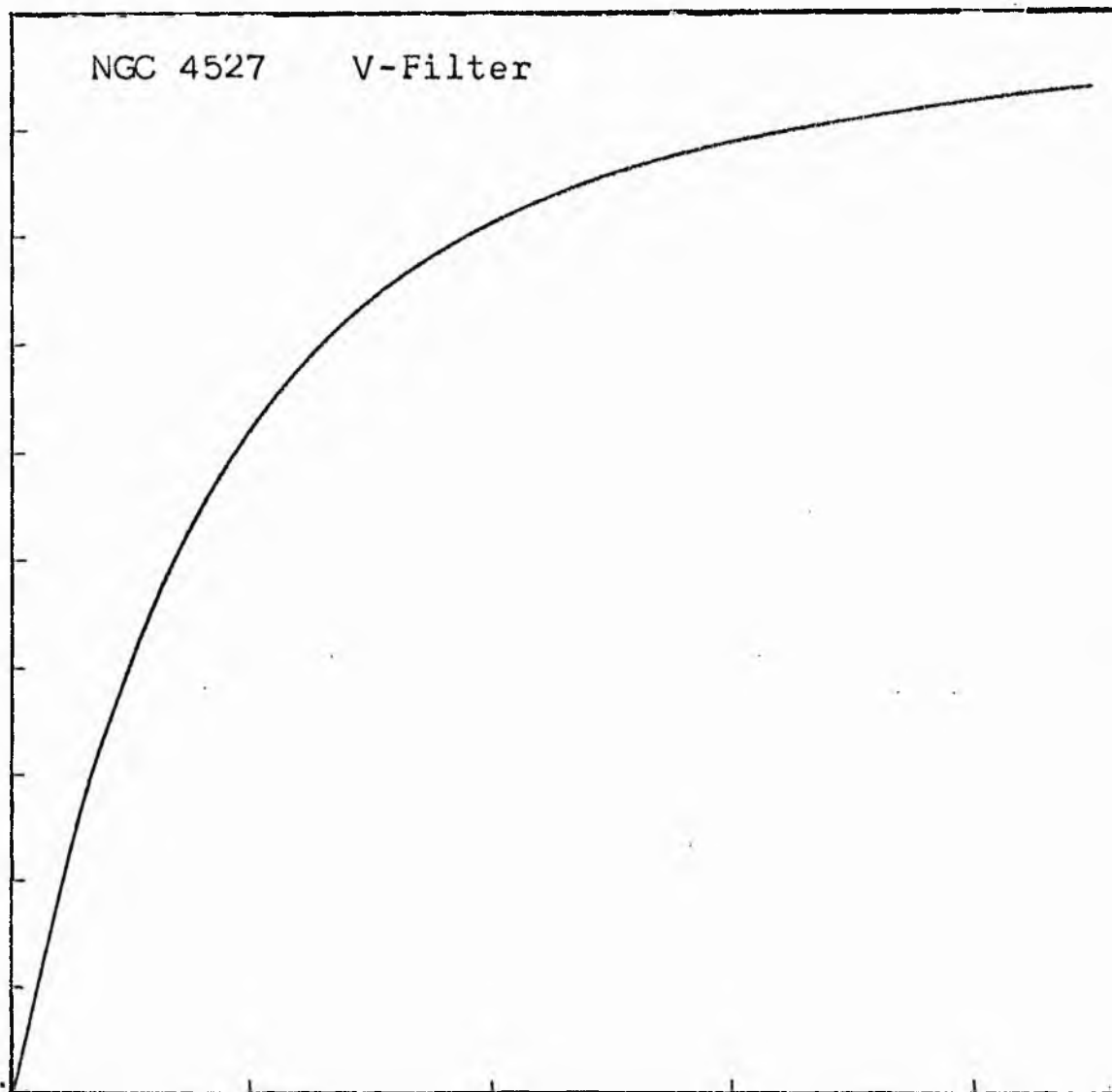
NGC 4527
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4527 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4527
V COLOUR

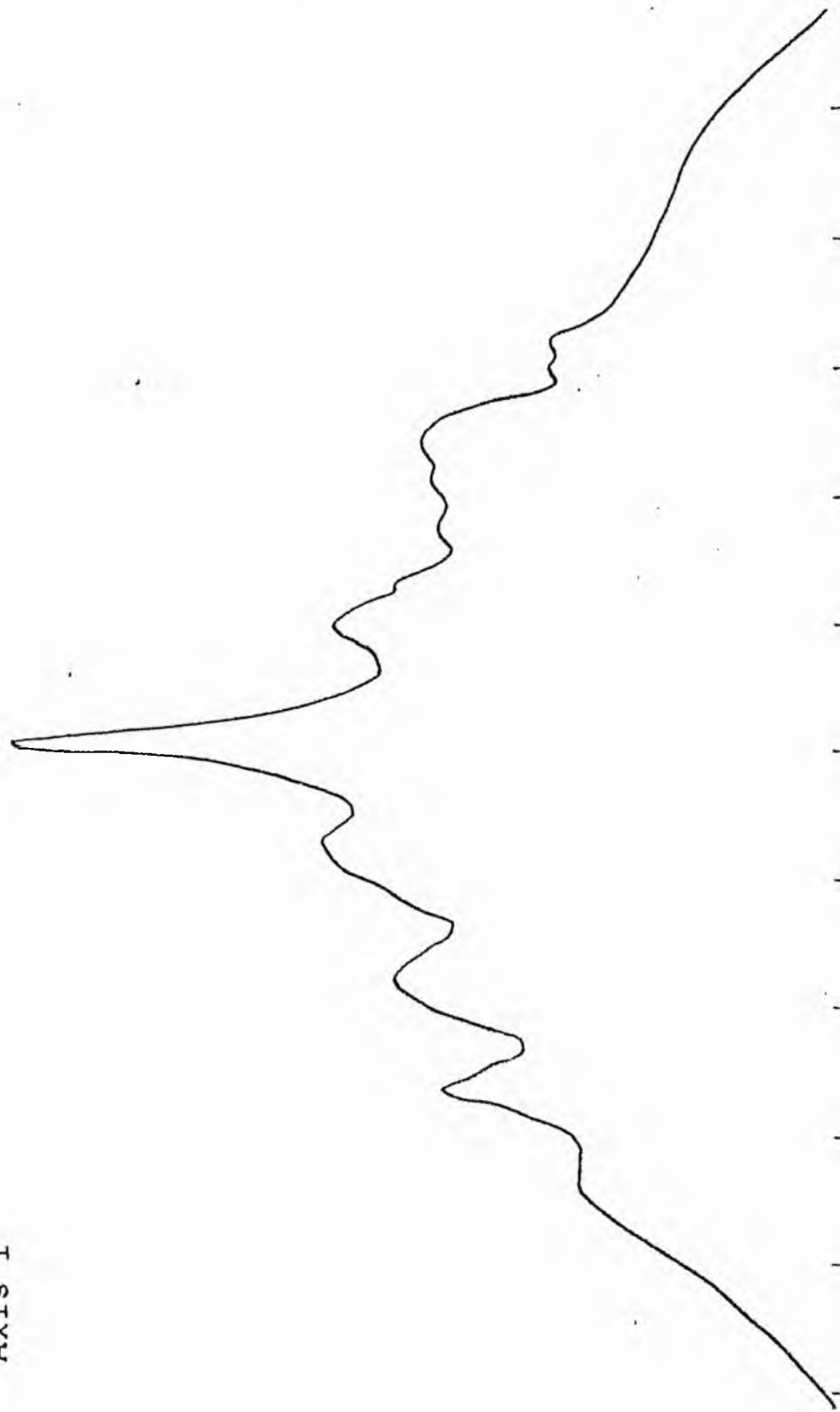
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.65	44.668		0.0	0.0			0.0	0.0	0.0	1.537	17.65
1.60	39.811	42.239	1.50	7.07	7.07	298.5730	298.57	0.01	0.04	1.487	17.78
1.50	31.623	35.717	2.40	18.10	11.03	393.8469	692.42	0.02	0.07	1.387	18.03
1.40	25.119	28.371	3.60	40.72	22.62	641.7307	1334.15	0.04	0.11	1.287	18.28
1.30	19.953	22.536	4.70	69.40	28.68	646.3843	1980.53	0.07	0.14	1.187	18.53
1.20	15.849	17.901	7.50	176.71	107.32	1921.0459	3901.58	0.13	0.22	1.087	18.78
1.10	12.589	14.219	8.50	226.98	50.27	714.7253	4616.30	0.15	0.25	0.987	19.03
1.00	10.000	11.295	10.70	359.68	132.70	1498.7966	6115.10	0.20	0.32	0.887	19.28
0.90	7.943	8.972	12.03	454.65	94.97	852.0591	6967.16	0.23	0.36	0.787	19.53
0.80	6.310	7.126	13.58	579.36	124.71	888.7102	7855.86	0.26	0.41	0.687	19.78
0.70	5.012	5.661	15.34	739.27	159.90	905.1699	8761.03	0.29	0.46	0.587	20.03
0.60	3.981	4.496	17.32	942.42	203.16	913.4766	9674.51	0.32	0.52	0.487	20.28
0.50	3.162	3.572	19.16	1153.29	210.87	753.1663	10427.67	0.35	0.57	0.387	20.53
0.40	2.512	2.837	22.04	1526.06	372.77	1057.5725	11485.24	0.38	0.66	0.287	20.78
0.30	1.995	2.254	25.49	2041.22	515.15	1160.9297	12646.17	0.42	0.76	0.187	21.03
0.20	1.585	1.790	29.63	2758.12	716.90	1283.3025	13929.47	0.46	0.89	0.087	21.28
0.10	1.259	1.422	33.91	3612.48	854.36	1214.8147	15144.29	0.51	1.02	-0.013	21.53
-0.00	1.000	1.129	38.76	4719.73	1107.25	1250.5940	16394.88	0.55	1.16	-0.113	21.78
-0.10	0.794	0.897	43.95	6068.30	1348.57	1209.8862	17634.76	0.59	1.32	-0.213	22.03
-0.20	0.631	0.713	49.07	7564.53	1496.22	1066.2671	18671.03	0.62	1.47	-0.313	22.28
-0.30	0.501	0.566	54.31	9266.36	1701.84	963.3574	19634.38	0.65	1.63	-0.413	22.53
-0.40	0.398	0.450	59.69	11193.16	1926.80	866.3765	20500.76	0.68	1.79	-0.513	22.78
-0.50	0.316	0.357	65.19	13350.93	2157.77	770.6816	21271.44	0.71	1.96	-0.613	23.03
-0.60	0.251	0.284	70.83	15761.02	2410.08	683.7573	21955.19	0.73	2.12	-0.713	23.28
-0.70	0.200	0.225	76.59	18428.67	2667.65	601.1729	22556.36	0.75	2.30	-0.813	23.53
-0.80	0.158	0.179	82.49	21377.27	2948.61	527.8210	23084.18	0.77	2.47	-0.913	23.78
-0.90	0.126	0.142	93.24	27312.05	5934.78	843.8682	23928.05	0.80	2.80	-1.013	24.03
-1.00	0.100	0.113	102.75	33167.56	5855.51	661.3560	24589.41	0.82	3.08	-1.113	24.28
-1.10	0.079	0.090	111.08	38763.37	5595.80	502.0344	25091.44	0.84	3.33	-1.213	24.53
-1.20	0.063	0.071	119.43	44810.17	6046.80	430.9207	25522.36	0.85	3.58	-1.313	24.78
-1.30	0.050	0.057	132.51	55162.91	10352.74	586.0403	26108.39	0.87	3.97	-1.413	25.03
-1.40	0.040	0.045	141.20	62635.31	7472.40	335.9951	26444.39	0.88	4.24	-1.513	25.28
-1.50	0.032	0.036	154.70	75184.81	12549.50	448.2290	26892.61	0.90	4.64	-1.613	25.53
-1.60	0.025	0.028	169.30	90045.81	14861.00	421.6211	27314.23	0.91	5.08	-1.713	25.78
-1.70	0.020	0.023	183.20	105438.87	15393.06	346.8962	27661.13	0.92	5.49	-1.813	26.03
-1.80	0.016	0.018	198.00	123163.00	17724.12	317.2778	27978.41	0.93	5.94	-1.913	26.28
-1.90	0.013	0.014	213.00	142530.87	19367.87	275.3960	28253.80	0.94	6.39	-2.013	26.53
-2.00	0.010	0.011	222.40	155388.69	12857.81	145.2257	28399.02	0.95	6.67	-2.113	26.78
-∞							29979.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4527

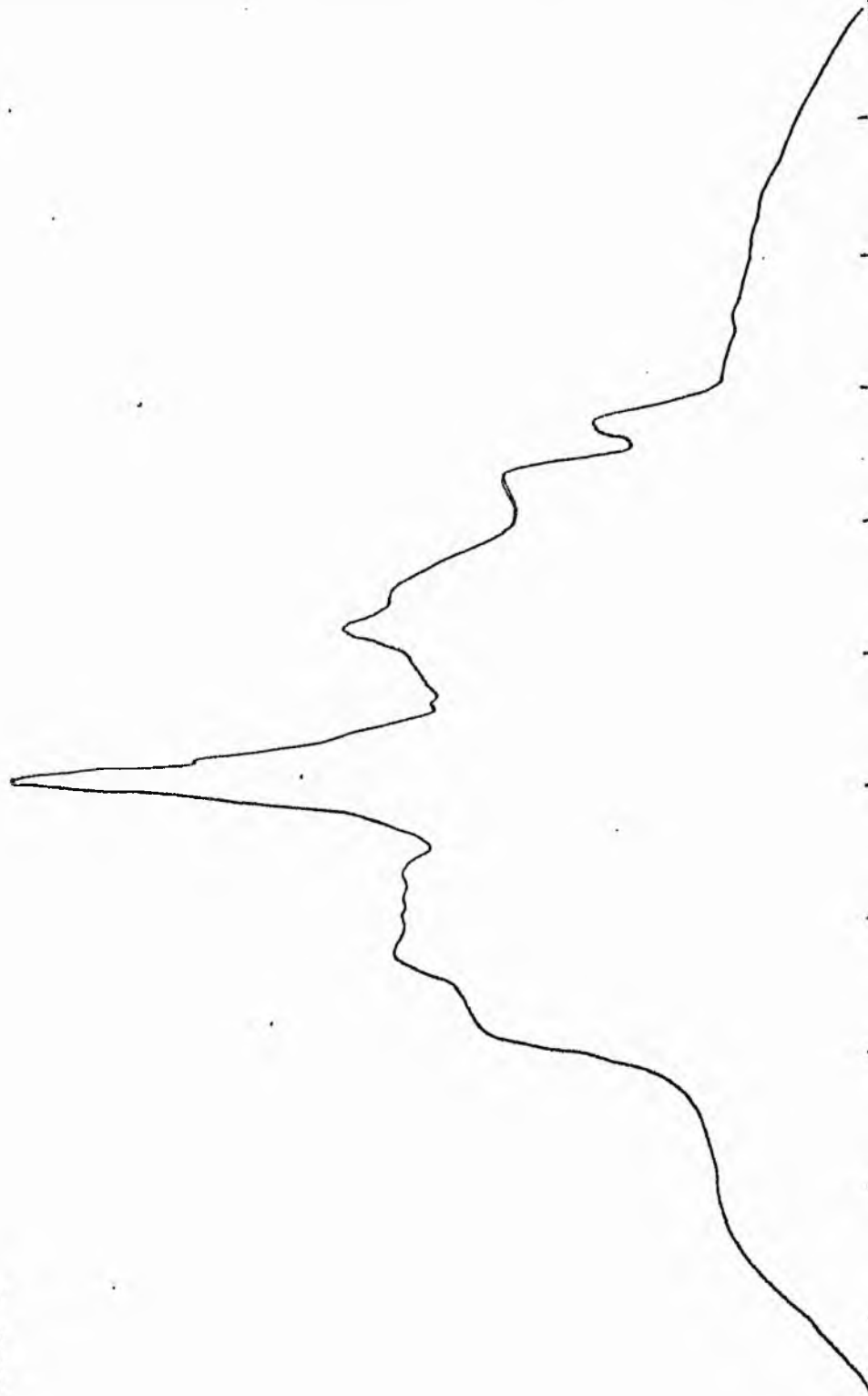
V-FILTER

Total luminosity	L_T	= 8.33
Total apparent magnitude	m_T	= 10.59
Apparent central surface brightness	μ_o	= 17.65
Major axis at threshold	$2a_m$	= 8.24
Minor axis at threshold	$2b_m$	= 5.52
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 6.28
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.87
Gradient of exponential component	$G(a)$	= -0.68
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.63
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.63
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.21
Axis ratio	b/a	= 0.59
Equivalent radius	r_1^*	= 0.22
Surface brightness	μ_1	= 19.70
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.85
Axis ratio	b/a	= 0.33
Equivalent radius	r_e^*	= 0.56
Surface brightness	μ_e	= 21.48
Mean surface brightness	μ_e'	= 11.34
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.70
Axis ratio	b/a	= 0.24
Equivalent radius	r_3^*	= 1.26
Surface brightness	μ_3	= 23.53
Concentration indices	$\left\{ \begin{array}{l} C_{21} \\ C_{32} \end{array} \right.$	= 2.58
		= 2.27

NGC 4535
B-Filter
Axis 1

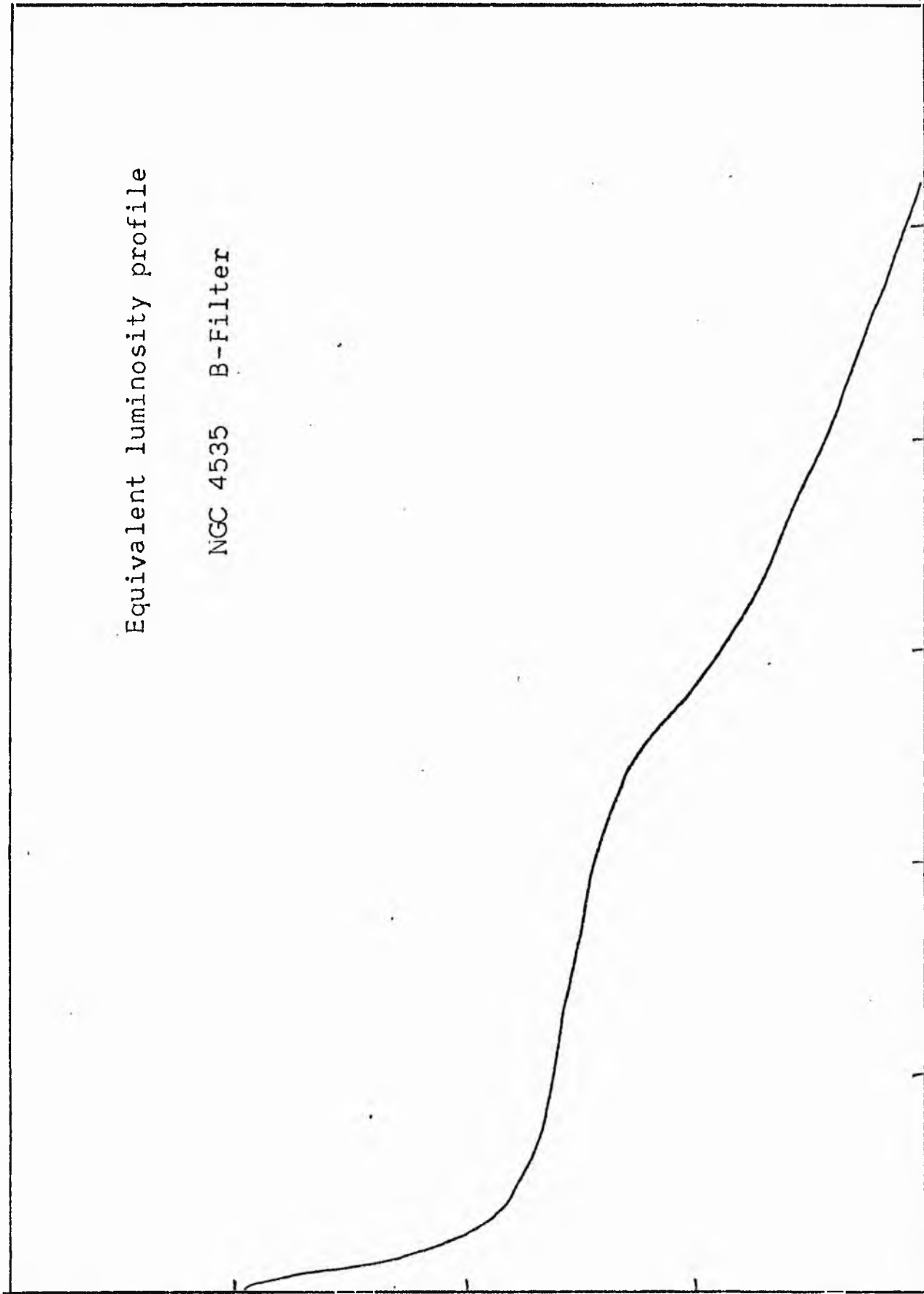


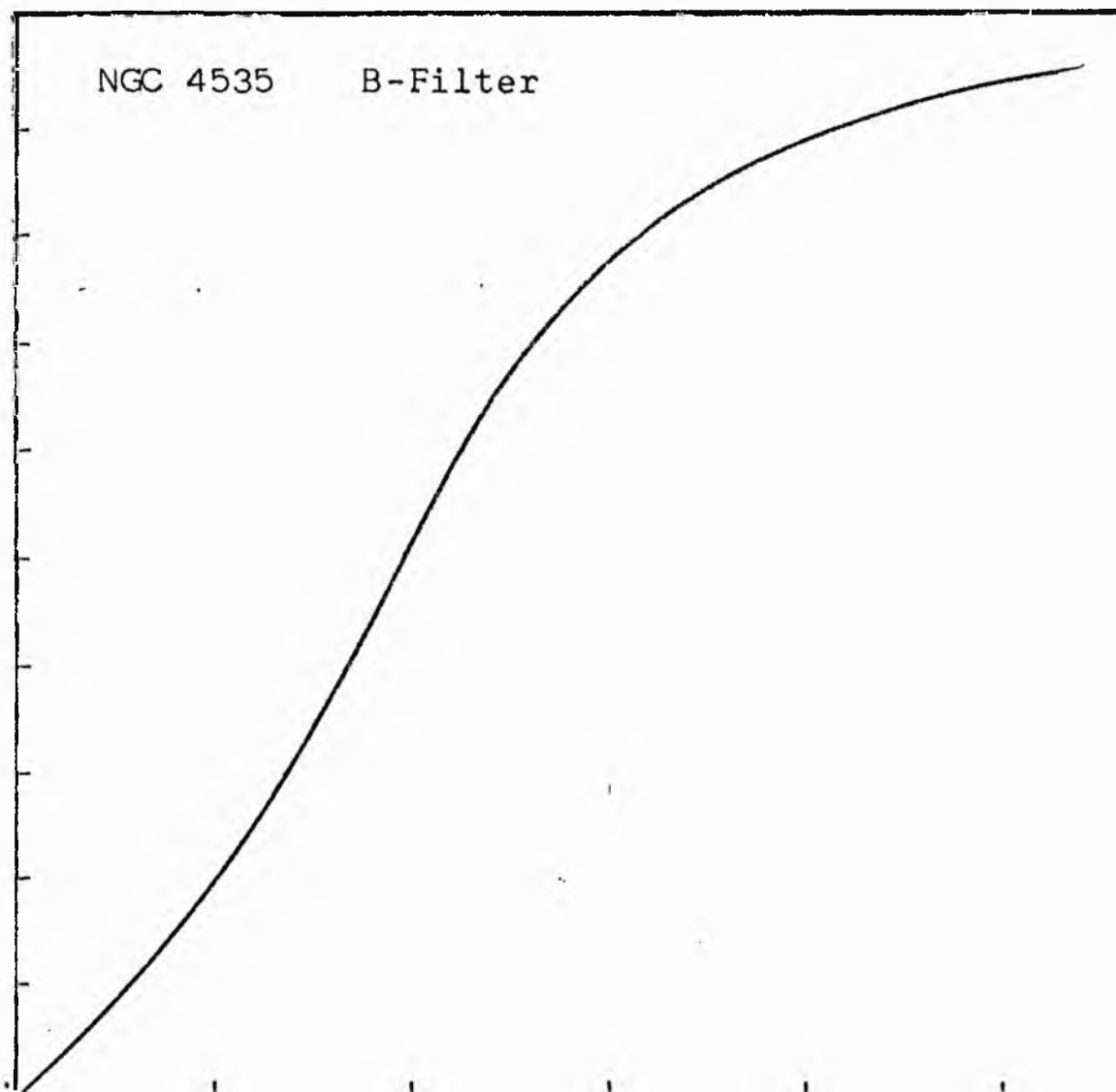
NGC 4535
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4535 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

D COLOUR

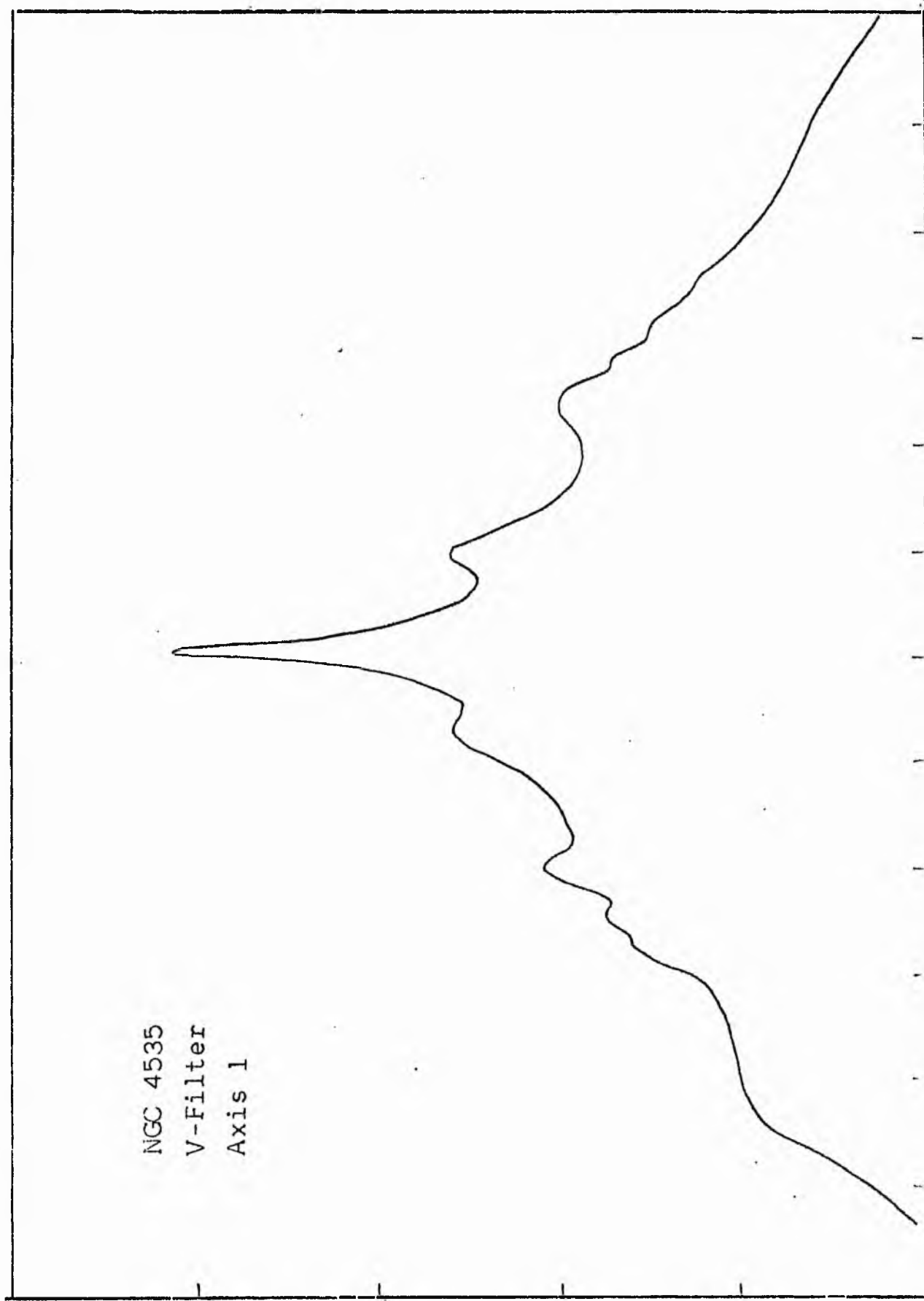
LOG I	I	I	R	AREA	ΔA	P	ΣP	KIRI	P	LOG J	μ
0.92	8.318	8.130	0.0	0.0	10.41	84.6071	0.0	0.0	0.0	1.465	19.16
0.90	7.943	7.126	1.02	10.41	10.18	72.5650	84.61	0.00	0.02	1.445	19.21
0.88	6.310	5.661	2.56	20.59	7.12	40.3209	157.17	0.01	0.03	1.345	19.46
0.70	5.012	4.496	2.97	27.71	17.89	80.4509	197.49	0.01	0.03	1.245	19.71
0.60	3.981	3.572	3.81	45.60	23.20	62.0797	277.94	0.01	0.04	1.145	19.96
0.50	3.162	2.837	4.68	60.01	16.14	45.7912	360.02	0.01	0.05	1.045	20.21
0.40	2.512	2.254	5.20	84.95	65.93	148.5692	406.61	0.02	0.05	0.945	20.46
0.30	1.995	1.790	6.93	150.87	95.74	171.3805	555.18	0.02	0.07	0.845	20.71
0.20	1.585	1.422	8.86	246.61	99.75	141.8311	726.56	0.03	0.09	0.745	20.96
0.10	1.259	1.129	10.50	346.36	201.03	227.0566	868.39	0.04	0.11	0.645	21.21
0.00	1.000	0.897	13.20	547.39	131.40	117.9553	1095.45	0.04	0.14	0.545	21.46
-0.10	0.794	0.713	14.70	678.87	1100.65	784.3730	1213.41	0.05	0.16	0.445	21.71
-0.20	0.631	0.566	23.80	1779.52	2661.93	1506.8462	1947.78	0.08	0.25	0.345	21.96
-0.30	0.501	0.450	37.60	4441.45	6199.07	2787.7571	3504.63	0.14	0.40	0.245	22.21
-0.40	0.398	0.357	58.20	10641.32	12485.94	4459.5742	6242.38	0.26	0.62	0.145	22.46
-0.50	0.316	0.284	85.80	23127.27	10610.04	3010.3860	10751.96	0.44	0.91	0.045	22.71
-0.60	0.251	0.225	103.63	33738.11	9120.24	2055.3174	13762.34	0.56	1.10	-0.055	22.96
-0.70	0.200	0.179	116.80	42858.35	6661.94	1192.5410	15817.66	0.65	1.23	-0.155	23.21
-0.80	0.158	0.142	125.55	49520.29	7202.09	1024.0730	17010.20	0.70	1.33	-0.255	23.46
-0.90	0.126	0.113	134.37	56722.38	3210.21	362.5815	18034.27	0.74	1.42	-0.355	23.71
-1.00	0.100	0.090	138.12	59932.59	7889.41	707.8118	18346.84	0.75	1.46	-0.455	23.96
-1.10	0.079	0.071	146.93	67822.00	9763.12	695.7646	19104.65	0.78	1.55	-0.555	24.21
-1.20	0.063	0.057	157.15	77585.12	13004.06	736.1279	19800.41	0.81	1.66	-0.655	24.46
-1.30	0.050	0.045	169.81	90589.19	12264.25	551.4622	20536.54	0.84	1.80	-0.755	24.71
-1.40	0.040	0.036	180.94	102853.44	12127.31	433.1516	21088.00	0.86	1.91	-0.855	24.96
-1.50	0.032	0.028	191.31	114980.75	18374.69	521.3098	21521.15	0.88	2.02	-0.955	25.21
-1.60	0.025	0.023	206.03	133355.44	24488.25	551.8667	22042.46	0.90	2.18	-1.055	25.46
-1.70	0.020	0.018	224.15	157843.69	19242.37	344.4575	22544.32	0.92	2.37	-1.155	25.71
-1.80	0.016	0.014	237.42	177086.06	14809.87	211.7234	22938.70	0.94	2.51	-1.255	25.96
-1.90	0.013	0.011	247.20	191975.94	18765.25	211.9496	23150.50	0.95	2.61	-1.355	26.21
-2.00	0.010		259.00	210741.19			23362.45	0.96	2.74	-1.455	26.46
-∞							24442.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4535

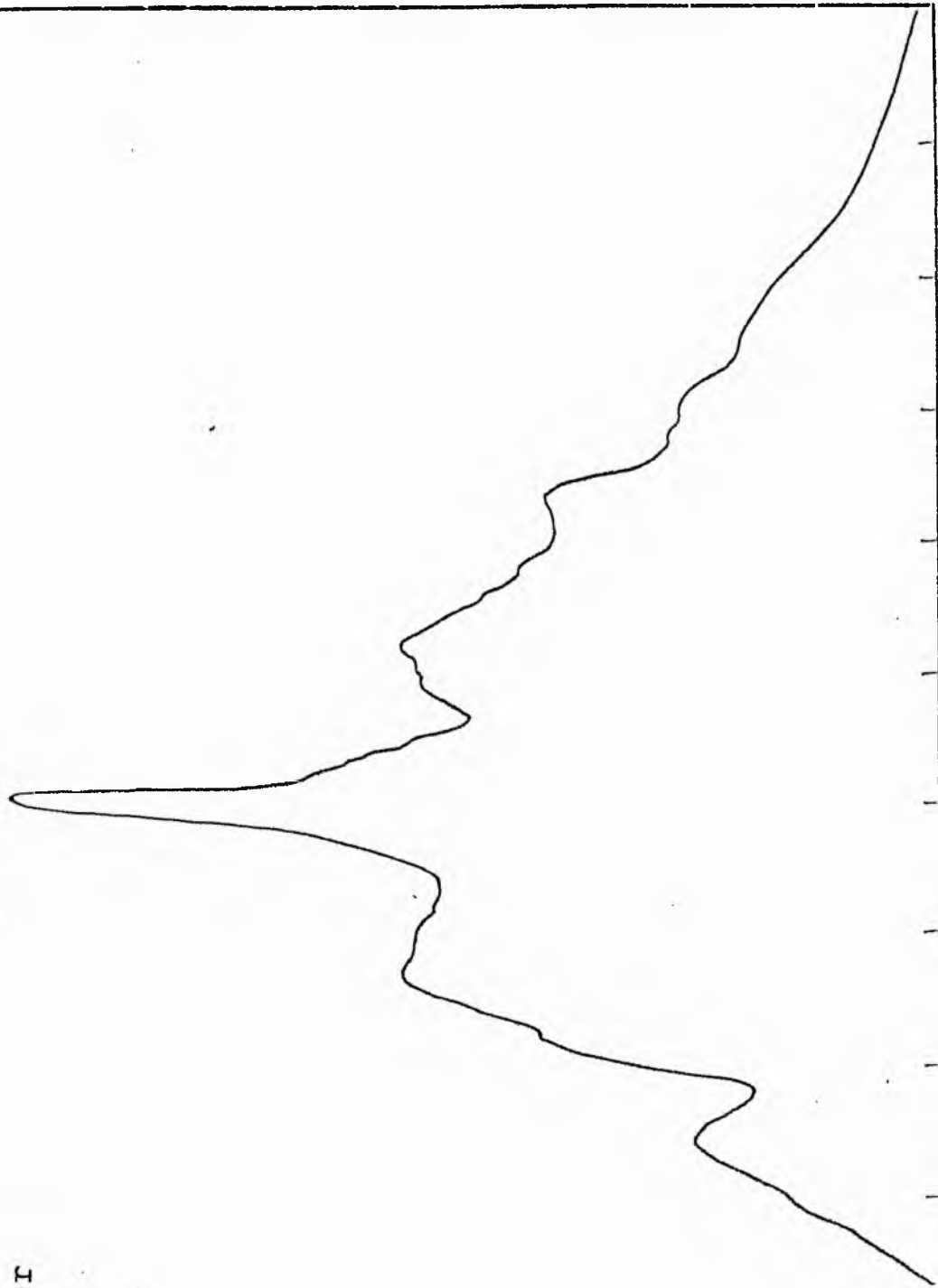
B-FILTER

Total luminosity	L_T	= 6.79
Total apparent magnitude	m_T	= 10.49
Apparent central surface brightness	μ_o	= 19.16
Major axis at threshold	$2a_m$	= 9.37
Minor axis at threshold	$2b_m$	= 8.85
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.95
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.87
Gradient of exponential component	$G(a)$	= -0.41
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.47
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.79
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.96
Axis ratio	b/a	= 0.74
Equivalent radius	r_1^*	= 0.94
Surface brightness	μ_1	= 22.44
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.61
Axis ratio	b/a	= 0.78
Equivalent radius	r_e^*	= 1.58
Surface brightness	μ_e	= 22.83
Mean surface brightness	μ_e'	= 13.48
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.68
Axis ratio	b/a	= 0.73
Equivalent radius	r_3^*	= 2.29
Surface brightness	μ_3	= 23.96
Concentration indices	$\left\{ \begin{matrix} C_{21} \\ C_{32} \end{matrix} \right.$	= 1.66
		= 1.45

NGC 4535
V-Filter
Axis 1

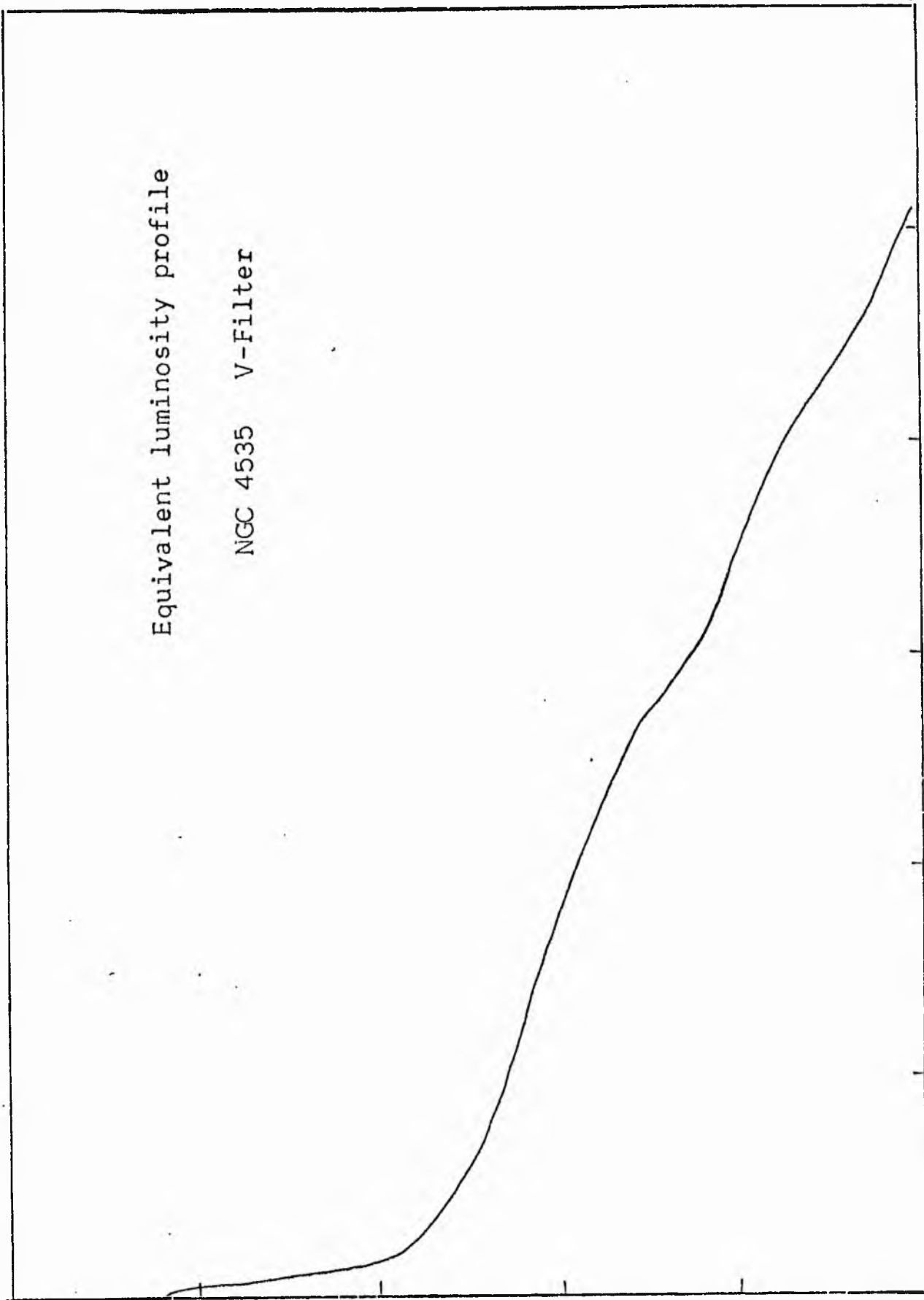


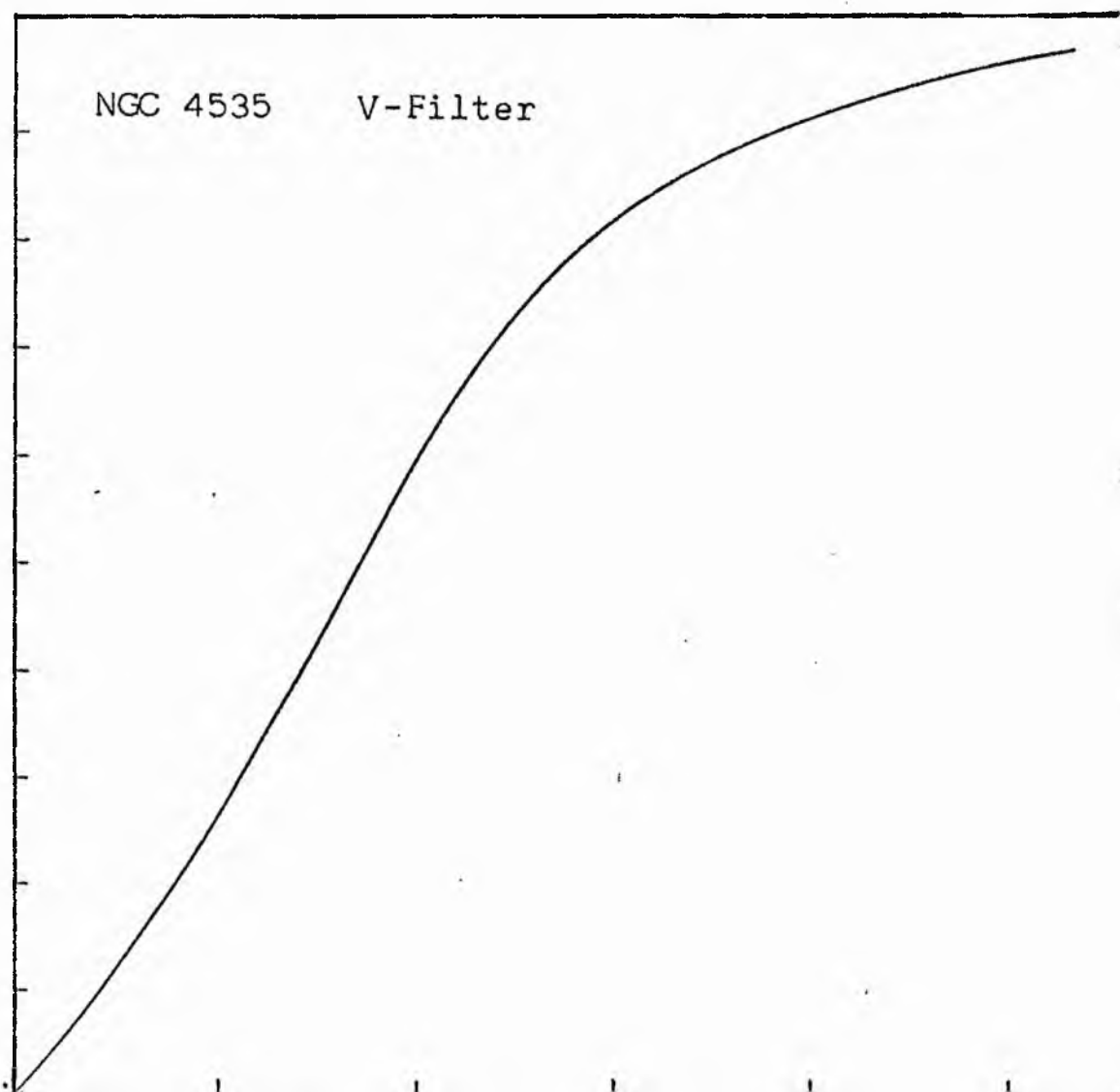
NGC 4535
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4535 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4335
V COLOUR

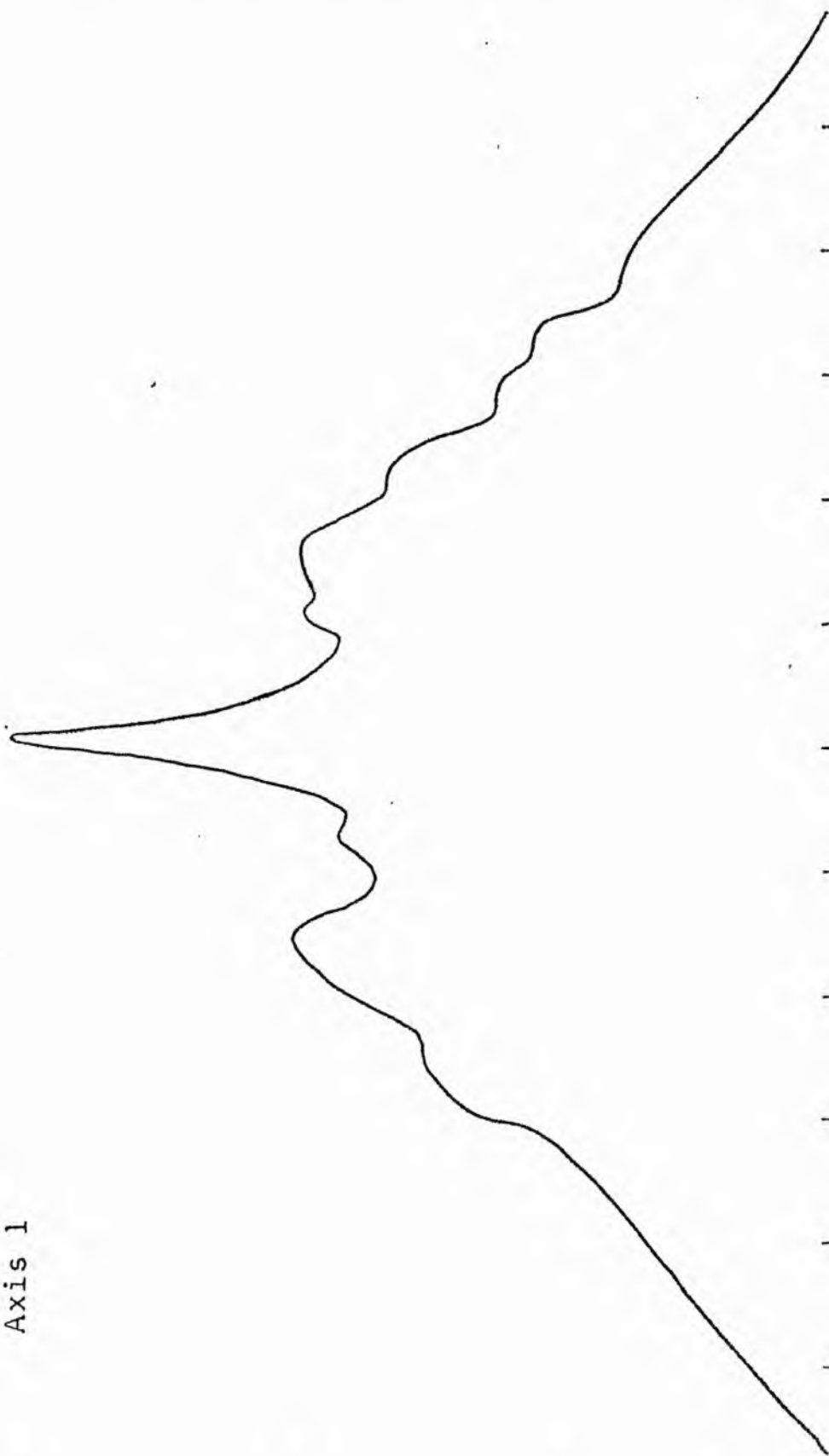
LOG I	I	T	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
1.23	16.982		0.0	0.0	10.41	170.8248	0.0	0.0	0.0	1.617	17.98
1.20	15.849	16.416	1.82	10.41	10.02	142.5033	170.82	0.01	0.02	1.587	18.06
1.10	12.589	14.219	2.55	20.43	15.89	179.4555	313.33	0.01	0.03	1.487	18.31
1.00	10.000	11.295	3.40	36.32	16.49	147.9719	492.78	0.02	0.04	1.387	18.56
0.90	7.943	8.972	4.10	52.81	13.67	97.3892	640.75	0.02	0.05	1.287	18.81
0.80	6.310	7.126	4.60	66.48	15.24	86.2506	738.14	0.02	0.05	1.187	19.06
0.70	5.012	5.661	5.10	81.71	55.13	247.9125	824.39	0.03	0.06	1.087	19.31
0.60	3.981	4.496	6.60	136.85	49.42	176.5022	1072.31	0.03	0.08	0.987	19.56
0.50	3.162	3.572	7.70	186.27	35.41	100.4488	1248.81	0.04	0.09	0.887	19.81
0.40	2.512	2.837	8.40	221.67	61.86	139.4011	1349.26	0.04	0.10	0.787	20.06
0.30	1.995	2.254	9.50	283.53	358.99	642.4502	1488.66	0.05	0.11	0.687	20.31
0.20	1.585	1.790	14.30	642.42	716.75	1019.1565	2131.11	0.07	0.17	0.587	20.56
0.10	1.259	1.422	20.80	1359.18	1246.58	1407.9678	3150.27	0.10	0.25	0.487	20.81
-0.00	1.000	1.129	28.80	2605.76	2050.86	1839.9607	4558.23	0.14	0.34	0.387	21.06
-0.10	0.794	0.897	38.50	4656.62	3643.33	2596.3921	6398.19	0.20	0.46	0.287	21.31
-0.20	0.631	0.713	51.40	8249.96	7940.84	4495.0820	8994.58	0.28	0.61	0.187	21.56
-0.30	0.501	0.566	71.90	16240.79	7048.48	3169.3306	13489.66	0.41	0.85	0.087	21.81
-0.40	0.398	0.450	86.10	23289.28	7067.58	2524.3093	16658.99	0.51	1.02	-0.013	22.06
-0.50	0.316	0.357	98.30	30356.86	9262.67	2627.8960	19183.30	0.59	1.16	-0.113	22.31
-0.60	0.251	0.284	112.30	39619.52	9075.95	2045.3342	21811.19	0.67	1.33	-0.213	22.56
-0.70	0.200	0.225	124.50	48695.48	8221.25	1471.6687	23856.52	0.73	1.47	-0.313	22.81
-0.80	0.158	0.179	134.60	56916.73	3433.13	488.1604	25328.19	0.78	1.59	-0.413	23.06
-0.90	0.126	0.142	138.60	60349.86	4613.36	521.0615	25816.35	0.79	1.64	-0.513	23.31
-1.00	0.100	0.113	143.80	64963.22	11392.53	1022.0979	26337.41	0.81	1.70	-0.613	23.56
-1.10	0.079	0.090	155.90	76355.75	16153.31	1151.1567	27359.51	0.84	1.85	-0.713	23.81
-1.20	0.063	0.071	171.60	92509.06	15710.50	889.3303	28510.66	0.87	2.03	-0.813	24.06
-1.30	0.050	0.057	185.60	108219.56	14682.31	660.1887	29399.99	0.90	2.20	-0.913	24.31
-1.40	0.040	0.045	197.79	122901.87	13120.50	468.6240	30060.18	0.92	2.34	-1.013	24.56
-1.50	0.032	0.036	208.08	136022.37	3552.94	100.8005	30528.80	0.94	2.47	-1.113	24.81
-1.60	0.025	0.028	210.78	139575.31	12560.69	283.0669	30629.60	0.94	2.50	-1.213	25.06
-1.70	0.020	0.023	220.06	152136.00	16505.44	295.4631	30912.66	0.95	2.61	-1.313	25.31
-1.80	0.016	0.018	231.69	168641.44	20965.44	298.1128	31208.12	0.96	2.74	-1.413	25.56
-1.90	0.013	0.014	245.67	189606.87	13076.12	147.6918	31506.23	0.97	2.91	-1.513	25.81
-2.00	0.010	0.011	254.00	202683.00			31653.93	0.97	3.01	-1.613	26.06
- ∞							32593.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NGC 4535

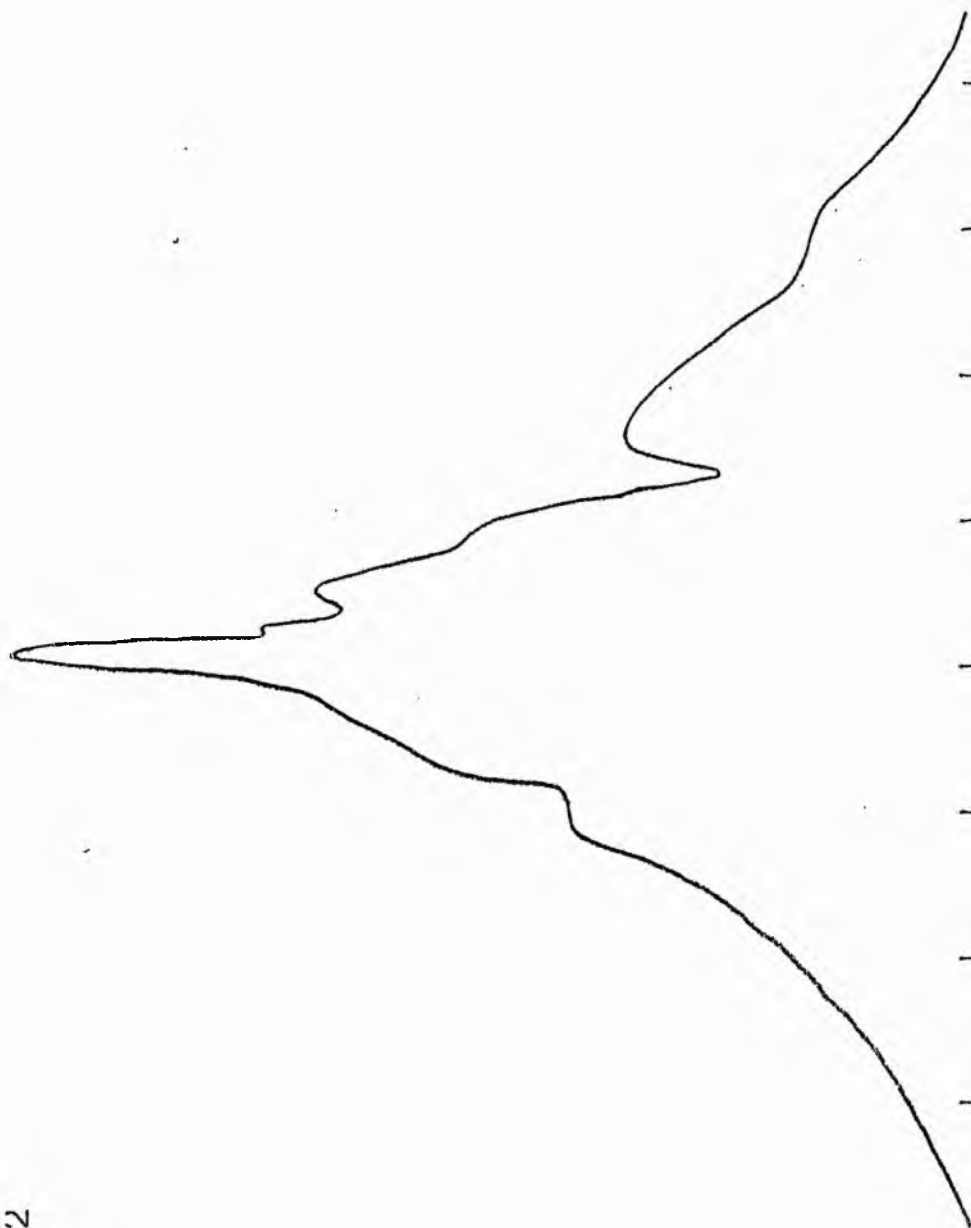
V-FILTER

Total luminosity	L_T	=	9.05
Total apparent magnitude	m_T	=	9.78
Apparent central surface brightness	μ_0	=	17.98
Major axis at threshold	$2a_m$	=	10.07
Minor axis at threshold	$2b_m$	=	7.97
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	=	8.57
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	=	0.94
Gradient of exponential component	$G(a)$	=	-0.45
Equivalent gradient of exponential comp....	$G(r^*)$	=	-0.50
Equivalent gradient of reduced exp. comp....	$G(\rho)$	=	-0.75
Parameters at $k = \frac{1}{4}$:			
Semi-major axis	a_1	=	0.85
Axis ratio	b/a	=	0.92
Equivalent radius	r_1^*	=	0.78
Surface brightness	μ_1	=	21.46
Parameters at $k = \frac{1}{2}(\text{effective})$:			
Semi-major axis	a_e	=	1.36
Axis ratio	b/a	=	0.93
Equivalent radius	r_e^*	=	1.41
Surface brightness	μ_e	=	22.03
Mean surface brightness	μ_e'	=	12.52
Parameters at $k = \frac{3}{4}$:			
Semi-major axis	a_3	=	2.23
Axis ratio	b/a	=	0.85
Equivalent radius	r_3^*	=	2.14
Surface brightness	μ_3	=	22.91
Concentration indices	$\begin{Bmatrix} C_{21} \\ C_{32} \end{Bmatrix}$	=	1.79
		=	1.52

NGC 4536
B-Filter
Axis 1

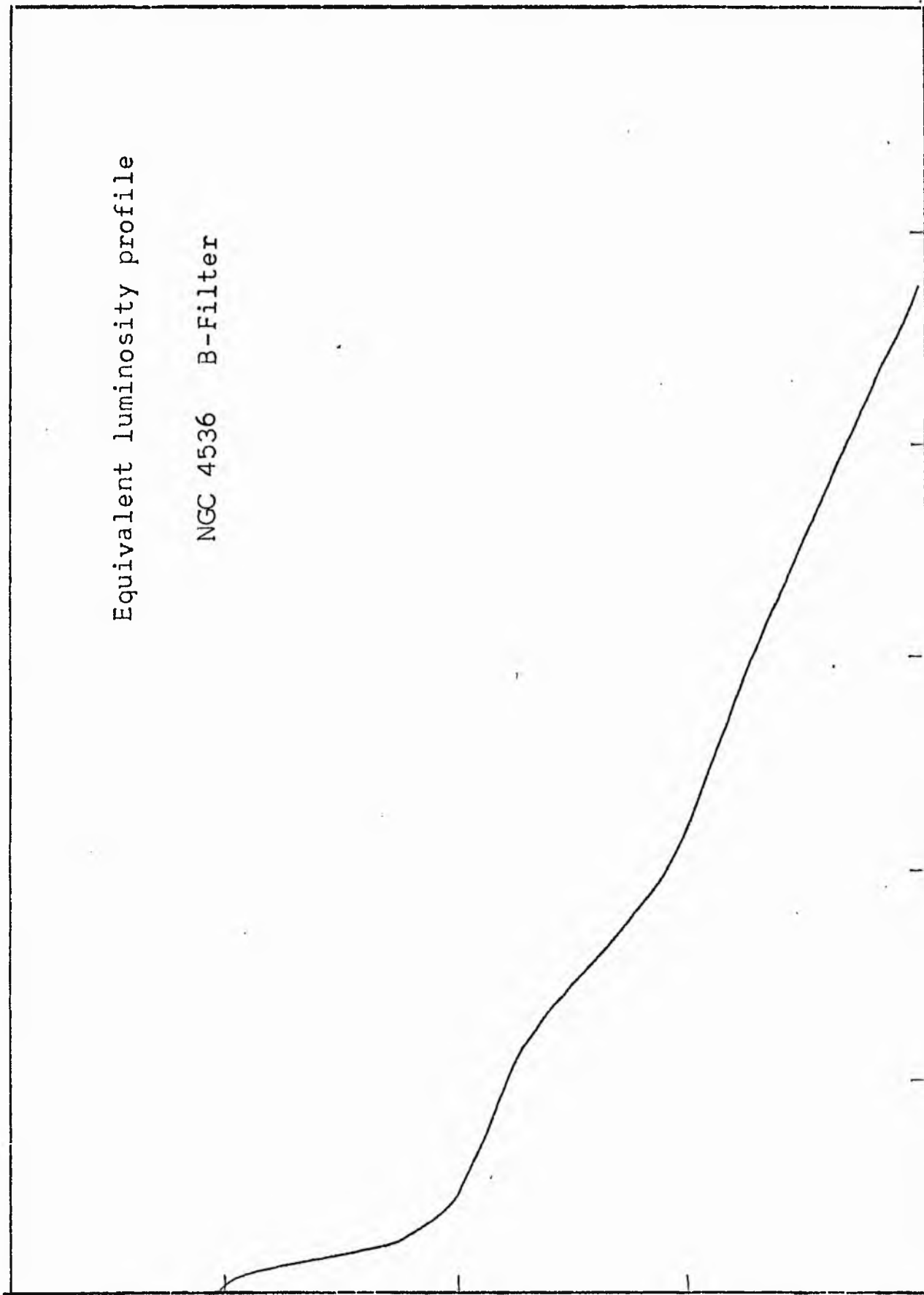


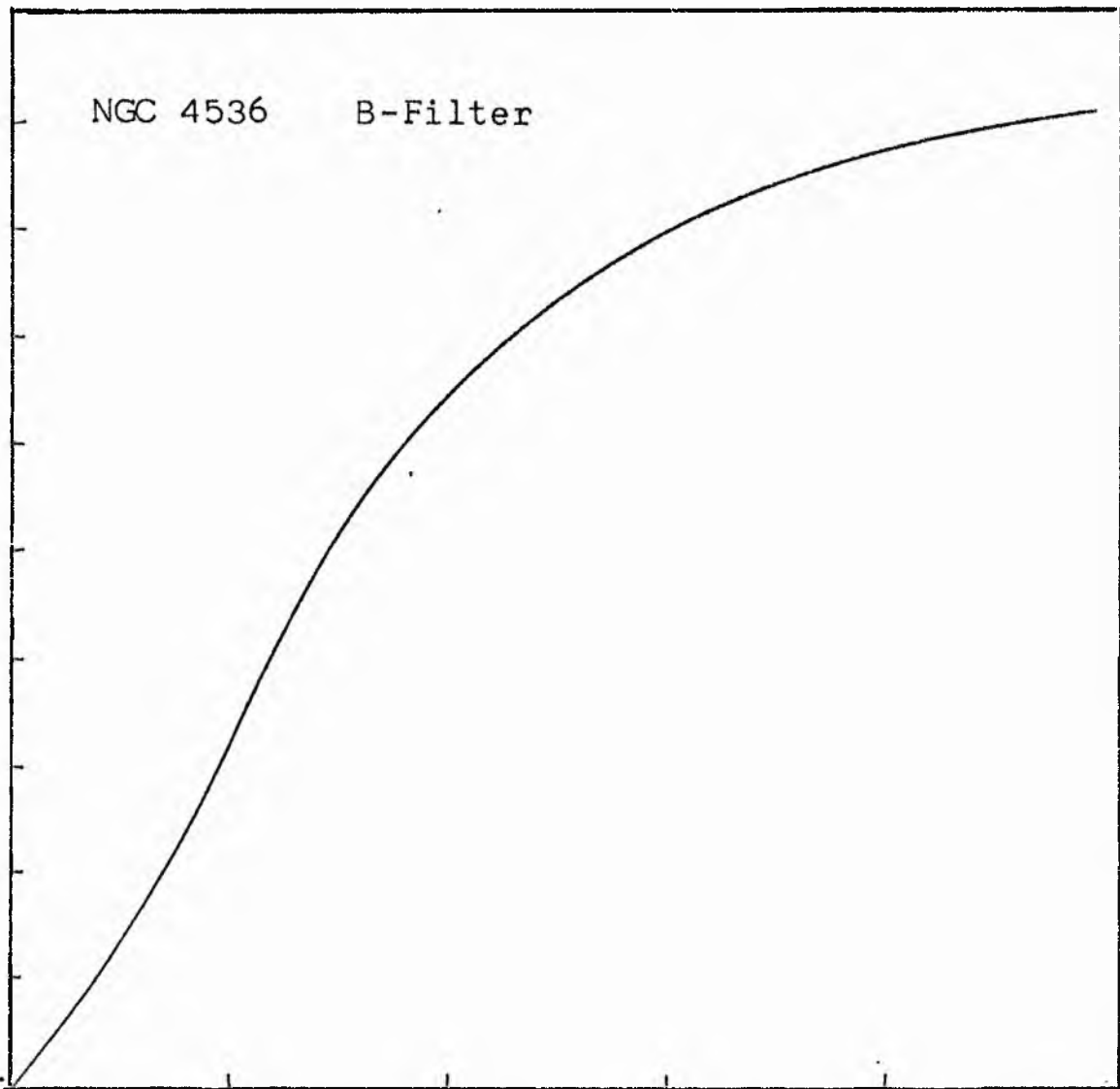
NGC 4536
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4536 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

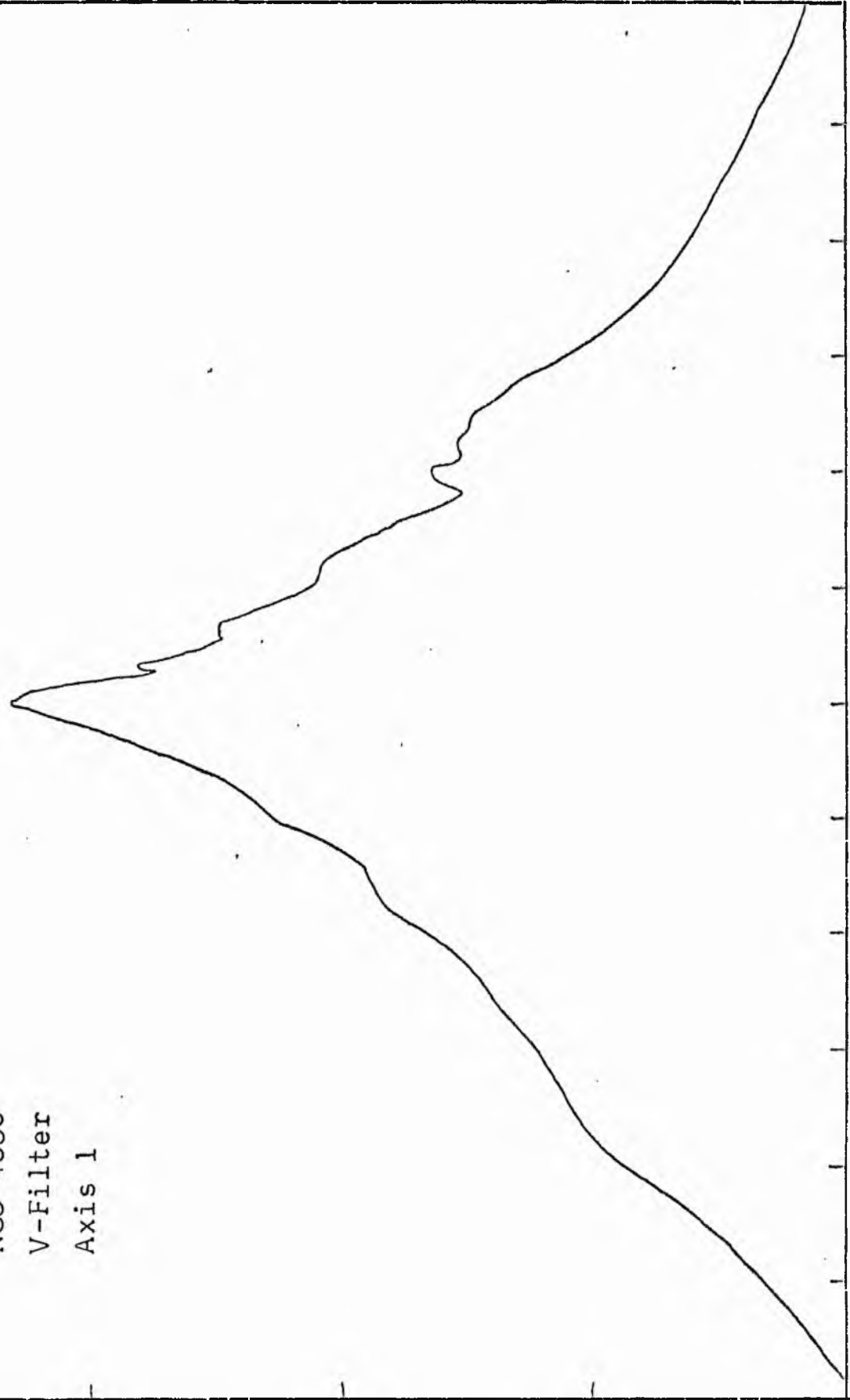
LOG I	I	T	R	AREA	ΔA	P	ΣP	KIR1	ρ	LOG J	μ
0.98	9.550	8.747	0.0	0.0	70.29	614.7676	0.0	0.0	0.0	1.402	19.48
0.90	7.943	7.126	4.73	70.29	2.10	14.9350	614.77	0.03	0.07	1.322	19.68
0.80	6.310	5.661	4.80	72.38	38.98	209.3140	629.70	0.03	0.07	1.222	19.93
0.70	5.012	4.496	5.90	109.36	67.36	302.8633	839.02	0.03	0.09	1.122	20.18
0.60	3.981	3.572	7.50	176.71	19.35	69.1196	1141.88	0.05	0.11	1.022	20.43
0.50	3.162	2.837	7.90	196.07	10.05	28.5214	1211.00	0.05	0.12	0.922	20.68
0.40	2.512	2.254	8.10	206.12	120.73	272.0767	1239.52	0.05	0.12	0.822	20.93
0.30	1.995	1.790	10.20	324.85	195.94	350.7498	1511.60	0.06	0.15	0.722	21.18
0.20	1.585	1.422	12.90	522.79	395.64	562.8486	1862.35	0.08	0.19	0.622	21.43
0.10	1.259	1.129	17.10	918.63	1108.20	1251.6672	2425.20	0.10	0.25	0.522	21.68
0.00	1.000	0.897	25.40	2026.83	3228.46	2896.4573	3676.88	0.15	0.37	0.422	21.93
-0.10	0.794	0.713	40.90	5255.29	4110.30	2929.1775	6573.32	0.27	0.60	0.322	22.18
-0.20	0.631	0.566	54.60	9365.59	1790.11	1013.3293	9502.50	0.39	0.80	0.222	22.43
-0.30	0.501	0.450	59.59	11155.69	2795.77	1257.1091	10515.82	0.44	0.87	0.122	22.68
-0.40	0.398	0.357	66.64	13951.46	4385.88	1566.4963	11772.93	0.49	0.97	0.022	22.93
-0.50	0.316	0.284	76.40	18337.34	2941.56	834.5464	13339.43	0.55	1.12	-0.078	23.18
-0.60	0.251	0.225	82.30	21278.90	2828.92	637.5186	14173.97	0.59	1.20	-0.178	23.43
-0.70	0.200	0.179	87.60	24107.82	2540.45	454.7600	14811.48	0.61	1.28	-0.278	23.68
-0.80	0.158	0.142	92.10	26648.27	8252.18	1173.3865	15266.24	0.63	1.35	-0.378	23.93
-0.90	0.126	0.113	105.40	34900.45	10037.39	1133.6882	16439.62	0.68	1.54	-0.478	24.18
-1.00	0.100	0.090	119.60	44937.84	6856.21	615.1160	17573.31	0.73	1.75	-0.578	24.43
-1.10	0.079	0.071	128.40	51794.04	9517.55	678.2642	18188.43	0.75	1.88	-0.678	24.68
-1.20	0.063	0.057	139.70	61311.60	11558.53	654.2998	18866.69	0.78	2.04	-0.778	24.93
-1.30	0.050	0.045	152.30	72870.12	13074.87	587.9121	19520.98	0.81	2.23	-0.878	25.18
-1.40	0.040	0.036	165.40	85945.00	11933.75	426.2380	20108.89	0.83	2.42	-0.978	25.43
-1.50	0.032	0.028	176.51	97878.75	14781.87	419.3777	20535.13	0.85	2.58	-1.078	25.68
-1.60	0.025	0.023	189.37	112660.62	10751.31	242.2914	20954.50	0.87	2.77	-1.178	25.93
-1.70	0.020	0.018	198.20	123411.94	19212.62	343.9250	21196.79	0.88	2.90	-1.278	26.18
-1.80	0.016	0.014	213.07	142624.56	18403.87	261.6899	21540.71	0.89	3.12	-1.378	26.43
-1.90	0.013	0.011	226.40	161028.44	27699.62	312.8613	21802.40	0.90	3.31	-1.478	26.68
-2.00	0.010		245.10	188728.06			22115.26	0.92	3.58	-1.578	26.93
-∞							24135.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4536

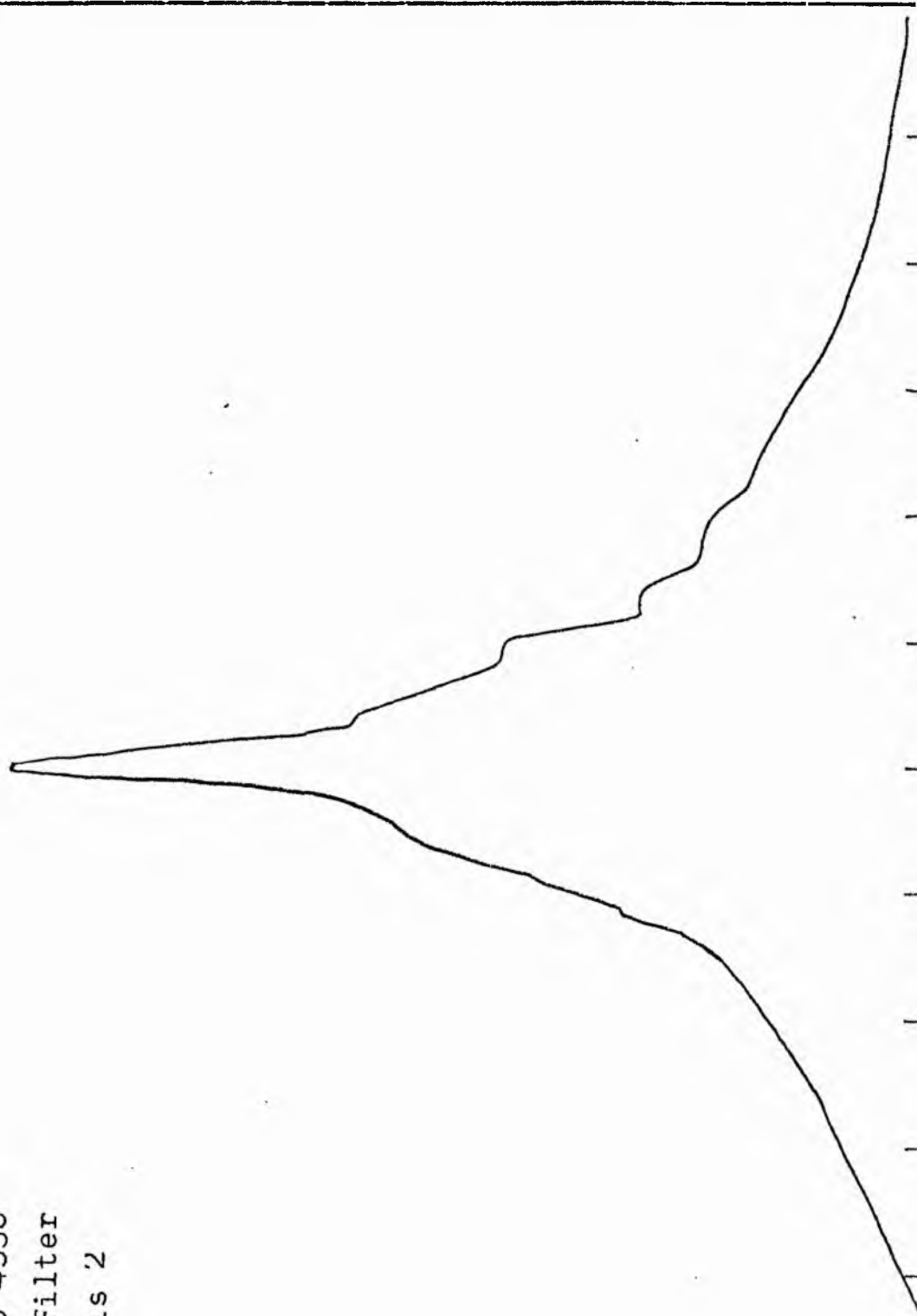
B-FILTER

Total luminosity	L_T	= 6.70
Total apparent magnitude	m_T	= 10.96
Apparent central surface brightness	μ_0	= 19.48
Major axis at threshold	$2a_m$	= 9.75
Minor axis at threshold	$2b_m$	= 7.06
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.64
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.79
Gradient of exponential component	$G(a)$	= -0.44
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.57
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.67
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.90
Axis ratio	b/a	= 0.41
Equivalent radius	r_1^*	= 0.64
Surface brightness	μ_1	= 22.14
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.93
Axis ratio	b/a	= 0.31
Equivalent radius	r_e^*	= 1.14
Surface brightness	μ_e	= 22.98
Mean surface brightness	μ_e'	= 13.24
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.92
Axis ratio	b/a	= 0.41
Equivalent radius	r_3^*	= 2.12
Surface brightness	μ_3	= 24.68
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.79 \\ = 1.86 \end{matrix}$

NGC 4536
V-Filter
Axis 1

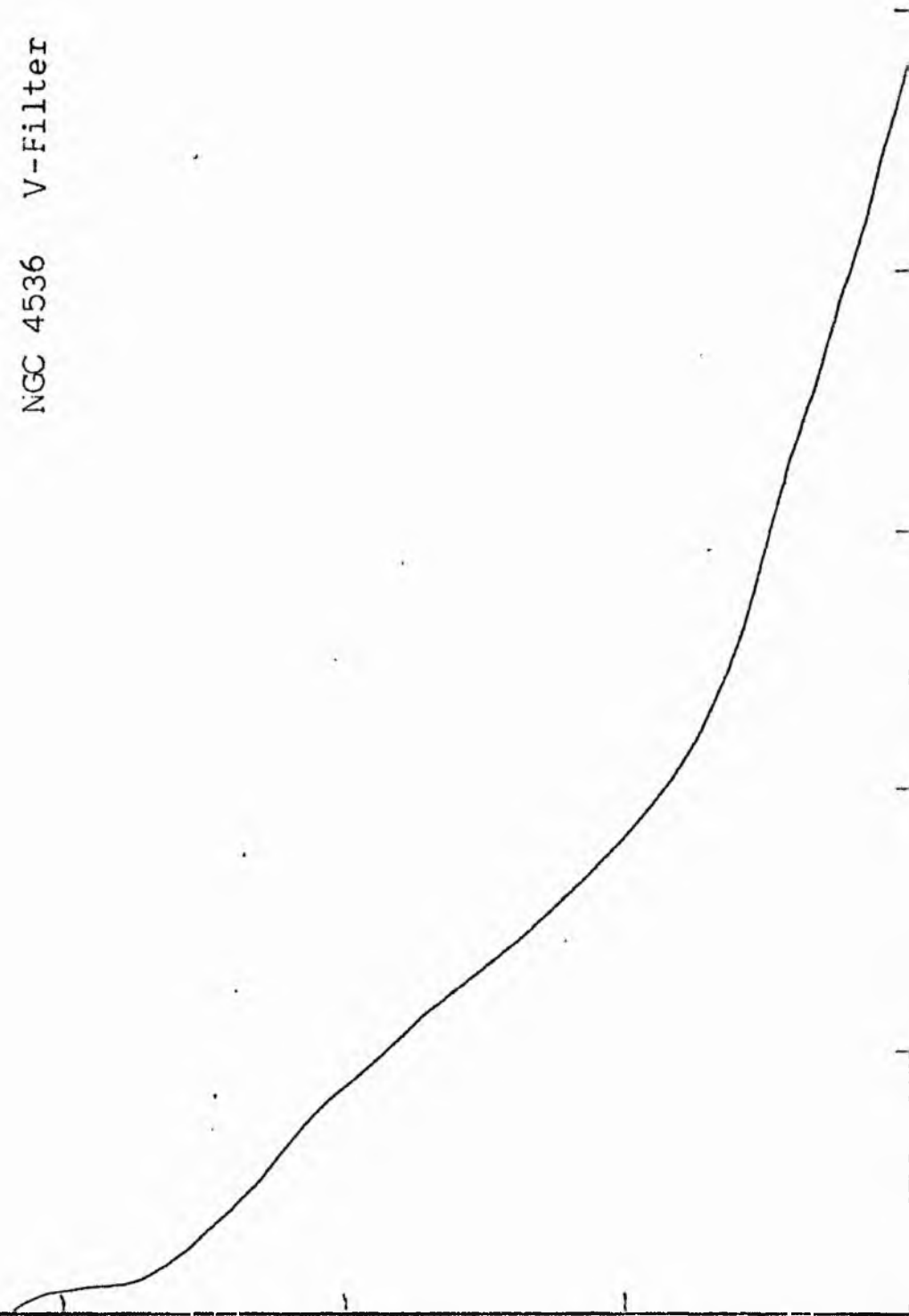


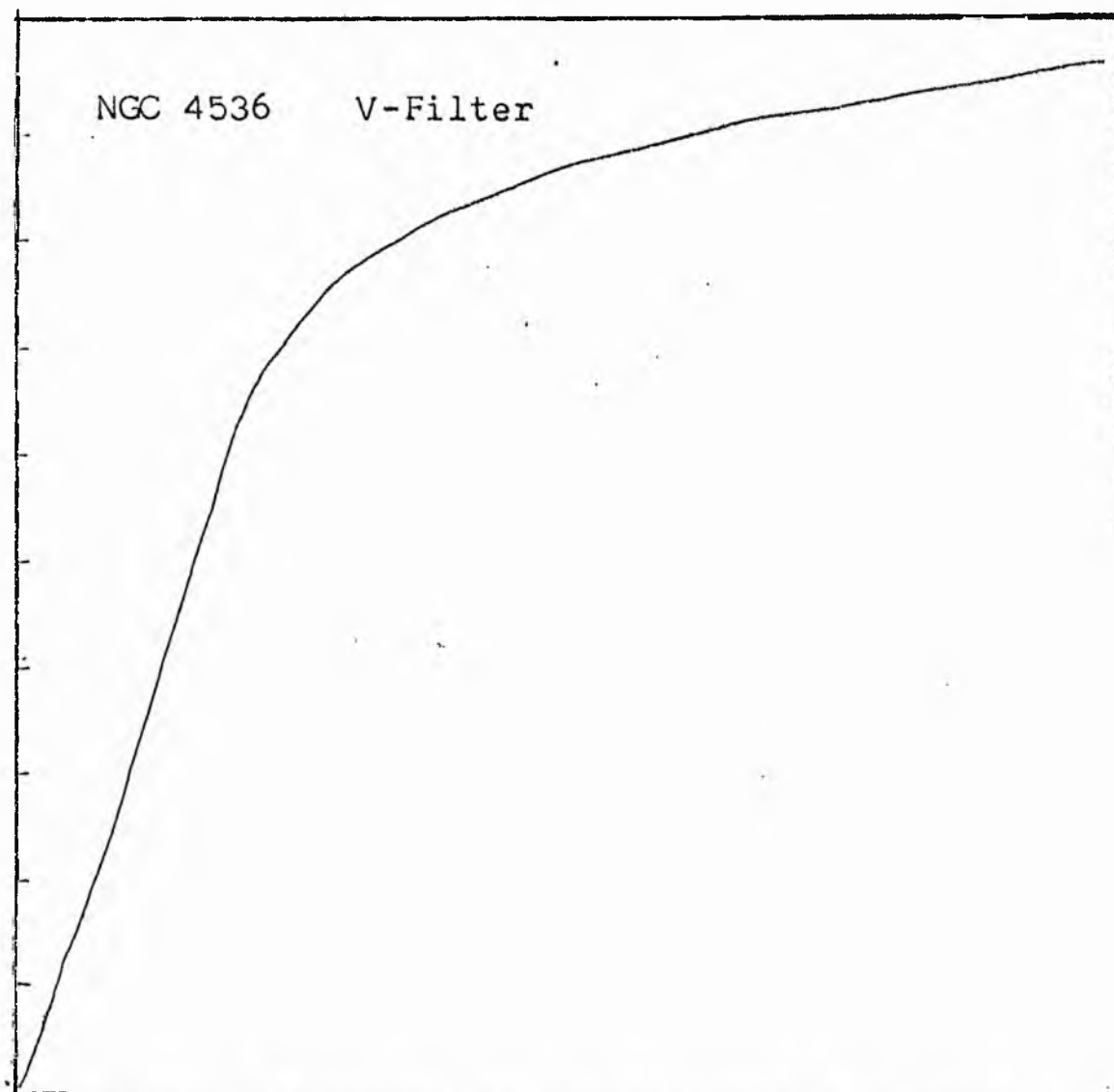
NGC 4536
V-Filter
Axis '2



Equivalent luminosity profile

NGC 4536 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4536
V COLOUR

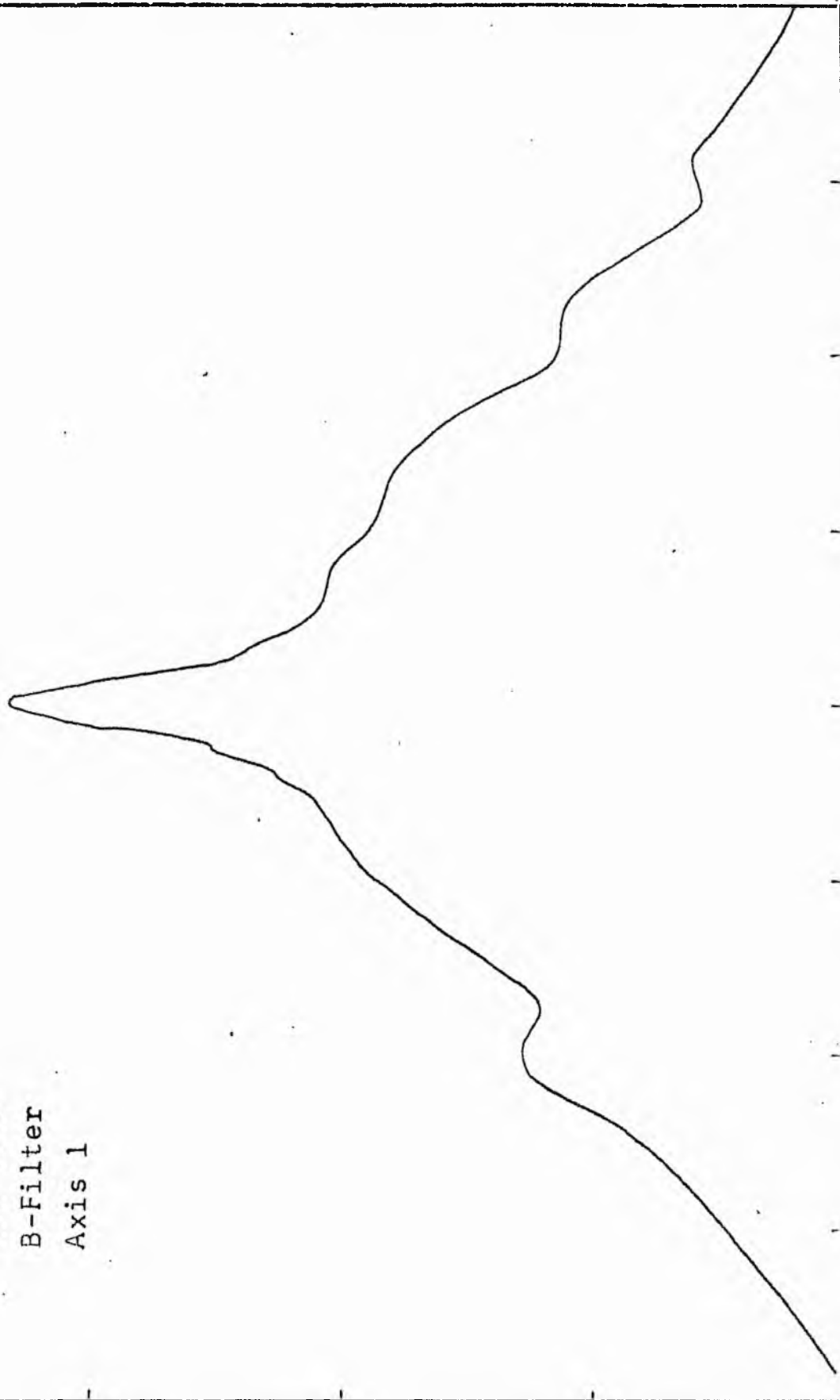
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.26	18.197		0.0	0.0			0.0	0.0	0.0	1.059	18.60
1.20	15.849	17.023	4.60	66.48	66.48	1131.6184	1131.62	0.03	0.11	0.999	18.75
1.10	12.589	14.219	6.90	149.57	83.10	1181.5354	2313.15	0.07	0.16	0.899	19.00
1.00	10.000	11.295	7.70	186.27	36.69	414.4421	2727.60	0.08	0.18	0.799	19.25
0.90	7.943	8.972	8.10	206.12	19.85	178.1302	2905.73	0.08	0.19	0.699	19.50
0.80	6.310	7.126	8.90	248.85	42.73	304.4810	3210.21	0.09	0.21	0.599	19.75
0.70	5.012	5.661	9.90	307.91	59.06	334.3311	3544.54	0.10	0.23	0.499	20.00
0.60	3.981	4.496	14.60	669.66	361.75	1626.6169	5171.15	0.15	0.34	0.399	20.25
0.50	3.162	3.572	21.10	1398.67	729.00	2603.7644	7774.91	0.22	0.50	0.299	20.50
0.40	2.512	2.837	28.90	2623.89	1225.22	3416.0488	11250.96	0.32	0.68	0.199	20.75
0.30	1.995	2.254	36.70	4231.38	1607.49	3622.5467	14873.55	0.43	0.87	0.099	21.00
0.20	1.585	1.790	42.40	5647.82	1416.45	2535.5449	17409.10	0.50	1.00	-0.001	21.25
0.10	1.259	1.422	46.60	6822.15	1174.33	1669.7869	19078.88	0.55	1.10	-0.101	21.50
-0.00	1.000	1.129	50.30	7948.51	1126.36	1212.1748	20351.05	0.59	1.19	-0.201	21.75
-0.10	0.794	0.897	55.70	9746.76	1798.25	1613.3240	21964.37	0.63	1.32	-0.301	22.00
-0.20	0.631	0.713	59.30	11047.37	1300.62	926.8745	22891.25	0.66	1.40	-0.401	22.25
-0.30	0.501	0.566	62.70	12350.51	1303.13	737.6667	23628.91	0.68	1.48	-0.501	22.50
-0.40	0.398	0.450	66.80	14018.54	1668.03	750.0237	24378.93	0.70	1.58	-0.601	22.75
-0.50	0.316	0.357	72.30	16422.00	2403.46	858.4395	25237.37	0.73	1.71	-0.701	23.00
-0.60	0.251	0.284	78.50	19359.28	2937.28	833.3311	26070.70	0.75	1.85	-0.801	23.25
-0.70	0.200	0.225	83.60	21956.46	2597.18	585.2930	26655.99	0.77	1.97	-0.901	23.50
-0.80	0.158	0.179	88.20	24439.20	2482.74	444.4302	27100.42	0.78	2.08	-1.001	23.75
-0.90	0.126	0.142	94.30	27936.57	3497.37	497.2939	27597.71	0.79	2.23	-1.101	24.00
-1.00	0.100	0.113	99.60	31165.09	3228.52	364.6497	27962.36	0.80	2.35	-1.201	24.25
-1.10	0.079	0.090	111.30	38917.06	7751.97	695.4792	28657.84	0.82	2.63	-1.301	24.50
-1.20	0.063	0.071	120.70	45768.26	6851.20	488.2468	29146.09	0.84	2.85	-1.401	24.75
-1.30	0.050	0.057	136.40	58449.20	12680.94	717.8350	29863.92	0.86	3.22	-1.501	25.00
-1.40	0.040	0.045	151.20	71821.31	13372.11	601.2756	30465.19	0.88	3.57	-1.601	25.25
-1.50	0.032	0.036	167.40	88036.06	16214.75	579.1414	31044.33	0.89	3.95	-1.701	25.50
-1.60	0.025	0.028	186.30	109037.37	21001.31	545.8289	31640.16	0.91	4.40	-1.801	25.75
-1.70	0.020	0.023	202.40	128697.69	19660.31	443.0637	32083.22	0.92	4.78	-1.901	26.00
-1.80	0.016	0.018	226.40	161028.44	32330.75	518.7515	32601.97	0.94	5.35	-2.001	26.25
-1.90	0.013	0.014	251.30	198396.87	37368.44	531.3511	33193.32	0.95	5.93	-2.101	26.50
-2.00	0.010	0.011	274.80	237237.00	38840.12	438.6899	33632.01	0.97	6.49	-2.201	26.75
-∞							34782.00	11)			∞

PHOTOMETRIC PARAMETERS OF NGC 4536

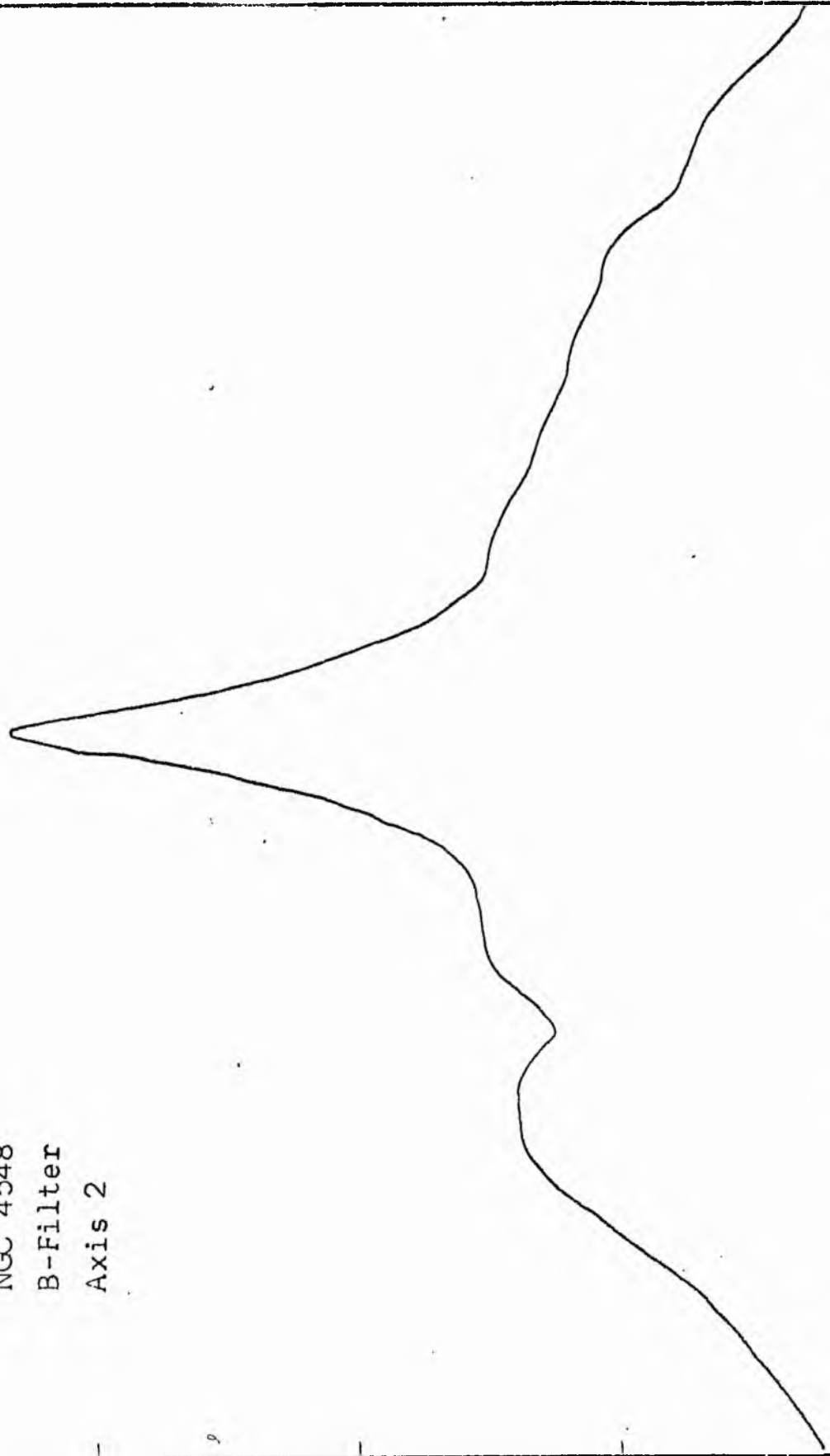
V-FILTER

Total luminosity	L_T	= 9.66
Total apparent magnitude	m_T	= 10.40
Apparent central surface brightness	μ_0	= 18.60
Major axis at threshold	$2a_m$	= 10.37
Minor axis at threshold	$2b_m$	= 8.18
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 6.34
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.86
Gradient of exponential component	$G(a)$	= -0.57
Equivalent gradient of exponential comp.	$G(r^*)$	= -0.52
Equivalent gradient of reduced exp. comp.	$G(\rho)$	= -0.80
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.43
Axis ratio	b/a	= 0.21
Equivalent radius	r_1^*	= 0.38
Surface brightness	μ_1	= 20.57
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.81
Axis ratio	b/a	= 0.19
Equivalent radius	r_e^*	= 0.71
Surface brightness	μ_e	= 21.25
Mean surface brightness	μ_e'	= 11.65
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.05
Axis ratio	b/a	= 0.39
Equivalent radius	r_3^*	= 1.31
Surface brightness	μ_3	= 23.25
Concentration indices	$\{C_{21}$	= 1.82
	C_{32}	= 1.86

NGC 4548
B-Filter
Axis 1

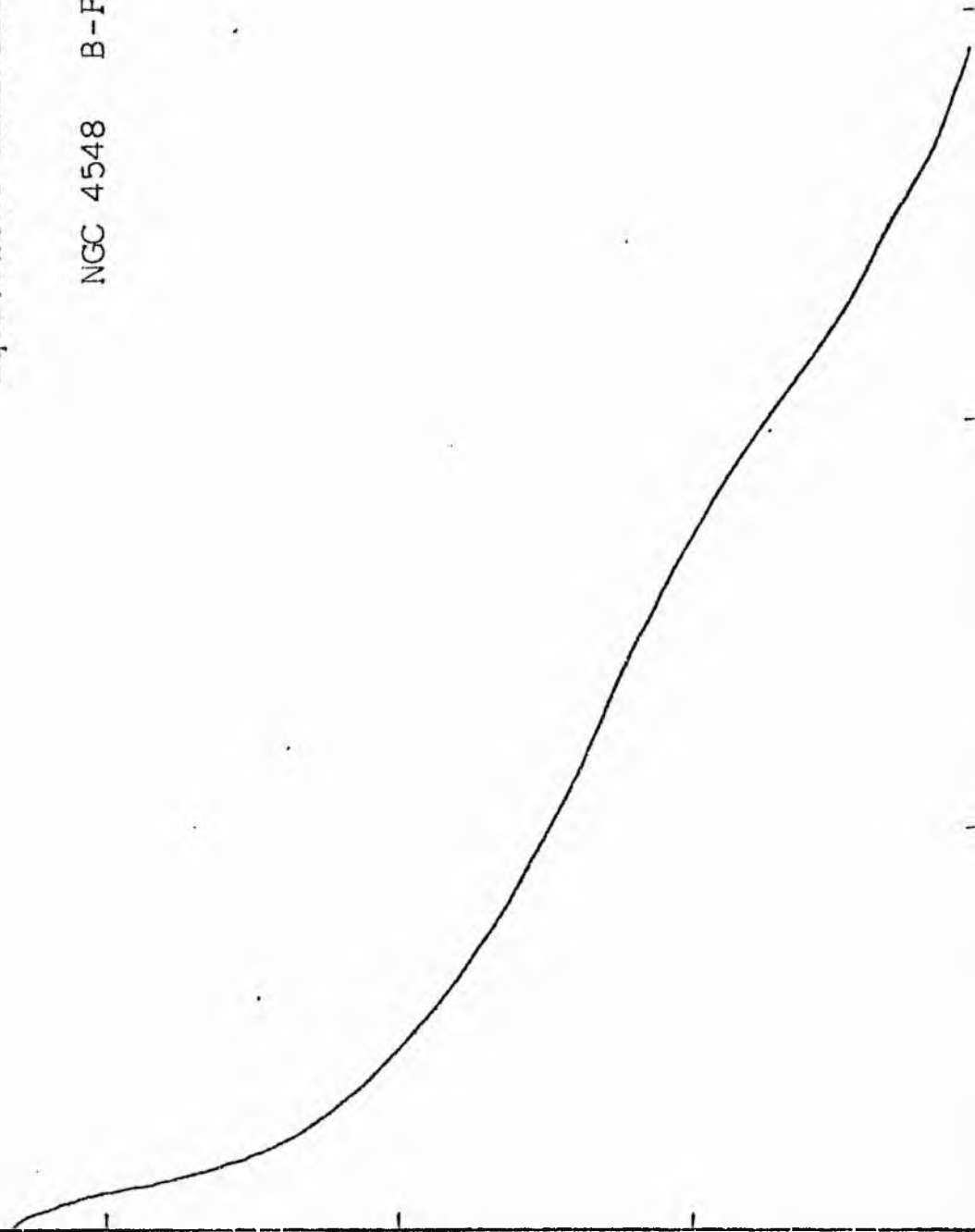


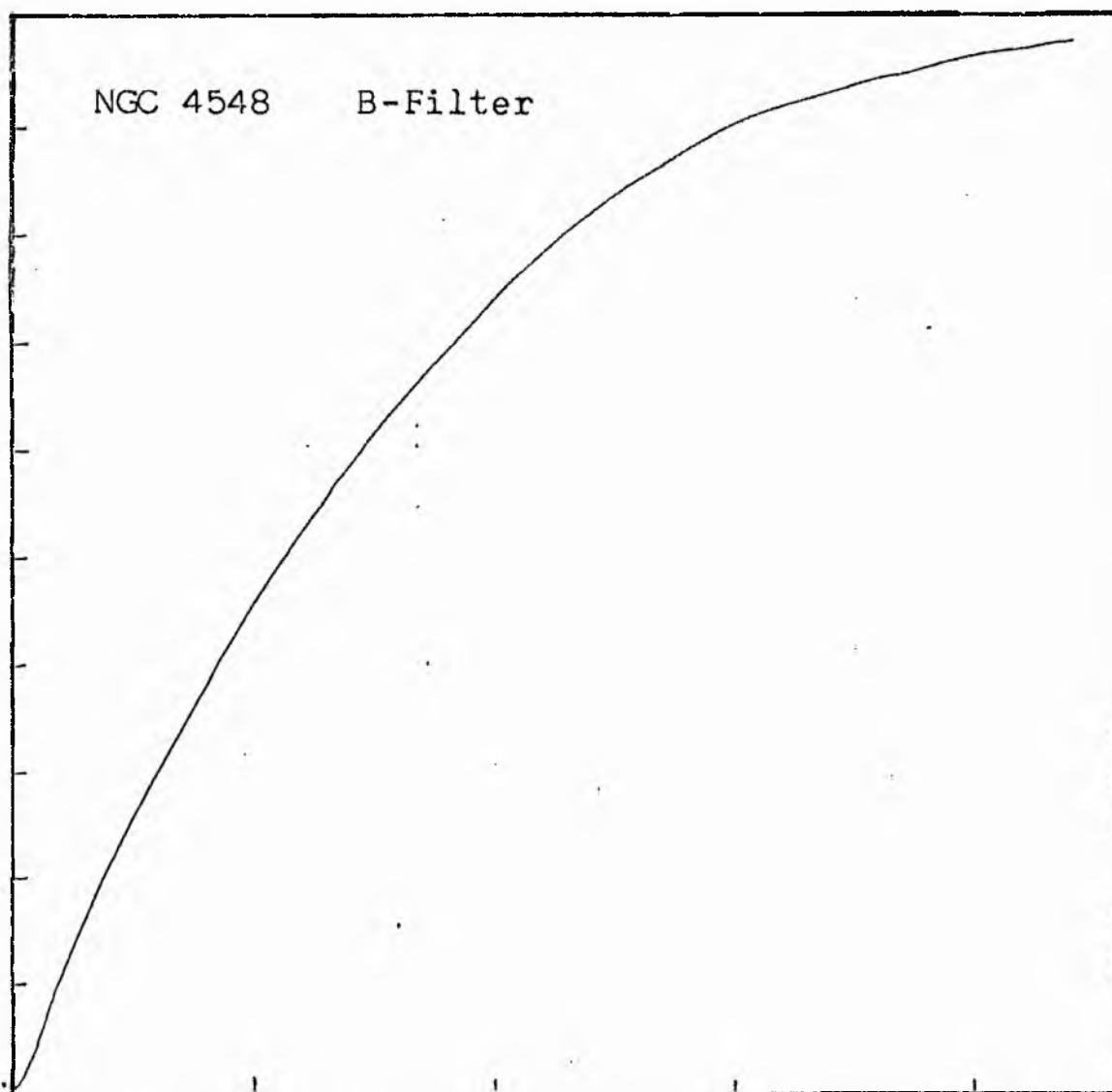
NGC 4548
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4548 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4548
BY COLOUR

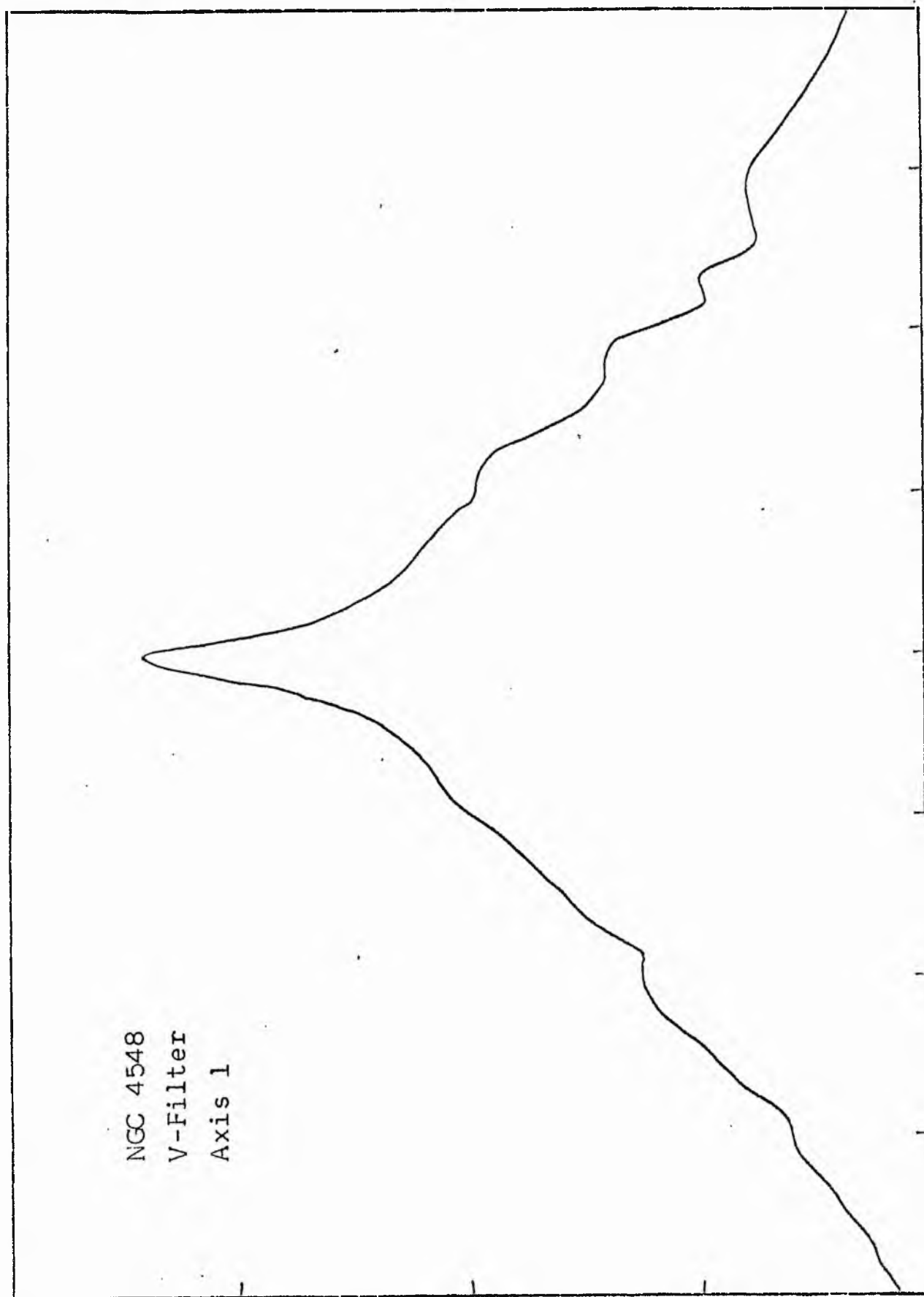
LOG I	I	T	R	AREA	Δ A	P	Σ P	KIR1	Q	LOG J	μ
1.29	19.498	17.674	0.0	0.0	38.48	680.1626	0.0	0.0	0.0	1.637	18.72
1.20	15.849	14.219	3.50	38.48	30.91	439.5376	680.16	0.03	0.06	1.547	18.95
1.10	12.509	11.295	4.70	69.40	55.29	624.5010	1119.72	0.04	0.09	1.447	19.20
1.00	10.000	8.972	6.30	124.69	33.68	302.1453	1744.22	0.07	0.11	1.347	19.45
0.90	7.943	7.126	7.10	158.37	95.98	398.9055	2046.37	0.08	0.13	1.247	19.70
0.80	6.310	5.661	8.26	214.34	85.53	484.1648	2445.27	0.09	0.15	1.147	19.95
0.70	5.012	4.496	9.77	299.87	92.80	417.2764	2929.44	0.11	0.18	1.047	20.20
0.60	3.981	3.572	11.18	392.88	153.06	548.6738	3346.71	0.13	0.20	0.947	20.45
0.50	3.162	2.837	13.18	545.73	189.68	538.1421	3891.34	0.15	0.24	0.847	20.70
0.40	2.512	2.254	15.30	735.42	289.26	651.8616	4431.55	0.17	0.28	0.747	20.95
0.30	1.995	1.790	18.06	1024.67	511.10	914.9067	5081.39	0.20	0.33	0.647	21.20
0.20	1.585	1.422	22.11	1535.77	1079.04	1534.3022	5998.29	0.23	0.40	0.547	21.45
0.10	1.259	1.129	28.85	2614.82	1674.41	1891.1772	7532.59	0.29	0.52	0.447	21.70
-0.00	1.060	0.897	36.95	4289.22	1387.95	1245.2136	9423.77	0.36	0.67	0.347	21.95
-0.10	0.794	0.713	42.51	5677.17	1136.20	809.7061	10668.98	0.41	0.77	0.247	22.20
-0.20	0.631	0.566	46.57	6813.37	1658.64	938.9077	11478.68	0.44	0.85	0.147	22.45
-0.30	0.501	0.450	51.93	8472.01	3811.62	1713.8843	12417.59	0.48	0.94	0.047	22.70
-0.40	0.398	0.357	62.53	12283.63	4680.33	1664.5178	14131.47	0.55	1.14	-0.053	22.95
-0.50	0.316	0.284	73.44	16943.96	5572.82	1581.0554	15795.99	0.61	1.33	-0.153	23.20
-0.60	0.251	0.225	84.66	22516.78	6714.26	1513.1084	17377.04	0.67	1.54	-0.253	23.45
-0.70	0.200	0.179	96.46	29231.04	6851.43	1226.4602	18890.15	0.73	1.75	-0.353	23.70
-0.80	0.158	0.142	107.17	36082.47	7276.37	1034.6343	20116.61	0.78	1.95	-0.453	23.95
-0.90	0.126	0.113	117.48	43358.84	7152.43	807.8401	21151.24	0.82	2.13	-0.553	24.20
-1.00	0.100	0.090	126.80	50511.27	6999.00	627.9260	21959.08	0.85	2.30	-0.653	24.45
-1.10	0.079	0.071	135.30	57510.27	5507.10	392.4604	22587.00	0.87	2.46	-0.753	24.70
-1.20	0.063	0.057	141.63	63017.37	7423.57	420.2285	22979.46	0.89	2.57	-0.853	24.95
-1.30	0.050	0.045	149.74	70440.94	8184.44	368.0125	23399.69	0.90	2.72	-0.953	25.20
-1.40	0.040	0.036	158.20	78625.37	9000.94	321.4858	23767.70	0.92	2.87	-1.053	25.45
-1.50	0.032	0.028	167.01	87626.31	7707.31	218.6644	24089.18	0.93	3.03	-1.153	25.70
-1.60	0.025	0.023	174.20	95333.62	12536.31	282.5176	24307.85	0.94	3.17	-1.253	25.95
-1.70	0.020	0.018	185.30	107869.94	10977.25	196.5033	24590.36	0.95	3.37	-1.353	26.20
-1.80	0.016	0.014	194.50	118847.19	17463.00	248.3108	24786.86	0.96	3.53	-1.453	26.45
-1.90	0.013	0.011	208.30	136310.19	12853.87	145.1815	25035.17	0.97	3.78	-1.553	26.70
-2.00	0.010		217.90	149164.06			25180.35	0.97	3.96	-1.653	26.95
-∞							25875.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4548

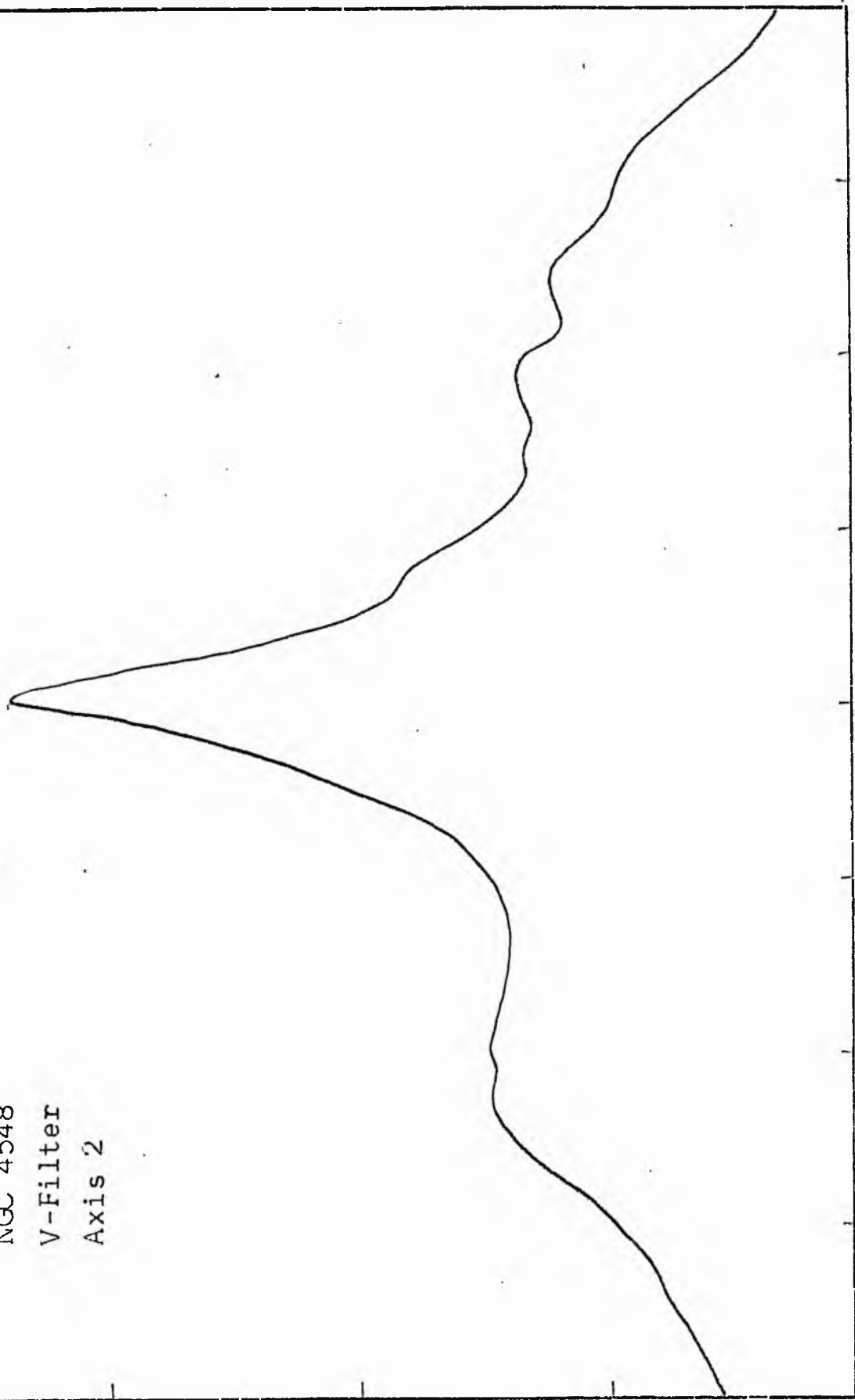
B-FILTER

Total luminosity	L_T	= 7.19
Total apparent magnitude	m_T	= 10.90
Apparent central surface brightness	μ_o	= 18.72
Major axis at threshold	$2a_m$	= 7.02
Minor axis at threshold	$2b_m$	= 6.94
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.23
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.76
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.70
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.62
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 1.16
Equivalent radius	r_1^*	= 0.40
Surface brightness	μ_1	= 21.52
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.92
Axis ratio	b/a	= 0.56
Equivalent radius	r_e^*	= 0.92
Surface brightness	μ_e	= 22.77
Mean surface brightness	μ_e'	= 12.70
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.55
Axis ratio	b/a	= 1.05
Equivalent radius	r_3^*	= 1.68
Surface brightness	μ_3	= 23.80
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.28 \\ = 1.83 \end{matrix}$

NGC 4548
V-Filter
Axis 1

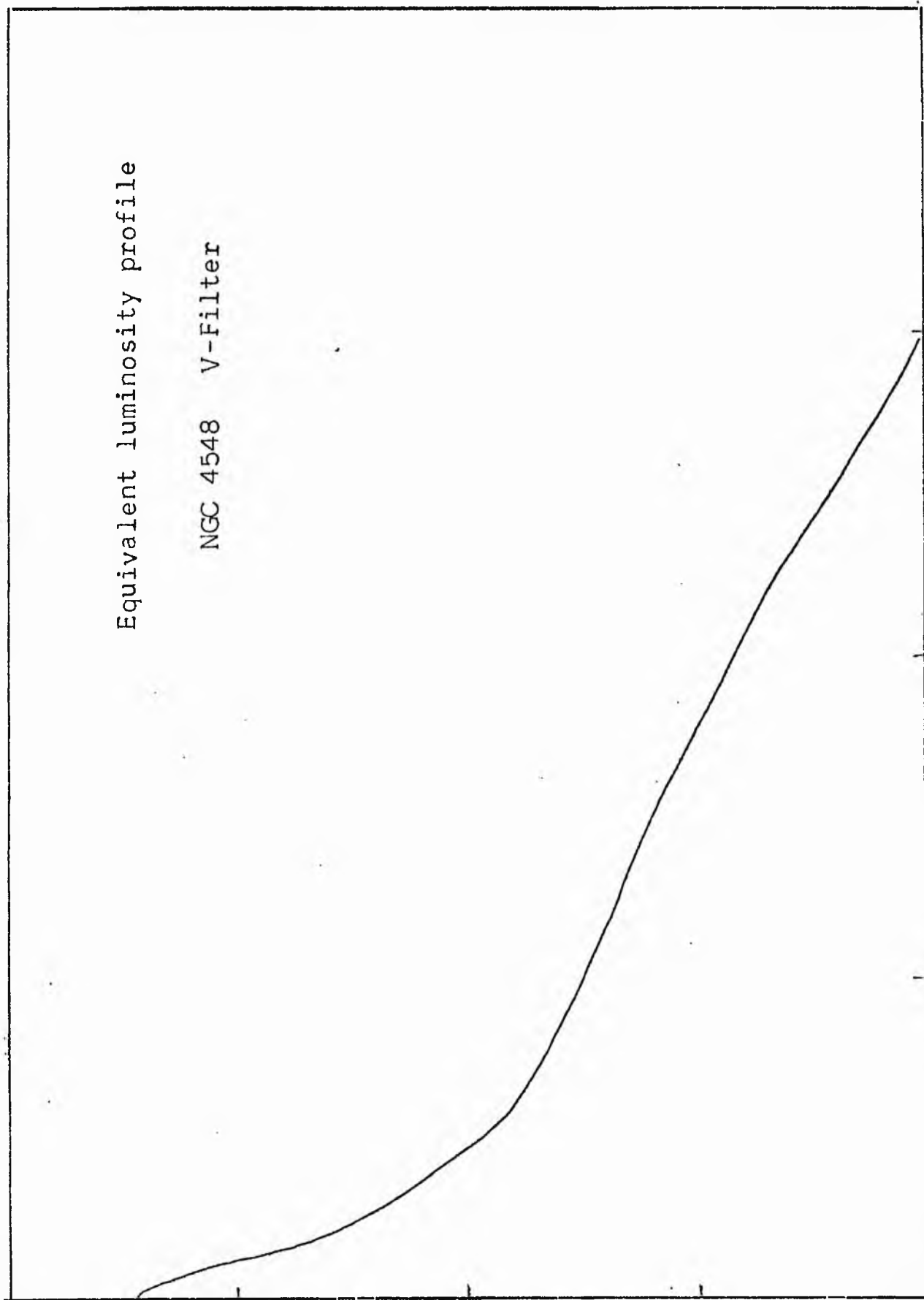


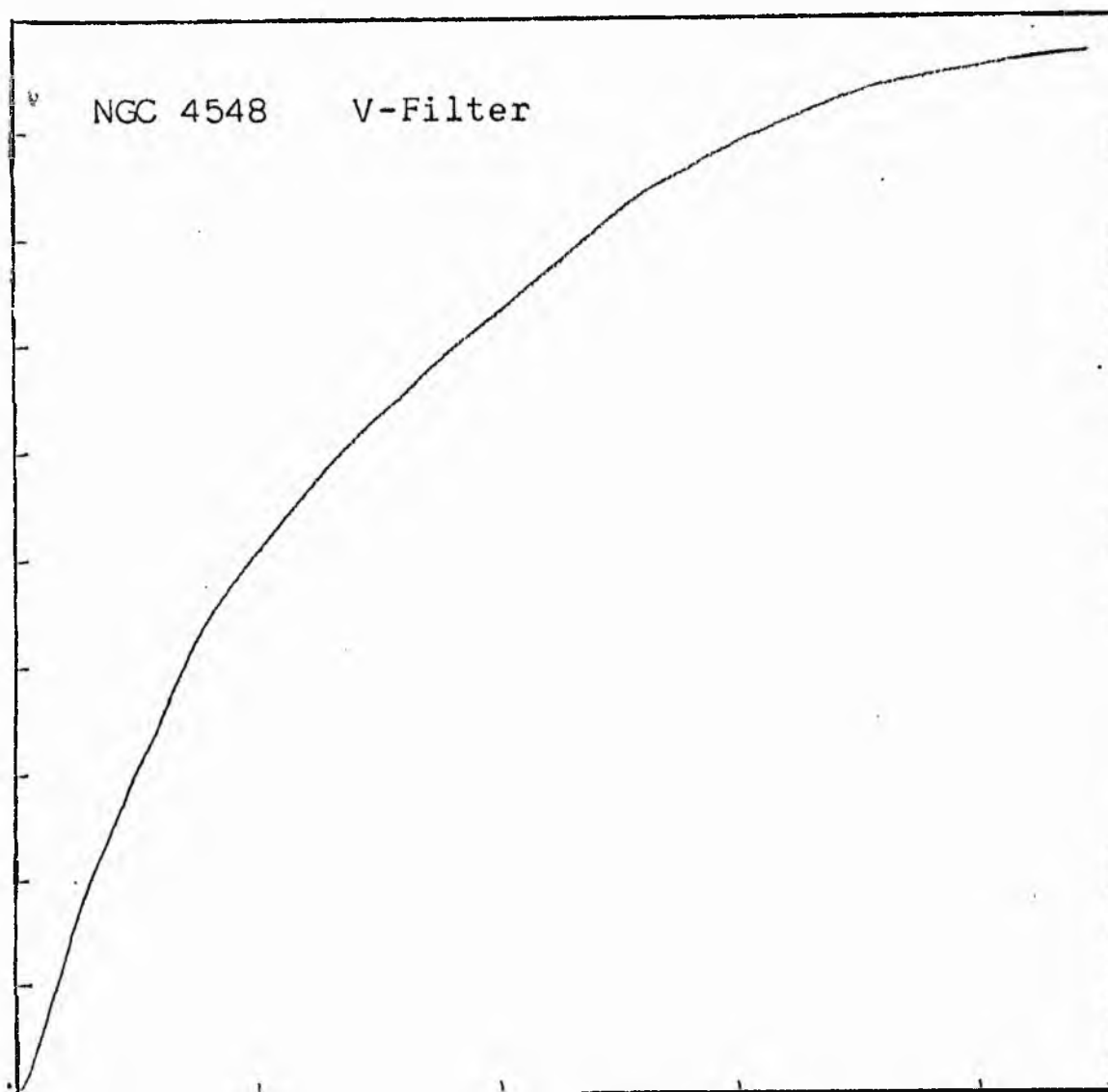
NGC 4548
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4548 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4548
V COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.37	23.442		0.0	0.0			0.0	0.0	0.0	1.630	17.83
		21.697			52.81	1145.8430					
1.30	19.953	17.901	4.10	52.81	28.90	517.3779	1145.84	0.04	0.08	1.560	18.01
1.20	15.849	14.219	5.10	81.71	42.98	611.0918	1663.22	0.06	0.10	1.460	18.26
1.10	12.589	11.295	6.30	124.69	56.77	641.1770	2274.31	0.08	0.13	1.360	18.51
1.00	10.000	8.972	7.60	181.46	53.06	476.0354	2915.49	0.10	0.15	1.260	18.76
0.90	7.943	7.126	8.64	234.52	100.07	713.1270	3391.53	0.12	0.17	1.160	19.01
0.80	6.310	5.661	10.32	334.59	106.56	613.2227	4104.65	0.14	0.21	1.060	19.26
0.70	5.012	4.496	11.85	441.15	180.77	812.8462	4717.87	0.16	0.24	0.960	19.51
0.60	3.981	3.572	14.07	621.93	167.31	597.5872	5520.71	0.19	0.28	0.860	19.76
0.50	3.162	2.837	15.85	789.24	396.79	1125.7200	6118.30	0.21	0.31	0.760	20.01
0.40	2.512	2.254	19.43	1186.03	679.75	1531.8696	7244.02	0.25	0.39	0.660	20.26
0.30	1.995	1.790	24.37	1865.78	838.61	1501.1775	8775.89	0.30	0.48	0.560	20.51
0.20	1.585	1.422	29.34	2704.39	716.80	1019.2222	10277.06	0.35	0.58	0.460	20.76
0.10	1.259	1.129	33.00	3421.19	810.18	915.0701	11296.28	0.38	0.66	0.360	21.01
-0.00	1.000	0.897	36.70	4231.38	586.27	525.9749	12211.35	0.41	0.73	0.260	21.26
-0.10	0.794	0.713	39.16	4817.64	1010.09	719.8284	12737.32	0.43	0.78	0.160	21.51
-0.20	0.631	0.566	43.07	5827.73	3483.04	1971.6479	13457.15	0.46	0.86	0.060	21.76
-0.30	0.501	0.450	54.44	9310.77	3118.65	1452.2900	15428.79	0.52	1.08	-0.040	22.01
-0.40	0.398	0.357	62.90	12429.43	4065.35	1452.0066	16831.08	0.57	1.25	-0.140	22.26
-0.50	0.316	0.284	72.46	16494.77	5163.30	1464.8687	18283.09	0.62	1.44	-0.240	22.51
-0.60	0.251	0.225	83.03	21658.08	6563.62	1479.1577	19747.95	0.67	1.65	-0.340	22.76
-0.70	0.200	0.179	94.78	28221.70	7746.40	1386.6626	21227.11	0.72	1.88	-0.440	23.01
-0.80	0.158	0.142	107.00	35968.10	9792.57	1392.4116	22613.77	0.77	2.13	-0.540	23.26
-0.90	0.126	0.113	120.69	45760.66	8498.46	959.8689	24006.18	0.81	2.40	-0.640	23.51
-1.00	0.100	0.090	131.42	54259.12	11865.69	1064.5479	24966.05	0.85	2.61	-0.740	23.76
-1.10	0.079	0.071	145.08	66124.81	8613.56	613.8406	26030.59	0.88	2.88	-0.840	24.01
-1.20	0.063	0.057	154.24	74738.37	10480.69	593.2842	26644.43	0.90	3.06	-0.940	24.26
-1.30	0.050	0.045	164.70	85219.06	4826.75	217.0344	27237.71	0.92	3.27	-1.040	24.51
-1.40	0.040	0.036	169.30	90045.81	7711.00	275.4133	27454.75	0.93	3.36	-1.140	24.76
-1.50	0.032	0.028	176.40	97756.81	8836.25	250.6935	27730.16	0.94	3.50	-1.240	25.01
-1.60	0.025	0.023	184.20	106593.06	11156.81	251.4293	27980.85	0.95	3.66	-1.340	25.26
-1.70	0.020	0.018	193.60	117749.87	10947.81	195.9764	28232.27	0.96	3.85	-1.440	25.51
-1.80	0.016	0.014	202.40	128647.69	11567.12	164.4759	28428.25	0.97	4.02	-1.540	25.76
-1.90	0.013	0.011	211.30	140264.81	12757.37	144.0916	28592.72	0.97	4.20	-1.640	26.01
-2.00	0.010		220.70	153022.19			28736.81	0.98	4.39	-1.740	26.26
-∞							29456.00	(1)		∞	

PHOTOMETRIC PARAMETERS OF NGC 4548

V-FILTER

Total luminosity	L_T	= 7.19
Total apparent magnitude	m_T	= 10.90
Apparent central surface brightness	μ_o	= 18.72
Major axis at threshold	$2a_m$	= 7.02
Minor axis at threshold	$2b_m$	= 6.94
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.23
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.76
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.70
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.63

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.28
Axis ratio	b/a	= 1.16
Equivalent radius	r_1^*	= 0.40
Surface brightness	μ_1	= 21.52

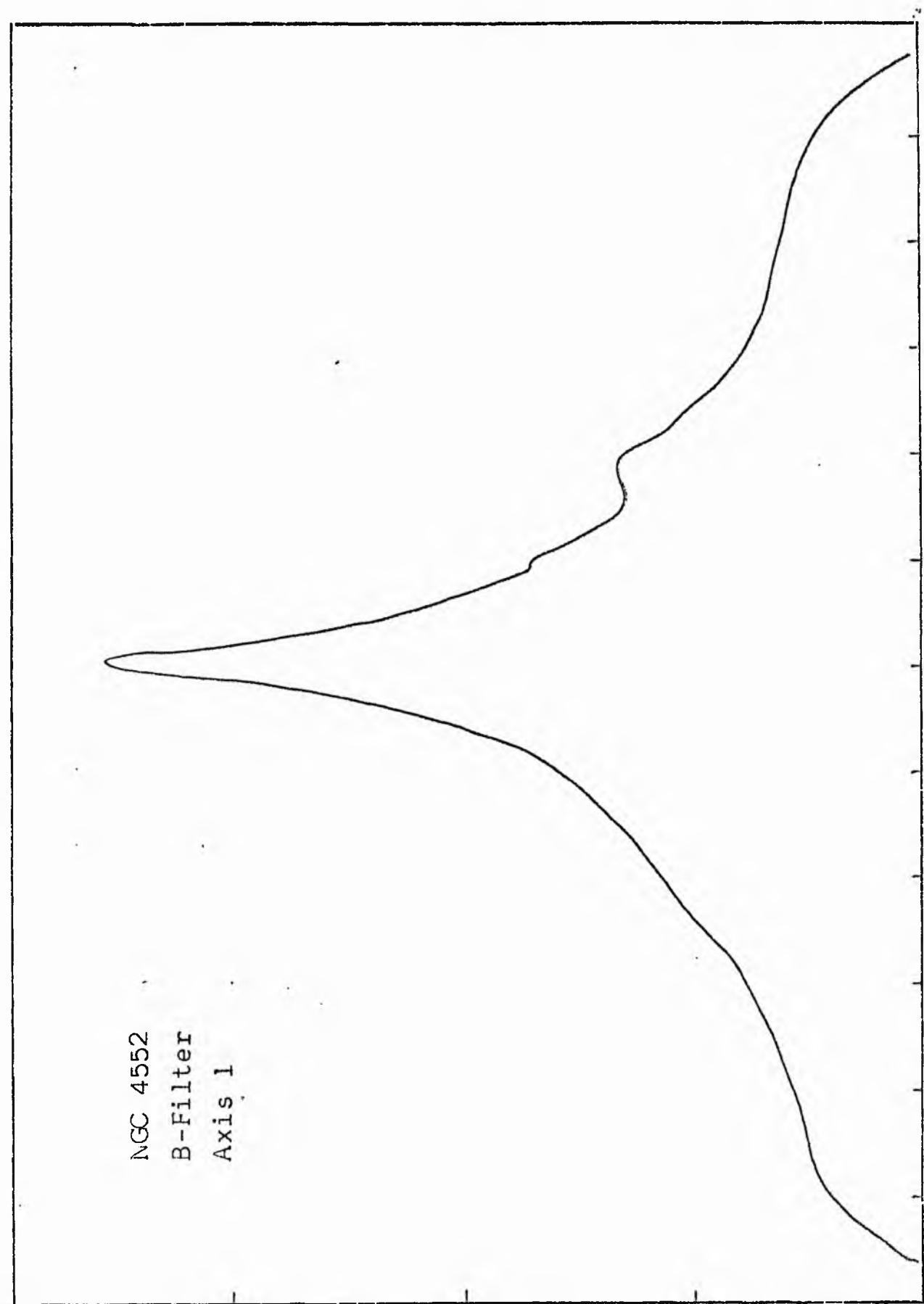
Parameters at $k = \frac{1}{2}$ (effective) :

Semi-major axis	a_e	= 0.92
Axis ratio	b/a	= 0.56
Equivalent radius	r_e^*	= 0.92
Surface brightness	μ_e	= 22.77
Mean surface brightness	μ_e'	= 12.70

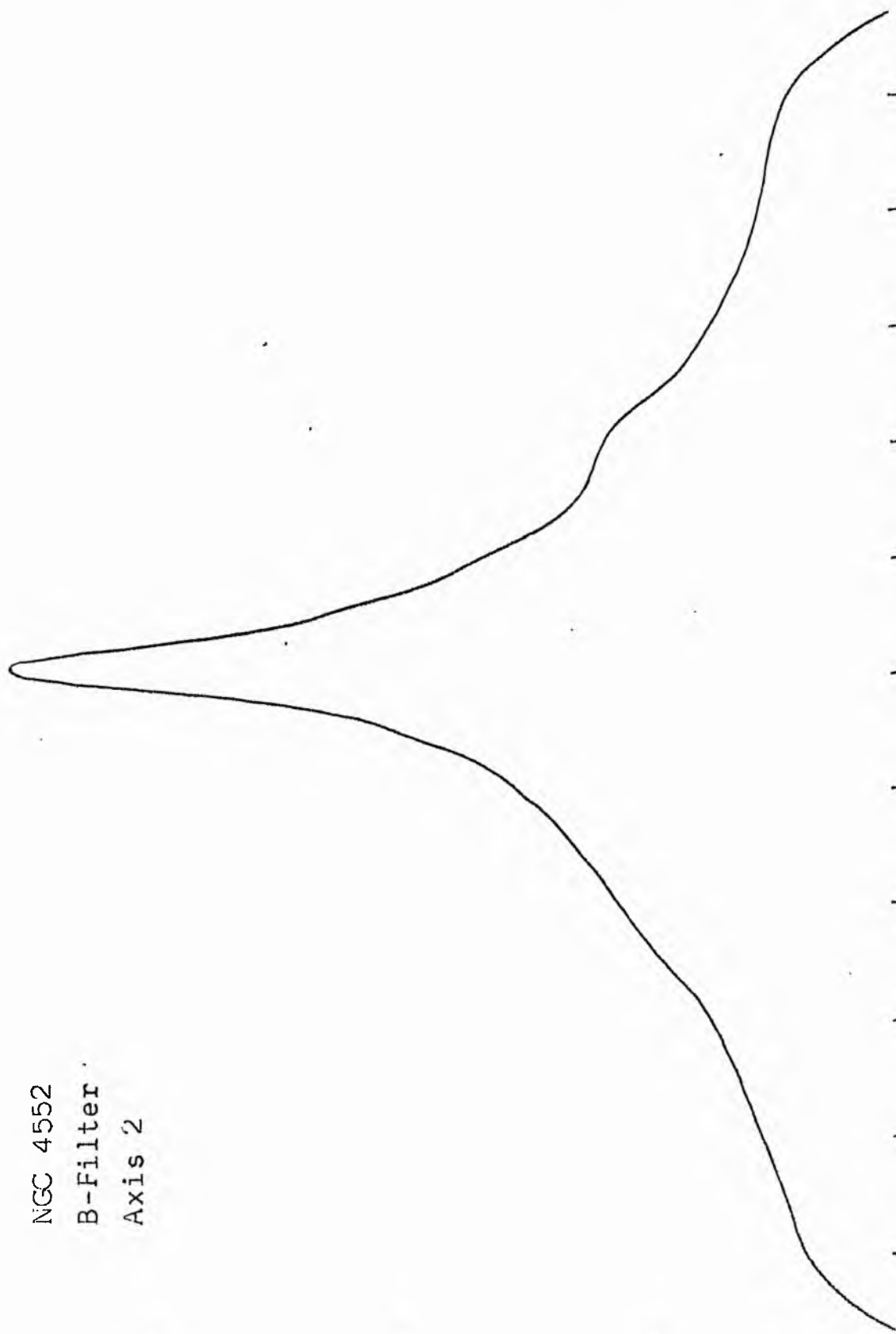
Parameters at $k = \frac{3}{4}$:

Semi-major axis	a_3	= 1.55
Axis ratio	b/a	= 1.05
Equivalent radius	r_3^*	= 1.68
Surface brightness	μ_3	= 23.80

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.28 \\ = 1.83 \end{cases}$
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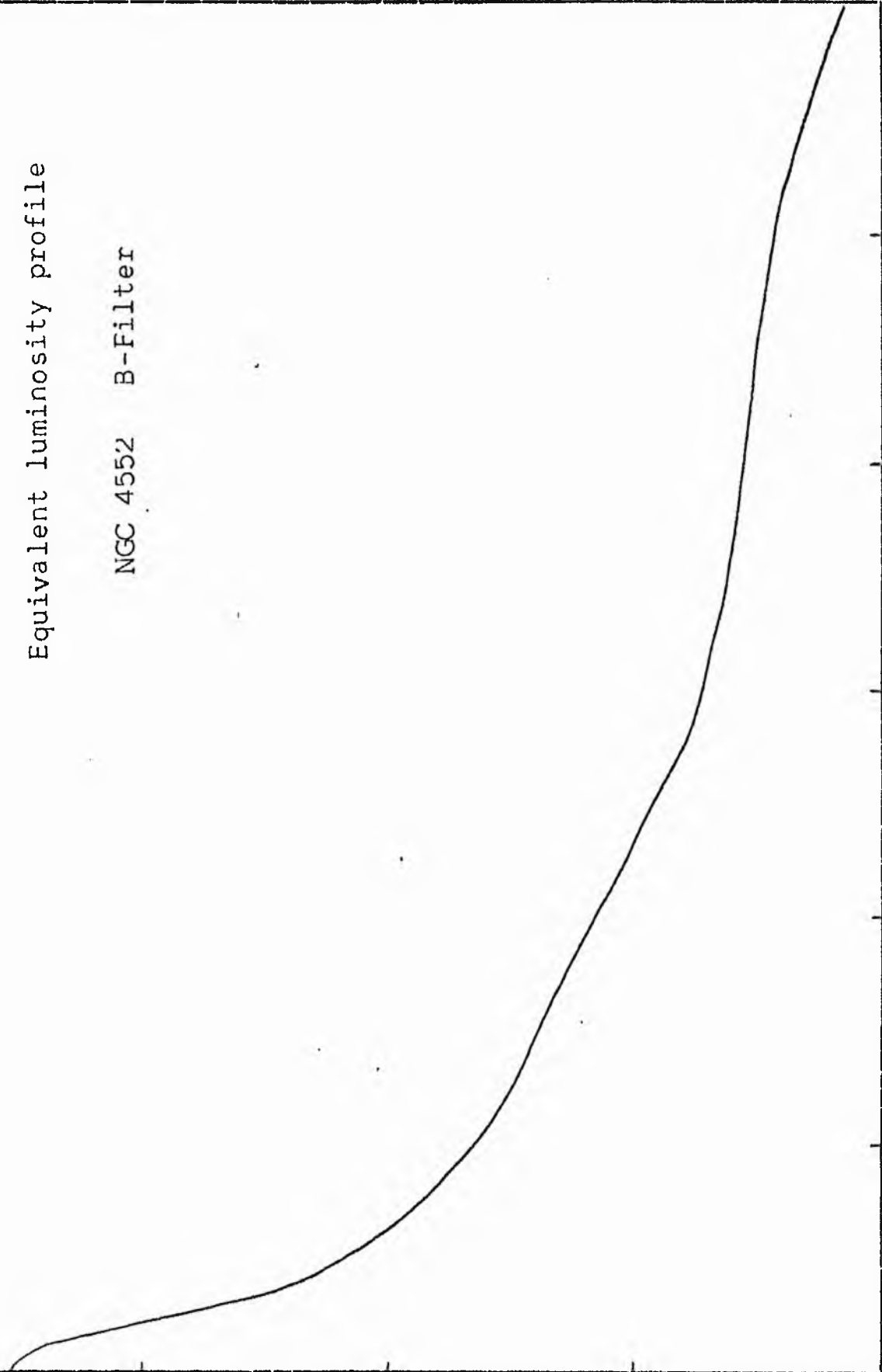


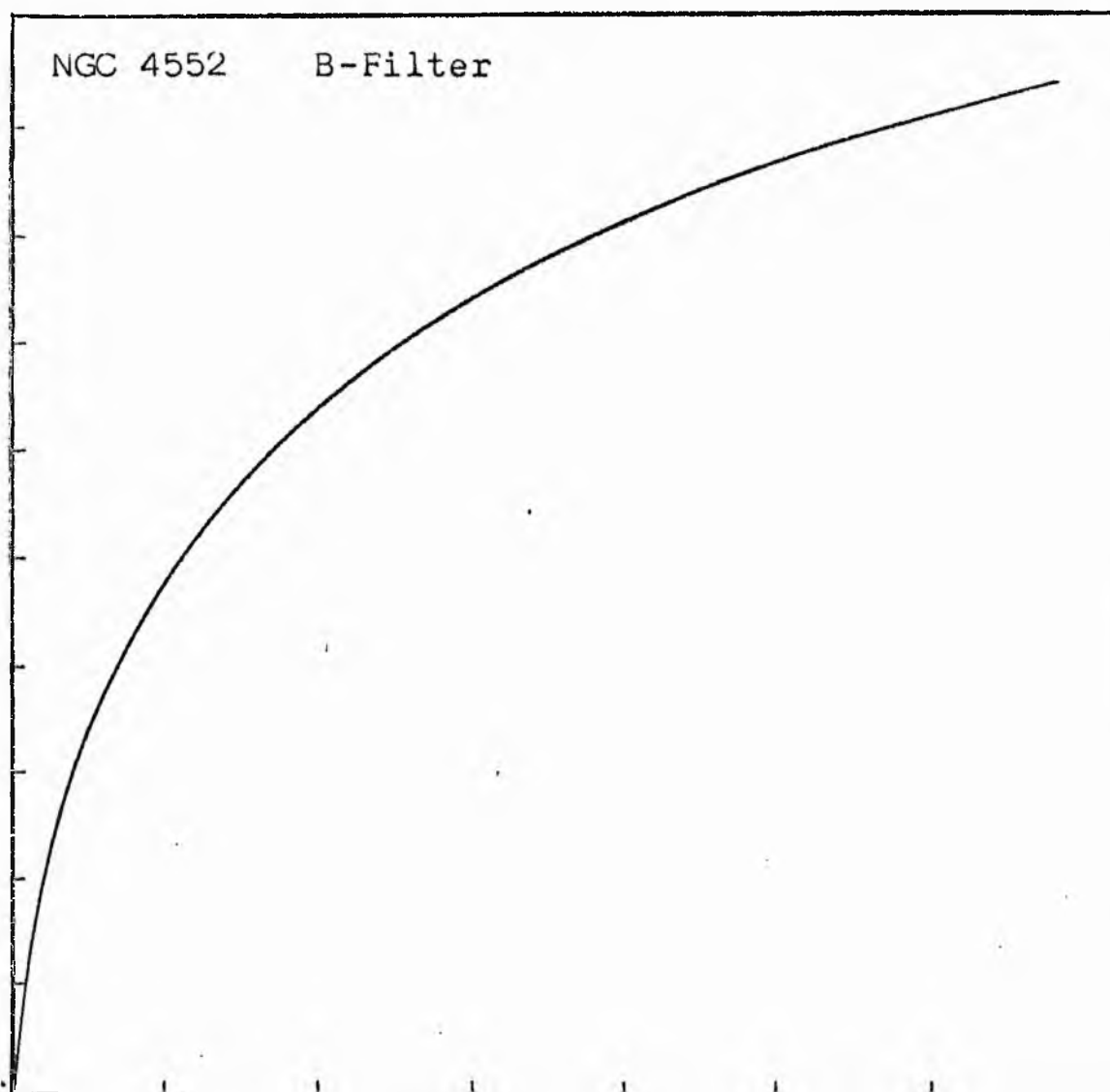
NGC 4552
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4552 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4552
B COLOR

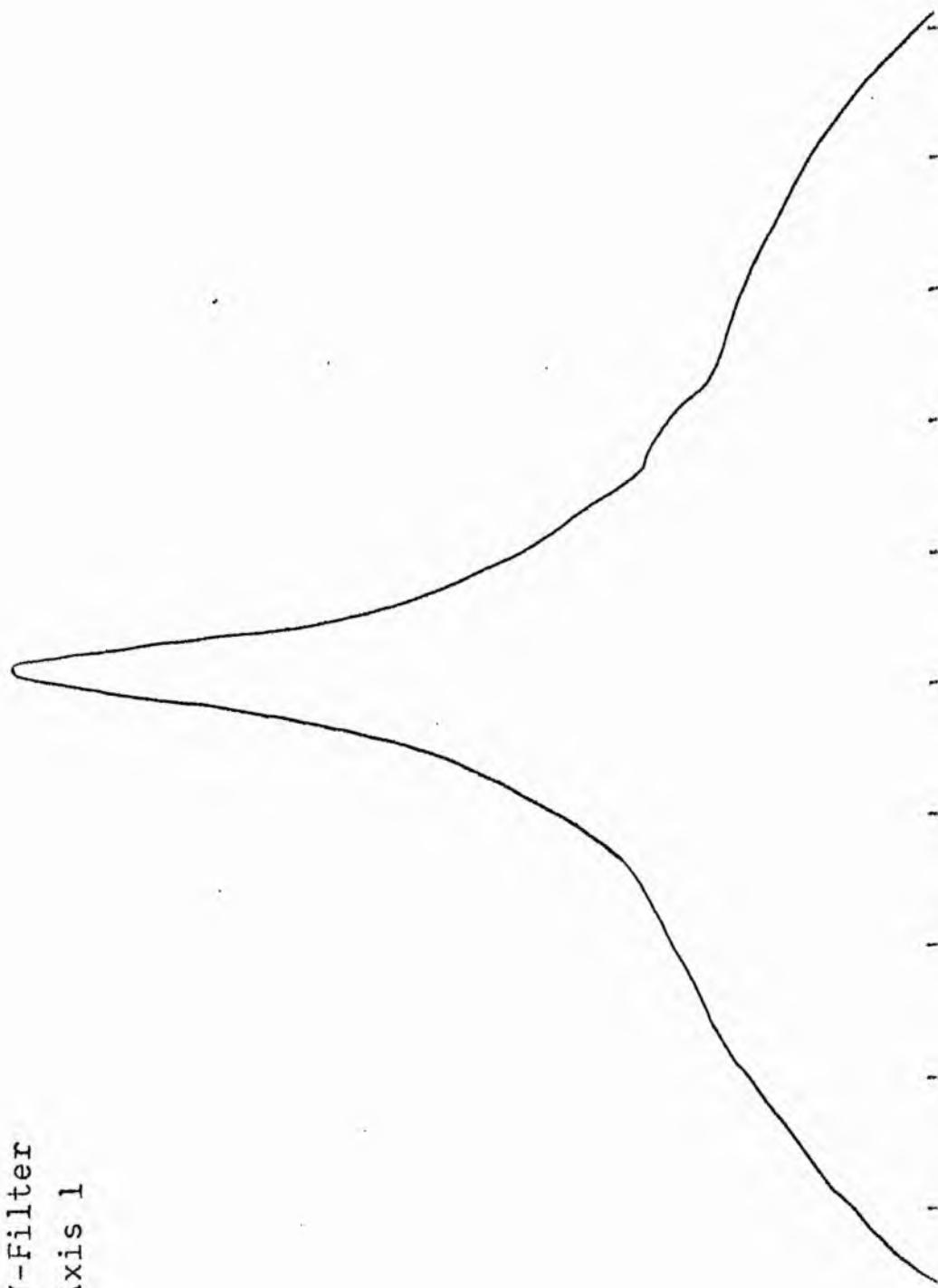
LOG I	I	T	R	AREA	ΔA	P	ΣP	KIRI	ρ	LOG J	μ
1.54	34.674		0.0	0.0			0.0	0.0	0.0	1.950	18.14
1.50	31.623	33.148	3.57	40.04	40.04	1327.2302	1327.23	0.04	0.06	1.910	18.24
1.40	25.119	28.371	5.86	107.88	67.84	1924.7229	3251.95	0.09	0.10	1.810	18.49
1.30	19.953	22.536	6.93	150.87	42.99	968.8911	4220.84	0.12	0.12	1.710	18.74
1.20	15.849	17.901	7.97	199.56	48.68	871.4446	5092.29	0.14	0.14	1.610	18.99
1.10	12.589	14.219	9.28	270.55	70.49	1009.4395	6101.72	0.17	0.16	1.510	19.24
1.00	10.000	11.295	10.44	342.41	71.86	811.6792	6913.40	0.19	0.18	1.410	19.49
0.90	7.943	8.972	11.12	388.47	46.06	413.2170	7326.61	0.20	0.19	1.310	19.74
0.80	6.310	7.126	12.18	466.06	77.59	552.9451	7879.55	0.22	0.21	1.210	19.99
0.70	5.012	5.661	13.80	548.28	132.22	748.4690	8628.02	0.24	0.24	1.110	20.24
0.60	3.981	4.496	15.60	764.54	166.25	747.5500	9375.57	0.26	0.27	1.010	20.49
0.50	3.162	3.572	16.46	851.16	86.62	309.3682	9684.94	0.27	0.29	0.910	20.74
0.40	2.512	2.837	19.20	1158.12	306.96	870.8689	10555.82	0.29	0.33	0.810	20.99
0.30	1.995	2.254	21.80	1493.01	334.89	754.7029	11310.50	0.31	0.38	0.710	21.24
0.20	1.585	1.790	25.10	1979.23	486.23	870.3779	12180.88	0.34	0.44	0.610	21.49
0.10	1.259	1.422	28.80	2605.76	626.53	890.8613	13071.74	0.36	0.50	0.510	21.74
-0.00	1.000	1.129	32.80	3379.85	774.09	874.3010	13946.04	0.39	0.57	0.410	21.99
-0.10	0.794	0.897	37.40	4512.61	1132.76	1016.2729	14962.31	0.41	0.66	0.310	22.24
-0.20	0.631	0.713	41.90	5515.41	1002.79	714.6311	15676.94	0.43	0.73	0.210	22.49
-0.30	0.501	0.566	47.10	6969.34	1453.93	823.0266	16499.96	0.46	0.82	0.110	22.74
-0.40	0.398	0.450	56.80	10135.53	3166.19	1423.6658	17923.62	0.50	0.99	0.010	22.99
-0.50	0.316	0.357	65.12	13322.28	3186.75	1138.2007	19061.82	0.53	1.13	-0.090	23.24
-0.60	0.251	0.284	73.24	16851.80	3529.52	1001.3516	20063.18	0.56	1.27	-0.190	23.49
-0.70	0.200	0.225	86.53	23522.48	6670.68	1503.2852	21566.46	0.60	1.50	-0.290	23.74
-0.80	0.158	0.179	101.42	32314.47	8791.99	1573.8315	23140.29	0.64	1.76	-0.390	23.99
-0.90	0.126	0.142	109.54	37646.00	5381.52	765.2024	23905.49	0.66	1.90	-0.490	24.24
-1.00	0.100	0.113	118.48	44100.13	6404.14	723.3228	24628.81	0.68	2.05	-0.590	24.49
-1.10	0.079	0.090	128.23	51656.59	7556.86	677.9749	25306.78	0.70	2.22	-0.690	24.74
-1.20	0.063	0.071	139.61	61232.61	9575.62	682.4011	25989.18	0.72	2.42	-0.790	24.99
-1.30	0.050	0.057	150.82	71460.75	10228.14	578.9880	26568.16	0.74	2.62	-0.890	25.24
-1.40	0.040	0.045	194.53	118883.87	47423.12	2132.3765	28700.54	0.80	3.37	-0.990	25.49
-1.50	0.032	0.036	231.91	168961.87	50078.00	1788.6333	30489.17	0.84	4.02	-1.090	25.74
-1.60	0.025	0.028	255.76	205501.50	36539.62	1036.6667	31525.84	0.87	4.44	-1.190	25.99
-1.70	0.020	0.023	274.13	236081.87	30580.37	689.1575	32214.99	0.89	4.75	-1.290	26.24
-1.80	0.016	0.018	288.34	261191.44	25109.56	449.4851	32664.48	0.91	5.00	-1.390	26.49
-1.90	0.013	0.014	310.85	303564.62	42373.19	602.5149	33266.99	0.92	5.39	-1.490	26.74
-2.00	0.010	0.011	325.10	332034.56	28469.94	321.5613	33588.55	0.93	5.64	-1.590	26.99
-∞							36088.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4552

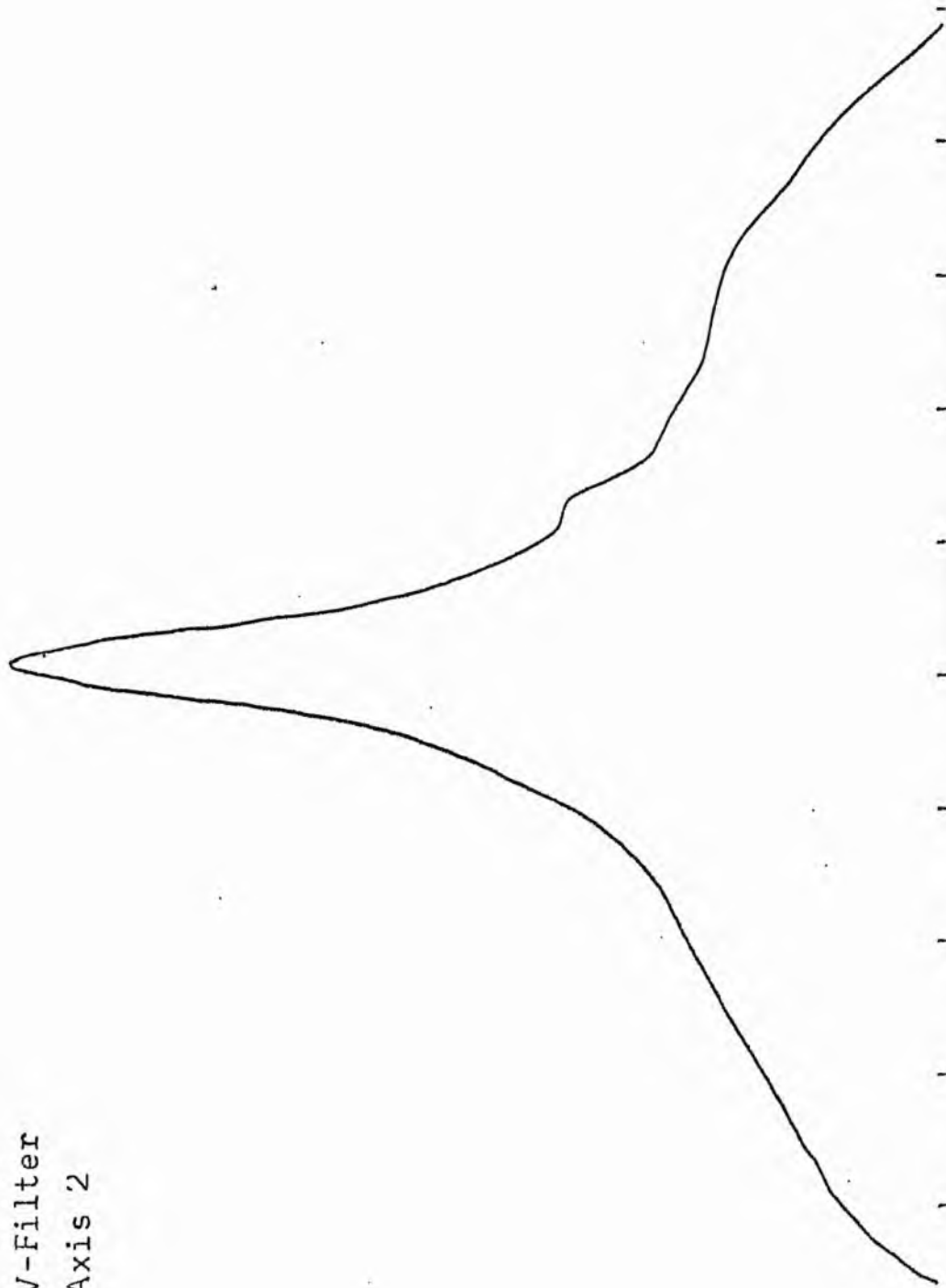
B-FILTER

Total luminosity	L_T	= 10.02
Total apparent magnitude	m_T	= 10.63
Apparent central surface brightness	μ_0	= 18.14
Major axis at threshold	$2a_m$	= 9.47
Minor axis at threshold	$2b_m$	= 9.33
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 4.95
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.72
Gradient of exponential component	$G(a)$	= -0.38
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.43
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -----
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.26
Axis ratio	b/a	= 0.93
Equivalent radius	r_1^*	= 0.25
Surface brightness	μ_1	= 20.36
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.94
Axis ratio	b/a	= 0.96
Equivalent radius	r_e^*	= 0.96
Surface brightness	μ_e	= 22.99
Mean surface brightness	μ_e'	= 12.54
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.92
Axis ratio	b/a	= 0.94
Equivalent radius	r_3^*	= 2.68
Surface brightness	μ_3	= 25.28
Concentration indices	$\{C_{21}$	= 3.92
	C_{32}	= 2.79

NGC 4552
V-Filter
Axis 1

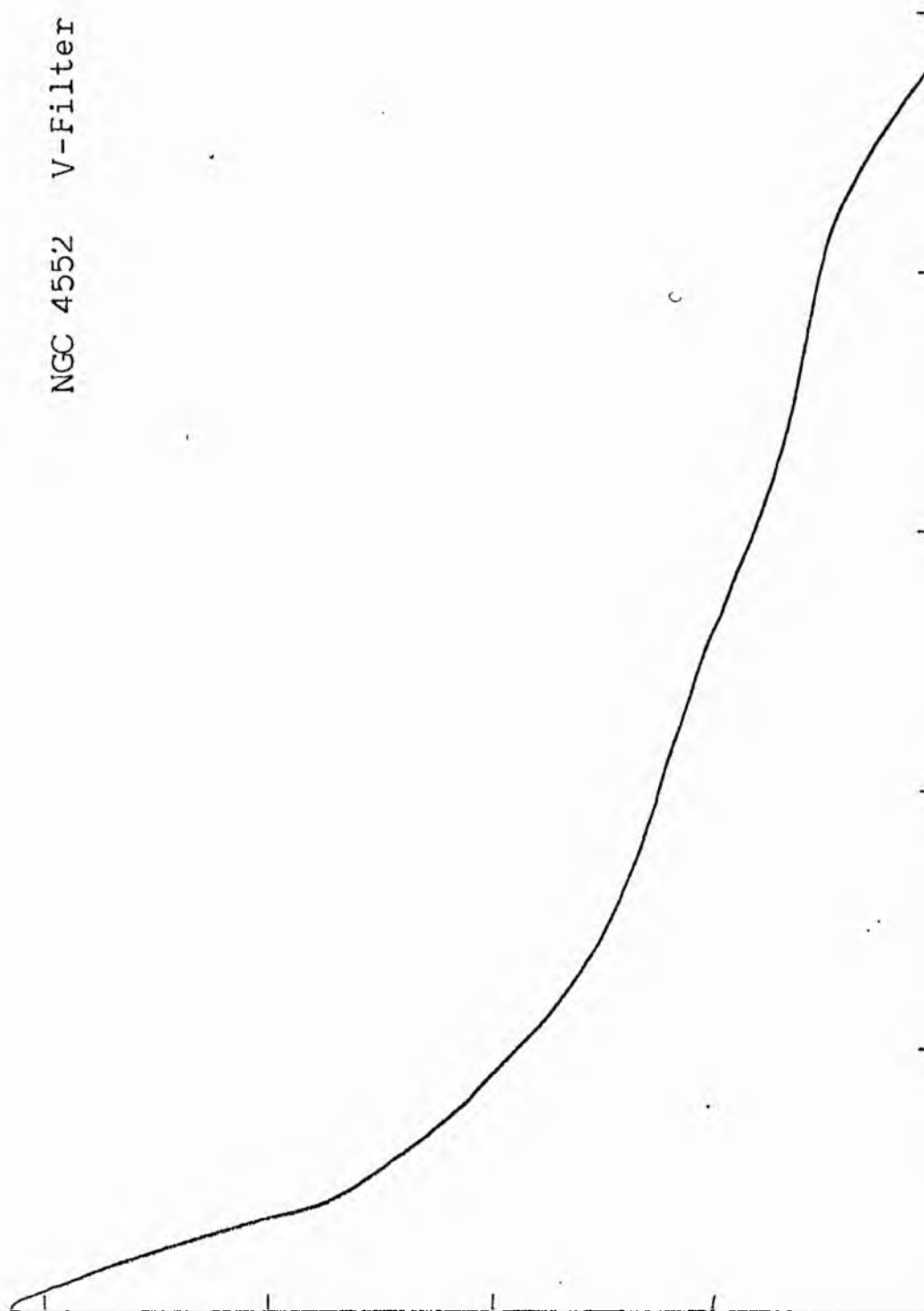


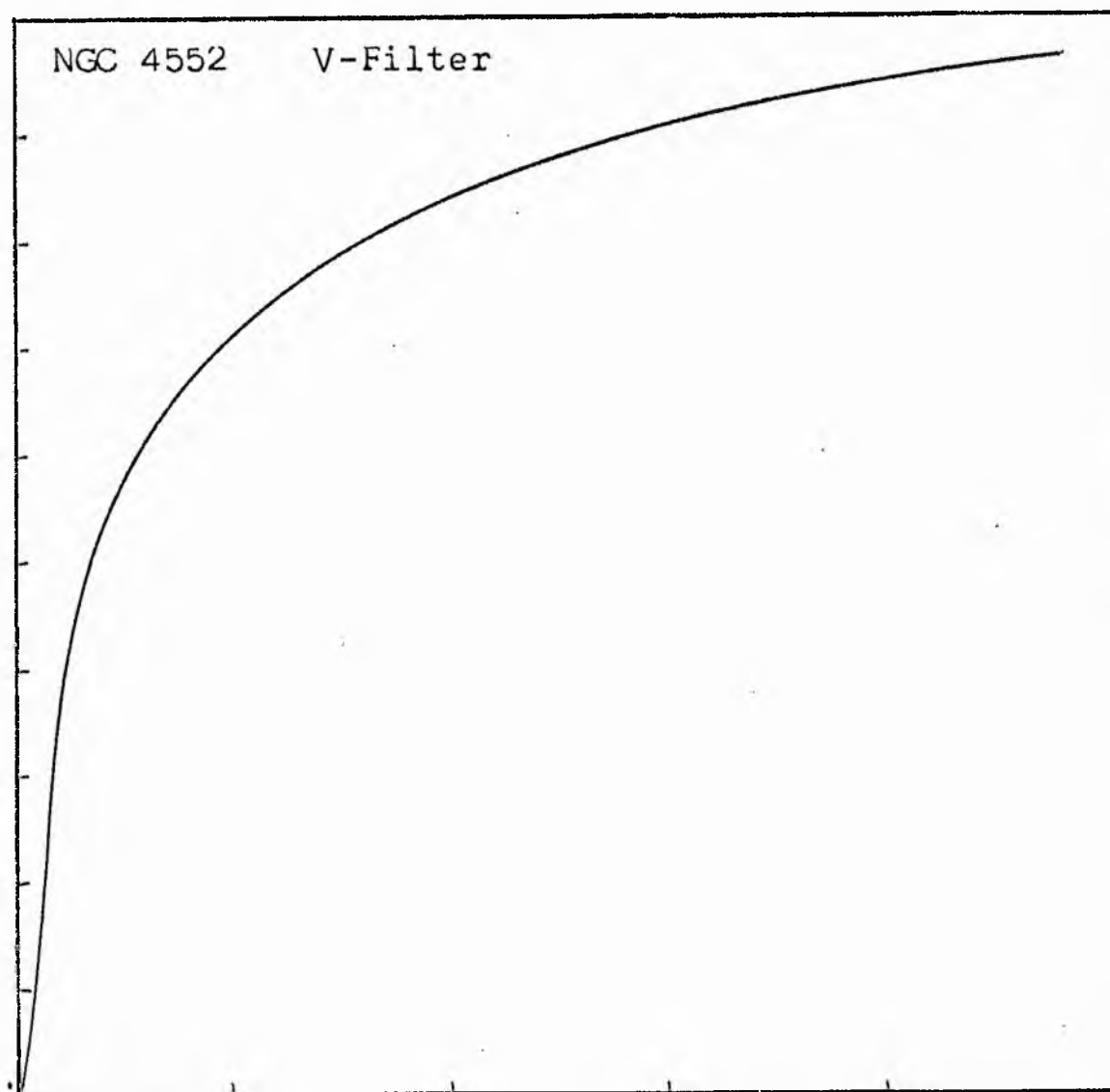
NGC 4552
V-Filter
Axis '2



Equivalent luminosity profile

NGC 4552 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4552
V COLOUR

LOG I	I	I	R	AREA	ΔA	P	Σ P	K(R)	ρ	LOG J	μ
2.12	131.826		0.0	0.0	8.04	1036.3445	0.0	0.0	0.0	1.357	16.56
2.10	125.892	128.859	1.60	8.04	37.32	4215.3828	1036.34	0.02	0.09	1.337	16.61
2.00	100.000	112.946	3.80	45.36	36.35	3261.0178	5251.73	0.09	0.22	1.237	16.86
1.90	79.433	89.716	5.10	81.71	39.05	2782.8618	8512.74	0.15	0.30	1.137	17.11
1.80	63.095	71.264	6.20	120.76	46.65	2640.8630	11295.60	0.20	0.37	1.037	17.36
1.70	50.118	56.607	7.30	167.42	64.94	2919.8440	13936.46	0.24	0.43	0.937	17.61
1.60	39.810	44.964	8.60	232.35	114.01	4071.9841	16856.30	0.29	0.51	0.837	17.86
1.50	31.623	35.717	10.50	346.36	54.79	1554.4558	20928.29	0.36	0.62	0.737	18.11
1.40	25.119	28.371	11.30	401.15	36.29	817.7114	22482.69	0.39	0.67	0.637	18.36
1.30	19.952	22.536	11.80	437.44	53.44	956.5818	23300.40	0.40	0.70	0.537	18.61
1.20	15.849	17.901	12.50	490.87	133.71	1901.1594	24256.98	0.42	0.74	0.437	18.86
1.10	12.589	14.219	14.10	624.58	156.71	1769.9841	26158.13	0.45	0.83	0.337	19.11
1.00	10.000	11.295	15.77	781.29	43.19	387.4521	27928.11	0.48	0.93	0.237	19.36
0.90	7.943	8.972	16.20	824.48	211.57	1507.7373	28315.56	0.49	0.96	0.137	19.61
0.80	6.310	7.126	18.16	1036.05	229.40	1298.5393	29823.30	0.52	1.07	0.037	19.86
0.70	5.012	5.661	20.07	1265.45	162.54	730.8416	31121.84	0.54	1.18	-0.063	20.11
0.60	3.981	4.496	21.32	1427.99	315.83	1128.0239	31852.68	0.55	1.26	-0.163	20.36
0.50	3.162	3.572	23.56	1743.81	512.60	1454.2776	32980.70	0.57	1.39	-0.263	20.61
0.40	2.512	2.837	26.80	2256.42	639.28	1440.6521	34434.98	0.59	1.58	-0.363	20.86
0.30	1.995	2.254	30.36	2895.70	787.43	1409.5549	35875.62	0.62	1.79	-0.463	21.11
0.20	1.585	1.790	34.24	3683.13	1009.85	1435.9016	37285.18	0.64	2.02	-0.563	21.36
0.10	1.259	1.422	38.65	4692.98	1032.37	1166.0088	38721.08	0.67	2.28	-0.663	21.61
-0.00	1.000	1.129	42.69	5725.35	1126.12	1010.3025	39887.09	0.69	2.52	-0.763	21.86
-0.10	0.794	0.897	46.70	6851.46	1220.78	869.9697	40897.39	0.71	2.75	-0.863	22.11
-0.20	0.631	0.713	50.69	8072.24	1503.78	851.2405	41767.36	0.72	2.99	-0.963	22.36
-0.30	0.501	0.566	55.21	9576.02	3179.59	1429.6790	42618.59	0.74	3.26	-1.063	22.61
-0.40	0.398	0.450	63.72	12755.61	2726.28	973.7307	44048.27	0.76	3.76	-1.163	22.86
-0.50	0.316	0.357	70.20	15481.89	5797.01	1644.6462	45022.00	0.78	4.14	-1.263	23.11
-0.60	0.251	0.284	82.30	21278.90	7915.79	1783.8667	46666.64	0.81	4.85	-1.363	23.36
-0.70	0.200	0.225	96.40	29194.69	5112.29	915.1328	48450.51	0.84	5.68	-1.463	23.61
-0.80	0.158	0.179	104.50	34306.98	6664.53	947.6282	49365.64	0.85	6.16	-1.563	23.86
-0.90	0.126	0.142	114.20	40971.51	11008.25	1243.3320	50313.27	0.87	6.73	-1.663	24.11
-1.00	0.100	0.113	128.63	51979.77	1998.96	179.3386	51556.60	0.89	7.58	-1.763	24.36
-1.10	0.079	0.090	131.08	53978.73	12969.09	924.2285	51735.93	0.89	7.73	-1.863	24.61
-1.20	0.063	0.071	145.98	66947.81	24883.25	1438.5669	52660.16	0.91	8.61	-1.963	24.86
-1.30	0.050	0.057	170.97	91831.06	21711.75	976.2603	54068.73	0.93	10.08	-2.063	25.11
-1.40	0.040	0.045	190.11	113542.81	14786.37	528.1206	55044.98	0.95	11.21	-2.163	25.36
-1.50	0.032	0.036	202.11	128329.19	13827.19	392.2888	55573.10	0.96	11.92	-2.263	25.61
-1.60	0.025	0.028	212.72	142156.37	4662.12	105.0647	55965.39	0.97	12.54	-2.363	25.86
-1.70	0.020	0.023	216.18	146818.50	9676.06	173.2097	56070.45	0.97	12.75	-2.463	26.11
-1.80	0.016	0.018	223.19	156494.56	9580.06	136.2204	56243.66	0.97	13.16	-2.563	26.36
-1.90	0.013	0.014	229.92	166074.62	9493.12	107.2219	56379.87	0.97	13.56	-2.663	26.61
-2.00	0.010	0.011	236.40	175567.75			56487.09	0.98	13.94	-2.763	26.86
-∞							57888.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4552

V-FILTER

Total luminosity	L_T	= 16.08
Total apparent magnitude	m_T	= 9.95
Apparent central surface brightness	μ_0	= 16.56
Major axis at threshold	$2a_m$	= 8.12
Minor axis at threshold	$2b_m$	= 7.97
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.89
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.92
Gradient of exponential component	$G(a)$	= -0.48
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.54
Equivalent gradient of reduced exp. comp....	$G(p)$	= -----

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.16
Axis ratio	b/a	= 0.86
Equivalent radius	r_1^*	= 0.13
Surface brightness	μ_1	= 17.66

Parameters at $k = \frac{1}{2}$ (effective) :

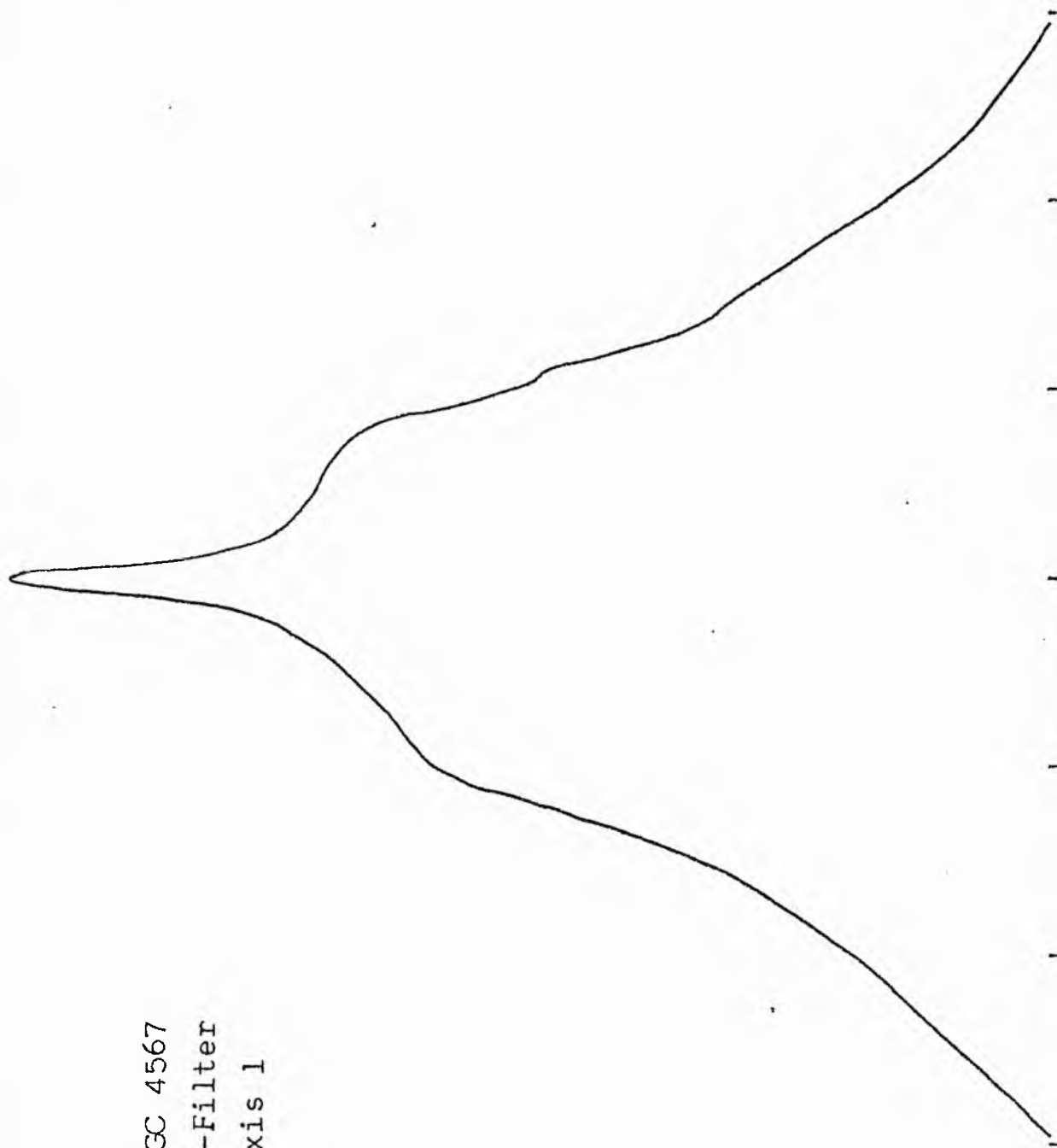
Semi-major axis	a_e	= 0.27
Axis ratio	b/a	= 0.91
Equivalent radius	r_e^*	= 0.28
Surface brightness	μ_e	= 19.69
Mean surface brightness	μ_e'	= 9.20

Parameters at $k = \frac{3}{4}$:

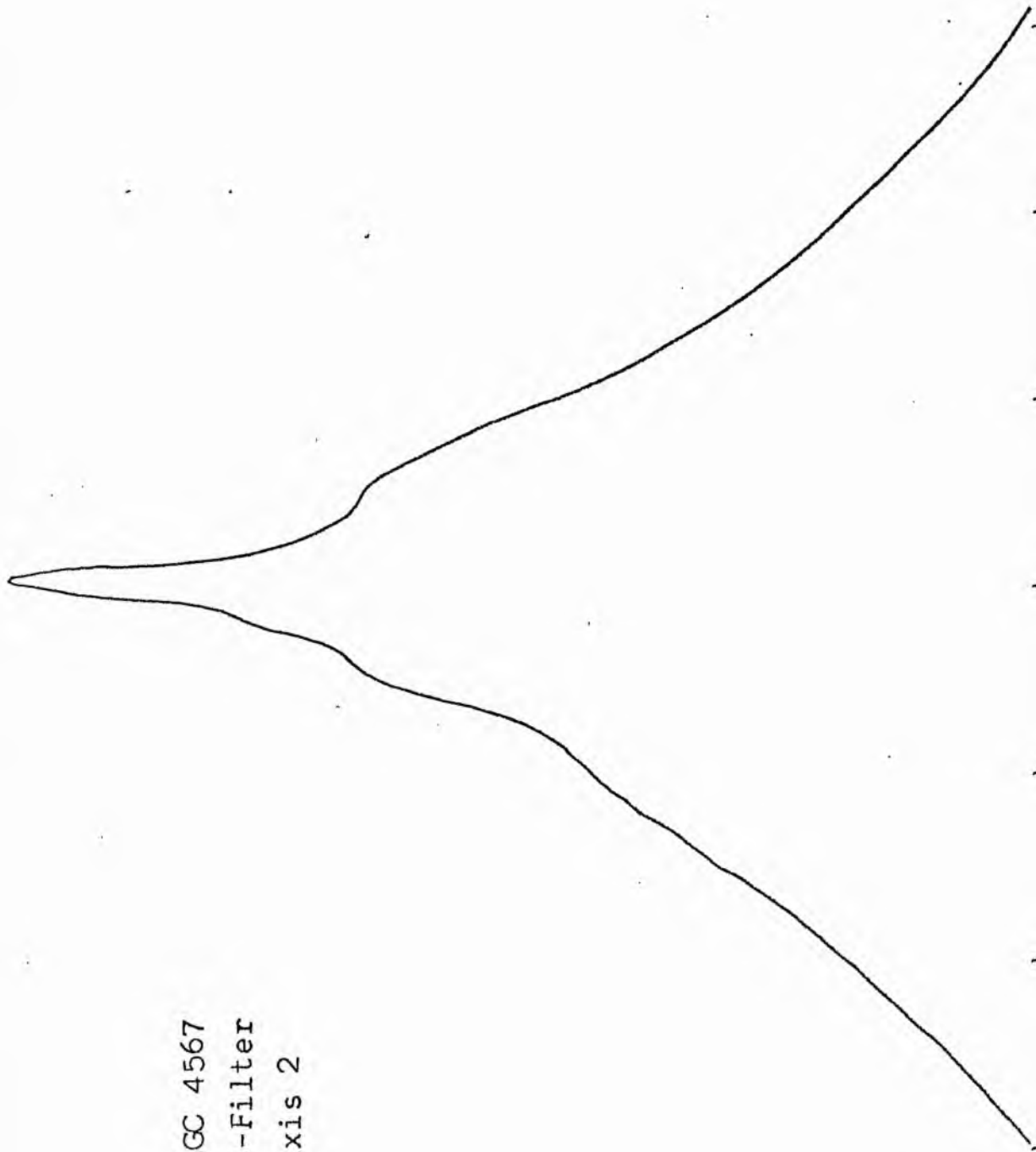
Semi-major axis	a_3	= 1.02
Axis ratio	b/a	= 0.93
Equivalent radius	r_3^*	= 1.00
Surface brightness	μ_3	= 22.73

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.25 \\ = 3.53 \end{matrix}$
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NGC 4567
B-Filter
Axis 1

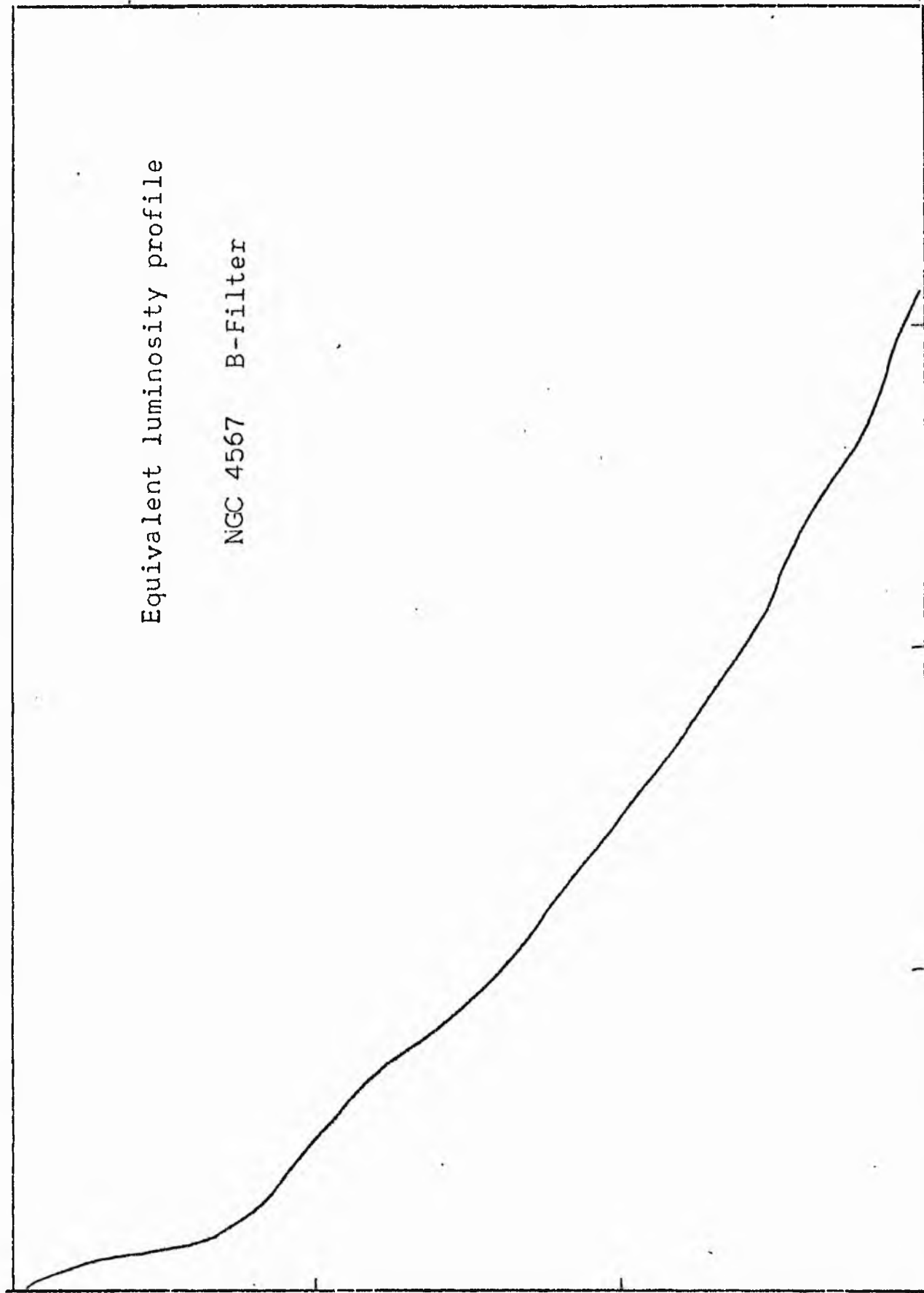


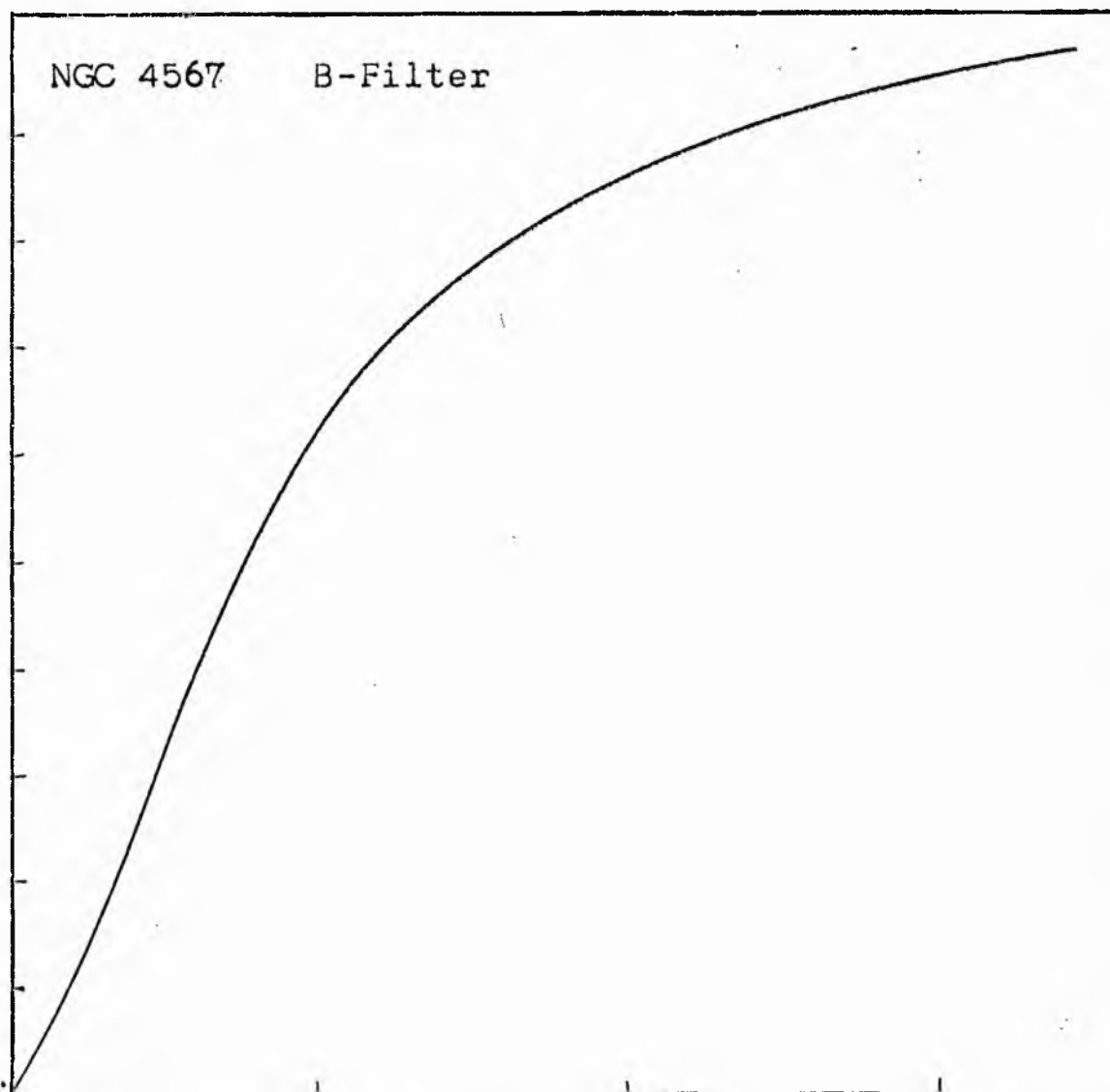
NGC 4567
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4567 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4567
B COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	P	LOG J	μ
0.94	8.710	8.326	0.0	0.0	16.62	138.3774	0.0	0.0	0.0	1.241	19.60
0.90	7.943	7.126	2.30	16.62	28.75	204.8531	138.38	0.01	0.06	1.201	19.70
0.80	6.310	5.661	3.80	45.36	24.03	136.0452	343.23	0.04	0.10	1.101	19.95
0.70	5.012	4.496	4.70	69.40	12.31	55.3740	479.28	0.05	0.13	1.001	20.20
0.60	3.981	3.572	5.10	81.71	46.97	167.7503	534.65	0.06	0.14	0.901	20.45
0.50	3.162	2.837	6.40	128.68	43.35	122.9987	702.40	0.08	0.17	0.801	20.70
0.40	2.512	2.254	7.40	172.03	44.39	100.0379	825.40	0.09	0.20	0.701	20.95
0.30	1.995	1.790	8.30	216.42	124.94	232.5961	925.44	0.10	0.23	0.601	21.20
0.20	1.585	1.422	10.50	346.36	508.94	723.6641	1158.03	0.12	0.29	0.501	21.45
0.10	1.259	1.129	16.50	655.30	763.53	862.3811	1881.70	0.20	0.45	0.401	21.70
0.00	1.000	0.897	22.70	1618.83	1005.06	901.7021	2744.08	0.29	0.62	0.301	21.95
-0.10	0.794	0.713	28.90	2623.89	797.31	568.1948	3645.78	0.39	0.79	0.201	22.20
-0.20	0.631	0.566	33.00	3421.19	810.18	458.6228	4213.97	0.45	0.90	0.101	22.45
-0.30	0.501	0.450	36.70	4231.38	820.33	368.8584	4672.59	0.50	1.00	0.001	22.70
-0.40	0.398	0.357	40.10	5051.71	703.18	251.1546	5041.45	0.54	1.09	-0.099	22.95
-0.50	0.316	0.284	42.80	5754.89	691.93	196.3075	5292.60	0.57	1.17	-0.199	23.20
-0.60	0.251	0.225	45.30	6446.82	2578.84	581.1614	5488.91	0.59	1.23	-0.299	23.45
-0.70	0.200	0.179	53.60	9025.66	2397.45	429.1619	6070.07	0.65	1.46	-0.399	23.70
-0.80	0.158	0.142	60.30	11423.11	2511.61	357.1284	6499.23	0.70	1.64	-0.499	23.95
-0.90	0.126	0.113	66.60	13934.72	2715.21	306.6731	6856.35	0.73	1.81	-0.599	24.20
-1.00	0.100	0.090	72.80	16649.93	3056.14	274.1870	7163.02	0.77	1.98	-0.699	24.45
-1.10	0.079	0.071	79.20	19706.07	3152.43	224.6565	7437.21	0.80	2.16	-0.799	24.70
-1.20	0.063	0.057	85.30	22858.50	3501.21	198.1947	7661.86	0.82	2.32	-0.899	24.95
-1.30	0.050	0.045	91.60	26359.71	3689.11	165.8811	7860.05	0.84	2.49	-0.999	25.20
-1.40	0.040	0.036	97.80	30048.82	8102.80	289.4080	8025.93	0.86	2.66	-1.099	25.45
-1.50	0.032	0.028	110.20	38151.62	3903.27	110.7400	8315.34	0.89	3.00	-1.199	25.70
-1.60	0.025	0.023	115.70	42054.89	8138.20	183.4023	8426.00	0.90	3.15	-1.299	25.95
-1.70	0.020	0.018	126.40	50193.09	4049.52	72.4904	8609.48	0.92	3.44	-1.399	26.20
-1.80	0.016	0.014	131.40	54242.61	13828.89	196.6370	8681.97	0.93	3.58	-1.499	26.45
-1.90	0.013	0.011	147.20	68071.50	6434.50	72.6763	8878.61	0.95	4.01	-1.599	26.70
-2.00	0.010		154.00	74506.00			8951.28	0.96	4.19	-1.699	26.95
-∞							9351.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4567

B-FILTER

Total luminosity	L_T	= 2.60
Total apparent magnitude	m_T	= 12.04
Apparent central surface brightness	μ_o	= 19.60
Major axis at threshold	$2a_m$	= 4.95
Minor axis at threshold	$2b_m$	= 4.75
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.12
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -1.18
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.89
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.84

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.37
Axis ratio	b/a	= 0.82
Equivalent radius	r_1^*	= 0.33
Surface brightness	μ_1	= 21.81

Parameters at $k = \frac{1}{2}$ (effective) :

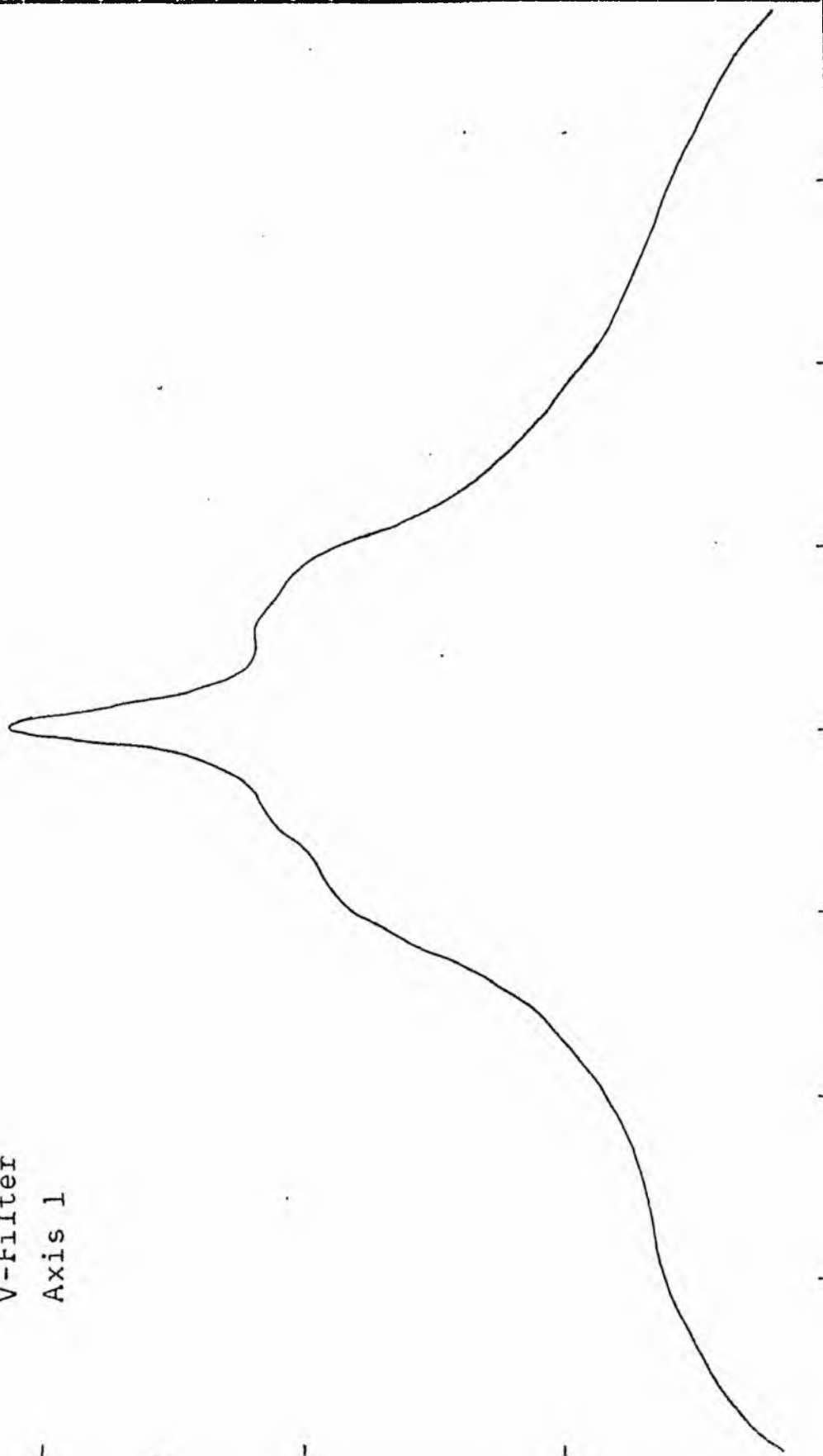
Semi-major axis	a_e	= 0.81
Axis ratio	b/a	= 0.73
Equivalent radius	r_e^*	= 0.61
Surface brightness	μ_e	= 22.70
Mean surface brightness	μ_e'	= 12.96

Parameters at $k = \frac{3}{4}$:

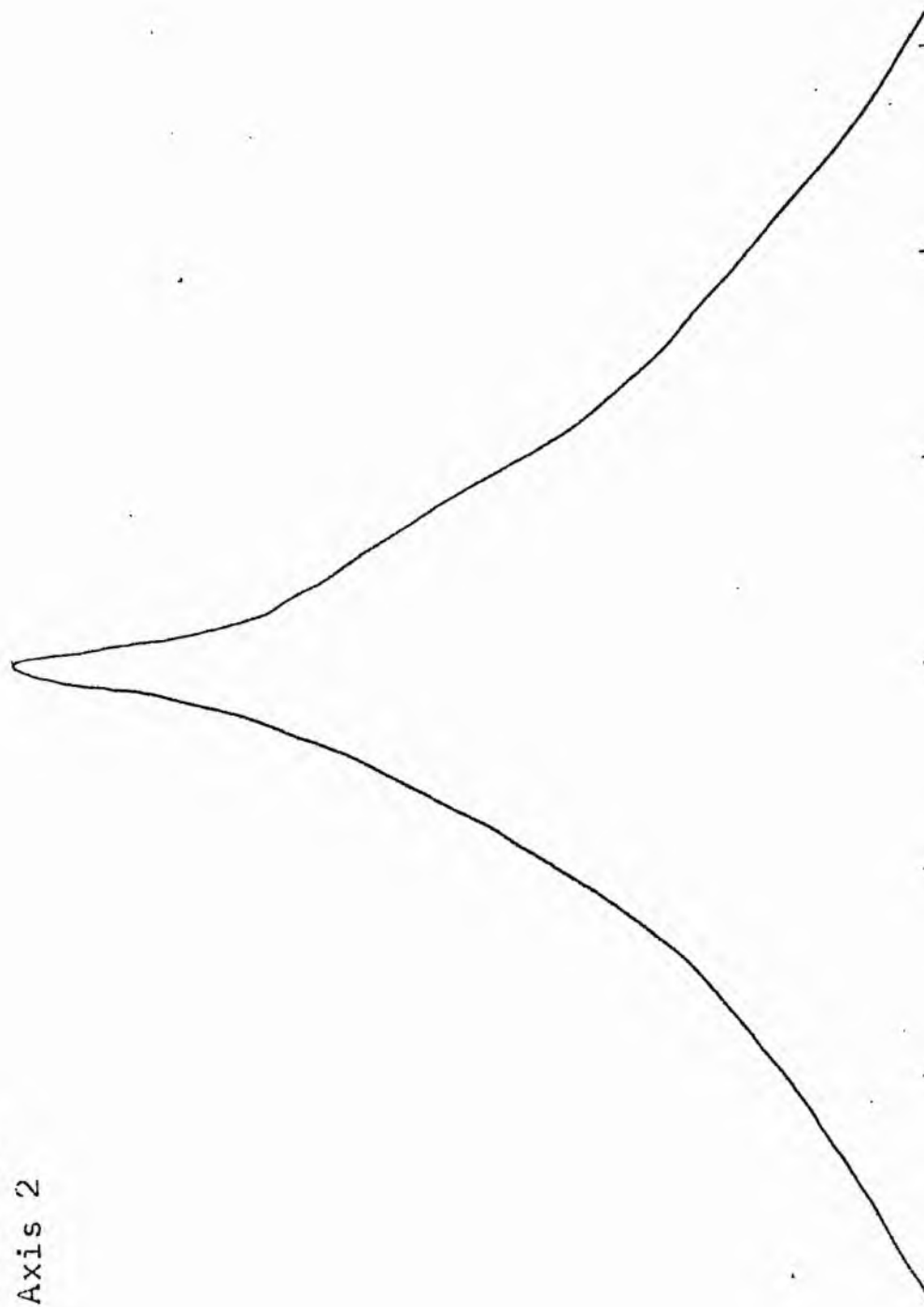
Semi-major axis	a_3	= 1.24
Axis ratio	b/a	= 0.89
Equivalent radius	r_3^*	= 1.16
Surface brightness	μ_3	= 24.32

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.85 \\ = 1.90 \end{matrix}$
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NGC 4567
V-Filter
Axis 1

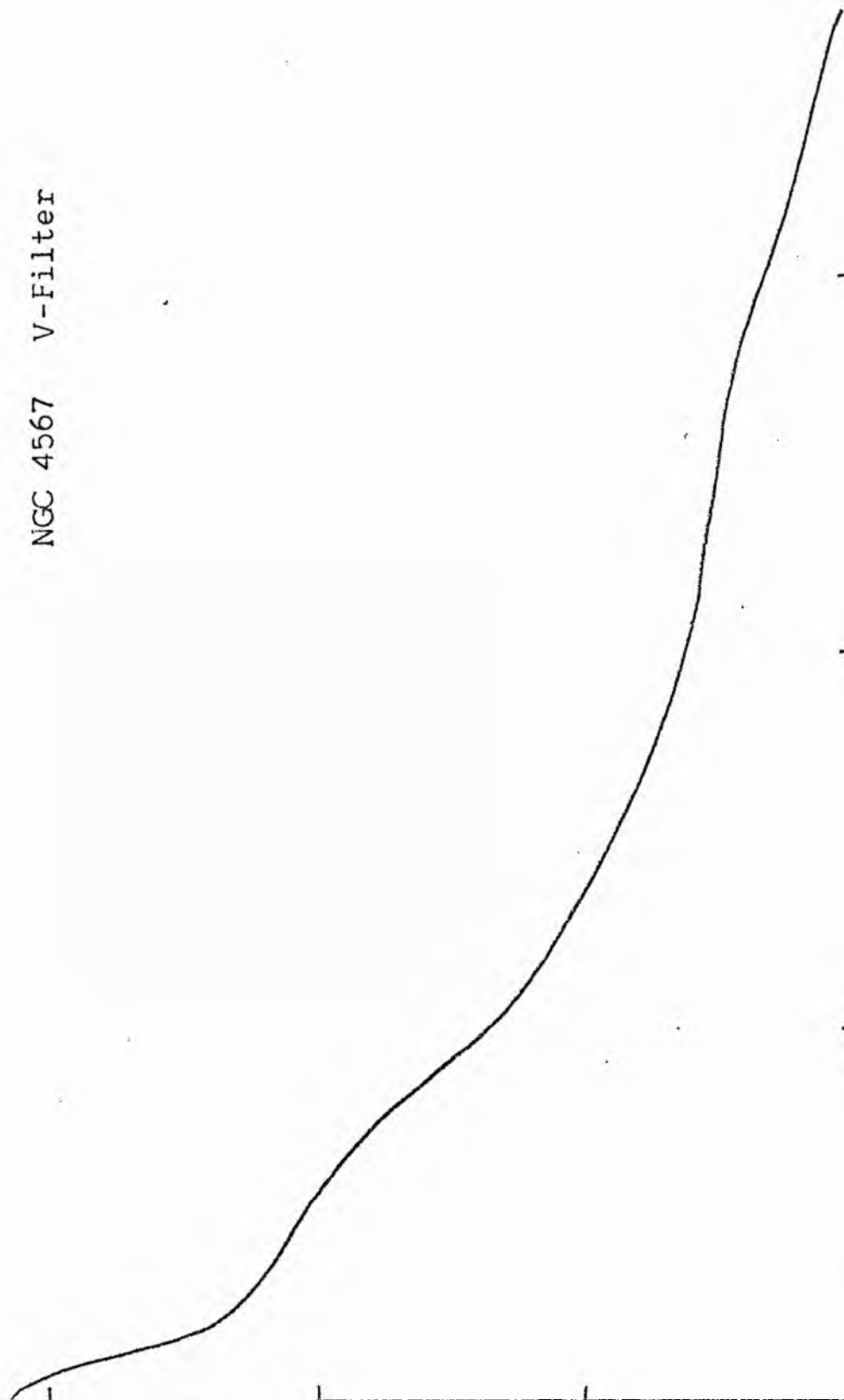


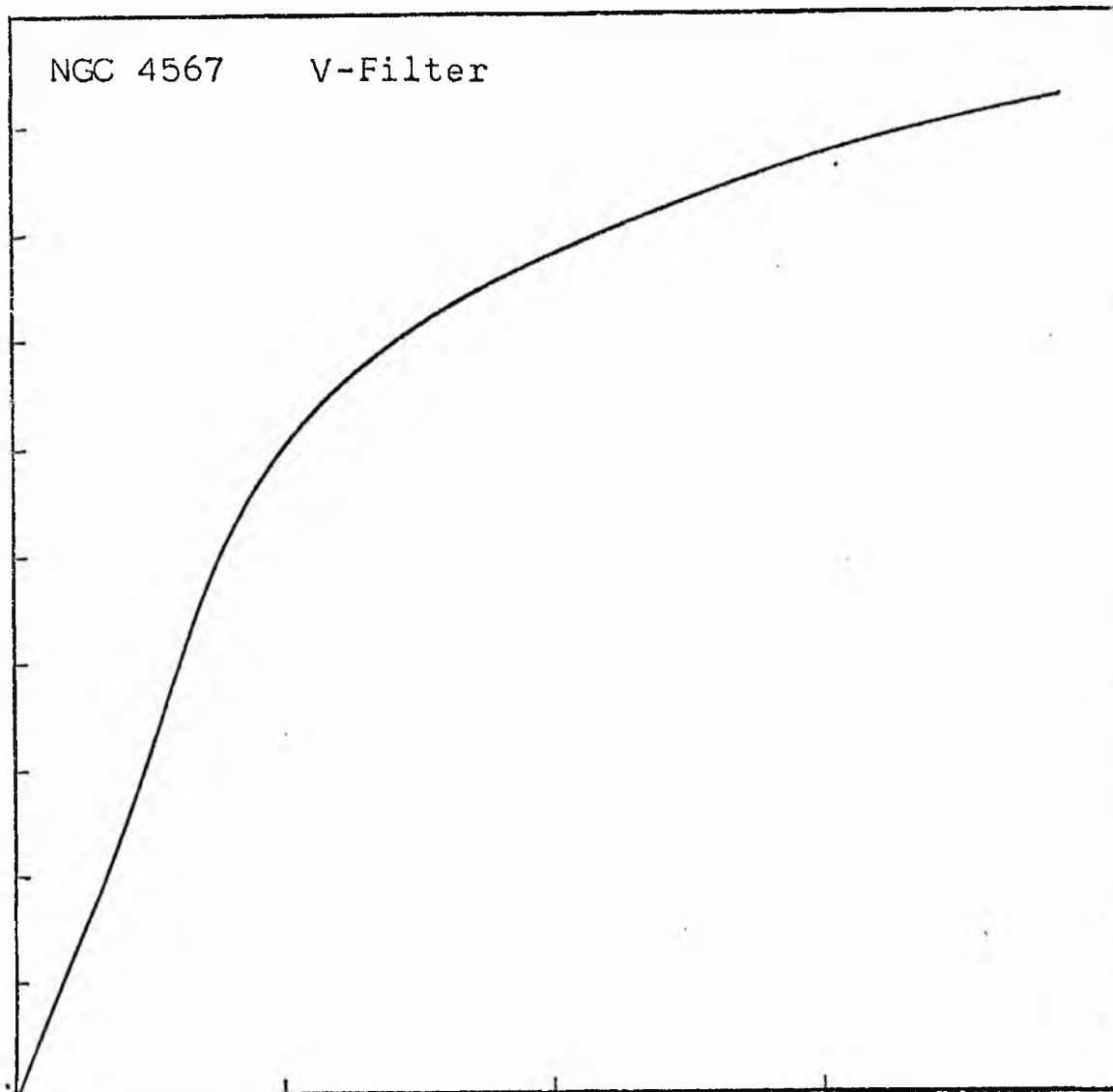
NGC 4567
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4567 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4567
V COLOUR

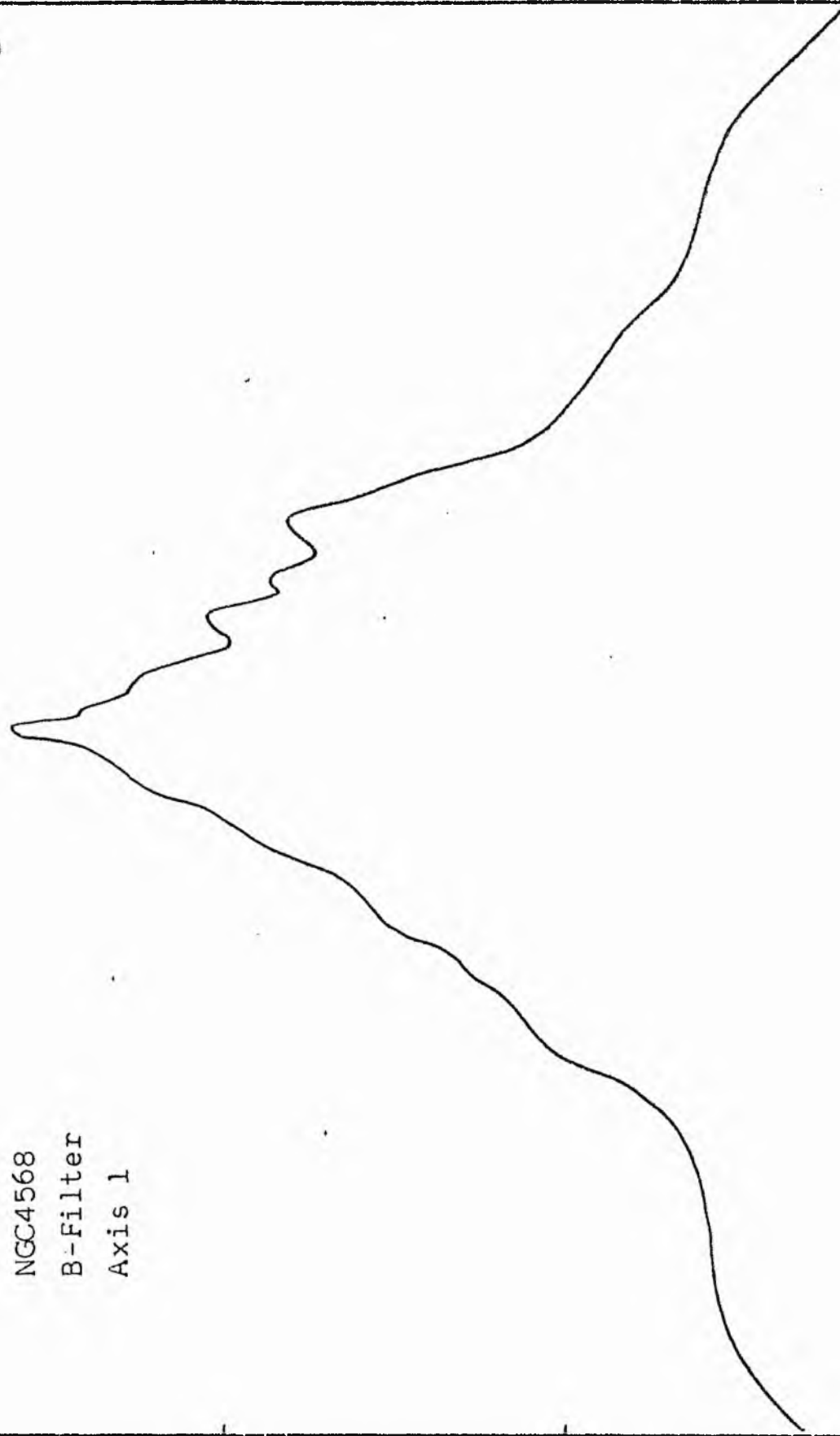
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.14	13.804		0.0	0.0			0.0	0.0	0.0	1.340	18.55
		13.197			16.62	219.3131					
1.10	12.589		2.30	16.62			219.31	0.02	0.06	1.300	18.65
		11.295			26.39	298.0576					
1.00	10.000		3.70	43.01			517.37	0.04	0.10	1.200	18.90
		8.972			12.41	111.3312					
0.90	7.943		4.20	55.42			628.70	0.05	0.11	1.100	19.15
		7.126			20.01	142.6137					
0.80	6.310		4.90	75.43			771.32	0.06	0.13	1.000	19.40
		5.661			53.25	371.4329					
0.70	5.012		6.40	128.68			1072.75	0.09	0.17	0.900	19.65
		4.496			34.18	153.6917					
0.60	3.981		7.20	162.86			1226.44	0.10	0.20	0.800	19.90
		3.572			38.20	136.4441					
0.50	3.162		8.00	211.06			1362.88	0.11	0.22	0.700	20.15
		2.837			88.47	250.9881					
0.40	2.512		9.60	289.53			1613.87	0.13	0.26	0.600	20.40
		2.254			233.26	525.6753					
0.30	1.995		12.90	522.79			2139.55	0.17	0.35	0.500	20.65
		1.790			439.32	786.4172					
0.20	1.585		17.50	962.11			2925.96	0.23	0.48	0.400	20.90
		1.422			772.83	1098.8958					
0.10	1.259		23.50	1734.94			4024.86	0.32	0.64	0.300	21.15
		1.129			198.94	902.3687					
-0.00	1.000		28.40	2533.88			4927.23	0.39	0.77	0.200	21.40
		0.897			184.43	703.7578					
-0.10	0.794		32.50	3318.31			5630.98	0.45	0.89	0.100	21.65
		0.713			913.07	650.6931					
-0.20	0.631		36.70	4231.38			6281.68	0.50	1.00	0.000	21.90
		0.566			596.12	337.4451					
-0.30	0.501		39.20	4827.50			6619.12	0.53	1.07	-0.100	22.15
		0.450			793.72	356.8933					
-0.40	0.398		42.30	5621.21			6976.01	0.56	1.15	-0.200	22.40
		0.357			939.97	335.7263					
-0.50	0.316		45.70	6561.18			7311.73	0.58	1.25	-0.300	22.65
		0.284			767.80	217.8315					
-0.60	0.251		48.30	7328.98			7529.56	0.60	1.32	-0.400	22.90
		0.225			1696.68	382.3594					
-0.70	0.200		53.60	9025.66			7911.92	0.63	1.46	-0.500	23.15
		0.179			1652.26	295.7668					
-0.80	0.158		58.30	10677.92			8207.68	0.65	1.59	-0.600	23.40
		0.142			1949.89	1277.2573					
-0.90	0.126		63.40	12627.82			8484.94	0.68	1.73	-0.700	23.65
		0.113			1941.65	219.3025					
-1.00	0.100		68.10	14569.47			8704.24	0.69	1.86	-0.800	23.90
		0.090			3719.90	333.7366					
-1.10	0.079		76.30	18289.37			9037.97	0.72	2.08	-0.900	24.15
		0.071			3562.16	253.8554					
-1.20	0.063		83.40	21851.53			9291.82	0.74	2.27	-1.000	24.40
		0.057			6620.85	374.7888					
-1.30	0.050		95.20	28472.38			9666.61	0.77	2.59	-1.100	24.65
		0.045			9748.51	438.3406					
-1.40	0.040		110.30	38220.89			10104.95	0.80	3.00	-1.200	24.90
		0.036			12210.75	436.1306					
-1.50	0.032		126.70	50431.63			10541.08	0.84	3.45	-1.300	25.15
		0.028			12026.37	341.2004					
-1.60	0.025		141.00	62458.01			10882.28	0.87	3.84	-1.400	25.40
		0.023			11661.43	262.8013					
-1.70	0.020		153.60	74119.44			11145.08	0.89	4.18	-1.500	25.65
		0.018			11099.62	198.6940					
-1.80	0.016		164.70	85219.06			11343.77	0.90	4.49	-1.600	25.90
		0.014			12426.87	176.7009					
-1.90	0.013		176.30	97645.94			11520.47	0.92	4.80	-1.700	26.15
		0.011			14574.87	164.6198					
-2.00	0.010		189.00	112220.81			11685.09	0.93	5.15	-1.800	26.40
-∞							12565.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4567

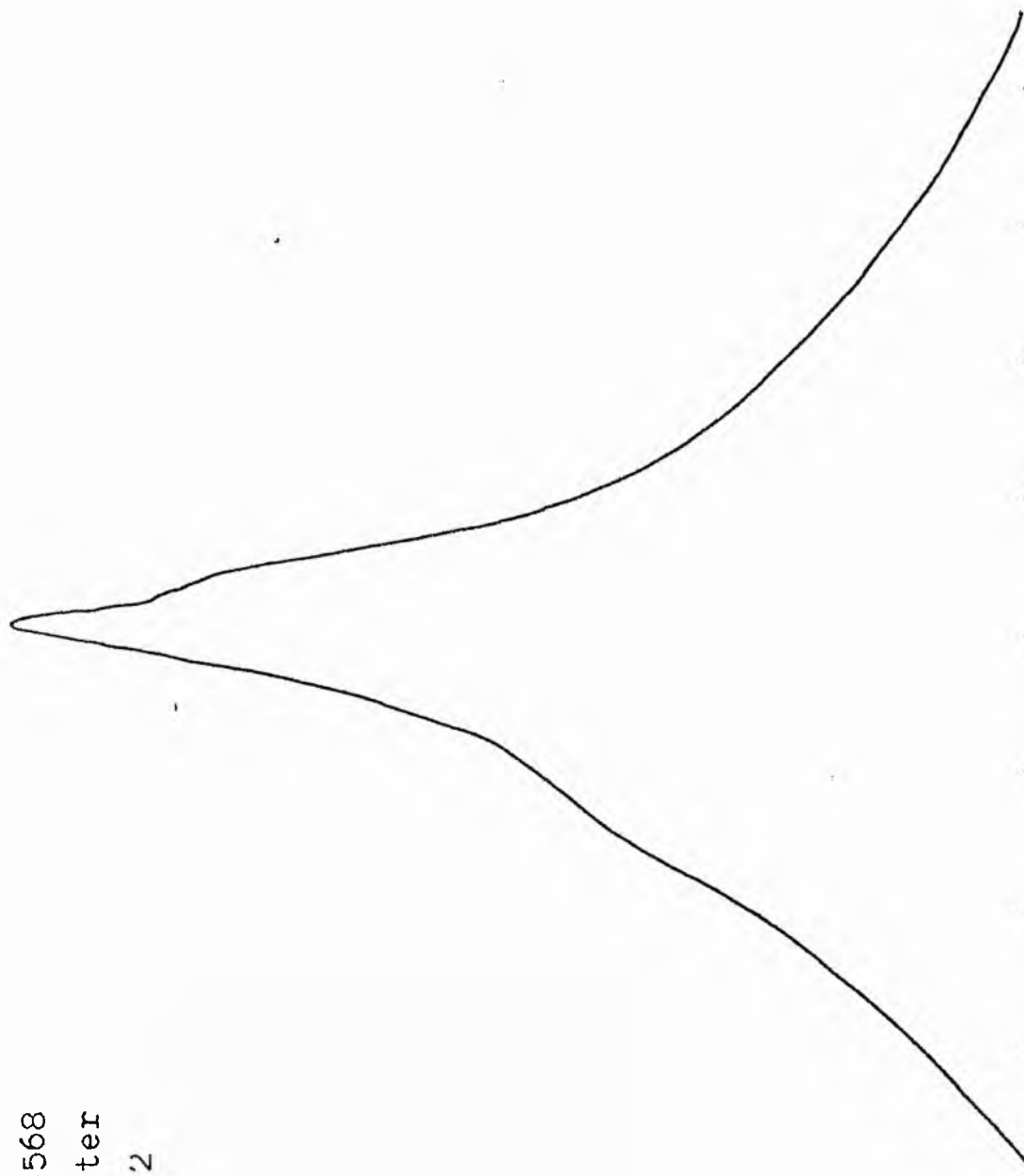
V-FILTER

Total luminosity	L_T	= 3.49
Total apparent magnitude	m_T	= 11.15
Apparent central surface brightness	μ_o	= 18.55
Major axis at threshold	$2a_m$	= 6.43
Minor axis at threshold	$2b_m$	= 5.28
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.30
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.81
Gradient of exponential component	$G(a)$	= -0.76
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.91
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.91
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.31
Axis ratio	b/a	= 0.89
Equivalent radius	r_1^*	= 0.32
Surface brightness	μ_1	= 20.96
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.82
Axis ratio	b/a	= 0.76
Equivalent radius	r_e^*	= 0.62
Surface brightness	μ_e	= 21.90
Mean surface brightness	μ_e'	= 12.11
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.77
Axis ratio	b/a	= 0.76
Equivalent radius	r_3^*	= 1.46
Surface brightness	μ_3	= 24.48
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.96 \\ = 2.38 \end{cases}$

NGC4568
B-Filter
Axis 1

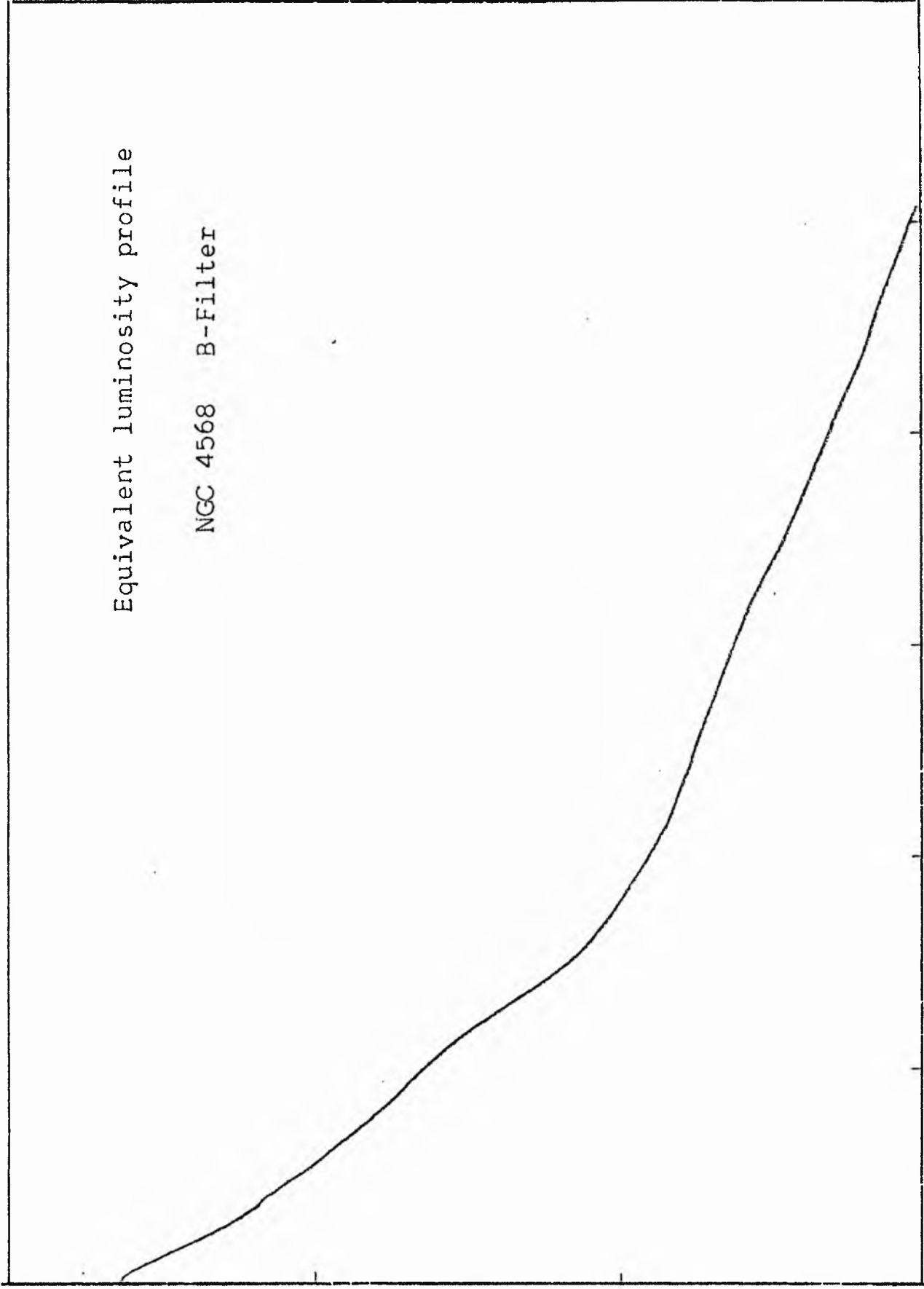


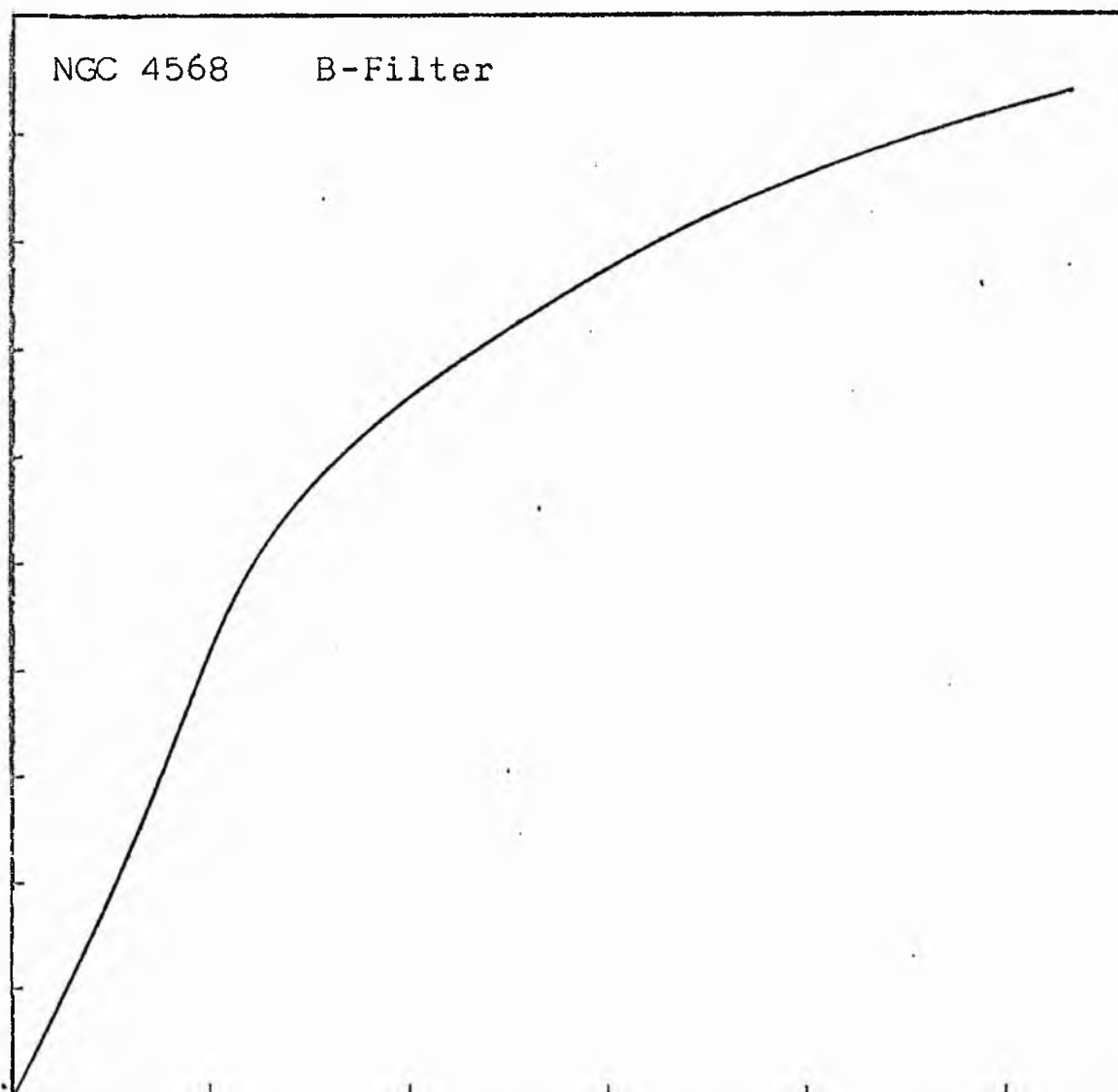
NGC 4568
B-Filter
Axis '2



Equivalent luminosity profile

NGC 4568 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4368
& COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.61	4.074	4.027	0.0	0.0	14.66	59.0318	0.0	0.0	0.0	1.134	20.50
0.60	3.981	3.572	2.16	14.66	57.72	206.1741	59.03	0.00	0.04	1.124	20.53
0.50	3.162	2.837	4.80	72.30	165.40	469.2668	265.21	0.02	0.08	1.024	20.78
0.40	2.512	2.254	8.70	237.79	45.74	103.0817	734.47	0.04	0.14	0.924	21.03
0.30	1.995	1.790	9.50	283.53	551.16	986.6184	837.55	0.05	0.16	0.824	21.28
0.20	1.585	1.422	16.30	834.69	561.33	798.1575	1024.17	0.11	0.21	0.724	21.53
0.10	1.259	1.129	21.08	1396.02	911.20	1029.1650	2622.33	0.15	0.35	0.624	21.78
0.00	1.000	0.897	27.10	2307.22	956.18	857.8542	3651.50	0.21	0.45	0.524	22.03
-0.10	0.794	0.713	32.23	3263.40	1512.51	1077.8770	4509.35	0.26	0.53	0.424	22.28
-0.20	0.631	0.566	38.99	4775.91	2315.29	1310.6235	5507.22	0.32	0.64	0.324	22.53
-0.30	0.501	0.450	47.51	7091.20	2113.84	950.4027	6097.84	0.40	0.78	0.224	22.78
-0.40	0.398	0.357	54.13	9205.04	1954.39	698.0466	7848.32	0.45	0.89	0.124	23.03
-0.50	0.316	0.284	59.60	11159.43	1628.23	461.9417	8546.37	0.49	0.98	0.024	23.28
-0.60	0.251	0.225	63.80	12787.66	1399.26	315.3335	9008.31	0.52	1.05	-0.076	23.53
-0.70	0.200	0.179	67.20	14186.92	1783.96	319.3425	9323.64	0.54	1.11	-0.176	23.78
-0.80	0.158	0.142	71.30	15970.87	2947.04	419.0420	9642.98	0.56	1.17	-0.276	24.03
-0.90	0.126	0.113	77.60	18917.91	5079.96	573.7629	10062.02	0.58	1.28	-0.376	24.28
-1.00	0.100	0.090	87.40	23997.87	6482.65	581.6011	10635.78	0.61	1.44	-0.476	24.53
-1.10	0.079	0.071	98.50	30480.52	9139.00	651.2859	11217.38	0.65	1.62	-0.576	24.78
-1.20	0.063	0.057	112.30	39619.52	16959.42	960.0286	11868.66	0.68	1.85	-0.676	25.03
-1.30	0.050	0.045	134.20	56578.95	14768.12	664.0469	12028.69	0.74	2.21	-0.776	25.28
-1.40	0.040	0.036	150.70	71347.06	15535.81	554.8918	13492.74	0.78	2.48	-0.876	25.53
-1.50	0.032	0.028	166.30	86882.87	16494.19	467.9570	14047.63	0.81	2.74	-0.976	25.78
-1.60	0.025	0.023	181.40	103377.06	22286.62	502.2500	14515.59	0.84	2.99	-1.076	26.03
-1.70	0.020	0.018	200.00	125663.69	23363.50	418.2290	15017.84	0.87	3.30	-1.176	26.28
-1.80	0.016	0.014	217.80	149027.19	25058.37	356.3113	15436.06	0.89	3.59	-1.276	26.53
-1.90	0.013	0.011	235.40	174085.56	22263.94	251.4660	15792.37	0.91	3.88	-1.376	26.78
-2.00	0.010		250.00	196349.50			16043.84	0.93	4.12	-1.476	27.03
-∞							17343.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4568

B-FILTER

Total luminosity	L_T	= 4.82
Total apparent magnitude	m_T	= 11.50
Apparent central surface brightness	μ_o	= 20.50
Major axis at threshold	$2a_m$	= 11.02
Minor axis at threshold	$2b_m$	= 7.17
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 4.77
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.67
Gradient of exponential component	$G(a)$	= -0.41
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.38
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -1.05

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.63
Axis ratio	b/a	= 0.53
Equivalent radius	r_1^*	= 0.53
Surface brightness	μ_1	= 22.23

Parameters at $k = \frac{1}{2}$ (effective) :

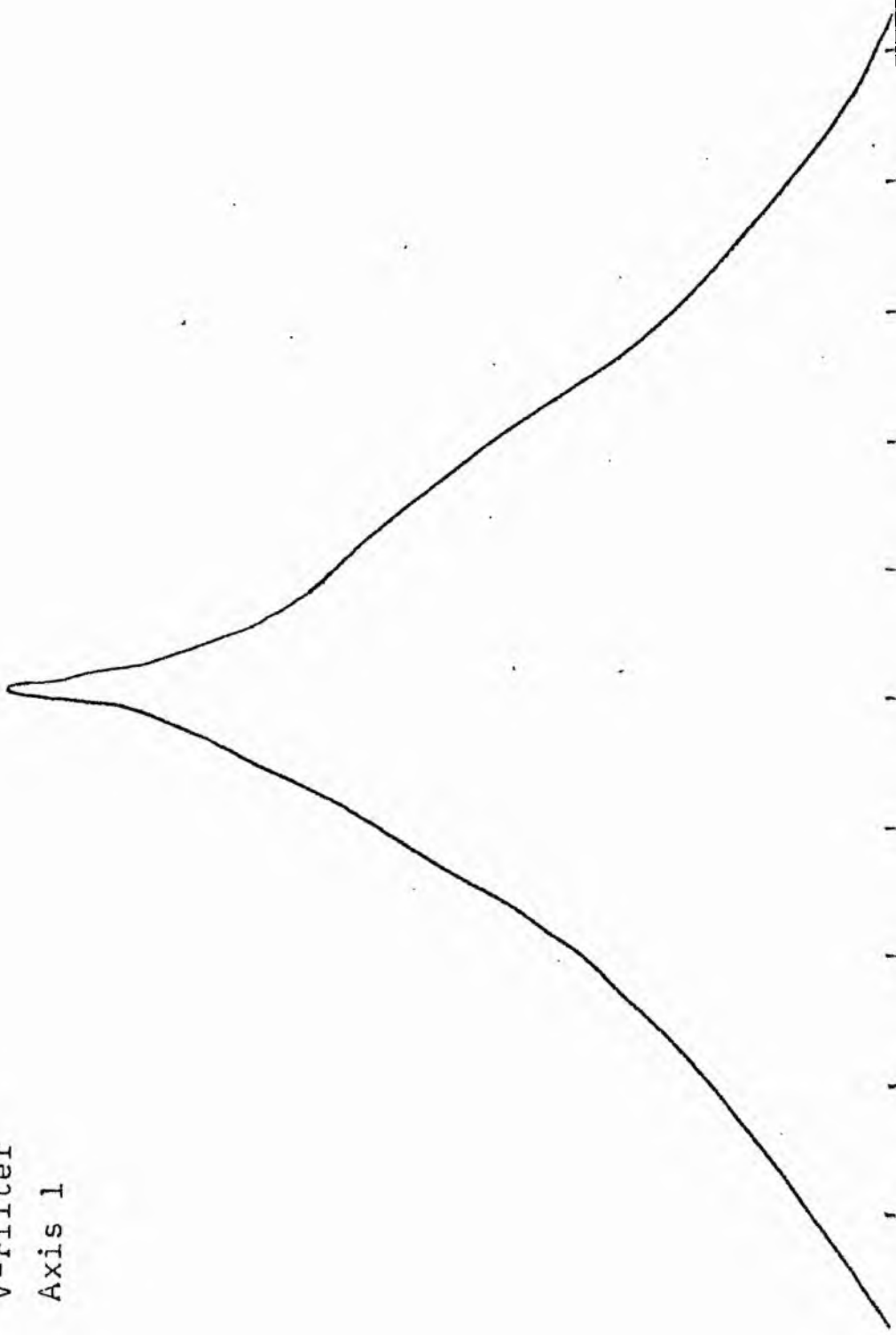
Semi-major axis	a_e	= 1.59
Axis ratio	b/a	= 0.38
Equivalent radius	r_e^*	= 1.01
Surface brightness	μ_e	= 23.36
Mean surface brightness	μ_e'	= 13.52

Parameters at $k = \frac{3}{4}$:

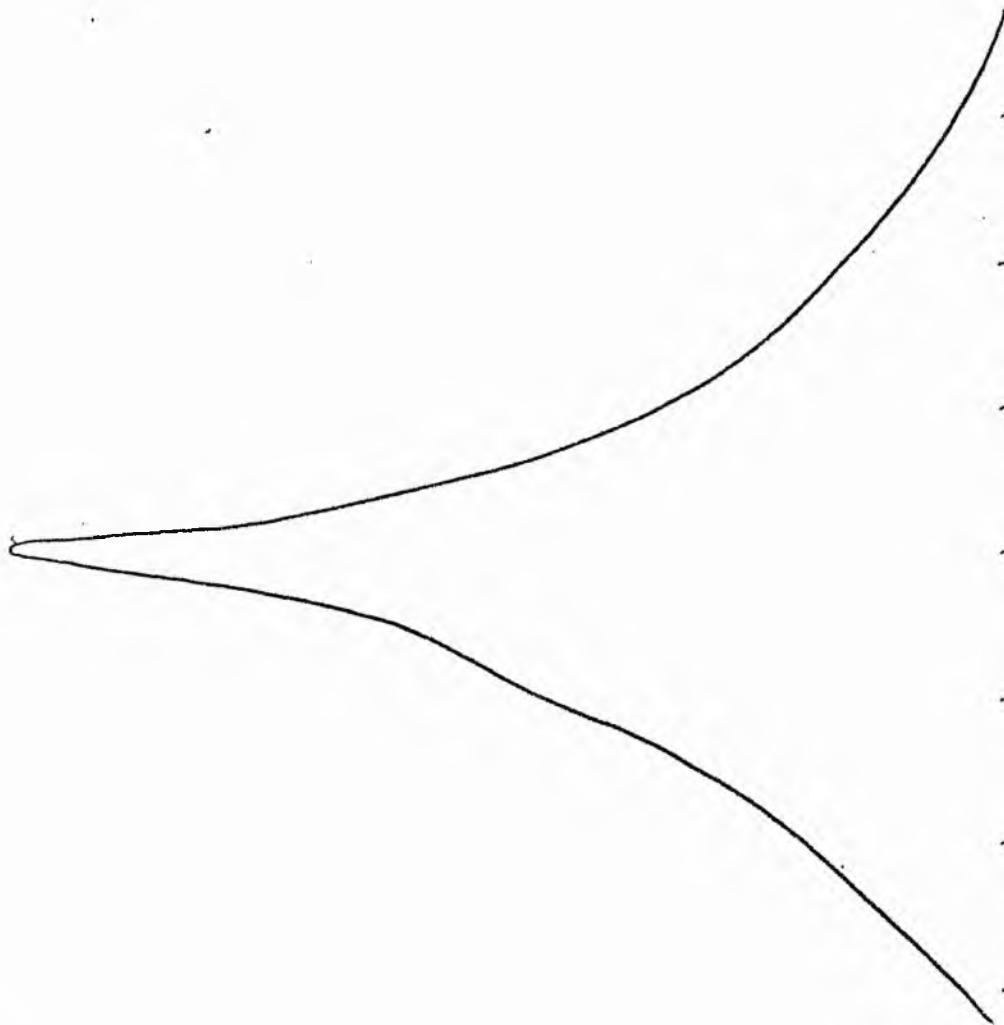
Semi-major axis	a_3	= 2.74
Axis ratio	b/a	= 0.53
Equivalent radius	r_3^*	= 2.31
Surface brightness	μ_3	= 25.34

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.95 \\ = 2.28 \end{matrix}$
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NGC 4568
V-Filter
Axis 1

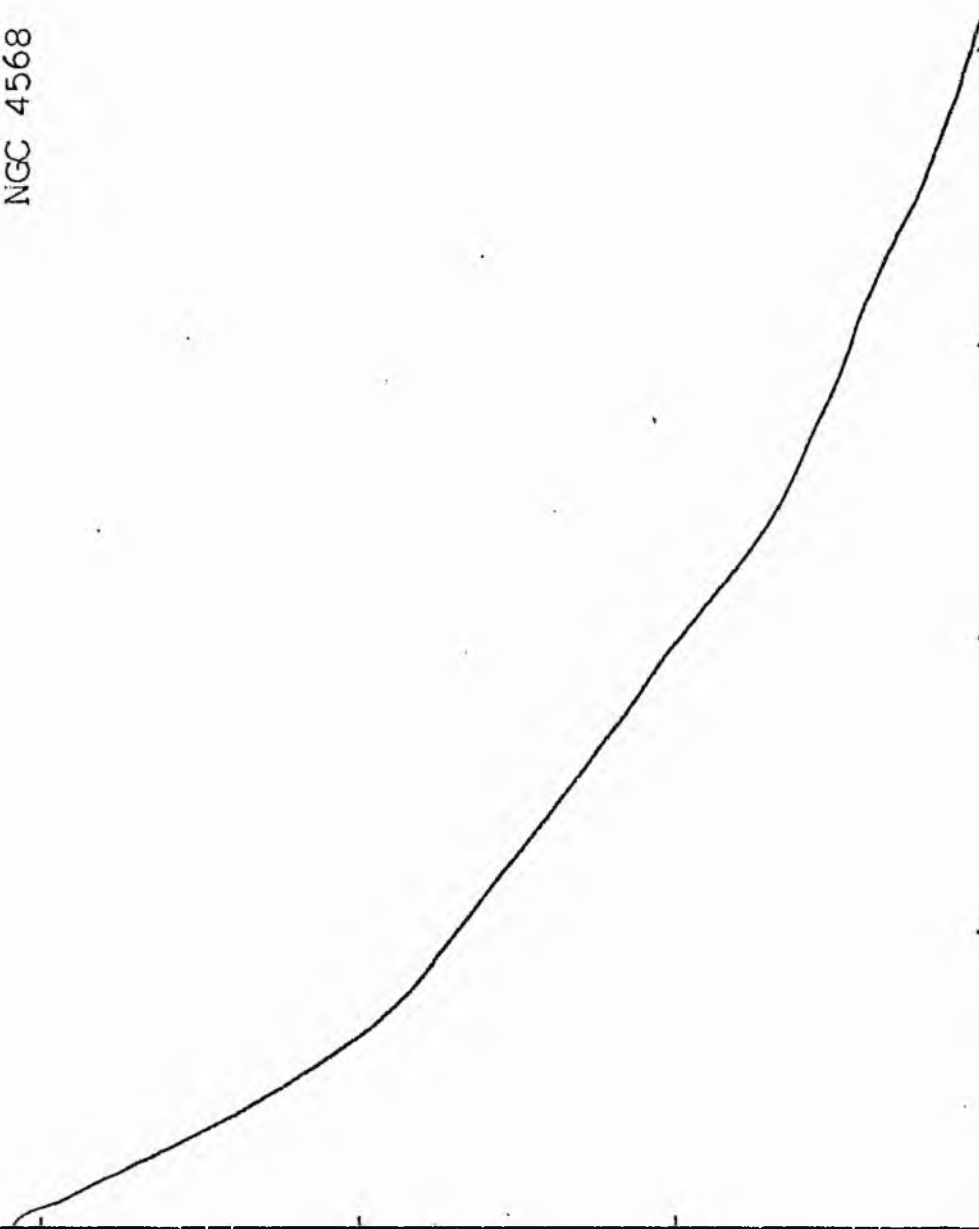


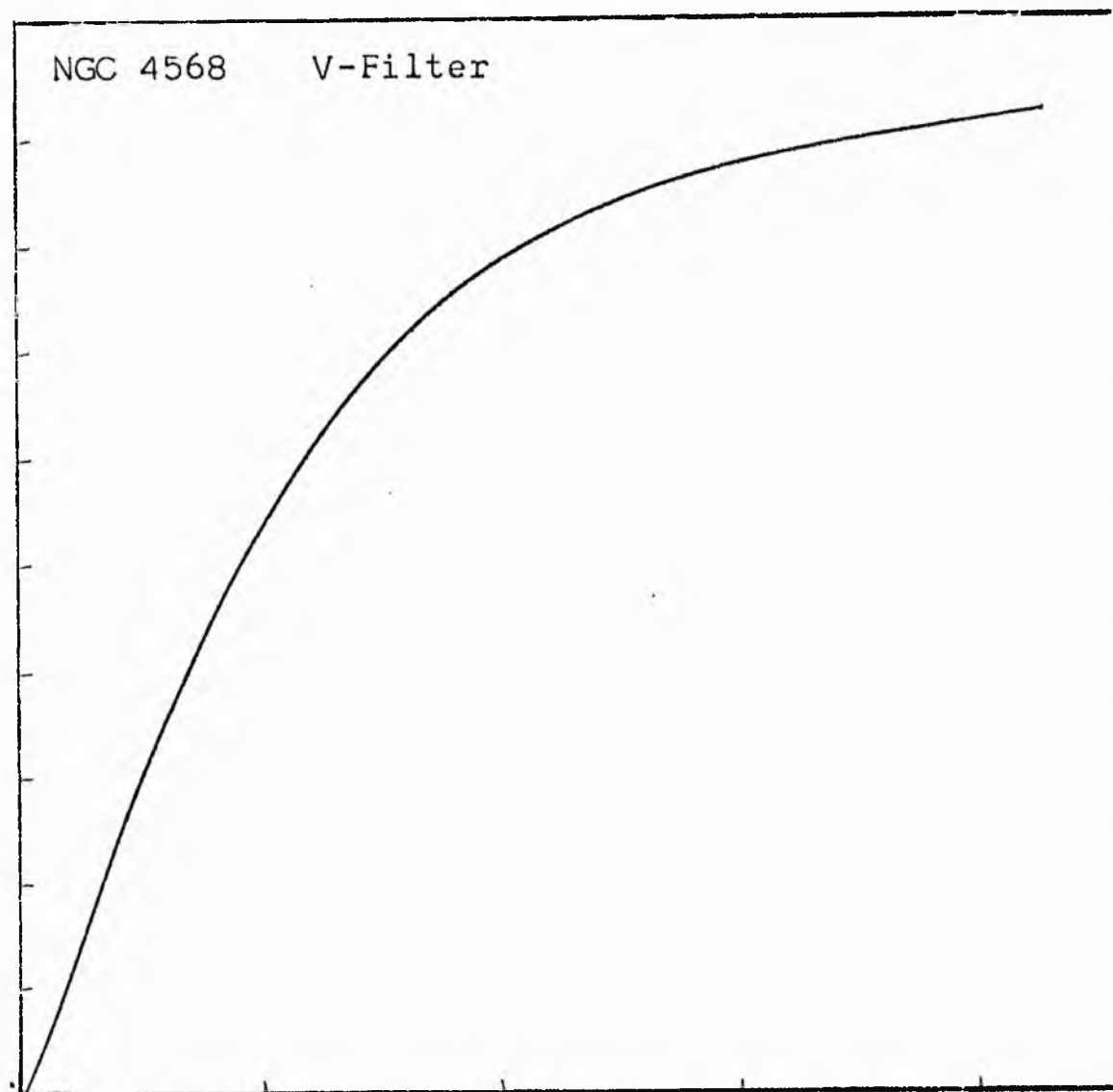
NGC 4568
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4568 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4568
V COLOUR

LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.09	12.303		0.0	0.0			0.0	0.0	0.0	1.314	18.78
1.00	10.000	11.151	3.30	34.21	34.21	301.5081	301.51	0.02	0.07	1.224	19.01
0.90	7.943	8.972	4.20	55.42	21.21	190.2499	571.76	0.03	0.09	1.124	19.26
0.80	6.310	7.126	7.40	172.03	116.62	831.0547	1432.81	0.07	0.16	1.024	19.51
0.70	5.012	5.661	8.60	232.35	60.32	341.4465	1744.26	0.08	0.19	0.924	19.76
0.60	3.981	4.496	11.50	415.48	183.12	823.4092	2567.67	0.12	0.25	0.824	20.01
0.50	3.162	3.572	14.00	615.75	200.28	715.3228	3282.99	0.16	0.30	0.724	20.26
0.40	2.512	2.837	18.01	1019.01	403.25	1144.0659	4427.05	0.21	0.39	0.624	20.51
0.30	1.995	2.254	20.33	1298.45	279.44	629.7388	5056.79	0.25	0.44	0.524	20.76
0.20	1.585	1.790	23.85	1787.01	488.56	874.5615	5931.35	0.29	0.52	0.424	21.01
0.10	1.259	1.422	27.70	2410.51	623.51	886.5681	6817.92	0.33	0.60	0.324	21.26
0.00	1.000	1.129	30.89	2997.68	587.17	663.1870	7481.10	0.36	0.67	0.224	21.51
-0.10	0.794	0.897	36.60	4208.35	1210.67	1086.1663	8567.27	0.42	0.79	0.124	21.76
-0.20	0.631	0.713	44.47	6212.75	2004.40	1428.4236	9995.69	0.48	0.96	0.024	22.01
-0.30	0.501	0.566	51.70	8397.13	2184.38	1236.5166	11232.20	0.54	1.12	-0.076	22.26
-0.40	0.398	0.450	58.35	10696.25	2299.12	1033.7927	12265.99	0.59	1.26	-0.176	22.51
-0.50	0.316	0.357	63.24	12564.16	1867.91	667.1589	12933.15	0.63	1.37	-0.276	22.76
-0.60	0.251	0.284	70.15	15459.84	2895.68	821.5293	13754.68	0.67	1.52	-0.376	23.01
-0.70	0.200	0.225	77.56	18898.41	3438.57	774.9089	14529.58	0.70	1.68	-0.476	23.26
-0.80	0.158	0.179	84.20	22272.75	3374.34	604.0342	15133.61	0.73	1.82	-0.576	23.51
-0.90	0.126	0.142	91.30	26187.34	3914.59	556.6196	15690.23	0.76	1.97	-0.676	23.76
-1.00	0.100	0.113	96.50	29255.30	3067.96	346.5154	16036.74	0.78	2.09	-0.776	24.01
-1.10	0.079	0.090	100.20	31541.71	2286.42	205.1298	16241.87	0.79	2.17	-0.876	24.26
-1.20	0.063	0.071	104.70	34438.41	2896.70	206.4317	16448.30	0.80	2.26	-0.976	24.51
-1.30	0.050	0.057	117.80	43595.37	9156.96	518.3528	16966.65	0.82	2.55	-1.076	24.76
-1.40	0.040	0.045	125.40	49402.04	5806.67	261.0969	17227.75	0.84	2.71	-1.176	25.01
-1.50	0.032	0.036	138.50	60262.82	10860.78	387.9148	17615.66	0.85	2.99	-1.276	25.26
-1.60	0.025	0.028	152.30	72870.12	12607.30	357.6829	17973.34	0.87	3.29	-1.376	25.51
-1.70	0.020	0.023	164.70	85219.06	12348.94	278.2954	18251.63	0.89	3.56	-1.476	25.76
-1.80	0.016	0.018	179.63	101369.50	16150.44	287.1089	18540.74	0.90	3.88	-1.576	26.01
-1.90	0.013	0.014	195.40	119949.62	18580.12	264.1960	18804.93	0.91	4.22	-1.676	26.26
-2.00	0.010	0.011	213.60	143335.00	23385.37	264.1328	19069.07	0.92	4.62	-1.776	26.51
-∞							20619.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4568

V-FILTER

Total luminosity	L_T	= 5.73
Total apparent magnitude	m_T	= 10.72
Apparent central surface brightness	μ_0	= 18.78
Major axis at threshold	$2a_m$	= 8.03
Minor axis at threshold	$2b_m$	= 5.68
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.61
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.84
Gradient of exponential component	$G(a)$	= -0.57
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.60
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.74

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.43
Axis ratio	b/a	= 0.62
Equivalent radius	r_1^*	= 0.35
Surface brightness	μ_1	= 20.76

Parameters at $k = \frac{1}{2}$ (effective) :

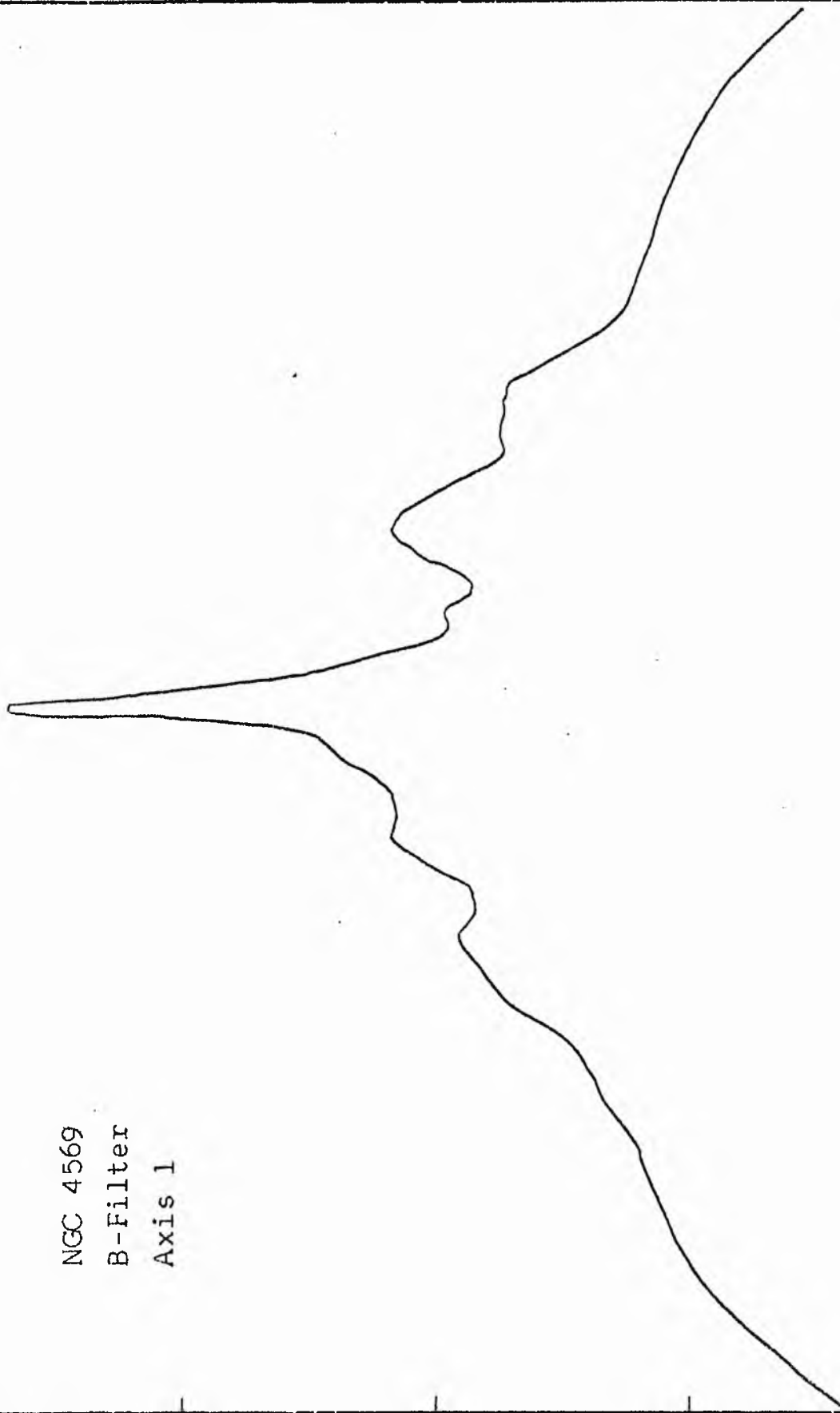
Semi-major axis	a_e	= 0.97
Axis ratio	b/a	= 0.57
Equivalent radius	r_e^*	= 0.77
Surface brightness	μ_e	= 22.09
Mean surface brightness	μ_e'	= 12.15

Parameters at $k = \frac{3}{4}$:

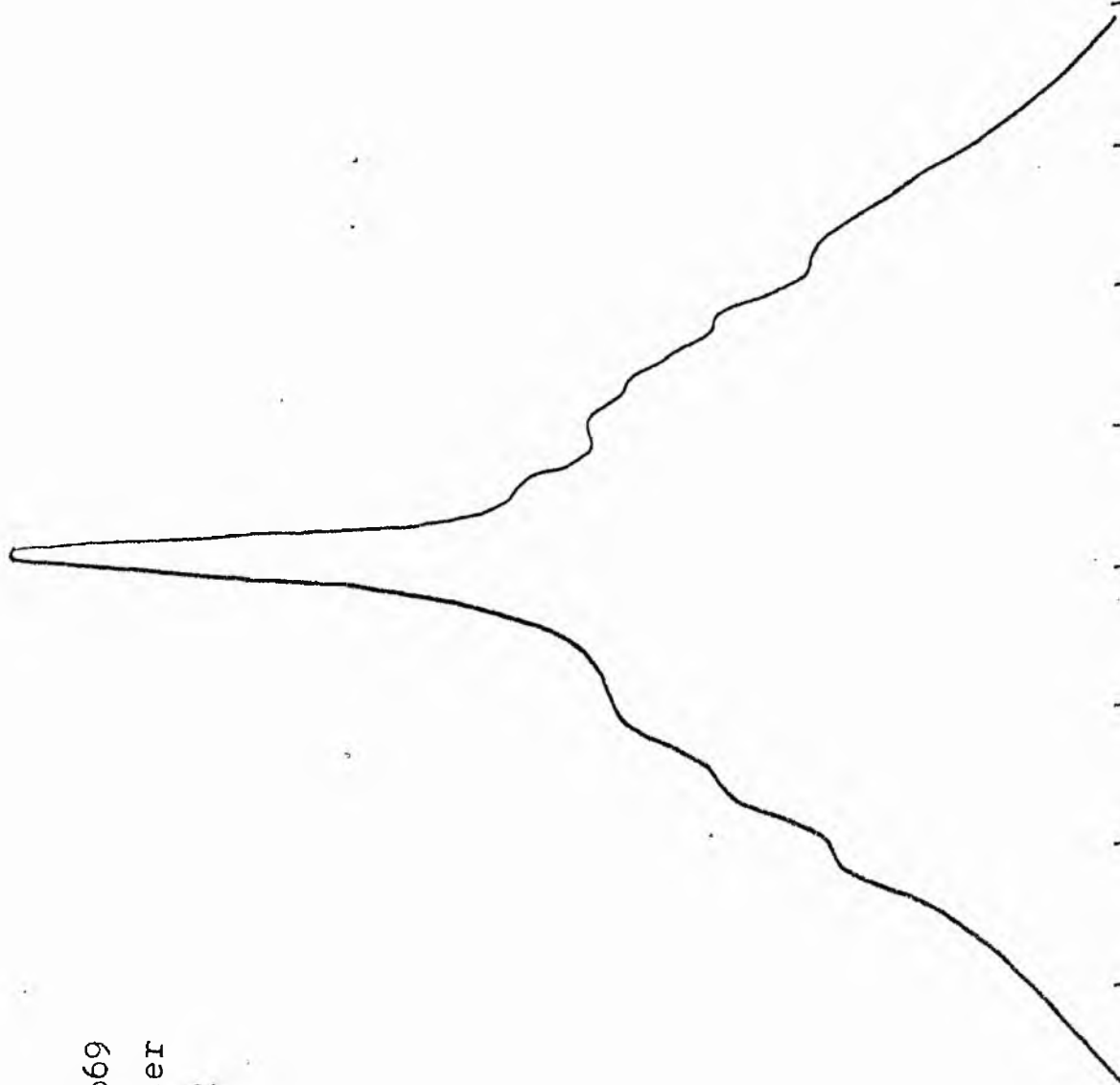
Semi-major axis	a_3	= 1.87
Axis ratio	b/a	= 0.49
Equivalent radius	r_3^*	= 1.47
Surface brightness	μ_3	= 23.68

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.23 \\ = 1.91 \end{cases}$
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NGC 4569
B-Filter
Axis 1

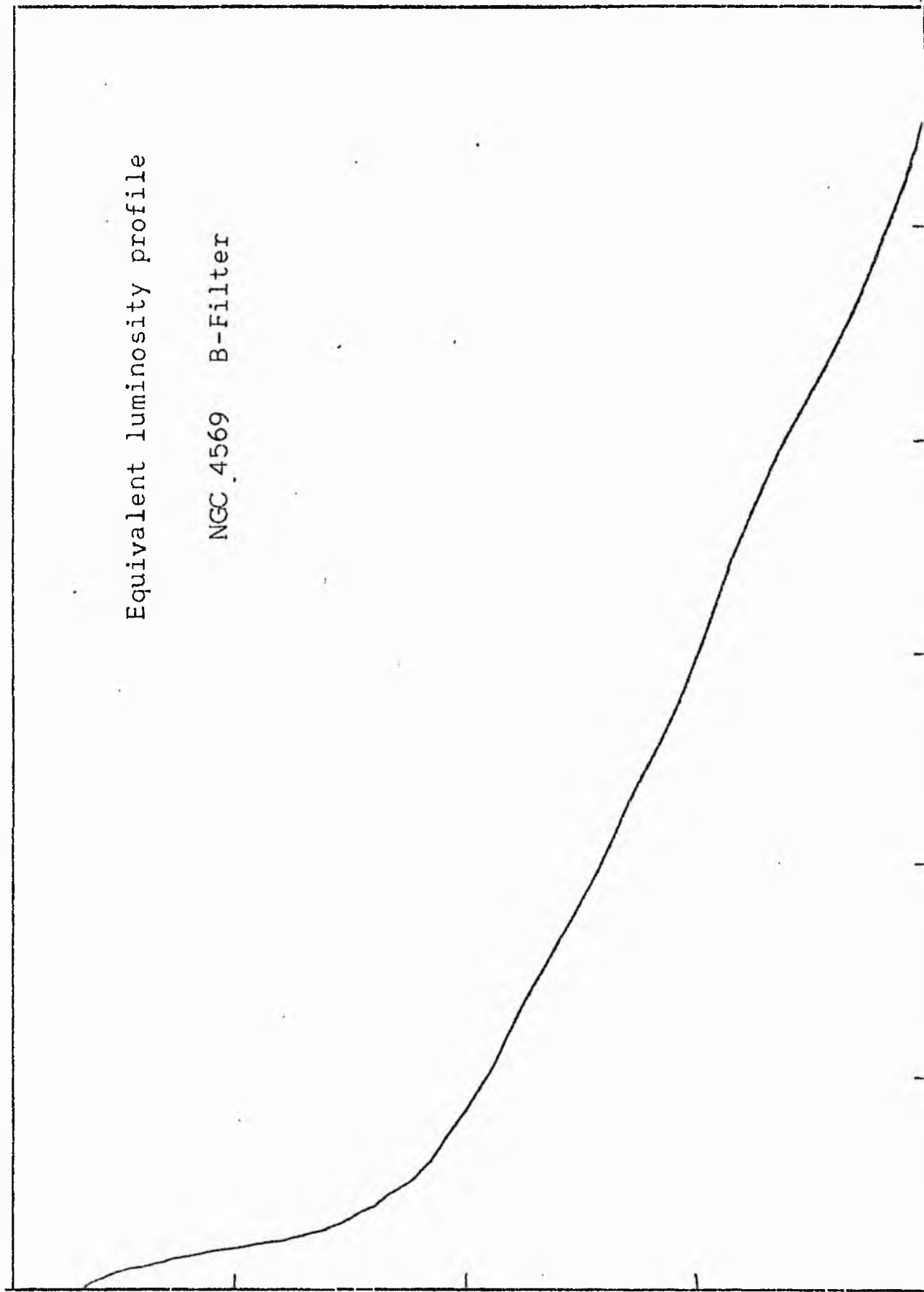


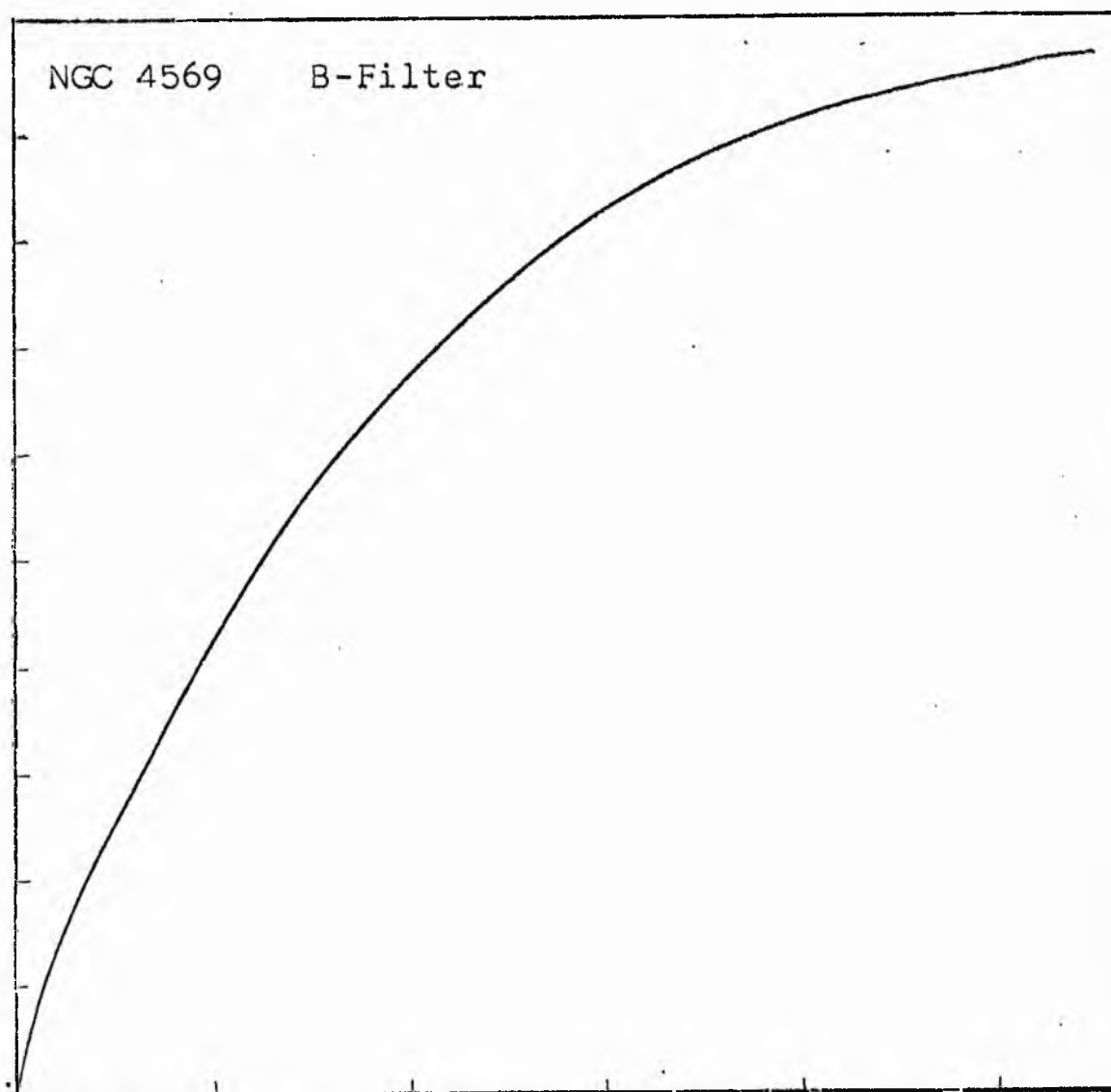
NGC 4569
B-Filter
Axis '2



Equivalent luminosity profile

NGC 4569 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

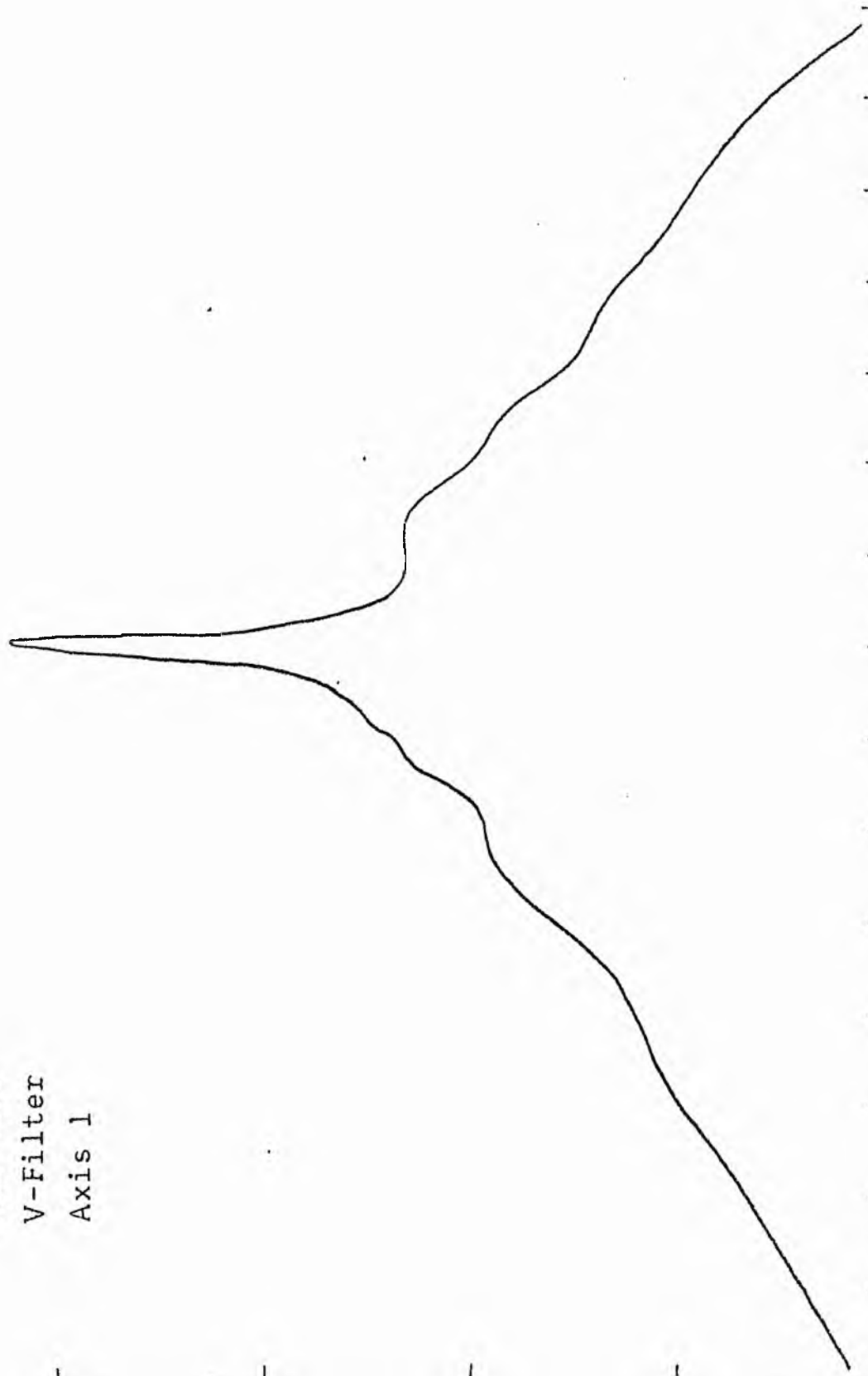
MEAN LUMINOSITY DISTRIBUTION IN MGC 4549
B COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(M)	ρ	LOG J	μ
1.58	38.019		0.0	0.0			0.0	0.0	0.0	1.807	17.73
1.50	31.623	34.821	3.33	34.84	34.84	1213.0464	1213.05	0.03	0.05	1.727	17.93
1.40	25.119	28.371	4.20	55.42	20.58	583.8950	1746.94	0.05	0.07	1.627	18.18
1.30	19.953	22.536	6.00	113.10	57.68	1299.8516	3096.79	0.08	0.09	1.527	18.43
1.20	15.849	17.901	6.30	124.69	11.59	207.5130	3304.31	0.08	0.10	1.427	18.68
1.10	12.589	14.219	7.20	162.86	38.17	542.7463	3847.05	0.10	0.11	1.327	18.93
1.00	10.000	11.295	8.00	201.06	38.20	431.4729	4278.52	0.11	0.13	1.227	19.18
0.90	7.943	8.972	8.80	243.28	42.22	378.8071	4657.33	0.12	0.14	1.127	19.43
0.80	6.310	7.126	9.60	289.53	46.24	329.5547	4986.88	0.13	0.15	1.027	19.68
0.70	5.012	5.661	10.70	359.68	70.15	397.1086	5383.99	0.14	0.17	0.927	19.93
0.60	3.981	4.496	13.20	547.39	187.71	844.0305	6228.02	0.16	0.21	0.827	20.18
0.50	3.162	3.572	15.80	784.27	236.88	846.0410	7074.05	0.18	0.25	0.727	20.43
0.40	2.512	2.837	16.83	889.85	105.58	249.5486	7373.60	0.19	0.26	0.627	20.68
0.30	1.995	2.254	20.70	1346.14	456.29	1028.2788	8401.88	0.24	0.32	0.527	20.93
0.20	1.585	1.790	26.90	2273.29	927.15	1659.6597	10061.54	0.25	0.42	0.427	21.18
0.10	1.259	1.422	34.10	3653.07	1379.79	1961.9248	12023.46	0.30	0.53	0.327	21.43
-0.00	1.000	1.129	42.30	5621.21	1968.14	2222.9355	14246.39	0.36	0.66	0.227	21.68
-0.10	0.794	0.897	52.20	8560.34	2939.12	2636.8662	16883.25	0.43	0.82	0.127	21.93
-0.20	0.631	0.713	61.30	11805.12	3244.79	2312.3689	19195.62	0.49	0.96	0.027	22.18
-0.30	0.501	0.566	70.40	15570.22	3765.10	2131.3108	21326.93	0.54	1.10	-0.073	22.43
-0.40	0.398	0.450	77.47	18854.57	3284.35	1476.7959	22803.72	0.58	1.22	-0.173	22.68
-0.50	0.316	0.357	87.18	23877.20	5022.62	1793.9126	24597.63	0.62	1.37	-0.273	22.93
-0.60	0.251	0.284	97.66	29962.86	6085.66	1726.5481	26324.18	0.67	1.53	-0.373	23.18
-0.70	0.200	0.225	108.92	37270.48	7307.63	1646.8252	27971.00	0.71	1.71	-0.473	23.43
-0.80	0.158	0.179	120.95	45958.04	8687.56	1555.1372	29526.14	0.75	1.90	-0.573	23.68
-0.90	0.126	0.142	136.70	58706.59	12748.55	1812.7246	31338.86	0.79	2.14	-0.673	23.93
-1.00	0.100	0.113	152.70	73253.37	14546.79	1643.0042	32981.87	0.83	2.40	-0.773	24.18
-1.10	0.079	0.090	165.30	85841.12	12587.75	1129.3286	34111.20	0.86	2.59	-0.873	24.43
-1.20	0.063	0.071	179.40	101110.12	15269.00	1080.1367	35199.33	0.89	2.81	-0.973	24.68
-1.30	0.050	0.057	184.20	106593.06	5482.94	310.3748	35509.70	0.90	2.89	-1.073	24.93
-1.40	0.040	0.045	196.20	120933.81	14340.75	644.8303	36154.53	0.91	3.08	-1.173	25.18
-1.50	0.032	0.036	207.10	134744.12	13810.31	493.2622	36647.79	0.93	3.25	-1.273	25.43
-1.60	0.025	0.028	216.80	147661.87	12917.75	366.4897	37014.28	0.94	3.40	-1.373	25.68
-1.70	0.020	0.023	228.00	163312.56	15650.69	352.7029	37366.98	0.94	3.58	-1.473	25.93
-1.80	0.016	0.018	240.00	180955.75	17643.19	315.8298	37682.81	0.95	3.77	-1.573	26.18
-1.90	0.013	0.014	259.40	211392.25	30436.50	432.7842	38115.59	0.96	4.07	-1.673	26.43
-2.00	0.010	0.011	272.70	233625.31	22233.06	251.1173	38366.71	0.97	4.28	-1.773	26.68
-∞							39566.00	(1)			∞

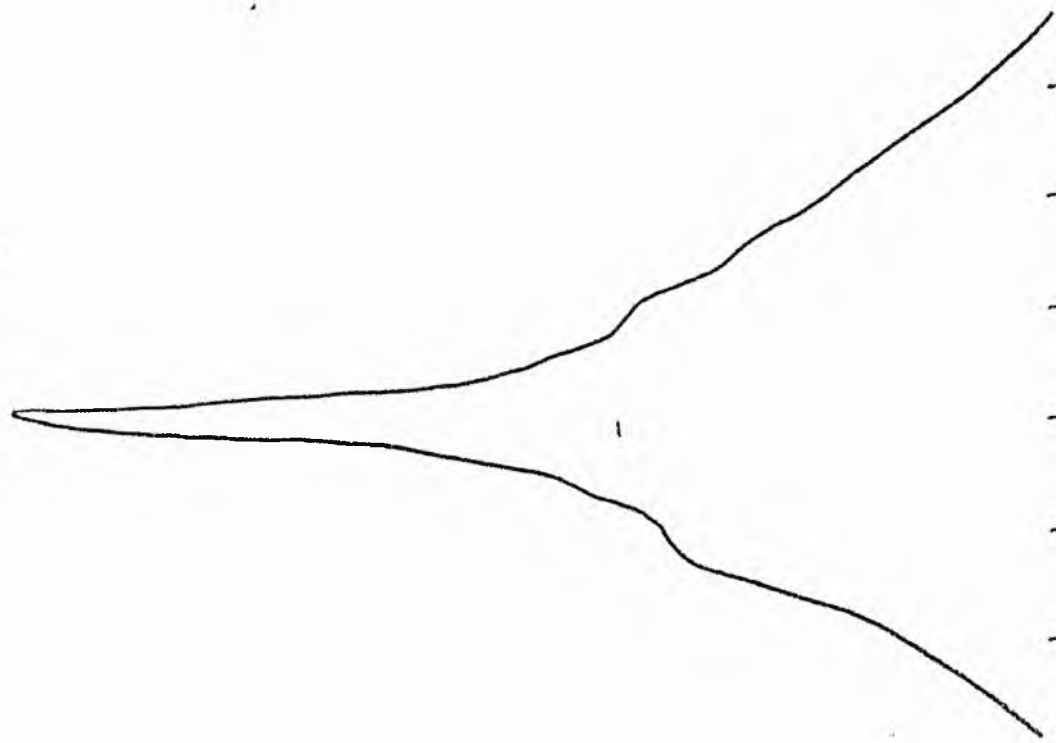
PHOTOMETRIC PARAMETERS OF NGC 4569
B-FILTER

Total luminosity	L_T	=	10.99
Total apparent magnitude	m_T	=	10.19
Apparent central surface brightness	μ_0	=	17.73
Major axis at threshold	$2a_m$	=	12.76
Minor axis at threshold	$2b_m$	=	6.28
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	=	8.89
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	=	0.90
Gradient of exponential component	$G(a)$	=	-0.43
Equivalent gradient of exponential comp....	$G(r^*)$	=	-0.53
Equivalent gradient of reduced exp. comp....	$G(\rho)$	=	-0.70
Parameters at $k = \frac{1}{4}$:			
Semi-major axis	a_1	=	0.57
Axis ratio	b/a	=	0.56
Equivalent radius	r_1^*	=	0.44
Surface brightness	μ_1	=	21.18
Parameters at $k = \frac{1}{2}$ (effective) :			
Semi-major axis	a_e	=	1.73
Axis ratio	b/a	=	0.38
Equivalent radius	r_e^*	=	1.066
Surface brightness	μ_e	=	22.23
Mean surface brightness	μ_e'	=	12.31
Parameters at $k = \frac{3}{4}$:			
Semi-major axis	a_3	=	2.77
Axis ratio	b/a	=	0.55
Equivalent radius	r_3^*	=	2.03
Surface brightness	μ_3	=	23.68
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	=	$\begin{matrix} 2.43 \\ 1.92 \end{matrix}$

NGC 4569
V-Filter
Axis 1

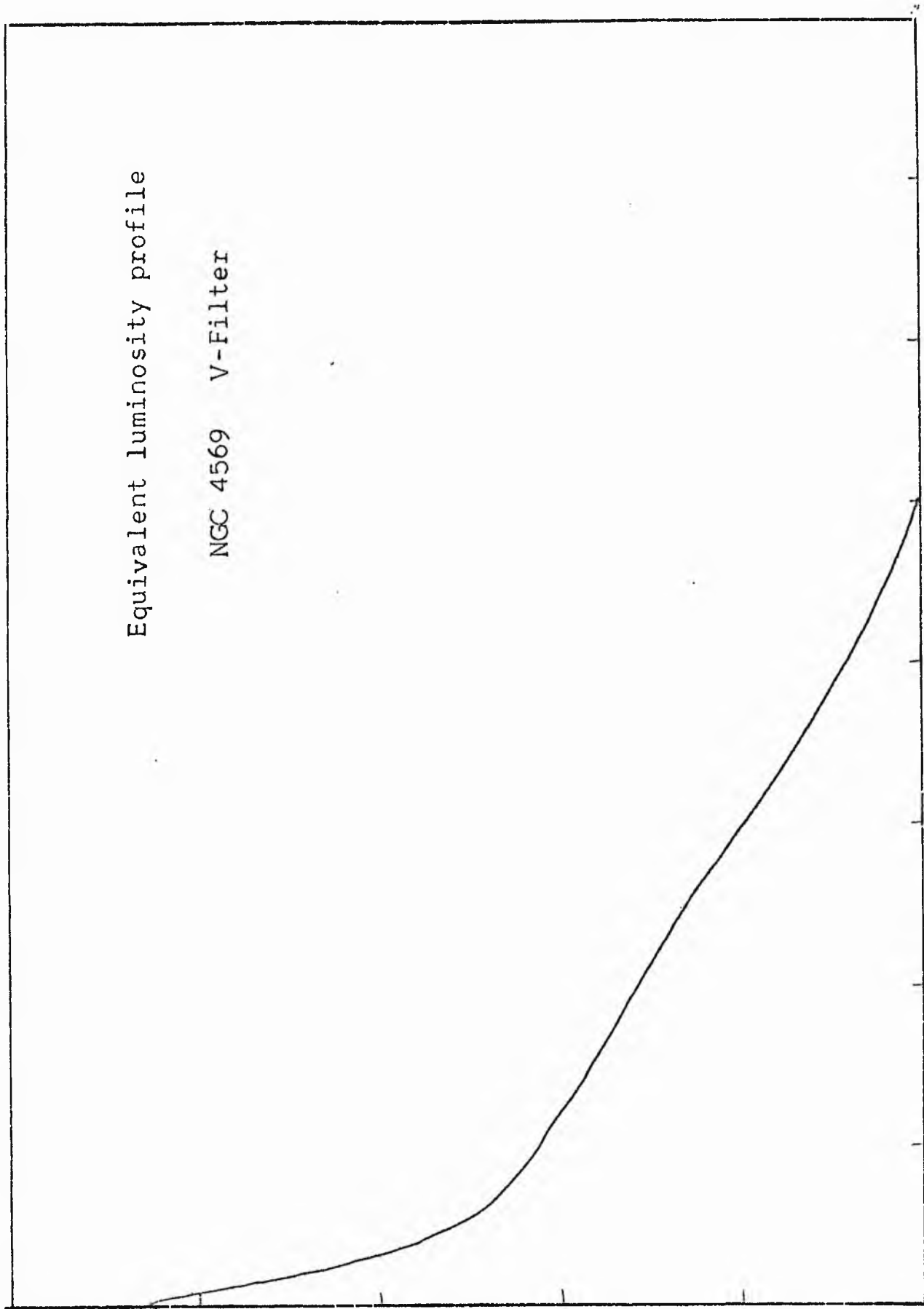


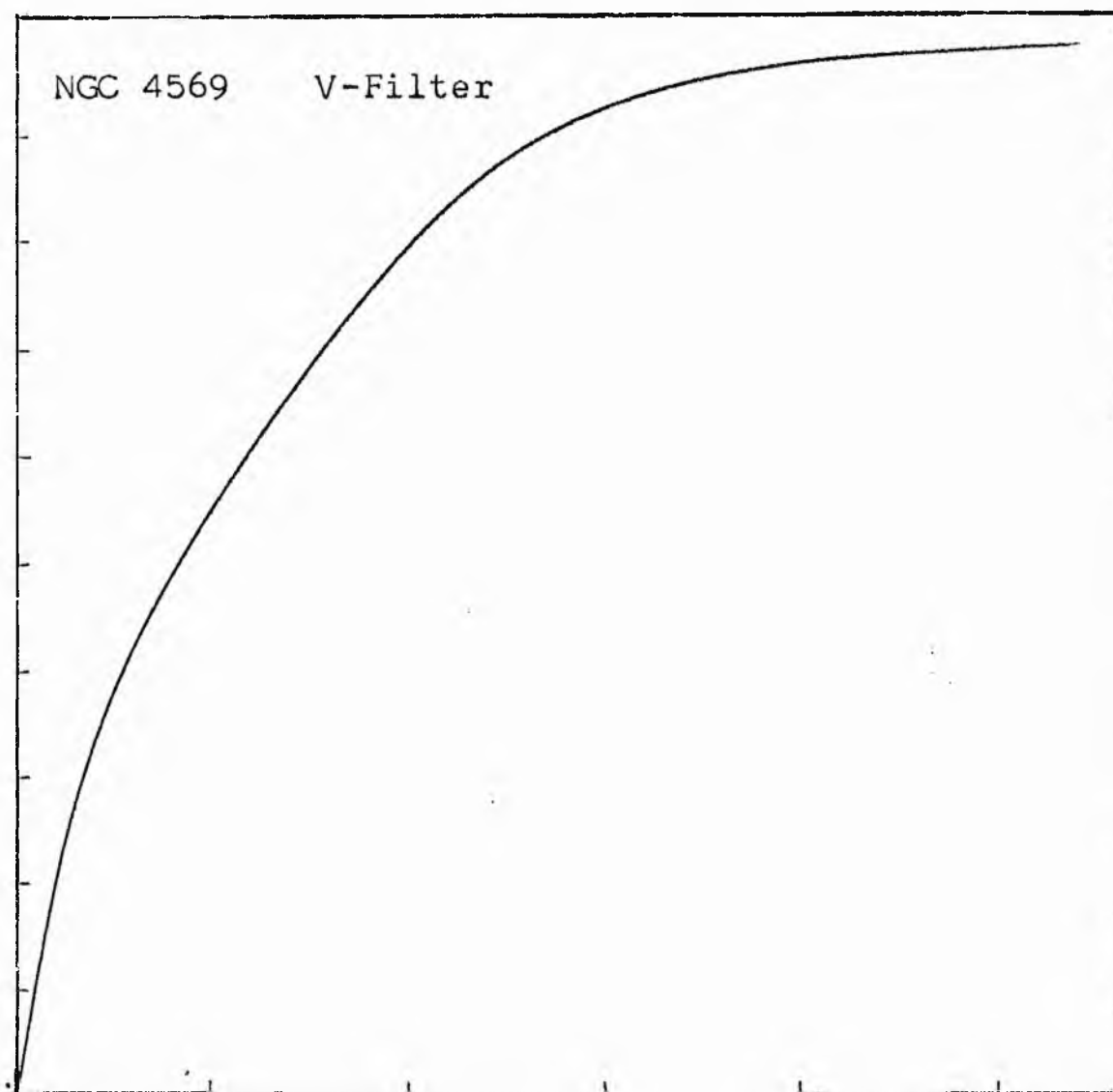
NGC 4569
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4569 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4569
V COLOUR

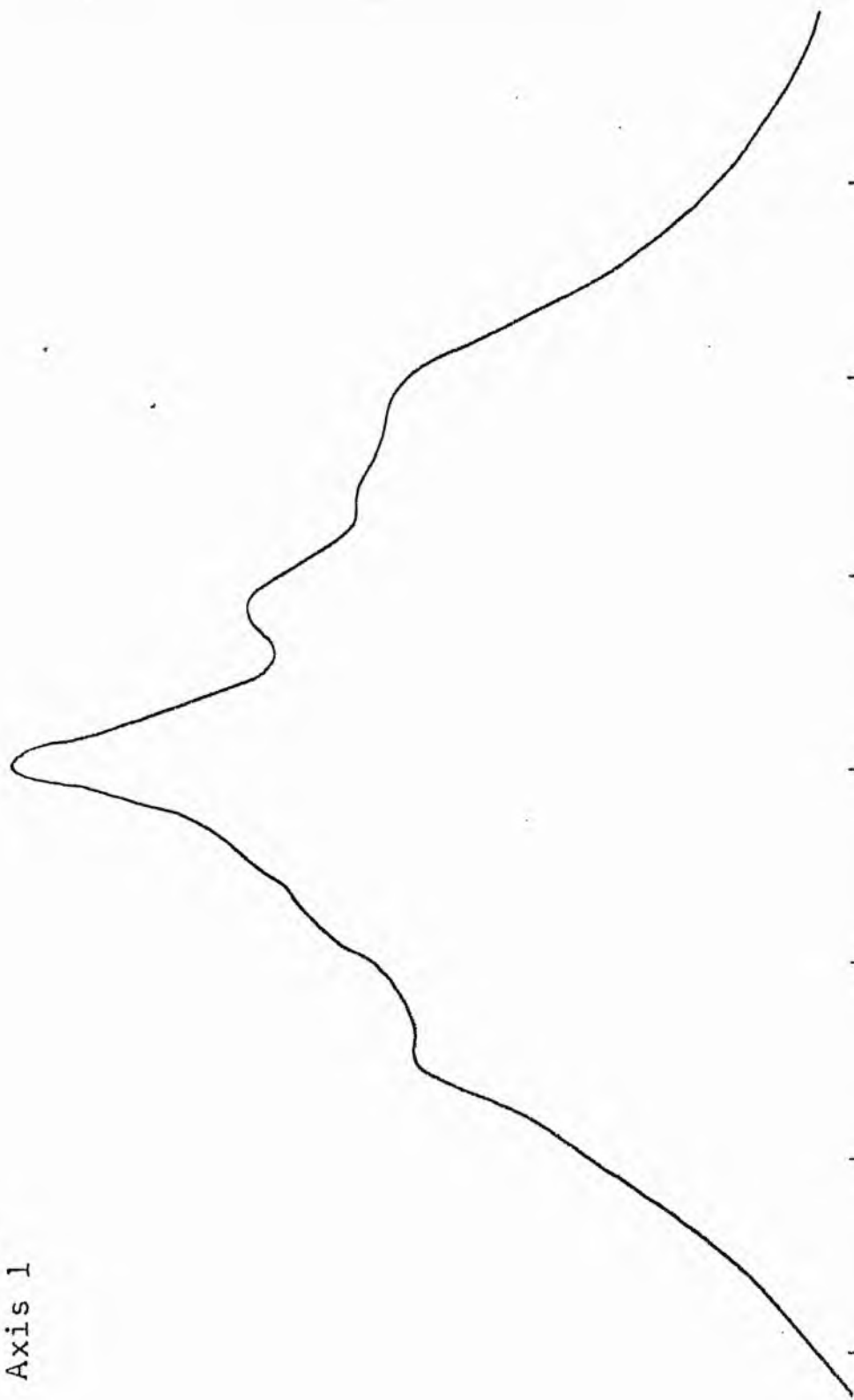
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.17	147.911		0.0	0.0			0.0	0.0	0.0	1.972	16.25
		153.200			4.52	693.0591					
2.20	158.489	142.191	1.20	4.52	5.65	804.0684	693.06	0.01	0.03	2.002	16.18
2.10	125.892	112.946	1.80	10.18	7.92	894.1729	1497.13	0.02	0.04	1.902	16.43
2.00	100.000	89.716	2.40	18.10	12.10	1085.1277	2371.30	0.04	0.06	1.802	16.68
1.90	79.433	71.264	3.10	30.19	15.17	1081.3521	3476.43	0.05	0.08	1.702	16.93
1.80	63.095	56.607	3.80	45.36	12.72	720.2346	4557.78	0.07	0.09	1.602	17.18
1.70	50.118	44.964	4.30	58.09	62.67	2818.1375	5278.01	0.08	0.11	1.502	17.43
1.60	39.810	35.717	6.20	120.76	51.27	1831.2126	8096.15	0.12	0.15	1.402	17.68
1.50	31.623	28.371	7.40	172.03	65.75	1865.4685	9927.36	0.15	0.18	1.302	17.93
1.40	25.114	22.536	8.70	237.79	63.93	1440.7290	11742.82	0.18	0.21	1.202	18.18
1.30	19.952	17.901	9.80	301.72	128.33	2297.2578	13233.55	0.20	0.24	1.102	18.43
1.20	15.849	14.219	11.70	430.05	76.65	1089.9502	15530.81	0.24	0.29	1.002	18.68
1.10	12.509	11.295	12.70	536.71	172.16	1944.4563	16620.76	0.25	0.31	0.902	18.93
1.00	10.000	8.972	14.70	678.87	182.66	1638.7688	18565.21	0.28	0.36	0.802	19.18
0.90	7.943	7.126	16.56	861.53	156.35	1114.1802	20203.98	0.31	0.41	0.702	19.43
0.80	6.310	5.661	18.00	1017.88	289.53	1638.9214	21318.16	0.33	0.44	0.602	19.68
0.70	5.012	4.496	20.40	1307.40	387.90	1744.1709	22957.07	0.35	0.50	0.502	19.93
0.60	3.981	3.572	23.23	1695.31	236.90	846.1121	24701.24	0.38	0.57	0.402	20.18
0.50	3.162	2.837	24.80	1932.20	750.11	2128.1084	25547.35	0.39	0.61	0.302	20.43
0.40	2.512	2.254	29.22	2682.32	1122.28	2529.1091	27675.46	0.42	0.72	0.202	20.68
0.30	1.995	1.790	34.80	3604.59	1348.40	2413.7246	30204.56	0.46	0.86	0.102	20.93
0.20	1.585	1.422	40.50	5153.00	2763.94	3930.0371	32618.29	0.50	1.00	0.002	21.18
0.10	1.259	1.129	50.20	7916.94	2871.16	3242.8367	36548.32	0.56	1.24	-0.098	21.43
-0.00	1.000	0.897	58.60	10780.10	3568.22	3201.2478	39791.16	0.61	1.44	-0.198	21.68
-0.10	0.794	0.713	67.60	14356.32	5499.34	3919.0247	42992.40	0.66	1.66	-0.298	21.93
-0.20	0.631	0.566	79.50	19855.65	5084.84	2878.3586	46911.43	0.72	1.96	-0.398	22.18
-0.30	0.501	0.450	89.10	24940.49	5354.64	2407.6758	49789.78	0.76	2.19	-0.498	22.43
-0.40	0.398	0.357	98.20	30295.13	5807.54	2074.2463	52197.46	0.80	2.42	-0.598	22.68
-0.50	0.316	0.284	107.20	36102.67	5952.23	1688.6816	54271.70	0.83	2.64	-0.698	22.93
-0.60	0.251	0.225	115.70	42054.69	6094.52	1373.4341	55960.38	0.86	2.85	-0.798	23.18
-0.70	0.200	0.179	123.80	48149.41	6341.16	1135.1077	57333.82	0.88	3.05	-0.898	23.43
-0.80	0.158	0.142	131.70	54490.57	9303.39	1322.8477	58468.92	0.89	3.24	-0.998	23.68
-0.90	0.126	0.113	142.50	63793.97	4833.59	545.9321	59791.77	0.92	3.51	-1.098	23.93
-1.00	0.100	0.090	147.80	68627.56	6557.25	588.2896	60337.70	0.92	3.64	-1.198	24.18
-1.10	0.079	0.071	154.70	75184.81	6552.12	466.9304	60925.99	0.93	3.81	-1.298	24.43
-1.20	0.063	0.057	161.30	81736.94	11528.37	652.5872	61392.92	0.94	3.97	-1.398	24.68
-1.30	0.050	0.045	172.30	93265.31	9088.56	408.6636	62045.50	0.95	4.24	-1.498	24.93
-1.40	0.040	0.036	180.50	102353.87	12254.56	437.6426	62454.16	0.96	4.44	-1.598	25.18
-1.50	0.032	0.028	191.00	114608.44	11936.44	338.6467	62891.86	0.96	4.70	-1.698	25.43
-1.60	0.025	0.023	200.70	126544.87	20572.62	463.6204	63230.50	0.97	4.94	-1.798	25.68
-1.70	0.020	0.018	216.40	147117.50	21538.56	385.5583	63694.12	0.97	5.33	-1.898	25.93
-1.80	0.016	0.014	231.70	168656.06	17922.12	254.8376	64079.67	0.98	5.70	-1.998	26.18
-1.90	0.013	0.011	243.70	186578.19	13400.75	151.3573	64334.51	0.98	6.00	-2.098	26.43
-2.00	0.010		252.30	199978.94			64485.86	0.99	6.21	-2.198	26.68
-∞							65345.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4569

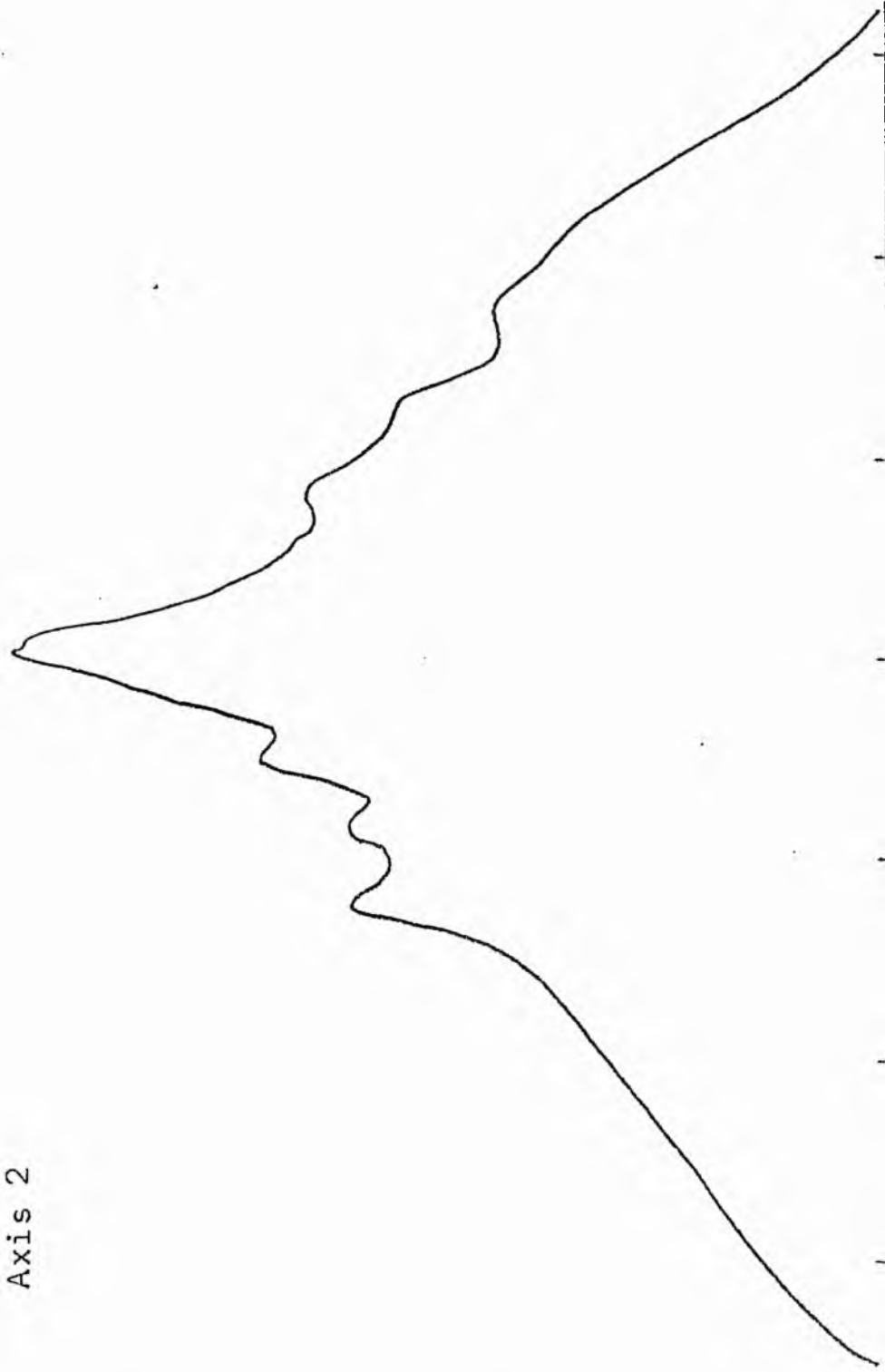
V-FILTER

Total luminosity	L_T	= 18.15
Total apparent magnitude	m_T	= 9.62
Apparent central surface brightness	μ_0	= 16.25
Major axis at threshold	$2a_m$	= 12.65
Minor axis at threshold	$2b_m$	= 5.77
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.75
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.95
Gradient of exponential component	$G(a)$	= -0.46
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.46
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.48
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 0.82
Equivalent radius	r_1^*	= 0.21
Surface brightness	μ_1	= 18.93
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.08
Axis ratio	b/a	= 0.34
Equivalent radius	r_e^*	= 0.68
Surface brightness	μ_e	= 21.18
Mean surface brightness	μ_e'	= 10.76
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.36
Axis ratio	b/a	= 0.47
Equivalent radius	r_3^*	= 1.44
Surface brightness	μ_3	= 22.36
Concentration indices	$\{C_{21}$	= 3.27
	C_{32}	= 2.13

NGC 4571
B-Filter
Axis 1

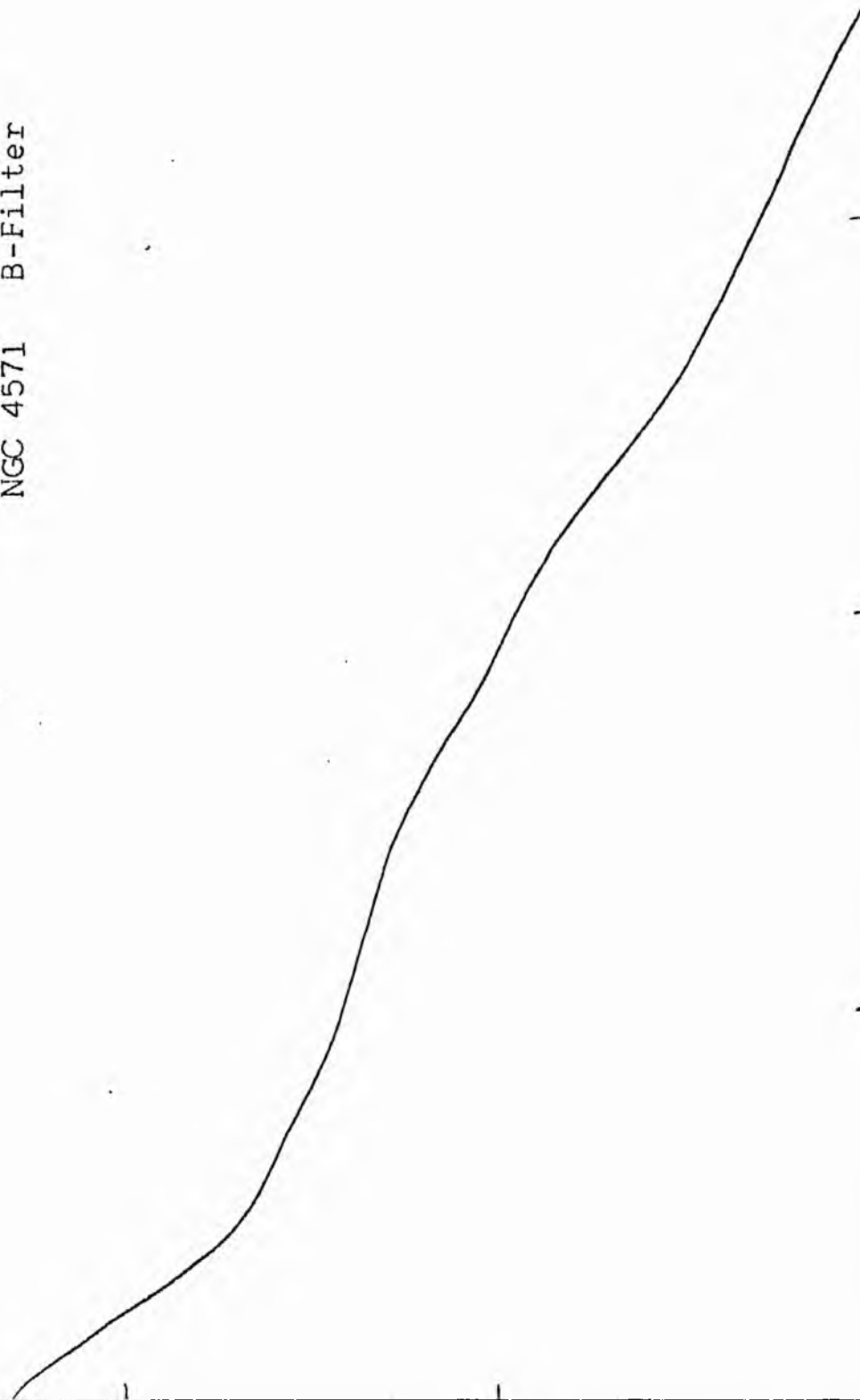


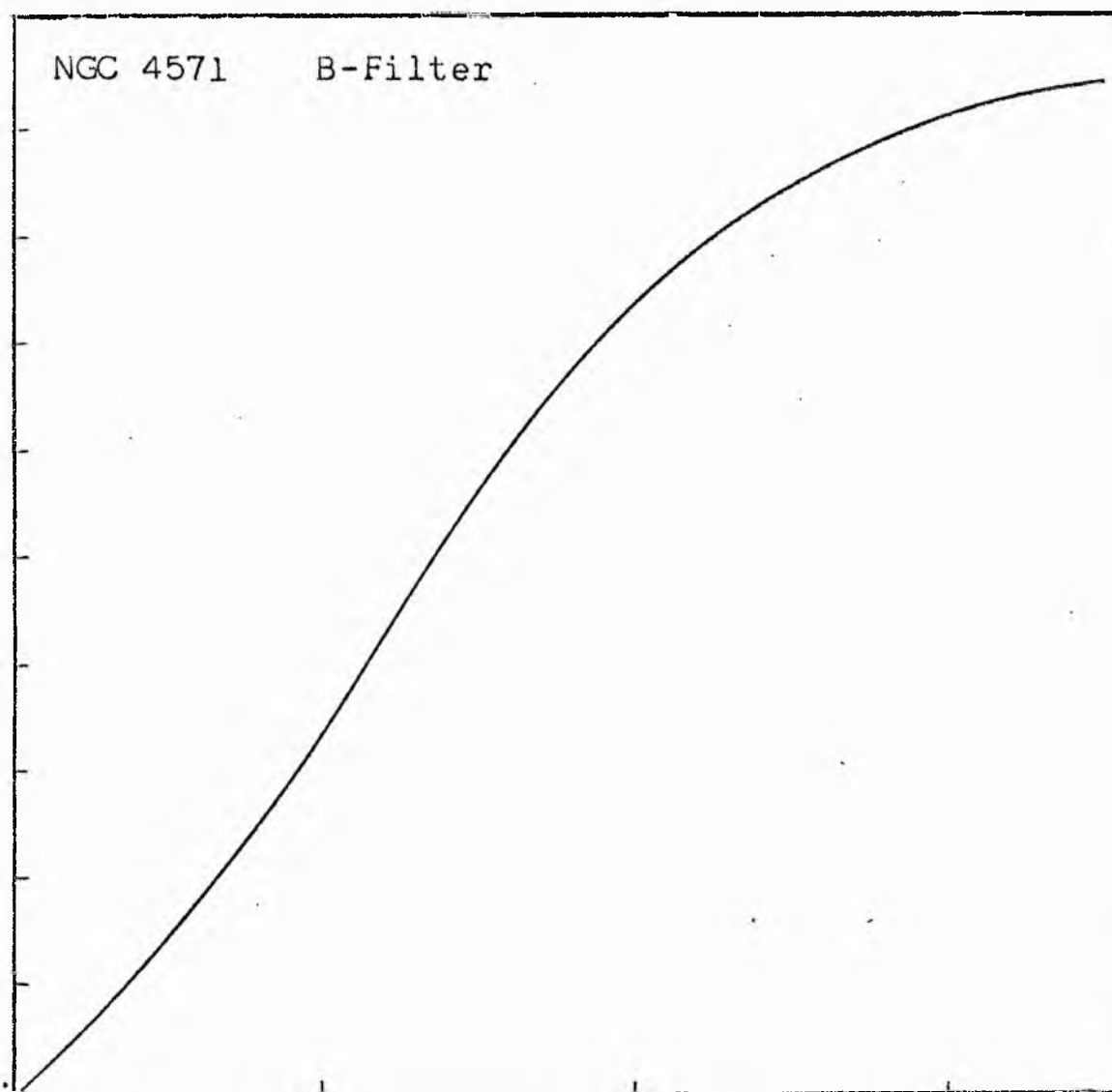
NGC 4571
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4571 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

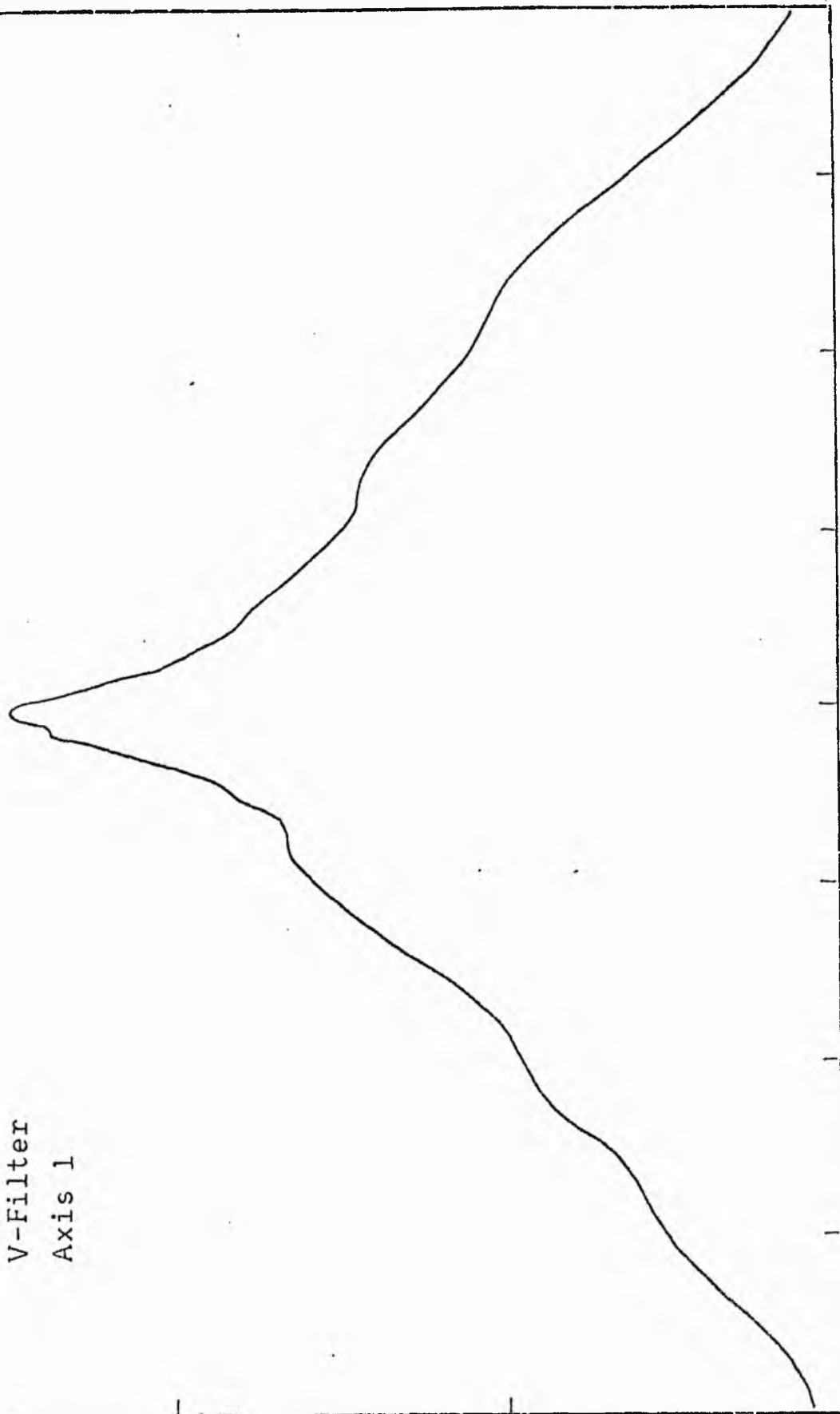
MEAN LUMINOSITY DISTRIBUTION IN NGC 4571
B COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.23	1.698	1.642	0.0	0.0	56.21	92.2761	0.0	0.0	0.0	0.942	21.06
0.20	1.585	1.422	4.23	56.21	145.46	235.2671	92.28	0.01	0.06	0.912	21.14
0.10	1.259	1.129	8.40	221.67	131.32	148.3193	327.54	0.03	0.12	0.812	21.39
0.00	1.000	0.897	10.60	352.99	194.40	174.4103	475.86	0.05	0.16	0.712	21.64
-0.10	0.794	0.713	13.20	547.39	328.77	234.2936	650.27	0.07	0.20	0.612	21.89
-0.20	0.631	0.566	16.70	876.16	277.14	156.8789	884.57	0.09	0.25	0.512	22.14
-0.30	0.501	0.450	19.16	1153.29	1416.40	636.8811	1041.44	0.11	0.28	0.412	22.39
-0.40	0.398	0.357	28.60	2569.70	1824.64	651.7009	1678.33	0.18	0.42	0.312	22.64
-0.50	0.316	0.284	37.40	4394.33	4284.48	1215.5444	2330.03	0.24	0.56	0.212	22.89
-0.60	0.251	0.225	52.56	8678.82	4993.50	1125.3242	3545.57	0.37	0.78	0.112	23.14
-0.70	0.200	0.179	65.97	13672.32	4612.25	825.6299	4670.89	0.49	0.98	0.012	23.39
-0.80	0.158	0.142	76.29	18284.57	5237.91	744.7832	5496.52	0.57	1.13	-0.088	23.64
-0.90	0.126	0.113	86.53	23522.48	6097.73	688.7170	6241.30	0.65	1.28	-0.188	23.89
-1.00	0.100	0.090	97.10	29620.21	5353.11	480.2627	6930.02	0.72	1.44	-0.288	24.14
-1.10	0.079	0.071	105.51	34973.33	4886.46	348.2312	7410.28	0.77	1.57	-0.388	24.39
-1.20	0.063	0.057	112.64	39859.79	312.01	17.6622	7758.51	0.81	1.67	-0.488	24.64
-1.30	0.050	0.045	113.08	40171.80	5407.05	243.1275	7776.17	0.81	1.68	-0.588	24.89
-1.40	0.040	0.036	120.45	45578.86	6854.49	244.8214	8019.30	0.84	1.79	-0.688	25.14
-1.50	0.032	0.028	129.19	52433.34	8501.40	241.1934	8264.12	0.86	1.92	-0.788	25.39
-1.60	0.025	0.023	139.27	60934.74	7136.76	160.8336	8505.31	0.89	2.07	-0.888	25.64
-1.70	0.020	0.018	147.20	68071.50	7697.69	137.7960	8666.14	0.91	2.19	-0.988	25.89
-1.80	0.016	0.014	155.30	75769.19	11532.12	163.9782	8803.93	0.92	2.31	-1.088	26.14
-1.90	0.013	0.011	166.70	87301.31	8065.12	91.0937	8967.91	0.94	2.47	-1.188	26.39
-2.00	0.010		174.23	95366.44			9059.00	0.95	2.59	-1.288	26.64
-∞							9569.00	(1)			∞

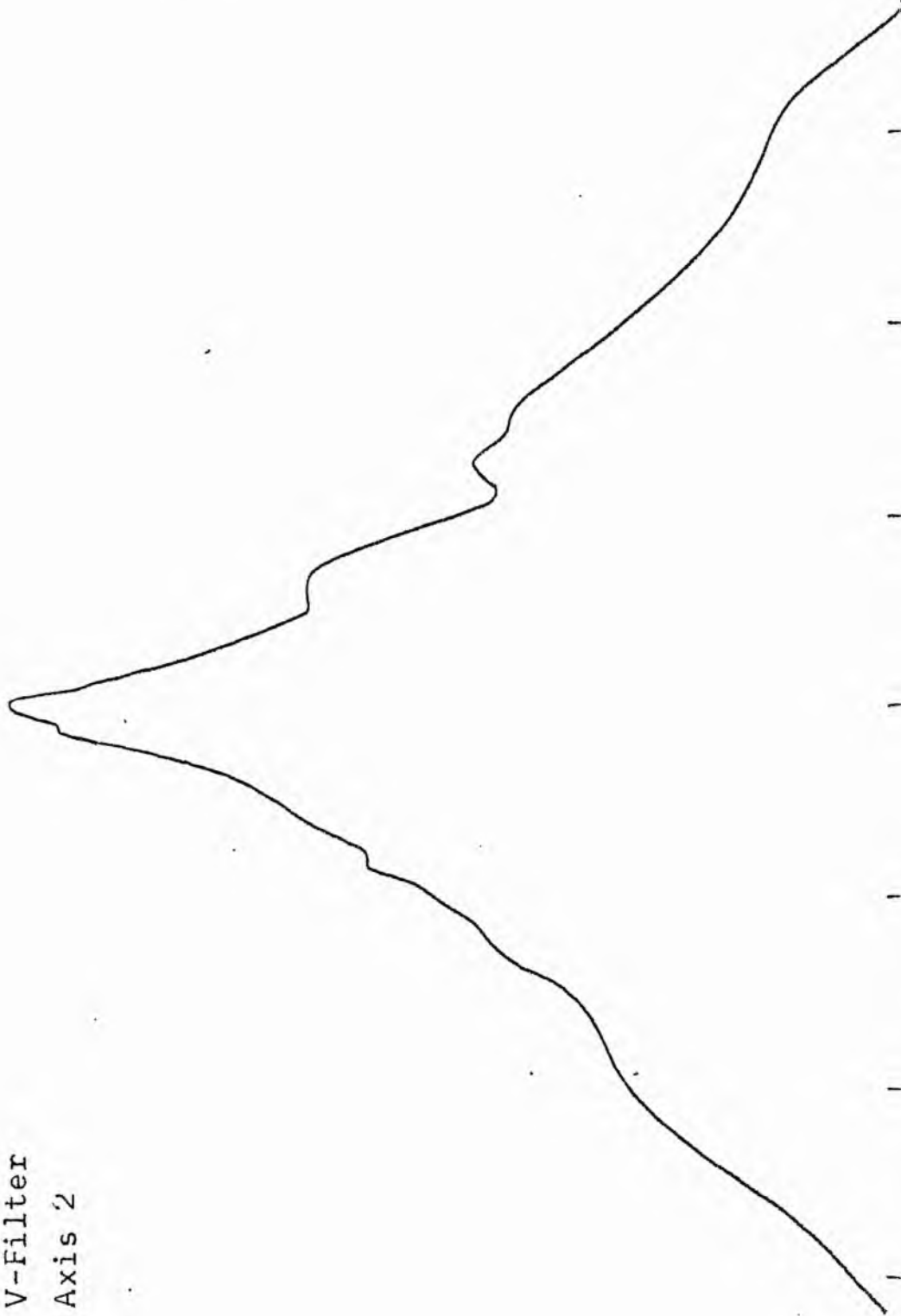
PHOTOMETRIC PARAMETERS OF NGC 4571
B-FILTER

Total luminosity	L_T	= 2.66
Total apparent magnitude	m_T	= 12.01
Apparent central surface brightness	μ_0	= 21.06
Major axis at threshold	$2a_m$	= 6.28
Minor axis at threshold	$2b_m$	= 5.80
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.86
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.68
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.73
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.78
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.71
Axis ratio	b/a	= 0.79
Equivalent radius	r_1^*	= 0.64
Surface brightness	μ_1	= 22.90
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.13
Axis ratio	b/a	= 0.84
Equivalent radius	r_e^*	= 1.12
Surface brightness	μ_e	= 23.42
Mean surface brightness	μ_e'	= 14.21
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.63
Axis ratio	b/a	= 0.94
Equivalent radius	r_3^*	= 1.69
Surface brightness	μ_3	= 24.29
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 1.76 \\ = 1.50 \end{matrix}$

NGC 4571
V-Filter
Axis 1

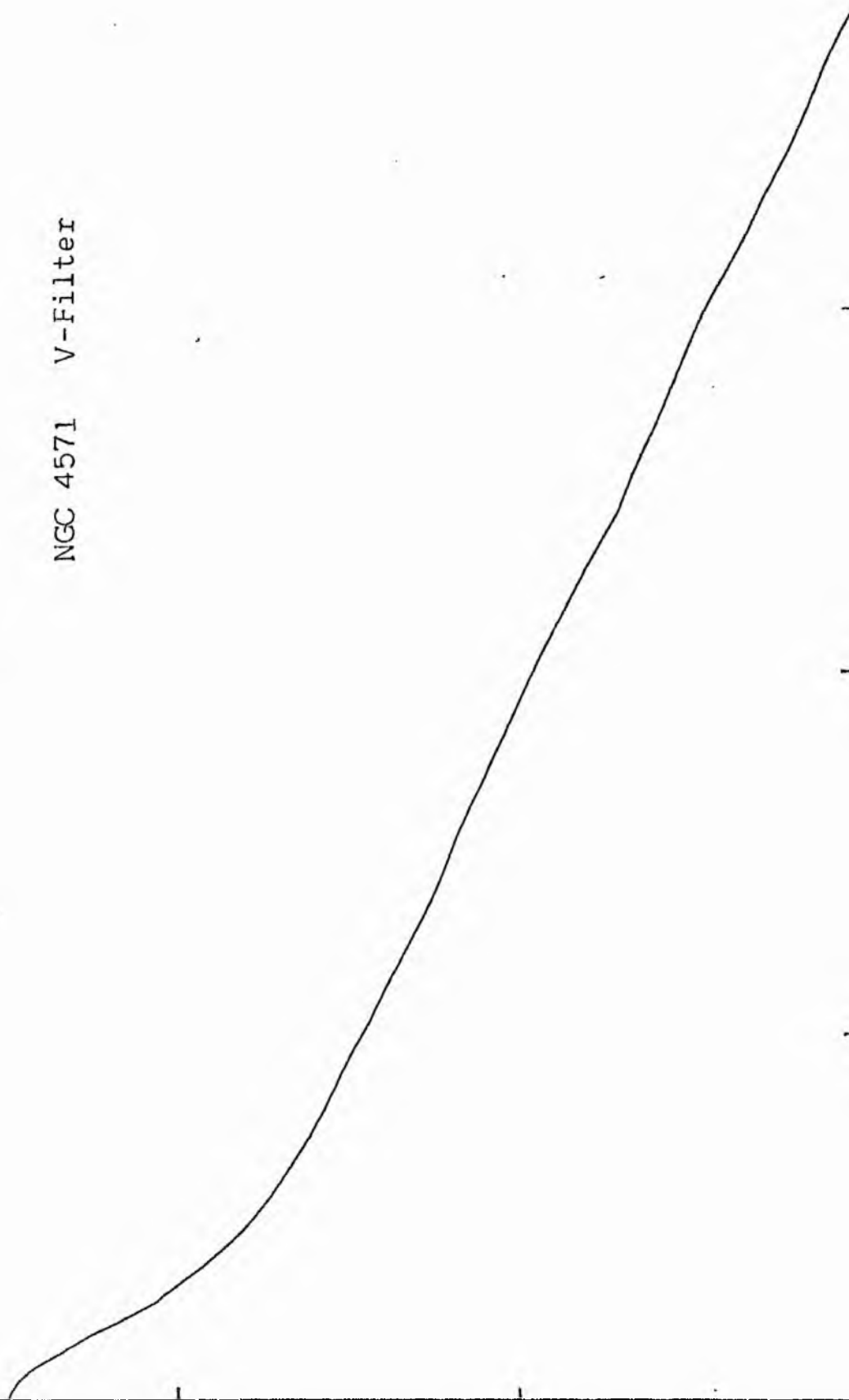


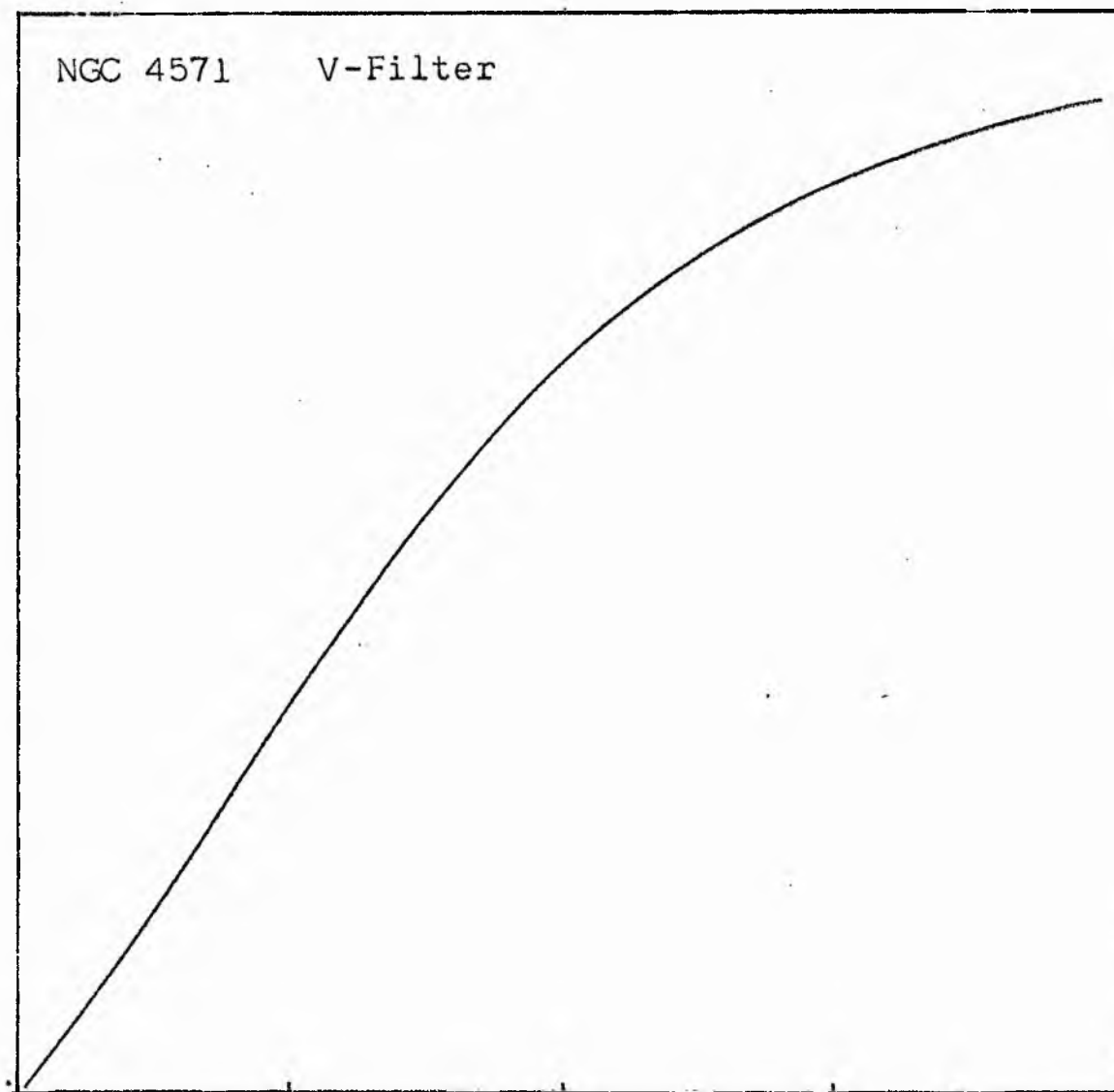
NGC 4571
V-Filter
Axis '2



Equivalent luminosity profile

NGC 4571 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4571
V COLOUR

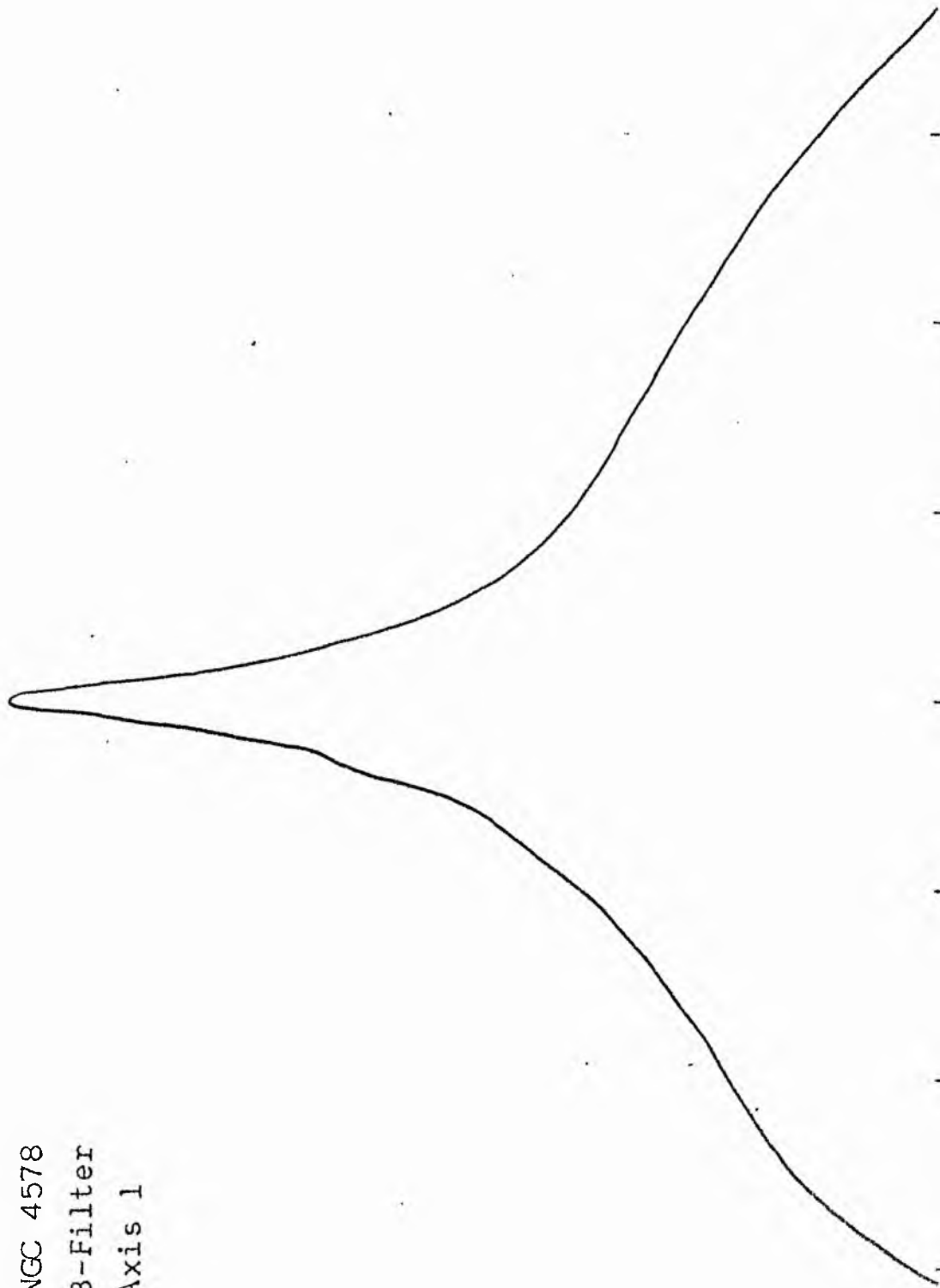
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.43	2.692	2.602	0.0	0.0	45.84	119.2712	0.0	0.0	0.0	1.189	20.14
0.40	2.512	2.254	3.82	45.84	78.45	176.7949	119.27	0.01	0.06	1.159	20.22
0.30	1.995	1.793	6.29	124.29	106.44	190.5352	296.07	0.03	0.09	1.059	20.47
0.20	1.585	1.422	8.57	230.73	142.52	202.6484	486.60	0.04	0.12	0.959	20.72
0.10	1.259	1.129	10.90	373.25	212.95	240.5244	689.25	0.06	0.16	0.859	20.97
0.00	1.000	0.897	13.66	586.21	269.09	241.4190	929.77	0.08	0.20	0.759	21.22
-0.10	0.794	0.713	16.50	855.30	547.35	370.0635	1171.19	0.10	0.24	0.659	21.47
-0.20	0.631	0.566	21.13	1402.65	892.67	505.3135	1561.26	0.14	0.31	0.559	21.72
-0.30	0.501	0.450	27.03	2295.31	1379.22	620.1621	2066.57	0.18	0.39	0.459	21.97
-0.40	0.398	0.357	34.20	3674.53	2270.14	810.8220	2686.73	0.24	0.50	0.359	22.22
-0.50	0.316	0.284	43.50	5944.68	3462.12	982.2329	3497.55	0.31	0.63	0.259	22.47
-0.60	0.251	0.225	54.72	9406.80	3169.29	714.2227	4479.79	0.40	0.80	0.159	22.72
-0.70	0.200	0.179	63.27	12576.08	4041.84	723.5212	5194.01	0.46	0.92	0.059	22.97
-0.80	0.158	0.142	72.73	16617.92	5071.46	721.1165	5917.53	0.53	1.06	-0.041	23.22
-0.90	0.126	0.113	83.09	21689.38	6699.31	756.6633	6638.64	0.59	1.21	-0.141	23.47
-1.00	0.100	0.090	95.06	28388.70	6458.79	579.4597	7395.30	0.66	1.38	-0.241	23.72
-1.10	0.079	0.071	105.32	34847.48	5452.32	388.5564	7974.76	0.71	1.53	-0.341	23.97
-1.20	0.063	0.057	113.26	40299.80	7376.29	417.5525	8363.31	0.74	1.65	-0.441	24.22
-1.30	0.050	0.045	123.19	47676.09	8818.56	396.5256	8780.86	0.78	1.79	-0.541	24.47
-1.40	0.040	0.036	134.10	56494.65	10471.54	374.0112	9177.39	0.82	1.95	-0.641	24.72
-1.50	0.032	0.028	146.00	66966.19	7762.50	220.2301	9551.39	0.85	2.12	-0.741	24.97
-1.60	0.025	0.023	154.23	74728.69	9561.56	215.4789	9771.62	0.87	2.24	-0.841	25.22
-1.70	0.020	0.018	163.80	84290.25	9843.19	176.2025	9987.10	0.89	2.38	-0.941	25.47
-1.80	0.016	0.014	173.10	94133.44	13853.00	196.9793	10163.30	0.90	2.52	-1.041	25.72
-1.90	0.013	0.011	185.40	107986.44	15550.00	175.6336	10360.27	0.92	2.69	-1.141	25.97
-2.00	0.010		198.30	123536.44			10535.91	0.94	2.88	-1.241	26.22
-∞							11255.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4571

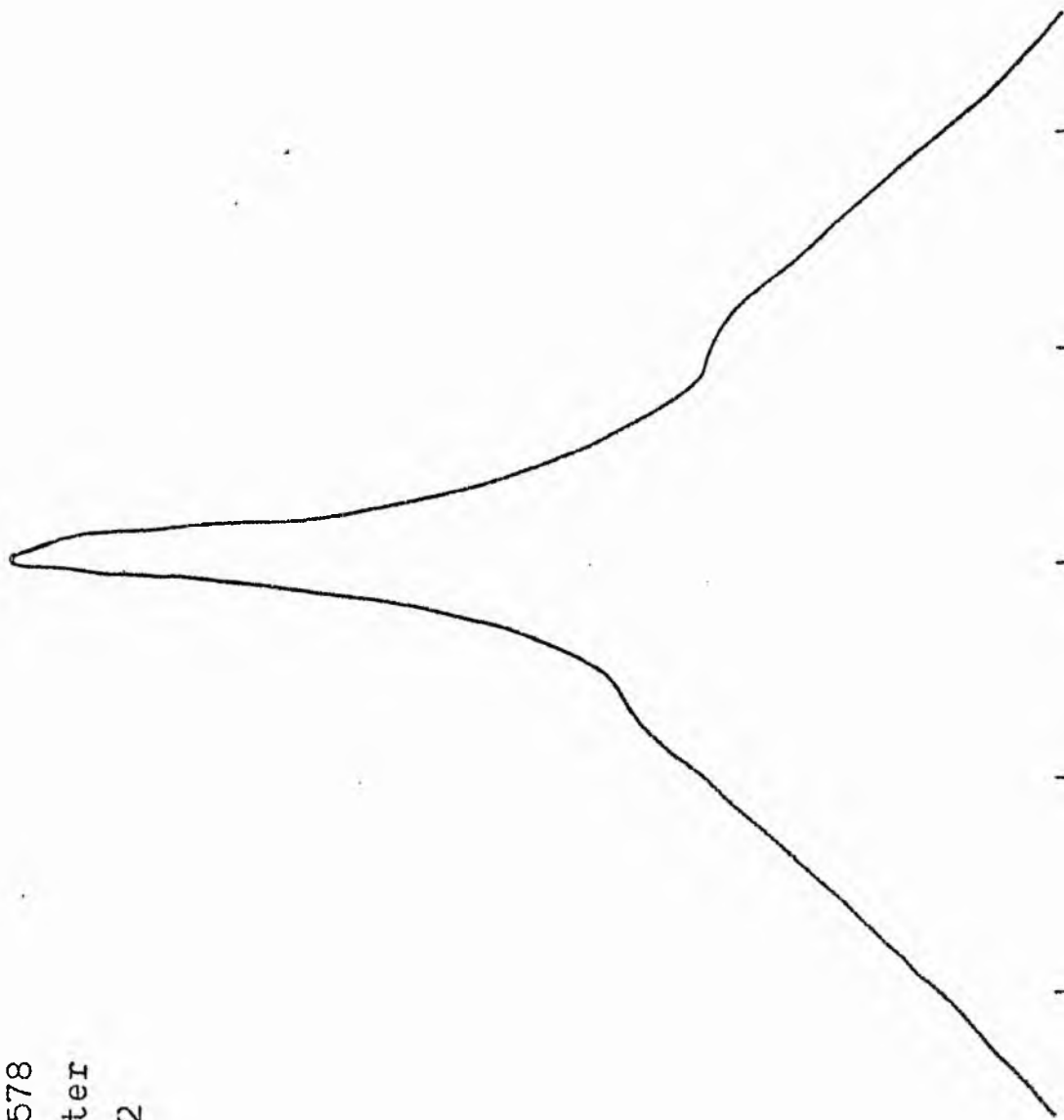
V-FILTER

Total luminosity	L_T	= 3.13
Total apparent magnitude	m_T	= 11.09
Apparent central surface brightness	μ_0	= 20.14
Major axis at threshold	$2a_m$	= 6.93
Minor axis at threshold	$2b_m$	= 6.04
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.05
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.85
Gradient of exponential component	$G(a)$	= -0.53
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.62
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= +0.72
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.72
Axis ratio	b/a	= 0.77
Equivalent radius	r_1^*	= 0.60
Surface brightness	μ_1	= 22.25
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.32
Axis ratio	b/a	= 0.76
Equivalent radius	r_e^*	= 1.15
Surface brightness	μ_e	= 23.11
Mean surface brightness	μ_e'	= 13.39
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.18
Axis ratio	b/a	= 0.75
Equivalent radius	r_3^*	= 1.92
Surface brightness	μ_3	= 24.28
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 1.93 \\ = 1.67 \end{cases}$

NGC 4578
B-Filter
Axis 1

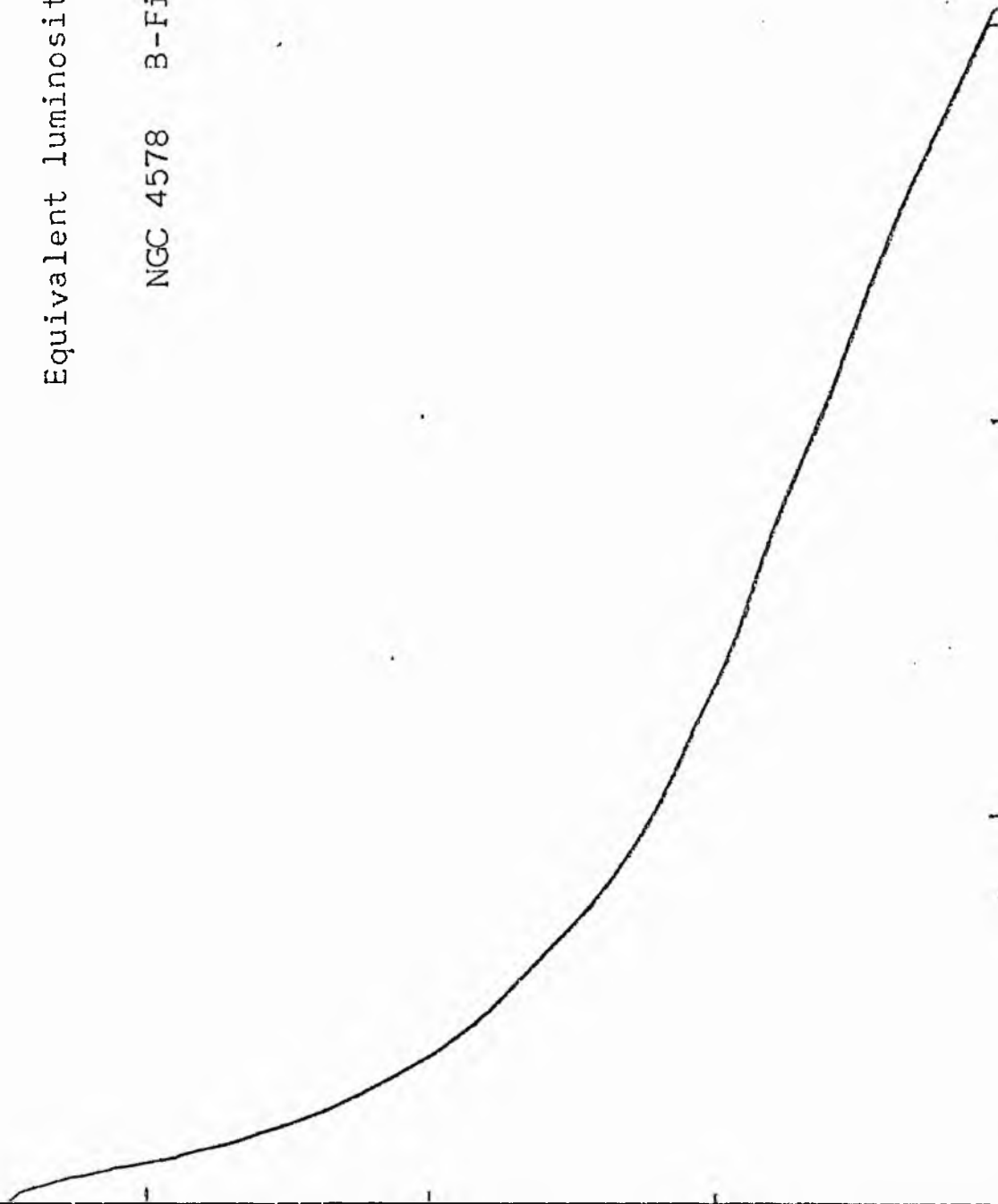


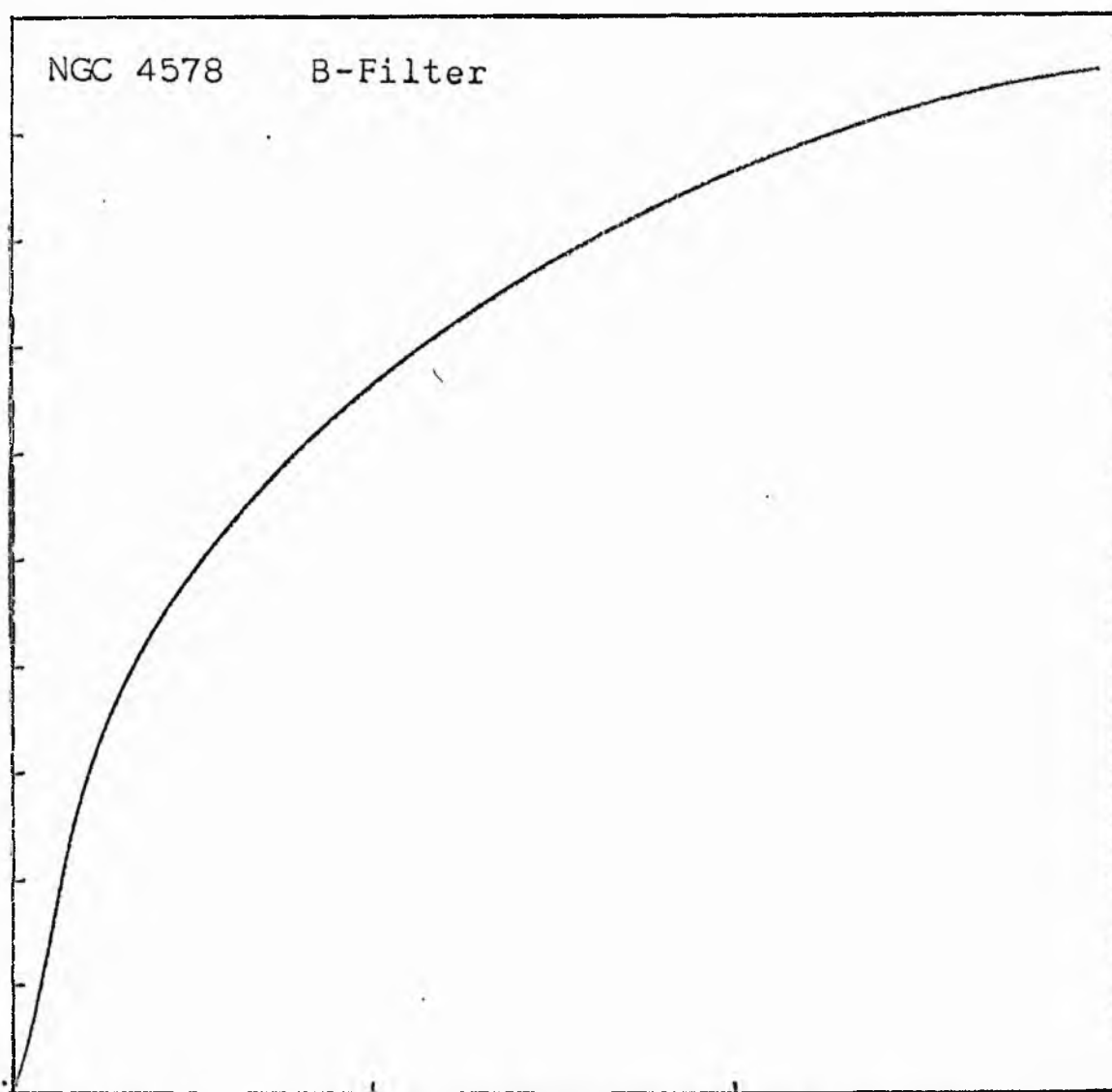
NGC 4578
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4578 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN MGC 4578
& COLOUR

LOG I	I	T	R	AREA	ΔA	P	ΣP	KIRI	ρ	LOG J	μ
1.43	26.915		0.0	0.0			0.0	0.0	0.0	1.717	18.61
		26.017			24.45	636.2332					
1.40	25.119	22.536	2.79	24.45	14.03	316.1765	636.23	0.07	0.11	1.667	18.69
1.30	19.953	17.901	3.50	38.40	19.60	350.9167	952.41	0.10	0.14	1.567	18.94
1.20	15.849	14.219	4.30	58.09	17.34	246.5811	1303.33	0.14	0.17	1.487	19.19
1.10	12.589	11.295	4.90	75.43	16.18	182.7371	1549.91	0.16	0.19	1.387	19.44
1.00	10.000	8.972	5.40	91.61	33.08	296.7891	1732.64	0.18	0.21	1.287	19.69
0.90	7.943	7.126	6.30	124.69	33.68	240.0022	2029.43	0.21	0.24	1.187	19.94
0.80	6.310	5.661	7.10	158.37	16.35	103.4566	2269.44	0.24	0.28	1.087	20.19
0.70	5.012	4.496	7.50	176.71	55.64	250.1717	2373.29	0.25	0.29	0.987	20.44
0.60	3.981	3.572	8.60	232.35	45.24	161.5782	2623.46	0.28	0.33	0.887	20.69
0.50	3.162	2.837	9.40	277.59	62.20	176.4754	2785.04	0.29	0.36	0.787	20.94
0.40	2.512	2.254	10.40	339.79	75.68	170.5526	2961.52	0.31	0.40	0.687	21.19
0.30	1.995	1.790	11.50	415.48	272.66	488.0786	3132.07	0.33	0.45	0.587	21.44
0.20	1.585	1.422	14.80	688.13	86.24	122.6205	3620.15	0.38	0.57	0.487	21.69
0.10	1.259	1.129	15.70	774.37	209.86	237.0260	3742.77	0.39	0.61	0.387	21.94
-0.00	1.000	0.897	17.70	984.23	297.67	267.0542	3979.79	0.42	0.69	0.287	22.19
-0.10	0.794	0.713	20.20	1281.89	365.59	260.5320	4246.85	0.45	0.78	0.187	22.44
-0.20	0.631	0.566	22.90	1647.48	509.03	268.1479	4507.38	0.47	0.89	0.087	22.69
-0.30	0.501	0.450	26.20	2156.51	689.80	310.1650	4795.52	0.50	1.02	-0.013	22.94
-0.40	0.398	0.357	30.10	2846.31	871.32	311.2068	5105.69	0.54	1.17	-0.113	23.19
-0.50	0.316	0.284	34.40	3717.63	1134.52	321.8716	5416.89	0.57	1.34	-0.213	23.44
-0.60	0.251	0.225	39.30	4852.15	1425.03	321.1406	5738.76	0.60	1.53	-0.313	23.69
-0.70	0.200	0.179	44.70	6277.18	1451.63	259.8530	6059.90	0.64	1.74	-0.413	23.94
-0.80	0.158	0.142	49.60	7728.82	3059.28	435.0012	6319.75	0.66	1.93	-0.513	24.19
-0.90	0.126	0.113	58.60	10788.10	2896.68	327.1689	6754.75	0.71	2.27	-0.613	24.44
-1.00	0.100	0.090	66.00	13684.78	4845.07	434.6829	7081.92	0.74	2.56	-0.713	24.69
-1.10	0.079	0.071	76.80	18529.85	4489.71	319.9570	7516.60	0.79	2.98	-0.813	24.94
-1.20	0.063	0.057	85.60	23019.57	5154.51	291.7832	7836.56	0.82	3.32	-0.913	25.19
-1.30	0.050	0.045	94.70	28174.07	4000.35	179.8754	8128.34	0.85	3.68	-1.013	25.44
-1.40	0.040	0.036	101.20	32174.43	5700.73	203.6126	8308.21	0.87	3.93	-1.113	25.69
-1.50	0.032	0.028	109.80	37875.15	4909.84	139.2973	8511.82	0.89	4.26	-1.213	25.94
-1.60	0.025	0.023	116.70	42785.00	6853.70	154.4547	8651.12	0.91	4.53	-1.313	26.19
-1.70	0.020	0.018	125.70	49638.70	7193.50	128.7706	8805.57	0.93	4.88	-1.413	26.44
-1.80	0.016	0.014	134.50	56832.20	7590.05	107.9249	8934.34	0.94	5.22	-1.513	26.69
-1.90	0.013	0.011	143.20	64422.25	5324.25	60.1361	9042.27	0.95	5.56	-1.613	26.94
-2.00	0.010		149.00	69746.50			9102.40	0.96	5.78	-1.713	27.19
-∞							9517.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4578

B-FILTER

Total luminosity	L_T	= 2.64
Total apparent magnitude	m_T	= 12.24
Apparent central surface brightness	μ_0	= 18.61
Major axis at threshold	$2a_m$	= 5.77
Minor axis at threshold	$2b_m$	= 4.25
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.88
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.80
Gradient of exponential component	$G(a)$	= -0.63
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.83
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.36

Parameters at $k = \frac{1}{4}$:

Semi-major axis	a_1	= 0.16
Axis ratio	b/a	= 0.80
Equivalent radius	r_1^*	= 0.13
Surface brightness	μ_1	= 20.44

Parameters at $k = \frac{1}{2}$ (effective) :

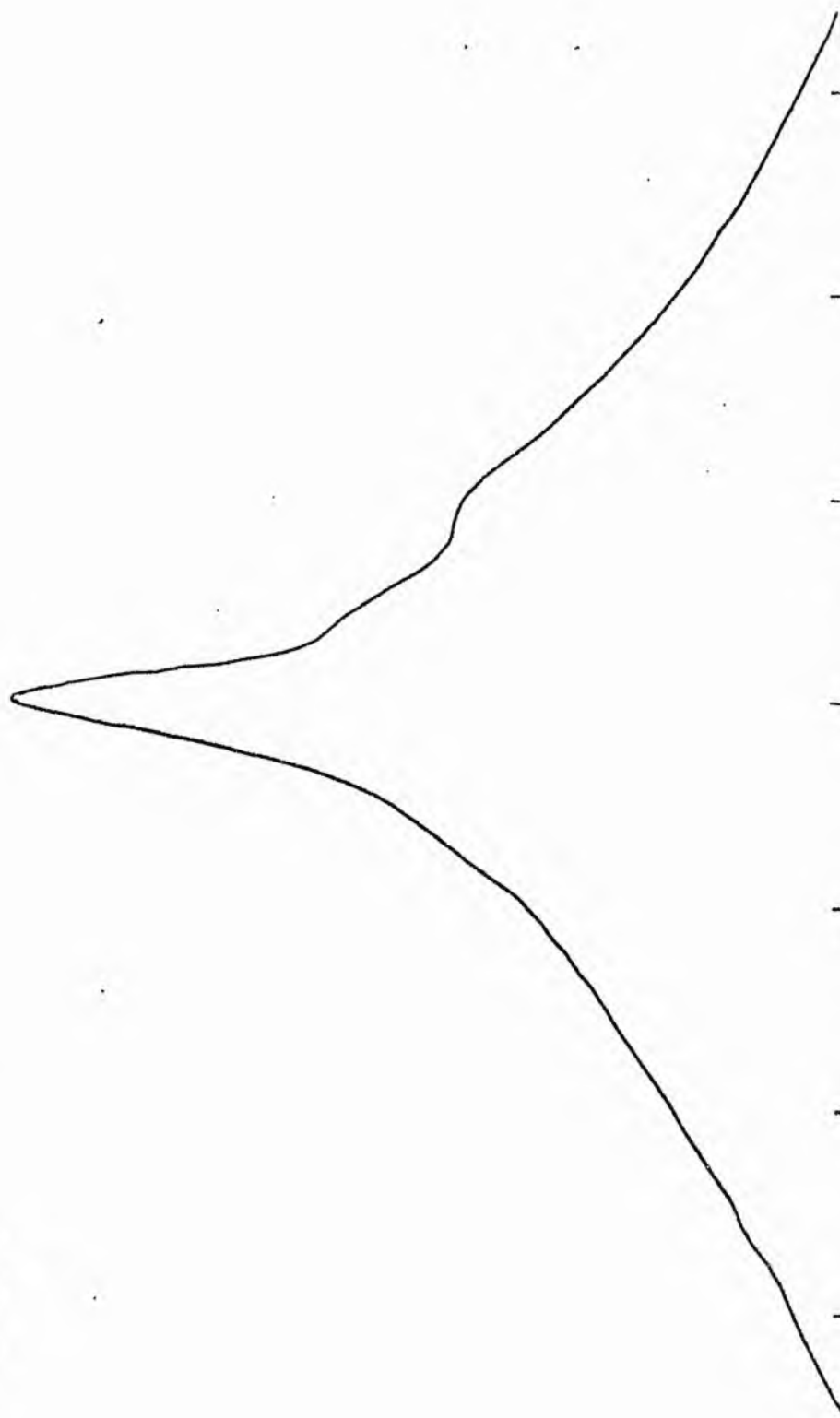
Semi-major axis	a_e	= 0.55
Axis ratio	b/a	= 0.70
Equivalent radius	r_e^*	= 0.43
Surface brightness	μ_e	= 22.94
Mean surface brightness	μ_e'	= 12.38

Parameters at $k = \frac{3}{4}$:

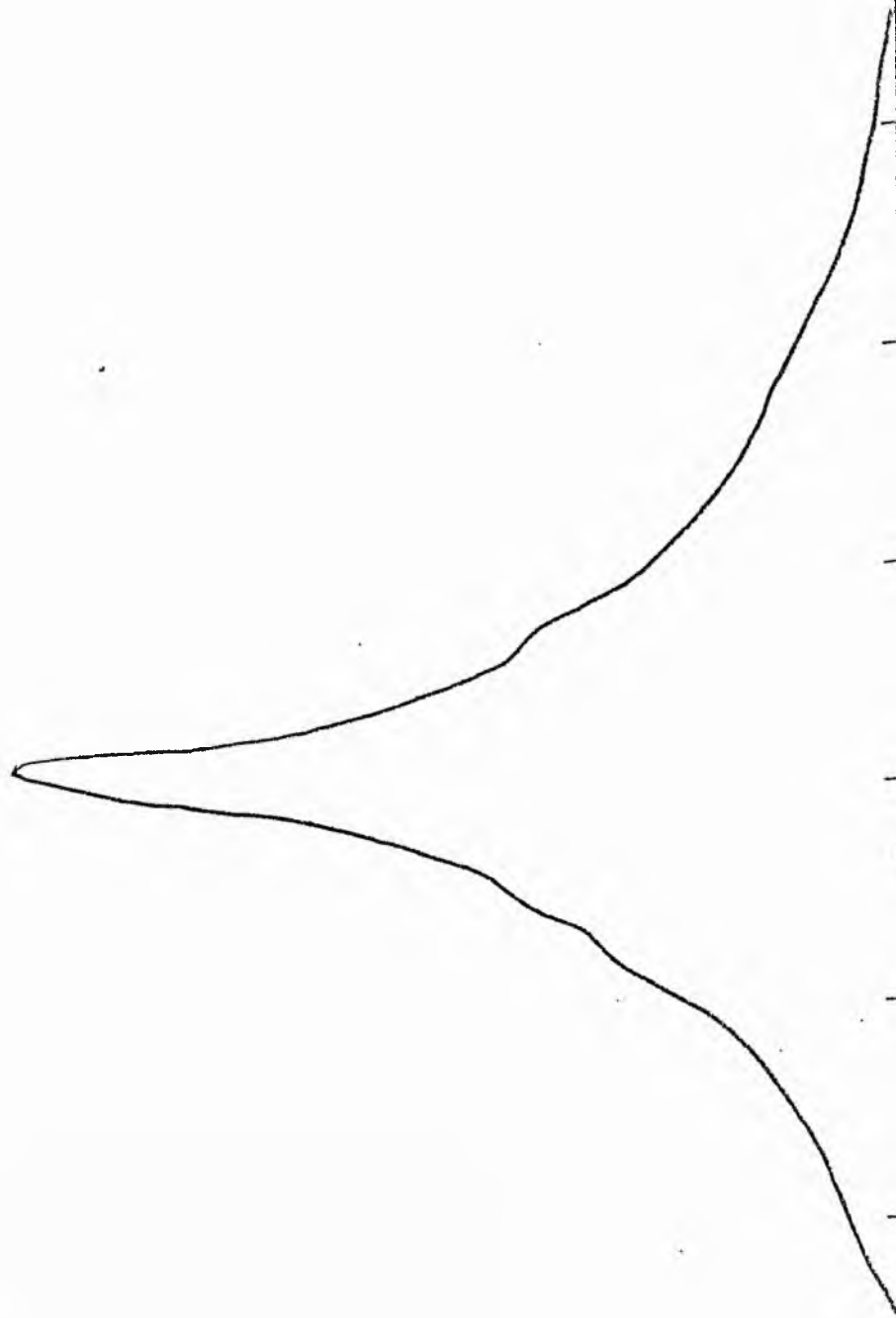
Semi-major axis	a_3	= 1.51
Axis ratio	b/a	= 0.66
Equivalent radius	r_3^*	= 1.12
Surface brightness	μ_3	= 24.74

Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 3.42 \\ = 2.61 \end{matrix}$
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NGC 4578
V-Filter
Axis 1



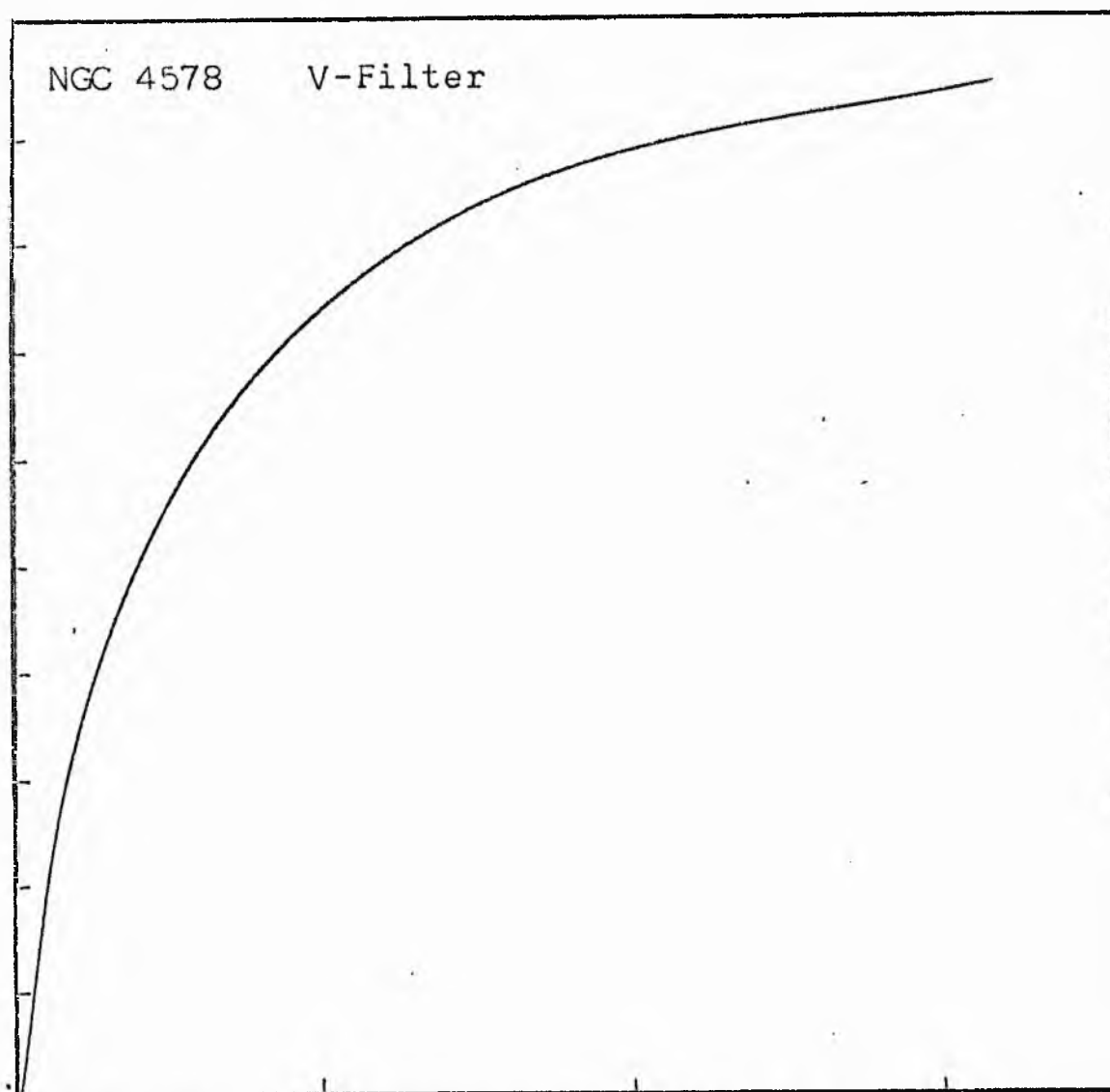
NGC 4578
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4578 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4578
V COLOUR

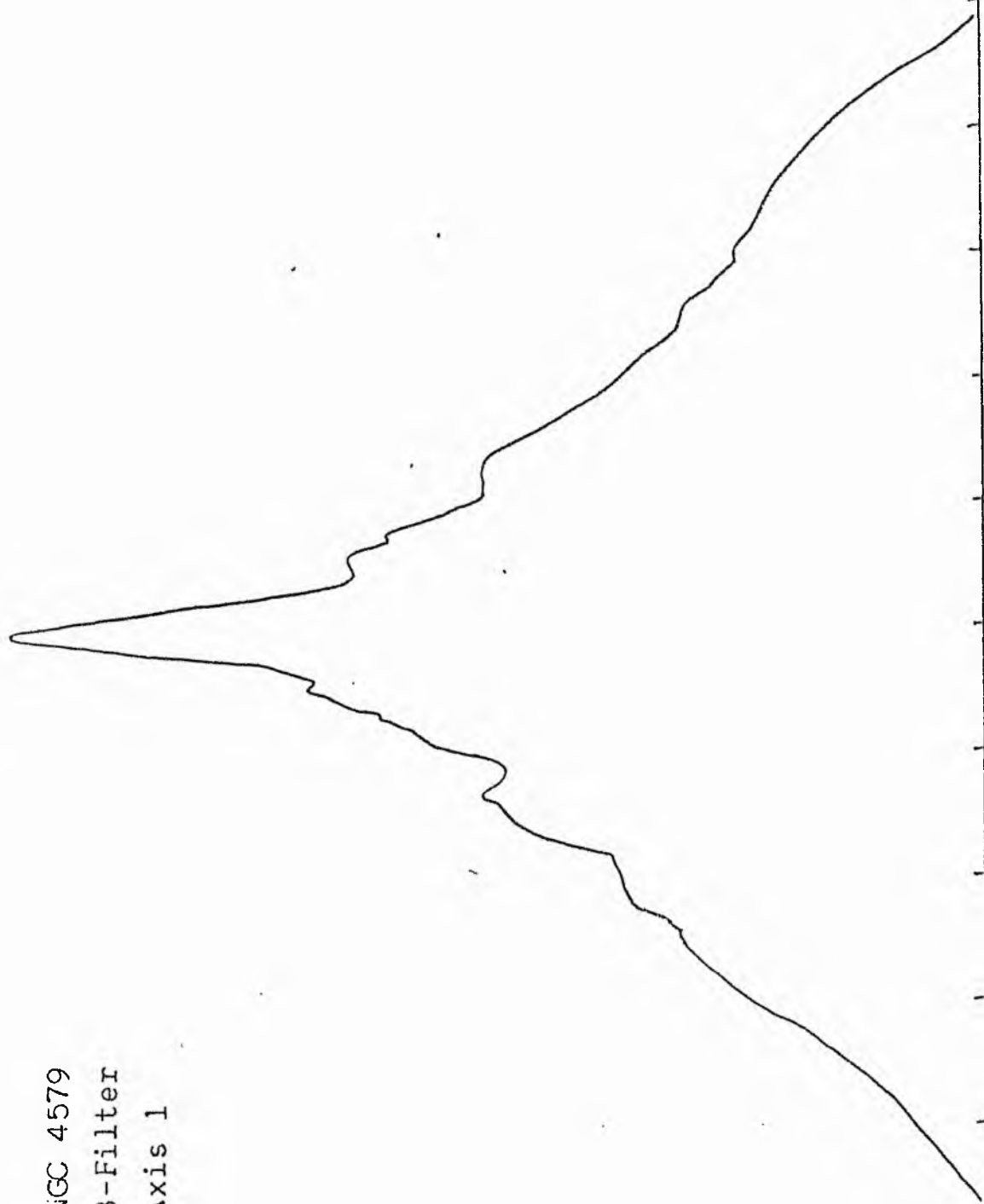
LOG I	I	I	R	AREA	A	P	P	K(R)	I	LOG J	J
1.56	36.308	33.965	0.0	0.0	4.52	153.6551	0.0	0.0	0.0	1.573	17.48
1.50	31.623	28.371	1.20	4.52	18.38	521.4570	153.66	0.01	0.06	1.513	17.63
1.40	25.119	22.536	2.70	22.90	20.11	453.1067	675.06	0.06	0.13	1.413	17.88
1.30	19.953	17.901	3.70	43.01	41.94	750.7607	1128.17	0.11	0.18	1.313	18.13
1.20	15.849	14.219	5.20	84.95	31.95	454.2981	1878.93	0.18	0.25	1.213	18.38
1.10	12.589	11.295	6.10	116.90	20.37	320.4104	2333.23	0.22	0.30	1.113	18.63
1.00	10.000	8.972	6.80	145.27	26.77	240.1374	2653.64	0.25	0.33	1.013	18.88
0.90	7.943	7.126	7.40	172.03	24.03	171.2704	2873.78	0.28	0.36	0.913	19.13
0.80	6.310	5.661	7.90	196.07	41.72	236.1665	3065.05	0.29	0.38	0.813	19.38
0.70	5.012	4.496	8.70	237.79	45.74	205.6744	3371.21	0.31	0.42	0.713	19.63
0.60	3.981	3.572	9.50	283.53	56.27	290.9627	3506.89	0.33	0.46	0.613	19.88
0.50	3.162	2.837	10.40	339.79	61.36	174.0699	3707.85	0.35	0.50	0.513	20.13
0.40	2.512	2.254	11.30	401.15	105.56	237.8807	3881.92	0.37	0.55	0.413	20.38
0.30	1.995	1.790	12.70	506.71	172.16	308.1777	4119.80	0.39	0.62	0.313	20.63
0.20	1.585	1.422	14.70	678.87	283.25	402.7488	4427.97	0.42	0.71	0.213	20.88
0.10	1.259	1.129	17.50	962.11	332.50	375.5498	4830.72	0.46	0.85	0.113	21.13
-0.00	1.000	0.897	20.30	1294.62	295.81	265.3933	5206.27	0.50	0.98	0.013	21.38
-0.10	0.794	0.713	22.50	1590.43	500.74	356.8447	5471.66	0.52	1.09	-0.087	21.63
-0.20	0.631	0.566	25.80	2091.17	680.00	384.9272	5820.50	0.56	1.25	-0.187	21.88
-0.30	0.501	0.450	29.70	2771.17	989.82	445.0688	6213.43	0.59	1.44	-0.287	22.13
-0.40	0.398	0.357	34.60	3760.99	1190.44	425.1863	6658.49	0.63	1.68	-0.387	22.38
-0.50	0.316	0.284	39.70	4951.43	1495.39	424.2546	7083.68	0.68	1.93	-0.487	22.63
-0.60	0.251	0.225	45.30	6446.82	1004.06	226.2716	7507.93	0.72	2.20	-0.587	22.88
-0.70	0.200	0.179	48.70	7450.88	1574.78	281.8975	7734.20	0.74	2.36	-0.687	23.13
-0.80	0.158	0.142	53.60	9025.66	1109.86	157.8122	8016.09	0.76	2.60	-0.787	23.38
-0.90	0.126	0.113	56.80	10135.53	1061.39	119.8001	8173.90	0.78	2.75	-0.887	23.63
-1.00	0.100	0.090	59.70	11196.92	1751.59	157.1461	8293.78	0.79	2.90	-0.987	23.88
-1.10	0.079	0.071	64.20	12948.50	3473.50	247.5368	8450.93	0.81	3.11	-1.087	24.13
-1.20	0.063	0.057	72.30	16422.00	3583.77	202.8681	8678.46	0.83	3.51	-1.187	24.38
-1.30	0.050	0.045	79.80	20005.77	2906.35	130.6838	8931.33	0.85	3.87	-1.287	24.63
-1.40	0.040	0.036	85.40	22912.12	4965.23	177.3427	9032.01	0.86	4.14	-1.387	24.88
-1.50	0.032	0.028	94.20	27877.35	5000.31	141.8640	9209.35	0.88	4.57	-1.487	25.13
-1.60	0.025	0.023	102.30	32877.66	8022.12	180.7861	9351.21	0.89	4.96	-1.587	25.38
-1.70	0.020	0.018	114.10	40899.79	10572.07	189.2503	9532.00	0.91	5.53	-1.687	25.63
-1.80	0.016	0.014	128.00	51471.86	7837.51	111.4436	9721.25	0.93	6.21	-1.787	25.88
-1.90	0.013	0.011	137.40	59309.37	11565.07	130.6247	9832.69	0.94	6.66	-1.887	26.13
-2.00	0.010		150.20	70874.44			9961.31	0.95	7.28	-1.987	26.38
-∞							1.493.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NGC 4578

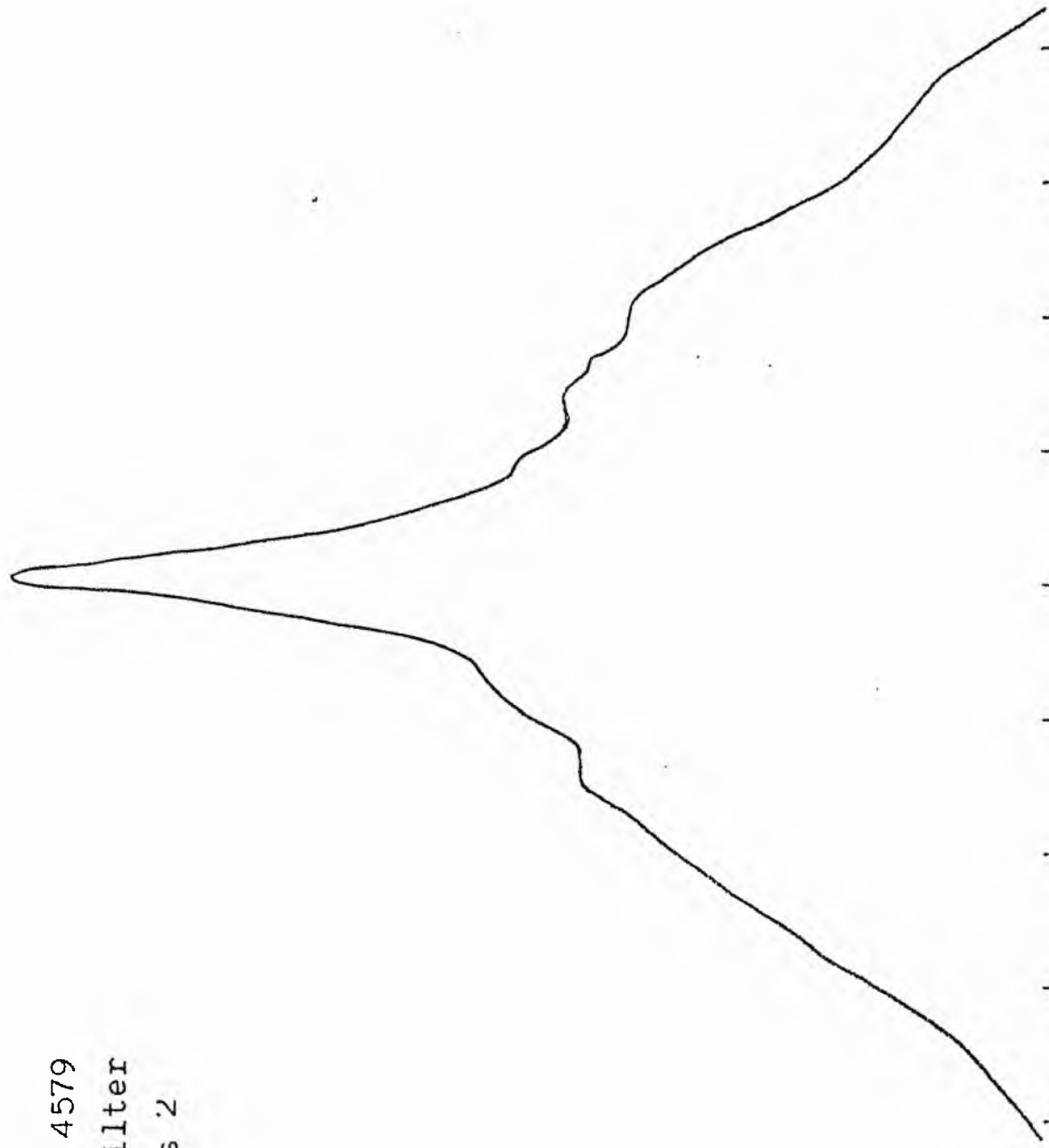
V-FILTER

Total luminosity	L_T	= 2.91
Total apparent magnitude	m_T	= 11.33
Apparent central surface brightness	μ_o	= 17.48
Major axis at threshold	$2a_m$	= 5.42
Minor axis at threshold	$2b_m$	= 4.95
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.76
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.87
Gradient of exponential component	$G(a)$	= -0.81
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.92
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.46
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.12
Axis ratio	b/a	= 1.00
Equivalent radius	r_1^*	= 0.11
Surface brightness	μ_1	= 18.88
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.37
Axis ratio	b/a	= 0.73
Equivalent radius	r_e^*	= 0.34
Surface brightness	μ_e	= 21.38
Mean surface brightness	μ_e'	= 10.98
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.11
Axis ratio	b/a	= 0.59
Equivalent radius	r_3^*	= 0.85
Surface brightness	μ_3	= 23.25
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 3.06 \\ = 2.47 \end{matrix}$

NGC 4579
B-Filter
Axis 1

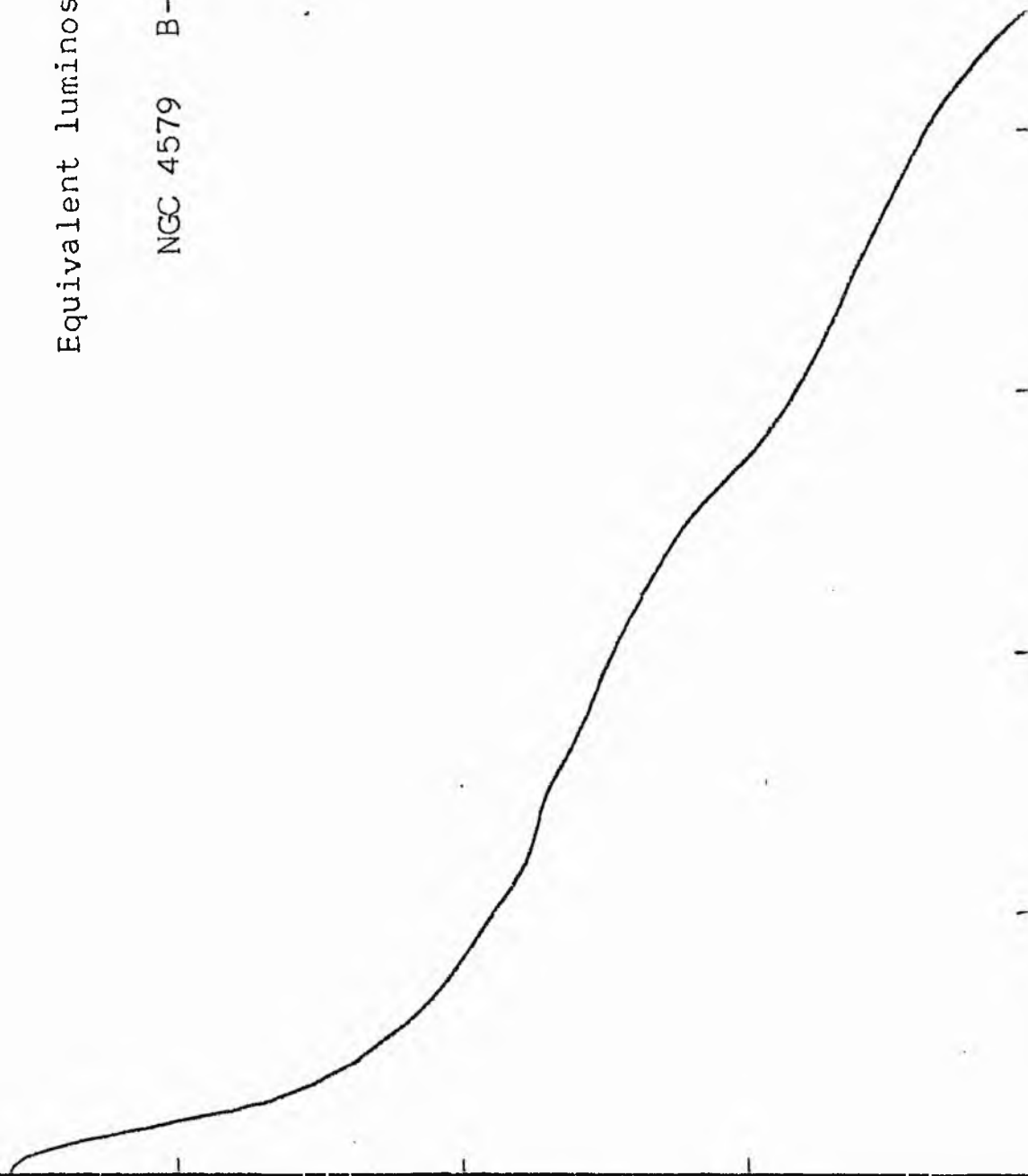


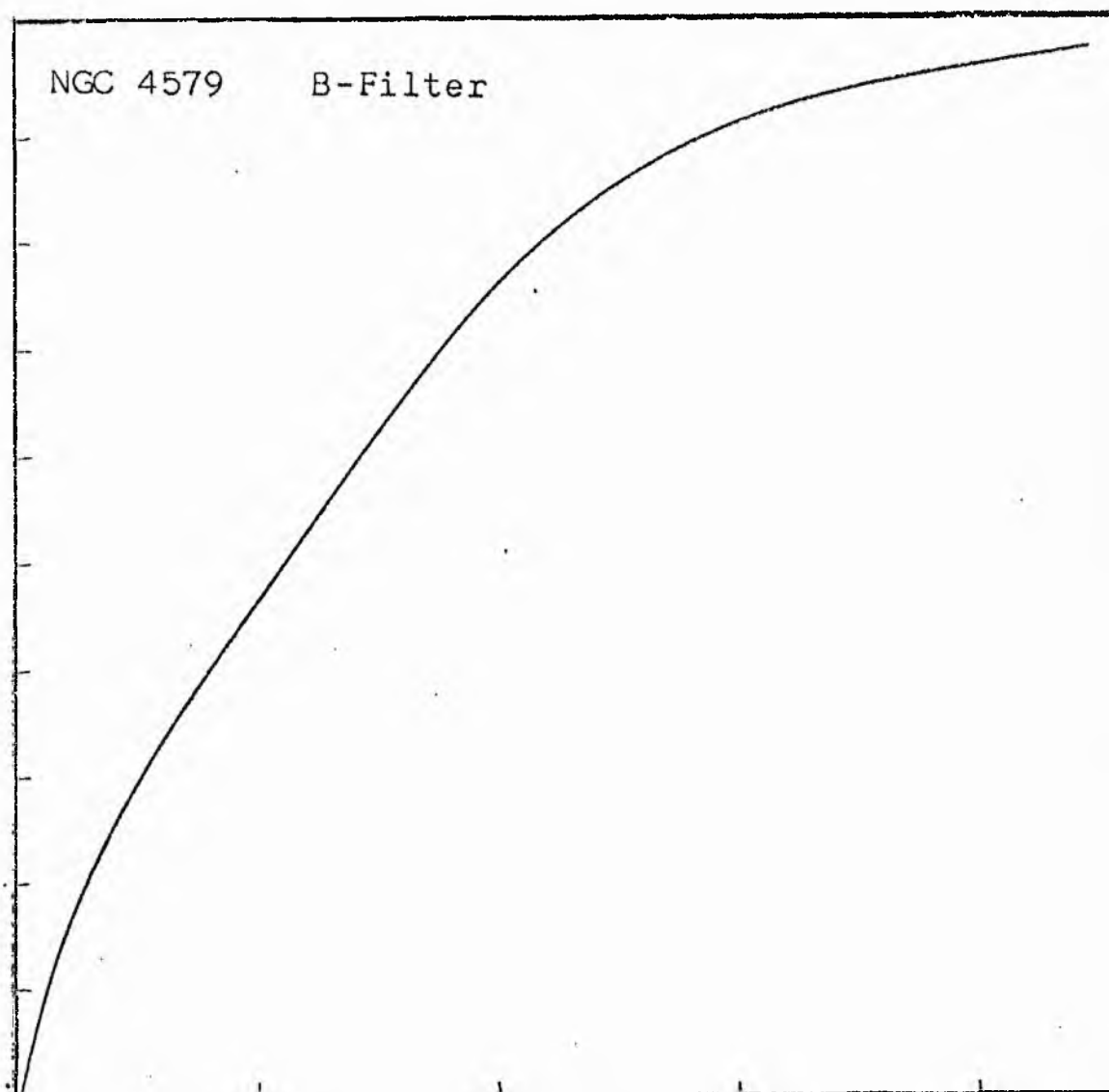
NGC 4579
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4579 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NCC 4579
O COLOUR

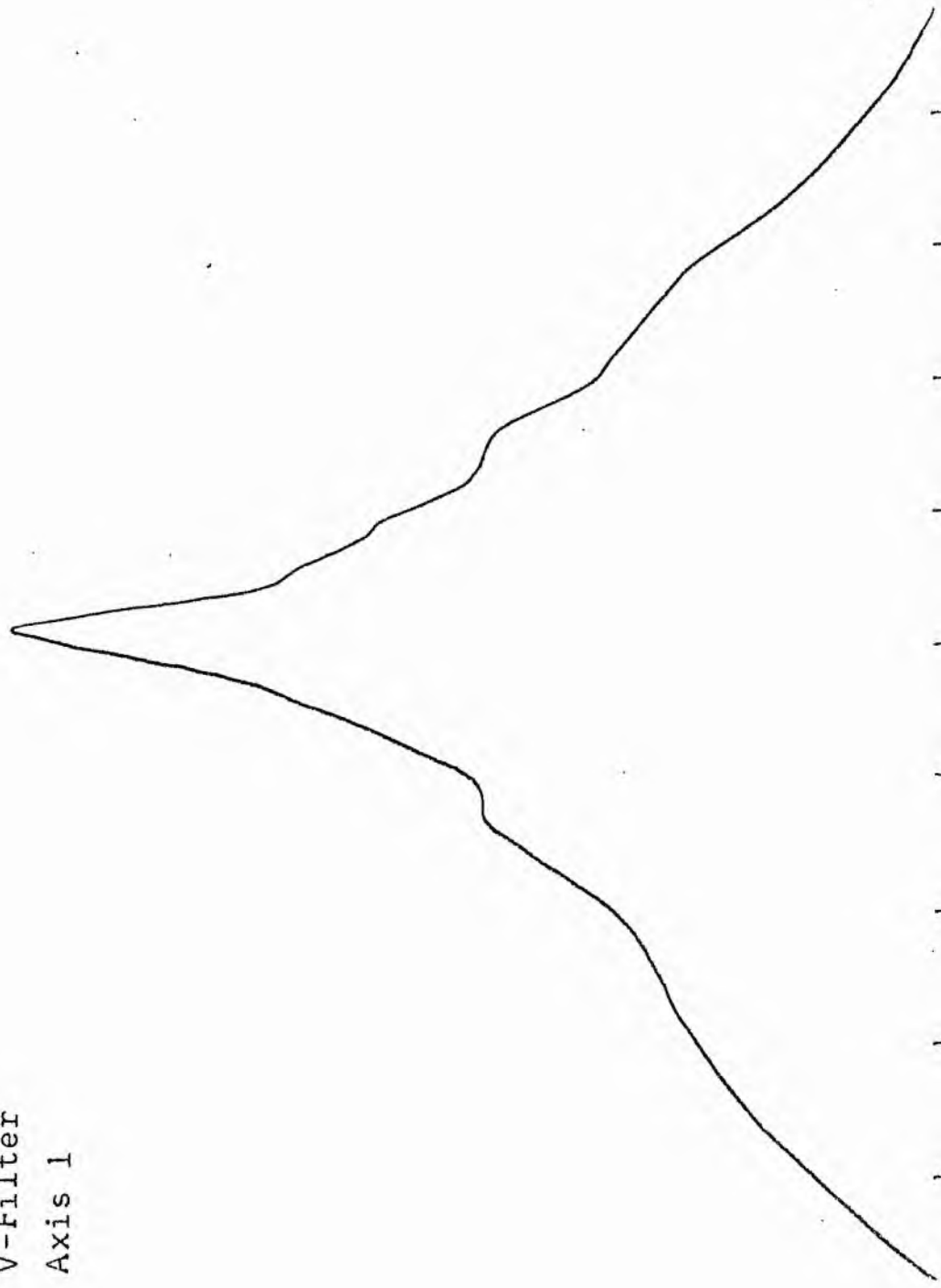
LCG 1	I	I	R	AREA	ΔA	P	ΣP	KINI	Q	LUG J	LL
1.56	36.308		0.0	0.0			0.0	0.0	0.0	1.728	17.09
1.50	31.623	33.965	3.29	34.00	34.00	1154.9854					
1.40	25.114	28.371	4.20	55.42	21.41	607.4961	1154.99	0.03	0.06	1.668	16.14
1.30	19.953	22.536	5.10	81.71	26.30	592.5781	1762.48	0.05	0.07	1.560	15.39
1.20	15.849	17.901	7.80	191.13	109.42	1958.1275	2355.06	0.06	0.09	1.468	14.64
1.10	12.589	14.219	8.50	226.98	35.85	509.6904	4313.79	0.11	0.14	1.368	14.09
1.00	10.000	11.295	9.50	283.53	56.55	638.6912	4823.47	0.13	0.15	1.268	14.14
0.90	7.943	8.972	10.40	339.79	56.27	504.7949	5462.16	0.14	0.17	1.168	14.34
0.80	6.310	7.126	11.60	422.73	82.94	591.0510	5966.95	0.15	0.19	1.068	14.64
0.70	5.012	5.661	13.30	555.72	132.98	752.7795	6558.00	0.17	0.21	0.968	14.69
0.60	3.981	4.496	15.10	716.31	160.60	722.1235	7310.18	0.19	0.24	0.868	20.14
0.50	3.162	3.572	17.50	962.11	245.80	877.9087	8032.90	0.21	0.27	0.768	20.39
0.40	2.512	2.837	20.80	1359.18	397.06	1126.4990	8910.81	0.23	0.31	0.668	20.64
0.30	1.995	2.254	24.30	1855.08	495.90	1117.5444	10037.10	0.26	0.37	0.568	20.69
0.20	1.585	1.790	27.20	2324.27	469.20	839.8479	11154.45	0.29	0.43	0.468	21.14
0.10	1.259	1.422	32.30	3277.59	953.31	1355.5225	11944.74	0.31	0.49	0.368	21.39
-0.00	1.000	1.129	39.70	4951.43	1673.84	1870.5344	13350.76	0.35	0.58	0.268	21.64
-0.10	0.794	0.897	48.90	7512.20	2560.77	2297.4272	15240.19	0.40	0.71	0.168	21.69
-0.20	0.631	0.713	59.60	11159.43	3647.23	2599.1650	17538.27	0.46	0.87	0.068	22.14
-0.30	0.501	0.566	74.56	17464.71	6305.28	3569.2339	20137.38	0.52	1.06	-0.032	22.39
-0.40	0.398	0.450	86.67	23548.66	8133.95	2758.1067	23706.61	0.62	1.33	-0.132	22.64
-0.50	0.316	0.357	98.34	30381.57	6782.91	2422.6277	26464.72	0.69	1.55	-0.232	22.69
-0.60	0.251	0.284	109.57	37716.65	7335.08	2081.6173	28887.34	0.75	1.76	-0.332	23.14
-0.70	0.200	0.225	116.27	42470.27	4753.62	1071.2625	30968.36	0.80	1.96	-0.432	23.39
-0.80	0.158	0.179	122.29	46982.02	4511.75	807.6360	32039.62	0.83	2.08	-0.532	23.64
-0.90	0.126	0.142	126.79	50503.30	3521.28	500.6934	32847.25	0.85	2.18	-0.632	23.69
-1.00	0.100	0.113	135.03	57280.97	6777.67	765.5117	33347.95	0.87	2.26	-0.732	24.14
-1.10	0.079	0.090	140.85	62325.18	5044.21	452.5483	34113.46	0.89	2.41	-0.832	24.39
-1.20	0.063	0.071	151.07	71697.87	9372.70	667.9402	34566.00	0.90	2.51	-0.932	24.64
-1.30	0.050	0.057	164.88	85405.44	13707.56	775.9492	35233.94	0.91	2.70	-1.032	24.69
-1.40	0.040	0.045	176.62	98000.75	12595.31	566.3469	36009.89	0.93	2.94	-1.132	25.14
-1.50	0.032	0.036	186.60	109388.81	11388.06	406.7468	36576.23	0.95	3.15	-1.232	25.39
-1.60	0.025	0.028	194.00	118236.94	8848.12	251.0304	36982.98	0.96	3.33	-1.332	25.64
-1.70	0.020	0.023	201.69	127746.37	9559.44	215.4310	37234.01	0.97	3.46	-1.432	25.89
-1.80	0.016	0.018	208.65	136768.62	8972.25	160.6119	37449.44	0.97	3.60	-1.532	26.14
-1.90	0.013	0.014	214.88	145058.00	8289.37	117.8687	37610.05	0.98	3.72	-1.632	26.39
-2.00	0.010	0.011	220.36	152551.06	7493.06	84.6324	37727.91	0.98	3.84	-1.732	26.64
							37812.54	0.98	3.93	-1.832	26.69
							38542.00	(1)			00

PHOTOMETRIC PARAMETERS OF NGC 4579

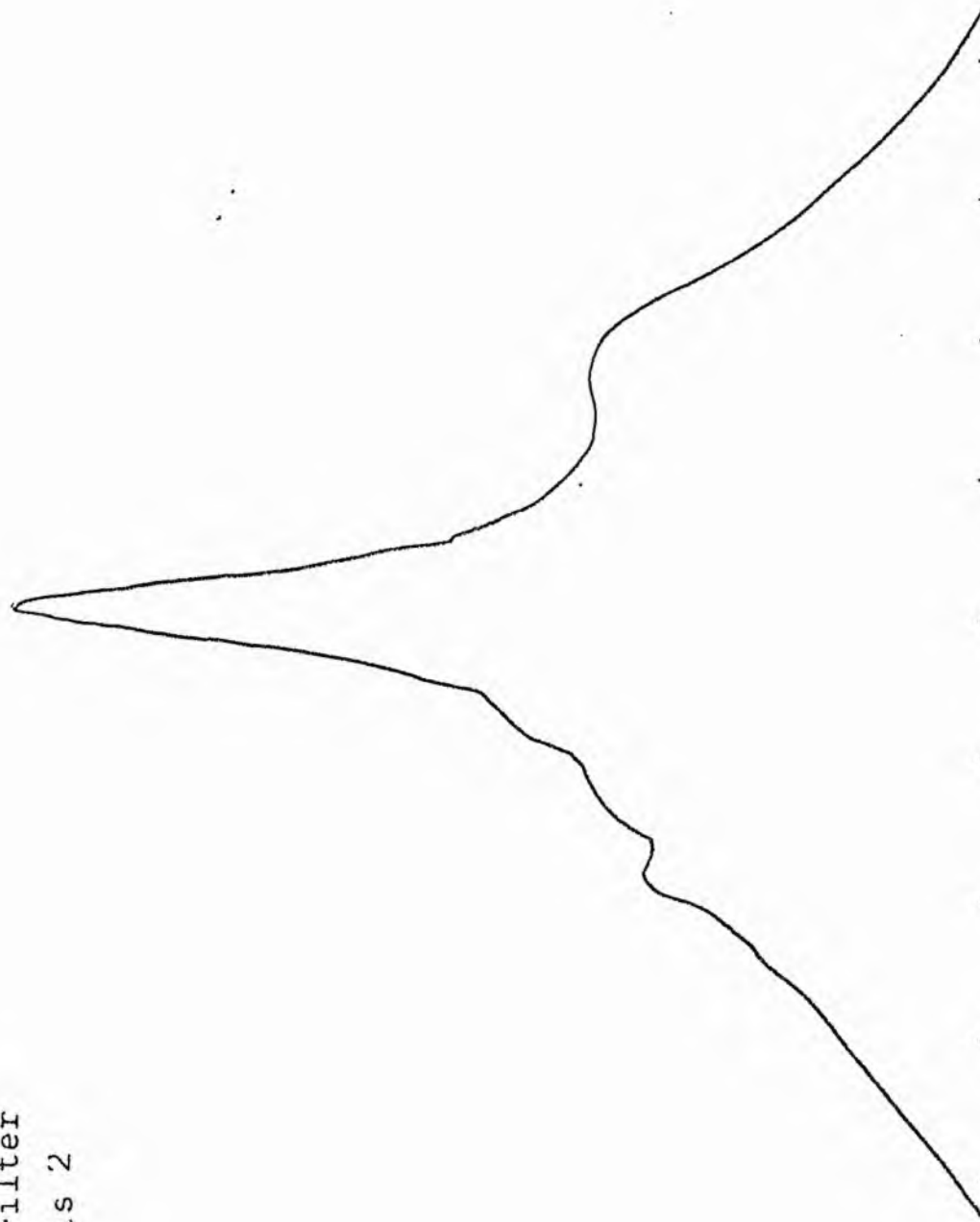
B-FILTER

Total luminosity	L_T	= 10.71
Total apparent magnitude	m_T	= 10.42
Apparent central surface brightness	μ_o	= 17.99
Major axis at threshold	$2a_m$	= 8.02
Minor axis at threshold	$2b_m$	= 6.75
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.38
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.92
Gradient of exponential component	$G(a)$	= -0.54
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.59
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.64
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.32
Axis ratio	b/a	= 0.94
Equivalent radius	r_1^*	= 0.33
Surface brightness	μ_1	= 20.81
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.92
Axis ratio	b/a	= 0.89
Equivalent radius	r_e^*	= 0.93
Surface brightness	μ_e	= 22.31
Mean surface brightness	μ_e'	= 12.26
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.56
Axis ratio	b/a	= 0.91
Equivalent radius	r_3^*	= 1.64
Surface brightness	μ_3	= 23.14
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.86 \\ = 1.76 \end{cases}$

NGC 4579
V-Filter
Axis 1

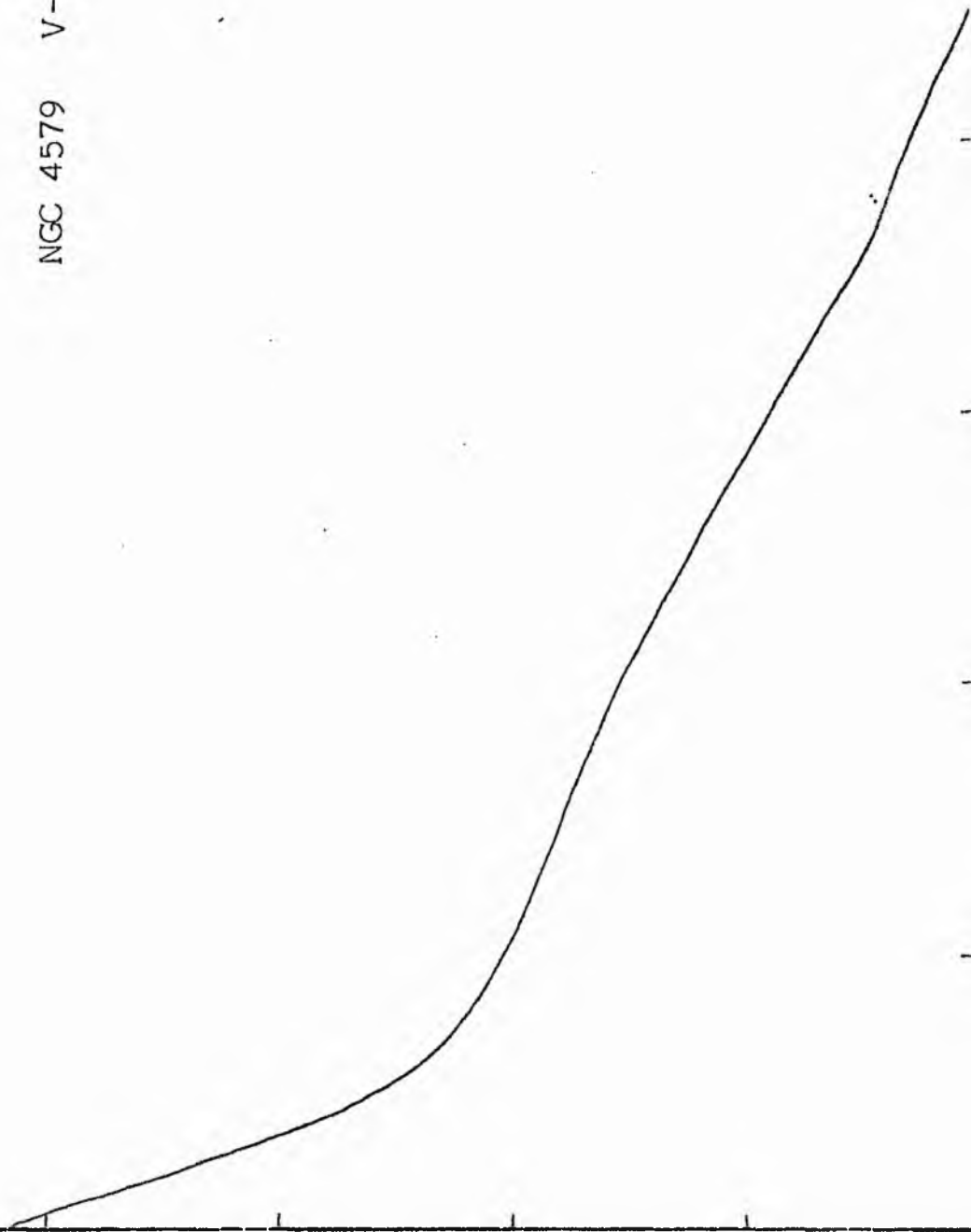


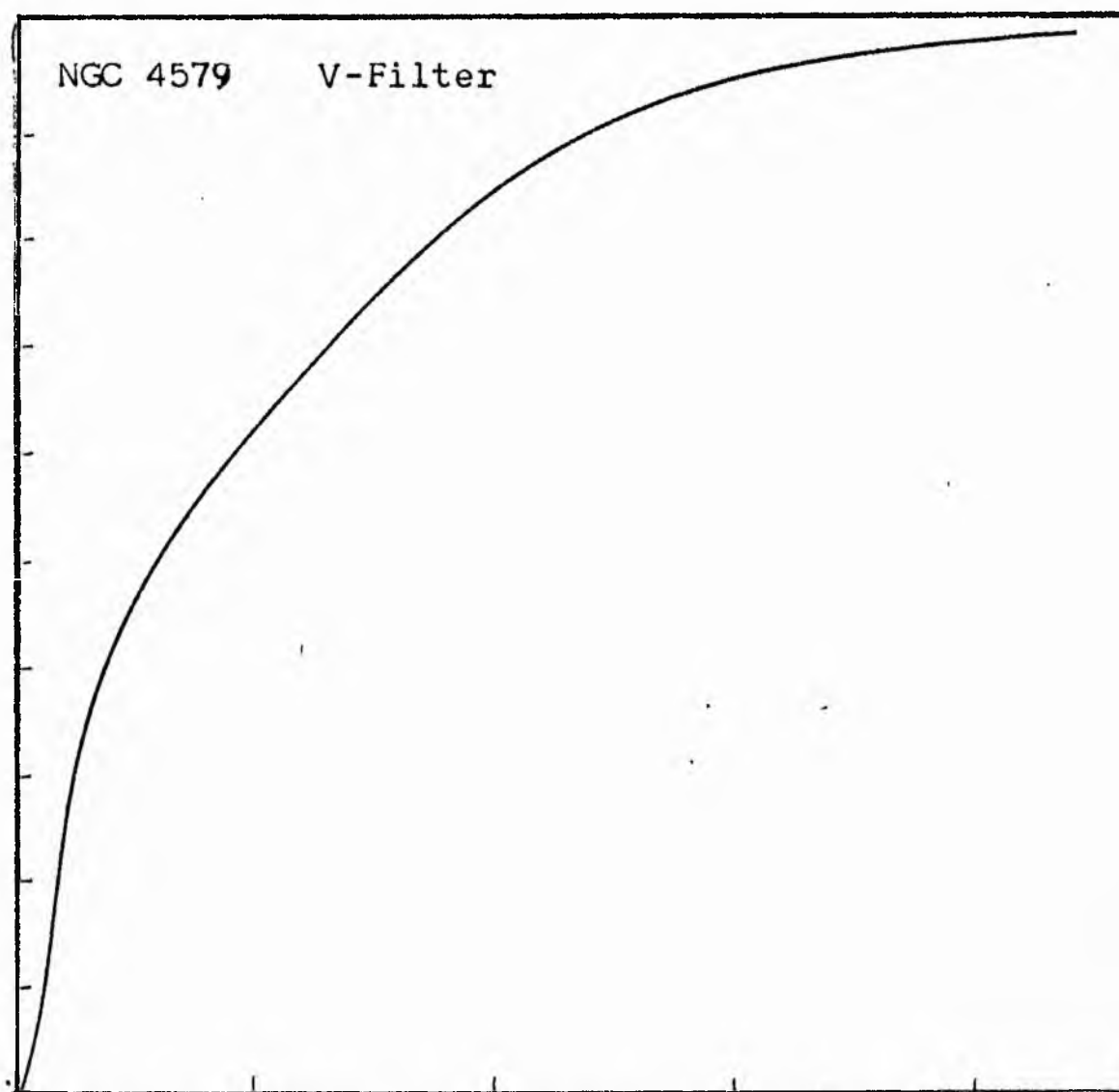
NGC 4579
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4579 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4579
V COLOUR

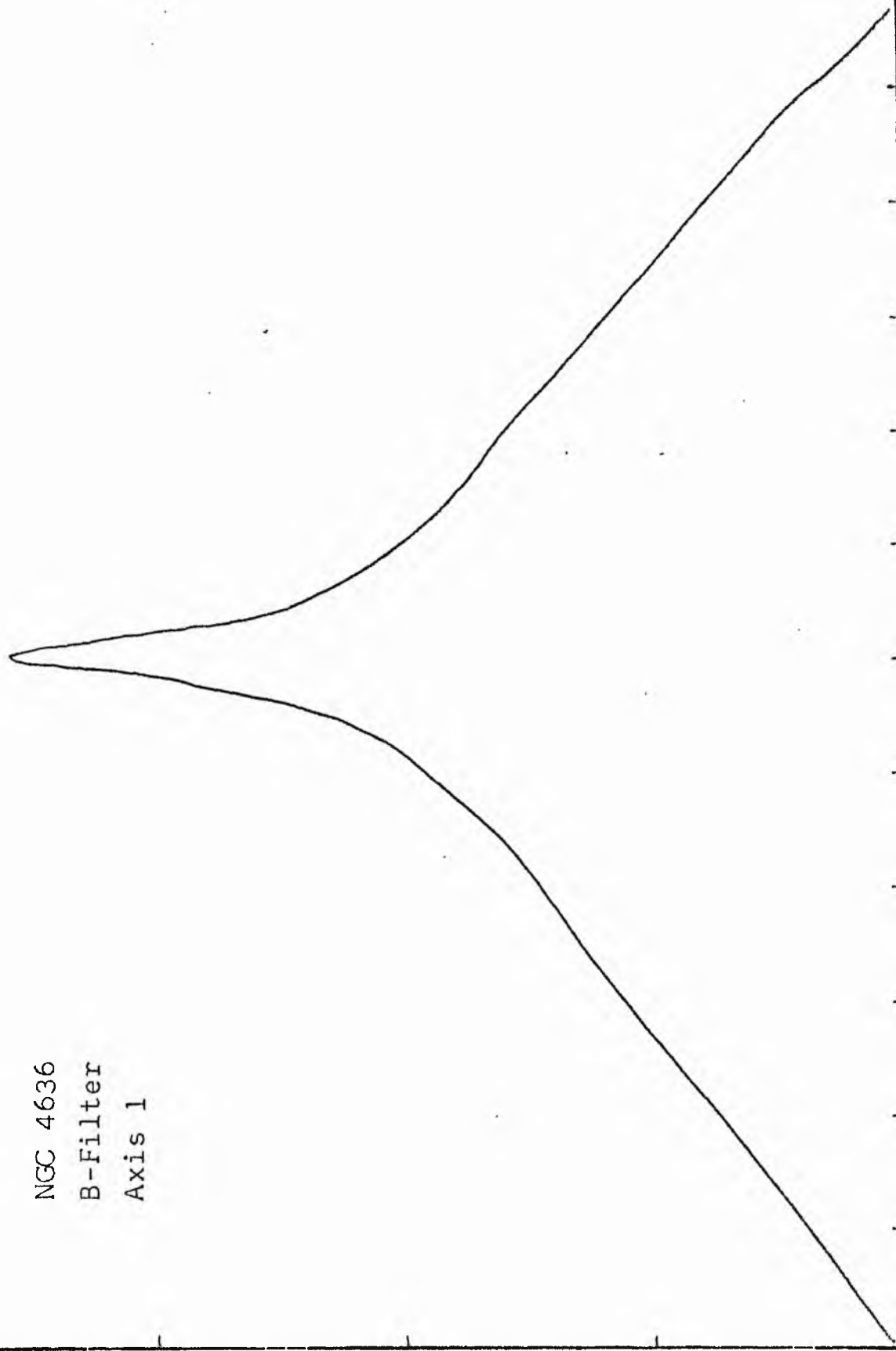
LOG I	I	I	R	AREA	AA	P	Σ P	K(R)	p	LOG J	μ
2.08	120.226		0.0	0.0			0.0	0.0	0.0	1.737	16.37
		110.113			4.52	478.1376					
2.00	100.000		1.20	4.52			448.14	0.01	0.04	1.657	16.57
		89.716			20.11	1803.8523					
1.90	79.433		2.80	24.63			2301.99	0.04	0.10	1.557	16.82
		71.264			16.08	1146.2812					
1.80	63.096		3.60	40.72			3440.27	0.06	0.13	1.457	17.07
		56.607			20.68	1623.6462					
1.70	50.119		4.70	69.40			5071.92	0.09	0.17	1.357	17.32
		44.965			32.67	1469.1069					
1.60	39.811		5.70	102.07			6541.02	0.11	0.20	1.257	17.57
		35.717			130.28	4653.2227					
1.50	31.623		8.60	232.35			11194.25	0.19	0.30	1.157	17.82
		28.371			114.01	3234.4961					
1.40	25.119		10.50	346.36			14428.74	0.25	0.37	1.057	18.07
		22.536			60.49	1363.1616					
1.30	19.952		11.38	406.85			15791.90	0.27	0.40	0.957	18.32
		17.901			84.02	1504.0828					
1.20	15.849		12.50	490.87			17295.98	0.30	0.44	0.857	18.57
		14.219			133.71	1901.1633					
1.10	12.589		14.10	624.58			19197.14	0.33	0.50	0.757	18.82
		11.295			63.55	717.8171					
1.00	10.000		14.80	688.13			19914.96	0.35	0.52	0.657	19.07
		8.972			146.55	1314.8162					
0.90	7.943		16.30	834.69			21229.77	0.37	0.57	0.557	19.32
		7.126			311.40	2219.1155					
0.80	6.310		19.10	1146.08			23448.89	0.41	0.67	0.457	19.57
		5.661			306.12	1732.8357					
0.70	5.012		21.50	1452.20			25181.72	0.44	0.76	0.357	19.82
		4.496			297.54	1337.8628					
0.60	3.981		23.60	1749.74			26519.58	0.46	0.83	0.257	20.07
		3.572			277.09	989.6643					
0.50	3.162		25.40	2026.83			27509.24	0.48	0.89	0.157	20.32
		2.837			179.35	508.8401					
0.40	2.512		26.50	2206.18			28018.08	0.49	0.93	0.057	20.57
		2.254			602.43	1357.6118					
0.30	1.995		29.90	2808.61			29375.69	0.51	1.05	-0.043	20.82
		1.790			974.15	1743.7830					
0.20	1.585		34.70	3782.76			31119.47	0.54	1.22	-0.143	21.07
		1.422			1706.35	2426.2571					
0.10	1.259		41.80	5489.11			33545.73	0.58	1.47	-0.243	21.32
		1.129			1931.20	2181.2061					
-0.00	1.000		48.60	7420.31			35726.93	0.62	1.71	-0.343	21.57
		0.897			2930.48	2629.0994					
-0.10	0.794		57.40	10350.79			38356.03	0.67	2.02	-0.443	21.82
		0.713			4911.34	3510.0039					
-0.20	0.631		69.70	15262.12			41856.03	0.73	2.45	-0.543	22.07
		0.566			4593.53	2600.2495					
-0.30	0.501		79.50	19855.65			44456.28	0.77	2.80	-0.643	22.32
		0.450			5140.86	2311.5571					
-0.40	0.398		89.20	24996.51			46767.83	0.81	3.14	-0.743	22.57
		0.357			5298.62	1892.4812					
-0.50	0.316		98.20	30295.13			48660.31	0.85	3.46	-0.843	22.82
		0.284			5538.61	1571.3401					
-0.60	0.251		106.80	35833.74			50231.65	0.87	3.76	-0.943	23.07
		0.225			5786.11	1303.9336					
-0.70	0.200		115.10	41619.85			51535.59	0.90	4.05	-1.043	23.32
		0.179			5754.86	1030.1575					
-0.80	0.158		122.80	47374.70			52565.74	0.91	4.32	-1.143	23.57
		0.142			5881.70	836.3196					
-0.90	0.126		130.20	53256.41			53402.06	0.93	4.59	-1.243	23.82
		0.113			5450.18	615.5745					
-1.00	0.100		136.70	58706.59			54017.63	0.94	4.81	-1.343	24.07
		0.090			8443.16	757.4871					
-1.10	0.079		146.20	67149.75			54775.12	0.95	5.15	-1.443	24.32
		0.071			7840.81	558.7683					
-1.20	0.063		154.50	74990.56			55333.88	0.96	5.44	-1.543	24.57
		0.057			6746.37	381.8931					
-1.30	0.050		161.30	81736.94			55715.77	0.97	5.68	-1.643	24.82
		0.045			4624.25	207.9281					
-1.40	0.040		165.80	86361.19			55923.70	0.97	5.84	-1.743	25.07
		0.036			8098.81	289.2637					
-1.50	0.032		173.40	94460.00			56212.96	0.98	6.11	-1.843	25.32
		0.028			8860.12	251.3697					
-1.60	0.025		181.35	103320.12			56464.33	0.98	6.39	-1.943	25.57
		0.023			15734.94	354.5999					
-1.70	0.020		194.67	119055.06			56818.93	0.99	6.86	-2.043	25.82
		0.018			6219.37	111.3323					
-1.80	0.016		199.69	125274.44			56930.26	0.99	7.03	-2.143	26.07
		0.014			10225.50	145.3984					
-1.90	0.013		207.68	135499.94			57075.65	0.99	7.31	-2.243	26.32
		0.011			9679.62	109.3287					
-2.00	0.010		214.97	145179.56			57184.98	0.99	7.57	-2.343	26.57
							57579.00	(1)			

PHOTOMETRIC PARAMETERS OF NGC 4579

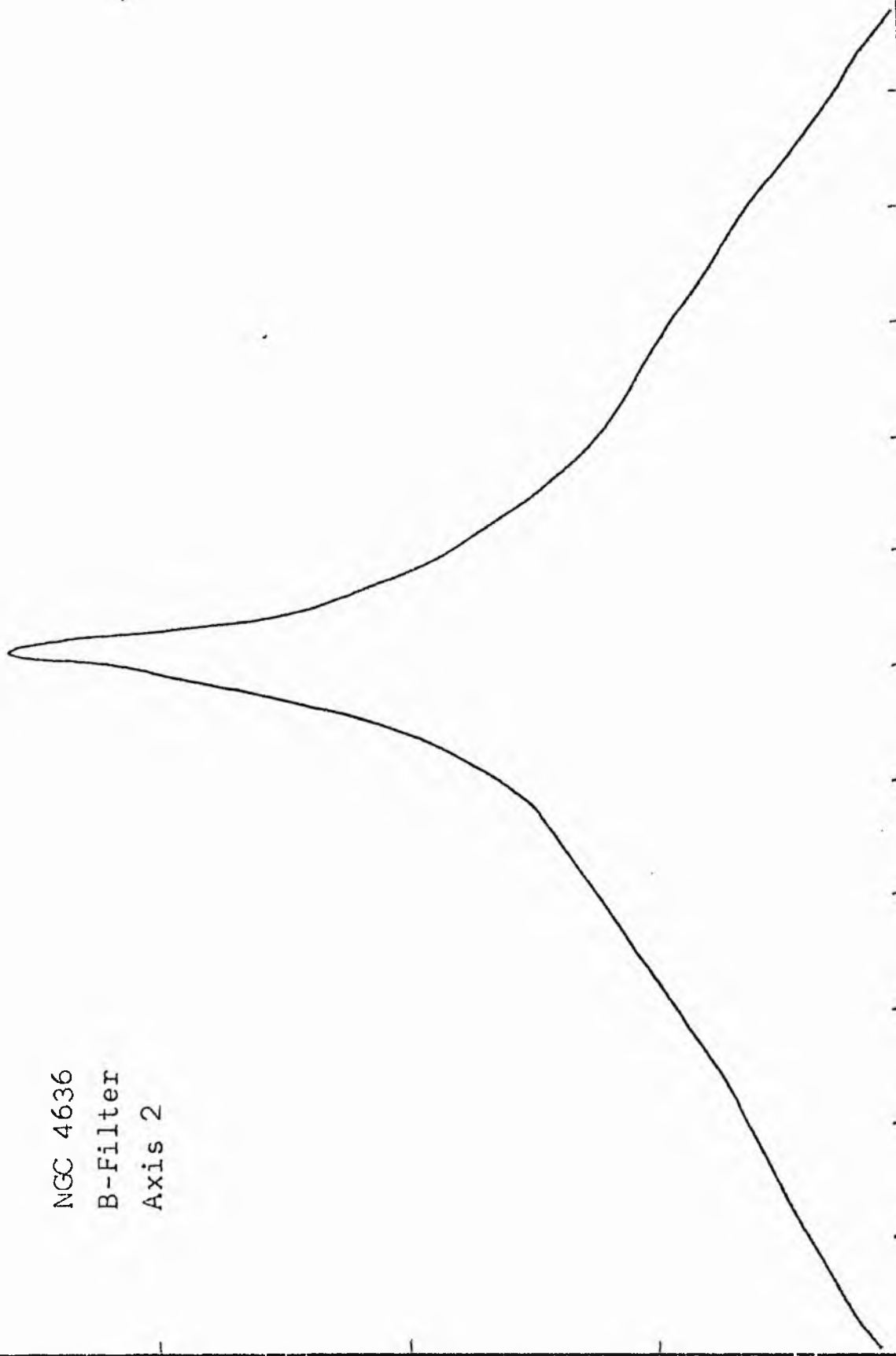
V-FILTER

Total luminosity	L_T	= 15.99
Total apparent magnitude	m_T	= 9.67
Apparent central surface brightness	μ_0	= 16.37
Major axis at threshold	$2a_m$	= 7.97
Minor axis at threshold	$2b_m$	= 6.95
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.62
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.97
Gradient of exponential component	$G(a)$	= -0.67
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.73
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.34
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.20
Axis ratio	b/a	= 0.83
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 18.07
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.75
Axis ratio	b/a	= 0.51
Equivalent radius	r_e^*	= 0.47
Surface brightness	μ_e	= 20.69
Mean surface brightness	μ_e'	= 10.02
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.37
Axis ratio	b/a	= 0.71
Equivalent radius	r_3^*	= 1.25
Surface brightness	μ_3	= 22.19
Concentration indices	$\{C_{21}$	= 2.71
	C_{32}	= 2.63

NGC 4636
B-Filter
Axis 1

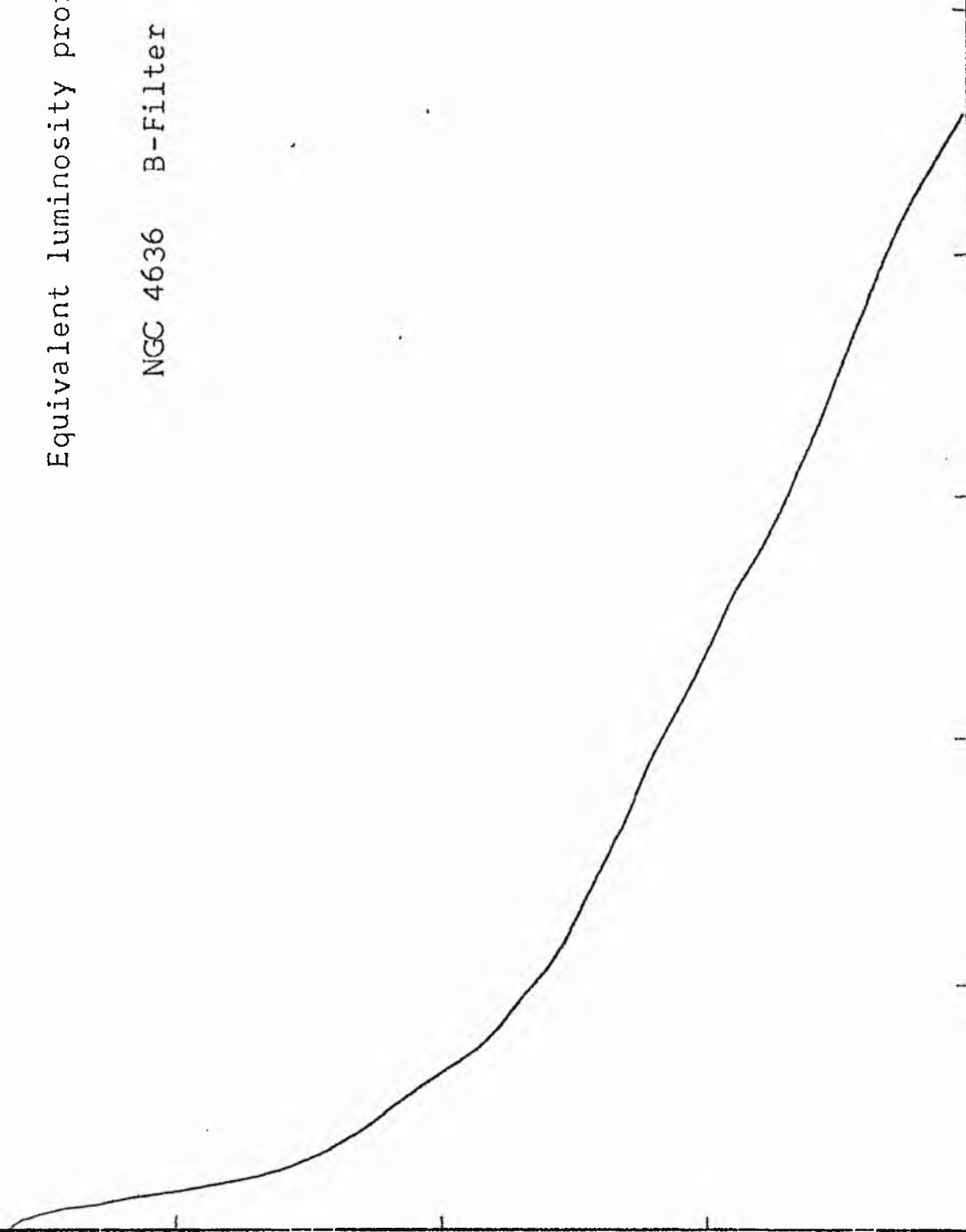


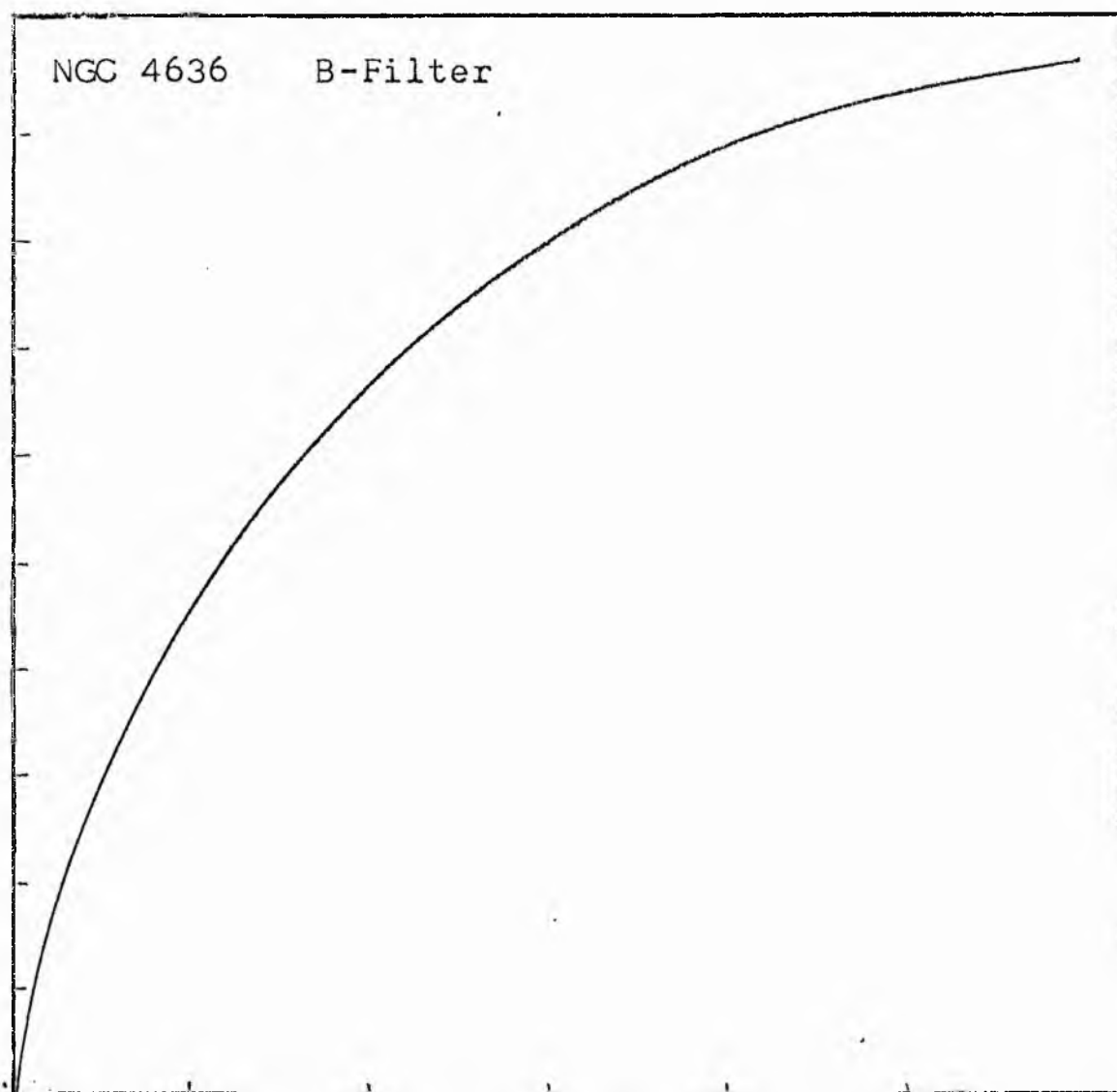
NGC 4636
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4636 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4636
B COLOUR

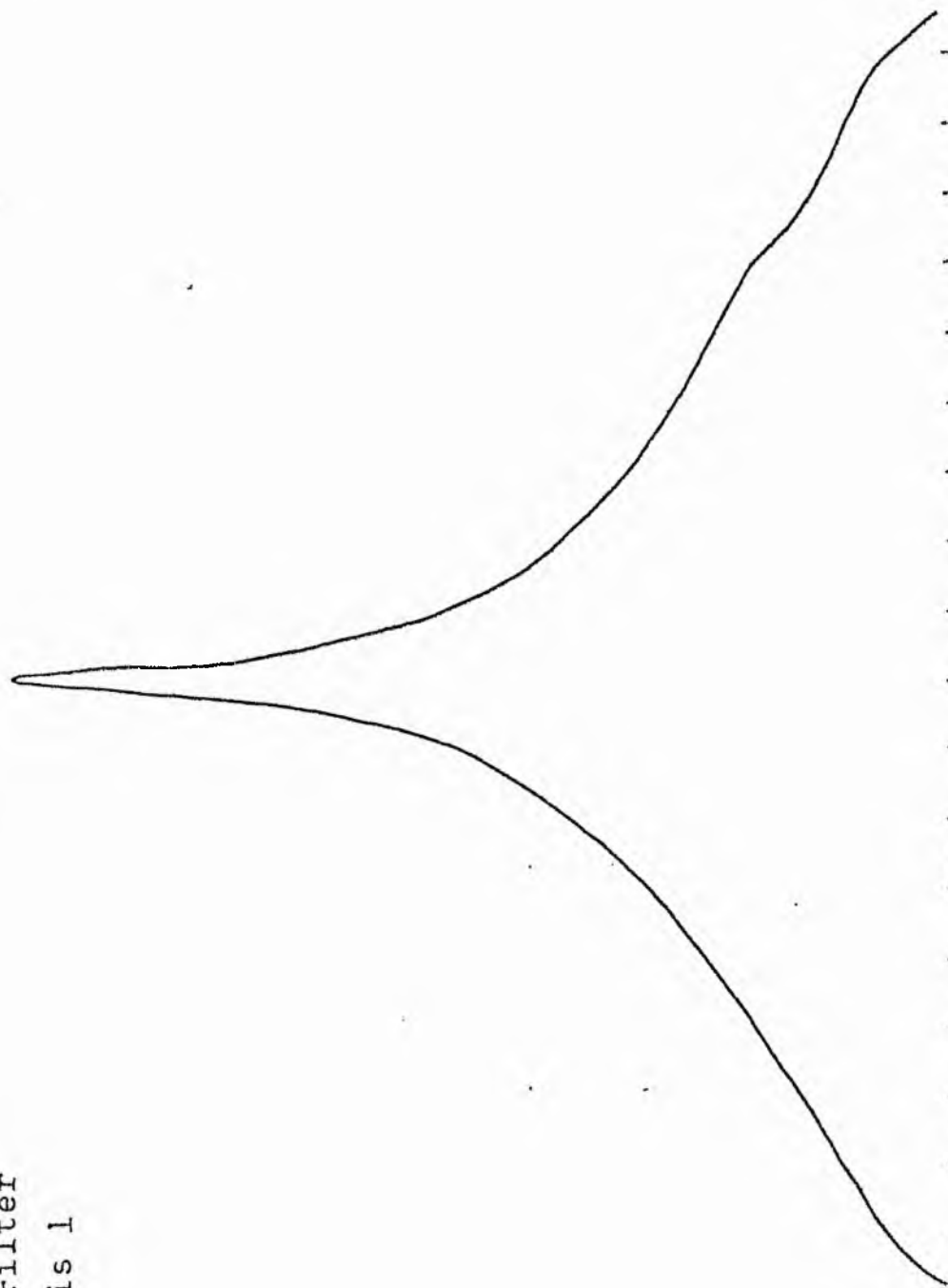
LOG I	I	T	R	AREA	ΔA	P	ΣP	KIRI	ρ	LOG J	μ
1.57	37.154		0.0	0.0	49.40	2386.4597	0.0	0.0	0.0	1.829	18.24
1.50	31.623	34.388	4.70	69.40	63.33	1796.8506	2386.46	0.05	0.08	1.759	18.42
1.40	25.119	28.371	6.50	132.73	34.68	781.6079	4183.31	0.09	0.11	1.659	18.67
1.30	19.953	22.536	7.30	167.42	49.01	877.2942	4964.91	0.11	0.13	1.559	18.92
1.20	15.849	17.901	8.30	216.42	67.10	954.1582	5842.21	0.13	0.14	1.459	19.17
1.10	12.589	14.219	9.50	283.53	49.76	562.0513	6796.36	0.15	0.16	1.359	19.42
1.00	10.000	11.295	10.30	333.29	134.30	1204.9128	7358.41	0.16	0.18	1.259	19.67
0.90	7.943	8.972	12.20	467.59	202.07	1440.0134	8563.32	0.19	0.21	1.159	19.92
0.80	6.310	7.126	14.60	669.66	144.67	818.9265	10003.34	0.22	0.25	1.059	20.17
0.70	5.012	5.661	16.10	814.33	125.92	566.1724	10822.26	0.24	0.28	0.959	20.42
0.60	3.981	4.496	17.30	940.25	111.84	399.4558	11388.43	0.25	0.30	0.859	20.67
0.50	3.162	3.572	18.30	1052.09	406.88	1154.3345	11787.89	0.26	0.32	0.759	20.92
0.40	2.512	2.837	21.55	1458.96	742.23	1672.6599	12942.22	0.28	0.37	0.659	21.17
0.30	1.995	2.254	26.47	2201.19	898.27	1607.9612	14614.87	0.32	0.46	0.559	21.42
0.20	1.585	1.790	31.41	3099.46	1051.60	1495.2742	16222.84	0.35	0.55	0.459	21.67
0.10	1.259	1.422	36.35	4151.05	1259.55	1422.6113	17718.11	0.39	0.63	0.359	21.92
-0.00	1.000	1.129	41.50	5410.61	1588.36	1425.0146	19140.72	0.42	0.72	0.259	22.17
-0.10	0.794	0.897	47.20	6948.96	2094.18	1492.3477	20565.73	0.45	0.82	0.159	22.42
-0.20	0.631	0.713	53.80	9093.14	2367.89	1340.3909	22058.12	0.48	0.93	0.059	22.67
-0.30	0.501	0.566	60.40	11461.03	3683.09	1656.0872	23398.52	0.51	1.05	-0.041	22.92
-0.40	0.398	0.450	69.43	15144.12	5113.14	1826.2417	25054.60	0.55	1.20	-0.141	23.17
-0.50	0.316	0.357	80.30	20257.26	8334.87	2364.6653	26880.84	0.58	1.39	-0.241	23.42
-0.60	0.251	0.284	95.40	28592.12	12375.39	2789.7812	29245.50	0.64	1.66	-0.341	23.67
-0.70	0.203	0.225	114.20	40971.51	12694.71	2272.4473	32035.29	0.70	1.98	-0.441	23.92
-0.80	0.158	0.179	130.70	53666.22	11695.16	1662.9424	34307.73	0.75	2.27	-0.541	24.17
-0.90	0.126	0.142	144.24	65361.38	11357.25	1282.7578	35970.67	0.79	2.50	-0.641	24.42
-1.00	0.100	0.113	156.27	76718.62	10300.12	924.0911	37253.43	0.82	2.71	-0.741	24.67
-1.10	0.079	0.090	166.43	87018.75	14395.94	1025.9185	38177.52	0.84	2.89	-0.841	24.92
-1.20	0.063	0.071	179.67	101414.69	18007.56	1019.3611	39203.44	0.86	3.12	-0.941	25.17
-1.30	0.050	0.057	194.97	119422.25	21081.62	947.9333	40222.80	0.88	3.38	-1.041	25.42
-1.40	0.040	0.045	211.48	140503.87	24489.06	874.6746	41170.73	0.90	3.67	-1.141	25.67
-1.50	0.032	0.036	229.17	164992.94	27573.62	782.2920	42045.40	0.92	3.98	-1.241	25.92
-1.60	0.025	0.028	247.58	192566.56	23216.25	523.2000	42827.69	0.94	4.30	-1.341	26.17
-1.70	0.020	0.023	262.08	215782.81	17893.94	320.3186	43350.89	0.95	4.55	-1.441	26.42
-1.80	0.016	0.018	272.73	233676.75	19212.00	273.1802	43671.20	0.96	4.73	-1.541	26.67
-1.90	0.013	0.014	283.72	252888.75	18251.12	206.1422	43944.38	0.96	4.92	-1.641	26.92
-2.00	0.010	0.011	293.78	271139.87			44150.52	0.97	5.10	-1.741	27.17
-∞							45700.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4636

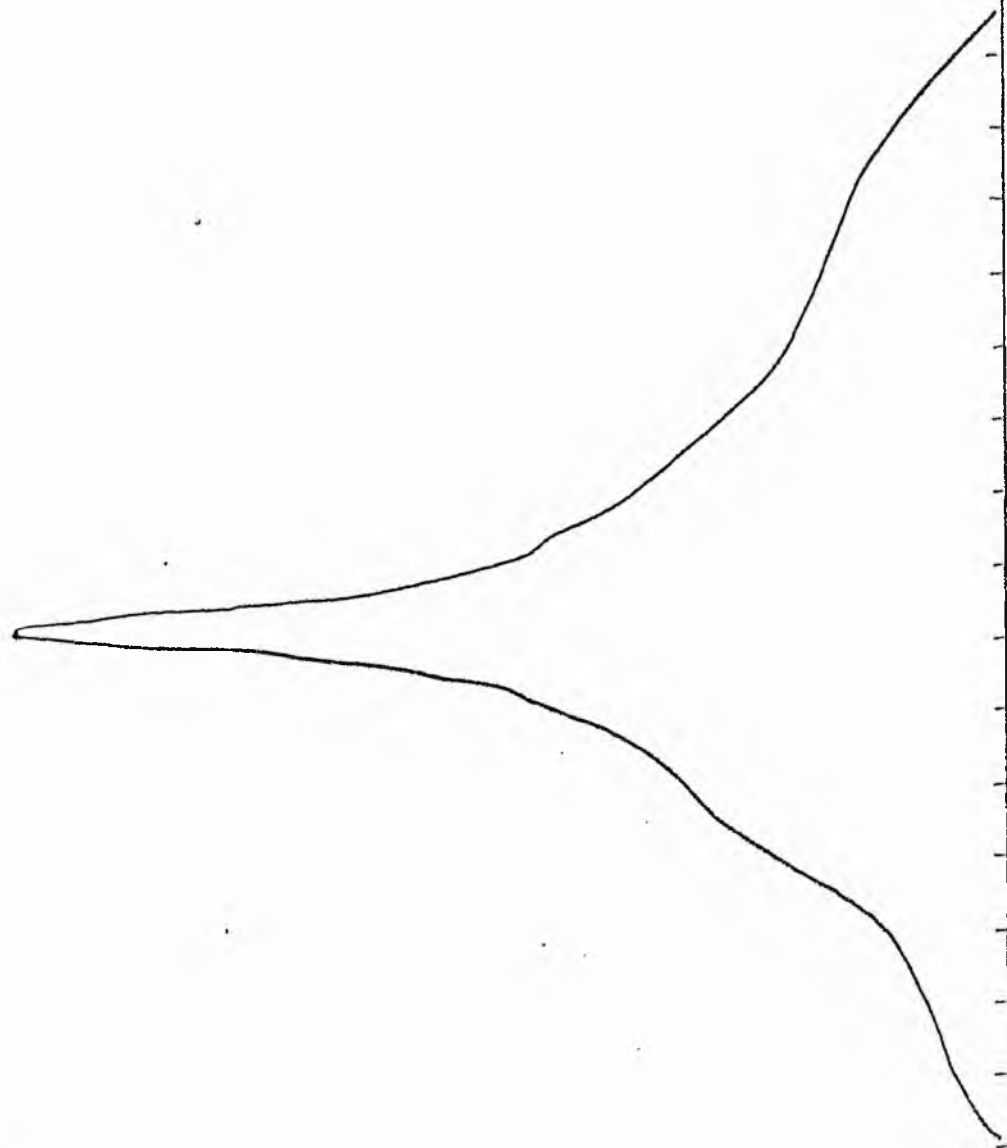
B-FILTER

Total luminosity	L_T	= 12.70
Total apparent magnitude	m_T	= 10.52
Apparent central surface brightness	μ_0	= 18.24
Major axis at threshold	$2a_m$	= 10.20
Minor axis at threshold	$2b_m$	= 9.97
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 6.17
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.85
Gradient of exponential component	$G(a)$	= -0.42
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.41
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.41
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.42
Axis ratio	b/a	= 0.44
Equivalent radius	r_1^*	= 0.29
Surface brightness	μ_1	= 20.67
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.17
Axis ratio	b/a	= 0.60
Equivalent radius	r_e^*	= 0.96
Surface brightness	μ_e	= 22.75
Mean surface brightness	μ_e'	= 12.43
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.35
Axis ratio	b/a	= 0.73
Equivalent radius	r_3^*	= 2.17
Surface brightness	μ_3	= 24.17
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.31 \\ = 2.26 \end{cases}$

NGC 4636
V-Filter
Axis 1

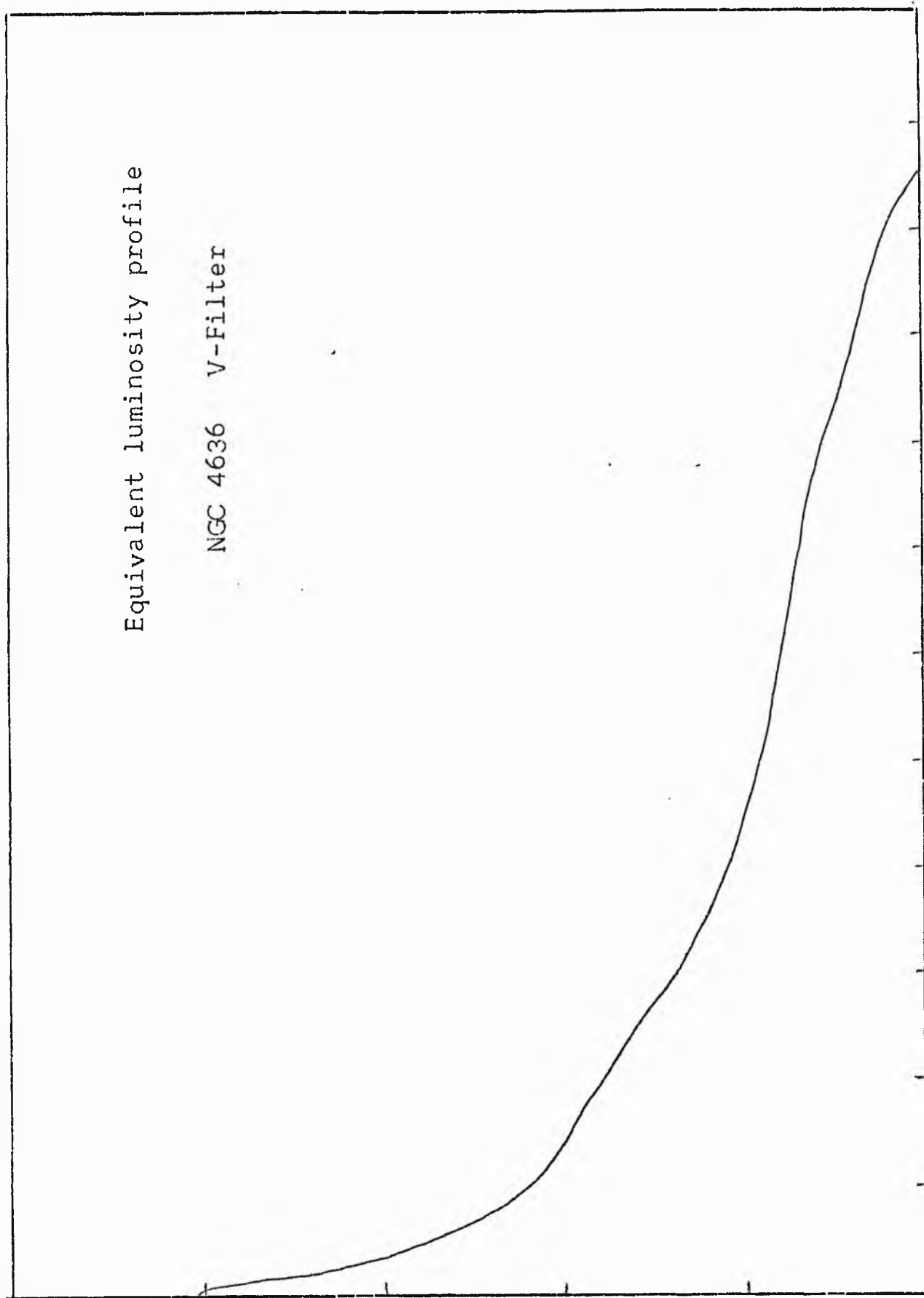


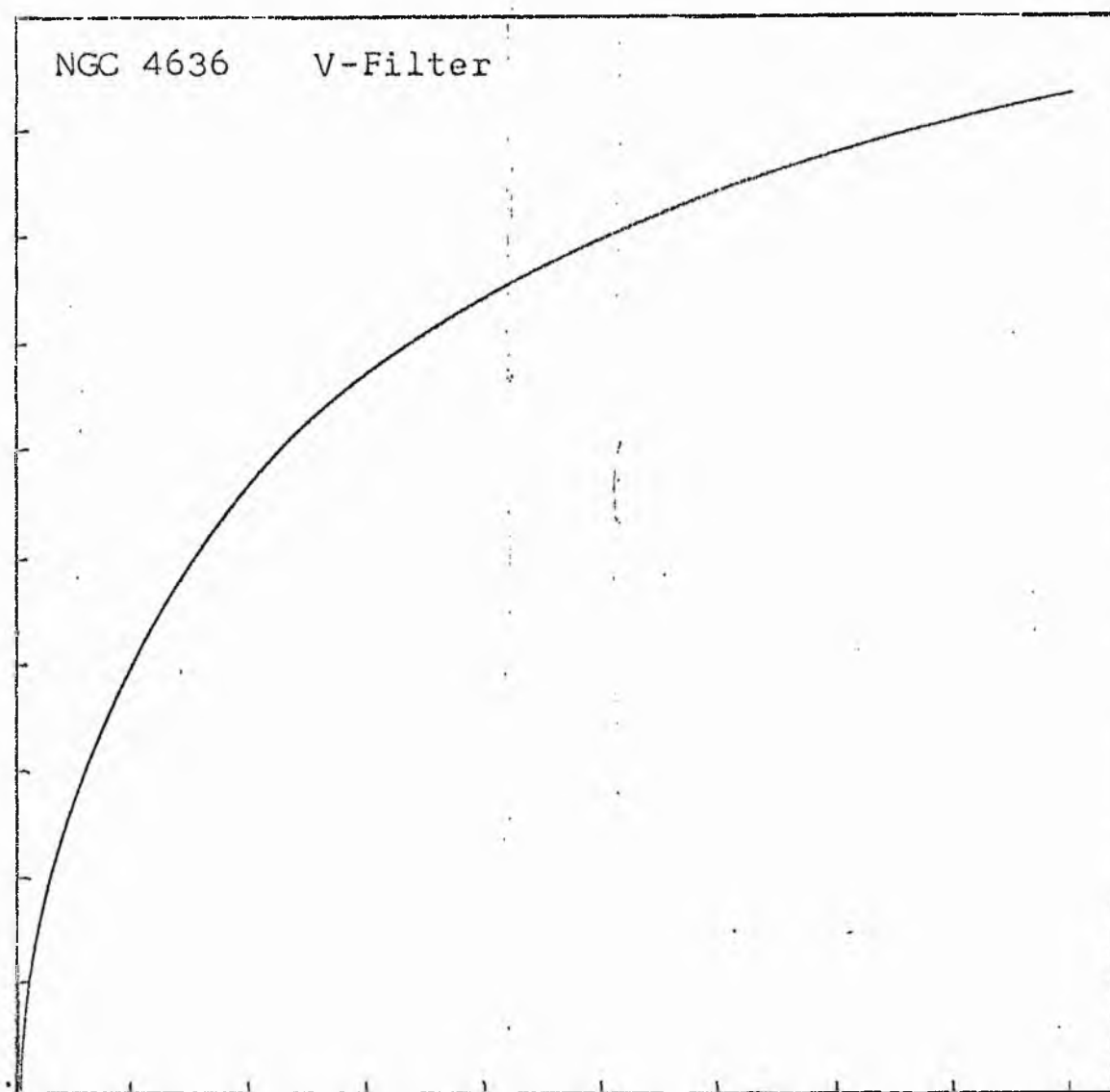
NGC 4636
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4636 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4636
V COLOUR

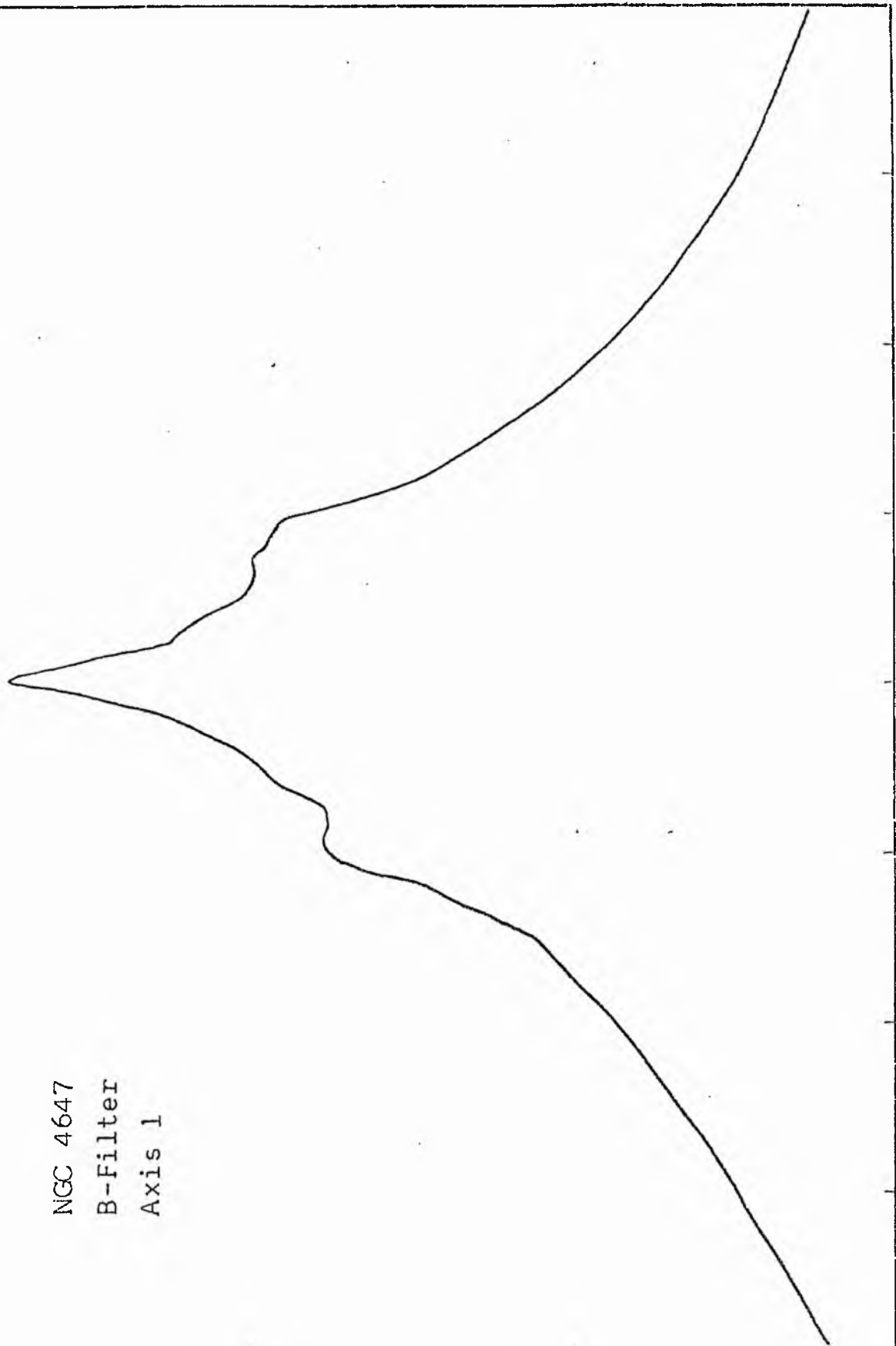
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	P	LOG J	μ
1.96	91.201		0.0	0.0			0.0	0.0	0.0	2.120	16.91
		85.317			6.16	525.3394					
1.90	79.433		1.40	6.16			525.34	0.01	0.02	2.060	17.06
		71.264			39.21	2794.0588					
1.80	63.096		3.80	45.36			3319.40	0.04	0.05	1.960	17.31
		56.607			42.88	2427.4648					
1.70	50.119		5.30	88.25			5746.86	0.07	0.07	1.860	17.56
		44.965			24.85	1117.3718					
1.60	39.811		6.00	113.10			6864.23	0.08	0.08	1.760	17.81
		35.717			68.36	2441.6226					
1.50	31.623		7.60	181.46			9305.85	0.11	0.10	1.660	18.06
		28.371			34.97	992.0071					
1.40	25.119		8.30	216.42			10297.86	0.12	0.11	1.560	18.31
		22.536			67.10	1512.2349					
1.30	19.952		9.50	283.53			11810.09	0.14	0.12	1.460	18.56
		17.901			62.83	1124.7341					
1.20	15.849		10.50	346.36			12934.82	0.15	0.14	1.360	18.81
		14.219			160.35	2279.9658					
1.10	12.589		12.70	506.71			15214.79	0.18	0.17	1.260	19.06
		11.295			207.15	2260.6118					
1.00	10.000		15.00	706.86			17475.39	0.20	0.20	1.160	19.31
		8.972			117.62	1055.2363					
0.90	7.943		16.20	824.48			18530.63	0.21	0.21	1.060	19.56
		7.126			137.63	980.8325					
0.80	6.310		17.50	962.11			19511.46	0.22	0.23	0.960	19.81
		5.661			278.24	1575.0276					
0.70	5.012		19.87	1240.35			21086.49	0.24	0.26	0.860	20.06
		4.496			468.11	2104.8452					
0.60	3.981		23.32	1708.47			23191.33	0.27	0.31	0.760	20.31
		3.572			539.54	1927.0420					
0.50	3.162		26.75	2248.01			25118.37	0.29	0.35	0.660	20.56
		2.837			573.77	1627.8274					
0.40	2.512		29.97	2821.78			26746.20	0.31	0.39	0.560	20.81
		2.254			758.81	1717.0251					
0.30	1.995		33.76	3580.59			28456.22	0.33	0.44	0.460	21.06
		1.790			409.89	733.7314					
0.20	1.585		35.64	3990.48			29189.95	0.34	0.47	0.360	21.31
		1.422			1839.95	2616.2263					
0.10	1.259		43.08	5830.43			31806.17	0.37	0.57	0.260	21.56
		1.129			4134.51	4669.7422					
-0.00	1.000		56.32	9964.95			36475.91	0.42	0.74	0.160	21.81
		0.897			4896.95	4393.3320					
-0.10	0.794		68.78	14861.89			40869.25	0.47	0.90	0.060	22.06
		0.713			5689.05	4054.2358					
-0.20	0.631		80.88	20550.95			44923.48	0.52	1.06	-0.040	22.31
		0.566			6399.08	3622.3167					
-0.30	0.501		92.62	26950.03			48545.80	0.56	1.22	-0.140	22.56
		0.450			4907.25	2206.5178					
-0.40	0.398		100.70	31857.29			50752.31	0.59	1.32	-0.240	22.81
		0.357			4542.36	1622.3738					
-0.50	0.316		107.64	36399.65			52374.68	0.60	1.41	-0.340	23.06
		0.284			7886.79	2237.5315					
-0.60	0.251		118.73	44286.43			54612.21	0.63	1.56	-0.440	23.31
		0.225			9232.06	2080.5010					
-0.70	0.200		130.52	53518.50			56692.71	0.65	1.72	-0.540	23.56
		0.179			6435.54	1510.0186					
-0.80	0.158		140.43	61954.04			58202.73	0.67	1.85	-0.640	23.81
		0.142			17728.46	2520.8105					
-0.90	0.126		159.26	79682.50			60723.54	0.70	2.09	-0.740	24.06
		0.113			24745.81	2794.9324					
-1.00	0.100		182.32	104428.31			63518.47	0.73	2.40	-0.840	24.31
		0.090			23064.06	2069.2170					
-1.10	0.079		201.45	127492.37			65587.62	0.76	2.65	-0.940	24.56
		0.071			44455.12	3168.0537					
-1.20	0.063		233.95	171947.50			68755.62	0.79	3.07	-1.040	24.81
		0.057			73649.81	4169.1055					
-1.30	0.050		279.60	245597.31			72924.69	0.84	3.67	-1.140	25.06
		0.045			49665.25	2233.1846					
-1.40	0.040		306.57	295262.56			75157.81	0.87	4.03	-1.240	25.31
		0.036			32577.37	1163.5596					
-1.50	0.032		323.04	327839.94			76321.31	0.88	4.25	-1.340	25.56
		0.028			68079.31	1931.4712					
-1.60	0.025		355.00	395919.25			78252.75	0.90	4.67	-1.440	25.81
		0.023			73208.69	1649.8191					
-1.70	0.020		386.43	469127.94			79902.56	0.92	5.08	-1.540	26.06
		0.018			70264.87	1257.8030					
-1.80	0.016		414.36	539392.81			81160.31	0.94	5.45	-1.640	26.31
		0.014			17505.00	248.9070					
-1.90	0.013		421.03	556897.81			81409.19	0.94	5.53	-1.740	26.56
		0.011			27662.50	312.4402					
-2.00	0.010		431.36	584560.31			81721.62	0.94	5.67	-1.840	26.81
-∞							86721.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4636

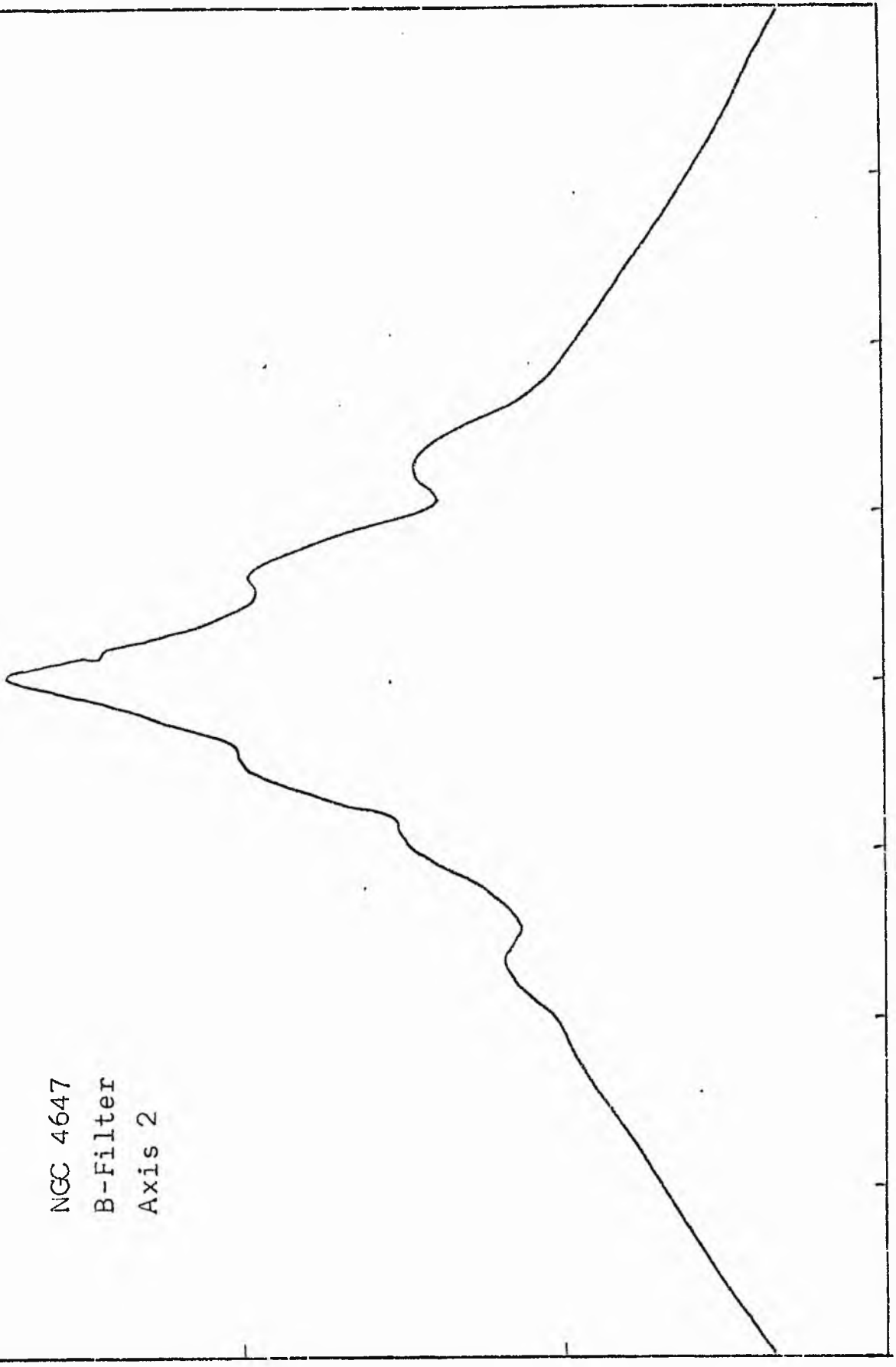
V-FILTER

Total luminosity	L_T	= 24.09
Total apparent magnitude	m_T	= 9.46
Apparent central surface brightness	μ_0	= 16.91
Major axis at threshold	$2a_m$	= 15.37
Minor axis at threshold	$2b_m$	= 13.41
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 10.27
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.23
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.26
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.74
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.48
Axis ratio	b/a	= 0.86
Equivalent radius	r_1^*	= 0.35
Surface brightness	μ_1	= 20.14
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.38
Axis ratio	b/a	= 0.72
Equivalent radius	r_e^*	= 1.27
Surface brightness	μ_e	= 22.21
Mean surface brightness	μ_e'	= 11.97
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 4.17
Axis ratio	b/a	= 0.70
Equivalent radius	r_3^*	= 3.27
Surface brightness	μ_3	= 24.48
Concentration indices	$\left\{ \begin{matrix} C_{21} \\ C_{32} \end{matrix} \right.$	= 3.66
		= 2.58

NGC 4647
B-Filter
Axis 1

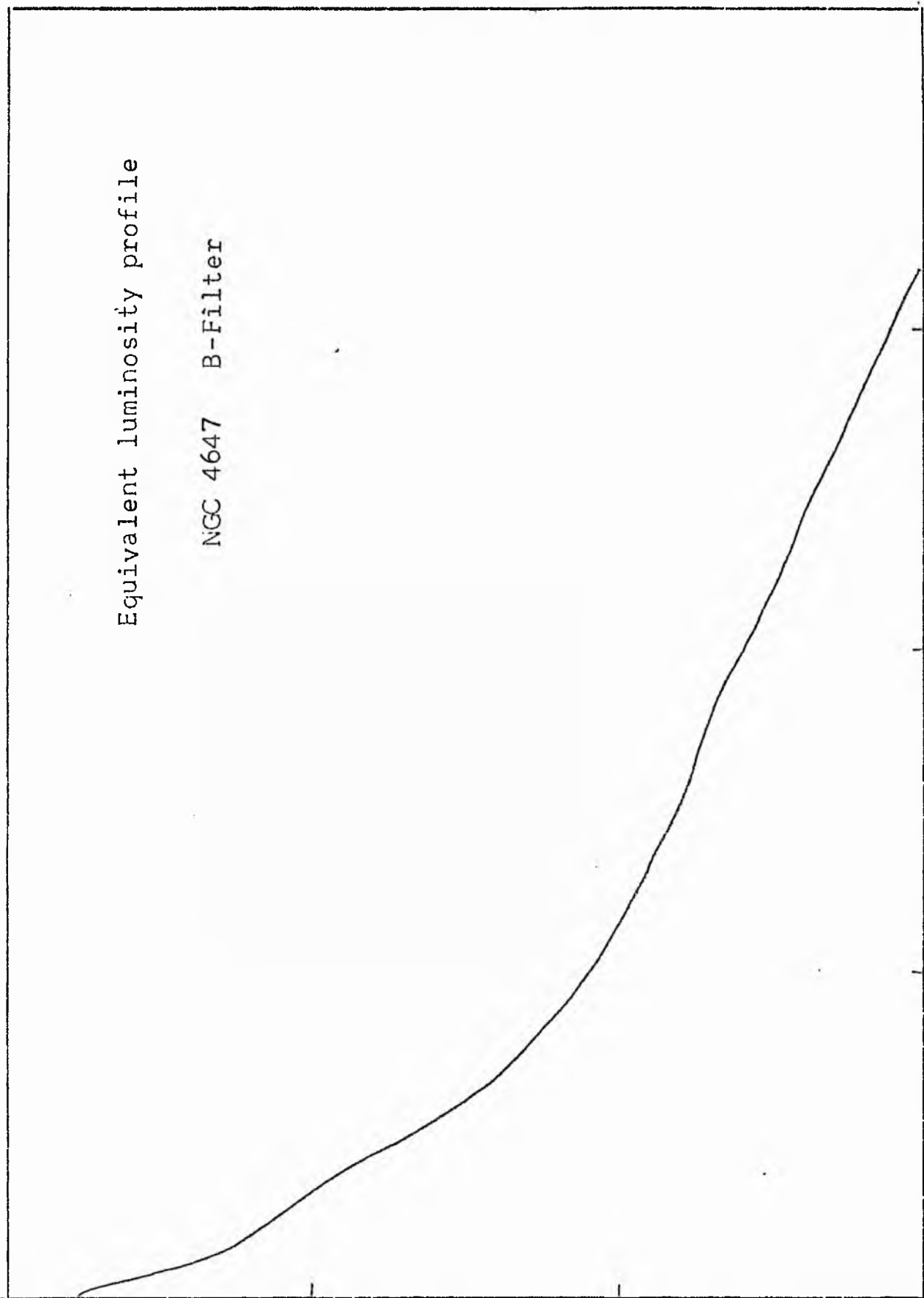


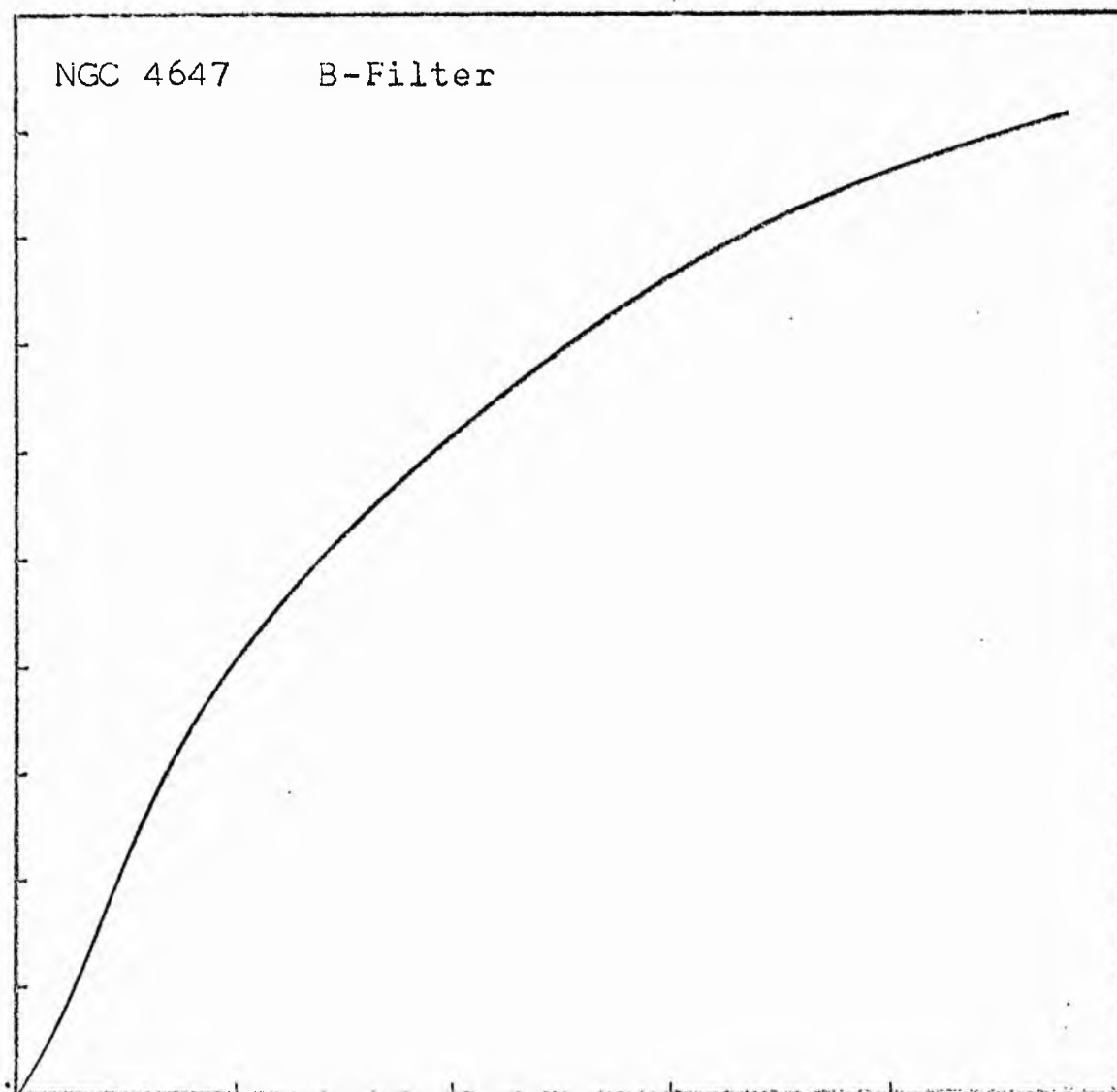
NGC 4647
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4647 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4647
B COLOUR

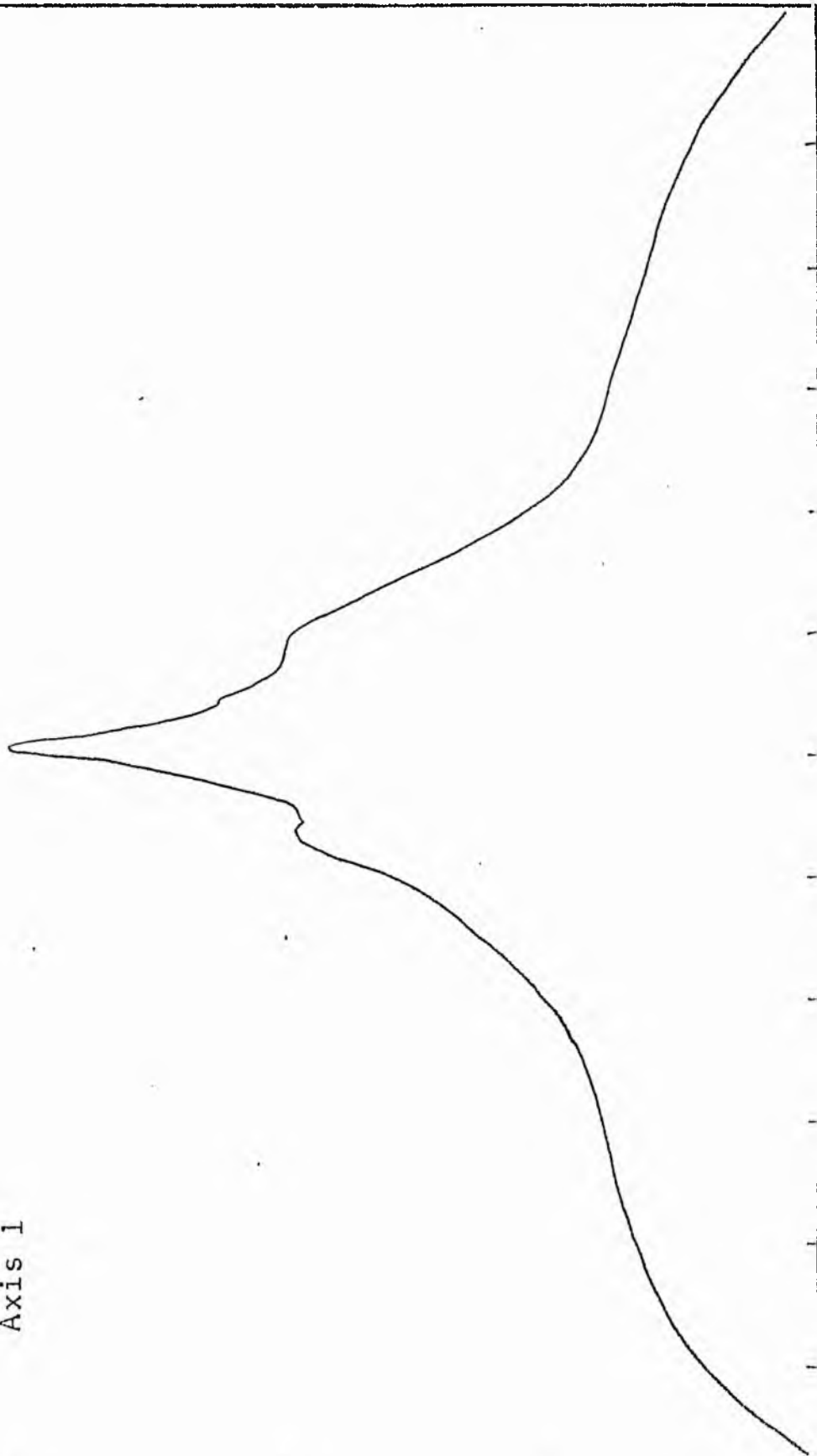
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
0.74	5.495		0.0	0.0			0.0	0.0	0.0	1.597	20.29
0.70	5.012	5.254	1.94	11.02	11.02	62.1174	62.12	0.00	0.03	1.557	20.39
0.60	3.481	4.496	4.62	67.06	55.23	248.3475	310.46	0.02	0.07	1.457	20.64
0.50	3.162	3.572	5.24	86.26	19.21	68.5947	379.06	0.03	0.08	1.357	20.89
0.40	2.512	2.837	7.83	192.61	106.35	301.7148	680.77	0.05	0.11	1.257	21.14
0.30	1.995	2.254	10.20	326.85	134.24	302.5276	983.30	0.07	0.15	1.157	21.39
0.20	1.585	1.790	13.02	532.56	205.71	368.2419	1351.54	0.09	0.19	1.057	21.64
0.10	1.254	1.422	18.19	1039.48	506.91	720.7834	2072.33	0.14	0.26	0.957	21.89
0.00	1.000	1.129	24.78	1929.09	889.61	1004.7852	3077.11	0.21	0.36	0.857	22.14
-0.10	0.794	0.897	29.48	2730.26	801.18	718.7861	3795.90	0.26	0.43	0.757	22.39
-0.20	0.631	0.713	32.27	3271.50	541.24	385.7107	4181.61	0.29	0.47	0.657	22.64
-0.30	0.501	0.566	35.84	4035.39	763.89	432.4158	4614.02	0.32	0.52	0.557	22.89
-0.40	0.398	0.450	39.61	4929.00	893.61	401.8103	5015.83	0.35	0.57	0.457	23.14
-0.50	0.316	0.357	43.50	5944.68	1015.67	362.7654	5378.59	0.37	0.63	0.357	23.39
-0.60	0.251	0.284	47.80	7178.03	1233.36	349.9133	5728.50	0.40	0.69	0.257	23.64
-0.70	0.200	0.225	56.20	9922.53	2744.50	618.4934	6347.00	0.44	0.81	0.157	23.89
-0.80	0.158	0.179	63.10	12508.59	2586.07	462.9265	6809.92	0.47	0.91	0.057	24.14
-0.90	0.126	0.142	74.00	17203.36	4694.77	667.5537	7477.47	0.52	1.07	-0.043	24.39
-1.00	0.100	0.113	87.90	24273.22	7069.86	798.5159	8275.99	0.58	1.27	-0.143	24.64
-1.10	0.079	0.090	98.30	30356.86	6083.63	545.8027	8821.79	0.61	1.42	-0.243	24.89
-1.20	0.063	0.071	112.40	39690.12	9333.27	665.1311	9486.92	0.66	1.63	-0.343	25.14
-1.30	0.050	0.057	132.50	55154.59	15464.47	875.4055	10362.32	0.72	1.92	-0.443	25.39
-1.40	0.040	0.045	149.60	70309.31	15154.72	681.4324	11043.75	0.77	2.16	-0.543	25.64
-1.50	0.032	0.036	159.30	79722.56	9413.25	336.2134	11379.96	0.79	2.30	-0.643	25.89
-1.60	0.025	0.028	176.40	97756.81	18034.25	511.6511	11891.61	0.83	2.55	-0.743	26.14
-1.70	0.020	0.023	188.30	111391.06	13634.25	307.2612	12198.87	0.85	2.72	-0.843	26.39
-1.80	0.016	0.018	202.10	128316.44	16925.37	302.9810	12501.85	0.87	2.92	-0.943	26.64
-1.90	0.013	0.014	220.60	152883.56	24567.12	349.3269	12851.17	0.89	3.19	-1.043	26.89
-2.00	0.010	0.011	237.00	176460.12	23576.56	266.2922	13117.46	0.91	3.43	-1.143	27.14
-∞							14367.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4647

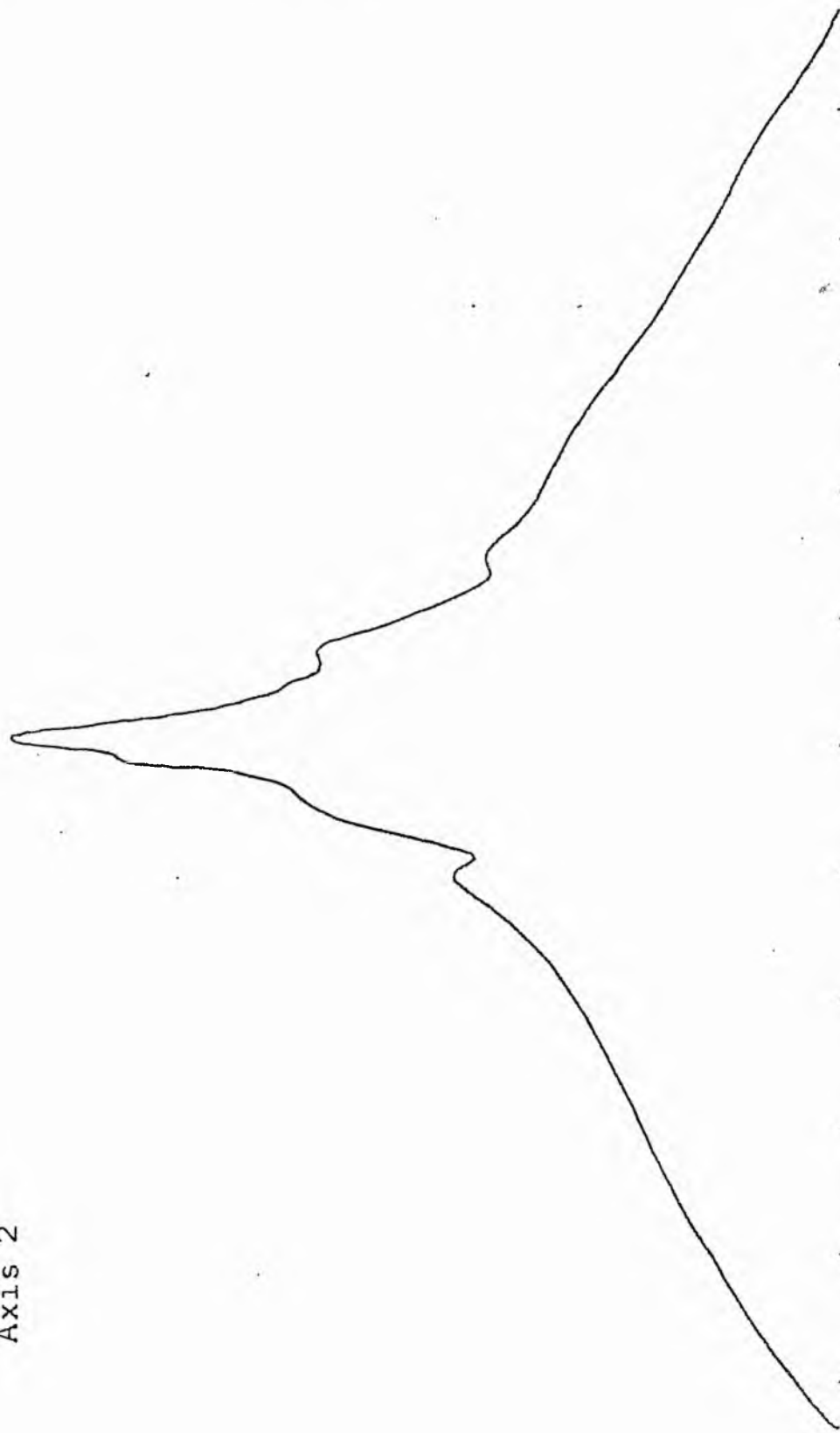
B-FILTER

Total luminosity	L_T	= 3.99
Total apparent magnitude	m_T	= 11.75
Apparent central surface brightness	μ_0	= 20.29
Major axis at threshold	$2a_m$	= 7.03
Minor axis at threshold	$2b_m$	= 7.13
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.37
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.64
Gradient of exponential component	$G(a)$	= -0.74
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.30
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.46
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.57
Axis ratio	b/a	= 0.82
Equivalent radius	r_1^*	= 0.47
Surface brightness	μ_1	= 22.34
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.25
Axis ratio	b/a	= 1.05
Equivalent radius	r_e^*	= 1.15
Surface brightness	μ_e	= 24.29
Mean surface brightness	μ_e'	= 14.04
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.15
Axis ratio	b/a	= 0.98
Equivalent radius	r_3^*	= 2.38
Surface brightness	μ_3	= 25.54
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.46 \\ = 2.06 \end{matrix}$

NGC 4647
V-Filter
Axis 1

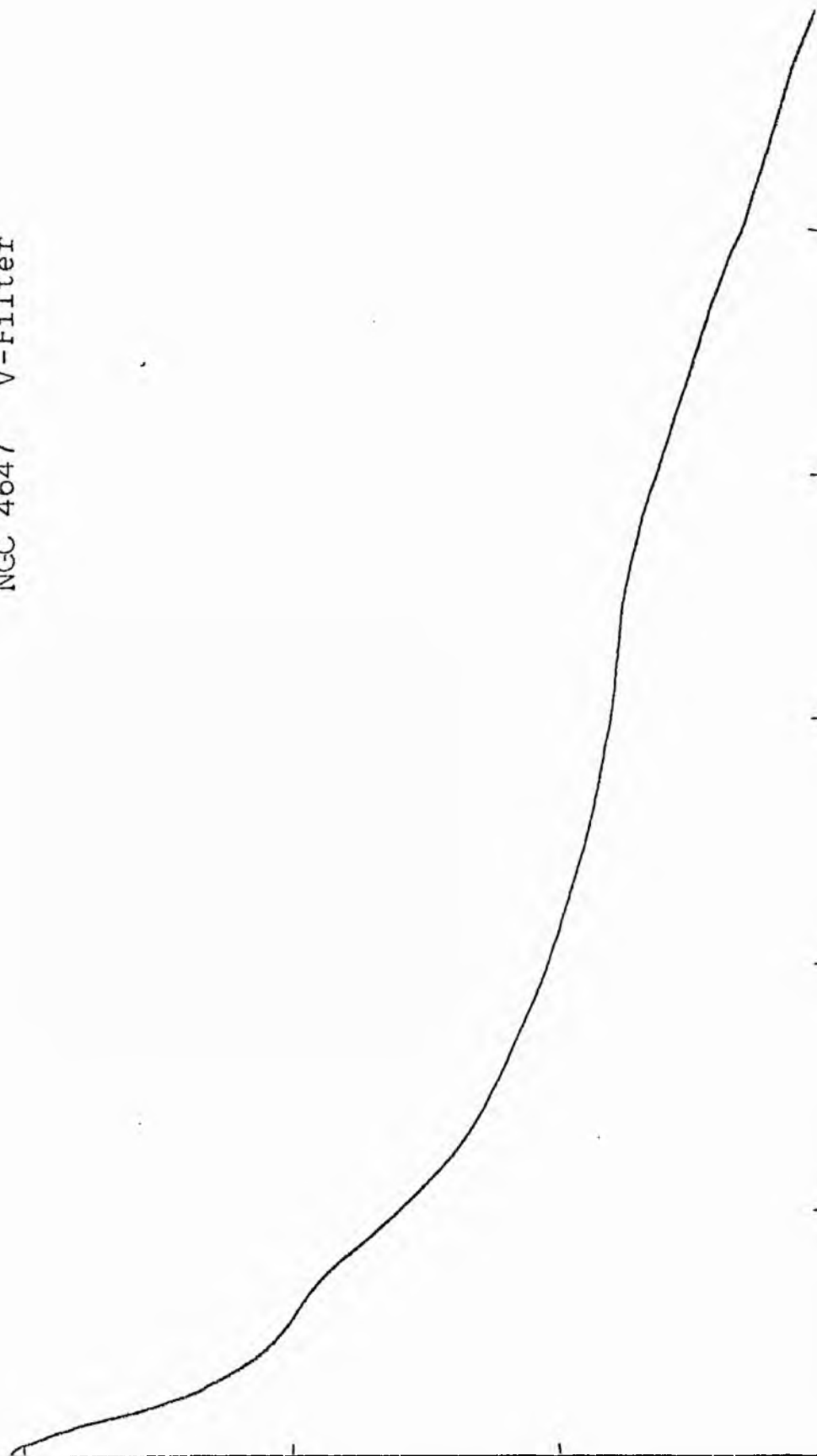


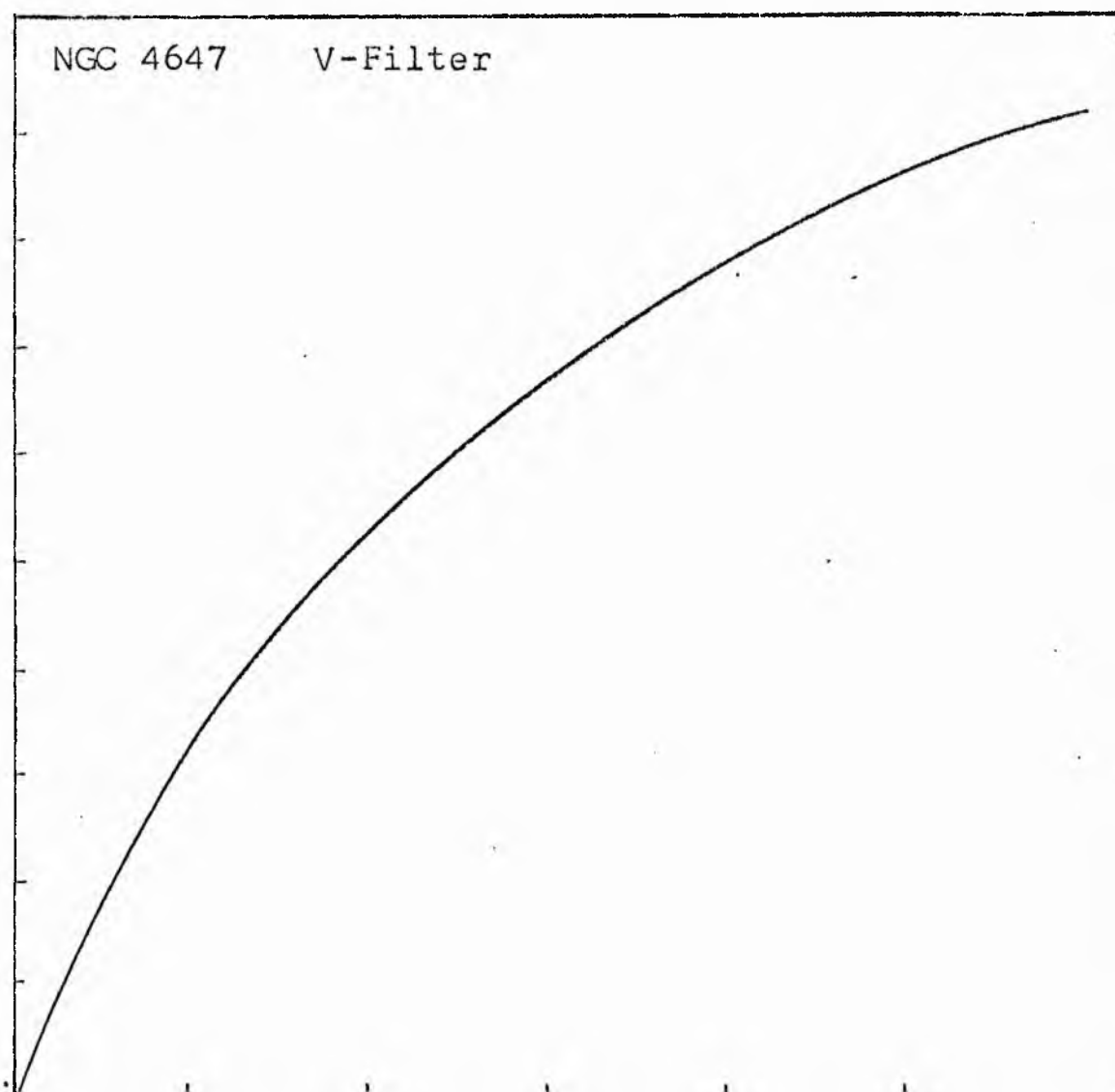
NGC 4647
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4647 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4647
V COLOUR

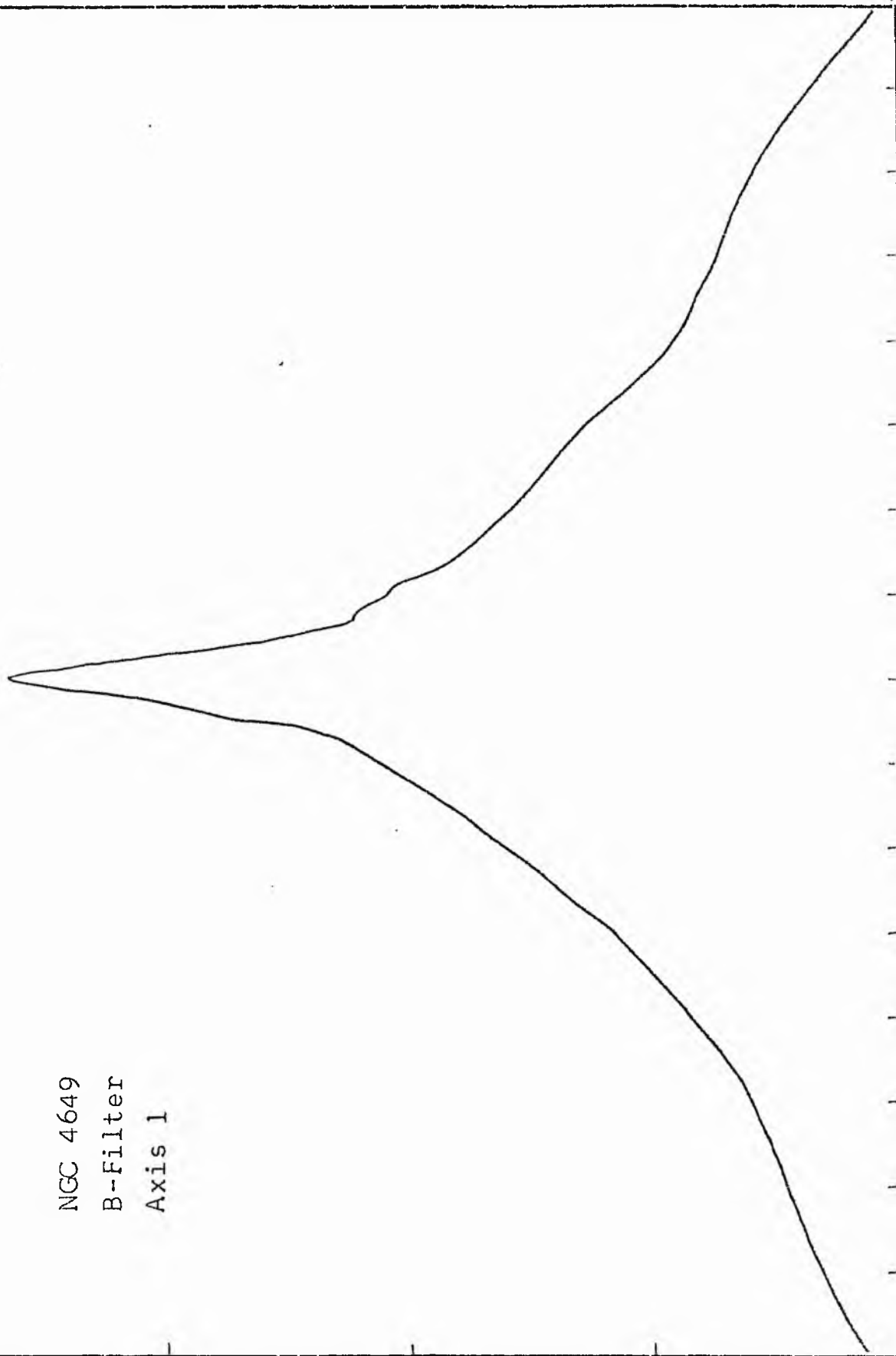
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(I)	ρ	LOG J	μ
1.06	11.482	10.741	0.0	0.0	18.55	199.2494	0.0	0.0	0.0	1.924	19.17
1.00	10.000	8.972	2.43	18.55	34.26	307.3621	199.25	0.01	0.03	1.864	19.32
0.90	7.943	7.126	4.10	52.81	22.62	161.1962	506.61	0.02	0.35	1.764	19.57
0.80	6.310	5.661	4.90	75.43	16.18	91.5858	667.81	0.03	0.05	1.664	19.82
0.70	5.012	4.496	5.40	91.61	124.82	561.2290	759.39	0.03	0.06	1.564	20.07
0.60	3.981	3.572	8.30	216.42	67.10	239.6746	1320.62	0.05	0.09	1.464	20.32
0.50	3.162	2.837	9.50	283.53	69.46	197.0657	1560.30	0.06	0.10	1.364	20.57
0.40	2.512	2.254	10.60	352.99	190.26	428.7712	1757.36	0.07	0.12	1.264	20.82
0.30	1.995	1.790	13.15	543.25	382.92	685.4504	2186.13	0.09	0.14	1.164	21.07
0.20	1.585	1.422	17.17	926.17	682.69	970.7268	2871.58	0.12	0.19	1.064	21.32
0.10	1.259	1.129	22.63	1608.86	1167.91	1319.1082	3842.31	0.16	0.25	0.964	21.57
0.00	1.000	0.897	29.73	2776.77	962.51	863.5330	5161.42	0.21	0.33	0.864	21.82
-0.10	0.794	0.713	34.50	3739.28	1014.61	723.0518	6024.95	0.25	0.38	0.764	22.07
-0.20	0.631	0.566	38.90	4753.89	1122.66	635.5046	6748.00	0.28	0.43	0.664	22.32
-0.30	0.501	0.450	43.25	5876.54	1513.27	680.4363	7383.50	0.31	0.48	0.564	22.57
-0.40	0.398	0.357	48.50	7389.81	2217.46	792.0059	8063.94	0.34	0.53	0.464	22.82
-0.50	0.316	0.284	55.30	9607.27	1815.84	515.1694	8855.94	0.37	0.61	0.364	23.07
-0.60	0.251	0.225	60.30	11423.11	4503.00	1014.7856	9371.11	0.39	0.66	0.264	23.32
-0.70	0.200	0.179	71.20	15926.11	5560.16	995.3130	10385.89	0.43	0.78	0.164	23.57
-0.80	0.158	0.142	82.70	21486.26	7225.87	1027.4553	11381.21	0.47	0.91	0.064	23.82
-0.90	0.126	0.113	95.60	28712.13	11544.98	1303.9656	12408.66	0.52	1.05	-0.036	24.07
-1.00	0.100	0.090	113.20	40257.12	15064.10	1351.5010	13712.62	0.57	1.25	-0.136	24.32
-1.10	0.079	0.071	132.70	55321.22	21525.09	1533.9763	15064.12	0.63	1.46	-0.236	24.57
-1.20	0.063	0.057	156.40	76846.31	19584.94	1108.6550	16598.10	0.69	1.72	-0.336	24.82
-1.30	0.050	0.045	175.20	96431.25	24625.87	1107.3030	17706.75	0.74	1.93	-0.436	25.07
-1.40	0.040	0.036	196.30	121057.12	26468.56	945.3787	18814.05	0.78	2.16	-0.536	25.32
-1.50	0.032	0.028	216.70	147525.69	20693.87	587.1077	19759.43	0.82	2.39	-0.636	25.57
-1.60	0.025	0.023	231.40	168219.56	24222.62	545.8806	20346.53	0.85	2.55	-0.736	25.82
-1.70	0.020	0.018	247.50	192442.19	30513.12	546.2151	20892.41	0.87	2.73	-0.836	26.07
-1.80	0.016	0.014	266.40	222955.31	14800.00	210.4454	21438.62	0.89	2.93	-0.936	26.32
-1.90	0.013	0.011	275.10	237755.31	43106.19	486.8752	21649.07	0.90	3.03	-1.036	26.57
-2.00	0.010		299.00	280861.50			22135.95	0.92	3.29	-1.136	26.82
-∞							24065.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4647

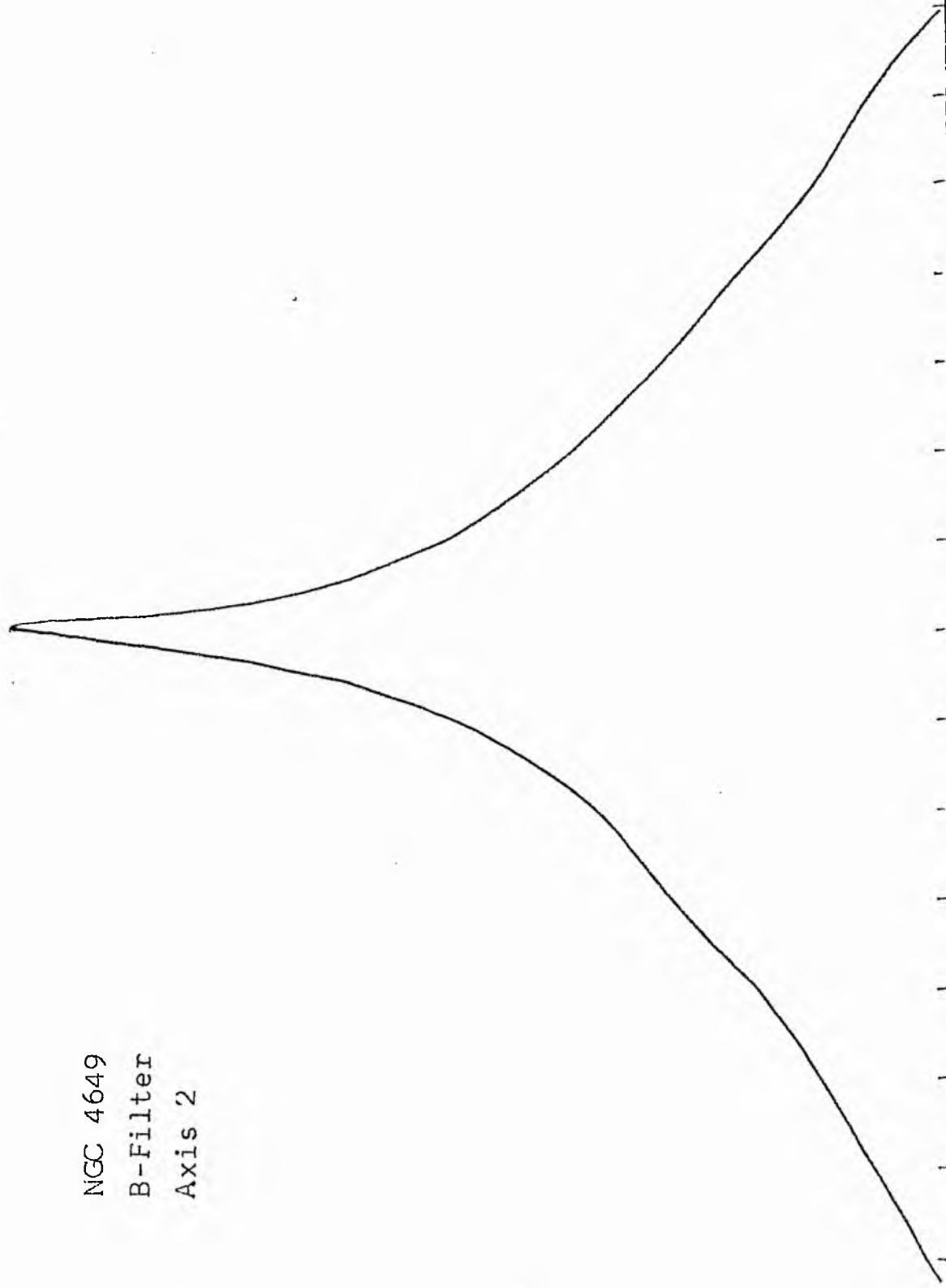
V-FILTER

Total luminosity	L_T	= 6.69
Total apparent magnitude	m_T	= 10.87
Apparent central surface brightness	μ_o	= 19.17
Major axis at threshold	$2a_m$	= 10.27
Minor axis at threshold	$2b_m$	= 9.37
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.33
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.73
Gradient of exponential component	$G(a)$	= -0.29
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.30
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.50
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.51
Axis ratio	b/a	= 0.98
Equivalent radius	r_1^*	= 0.57
Surface brightness	μ_1	= 22.07
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.38
Axis ratio	b/a	= 1.05
Equivalent radius	r_e^*	= 1.51
Surface brightness	μ_e	= 23.97
Mean surface brightness	μ_e'	= 13.76
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 3.38
Axis ratio	b/a	= 0.88
Equivalent radius	r_3^*	= 3.02
Surface brightness	μ_3	= 25.13
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{matrix} = 2.64 \\ = 2.00 \end{matrix}$

NGC 4649
B-Filter
Axis 1

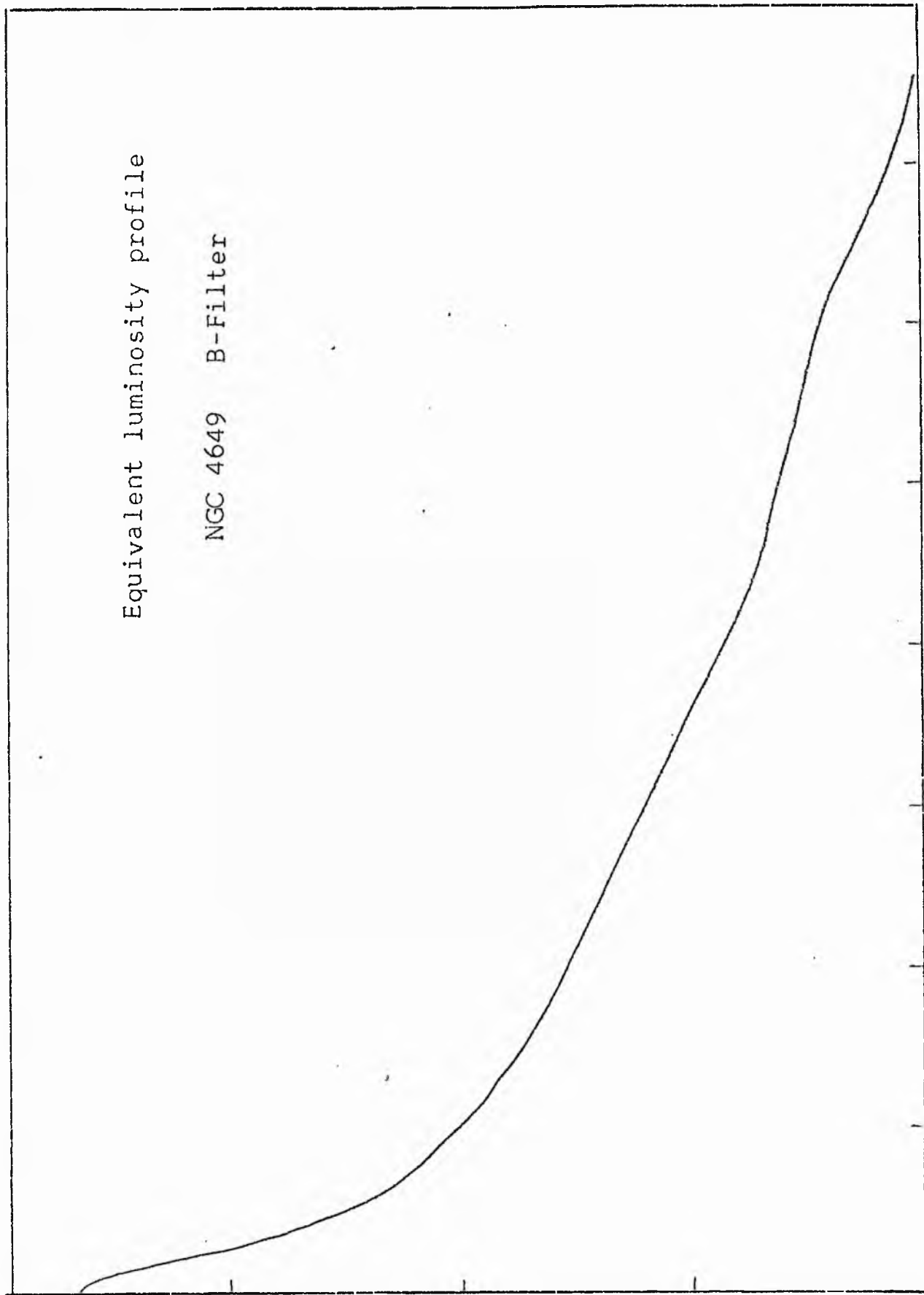


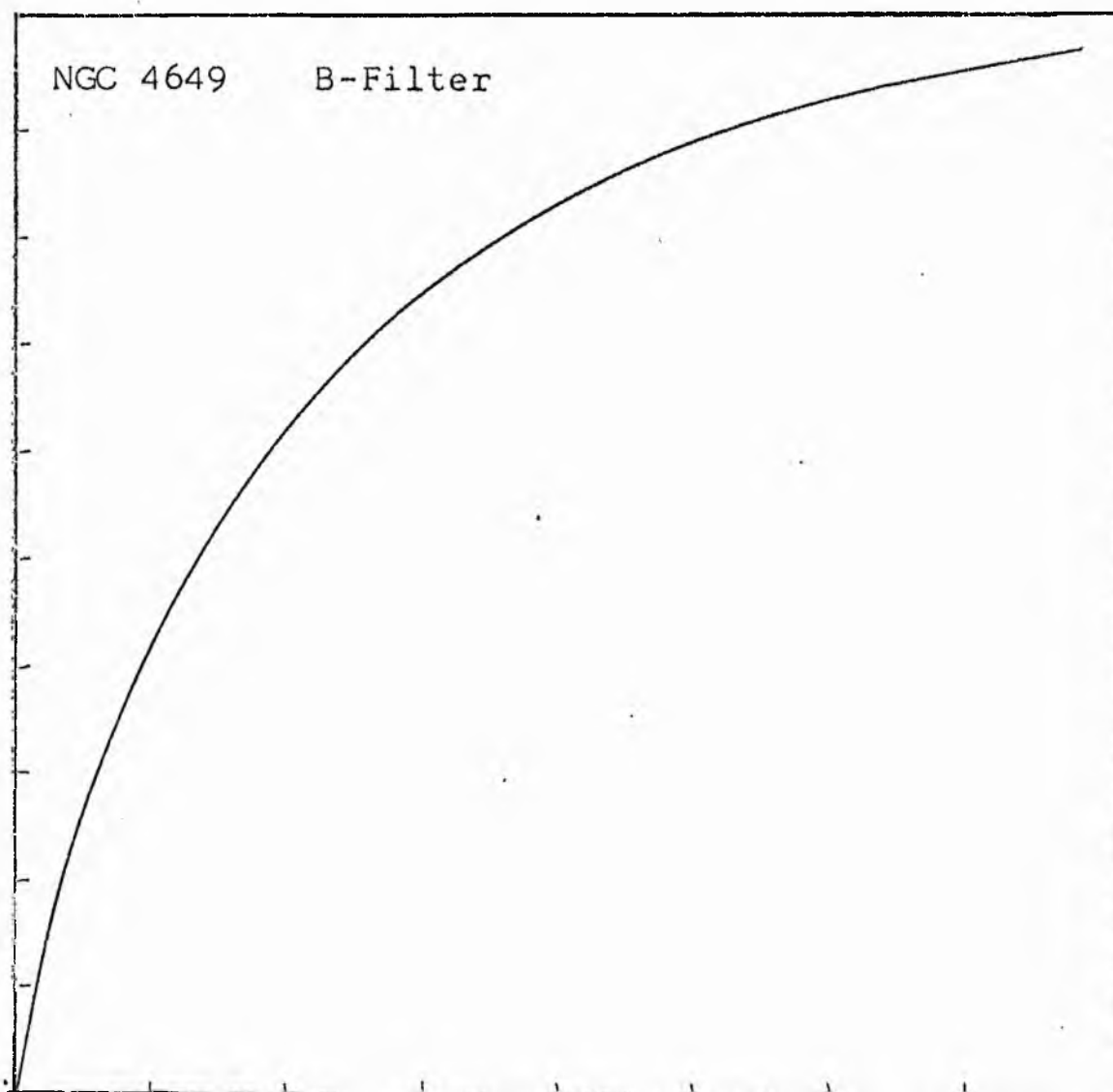
NGC 4649
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4649 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4649
B COLOUR

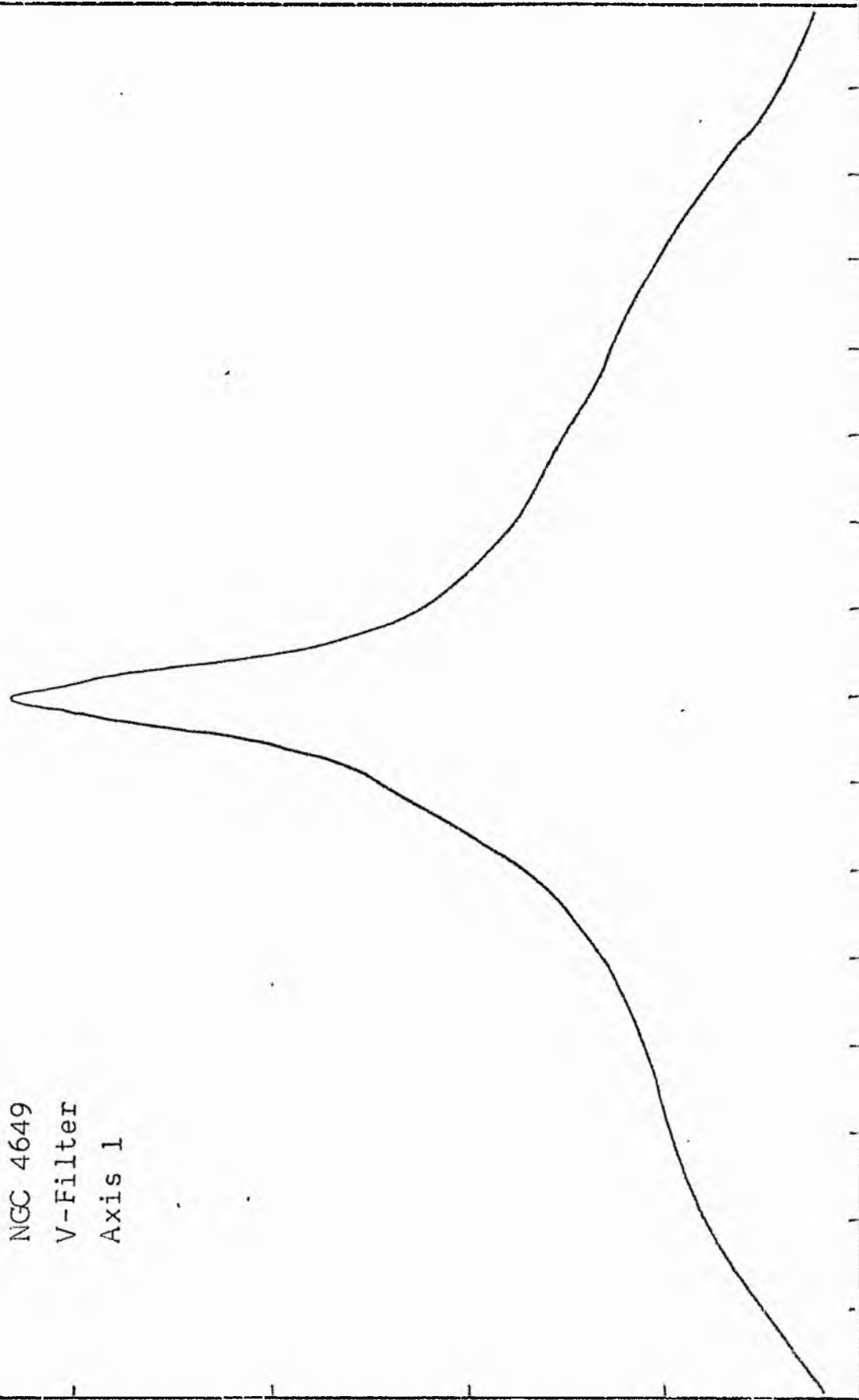
LOG I	I	T	R	AREA	Δ A	P	Σ P	K(R)	Q	LOG J	μ
1.66	45.709		0.0	0.0			0.0	0.0	0.0	1.829	17.82
1.60	39.811	42.760	3.36	35.47	35.47	1516.5123	1516.57	0.02	0.05	1.769	17.91
1.50	31.623	35.717	4.60	66.48	31.01	1107.5291	2624.10	0.04	0.07	1.669	18.23
1.40	25.119	28.371	6.80	145.27	78.79	2235.3621	4859.46	0.07	0.10	1.569	18.48
1.30	19.953	22.536	8.70	237.79	92.52	2084.9983	6944.46	0.10	0.13	1.469	18.73
1.20	15.849	17.901	10.30	333.29	95.50	1709.5947	8654.05	0.12	0.15	1.369	18.98
1.10	12.589	14.219	11.90	444.88	111.59	1586.6899	10240.74	0.15	0.17	1.269	19.23
1.00	10.000	11.295	13.50	572.56	127.67	1442.0247	11682.76	0.17	0.20	1.169	19.48
0.90	7.943	8.972	15.20	725.83	153.28	1315.1487	13057.91	0.19	0.22	1.069	19.73
0.80	6.310	7.126	17.20	929.41	203.57	1450.7517	14508.66	0.21	0.25	0.969	19.98
0.70	5.012	5.661	19.50	1144.59	265.16	1501.1150	16039.77	0.23	0.28	0.869	20.23
0.60	3.981	4.496	22.50	1590.43	355.84	1715.8767	17789.65	0.25	0.33	0.769	20.48
0.50	3.162	3.572	25.40	2026.83	436.40	1558.6804	19348.31	0.27	0.37	0.669	20.73
0.40	2.512	2.837	29.43	2814.25	787.42	2233.7774	21582.28	0.31	0.44	0.569	20.98
0.30	1.995	2.254	34.52	3743.62	925.36	2094.3702	23676.66	0.34	0.50	0.469	21.23
0.20	1.585	1.790	38.79	4775.91	1032.29	1847.8716	25524.53	0.36	0.57	0.369	21.48
0.10	1.259	1.422	44.97	6353.24	1577.33	2242.8110	27767.34	0.39	0.66	0.269	21.73
-0.00	1.000	1.129	51.86	8449.18	2095.94	2367.2747	30134.61	0.43	0.76	0.169	21.98
-0.10	0.794	0.897	60.74	11590.42	3141.24	2818.1953	32952.80	0.47	0.89	0.069	22.23
-0.20	0.631	0.713	71.93	16254.36	4663.93	3323.7024	36276.50	0.51	1.05	-0.031	22.48
-0.30	0.501	0.566	83.38	21841.05	5586.70	3162.4583	39438.96	0.56	1.22	-0.131	22.73
-0.40	0.398	0.450	95.65	28742.18	6901.12	3103.0574	42542.02	0.60	1.40	-0.231	22.98
-0.50	0.316	0.357	108.44	36942.71	8200.54	2928.9531	45470.97	0.64	1.58	-0.331	23.23
-0.60	0.251	0.284	122.71	4735.28	10362.57	2939.9333	48410.90	0.69	1.79	-0.431	23.48
-0.70	0.200	0.225	137.89	59733.14	12427.86	2800.6997	51211.60	0.73	2.01	-0.531	23.73
-0.80	0.158	0.179	147.14	68016.00	8282.86	1482.6902	52694.29	0.75	2.15	-0.631	23.98
-0.90	0.126	0.142	160.93	81362.37	13346.37	1897.7263	54592.01	0.77	2.35	-0.731	24.23
-1.00	0.100	0.113	172.79	94013.81	12651.44	1428.9280	56020.93	0.79	2.52	-0.831	24.48
-1.10	0.079	0.090	186.78	109600.00	15586.19	1378.3342	57419.27	0.81	2.73	-0.931	24.73
-1.20	0.063	0.071	202.82	129232.37	19632.37	1399.0854	58818.35	0.83	2.96	-1.031	24.98
-1.30	0.050	0.057	219.78	151749.12	22516.75	1274.6118	60092.96	0.85	3.21	-1.131	25.23
-1.40	0.040	0.045	253.58	207013.25	56264.12	2260.1172	62353.07	0.88	3.70	-1.231	25.48
-1.50	0.032	0.036	282.82	251266.56	49273.31	1759.8894	64112.96	0.91	4.13	-1.331	25.73
-1.60	0.025	0.028	309.10	300156.19	48865.62	1386.4790	65499.44	0.93	4.51	-1.431	25.98
-1.70	0.020	0.023	324.10	329995.06	29838.87	672.4456	66171.87	0.94	4.73	-1.531	26.23
-1.80	0.016	0.018	333.75	349939.06	19944.00	357.0156	66528.87	0.94	4.87	-1.631	26.48
-1.90	0.013	0.014	361.28	410050.31	60111.25	854.7351	67383.56	0.95	5.27	-1.731	26.73
-2.00	0.010	0.011	387.50	471729.81	61679.50	696.6541	68080.19	0.96	5.66	-1.831	26.98
-∞							70580.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4649

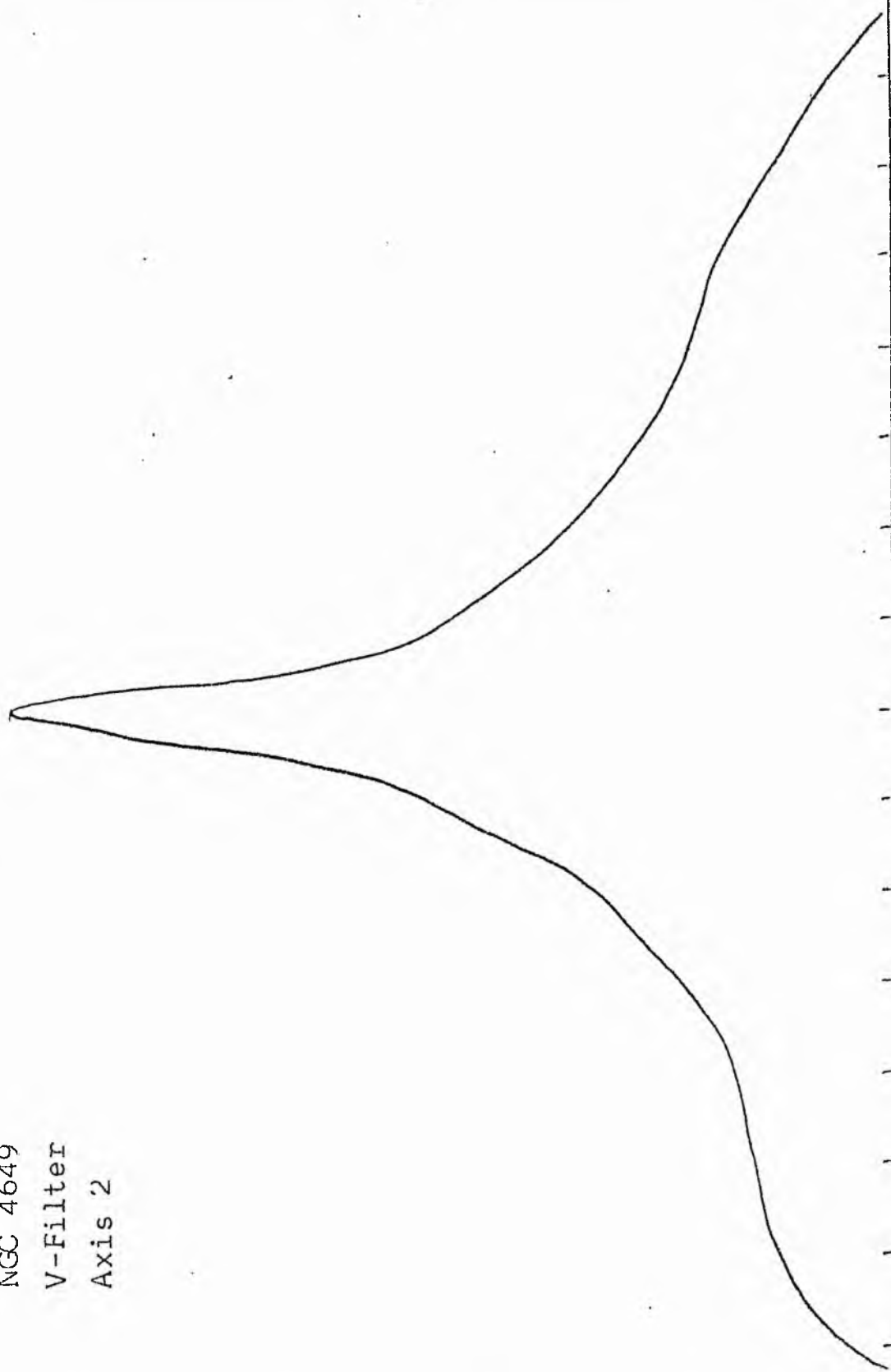
B-FILTER

Total luminosity	L_T	= 19.60
Total apparent magnitude	m_T	= 9.86
Apparent central surface brightness	μ_0	= 17.83
Major axis at threshold	$2a_m$	= 13.67
Minor axis at threshold	$2b_m$	= 12.10
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 8.17
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.30
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.35
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= +0.54
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.49
Axis ratio	b/a	= 0.90
Equivalent radius	r_1^*	= 0.37
Surface brightness	μ_1	= 20.48
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 1.30
Axis ratio	b/a	= 0.82
Equivalent radius	r_e^*	= 1.14
Surface brightness	μ_e	= 22.42
Mean surface brightness	μ_e'	= 12.14
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.61
Axis ratio	b/a	= 0.86
Equivalent radius	r_3^*	= 2.48
Surface brightness	μ_3	= 23.98
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.08 \\ = 2.17 \end{cases}$

NGC 4649
V-Filter
Axis 1

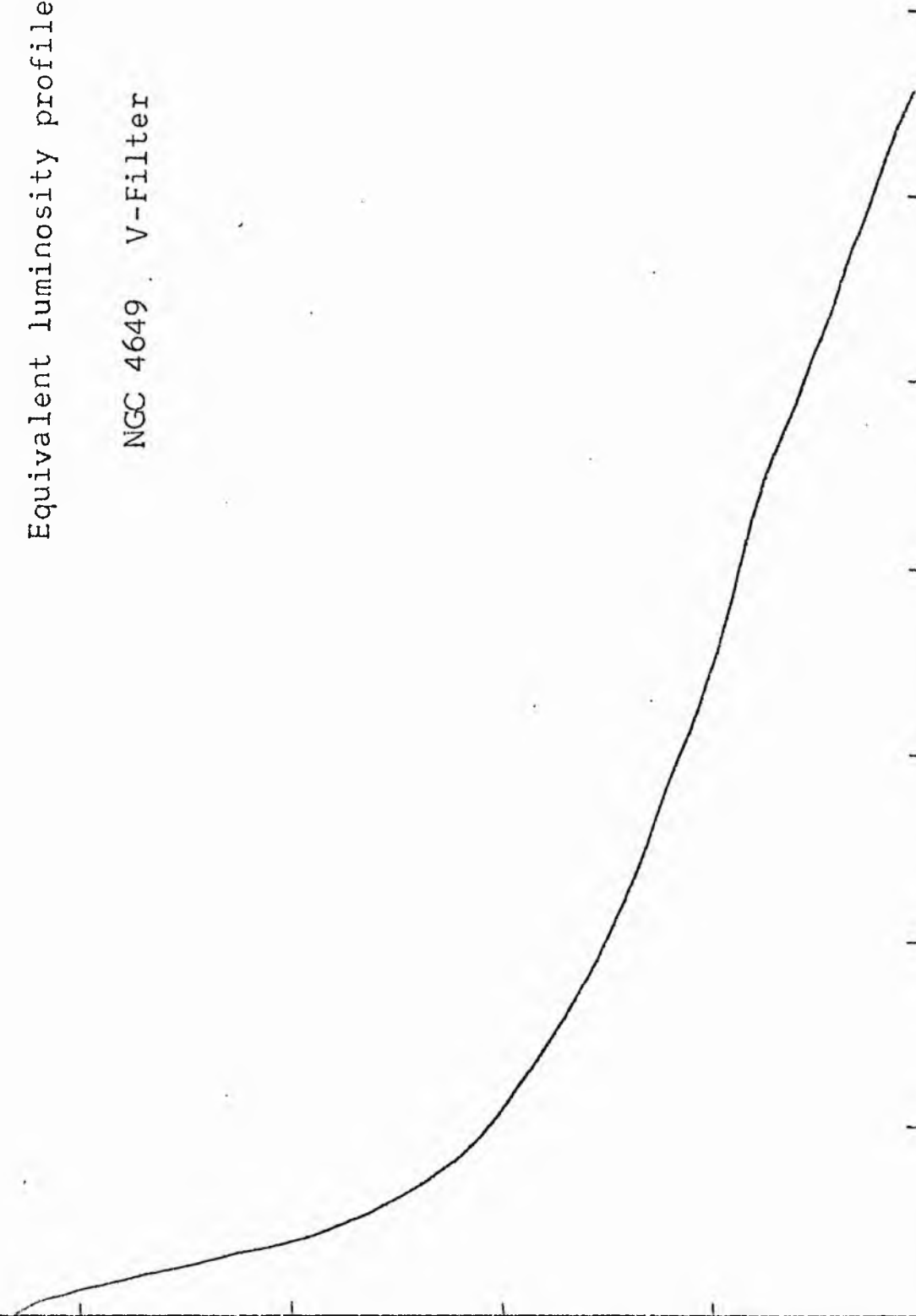


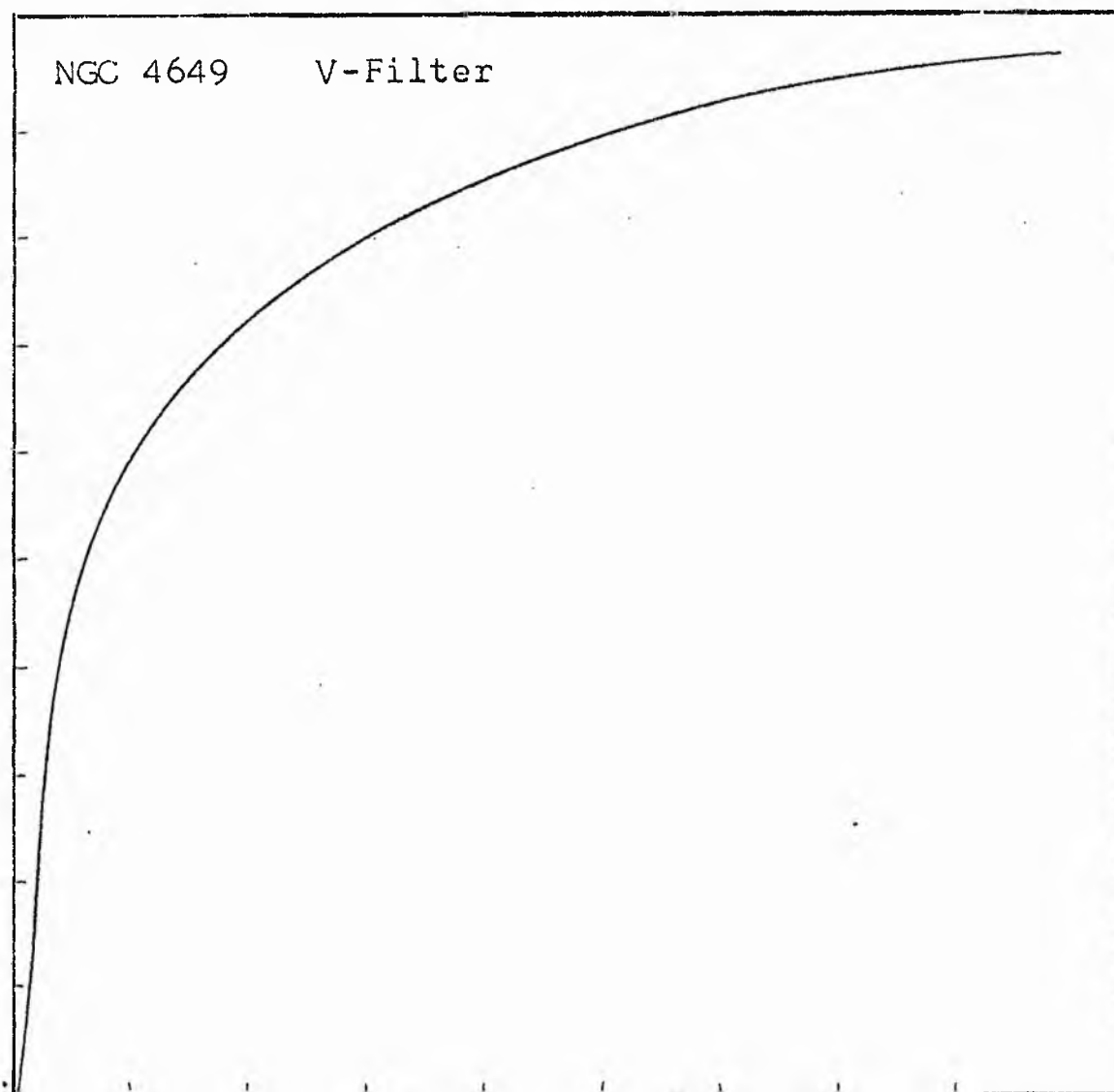
NGC 4649
V-Filter
Axis '2



Equivalent luminosity profile

NGC 4649 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4619
V COLOR

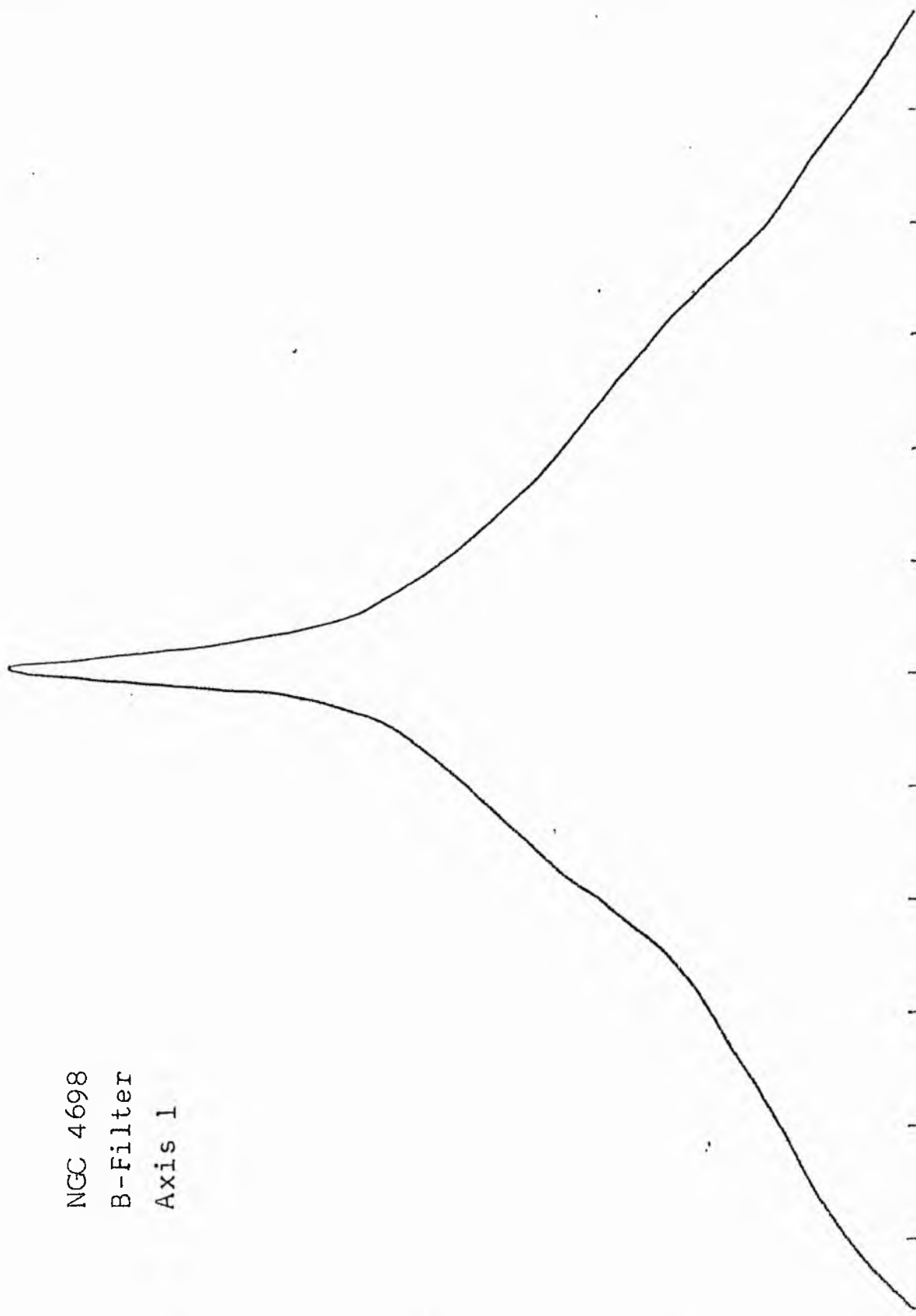
LOG I	I	\bar{I}	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
2.28	190.546		0.0	0.0			0.0	0.0	0.0	1.491	16.22
		174.517			8.04	1403.5513					
2.20	158.489	142.191	1.60	8.04	47.38	6736.3047	1403.55	0.01	0.05	1.411	16.42
2.10	125.092	112.946	4.20	55.42	89.85	10148.1445	8139.86	0.05	0.14	1.311	16.67
2.00	100.000	89.716	6.80	145.27	76.40	6854.6328	18288.00	0.12	0.23	1.211	16.92
1.90	79.433	71.264	8.40	221.67	124.69	8885.8984	25142.63	0.16	0.29	1.111	17.17
1.80	63.695	56.607	10.50	346.36	128.93	7298.3028	34028.53	0.22	0.36	1.011	17.42
1.70	50.118	44.964	12.30	475.29	137.82	6197.1992	41326.91	0.27	0.42	0.911	17.67
1.60	39.810	35.717	13.97	613.12	122.30	4368.1016	47524.11	0.31	0.48	0.811	17.92
1.50	31.623	28.371	15.30	735.42	226.70	6431.5430	51892.21	0.34	0.52	0.711	18.17
1.40	25.119	22.536	17.50	962.11	124.75	2811.3320	58323.76	0.38	0.60	0.611	18.42
1.30	19.952	17.901	18.60	1086.86	95.50	1709.5964	61135.09	0.39	0.63	0.511	18.67
1.20	15.849	14.219	19.40	1182.37	264.43	3759.9468	62844.68	0.41	0.66	0.411	18.92
1.10	12.589	11.295	21.46	1446.80	405.23	4576.8242	66604.62	0.43	0.73	0.311	19.17
1.00	10.000	8.972	24.28	1852.03	163.65	1468.1516	71181.44	0.46	0.83	0.211	19.42
0.90	7.943	7.126	25.33	2015.67	608.22	4334.3633	72649.56	0.47	0.86	0.111	19.67
0.80	6.310	5.661	28.90	2623.89	674.03	3815.4514	76983.87	0.50	0.99	0.011	19.92
0.70	5.012	4.496	32.40	3297.92	683.61	3073.7993	80799.31	0.52	1.11	-0.089	20.17
0.60	3.981	3.572	35.60	3981.53	507.30	1811.8987	83873.06	0.54	1.22	-0.189	20.42
0.50	3.162	2.837	37.80	4488.83	974.05	2763.4458	85684.94	0.55	1.29	-0.289	20.67
0.40	2.512	2.254	41.70	5462.88	1069.62	2410.4382	88448.37	0.57	1.42	-0.389	20.92
0.30	1.995	1.790	45.60	6532.50	1702.99	3048.4583	90858.81	0.59	1.56	-0.489	21.17
0.20	1.585	1.422	51.20	8235.49	2187.55	3110.4751	93907.25	0.61	1.75	-0.589	21.42
0.10	1.259	1.129	57.60	10423.05	3038.70	3432.0684	97017.69	0.63	1.97	-0.689	21.67
-0.00	1.000	0.897	65.46	13461.75	4017.02	3603.8887	100449.75	0.65	2.23	-0.789	21.92
-0.10	0.794	0.713	74.59	17478.77	5022.06	3578.8997	104053.62	0.67	2.55	-0.889	22.17
-0.20	0.631	0.566	84.63	22500.82	6301.50	3567.0671	107632.50	0.70	2.89	-0.989	22.42
-0.30	0.501	0.450	95.75	28802.32	7807.29	3510.4924	111199.56	0.72	3.27	-1.089	22.67
-0.40	0.398	0.357	107.95	36609.61	9637.67	3442.2336	114710.00	0.74	3.69	-1.189	22.92
-0.50	0.316	0.284	121.33	46247.28	11501.26	3262.9756	118152.19	0.76	4.14	-1.289	23.17
-0.60	0.251	0.225	135.58	57748.54	18538.71	4177.8008	121415.12	0.78	4.63	-1.389	23.42
-0.70	0.200	0.179	155.83	76287.25	25296.81	4528.2891	125592.87	0.81	5.32	-1.489	23.67
-0.80	0.158	0.142	179.82	101584.06	22214.19	3158.6309	130121.12	0.84	6.14	-1.589	23.92
-0.90	0.126	0.113	198.51	123798.25	29251.69	3303.8445	133279.75	0.86	6.78	-1.689	24.17
-1.00	0.100	0.090	220.72	153049.94	46137.19	4139.2383	136583.56	0.88	7.53	-1.789	24.42
-1.10	0.079	0.071	251.80	199187.12	35140.94	2504.2825	140722.75	0.91	8.60	-1.889	24.67
-1.20	0.063	0.057	273.11	234328.06	32034.25	1813.3638	143227.00	0.92	9.32	-1.989	24.92
-1.30	0.050	0.045	291.18	266362.31	32145.75	1445.4211	145040.31	0.94	9.94	-2.089	25.17
-1.40	0.040	0.036	308.25	298508.06	32200.31	1150.0891	146485.69	0.95	10.52	-2.189	25.42
-1.50	0.032	0.028	324.45	330718.37	39456.87	1119.4246	147635.75	0.95	11.08	-2.289	25.67
-1.60	0.025	0.023	343.26	370165.25	35809.62	816.9983	148755.12	0.96	11.72	-2.389	25.92
-1.70	0.020	0.018	359.48	405974.87	54957.87	983.7422	149562.06	0.97	12.27	-2.489	26.17
-1.80	0.016	0.014	383.04	460932.75	32714.75	465.1763	150545.81	0.97	13.08	-2.589	26.42
-1.90	0.013	0.011	396.40	493647.50	75673.31	854.7068	151010.94	0.98	13.53	-2.689	26.67
-2.00	0.010		425.70	569320.81			151865.62	0.98	14.53	-2.789	26.92
-∞							154865.00	{1}			∞

PHOTOMETRIC PARAMETERS OF NGC 4649

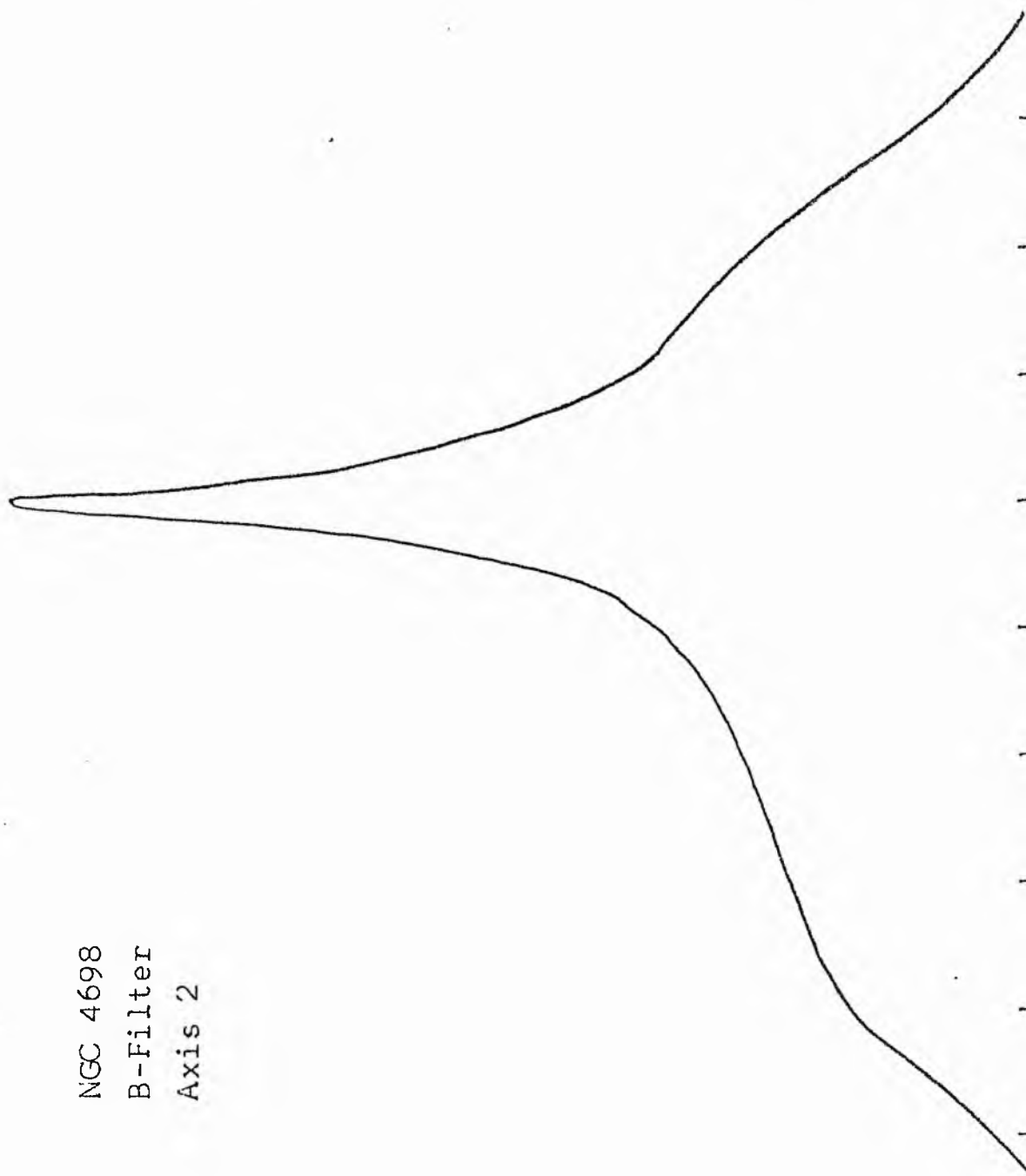
V-FILTER

Total luminosity	L_T	=	43.01
Total apparent magnitude	m_T	=	8.94
Apparent central surface brightness	μ_o	=	16.22
Major axis at threshold	$2a_m$	=	15.37
Minor axis at threshold	$2b_m$	=	12.56
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	=	9.65
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	=	0.93
Gradient of exponential component	$G(a)$	=	-0.31
Equivalent gradient of exponential comp....	$G(r^*)$	=	-0.32
Equivalent gradient of reduced exp. comp....	$G(\rho)$	=	-----
Parameters at $k = \frac{1}{4}$:			
Semi-major axis	a_1	=	0.20
Axis ratio	b/a	=	0.83
Equivalent radius	r_1^*	=	0.19
Surface brightness	μ_1	=	17.57
Parameters at $k = \frac{1}{2}$ (effective) :			
Semi-major axis	a_e	=	0.35
Axis ratio	b/a	=	0.91
Equivalent radius	r_e^*	=	0.49
Surface brightness	μ_e	=	19.92
Mean surface brightness	μ_e'	=	9.39
Parameters at $k = \frac{3}{4}$:			
Semi-major axis	a_3	=	2.43
Axis ratio	b/a	=	0.57
Equivalent radius	r_3^*	=	1.89
Surface brightness	μ_3	=	23.04
Concentration indices	$\left\{ \begin{matrix} C_{21} \\ C_{32} \end{matrix} \right.$	=	2.52
		=	3.87

NGC 4698
B-Filter
Axis 1

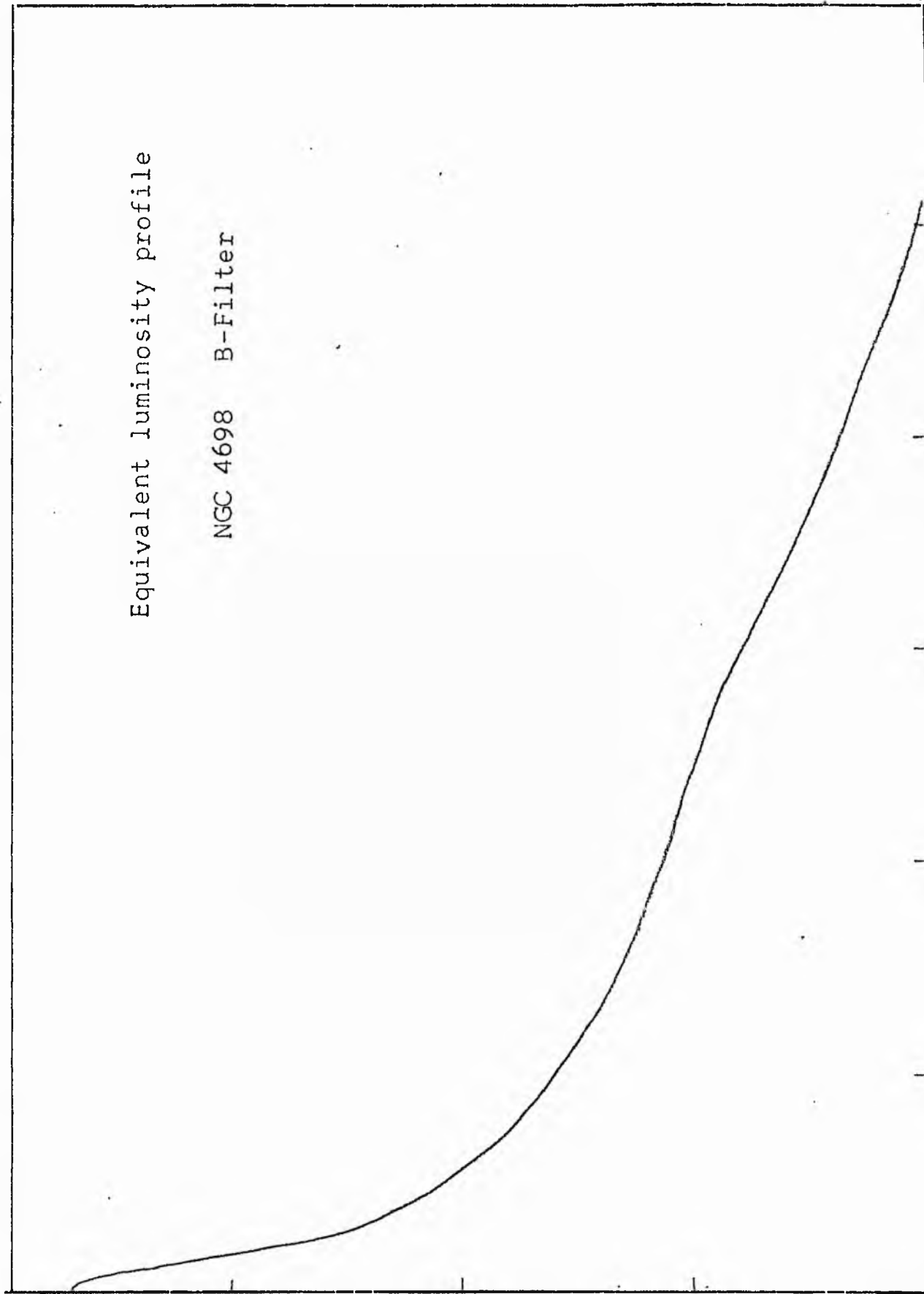


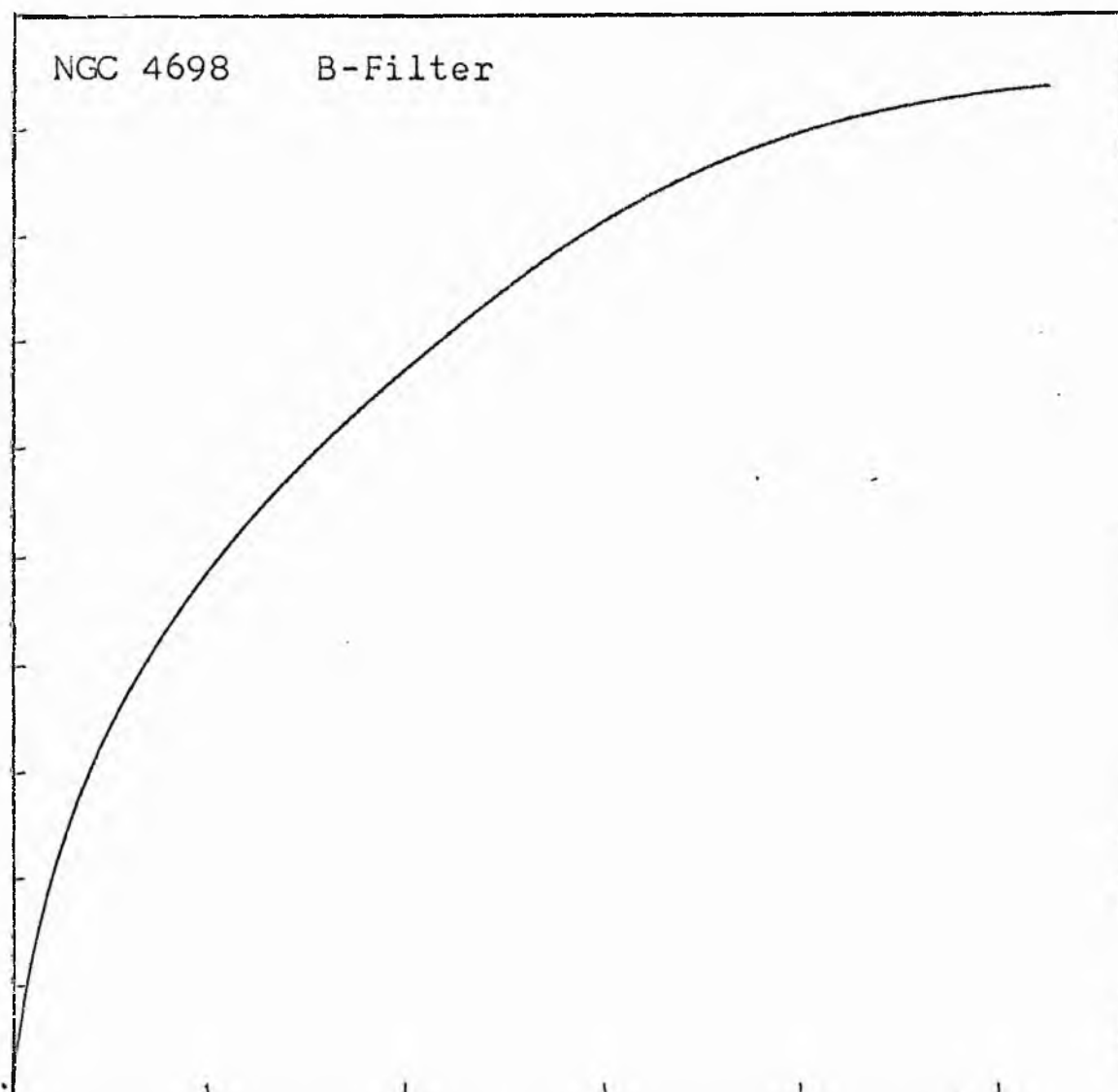
NGC 4698
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4698 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4698
B COLOUR

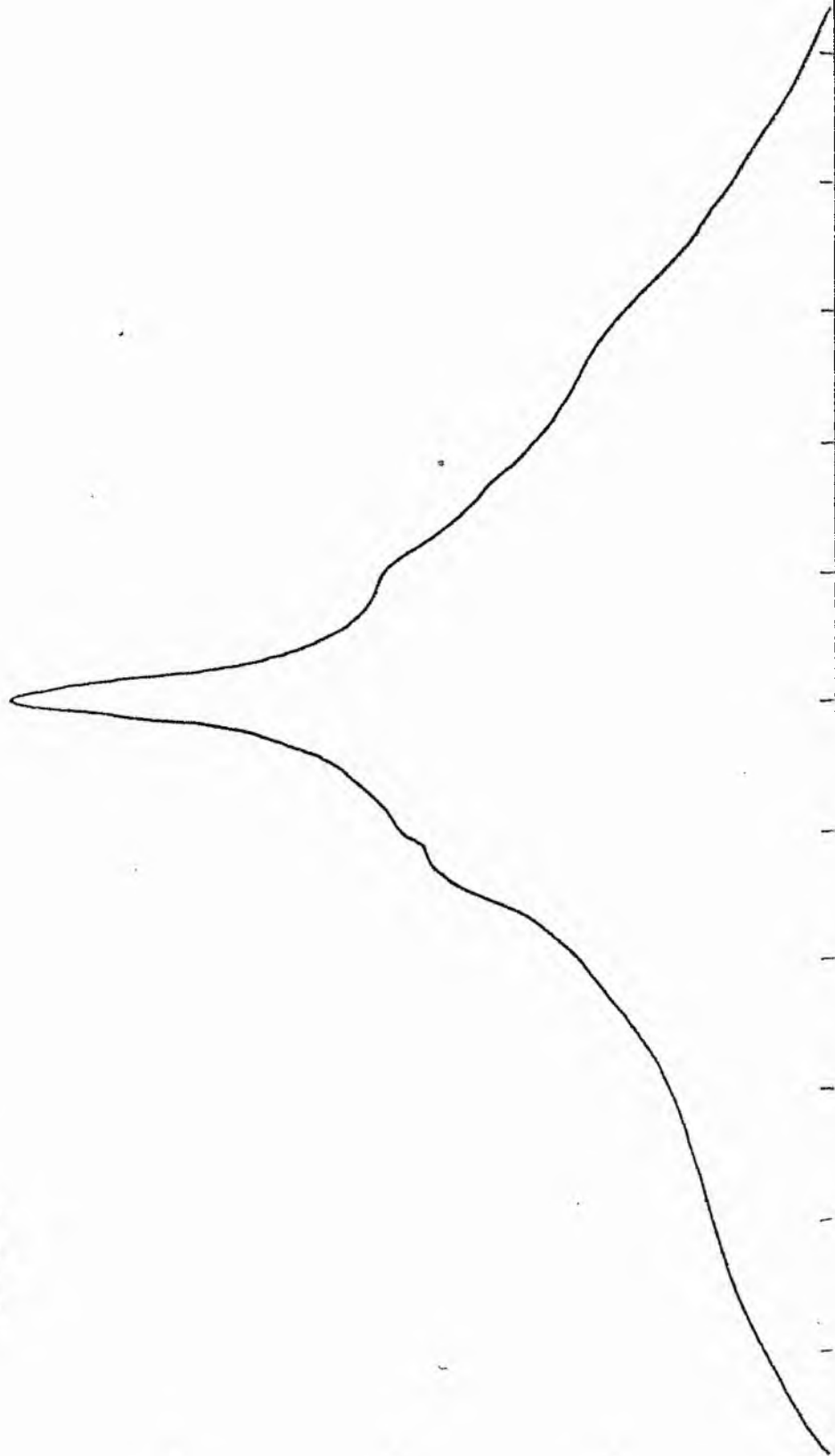
LOG I	I	I	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.69	48.978		0.0	0.0			0.0	0.2	0.0	2.121	18.22
1.60	39.811	44.394	3.13	30.78	30.78	1366.3591	1366.36	0.05	0.06	2.031	18.45
1.50	31.623	35.717	4.04	51.28	20.50	732.1187	2098.48	0.08	0.08	1.931	18.70
1.40	25.119	28.371	4.40	60.02	9.55	270.8093	2369.29	0.09	0.09	1.831	18.95
1.30	19.953	22.536	5.60	98.52	37.70	849.5745	3218.86	0.12	0.11	1.731	19.20
1.20	15.849	17.901	6.30	124.69	26.17	468.4514	3687.31	0.13	0.12	1.631	19.45
1.10	12.589	14.219	7.10	158.37	33.68	478.8660	4166.18	0.15	0.14	1.531	19.70
1.00	10.000	11.295	8.10	206.12	47.75	539.3398	4705.52	0.17	0.16	1.431	19.95
0.90	7.943	8.972	9.50	283.53	77.41	694.4783	5399.99	0.19	0.18	1.331	20.20
0.80	6.310	7.126	10.50	346.36	62.83	447.7651	5847.75	0.21	0.20	1.231	20.45
0.70	5.012	5.661	11.30	401.15	54.79	310.1450	6157.90	0.22	0.22	1.131	20.70
0.60	3.981	4.496	12.50	490.87	89.72	433.4385	6561.34	0.24	0.24	1.031	20.95
0.50	3.162	3.572	14.60	669.66	178.79	638.5688	7199.90	0.26	0.28	0.931	21.20
0.40	2.512	2.837	16.80	886.68	217.02	615.6987	7815.60	0.28	0.33	0.831	21.45
0.30	1.995	2.254	18.80	1110.36	223.68	504.0801	8319.68	0.30	0.36	0.731	21.70
0.20	1.585	1.790	21.80	1493.01	382.65	684.9617	9004.64	0.32	0.42	0.631	21.95
0.10	1.259	1.422	26.00	2123.72	630.71	896.8054	9901.44	0.36	0.50	0.531	22.20
-0.00	1.000	1.129	31.10	3038.58	914.86	1033.2954	10934.73	0.39	0.60	0.431	22.45
-0.10	0.794	0.897	35.50	3959.19	920.61	825.9382	11760.67	0.42	0.69	0.331	22.70
-0.20	0.631	0.713	36.10	4094.15	134.96	96.1781	11856.85	0.43	0.70	0.231	22.95
-0.30	0.501	0.566	40.60	5178.47	1084.32	613.8003	12470.64	0.45	0.79	0.131	23.20
-0.40	0.398	0.450	48.80	7481.51	2303.04	1035.5486	13506.19	0.49	0.94	0.031	23.45
-0.50	0.316	0.357	59.60	11159.43	3677.93	1313.6296	14819.82	0.53	1.15	-0.069	23.70
-0.60	0.251	0.284	68.40	14698.12	3538.69	1003.9517	15823.77	0.57	1.32	-0.169	23.95
-0.70	0.200	0.225	79.50	19855.65	5157.53	1162.2827	16986.05	0.61	1.54	-0.269	24.20
-0.80	0.158	0.179	91.20	26130.00	6274.35	1123.1523	18109.20	0.65	1.77	-0.369	24.45
-0.90	0.126	0.142	112.40	39690.12	13560.12	1928.1189	20037.32	0.72	2.18	-0.469	24.70
-1.00	0.100	0.113	129.70	52848.14	13158.02	1486.1448	21523.46	0.78	2.51	-0.569	24.95
-1.10	0.079	0.090	138.50	60262.82	7414.68	665.2168	22188.68	0.80	2.68	-0.669	25.20
-1.20	0.063	0.071	149.60	70309.31	10346.49	715.9551	22904.63	0.83	2.90	-0.769	25.45
-1.30	0.050	0.057	160.70	81130.00	10820.69	612.5295	23517.16	0.85	3.11	-0.869	25.70
-1.40	0.040	0.045	169.80	90578.50	9440.50	424.8501	23942.01	0.86	3.29	-0.969	25.95
-1.50	0.032	0.036	176.20	97535.25	6956.75	248.4735	24190.48	0.87	3.41	-1.069	26.20
-1.60	0.025	0.028	194.30	118602.94	21067.69	597.7109	24788.19	0.89	3.76	-1.169	26.45
-1.70	0.020	0.023	212.40	141729.00	23126.06	521.1663	25309.36	0.91	4.11	-1.269	26.70
-1.80	0.016	0.018	225.00	159043.12	17314.12	309.9385	25619.29	0.92	4.35	-1.369	26.95
-1.90	0.013	0.014	247.80	192908.94	33865.81	481.5454	26100.84	0.94	4.80	-1.469	27.20
-2.00	0.010	0.011	258.40	209765.69	16856.75	190.3927	26291.23	0.95	5.30	-1.569	27.45
-∞							27721.00	(1)			m

PHOTOMETRIC PARAMETERS OF NGC 4698

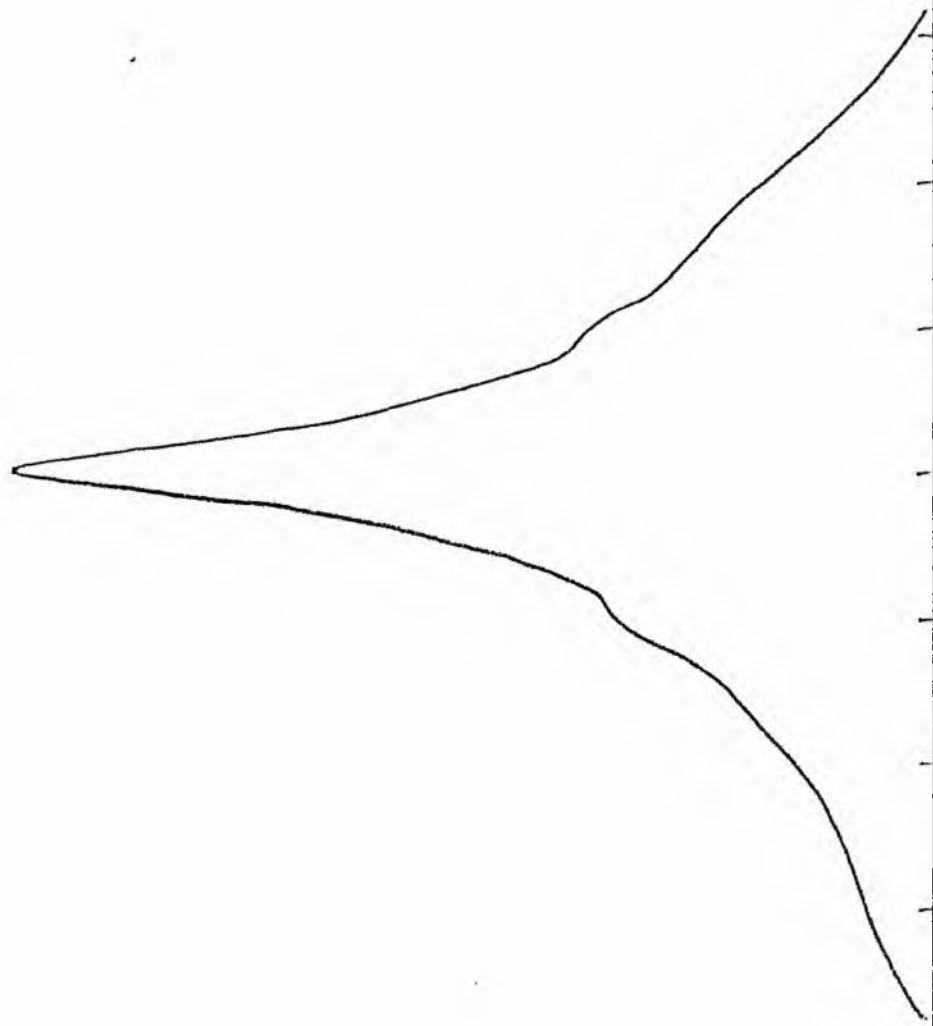
B-FILTER

Total luminosity	L_T	= 7.70
Total apparent magnitude	m_T	= 11.34
Apparent central surface brightness	μ_0	= 18.22
Major axis at threshold	$2a_m$	= 9.32
Minor axis at threshold	$2b_m$	= 7.40
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.30
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.78
Gradient of exponential component	$G(a)$	= -0.50
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.52
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.48
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.22
Axis ratio	b/a	= 0.92
Equivalent radius	r_1^*	= 0.23
Surface brightness	μ_1	= 21.09
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.28
Axis ratio	b/a	= 0.48
Equivalent radius	r_e^*	= 0.85
Surface brightness	μ_e	= 23.51
Mean surface brightness	μ_e'	= 12.98
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.16
Axis ratio	b/a	= 0.65
Equivalent radius	r_3^*	= 1.99
Surface brightness	μ_3	= 24.82
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 3.77 \\ = 2.34 \end{cases}$

NGC 4698
V-Filter
Axis 1

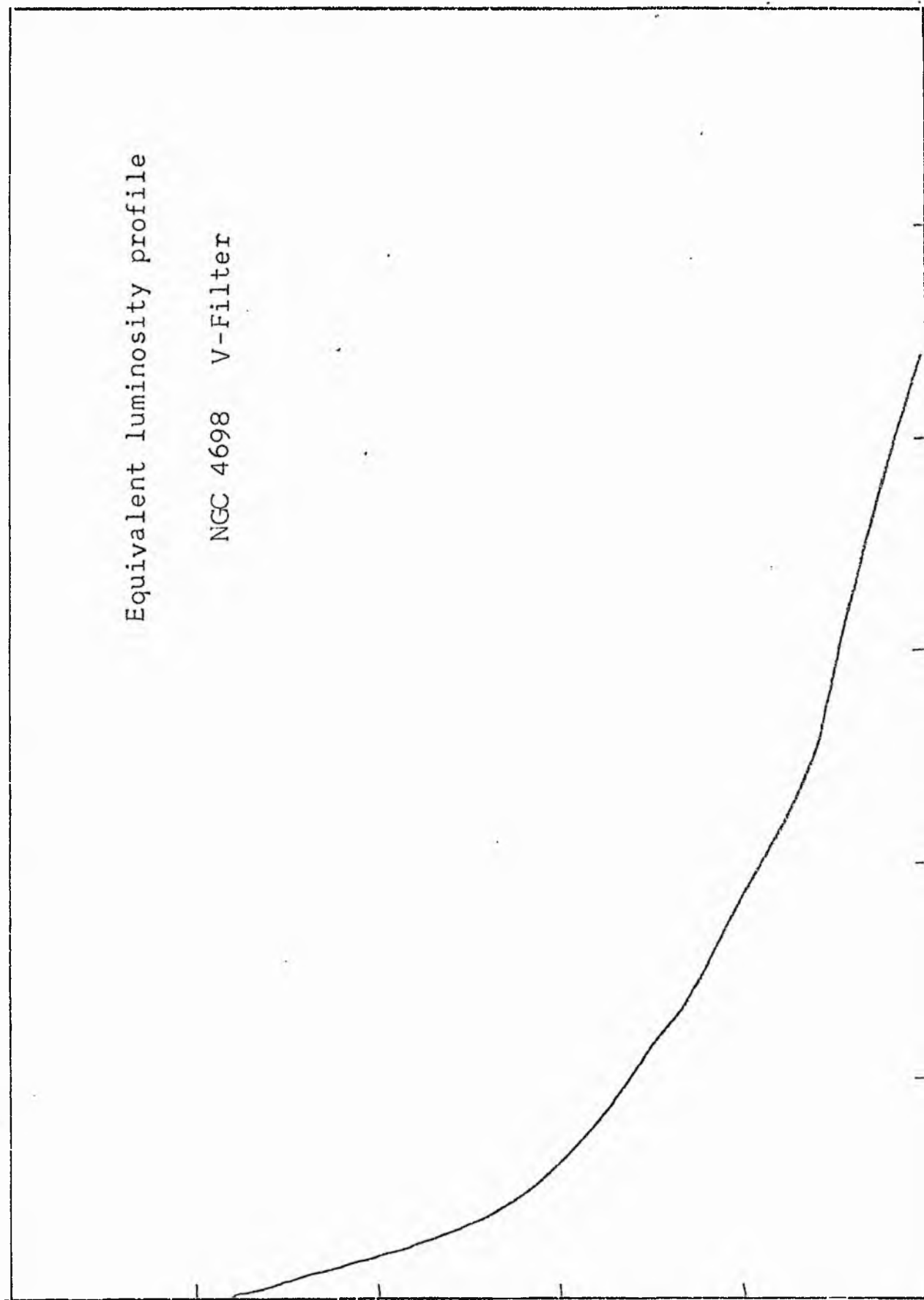


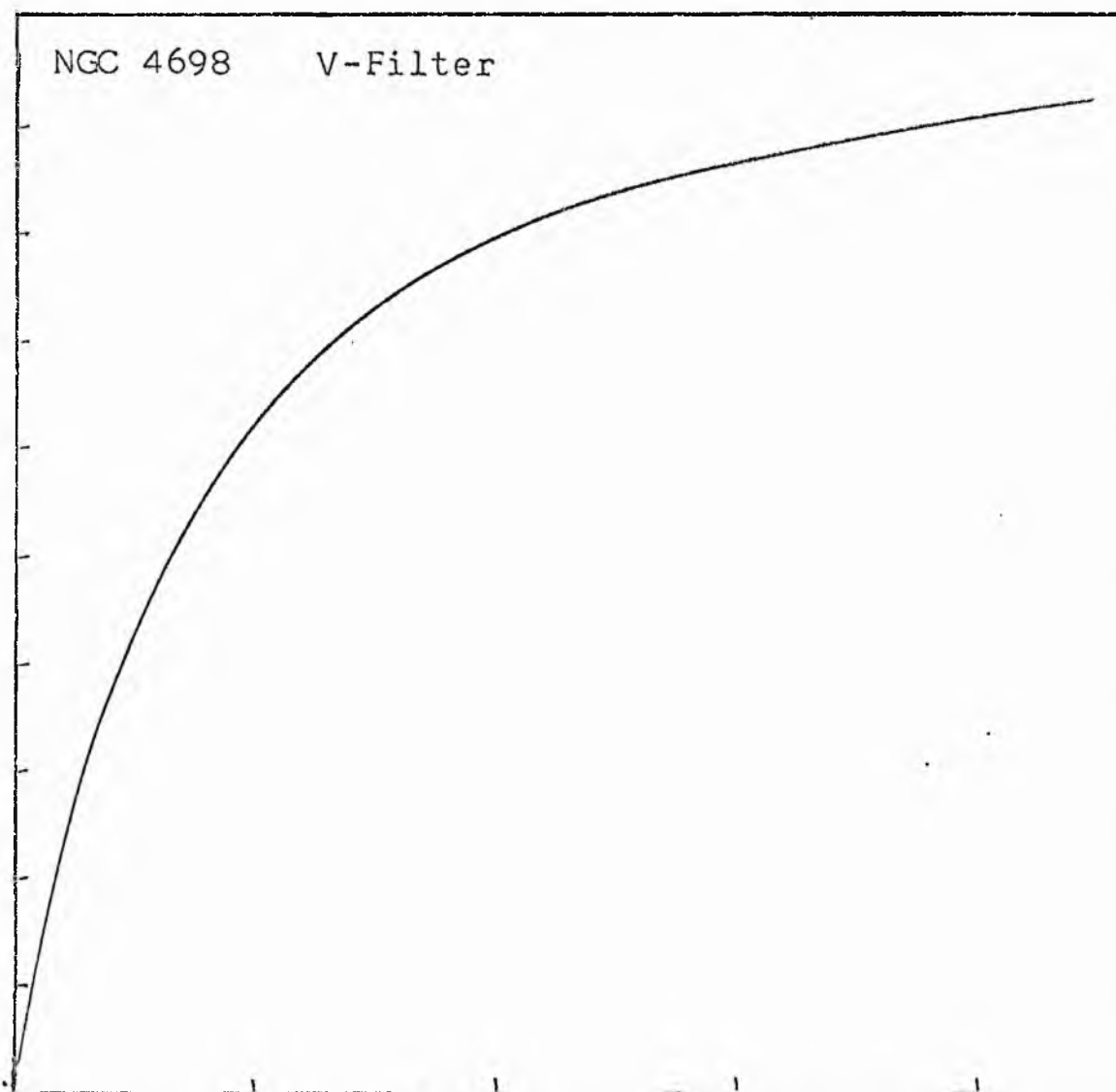
NGC 4698
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4698 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4698
V COLOUR

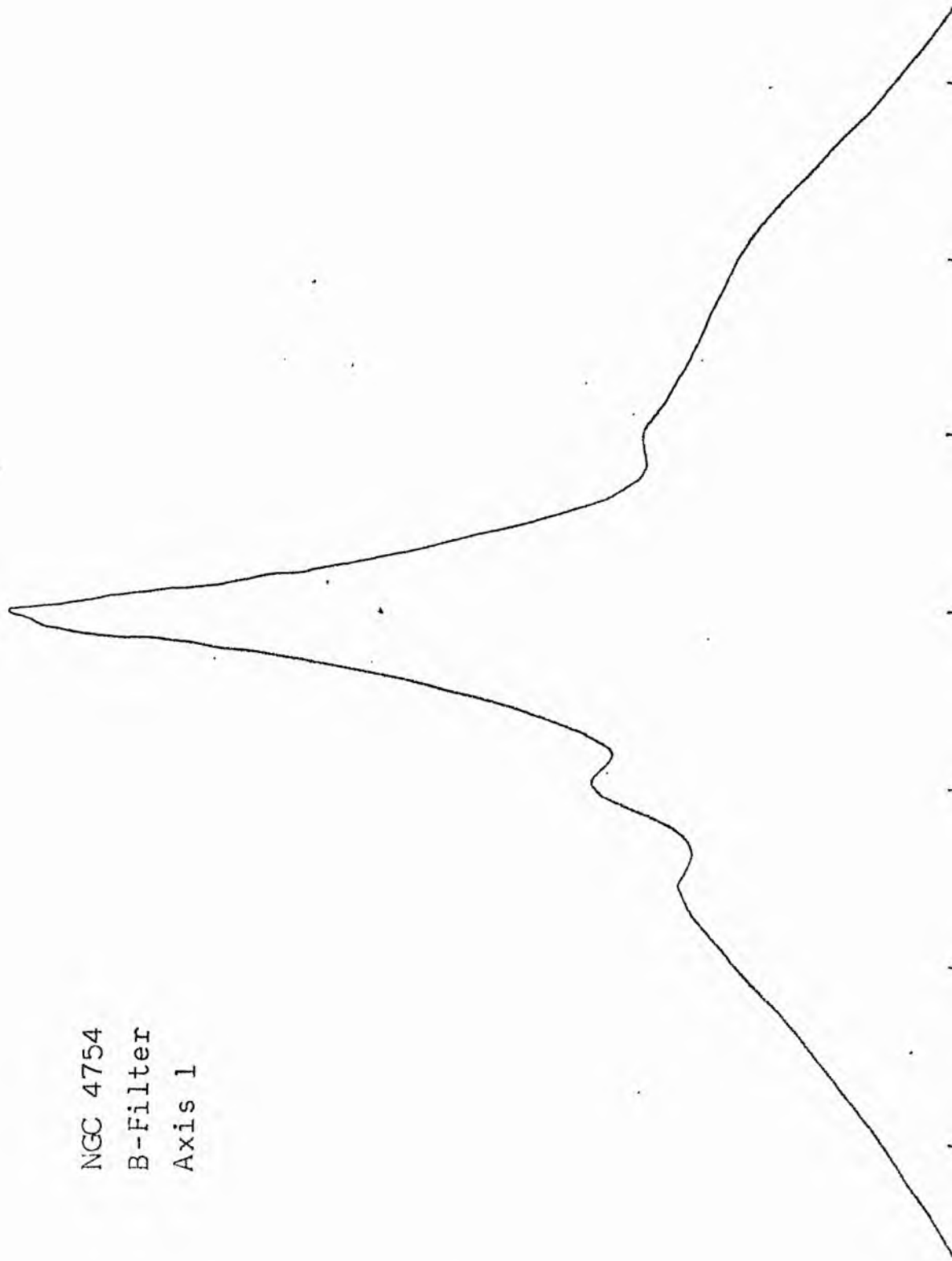
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.71	51.286		0.0	0.0			0.0	0.0	0.0	1.654	17.22
1.70	50.119	50.702	0.40	0.50	0.50	25.4858	25.49	0.00	0.01	1.644	17.25
1.60	39.811	44.965	1.50	7.07	6.57	295.2346	320.72	0.01	0.05	1.544	17.50
1.50	31.623	35.717	2.10	13.85	6.79	242.3673	563.09	0.02	0.07	1.444	17.75
1.40	25.119	28.371	3.50	38.48	24.63	698.7739	1261.86	0.05	0.11	1.344	18.00
1.30	19.953	22.536	5.20	84.95	46.46	1047.0996	2308.96	0.10	0.16	1.244	18.25
1.20	15.849	17.901	6.60	136.85	51.90	929.0298	3237.99	0.14	0.21	1.144	18.50
1.10	12.589	14.219	8.10	206.12	69.27	984.9810	4222.97	0.18	0.26	1.044	18.75
1.00	10.000	11.295	9.30	271.72	65.60	740.8799	4963.85	0.21	0.29	0.944	19.00
0.90	7.943	8.972	10.30	333.29	61.58	552.4290	5516.27	0.24	0.32	0.844	19.25
0.80	6.310	7.126	11.20	394.08	60.79	433.2114	5949.48	0.26	0.35	0.744	19.50
0.70	5.012	5.661	12.60	498.76	104.68	592.5481	6542.03	0.28	0.40	0.644	19.75
0.60	3.981	4.496	14.50	660.52	161.76	727.3481	7269.38	0.31	0.46	0.544	20.00
0.50	3.162	3.572	15.80	784.27	123.75	441.9824	7711.36	0.33	0.50	0.444	20.25
0.40	2.512	2.837	17.70	984.23	199.96	567.3052	8278.66	0.36	0.56	0.344	20.50
0.30	1.995	2.254	19.40	1182.37	198.14	446.5200	8725.18	0.38	0.61	0.244	20.75
0.20	1.585	1.790	23.30	1705.54	523.17	936.5078	9661.69	0.42	0.73	0.144	21.00
0.10	1.259	1.422	29.40	2715.47	1009.93	1436.0193	11097.71	0.48	0.93	0.044	21.25
-0.00	1.000	1.129	34.40	3717.63	1002.17	1131.9043	12229.61	0.53	1.08	-0.056	21.50
-0.10	0.794	0.897	38.70	4705.13	987.50	885.9409	13115.55	0.56	1.22	-0.156	21.75
-0.20	0.631	0.713	39.30	4852.15	147.02	104.7747	13220.32	0.57	1.24	-0.256	22.00
-0.30	0.501	0.566	43.60	5972.04	1119.89	633.9336	13854.25	0.60	1.37	-0.356	22.25
-0.40	0.398	0.450	48.90	7512.20	1540.16	692.5271	14546.78	0.63	1.54	-0.456	22.50
-0.50	0.316	0.357	55.00	9503.32	1991.11	711.1577	15257.93	0.66	1.73	-0.556	22.75
-0.60	0.251	0.284	61.80	11998.49	2495.17	707.8977	15965.83	0.69	1.95	-0.656	23.00
-0.70	0.200	0.225	66.70	13976.59	1978.10	445.7781	16411.61	0.71	2.10	-0.756	23.25
-0.80	0.158	0.179	76.80	18529.85	4553.26	815.0659	17226.67	0.74	2.42	-0.856	23.50
-0.90	0.126	0.142	82.30	21278.90	2749.05	390.8884	17617.55	0.76	2.59	-0.956	23.75
-1.00	0.100	0.113	90.40	25673.59	4394.68	496.3613	18113.91	0.78	2.85	-1.056	24.00
-1.10	0.079	0.090	98.70	30604.41	4930.83	442.3752	18556.29	0.80	3.11	-1.156	24.25
-1.20	0.063	0.071	104.50	34306.98	3702.57	263.8604	18820.15	0.81	3.29	-1.256	24.50
-1.30	0.050	0.057	113.50	40470.79	6163.80	348.9160	19169.06	0.82	3.58	-1.356	24.75
-1.40	0.040	0.045	123.70	48071.67	7600.89	341.7725	19510.83	0.84	3.90	-1.456	25.00
-1.50	0.032	0.036	142.70	63973.16	15901.49	567.9517	20078.78	0.86	4.50	-1.556	25.25
-1.60	0.025	0.028	161.20	81635.62	17662.46	501.1013	20579.88	0.88	5.08	-1.656	25.50
-1.70	0.020	0.023	176.30	97645.94	16010.31	360.8066	20940.68	0.90	5.56	-1.756	25.75
-1.80	0.016	0.018	190.30	113769.87	16123.94	288.6331	21229.32	0.91	6.00	-1.856	26.00
-1.90	0.013	0.014	208.40	136441.12	22671.25	322.3674	21551.68	0.93	6.57	-1.956	26.25
-2.00	0.010	0.011	220.90	153299.69	16858.56	190.4132	21742.09	0.93	6.96	-2.056	26.50
-∞							23262.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4698

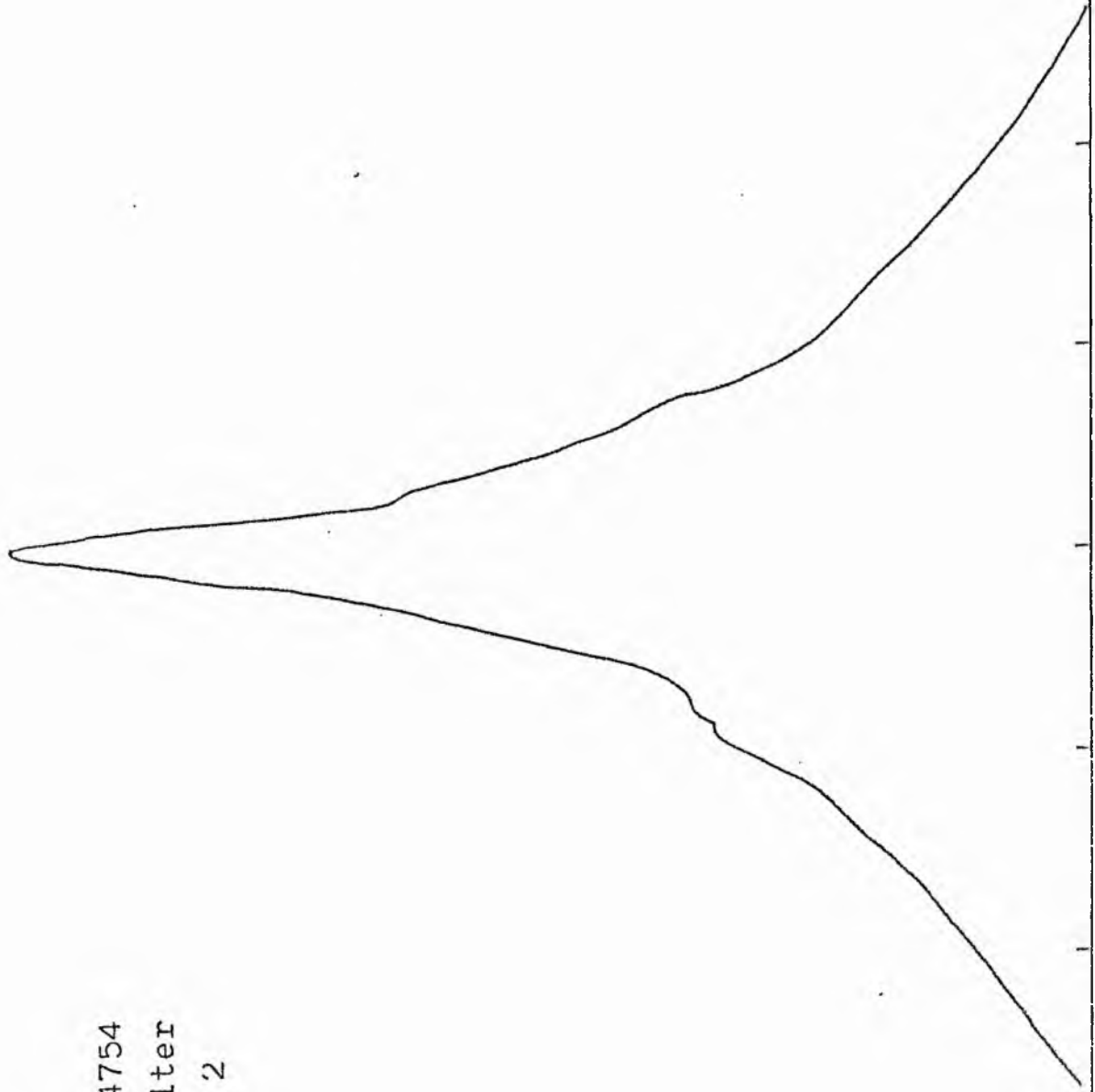
V -FILTER

Total luminosity	L_T	= 6.46
Total apparent magnitude	m_T	= 10.60
Apparent central surface brightness	μ_0	= 17.22
Major axis at threshold	$2a_m$	= 9.25
Minor axis at threshold	$2b_m$	= 6.05
Major axis at $\mu=25.0 \text{ mag sec}^{-2}$	$2a(25)$	= 5.67
Luminosity within $\mu=25.0 \text{ mag sec}^{-2}$	$k(25)$	= 0.84
Gradient of exponential component	$G(a)$	= -0.23
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.54
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.47
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.21
Axis ratio	b/a	= 0.92
Equivalent radius	r_1^*	= 0.19
Surface brightness	μ_1	= 19.37
Parameters at $k = \frac{1}{2}(\text{effective})$:		
Semi-major axis	a_e	= 0.80
Axis ratio	b/a	= 0.56
Equivalent radius	r_e^*	= 0.53
Surface brightness	μ_e	= 21.35
Mean surface brightness	μ_e'	= 11.22
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.83
Axis ratio	b/a	= 0.52
Equivalent radius	r_3^*	= 1.33
Surface brightness	μ_3	= 23.62
Concentration indices	$\left\{ \begin{array}{l} C_{21} \\ C_{32} \end{array} \right.$	= 2.91
		= 2.51

NGC 4754
B-Filter
Axis 1

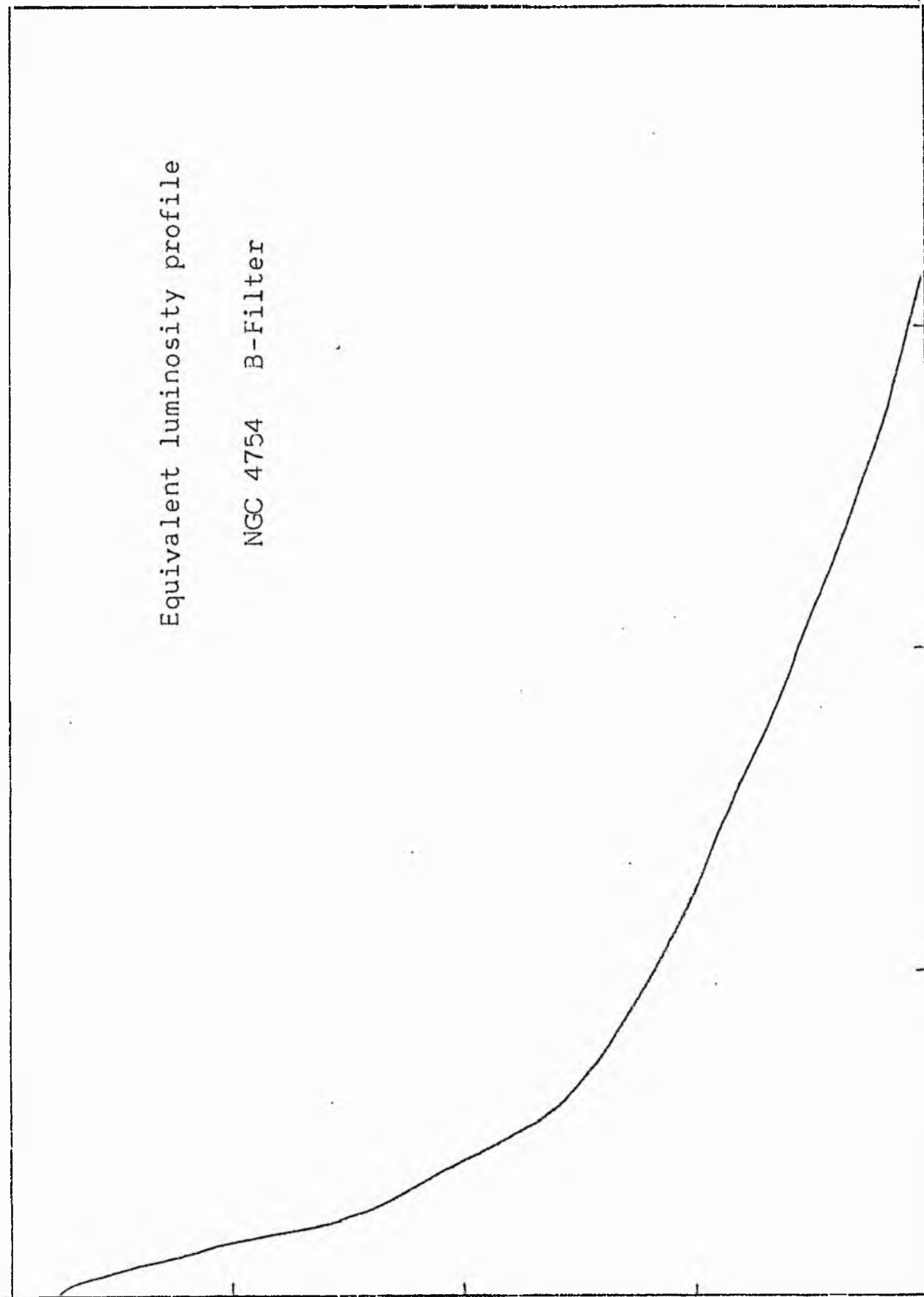


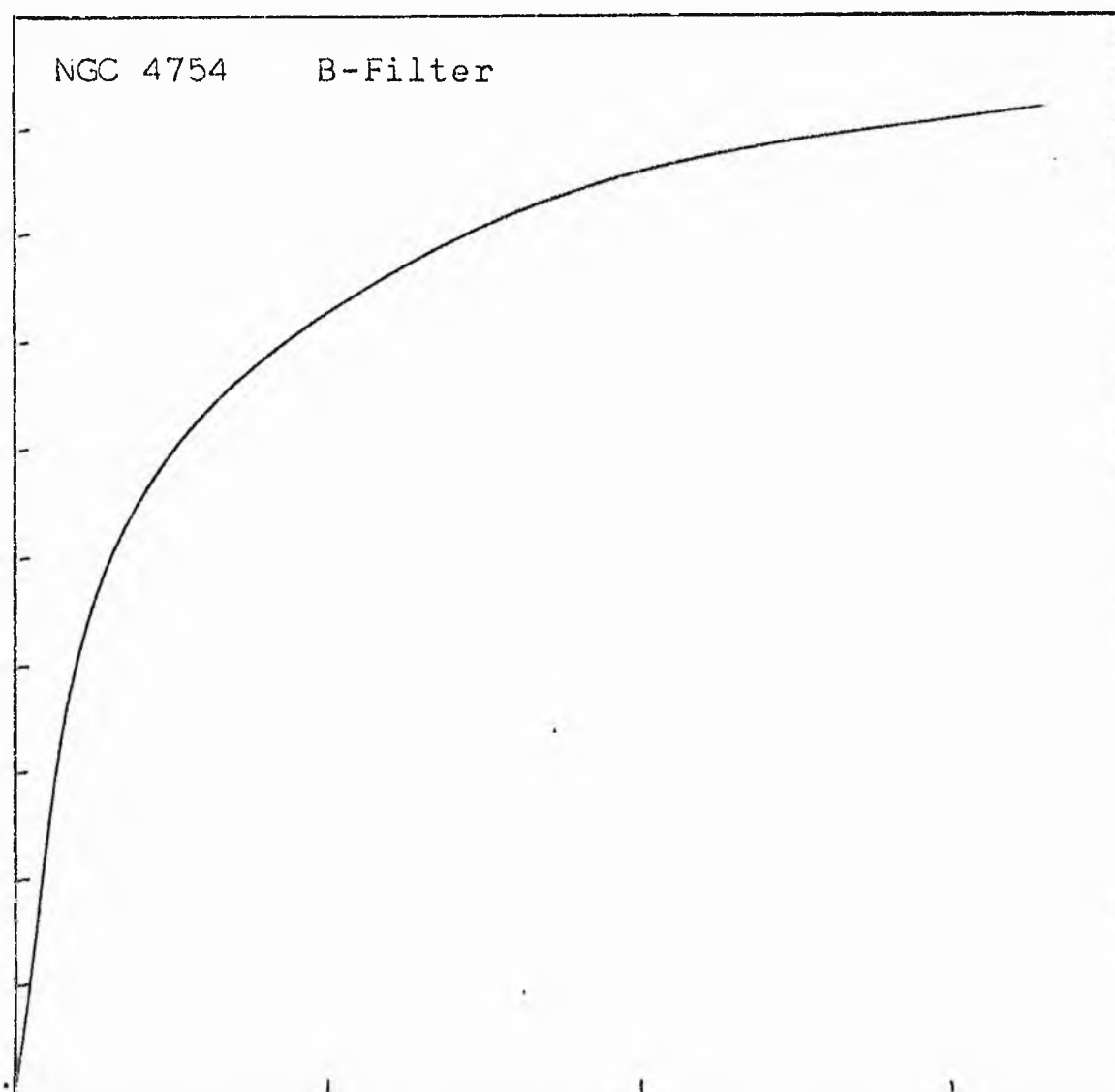
NGC 4754
B-Filter
Axis '2



Equivalent luminosity profile

NGC 4754 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

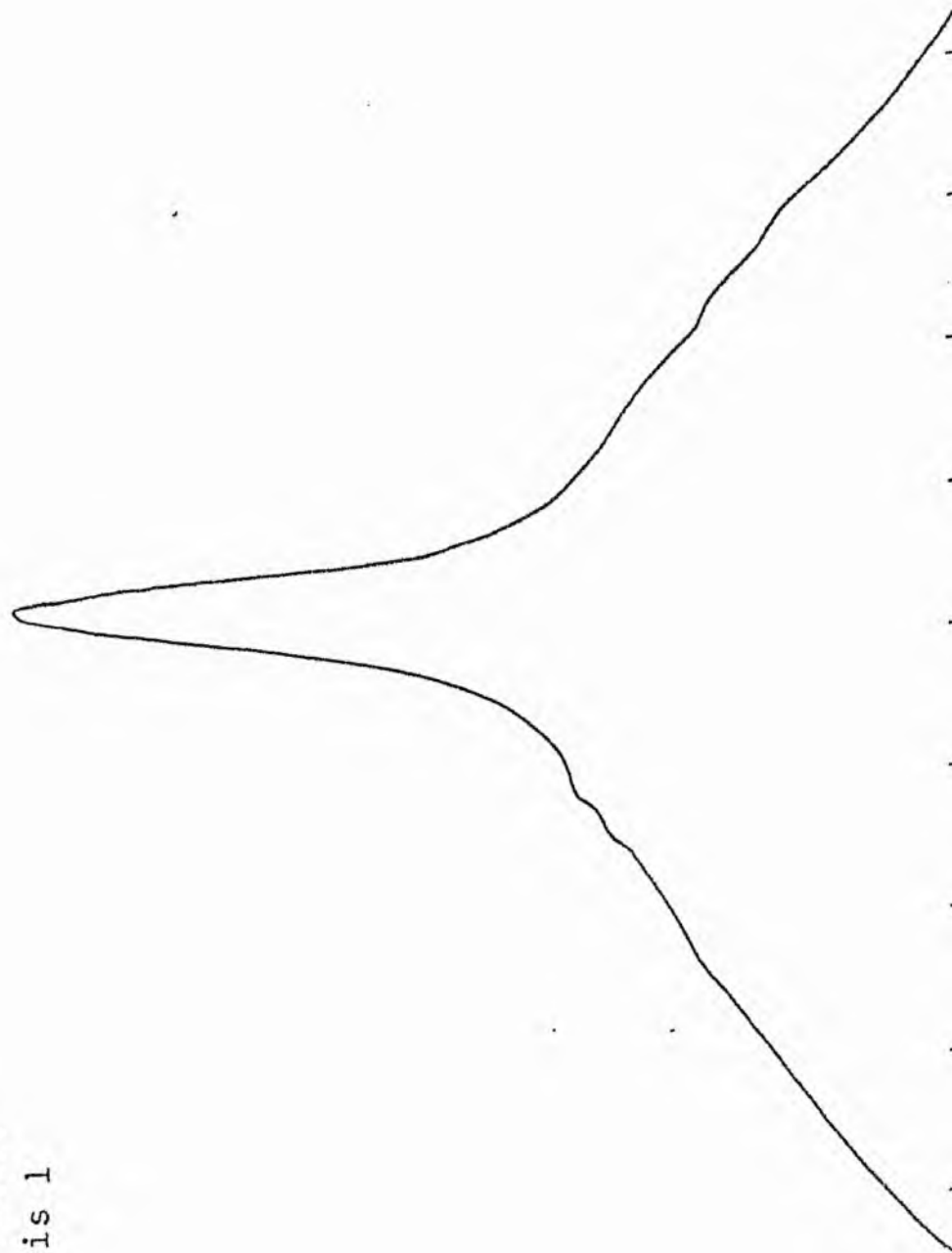
MEAN LUMINOSITY DISTRIBUTION IN NGC 4754
B COLUMN

LOG I	I	T	R	AREA	ΔA	P	ΣP	K(M)	ρ	LOG J	μ
1.72	52.401		0.0	0.0			0.0	0.0	0.0	1.411	17.45
1.70	50.119	51.300	2.01	12.69	12.69	651.1120	651.11	0.05	0.13	1.391	17.50
1.60	39.811	44.965	2.69	22.73	10.04	451.4690	1102.58	0.08	0.17	1.291	17.75
1.50	31.623	35.717	3.01	28.46	5.73	204.6656	1307.25	0.10	0.14	1.191	18.00
1.40	25.119	28.371	4.39	60.55	32.08	910.1875	2217.43	0.16	0.28	1.091	18.25
1.30	19.953	22.536	5.41	91.95	31.40	707.6965	2925.13	0.21	0.34	0.991	18.50
1.20	15.849	17.901	6.49	132.32	40.38	722.7534	3647.88	0.27	0.41	0.891	18.75
1.10	12.589	14.219	17.25	165.13	32.81	466.4661	4114.35	0.30	0.46	0.791	19.00
1.00	10.000	11.295	8.44	223.79	58.66	662.5039	4776.85	0.35	0.54	0.691	19.25
0.90	7.943	8.972	9.16	263.60	39.81	357.1609	5134.01	0.37	0.58	0.591	19.50
0.80	6.310	7.126	9.19	265.33	1.73	12.3233	5146.33	0.38	0.58	0.491	19.75
0.70	5.012	5.661	9.86	305.42	40.10	226.9818	5373.31	0.39	0.63	0.391	20.00
0.60	3.981	4.496	10.97	370.06	72.64	326.6113	5699.92	0.42	0.70	0.291	20.25
0.50	3.162	3.572	12.82	516.33	138.27	493.8386	6193.76	0.45	0.81	0.191	20.50
0.40	2.512	2.837	14.27	639.73	123.40	350.1033	6543.86	0.48	0.91	0.091	20.75
0.30	1.995	2.254	15.88	792.23	152.50	343.6628	6887.52	0.50	1.01	-0.009	21.00
0.20	1.585	1.790	17.68	982.01	189.78	339.7126	7227.23	0.53	1.12	-0.109	21.25
0.10	1.259	1.422	20.06	1264.19	282.18	401.2349	7628.46	0.56	1.27	-0.209	21.50
-0.00	1.000	1.129	22.28	1559.48	295.29	333.5212	7961.98	0.58	1.42	-0.309	21.75
-0.10	0.794	0.897	24.35	1862.72	303.24	272.0527	8234.04	0.60	1.55	-0.409	22.00
-0.20	0.631	0.713	25.00	1963.50	100.78	71.8174	8305.85	0.61	1.59	-0.509	22.25
-0.30	0.501	0.566	25.45	2034.82	71.32	40.3725	8346.22	0.61	1.62	-0.609	22.50
-0.40	0.398	0.450	29.03	2647.55	612.73	275.5117	8621.73	0.63	1.84	-0.709	22.75
-0.50	0.316	0.357	33.74	3576.35	928.80	331.7354	8953.47	0.65	2.14	-0.809	23.00
-0.60	0.251	0.284	39.59	4924.03	1347.68	382.3472	9335.81	0.68	2.52	-0.909	23.25
-0.70	0.200	0.225	46.57	6813.37	1889.34	425.7751	9761.59	0.71	2.96	-1.009	23.50
-0.80	0.158	0.179	54.68	9393.05	2579.68	461.7808	10223.36	0.75	3.47	-1.109	23.75
-0.90	0.126	0.142	60.50	11499.02	2105.96	299.4478	10522.81	0.77	3.84	-1.209	24.00
-1.00	0.100	0.113	66.00	13684.78	2185.76	246.8728	10769.68	0.79	4.19	-1.309	24.25
-1.10	0.079	0.090	72.50	16513.00	2828.22	253.7372	11023.41	0.80	4.61	-1.409	24.50
-1.20	0.063	0.071	75.50	17907.86	1394.87	99.4041	11122.82	0.81	4.80	-1.509	24.75
-1.30	0.050	0.057	86.50	23506.18	5598.32	316.9055	11439.72	0.83	5.50	-1.609	25.00
-1.40	0.040	0.045	97.00	29559.25	6053.06	272.1748	11711.89	0.85	6.16	-1.709	25.25
-1.50	0.032	0.036	105.60	35033.02	5473.77	195.5062	11907.39	0.87	6.71	-1.809	25.50
-1.60	0.025	0.028	116.70	42785.00	7751.98	219.9312	12127.32	0.88	7.42	-1.909	25.75
-1.70	0.020	0.023	125.80	49717.71	6932.71	156.2340	12283.56	0.90	7.99	-2.009	26.00
-1.80	0.016	0.018	136.70	58706.59	8988.88	160.9092	12444.46	0.91	8.69	-2.109	26.25
-1.90	0.013	0.014	147.20	68071.50	9364.91	133.1617	12577.62	0.92	9.35	-2.209	26.50
-2.00	0.010	0.011	160.00	80424.75	12353.25	139.5268	12717.15	0.93	10.17	-2.309	26.75
							13717.00	(1)			∞

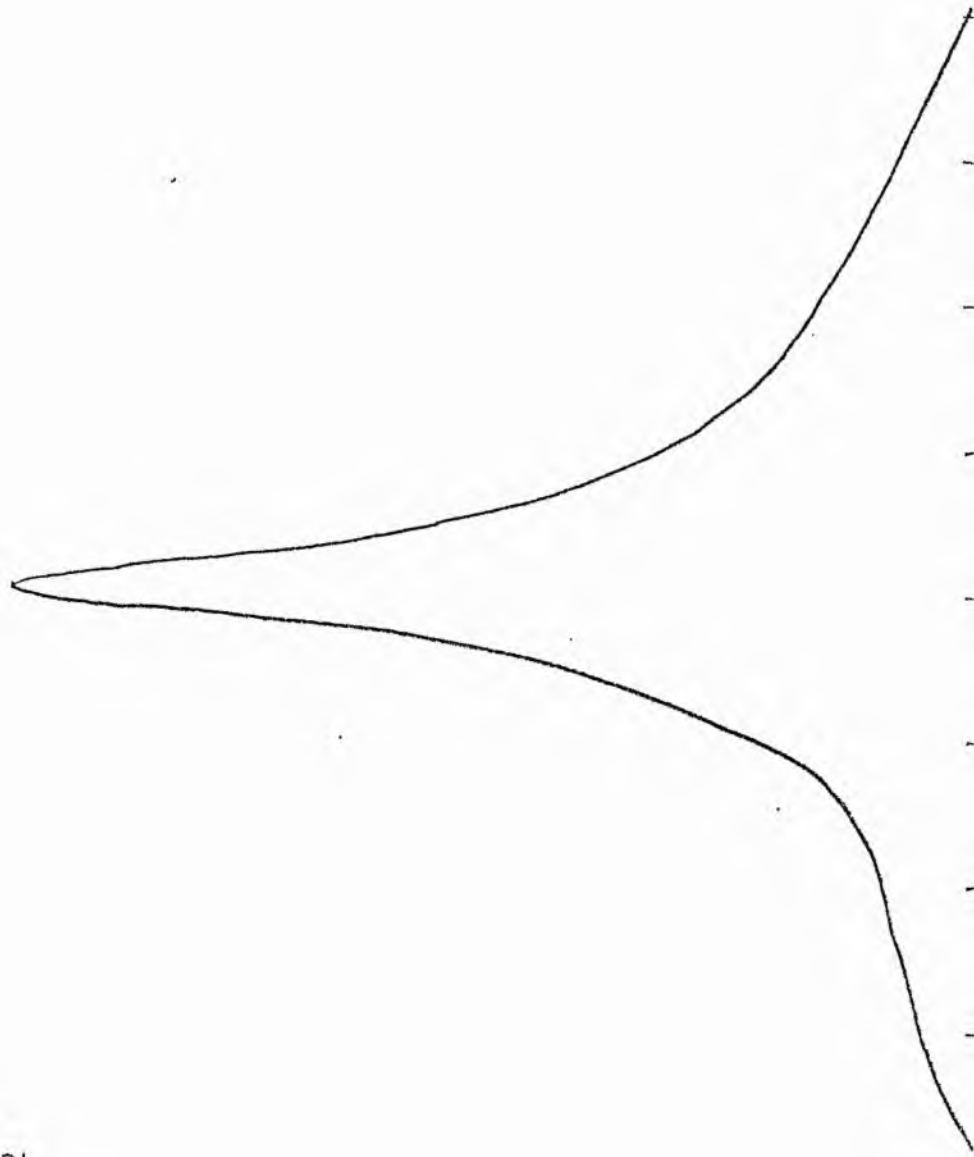
PHOTOMETRIC PARAMETERS OF NGC 4754
B-FILTER

Total luminosity	L_T	= 3.81
Total apparent magnitude	m_T	= 11.43
Apparent central surface brightness	μ_o	= 17.45
Major axis at threshold	$2a_m$	= 5.68
Minor axis at threshold	$2b_m$	= 4.37
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 3.53
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.83
Gradient of exponential component	$G(a)$	= -0.67
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.62
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.23
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.15
Axis ratio	b/a	= 0.77
Equivalent radius	r_1^*	= 0.10
Surface brightness	μ_1	= 18.67
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.28
Axis ratio	b/a	= 0.88
Equivalent radius	r_e^*	= 0.26
Surface brightness	μ_e	= 21.00
Mean surface brightness	μ_e'	= 10.50
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 0.96
Axis ratio	b/a	= 0.81
Equivalent radius	r_3^*	= 0.93
Surface brightness	μ_3	= 23.75
Concentration indices	$\begin{Bmatrix} C_{21} \\ C_{32} \end{Bmatrix}$	$\begin{matrix} = 2.56 \\ = 3.55 \end{matrix}$

NGC 4754
V-Filter
Axis 1

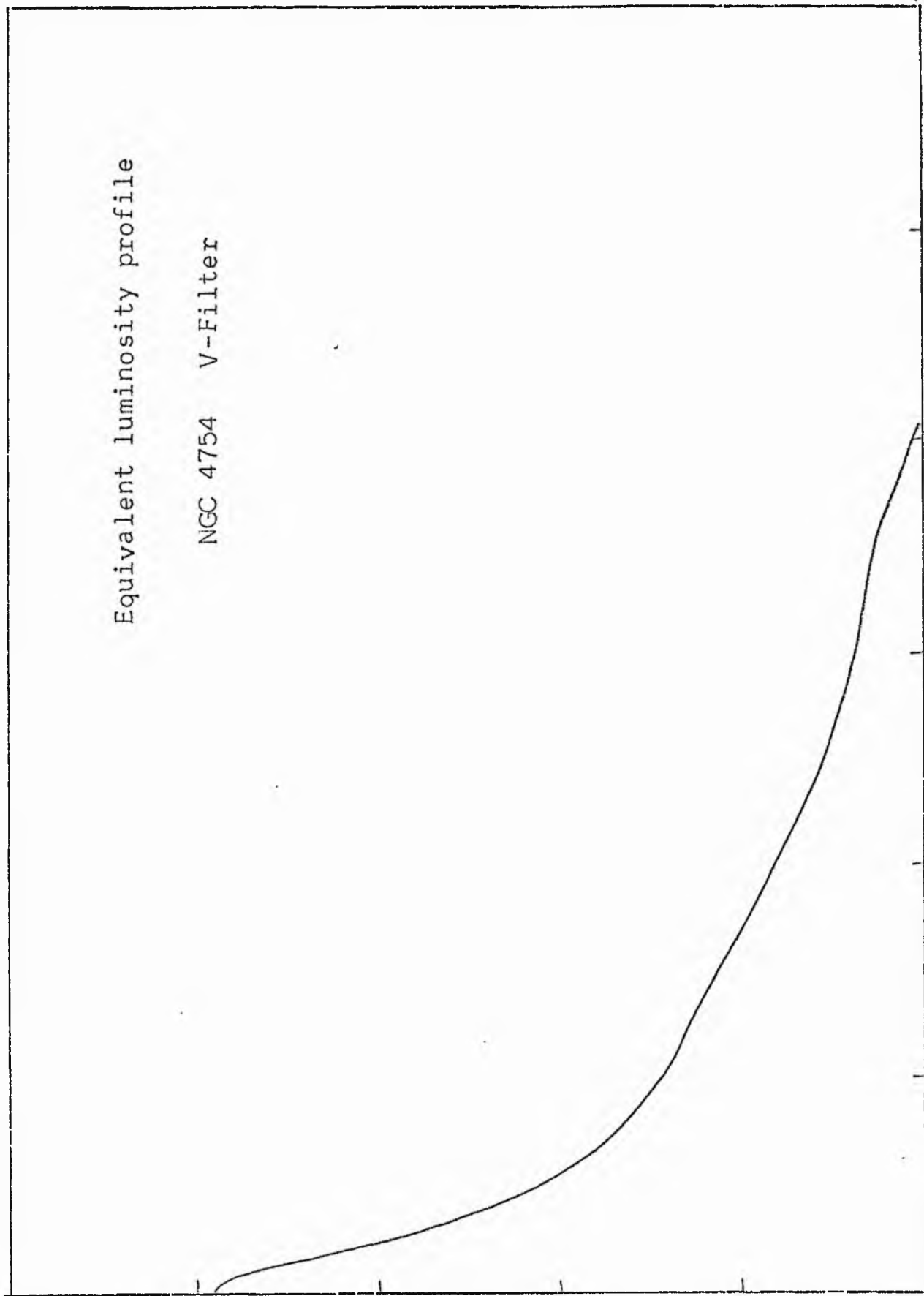


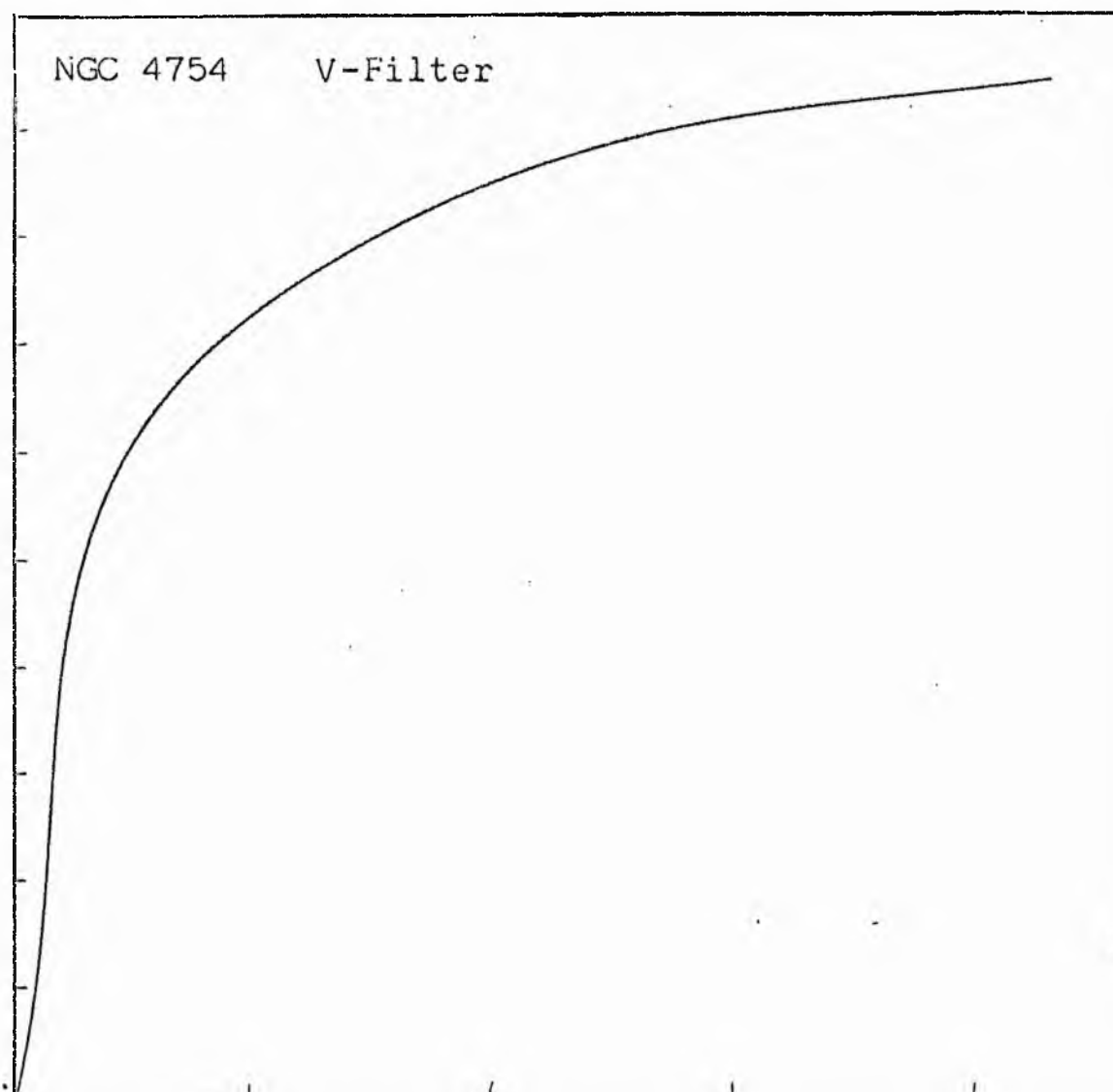
NGC 4754
V-Filter
Axis '2



Equivalent luminosity profile

NGC 4754 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4754
V COLOUR

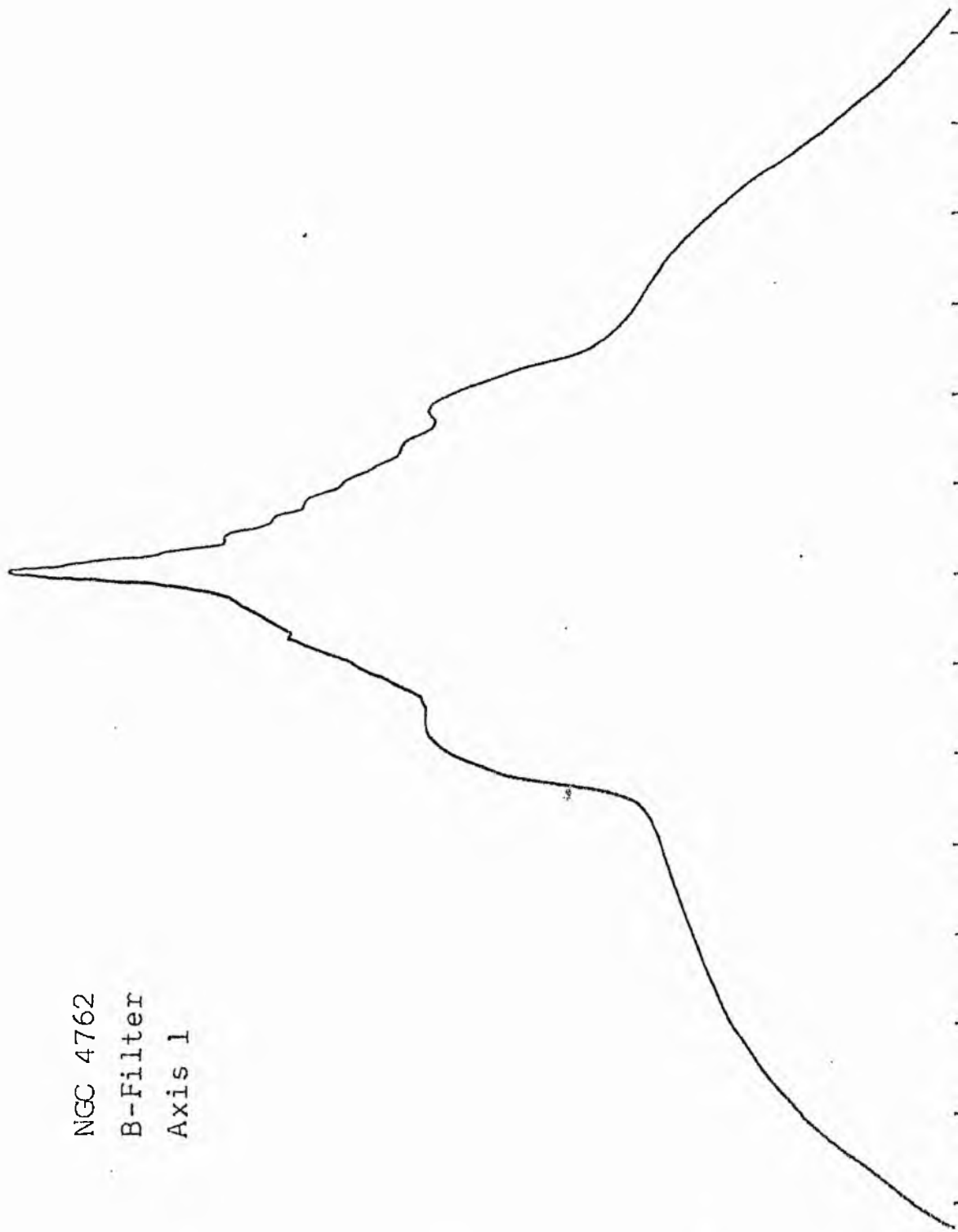
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	P	LOG J	μ
1.84	69.183		0.0	0.0			0.0	0.0	0.0	1.060	16.41
1.80	63.096	66.139	2.10	13.85	13.85	916.3210	916.32	0.03	0.15	1.020	17.01
1.70	50.119	56.607	3.90	47.78	33.93	1920.6331	2836.95	0.11	0.28	0.920	17.26
1.60	39.811	44.965	4.80	72.38	24.60	1106.0662	3943.02	0.15	0.34	0.820	17.51
1.50	31.623	35.717	6.20	120.76	48.38	1727.9888	5671.01	0.22	0.44	0.720	17.76
1.40	25.119	28.371	8.56	230.20	109.43	3104.6670	8775.69	0.34	0.61	0.620	18.01
1.30	19.952	22.536	9.04	256.74	26.54	598.0959	9373.79	0.36	0.64	0.520	18.26
1.20	15.849	17.901	9.20	265.90	9.17	164.1213	9537.91	0.36	0.65	0.420	18.51
1.10	12.589	14.219	9.98	312.90	47.00	668.2849	10206.19	0.39	0.71	0.320	18.76
1.00	10.000	11.295	11.40	408.28	95.38	1017.2427	11283.43	0.43	0.81	0.220	19.01
0.90	7.943	8.972	13.20	547.39	139.11	1248.0332	12531.46	0.48	0.93	0.120	19.26
0.80	6.310	7.126	13.97	613.12	65.72	468.3794	12999.84	0.50	0.99	0.020	19.51
0.70	5.012	5.661	14.89	696.53	83.41	472.1743	13472.01	0.51	1.05	-0.080	19.76
0.60	3.981	4.496	15.90	794.23	97.70	439.2871	13911.30	0.53	1.12	-0.180	20.01
0.50	3.162	3.572	16.70	876.16	81.93	292.6343	14203.93	0.54	1.18	-0.280	20.26
0.40	2.512	2.837	18.60	1086.86	210.71	597.7837	14801.71	0.57	1.32	-0.380	20.51
0.30	1.995	2.254	21.20	1411.96	325.09	732.6152	15534.32	0.59	1.50	-0.480	20.76
0.20	1.585	1.790	23.30	1705.54	293.58	525.5286	16059.85	0.61	1.65	-0.580	21.01
0.10	1.259	1.422	24.50	1885.74	180.20	256.2312	16316.08	0.62	1.73	-0.680	21.26
-0.00	1.000	1.129	25.70	2074.99	189.25	213.7479	16529.83	0.63	1.82	-0.780	21.51
-0.10	0.794	0.897	28.40	2533.88	458.89	411.6985	16941.52	0.65	2.01	-0.880	21.76
-0.20	0.631	0.713	31.50	3117.25	583.36	415.7271	17357.25	0.66	2.23	-0.980	22.01
-0.30	0.501	0.566	35.40	3936.92	819.67	463.9895	17821.24	0.68	2.50	-1.080	22.26
-0.40	0.398	0.450	39.70	4951.43	1014.51	456.1697	18277.41	0.70	2.81	-1.180	22.51
-0.50	0.316	0.357	44.80	6305.30	1353.87	483.5540	18760.96	0.72	3.17	-1.280	22.76
-0.60	0.251	0.284	54.50	9331.32	3026.02	858.5010	19619.46	0.75	3.86	-1.380	23.01
-0.70	0.200	0.225	62.50	12271.84	2940.53	662.6655	20282.12	0.77	4.42	-1.480	23.26
-0.80	0.158	0.179	68.90	14913.79	2641.95	472.9258	20755.05	0.79	4.87	-1.580	23.51
-0.90	0.126	0.142	74.50	17436.62	2522.84	358.7219	21113.77	0.81	5.27	-1.680	23.76
-1.00	0.100	0.113	90.50	25730.43	8293.80	936.7493	22050.51	0.84	6.40	-1.780	24.01
-1.10	0.079	0.090	95.60	28712.13	2981.70	267.5066	22318.02	0.85	6.76	-1.880	24.26
-1.20	0.063	0.071	99.30	30917.62	2265.49	161.4482	22479.46	0.86	7.02	-1.980	24.51
-1.30	0.050	0.057	108.60	37051.81	6074.19	343.8423	22823.30	0.87	7.68	-2.080	24.76
-1.40	0.040	0.045	115.40	41837.08	4785.27	215.1684	23038.47	0.88	8.16	-2.180	25.01
-1.50	0.032	0.036	132.50	55154.59	13317.51	475.6587	23514.12	0.90	9.37	-2.280	25.26
-1.60	0.025	0.028	149.20	69933.81	14779.22	419.2998	23933.42	0.91	10.56	-2.380	25.51
-1.70	0.020	0.023	172.80	93807.44	23873.62	538.0120	24471.43	0.93	12.22	-2.480	25.76
-1.80	0.016	0.018	180.30	102127.12	8319.69	148.9297	24620.36	0.94	12.76	-2.580	26.01
-1.90	0.013	0.014	194.70	117091.75	16964.62	241.2233	24861.59	0.95	13.77	-2.680	26.26
-2.00	0.010	0.011	208.00	135917.87	16826.12	190.0464	25051.63	0.96	14.71	-2.780	26.51
-∞							26181.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4754

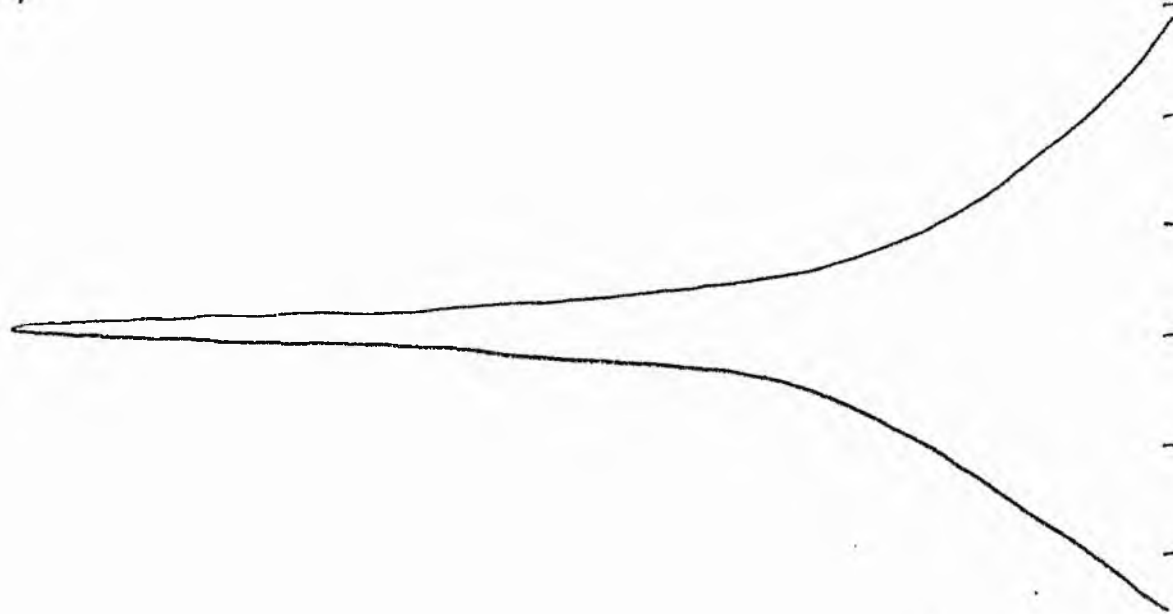
V-FILTER

Total luminosity	L_T	= 7.27
Total apparent magnitude	m_T	= 10.46
Apparent central surface brightness	μ_0	= 16.91
Major axis at threshold	$2a_m$	= 7.28
Minor axis at threshold	$2b_m$	= 6.48
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 5.41
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.88
Gradient of exponential component	$G(a)$	= -0.50
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.51
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.17
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.13
Axis ratio	b/a	= 0.93
Equivalent radius	r_1^*	= 0.11
Surface brightness	μ_1	= 17.82
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 0.25
Axis ratio	b/a	= 0.93
Equivalent radius	r_e^*	= 0.24
Surface brightness	μ_e	= 19.51
Mean surface brightness	μ_e'	= 9.36
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 1.21
Axis ratio	b/a	= 0.49
Equivalent radius	r_3^*	= 0.91
Surface brightness	μ_3	= 23.01
Concentration indices	$\begin{cases} C_{21} \\ C_{32} \end{cases}$	$\begin{cases} = 2.05 \\ = 3.87 \end{cases}$

NGC 4762
B-Filter
Axis 1

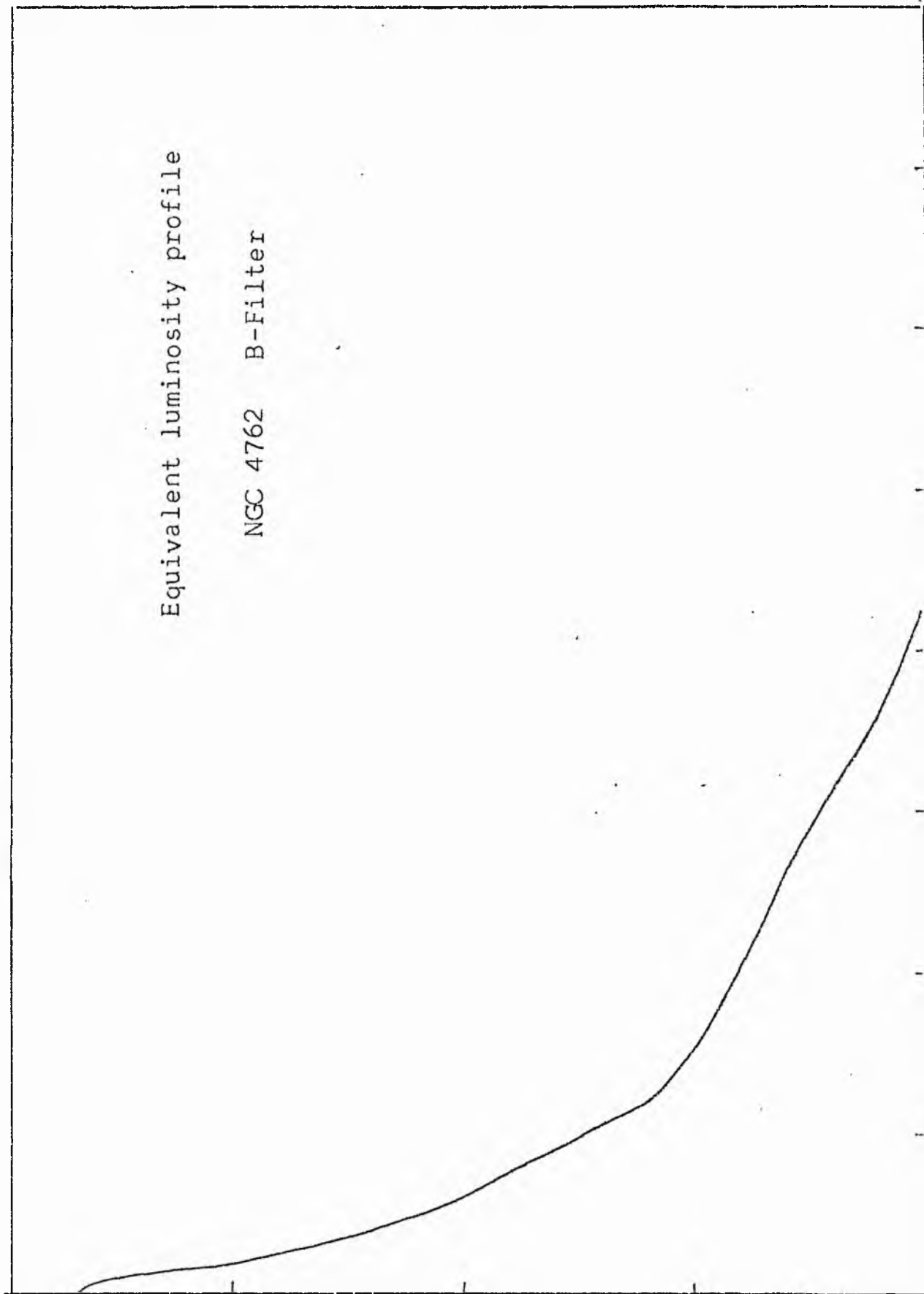


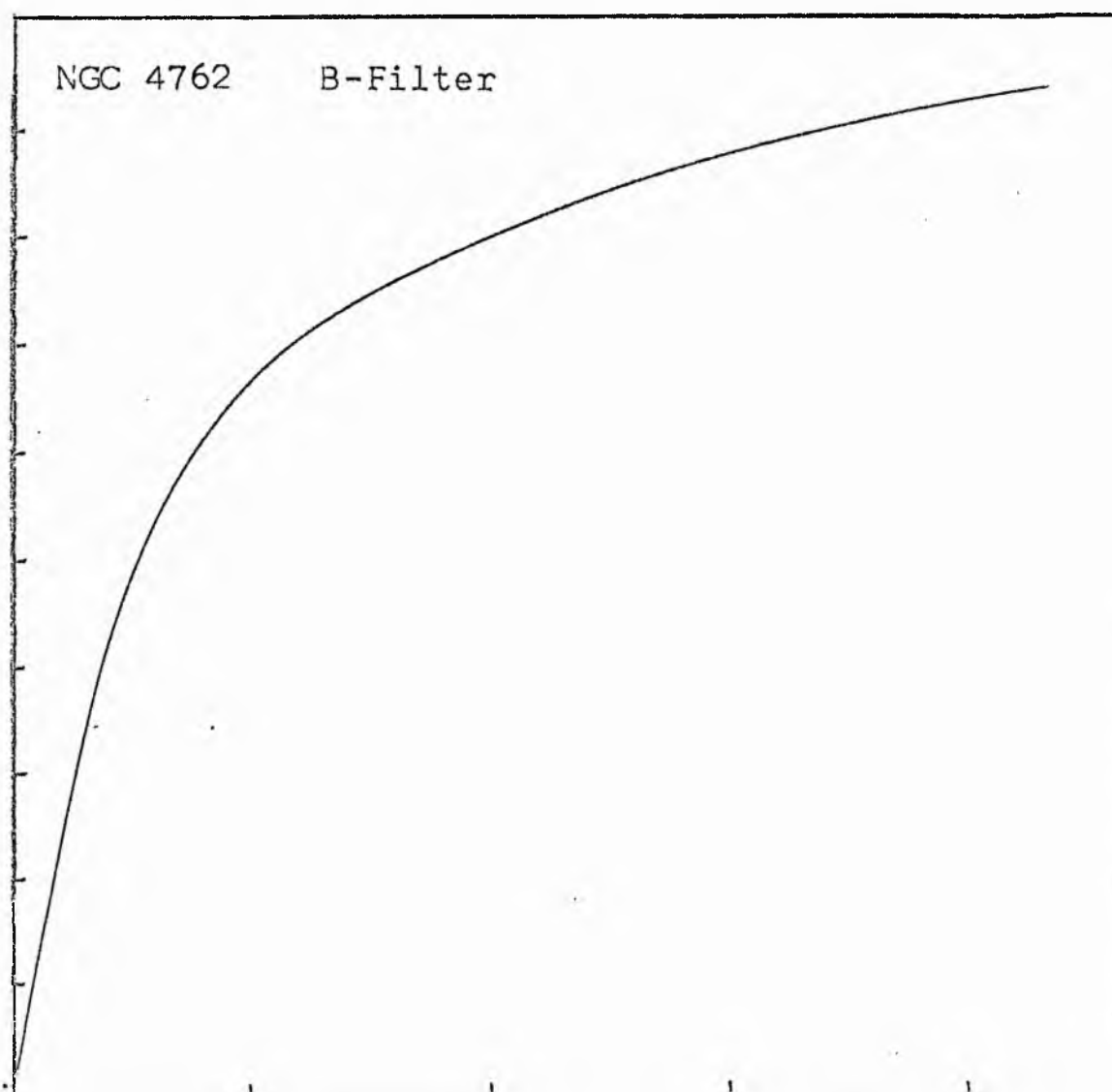
NGC 4762
B-Filter
Axis 2



Equivalent luminosity profile

NGC 4762 B-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4762
B COLOUR

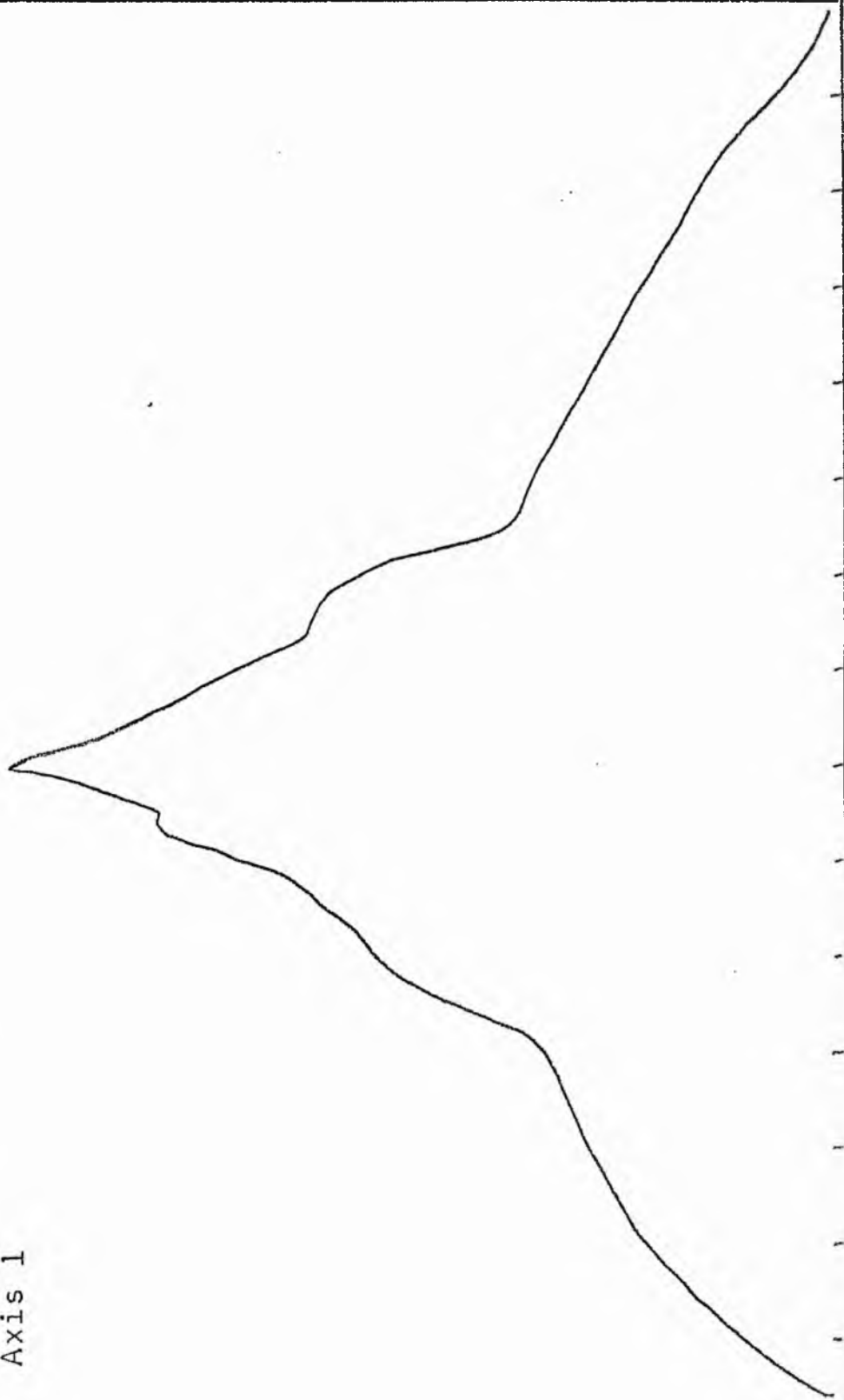
LOG I	I	T	R	AREA	ΔA	P	ΣP	K(R)	ρ	LOG J	μ
1.65	44.668	42.239	0.0	0.0	16.05	677.7734	0.0	0.0	0.0	1.518	17.73
1.60	39.811	35.717	2.26	16.05	16.33	583.0850	677.77	0.03	0.08	1.468	17.86
1.50	31.623	28.371	3.21	32.37	39.11	1109.5520	1260.86	0.06	0.12	1.368	18.11
1.40	25.119	22.536	4.77	71.48	16.77	377.8552	2370.41	0.12	0.18	1.268	18.36
1.30	19.953	17.901	5.30	88.25	13.82	247.4418	2748.27	0.14	0.20	1.168	18.61
1.20	15.849	14.219	5.70	102.07	7.29	103.6350	2995.71	0.15	0.21	1.068	18.86
1.10	12.589	11.295	5.90	109.36	128.43	1450.5408	3099.34	0.15	0.22	0.968	19.11
1.00	10.000	8.972	8.70	237.79	11.06	99.2115	4549.88	0.22	0.32	0.868	19.36
0.90	7.943	7.126	8.90	248.85	97.52	694.9297	4649.09	0.23	0.33	0.768	19.61
0.80	6.310	5.661	10.50	346.36	144.51	818.0442	5344.02	0.26	0.34	0.668	19.86
0.70	5.012	4.496	12.50	490.87	107.41	482.9670	6162.07	0.30	0.47	0.568	20.11
0.60	3.981	3.572	13.80	598.28	185.98	664.2646	6645.03	0.33	0.51	0.468	20.36
0.50	3.162	2.837	15.80	784.27	299.09	848.5488	7309.29	0.36	0.54	0.368	20.61
0.40	2.512	2.254	18.57	1083.36	211.26	476.0789	8157.84	0.40	0.64	0.268	20.86
0.30	1.995	1.790	20.30	1294.62	381.77	683.3679	8633.92	0.42	0.76	0.168	21.11
0.20	1.585	1.422	23.10	1676.38	834.68	1193.9485	9317.30	0.46	0.86	0.068	21.36
0.10	1.259	1.129	28.30	2516.07	761.52	860.1055	10511.25	0.52	1.06	-0.032	21.61
-0.00	1.000	0.897	32.30	3277.59	703.94	631.5437	11371.36	0.56	1.20	-0.132	21.86
-0.10	0.794	0.713	35.60	3981.53	365.93	260.7764	12002.90	0.59	1.33	-0.232	22.11
-0.20	0.631	0.566	37.20	4347.46	780.12	441.6028	12263.67	0.60	1.39	-0.332	22.36
-0.30	0.501	0.450	40.40	5127.58	817.10	367.4038	12705.27	0.62	1.51	-0.432	22.61
-0.40	0.398	0.357	43.50	5944.68	1054.29	376.5559	13072.68	0.64	1.62	-0.532	22.86
-0.50	0.316	0.284	47.20	6998.96	1076.46	305.4006	13449.23	0.66	1.76	-0.632	23.11
-0.60	0.251	0.225	50.70	8075.43	1290.16	290.7451	13754.63	0.68	1.89	-0.732	23.36
-0.70	0.200	0.179	54.60	9365.59	1422.51	254.6396	14045.37	0.69	2.04	-0.832	23.61
-0.80	0.158	0.142	58.60	10788.10	2772.55	394.2300	14300.01	0.70	2.19	-0.932	23.86
-0.90	0.126	0.113	65.70	13560.65	4489.80	507.1045	14694.23	0.72	2.45	-1.032	24.11
-1.00	0.100	0.090	75.80	18050.45	5947.42	533.5798	15201.34	0.75	2.83	-1.132	24.36
-1.10	0.079	0.071	87.40	23997.87	7733.00	551.0862	15734.91	0.77	3.26	-1.232	24.61
-1.20	0.063	0.057	100.50	31730.87	10542.40	596.7766	16286.00	0.80	3.75	-1.332	24.86
-1.30	0.050	0.045	116.00	42273.27	12465.84	560.5242	16882.77	0.83	4.33	-1.432	25.11
-1.40	0.040	0.036	132.00	54739.11	9054.86	323.4111	17443.30	0.86	4.92	-1.532	25.36
-1.50	0.032	0.028	142.50	63793.97	5671.91	160.9175	17766.71	0.87	5.32	-1.632	25.61
-1.60	0.025	0.023	148.70	69465.87	17940.25	404.2996	17927.62	0.88	5.55	-1.732	25.86
-1.70	0.020	0.018	166.80	87406.12	19186.94	343.4636	18331.92	0.90	6.22	-1.832	26.11
-1.80	0.016	0.014	184.20	106593.06	13602.25	193.4134	18675.38	0.92	6.87	-1.932	26.36
-1.90	0.013	0.011	195.60	120195.31	16507.81	186.4515	18868.79	0.93	7.30	-2.032	26.61
-2.00	0.010		208.60	136703.12			19055.24	0.94	7.78	-2.132	26.86
-∞							20355.00	(1)			CO

PHOTOMETRIC PARAMETERS OF NGC 4762

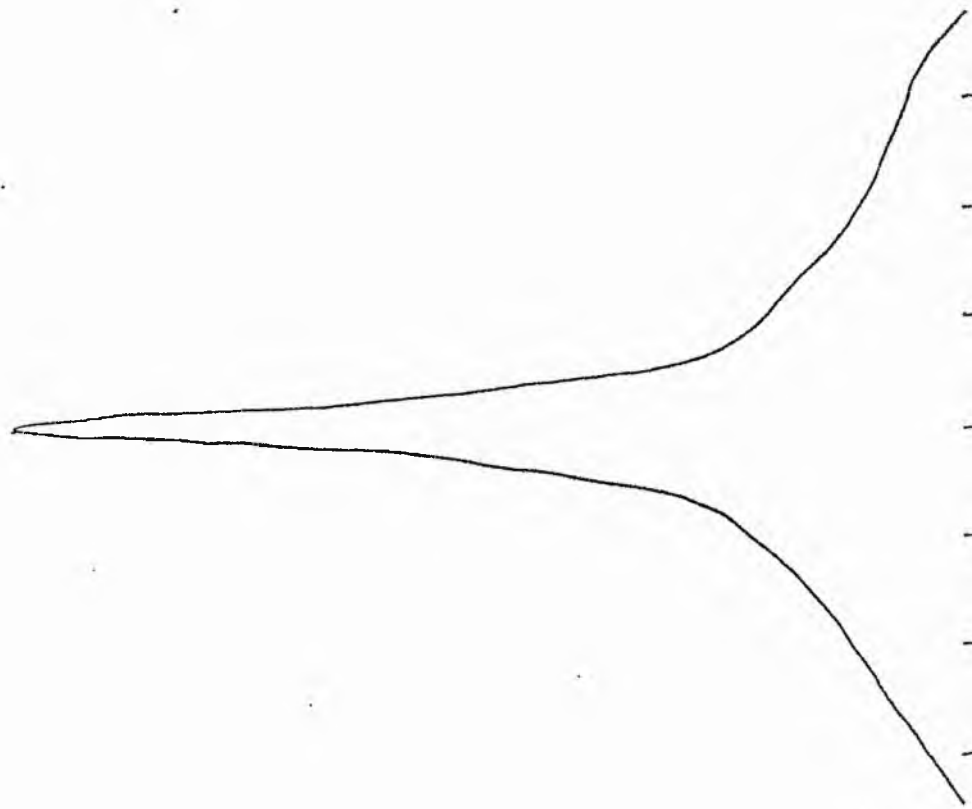
B-FILTER

Total luminosity	L_T	= 5.65
Total apparent magnitude	m_T	= 11.09
Apparent central surface brightness	μ_o	= 17.73
Major axis at threshold	$2a_m$	= 11.01
Minor axis at threshold	$2b_m$	= 4.06
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 7.92
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.82
Gradient of exponential component	$G(a)$	= -0.44
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.48
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.84
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.30
Axis ratio	b/a	= 0.44
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 19.78
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.12
Axis ratio	b/a	= 0.25
Equivalent radius	r_e^*	= 0.45
Surface brightness	μ_e	= 21.53
Mean surface brightness	μ_e'	= 11.34
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.83
Axis ratio	b/a	= 0.26
Equivalent radius	r_3^*	= 1.29
Surface brightness	μ_3	= 24.36
Concentration indices	$\begin{Bmatrix} C_{21} \\ C_{32} \end{Bmatrix}$	$\begin{matrix} = 2.71 \\ = 2.88 \end{matrix}$

NGC 4762
V-Filter
Axis 1

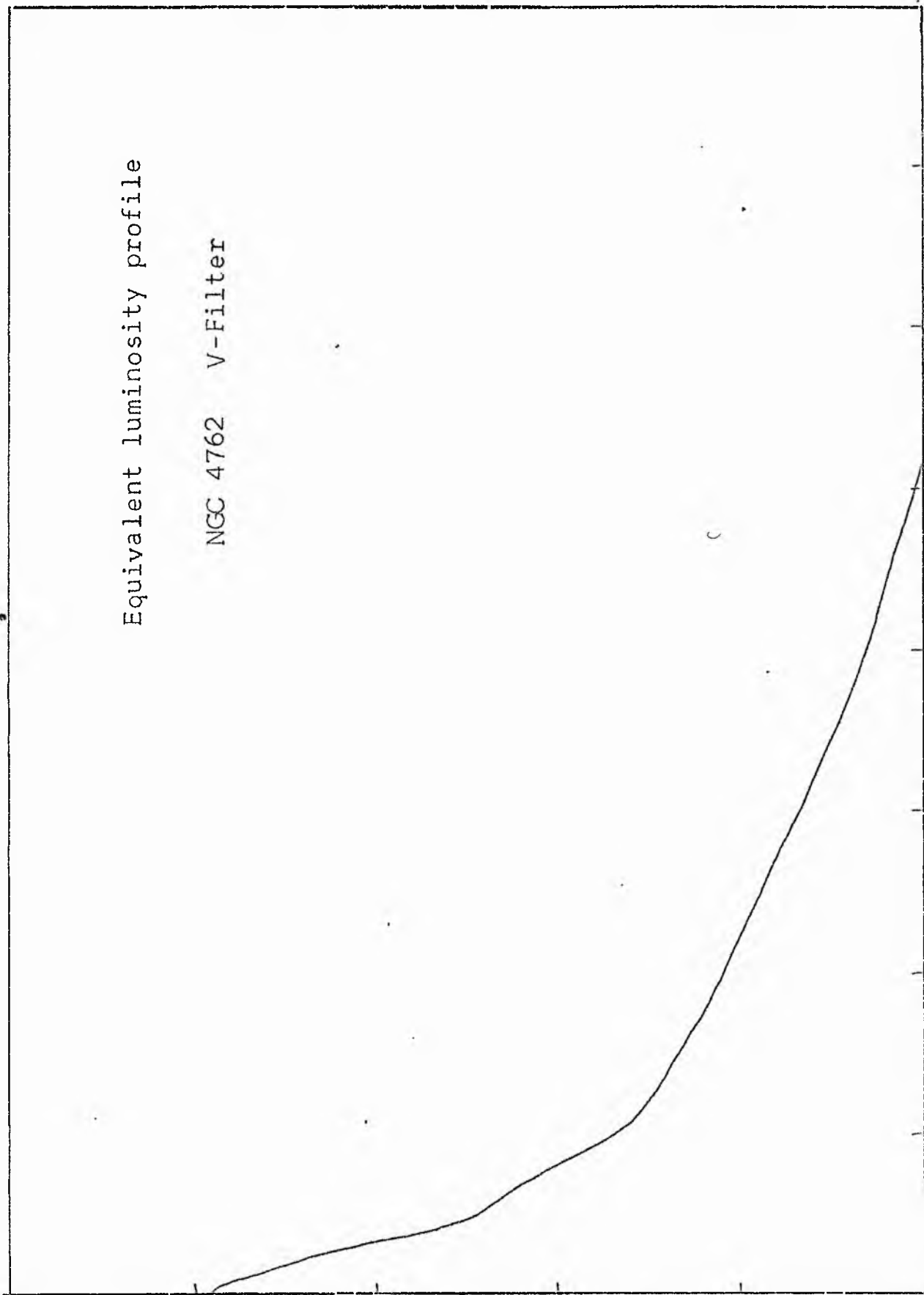


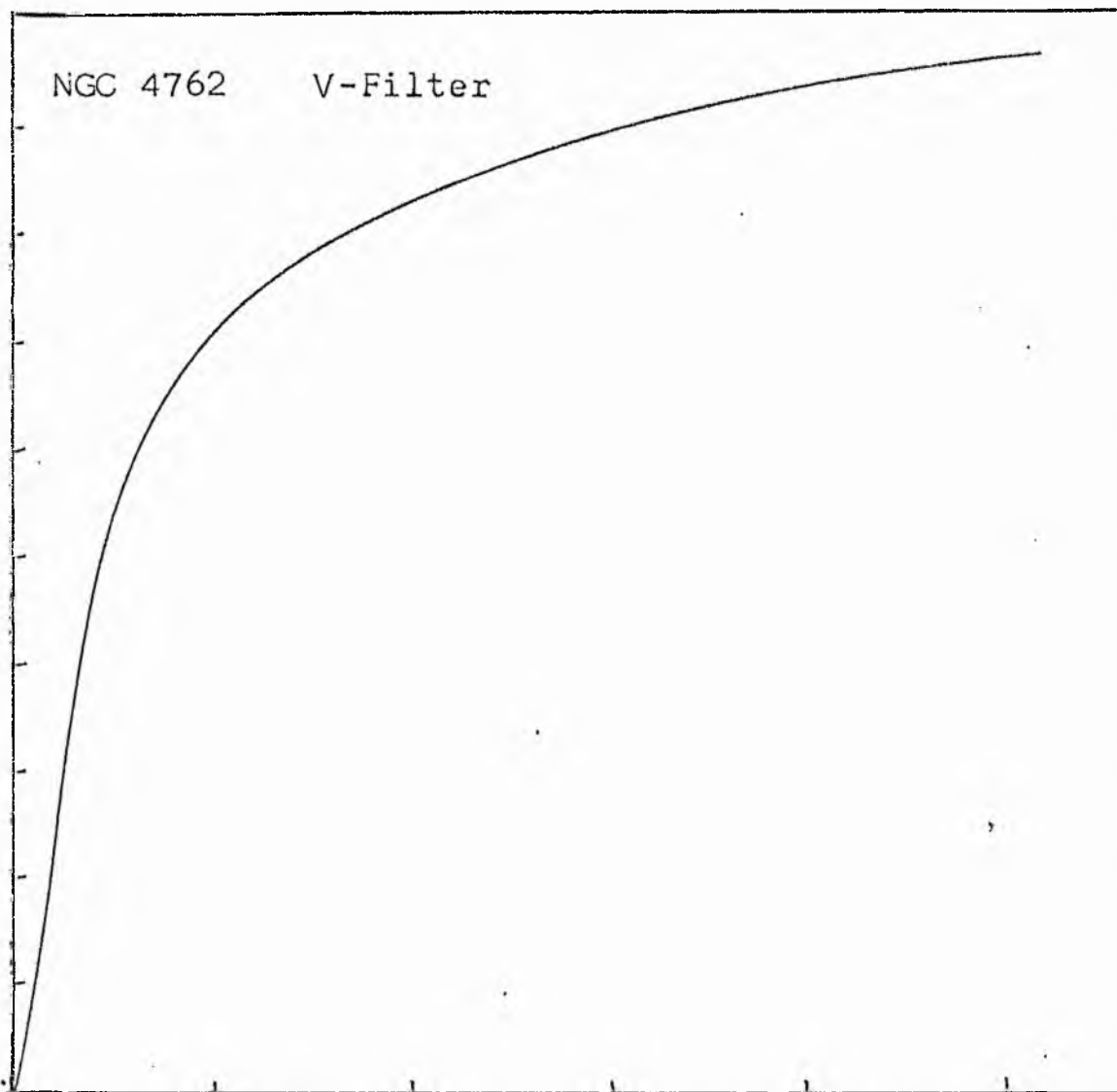
NGC 4762
V-Filter
Axis 2



Equivalent luminosity profile

NGC 4762 V-Filter





Relative integrated luminosity $k(r)$ versus
equivalent radius r^* .

MEAN LUMINOSITY DISTRIBUTION IN NGC 4762
V COLOUR

LOG I	I	T	R	AREA	AA	P	ΣP	K(R)	ρ	LOG J	μ
1.82	66.069		0.0	0.0			0.0	0.0	0.0	1.395	17.33
1.80	63.096	64.582			4.52	242.1638					
		56.607	1.20	4.52			292.16	0.01	0.06	1.375	17.38
1.70	50.119	44.965	2.70	22.90	18.38	1040.3425	1332.51	0.03	0.12	1.275	17.63
1.60	39.811	35.717	4.10	52.81	29.91	1344.7964	2677.30	0.07	0.19	1.175	17.88
1.50	31.623	28.371	5.40	91.61	38.80	1345.7571	4063.06	0.10	0.25	1.075	18.13
1.40	25.119	22.536	9.50	203.53	191.92	5444.8906	9507.95	0.24	0.44	0.975	18.38
1.30	19.952	17.901	10.50	346.36	62.83	1415.9573	10923.91	0.27	0.48	0.875	18.63
1.20	15.849	14.219	11.20	394.08	47.72	854.2314	11778.14	0.29	0.51	0.775	18.88
1.10	12.589	11.295	13.40	564.10	170.02	2417.5525	14195.69	0.35	0.61	0.675	19.13
1.00	10.000	8.972	14.90	697.46	133.36	1506.2444	15701.93	0.39	0.68	0.575	19.38
0.90	7.943	7.126	15.70	774.37	76.91	689.9692	16391.90	0.41	0.72	0.475	19.63
0.80	6.310	5.661	17.50	962.11	187.74	1337.9165	17729.81	0.44	0.80	0.375	19.88
0.70	5.012	4.496	19.66	1214.27	252.16	1427.3982	19157.21	0.48	0.90	0.275	20.13
0.60	3.981	3.572	19.97	1252.87	38.59	173.5398	19330.75	0.48	0.92	0.175	20.38
0.50	3.162	2.837	20.97	1381.48	128.62	459.3723	19790.12	0.49	0.96	0.075	20.63
0.40	2.512	2.254	25.70	2074.99	693.51	1967.5161	21757.63	0.54	1.18	-0.025	20.88
0.30	1.995	1.790	30.80	2980.24	905.25	2040.0295	23797.66	0.59	1.41	-0.125	21.13
0.20	1.585	1.422	34.30	3696.05	715.81	1281.3489	25079.01	0.62	1.57	-0.225	21.38
0.10	1.259	1.129	37.10	4324.12	628.07	893.0486	25972.05	0.64	1.70	-0.325	21.63
-0.00	1.000	0.897	39.80	4976.40	652.29	736.7263	26708.78	0.66	1.83	-0.425	21.88
-0.10	0.794	0.713	41.40	5384.56	408.16	366.1838	27074.96	0.67	1.90	-0.525	22.13
-0.20	0.631	0.566	45.50	6503.88	1119.32	797.6699	27872.63	0.69	2.09	-0.625	22.38
-0.30	0.501	0.450	47.80	7178.03	674.15	381.6138	28254.24	0.70	2.19	-0.725	22.63
-0.40	0.398	0.357	50.30	7948.51	770.48	346.4402	28600.68	0.71	2.31	-0.825	22.88
-0.50	0.316	0.284	55.80	9781.78	1833.27	654.7812	29255.46	0.73	2.56	-0.925	23.13
-0.60	0.251	0.225	64.50	13069.81	3288.03	932.8345	30188.29	0.75	2.96	-1.025	23.38
-0.70	0.200	0.179	75.80	18050.45	4980.64	1122.4172	31310.71	0.78	3.48	-1.125	23.63
-0.80	0.158	0.142	87.70	24162.90	6112.45	1094.1692	32404.87	0.80	4.02	-1.225	23.88
-0.90	0.126	0.113	94.60	28114.61	3951.71	561.8933	32966.77	0.82	4.34	-1.325	24.13
-1.00	0.100	0.090	106.30	35499.00	7384.40	834.0356	33800.80	0.84	4.87	-1.425	24.38
-1.10	0.079	0.071	121.50	46376.98	10877.98	975.9292	34776.73	0.86	5.57	-1.525	24.63
-1.20	0.063	0.057	137.80	59655.19	13278.21	946.2595	35722.98	0.89	6.32	-1.625	24.88
-1.30	0.050	0.045	143.50	64692.46	5037.27	285.1455	36008.13	0.89	6.58	-1.725	25.13
-1.40	0.040	0.036	152.40	72965.87	8273.41	372.0115	36380.14	0.90	6.99	-1.825	25.38
-1.50	0.032	0.028	168.90	89620.81	16654.94	594.8611	36975.00	0.92	7.75	-1.925	25.63
-1.60	0.025	0.023	185.70	108336.19	18715.37	530.9719	37505.96	0.93	8.52	-2.025	25.88
-1.70	0.020	0.018	204.10	130868.69	22532.50	517.7886	38013.75	0.94	9.36	-2.125	26.13
-1.80	0.016	0.014	222.00	154830.25	23961.56	428.9329	38442.68	0.95	10.18	-2.225	26.38
-1.90	0.013	0.011	236.70	176013.62	21183.37	301.2104	38743.89	0.96	10.85	-2.325	26.63
-2.00	0.010		250.00	196349.50	20335.87	229.6881	38973.57	0.97	11.46	-2.425	26.88
-∞							40323.00	(1)			∞

PHOTOMETRIC PARAMETERS OF NGC 4762

V-FILTER

Total luminosity	L_T	= 11.20
Total apparent magnitude	m_T	= 10.37
Apparent central surface brightness	μ_o	= 17.33
Major axis at threshold	$2a_m$	= 11.60
Minor axis at threshold	$2b_m$	= 6.03
Major axis at $\mu=25.0$ mag sec ⁻²	$2a(25)$	= 8.64
Luminosity within $\mu=25.0$ mag sec ⁻²	$k(25)$	= 0.89
Gradient of exponential component	$G(a)$	= -0.57
Equivalent gradient of exponential comp....	$G(r^*)$	= -0.44
Equivalent gradient of reduced exp. comp....	$G(\rho)$	= -0.74
Parameters at $k = \frac{1}{4}$:		
Semi-major axis	a_1	= 0.31
Axis ratio	b/a	= 0.37
Equivalent radius	r_1^*	= 0.17
Surface brightness	μ_1	= 18.46
Parameters at $k = \frac{1}{2}$ (effective) :		
Semi-major axis	a_e	= 1.16
Axis ratio	b/a	= 0.21
Equivalent radius	r_e^*	= 0.36
Surface brightness	μ_e	= 20.68
Mean surface brightness	μ_e'	= 10.15
Parameters at $k = \frac{3}{4}$:		
Semi-major axis	a_3	= 2.59
Axis ratio	b/a	= 0.22
Equivalent radius	r_3^*	= 1.08
Surface brightness	μ_3	= 23.38
Concentration indices	$\begin{Bmatrix} C_{21} \\ C_{32} \end{Bmatrix}$	$\begin{matrix} = 2.20 \\ = 2.98 \end{matrix}$